Exploring Time Domain Multi-Messenger Astronomy through the Virtual Observatory

Ada Nebot & the CDS team

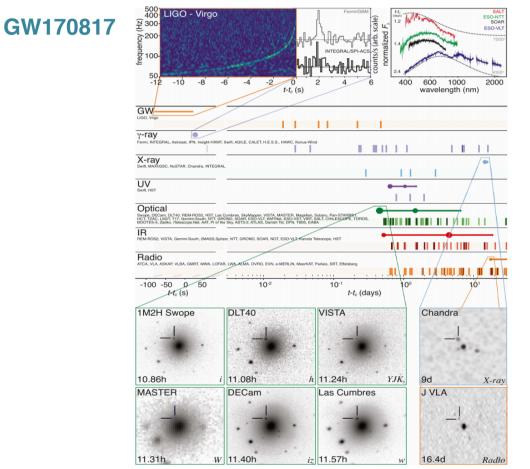
TS2020-III 26 September 2019

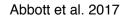


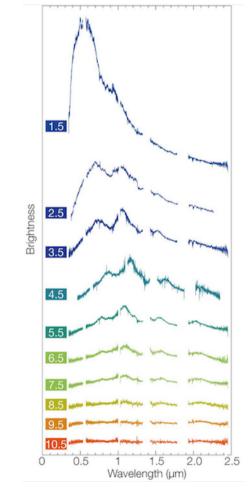
Time Domain Multi-messenger Astronomy

Abbott et al.

THE ASTROPHYSICAL JOURNAL LETTERS, 848:L12 (59pp), 2017 October 20







X-shooter spectra in the kilonova in NGC 4993 over 12 days. Image credit: ESO/Pian et al./Smartt & ePESSTO.

To characterise and classify sources...

- Multi-wavelength / messenger approach is (sometimes) needed
- Follow-up observations and reaction time for that can be crucial
- Visualisation & navigation thought the data
- Coordination & transmission of information

The VO (IVOA) should match user's needs So, what is available through the VO?

What is the VO?

• Astronomical datasets, tools, services should work seamlessly together

What is the IVOA?

- An organisation that debates and agrees the technical standards that are needed to make the VO possible
- A focal point for VO aspirations, a framework for discussing and sharing VO ideas and technology
- Promoting and publicising the VO

Who is the IVOA?

- 6 Working Groups, 7 Interest Groups
 - There is a Time Domain Interest Group
 - Chair & Vice-chair: A. Nebot & D. Morris
 - Completely open to participation

How to join the IVOA?

- 2 interoperability meetings per year
 - Next IVOA meeting in Groningen 11-13 Oct.
 - following ADASS 6-10 oct.
 - https://www.adass2019.nl/ivoa/ivoa-participants/
 - Register to email lists for discussion of topics
 - Asking me directly can work too <u>ada.nebot@asto.unistra.fr</u>

http://ivoa.net/



Who is the VO for?

- Research astronomers
- Data Centres and Archives
- Software developers
- Educators

. . .

Idea of the VO?

- In a seamless way for the user:
 - Data discovery & access
 - Visualisation & analysis
 - Through Services & tools
- Research astronomers Time Domain Astronomers

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- Coordination & transmission of information
- Multi-wavelength/messenger
- Combining data from missions covering different wavelength ranges
 - Source identification
 - Cross-matching techniques

Minimum information about objects

Which objects around this area are already known and have a classification?

Give me a minimum information about this object / list of objects (e.g. it's a Galaxy at redshift z)

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Overview Cross-IDs (6	65) Coordinates (47) Re	edshifts (29) Distances (101)	Classifications (117)	Notes (48) Diameters (8))
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UCAC4 354-117192 🏶 WR 110 2XMM J180756.9-192356 Gaia DR2 40951252208078

Identifiers (22) :

An access of full data is available using the icon Vizier near the identifier of the catalogue

HD 165688	Hen 3-1594	MR 83
ALS 4678	HIC 88828 🕮	PPN 718808
BD-19 4854 🅮	HIP 88828 🕮	SSTGLMC G010.8000+00.3943 🏶
CPD-19 6469 🕮	JP11 2931	TTC 6259-2666-1 🏶
GEN# +1.00165688	LS 4678	UBV 15399
GSC 06259-02666 🏶	2MASS J18075695-1923568	UCAC2 24414003

References (137 between 1850 and 2019) (Total 137) References (137 between 1850 and 2019) (Total 137) Simbad bibliographic survey began in 1850 for stars (at least brig Follow new references on this object

Reference summaries :

from: 1850 to: \$currentYea

Display or select by : (not exhaustive, explanation here) In table Title/Abstract/Keyword Score

Collections of Measurements

distance : 2 PM : 3 PLX : 3 MK : 5 display selected measurements display all measurements clear

Cross-matching — A key point

- CDS dev. F. X. Pineau et al.
- Positional cross-correlation of sources in 2 tables (VizieR tables, simbad, user uploaded lists)
- Result in different formats (VOTable, CSV or ASCII)
- Programatic access too (http API)
- New developments for a multi-catalogue cross-match
- Available through TOPCAT and stilts too

