LiteBIRD

Lite (light) satellite for the studies of B-mode polarization and Inflation from cosmic background Radiation Detection

Josquin Errard

LPNHE 4th of November, 2019



In May 2019, *LiteBIRD* was selected for JAXA's strategic L-class mission!

launch planned in 2027!

selected history/status

- **2008** first LiteBIRD proposal; Basic concept of a focused mission established;
- early 2015 LiteBIRD US MO proposal submitted and granted (July 2015): extended frequency range (HFT) proposed;
- mid-2015 LiteBIRD selected as one of the two candidate projects for JAXA's Large mission; LiteBIRD and SolarSail (more recently known as Okeanos);
- **fall 2015** an 'invitation' letter to the European CMB community by Saku Tsuneta (ISAS);
- mid-2017 LiteBIRD-Europe established;
 first meeting in Cardiff, UK;
- mid-2018 ESA-JAXA CDF study;
 - LiteBIRD-Europe collaboration representatives invited;
- mid/fall 2018 selection for the phase A in Italy and France;
- May 2019 LiteBIRD downselected by JAXA for the L3 large mission with a launch in FY 2027.



Introduction LiteBIRD design LiteBIRD science Challenges Collaboration



Hot Big Bang scenario







T = 2.7 K $\Delta T/T \sim 10^{-3}$

$$d\ell^2 = a^2(t)[1 + 2\zeta(\mathbf{x}, t)][\delta_{ij} + h_{ij}(\mathbf{x}, t)]dx^i dx^j$$

• Fluctuations we observe today in CMB and the matter distribution originate from quantum fluctuations during inflation

Mukhanov & Chibisov (1981) Guth & Pi (1982) Hawking (1982) Starobinsky (1982) Bardeen, Steinhardt & Turner (1983)

observations are already in remarkable agreement with single-field slow-roll inflation:

- super-horizon fluctuation
- adiabaticity
- gaussianity
- n_s < 1

```
but we want gravitational
waves in addition!
```

e.g. *The Best Inflationary Models After Planck* J. Martin, C. Ringeval, R. Trotta, V. Vennin, JCAP, 2014 e.g. *Exploring Cosmic Origins with CORE: Inflation* F. Finelli, M. Bucher et al., JCAP, 2017

$$d\ell^{2} = a^{2}(t)[1 + 2\zeta(\mathbf{x}, t)][\delta_{ij} + h_{ij}(\mathbf{x}, t)]dx^{i}dx^{j}$$
• There should also be ultra long-wavelength gravitational waves generated during inflation
$$\mathbf{h_{ij}} \longrightarrow \mathbf{h_{ij}} \bigoplus \mathbf{h_{$$

Starobinsky (1979)

- LIGO/Virgo detected gravitational waves from binary blackholes, with the wavelength of thousands of kilometers
- But the primordial GW affecting the CMB has a wavelength of billions of light-years!

how to detect them?

CMB POLARIZATION!

WMAP

measurements of r starts being limited by our own galaxy: the Milky Way

tensor-

r<0.

B

Introduction LiteBIRD design LiteBIRD science Challenges Collaboration

CMB B-modes: race to sensitivity

Space and ground: a powerful duo

LiteBIRD JAXA-led focused mission $\sigma(r) < 0.001$ $2 \le \ell \le 200$ focused but still with many byproducts

- best cost-effective way to characterize B-modes
- great synergy with two on-going projects

Why space?

• Superb environment !

- No statistical/systematic uncertainty due to atmosphere
- No limitation on the choice of observing bands (except CO lines), important for foreground separation
- No ground pickup

Rule of thumb: 1,000 detectors in space ~ 100,000 detectors on ground

- Only way to access lowest multipoles w/ $\delta r \sim O(0.001)$
 - Both B-mode bumps need to be observed for the firm confirmation of Cosmic Inflation → we need measurements from space.

