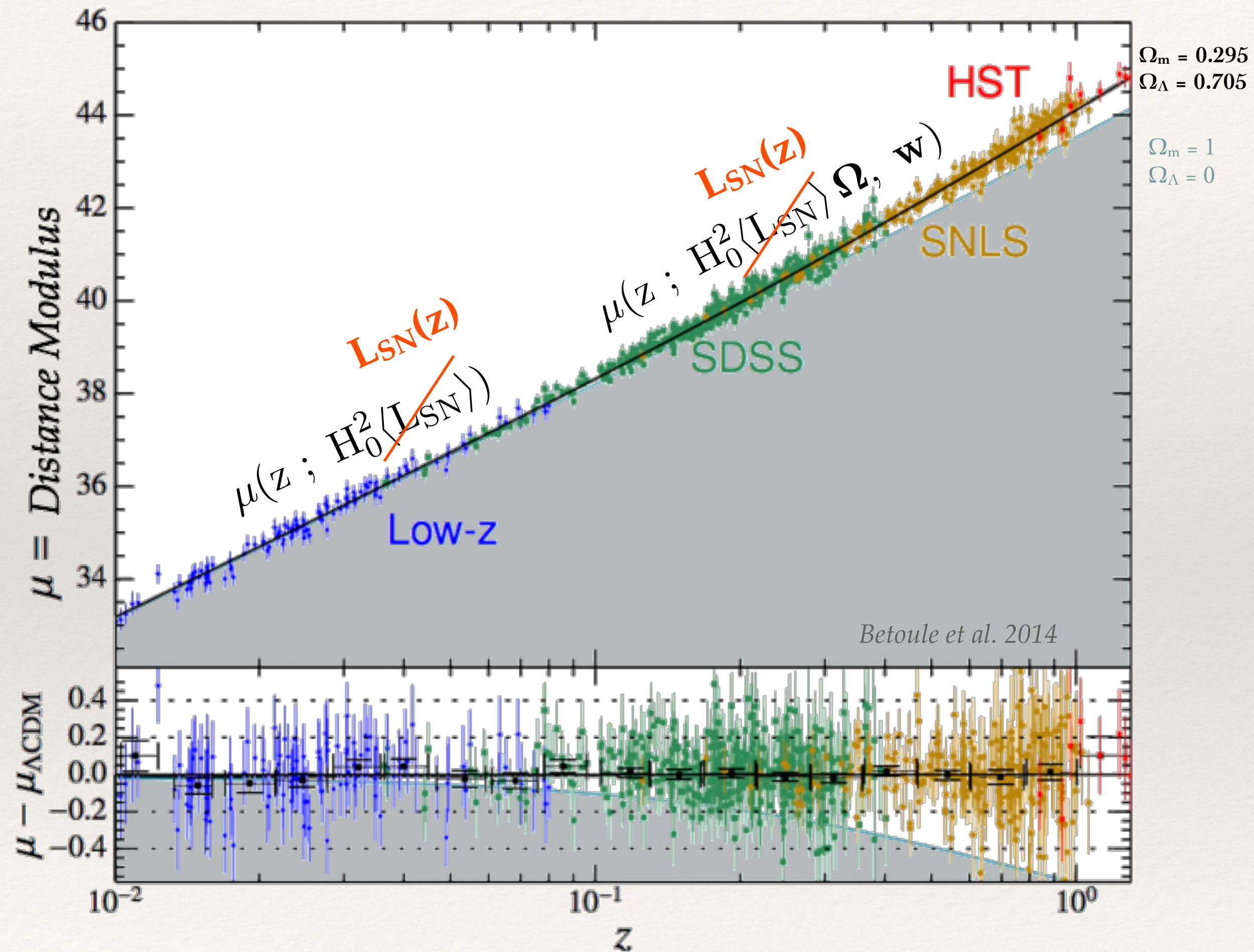




*Host*



# Type Ia Supernovae | Redshift drift



Is the average standardised  
SN luminosity constant ?