CDS X-Match Service X-match Tables management Documentation Select below the two tables to cross-match Image: Then, choose cross-match method and sky area in options. Finally, click on Begin the X-Match to launch the computation.	Login P
Choose tables to cross-match	
e.g. VII/260/dr7qso, or select in list VideR SIMBAD My store © Show options Begin the X-Match Visualize and manage your cross-match jobs	Table List Current Table Properties Label: anonymous1541509785078.xml Location: /Users/angm/Downloads/anonymous1541509785078.xml Name: anonymous1541509785078.xml Rows: 128 Columns: 12 Sort Order: C Row Subset: All
List of X-match jobs Table 1 Table 2 Options Begin Status Actions	Activation Actions: 0 / 0
laoie 1 1aoie 2 Options Begin Status Actions No job in list	SAMP- Messages: O Clients: • 🏟 🗋

Cross-matching

Positional cross-match performance, radius 5"

Table 1	Table 2	Computation time	Result size	Total time	
SDSS DR9 469M rows	2MASS 470M rows	3 min	7 min	19 GB	10 min
2MASS 470M	GAIA-DR1 1.1 billion	16 min	65 min	193 GB	81 min
Tycho-2 2M	SIMBAD 8M	6 sec	25 sec	1 GB	35 sec
List of 40k positions	SIMBAD 8M	1 second	4 seconds	10 MB	5 sec

Under dev.: add the time as a possible information to cross-matches

To characterise and classify sources...

- Multi-wavelength/messenger approach is (sometimes) needed
- Follow-up observations & reaction time for that can be crucial
- Visualisation & navigation thought the data
- Coordination & transmission of information
- Follow-up observations
 - Transmission of events: VOEvent, IVOA Standard
 - VOEvent Standard for Fast Radio Bursts FRBs, Petroff et al. 2017
 - Damien Dornic yesterday mentioned: VOEvent Standard for Neutrino ! :)
 - VOEvent Standard for other science / mission specific field?
 - Planning observations: visibility, available telescope time 2 IVOA Standards in process