Complementarity with ground-based CMB projects

- Foreground information from space will help foreground cleaning for ground CMB data
- High multipole information from ground will help to "delens" space
 CMB data

LiteBIRD operation

orbit: Sun-Earth L2 Lissajous

Three telescopes with TES arrays

- Polarization modulator with a rotating half-wave plate (HWP) for 1/f noise & systematics reduction
- Cryogenic system for 0.1K base temperature

Crossed Dragone

- Aperture diameter: 400 mm
- Angular resolution: 20 -70 arcmin.
- Freq. coverage: 34 161GHz
- Field of view: 20 deg x 10 deg
- F#3.0 & crossed angle of 90 degree
- All 5K parts are made of Aluminum → less than 150 kg
- New mirror design (anamorphic aspherical surfaces)

Three telescopes with TES arrays

- Polarization modulator with a rotating half-wave plate (HWP) for 1/f noise & systematics reduction
- Cryogenic system for 0.1K base temperature

Two F/2.3 refractive telescopes:

- 89-270 GHz
- 238-448 GHz

Apertures:

- 30mm
- 20mm

FoV: Φ 20mm

- Transmissive metal-mesh HWP
- Silicon lenses

- Three telescopes with TES arrays
- Polarization modulator with a rotating half-wave plate (HWP) for 1/f noise & systematics reduction
- Cryogenic system for 0.1K base temperature

Superconducting magnetic bearing system operational in a 4K cryostat.

We observed the stable rotation at cryogenic temperature (<10K).

- Three telescopes with TES arrays
- Polarization modulator with a rotating half-wave plate (HWP) for 1/f noise & systematics reduction
- Cryogenic system for 0.1K base temperature

Introduction LiteBIRD design LiteBIRD science Challenges Collaboration

LiteBIRD science goals

Full success:

- total uncertainty δr < 0.001 (for r=0)
- > 5 σ observation for each bump (for r \ge 0.01)

Rationale

- Large discovery potential for 0.005 < r < 0.05
- Simplest and well-motivated R+R² "Starobinsky" model will be tested.
- Clean sweep of single-field models with characteristic field variation scale of inflaton potential greater than M_{pl} (A. Linde, JCAP 1702 (2017) no.02, 006
LiteBIRD science goals

Full success:

- total uncertainty δr < 0.001 (for r=0)
- > 5 σ observation for each bump (for r \geq 0.01)



Rationale

- Large discovery potential for 0.005 < r < 0.05
- Simplest and well-motivated R+R² "Starobinsky" model will be tested.
- Clean sweep of single-field models with characteristic field variation scale of inflaton potential greater than M_{pl} (A. Linde, JCAP 1702 (2017) no.02, 006

LiteBIRD science goals

Full success:

- total uncertainty δr < 0.001 (for r=0)
- > 5 σ observation for each bump (for r \geq 0.01)



Statistical uncertainty includesforeground cleaning residuals

- lensing B-mode power
- 1/f noise

Systematic uncertainty includes

- Bias from 1/f noise
- Polarization efficiency & knowledge
- Disturbance to instrument
- Off-boresight pick up
- Calibration accuracy

















characterize reionization history



→ better *τ* and sum of neutrino masses
→ crucial information for many other surveys







elucidating anomalies with polarization



Fig. 1. The *Planck* 2015 temperature power spectrum. At multipoles $\ell \ge 30$ we show the maximum likelihood frequency averaged temperature spectrum computed from the Plik cross-half-mission likelihood with foreground and other nuisance parameters determined from the MCMC analysis of the base Λ CDM cosmology. In the multipole range $2 \le \ell \le 29$, we plot the power spectrum estimates from the Commander component-separation algorithm computed over 94% of the sky. The best-fit base Λ CDM theoretical spectrum fitted to the *Planck* TT+lowP likelihood is plotted in the upper panel. Residuals with respect to this model are shown in the lower panel. The error bars show $\pm 1 \sigma$ uncertainties.

+ Integrated Sachs Wolfe effect complementary to density measurements

 10^{3}











Introduction LiteBIRD design LiteBIRD science Challenges Collaboration

LiteBIRD — foregrounds cleaning



synchrotron and dust polarized emissions follow the galactic magnetic field



synchrotron and dust polarized emissions follow the galactic magnetic field



intensity @ 30GHz + B-field from polarization Planck 2015 results. X. Diffuse component separation: Foreground maps The Planck collaboration, A&A, 2015



amplitude of galactic foregrounds for different frequencies



removing galactic foregrounds

analogy interlude



removing galactic foregrounds



 $d_{\nu_0} = a_0 \operatorname{CMB} + b_0 \operatorname{dust} + n_{\nu_0}$ $d_{\nu_1} = a_1 \operatorname{CMB} + b_1 \operatorname{dust} + n_{\nu_1}$