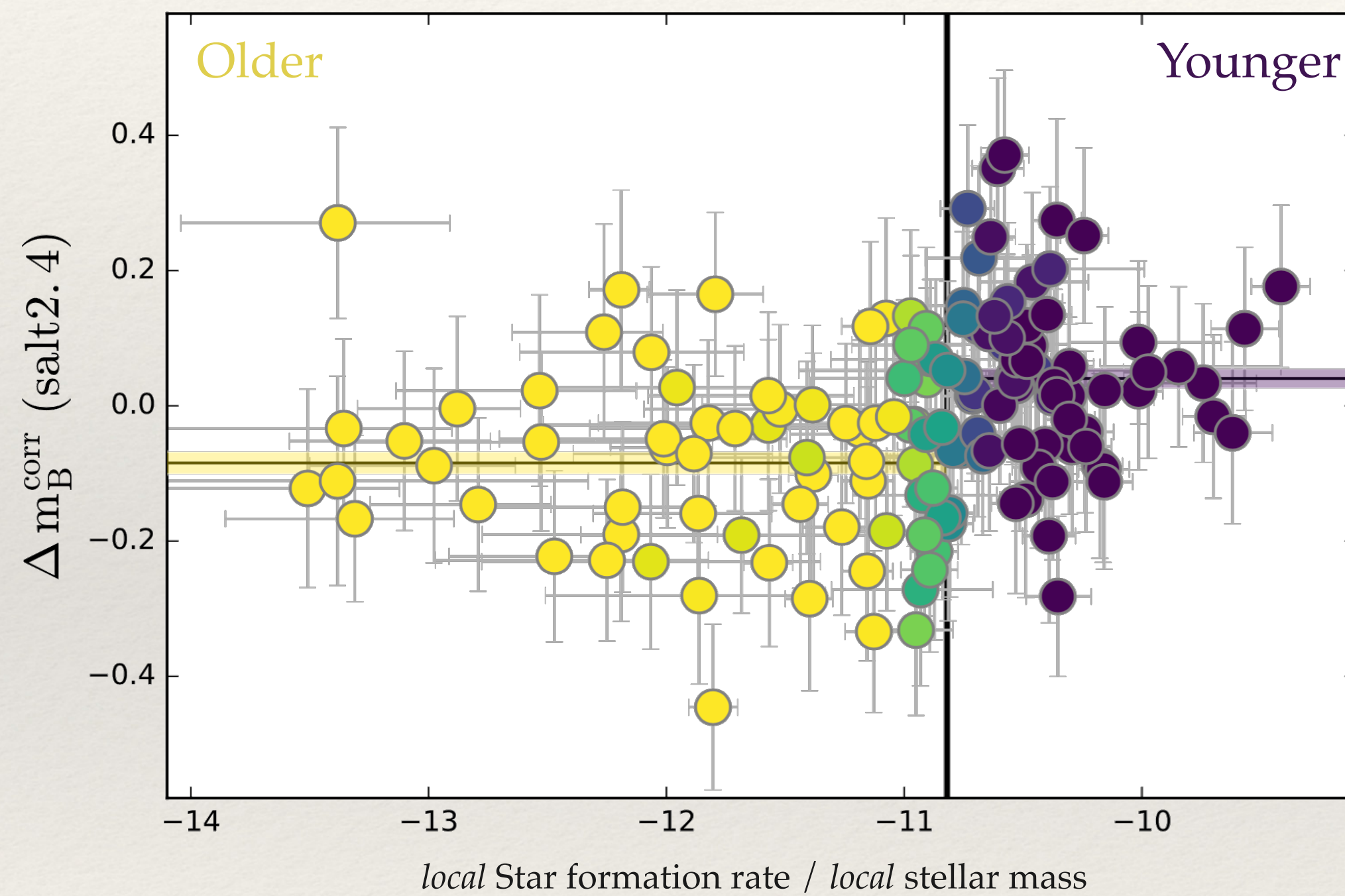
*Between sample | as a function of redshift*



