







Antiproton Decelerator (AD)

Only source of slow antiprotons

26 GeV/c PS beam onto Ir target ~30 million antiprotons 5.3 MeV kinetic energy (100 MeV/c)













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ELENA: a boost to the AD physics programme

AD:

p̄ caught in Penning traps using degraders → 99.9% are lost

ELENA:

p at 100 keV at improved beam emittance

all experiments gain a factor 10-100 in trapping efficiency (degrading at low particle energies is more efficient)

"simultaneous" delivery to almost all experiments \rightarrow Gain in total beam time

additional experimental zone

Energy scale (ev)



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BASE/STEP (p̄ in Penning trap), ASACUSA (p̄He) Fundamental properties of the antiproton



ALPHA Spectroscopy of 1S-2S in antihydrogen



ASACUSA, ALPHA Spectroscopy of GS-HFS in antihydrogen



ALPHA, AEgIS, GBAR Test free fall/equivalence principle with antihydrogen

AD community: ~60 research institues/universities - 400 researchers - 5 collaboration (+1 : connection to ISOLDE with the PUMA exp.)



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Gravity with <u>matter</u> scrutinized via different experimental methods

- Universality of free-fall established by Galileo (~450 years go) and Newton
- Weak equivalence principle starting point for Einstein's theory of general relativity (~100 years ago)

Einstein's equivalence principle (EEP) extensively tested experimentally

- WEP





go) and Newton eory of general





But gravity is a peculiar force

Very weak force Lack of consistent quantum treatment

Gravity on antimatter has "never" been directly tested

"Peculiarity" of antimatter :

non detection of primordial antimatter

&

lack of experimental hints for the justification of baryon asymmetry

Need for a free-fall experiment on antimatter



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What implications if antimatter behaves differently than matter in a gravitational field?

• GR and WEP would have to be broken?

validity well tested on matter but not on antimatter accelerating expansion of the universe requires dark energy (composition of the universe) but could it be a sign for the need of revised theory?

matter-matter but not for matter-antimatter)

Example:
$$V = -\frac{Gm_1m_2}{r}(1 \mp a e^{-r/v} + b e)$$



• New forces : scalar or vector mediators would not necessarily invalidate GR (if similar magnitude cancellation for

a: Gravivector, b: Graviscalar

-r/s

– attractive (matter-matter)

+: repulsive: matter-antimatter

matter experiments: |a-b| antimatter: a+b





Existing <u>indirect</u> bounds

GR effect of gravitational redshift

i.e. clocks frequency appear different for an observer in a gravitational potential

cyclotron frequencies of particles are like clocks, so if $\bar{g} \neq g$, the cyclotron frequency (of \bar{p} and p for example) experimentally observed will be different by $(1 - \frac{\bar{g}}{g})\frac{GM}{Rc^2}$

cyclotron frequencies of antiproton and proton measured to ppb precision but : - "arbitrariness" of the definition of the "absolute gravitational potential" (which sets the upper bound on anomalous antimatter gravity)

- assumes CPT invariance

Need for a free-fall experiment on antimatter

an experiment that drop, throw or deflect and measure resulting force (i.e. independent of framework)



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Michael Martin Nieto, T. Goldman, Physics Reports, Volume 205, Issue 5,1991, Pages 221-281, https://doi.org/10.1016/0370-1573(91)90138-C.

Mark Fischler, Joe Lykken, Tom Roberts arXiv.org > hep-th > arXiv:0808.3929v1

Need antimatter! preferentially long-lived, "easily" produced : **p**, **e**⁺ Attempted on those (charged) but too difficult to control stray-fields at a high enough level of precision $(F_G/F_{EM} \sim 10^{-36})$

anti-neutrons?

Cannot be easily cooled like neutrons (at T~1K, v~140 ms⁻¹)







Formation of antihydrogen atoms: several approaches





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Formation of antihydrogen atoms: several approaches

$$+e^+ \rightarrow \bar{H} + e^+$$



$$Ps^{\star} \to \bar{H}^{\star} + e^{-1}$$











Formation of antihydrogen atoms: several approaches

$$+e^+ \rightarrow \bar{H} + e^+$$



$$s^{\star} \to \bar{H}^{\star} + e^{-}$$



Antihydrogen ION !







ALPHA First direct measurement in 2012 (in a magnetic trap!)

CERN



$-65 < g/\bar{g} < 110$

C. Amole et al. Nature Communications 4, 1785 (2013)

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Vertical position of annihilation vertex during release of trapping field



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1d

Now commissioning a VERTICAL TRAP

- increase sensitivity in up/down direction (up to 1.3m trapping range)

- much improved field control

Sign measurement planned rapidly 1% targeted \overline{H} cooling to ~20 mK and advanced magnetometry



Article

Laser cooling of antihydrogen atoms

https://doi.org/10.1038/s41586-021-03289-6

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Check for updates

ARTICLE

https://doi.org/10.1038/s41467-021-26086-1

Sympathetic cooling of positrons to cryogenic temperatures for antihydrogen production



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W. A. Bertsche Phil. Trans. R. Soc. A 2018 376 20170265; DOI: 10.1098/rsta.2017.0265. (2018)





AEGIS : DEFLECTOMETER

S. Aghion et al. Nature Communications 5 (2014) 4538





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Cold \overline{H} production relying on sympathetic cooling of \overline{p} via laser cooled anions or molecules

E. Jordan et al., Phys. Rev. Lett. 115 113001 (2015)







GBAR : USING H+

- will produce first ever \overline{H}^+ ion
- will bring antimatter to the coldest temperature ever achieved (by several orders of magnitude)

Cooling below 1 m/s : Sympathetic cooling of \overline{H}^+





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neutron wavefunction $\psi_p^2(z)$. The vertical axis z provides the length scale for this phenomenon. E_n is the energy of the *n*th quantum state.