Planning observations: visibility services

	ESO - Reaching New Heights in Astronomy EVer Portal Contact Site Map Search Col get 220 Natures * Central College * Ce	Target Name M31 SIMBAD Loc	P SIMBAD OR NED AGAIN, OR RUN 1	THE VISIBILITY CI	IN MULTI-TARGET VISIBILITY		About ING	NEWTON GROUP OF TELESCOPES
Science Users Information Observing Facilities Future Facilities and Development Observing with ESO Telescopes	Object Observability See also: Coject Observability - Atmasses - Daily Amanac - Sily Calendar This for provides object observability tables based on site, object coordinates and observing period. Times are given for the local time, including daylight saving times when applicable.	SIMBAD LOOKUP I					altitude agains a particular nig	ogram that shows the observability of objects in various ways: either you can plot t time for a particular night (Staraft), or plot the path of your objects across the sky for Int (Startack), or plot how altitude changes over a year (Starobs), or get a table with ving date for each object (Starmult). For further information, click on the "help" button of the page.
Policies and Procedures	Select site, object coordinates and observing period; then press Compute.		s, complete the "Visibility Details" and Submit					
Telescope Time Allocation Phase 1 Proposals	More detailed information is provided is a separate document Notes for Skycalc by John Thorstensen.	TARGET DETAILS					Mode	Staralt
Phase 2 Preparation		Target Name M31	Target no Decimal degrees or HH:MM:SS.S (eg: 13:30:52.5)	ame or identifier for output (e	g; Abell 1750)		Night	12 V October V 2017 V or date when the local night starts. Staralt, Startrack only.
Phase 3 Public Surveys Observing Tools and Services ESO ETC's	Site: [Paratal Observatory (NT) • Date: drywy mm dd): From 20127 10 20 Tor 2017 11 15 Obsect: Coordinates (J2000) Obsect: Coordinates (J2000) Obsect: Coordinates (J2000)	Dec +41:16:07.50 VISIBILITY DETAILS Select either	Decimal degrees or HH MM ISS S (egr. 13:30:52.5) Decimal degrees or DD MM ISS S (egr. 101:50:27.0 obvition[3369] default is A017 revolution range: 338				Observatory	La Sill observatory (Chile) Select one above or specify your own site with this format Longitude("feast) Latitude(") Altitude(metres) UTC offset(hours) EX.: 289.2767 -30.2283 2725 -4
Instrumental Characteristics Archives and Catalogues Calendars and Calculators Weather Images Astroclimatology Meleo Information	Ar 10 52 34.5 Dec69 45 22 Compute StyCaic provided by courtery of John Thorstensen, Datmouth College John Thorstensen@dartmouth.adu	Last Rev or Date Range From Da To Date Minimum visibility 5000 Submit					Coordinates	Formats can be any of these: name hh mm ss tdd mm ss name dhd.ddd dd.ddd name ddd.ddd dd.ddd name ddd.ddd dd.ddd name thist be a single word with no dots, avoid using single numbers. Every entry mussibe in the same format, do not use different formats with different entries. We roministry of the same format in the same format is the same format is the same format is the same format. So the same same same same same same same sam
Visiting Astronomers Science Software Data Handling and Products Science Archive Facility Science Activities			VIEWING CO	17 TARGET VISIBILITY	N		Altitudes, La moon (dashed): Coordinates: s ^h 4 ^m + 18°15 Illumination: 42% Quarter: 4	Silla Observatory 289,2700E - 29,2567N, 2347 m above sea level
See also Object Observability - A	rmasses - Daily Almanac - Ephemerides		All four	2" x 2" 70" - 110" 42"				70°
Observability for 05 2 Paranal Observatory (VLT)	23 34.5 -69 45 22 The ESO Sky Calendar Tool		Min Via (5) Start Ort	CRITERIA FOR ALL TARGETS St End Orbit Start Date End Da 3551 D1-May-2018 29-Apr-2			Numbers below or are Moon distance (in degrees) at th corresponding times.	
RA & dec: 5 23 34.5, -6 Site long⪫: +4 41 36.8 (h.1		Targets that are or	Iy visible for a small fraction of an orbit are only visible at the start or end of a re	volution (see columns Visibility Start E	nd Phase) and therefore have a higher likelihood for increased bac	okground radiation.	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	
(2) natural center of night, an nighttime hours during which of	ase, be age ad acc. at (1) enc. bulligh, d(1) avoing bullight that coust number of feet is at sec. less than 3, 2, and 1.5. by our allithde - scile degrees.	Target Name RA Dec M31 10.6647 41.2607 Rev. yyyyemm dd hicreni	Vis. Window Vis. End Dutation (s) (yyyy-stm-4d htcmm)	Rounded Vis. (s)	Visibility Visibility Solar Aspect Start Phase End Phase Angle(*)	Mean Astronomical Position Angle(*)		20 ⁰ 100 8
2017 Nov 3 F -6 52 3.1	cent norm night hrughec.s: Hi sec.: HA sec.: G -Q -Q -L-5 -2.48 - L-6 -HI -L-5 -B-6 -B-3-3 -1.49 -L-5 - 42.67 -L-5 -7.8 -6.7 -3.8	3397 2014-06-28 02:56 3396 2014-06-28 02:56 3399 2014-06-20 12:42 3400 2014-07-03 12:35 3401 2014-07-05 12:29 3402 2014-07-05 12:29 3402 2014-07-05 12:29	27056 2914-04-29 10 29 781-06 2016-06-30 10 31 78065 2016-07-20 10 23 77806 2016-07-20 10 23 77804 2016-07-40 10 4 77804 2016-07-80 10 6 77775 2016-07-80 10 55 78052 2016-07-10 10 00	25000 75000 75000 75000 75000 75000 75000	0.76 0.82 71.3 0.47 0.92 72.6 0.47 0.92 74.2 0.47 0.82 75.7 0.47 0.82 75.7 0.47 0.82 77.3 0.47 0.92 78.8 0.47 0.93 80.4	742 728 717 705 605 665 674		
<u>SkyCalc</u> provided by courtesy of J	ohn Thorstensen, Dartmouth College. <u>John Thorstensen@dartmouth.edu</u>	3404 2018-07-11 12:07	78345 2016-07-12 09-53	75000	0.47 0.93 82.0	66.3		Mean Solar Zone Time, starting night 12 10 2017 Processed: 2017/10/12 at 10:21:34 UT. Heas Newton Group of Telescopes, La Patma.

Different services have different inputs / outputs Facilitate the work by having same inputs / outputs

Object Visibility Simple Access Protocol Aitor Ibarra, Richard Saxton, Jesús Salgado et al. 2019 http://www.ivoa.net/documents/ObjVisSAP/index.html

Planning observations: coordination of observations



Schedule for revolution 1872

(this list is also available in csv-format, click here to download)