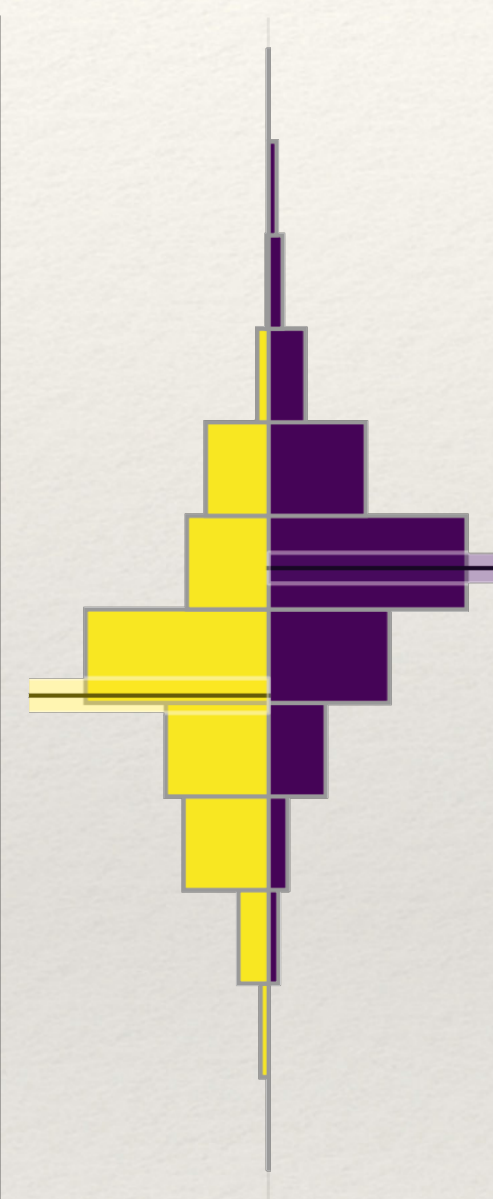
# SNeIa Astrophysical dependencies

Low-Z | SNf  
*Rigault et al. 2018*

non-zero at  $\sim 6\sigma$  level |  $\Delta_Y = 0.16 \pm 0.03$

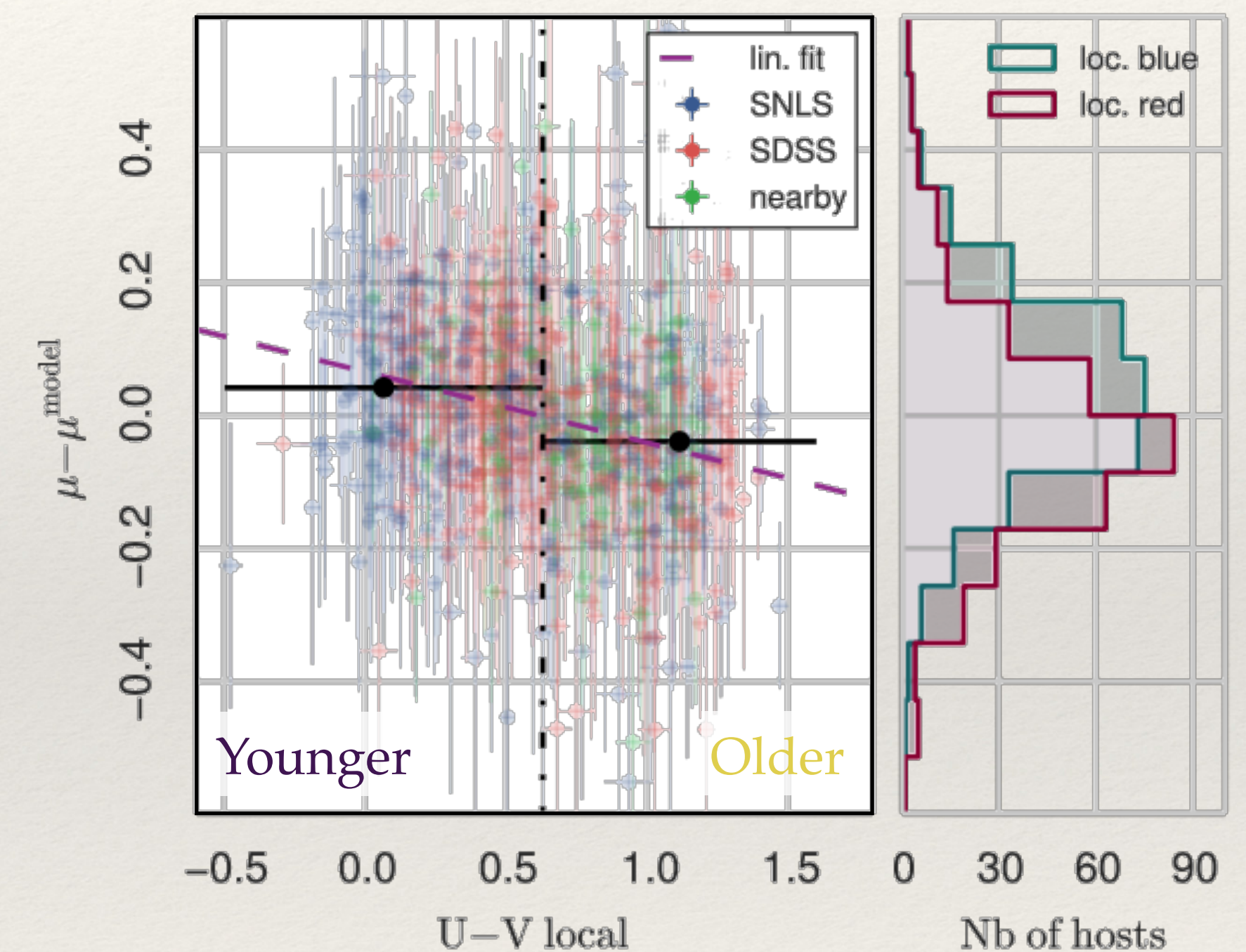