14

Study of alternative cooling mechanism in a neutral atom trap



C Malbrunot et al 2022 J. Phys. B: At. Mol. Opt. Phys. 55 044003



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Figure 4. Ratio of trapped atoms with and without THz-stimulated decay ($R_{st/sp}$) as a function of n' for different trap configurations. The stimulated rate is $\Gamma = 4 \times 10^7 \text{ s}^{-1}$. The lines are a guide for the eye.

Adiabatic-like cooling by acting on the internal structure of the atoms at an appropriate place in the trap - increase the phase space density

Based on previous work on fast stimulated deexcitation: e.g. Wolz T, Malbrunot C, Vieille-Grosjean M and Comparat D 2020 Phys. Rev. A 101 043412





VERTICAL TRAP

- increase up/down sensitivity (up to 1.3m trapping range) - much improved field control

Sign measurement planned soon 1% targeted \overline{H} cooling to ~20 mK and advanced magnetometry



Plurality of approaches

H BEAM





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- Sensitivity to ~10 µm deflection
 - needed
- cold antiproton translates in cold
 - **H** thanks to CE mechanism

cooling of \bar{H}^+ - opens new horizons











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- Uniqueness of the physics question addressed
- Calls for a direct measurement
- H is a tool of choice for such a measurement
- Three collaborations at CERN/AD are taking on the challenge
- A sign measurement is expected soon (~2022/2023)
- Diversity of approaches aims at tackling different sensitivities (with different systematics)
- Typical time-scales involved for new experiments and precision measurements are long (typically >10 years)
- Other "gravity" endeavours with antimatter : muonium (μ^+e^-), positronium (e^+e^-) Testing leptonic matter-antimatter systems Mu: Testing systems containing 2nd generation particles!



Summary





Where are the anti-atoms??



Strong baryon asymmetry in the universe

originating from a ~10⁻¹⁰ imbalance

CP violation in the SM is by far not enough to explain this imbalance



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baryon asymmetry:

Comparison of fundamental properties of simple baryonic and anti-baryonic systems at <u>low energy</u> and with <u>high precision</u>



Strong baryon asymmetry in the universe

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baryon asymmetry:

kaon Δm

 10^{-21}

Comparison of fundamental properties of simple baryonic and anti-baryonic relative



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(backups)



Strong baryon asymmetry in the universe

originating from a ~10⁻¹⁰ imbalance

CP violation in the SM is by far not enough to explain this imbalance

systems at <u>low energy</u> and with <u>high precision</u> relative





antimatter & gravity Indirect limits exists Universality of free-fall never tested directly on antimatter



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baryon asymmetry:

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(backups)

Low energy antiprotons for tests of baryon asymmetry

Workhorse Penning trap: В a) Degrading Solenoid - B = 3 Tesla t = 0 s Antiprotons Degrader Cold electron cloud [cooled by Synchrotron Radiation, $\tau \sim 0.4s$] b) Reflecting 99.9% lost 0.1% t = 200 ns Potential c) Trapping t = 500 ns Potential d) Cooling in a trap (beam) t~20 s Potential [through Coulomb interaction] Long trapping times require good vacuum!

BASE experiment: P < 2. 10⁻¹⁸ mbar $\tau(\bar{p}) > 10.2$ years



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Bohr





Some spectroscopy highlights with antihydrogen



In a TRAP:

Investigation of the **FINE STRUCTURE** of antihydrogen

(~10% precision)

Toward antimatter only determination of the antiproton charge radius (together with 1S-2S precision spectroscopy above)!

M. Ahmadi et al., Nature 578, 375–380 (2020)



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In a BEAM:



M. Ahmadi et al. Nature 548, 66–69 (2017)

M. Diermaier et al. Nature Communications 8, 15749 (2017)

$$\nu_{HF} = \frac{16}{3} \mathscr{R}_{y} \alpha^{2} c \left(\frac{m_{\bar{p}}}{m_{\bar{p}} + m_{e^{+}}} \right)^{3} \frac{m_{e^{+}}}{m_{\bar{p}}} \frac{\mu_{e^{+}}}{\mu_{B}} \frac{\mu_{p}}{\mu_{N}} (1 + \delta_{str} + \delta_{QED})$$
$$\Delta \nu(\text{Zemach}) = \nu_{HF} \frac{2Z \alpha m_{e^{+}}}{\pi^{2}} \int \frac{d^{3}p}{p^{4}} \left[\frac{G_{E(\bar{p})}(p^{2})G_{M(\bar{p})}(p^{2})}{1 + \kappa} \right]$$

(backups)

Some highlights on antiprotons





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first measurement more precise for antimatter than for matter