Rev	Start time (UTC)	End time (UTC)	Exp. time (s)	Target	Ra (J2000)	Dec (J2000)	Pattern	PI	Proposal	Observation	Notes
1872	2017-10-10 13:29:15	2017-10-10 17:10:51	12600	Gal. Bulge region	17:45:36.00	-28:56:00.0	<u>HEX</u>	Erik Kuulkers	1420001	1420001 / 0022	Public
1872	2017-10-10 17:13:34	2017-10-11 07:55:55	50000	Galactic Center	17:52:11.21	-25:21:49.7	5x5 Seq	Joern Wilms	1420009	1420009 / 0011	
1872	2017-10-11 08:16:46	2017-10-11 11:58:32	12600	Galaxy (I=0, b=0)	17:42:23.76	-29:38:02.4	HEX	Rashid Sunyaev	1420021	1420021 / 0039	
1872	2017-10-11 12:26:36	2017-10-11 12:56:36	1800	Galaxy (I=0, b=-30)	20:02:16.80	-41:20:31.2	HEX	Rashid Sunyaev	1420021	1420021 / 0038	
1872	2017-10-11 13:27:21	2017-10-11 14:29:17	3600	Galaxy (I=0, b=-30)	19:59:40.80	-41:05:16.8	HEX	Rashid Sunyaev	1420021	1420021 / 0040	
1872	2017-10-11 15:00:12	2017-10-11 17:38:07	9000	Galaxy (I=0, b=-30)	19:59:40.80	-41:05:16.8	HEX	Rashid Sunyaev	1420021	1420021 / 0040	
1872	2017-10-11 18:41:00	2017-10-12 08:01:56	45000	GRS 1915+105	19:15:11.79	+10:56:45.7	5x5 Seq	Jerome Rodriguez	1420029	1420029 / 0008	
1872	2017-10-12 09:06:18	2017-10-12 12:47:54	12600	Galaxy (I=0, b=0)	17:50:46.80	-28:55:30.0	HEX	Rashid Sunyaev	1420021	1420021 / 0041	
1872	2017-10-12 13:16:06	2017-10-12 14:49:58	5400	Galaxy (I=0, b=-30)	20:07:12.96	-40:00:10.8	HEX	Rashid Sunyaev	1420021	1420021 / 0042	

	10	Cesa
Charle Terrer Cabadada		

XMM-NEWTON SHORT-TERM SCHEDULE

```
The Short-term Schedule gives an overview of scheduled observations covering the time range from the past week until the uncoming ~2.4 weeks
```

Background: The planning and scheduling procedure is described in Sect. 8.2 of the Policies and Procedures. In addition, the process of scheduling XMM-Newton observations is described in A guided tour to the scheduling of an XMM-Newton orbit.

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where years, cannot use the residue instanting (price), Cases notice instanting (price), priority controlling to provide any price of the stanting of the price o
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Caveats: The scheduling of an XMM-Newton revolution may have to be revised (see Sects. 8.2, 8.3, and 5.2.2 of the Policies and Procedures). Contingencies of a

ate frequency: Every 8 hours or when the schedule is updated (new revolution planned or any existing updated). The latest available wention can be viewed at

ist updated on: 2017-10-10 12:42:00 UT (Current Rev =

Revn #	Obs Id.	Target Name	RA hh:mm:ss	DEC dd:mm:ss	PA ddd.dd	UTC Obs Start yyyy-mm- dd hh:mm:ss	UTC Obs End yyyy-mm- dd hh:mm:ss	Prime Instr.	PN Dur Ks	MOS1 Dur. Ks	MOS2 Dur. Ks	RGS1 Dur. Ks	RGS2 Dur. Ks	OM Dur. Ks	PI
3276	0805150401	ESO 018-G009	08:24:07	-77:46:57	88.63	2017-10-29 19:34:26	2017-10-30 00:54:26	EPIC	16.7	18.1	18.1	18.2	18.2	18.0	Peter Boorman
3276	0801870801	HD 81809	09:27:46	-06:04:17	92.00	2017-10-29 15:00:13	2017-10-29 18:20:13	EPIC	9.5	10.9	10.9	11.0	11.0	10.8	Fabio Favata
3276	0561381201	zeta Puppis	08:03:40	-40:00:36	112.00	2017-10-29 01:21:41	2017-10-29 14:08:21	RGS	44.5	44.9	44.9	45.0	45.0	37.3	Fred Janser XMM- Newton MN
3276	0803950401	SDSS 102714.77+35431	10:27:14	+35:43:17	119.93	2017-10-28 15:44:35	2017-10-28 23:31:15	EPIC	25.5	26.9	26.9	27.0	27.0	26.8	Guido Risaliti
3276	0803240201	J072637.95+394558.0	07:26:37	+39:45:58	91.37	2017-10-28 11:02:32	2017-10-28 14:55:52	EPIC	11.5	12.9	12.9	13.0	13.0	12.9	Nathan Secrest
3275	0801990201	0457-6739	04:57:33	-67:39:06	136.67	2017-10-27 12:22:47	2017-10-28 01:07:47	EPIC	43.4	44.8	44.8	44.9	44.9	43.7	Patrick Kavanagh
3275	0801990401	0449-6903	04:49:34	-69:03:34	138.62	2017-10-26 23:32:47	2017-10-27 12:02:47	EPIC	42.5	43.9	43.9	44.0	44.0	42.8	Patrick Kavanagh
3275	0803952601	SDSS	08:26:19	+31:48:48	101.78	2017-10-26	2017-10-26	EPIC	36.0	37.4	37.4	37.5	37.5	37.3	Guido