removing one or several

- components increase the noise variance in the final "clean" component
- misestimating a spectrum leaks components to the "clean" component (can be statistical or systematic misestimation)

$$d_{\nu_0}b_1 - d_{\nu_1}b_0 = CMB \ (b_1a_0 - b_0a_1) + n_{\nu_0}b_1 - n_{\nu_1}b_0$$

boosted variance

$$\sigma_{\rm CMB}^2 = \frac{\sigma_{\nu_0}^2 b_1^2 + \sigma_{\nu_1}^2 b_0^2}{\left(b_1 a_0 - b_0 a_1\right)^2}$$

statistical/systematic residuals in the cleaned signal

 $\delta \text{CMB} \propto \delta b_1 \left(\alpha \, d_{\nu_0} + \beta \, d_{\nu_1} \right)$





LiteBIRD — systematics and calibration

- One of the largest study groups in LiteBIRD
- Systematic approach for systematic uncertainties
 - List systematic error items 14 categories, 70 items listed
 - Assign each item $\sigma(r)_{sys} < 5.7 \times 10^{-6}$ as the budget (1% of total budget for systematic error)
 - Derive a requirement for each item, define method (incl. calibration methods) and estimate $\sigma(r)_{\text{sys}}$
 - Assign special budget allocations for outstanding items
 - Sum each contribution at the map level to estimate total $\sigma(r)_{sys}$ (some studies use TOD basis) to take positive correlations into account
 - Iterate procedure
- Example: studies of systematic errors due to HWP imperfection
 - Mueller matrix from RCWA simulations of electromagnetic wave propagation through realistic HWP for different frequencies and incident angles
 - 4f component from $M_{\text{IQ}},\,M_{\text{IU}}\,{\sim}10^{\text{-4}}$ in the worst case
 - Obtain leakage maps and BB power to estimate $\sigma(r)_{\text{sys}}$



Introduction LiteBIRD design LiteBIRD science Challenges Collaboration

LiteBIRD Joint Study Group

About 200 researchers from Japan, North America & Europe Team experiences: CMB exp., X-ray satellites, other large proj. (HEP, ALMA etc.)





LiteBIRD is the next-generation CMB satellite selected by JAXA as a Strategic Large Mission to be launched in 2027

MFT and HFT (100 - 402GHz) - 66

4700 multichroic TES detectors

50x Planck sensitivity on large angular scales

15 frequency bands $40 \le v \le 402$ GHz



telescopes + 3 instruments rotating half-wave plates year observation at L2





Signals Imply a 'Big Bang' Universe



Horn antenna, used in space exploration, at the Bell Laboratories in Holmdel, N. J.

By WALTER SULLIVAN

Scientists at the Bell Telephone Laboratories have observed what a group at Princeton University believes may be remnants of an explosion that gave birth to the universe.

These remnants are thought to have originated in the burst of light from that cataclysmic event.

Such a primordial explosion is embodied in the "big bang" theory of the universe. It seeks to explain the observation that virtually all distant galaxies are flying away from the earth. Their motion implies that they all originated at a single point 10 or 15 billion years ago.

The Bell observations, made by Drs. Arno A. Penzias and Robert W. Wilson from a hilltop in Holmdel, N. J., were of radio waves that appear to be flying in all directions through the universe. Since radio waves and light waves are identical, except for their wavelength, these are thought to be remnants of light waves from the primordial flash.

The waves were stretched into radio waves by the vast expansion of the universe that has occurred since the explosion and release of the waves from the expanding gas cloud born of the fireball. In what may prove to be one of the most remarkable coincidences in scientific history, the existence of such waves was predicted at

Continued on Page 18, Column 1

Ehe New Hork Eimes Published: May 21, 1965 Copyright © The New York Times





the hot Big Bang also predicts the good proportions of elements in the Universe!

Density of Ordinary Matter (Relative to Photons)





Pol 2

- Full success
- Extra success (see previous page)
- Characterization of B-mode and search for sources fields (e.g scale-invariance, non-Gaussianity, parity violation)
- Power spectrum features in polarization
- Large-scale E mode and its implications for reionization history and the neutrino mass
- Cosmic birefringence
- SZ effect (thermal and relativistic correction)
- Elucidating anomalies
- Galactic science