Observing schedules

Short Range Observatory Schedule Download

This is the confirmed schedule of NuCRA observations. This sequence of observations has been uploaded to the opaccent and will execute automonady unless interrupted by a new schedule. They of objectivity, or instrument and spacent's amounted, built own various time ranges depending on the exposure time goal of the observations, but will usually be for a period of the least one week. The times reports have ne the start and end of the on-target grant of glavet material three starts the start of the silver to could also and the SAA passage time where detector background is increased. The end time of the observation is the start of the silver to the net target. Researce anime the NuCRAA Act norm Timiting (ATT) for the (or g) part observation.

obs_start	obs_end	sequenceID	Name	J2000_RA	J2000_Dec	Exp	Notes
2017:281:19:05:02	2017:283:00:30:00	90201021006	Kepler	262.671620	-21.491957	60.6	DDT
2017:283:01:11:23	2017:283:02:40:00	90311211001	Sol_17282_AR2683_POS11	195.15715	-6.38520	3.4	ToO
2017:283:02:40:32	2017:283:04:20:00	90311212001	Sol_17282_AR2683_POS12	195.21879	-6.41062	3.4	ToO
2017:283:04:20:32	2017:283:05:50:00	90311213001	Sol_17282_AR2683_POS13	195.28046	-6.43604	3.4	ToO
2017:283:06:55:11	2017:284:09:20:00	60376001002	2MASXJ19301380p3410495	292.557500	34.180500	55.3	Extragalactic Legacy Survey
2017:284:09:45:09	2017:284:20:35:00	60360008002	SDSSJ152132d21p391206d9	230.3874232	39.2007671	22.0	Extragalactic Legacy Survey
2017:284:21:10:03	2017:285:21:00:00	90301320002	NGC_6440	267.218083	-20.358944	49.5	ToO
2017:285:21:20:06	2017:286:08:20:00	30302020004	GRS_1915p105	288.79813	10.94578	21.9	(2/4) coordinated with XMM and VLT
2017:286:08:35:06	2017:286:19:30:00	60160701002	2MASXJ18560128p1538059	284.00210000	15.63200000	23.3	BAT AGN
2017:286:20:05:11	2017:287:15:05:00	60376007002	UGC06728	176.316800	79.681500	61.4	Extragalactic Legacy Survey
2017:287:15:50:11	2017:288:03:20:00	60368001002	NGC_1144	43.80083	-0.18361	22.0	
2017:288:04:05:09	2017:288:23:00:00	60301004002	ESO_103m35	279.58458	-65.4275	50.3	
2017:288:23:30:08	2017:290:05:45:00	30301026002	AX_J1841d0m0536	280.25179	-5.59625	59.7	phase constrained
2017:290:06:00:04	2017:290:17:00:00	60160670002	2E1739d1m1210	265.47600000	-12.19700000	23.5	BAT AGN
2017:290:17:15:01	2017:291:04:20:00	30363001002	GX_3p1	266.98333	-26.56361	21.8	

Long Range Observatory Schedule	Download
This is the latest NuSTAR long-term schedule. Observ	vations have been sorted into one-week intervals, taking into account Sun, M
exposure time, and other constraints. So the date is t	the Monday of the week in which the observation is scheduled to begin.

	uling Unit UT End UT	SU Id	Principal Investigat	Exp #	Target	Science Instrume	Mode	Apertures	Spectral Elements	Exposure Time(sec)	08	AL	EX
2017.289 02:3	:56 03:08:18	14518F2	Colimowski	F2-001	BIAS	ACS/WFC	ACCUM	WFC	F502N	0.00	22	01	01
2017.289 02:3	:56 03:08:18	14518F2	Golimowski	F2-002	DARK	ACS/WFC	ACCUM	WFC	F660N F502N F660N	1000.50	F2		02

ACS/WEC ACCUM WEC

NIRVII G230L G230L G230L G130M G130M G130M G130M G130M NIRVII F373M NIRVII F373M NIRVII F373M NIRVII F373M NIRVII F373M NIRVII NIRVII

D9-Oct-2017 18:48:29 --- Preliminary EST Observing Timeline Report for SMS: 172888A4 ---SMS Start: 2017.288:22:10:00 (15-OCT-2017 22:10:00), End: 2017.296:00:00:00 (23-OCT-2017)

Scheduling Unit Principal Secin UT End UT SU Id Investigat Exp # Target

17.289 02:09:22 02:38:56 14518F1 Golimowski F1-001 DARK

12:09:22 02:38:56 14518F1 Colimowski F1-002 DARK

What object has been (or will be) observed when and in which wavelength?

Observation Locator Table Access Protocol, Aitor Ibarra, Jesús Salgado et al. 2019

http://www.ivoa.net/documents/ObsLocTAP/20190227/index.html

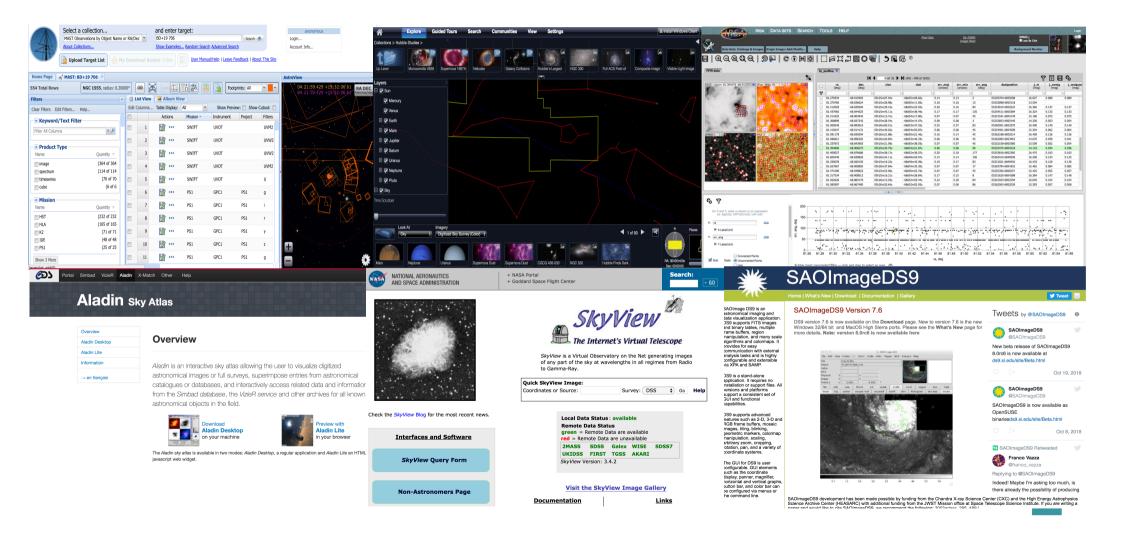
To characterise and classify sources...

- Multi-wavelength/messenger approach is (sometimes) needed
- Follow-up observations & reaction time for that can be crucial
- Visualisation & navigation thought the data
- Coordination & transmission of information

• Visualisation & navigation

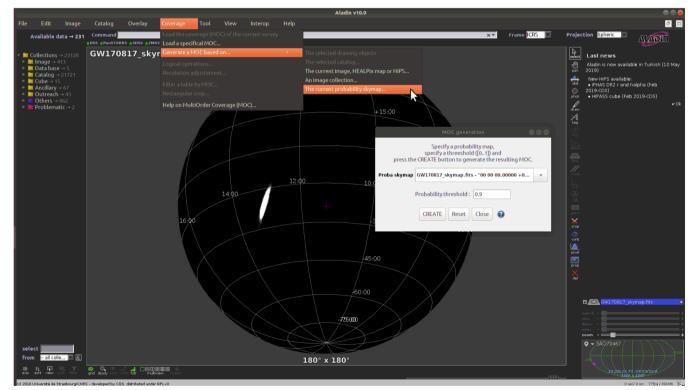
- sequences of images, spectra, photometry, positions, ... and all interoperable
- 🗭 tools

Visualisation of the sky



Visualisation of the sky in Aladin

- GW localisation in the sky in Aladin using MOC IVOA standard
- Background image can be DSS, 2MASS, WISE, XMM, Fermi,...
- We can overlay catalogues of interest, query Simbad, ... by region



https://emfollow.docs.ligo.org/userguide/tutorial/aladin.html

Visualisation of the sky in AladinLite

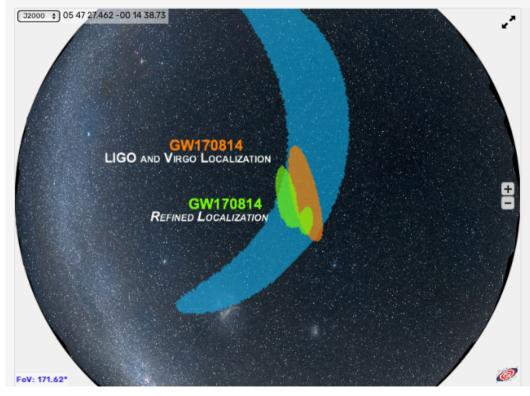
The CDS is very pleased that our Aladin and Aladin-Lite tools are being used to visualise the sky locations of GW170814, and that the images are on Astronomy Picture of the Day and the New York Times.

www.virgo-gw.eu/skymap.html



apod.nasa.gov/apod/astropix.html

www.nytimes.com/2017/09/27/science/black-holes-collision-ligo-virgo.html



Using the skymap

Click on the various options below to display information relating to each detection.

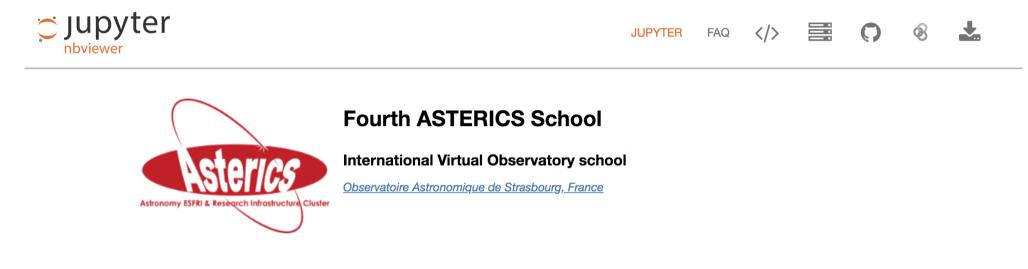
Detection	Sky localisation	Label	Pop-up info
GW170814 - L1/H1 only	×.		•
GW170814 - L1/H1/V1			
GW170814 - refined skymap			-
GW150914			
GW151226			
GW170104	-	-	-

Backgrounds

If you want to see the extension of these sky regions through the constellations you can select an artistic background image **Constellations.**

Visualisation of the sky in AladinLite

Tutorial on usage of VO tools for EM follow-up of GWs



Electromagnetic follow-up of gravitational-wave events

by G. Greco giuseppe.greco@uniurb.it, E. Chassande-Mottin ecm@apc.univ-paris7.fr and M.Branchesi marica.branchesi@gmail.com and many others

The tutorial focuses on some basic strategies for working with gravitational-wave sky localization maps in the context of electromagnetic follow-up activities. Here we propose the usage of <u>Aladin</u>, <u>TOPCAT</u> and <u>GWsky</u>. The following main topics are addressed.

- 1. Gravitational-Wave sky localization map: visualization and tiling
- 2. Access to existing catalogs using the Multi-Order Coverage map (MOC)
- 3. Planning for EM follow-up observations

Visualisation of the sky in AladinLite



LIGO Scientific Collaboration



User Guide

Primer on public alerts for astronomers from the LIGO and Virgo gravitational-wave observatories.

Navigation

Getting Started Checklist Observing Capabilities Data Analysis Alert Contents Sample Code • Prerequisites

- Receiving GCNs
- Working with Sky Maps
- Multi-Order Sky Maps (For Advanced Users)

Sky Map Visualizations and Credible Regions in Aladin

In this section, we demonstrate working with gravitational-wave sky localizations in <u>Aladin Desktop</u>. The following main topics are addressed.

- MOC and GW Sky Localizations
- Running Aladin Desktop
- Loading a GW Sky Localization
- Building a Credible Region
- Area Within a Credible Region

MOC and GW Sky Localizations

The enclosed area within a given probability level contour of a GW sky map can be effectively described with a Multi-Order Coverage (MOC) map [1]. MOC is a standard of the Virtual Observatory which provides a representation of arbitrary regions on the unit sphere using the HEALPix sky tessellation.

 $\leftarrow \underline{\text{Additional Tools}} \mid \underline{\text{Change Log}} \rightarrow$

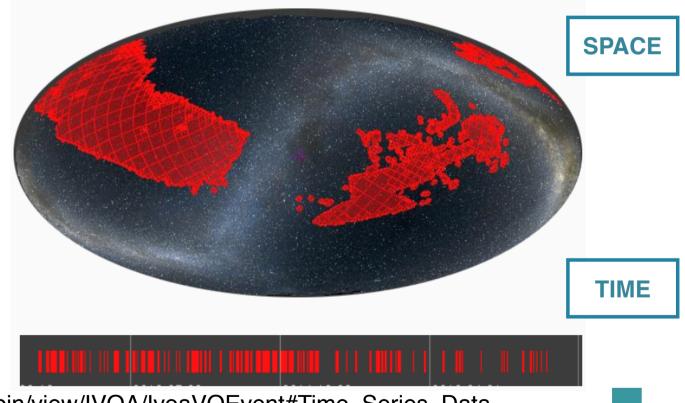
Visualisation of sky coverages

- Coverage maps based on MOC IVOA standard are created from positions
 - We know where but we don't know when!
 - Need to add the time dimension

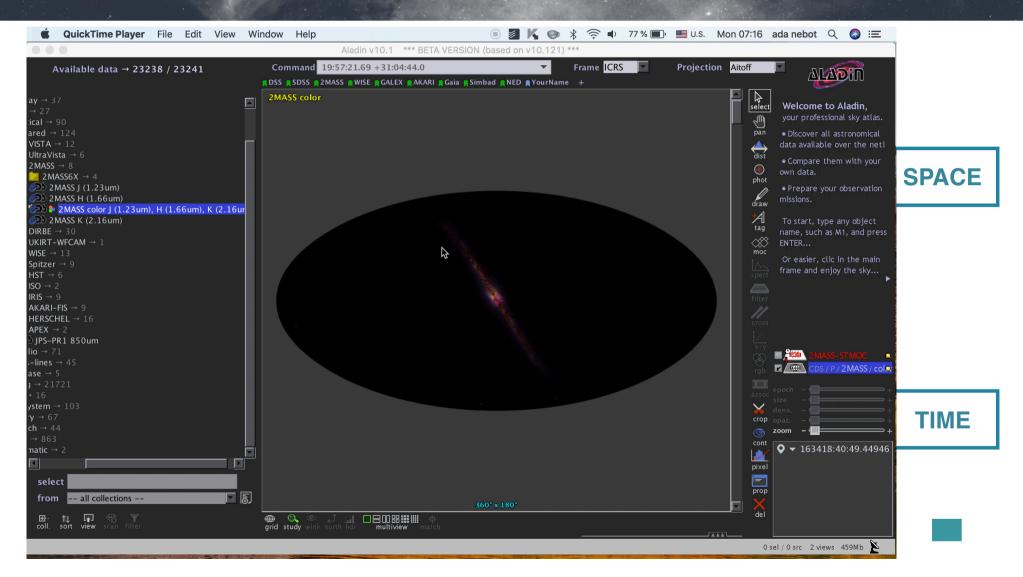


- The STMOC = Space Time MultiOrder Coverage
- Merge together both dimensions in a unique MOC in order to have simultaneously space and time coverage

2 years of CDS&Co R&D driven by the IVOA Time Domain Interest Group

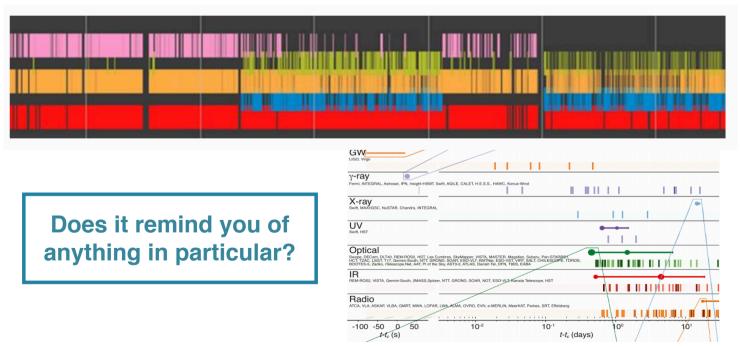


https://wiki.ivoa.net/twiki/bin/view/IVOA/IvoaVOEvent#Time_Series_Data



Possibility of extremely fast

- computations (generate from catalog, images, regions)
 - The only thing you need is a list of positions and times, spatial and temporal resolution)
- operations (unions, intersections,...)
 - E.g. Have these two telescopes observed the same sky region within this interval of time?



• Example use:

- What is the 2MASS mission coverage in space & time ?
- Are there common observations in space and time for XMM and Chandra missions ?
- Which observations are available for this Gravitational Wavelength probability area detected at this epoch ?
- What are the LSST alerts in this sky region and in this interval of time?

ST-MOC: Find where and when

- IVOA Note with technical details (Durand, Fernique, Nebot et al 2018)
- Precomputed for 150 catalogues in VizieR (and more to come!) <u>http://alasky.u-strasbg.fr/footprints/STMOC/</u>
- Precompute for solar system body ephemerides
- Available in MOCPy
- Jupyter Notebook showing examples

MOCpy:docs	astropy 🗹 Index Modules Search Branch-aster 🔨 cpy / notebooks / Space & Time coverages.ipynb	Find file	Copy path
Page Contents mocpy.STMOC	• previous least + Boma hieue use astronguery.vizier to get all the info of the matching observation 1 contributor	3227284	on 21 Aug
revious topic class mocpy. STMOC mocpy.TimeMOC Time-Spatial Coverage class. lext topic max_depth mocpyWCS max_time	[source] 603 lines (602 sloc) 157 KB ↔ E Raw Blame	History	
	Mo. row 1. How to create Space-Time coverages of the 2MASS image catalog and the XMM_DR8 catalog A cover om a u of times and usificons 2. How to compute the intersection of two Space-Time coverages		Θ

Where can I find more info? @ IVOA

https://wiki.ivoa.net/twiki/bin/view/IVOA/IvoaVOEvent#Time_Series_Data

Visualisation of data: search, find and retrieve

KEY POINT: IVOA Standardisation of time annotation

Time Scale: UTC, TT, TAI, TCB,...

Format: JD, MJD, ISO, truncated ISO,...

Offset: e.g. JD-XXX (e.g. Gaia...)

Reference position: Topocentre, Geocentre, Barycentre,... (light-travel correction)

TIMESYS element in VOTables (Demleitner, M., Nebot, A., Bonnarel, et al. 2018)

To characterise and classify sources...

- Multi-wavelength/messenger approach is (sometimes) needed
- Follow-up observations & reaction time for that can be crucial
- Visualisation & navigation
- Coordination & transmission of information

Coordination & transmission

- collect what was/will observed, when, in which wavelength, ...
- alerts, emails, webpages, references,...
 - A multi-messenger platform ?

Summary

- To enable access, discovery and interoperability the VO is based on standards
- The Time Domain standards needed for time domain multi-messenger astronomy are:
 - Existing (e.g. VOEvent, TAP, VOTable, MOC, HiPS...)
 - or under development:
 - Definition of the minimum metadata for time (TIMESYS in VOTables)
 - Space + time coverage (STMOC)
 - Visibility & Observation locator (ObjVisSAP & ObsLocTAP)