Fraction of young star at the SN location



All Z | SDSS & SNLS & Nearby  
*Roman et al. 2018*

non-zero at  $\sim 7\sigma$  level |  $\Delta_Y = 0.09 \pm 0.02$

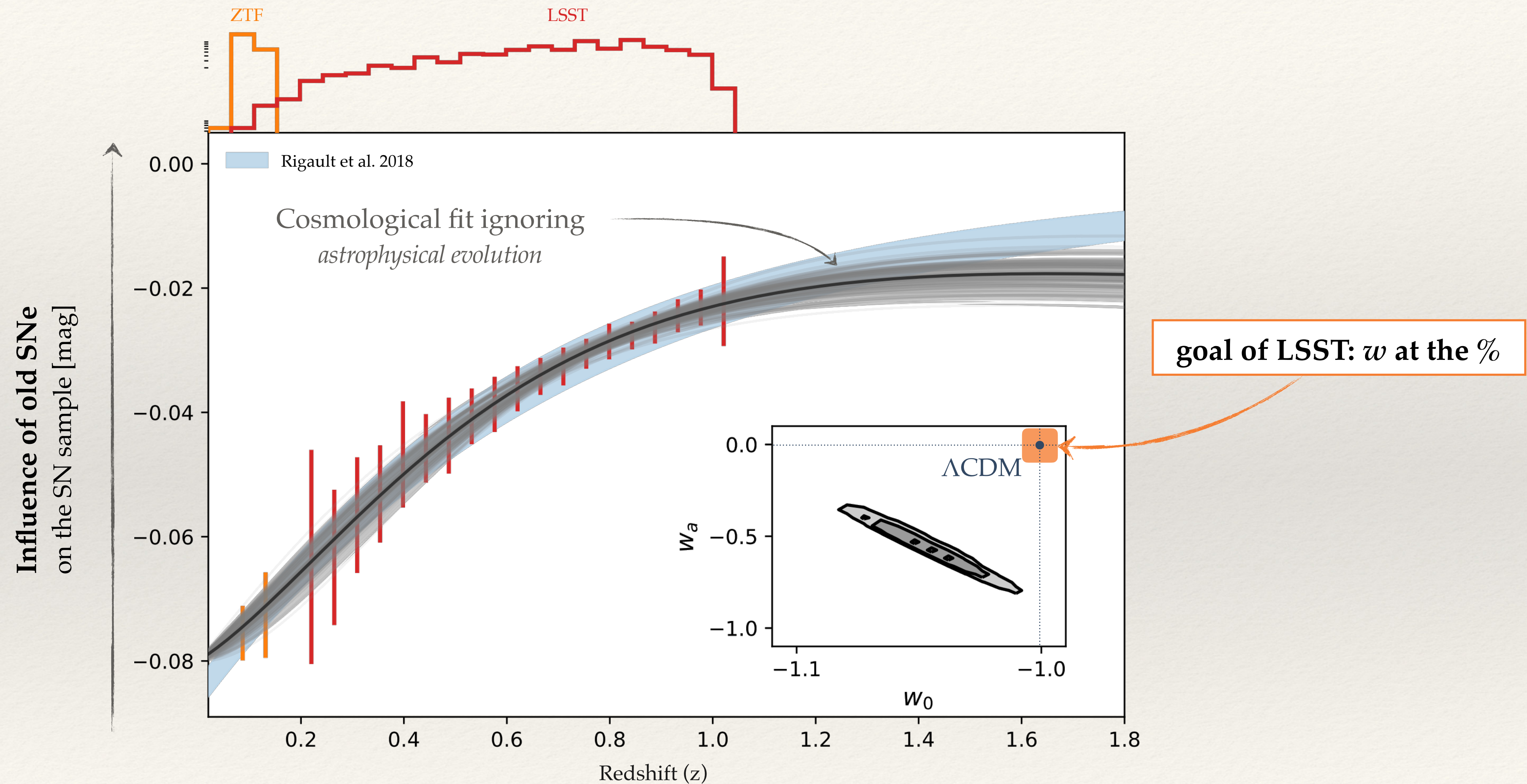


Fraction of young star at the SN location



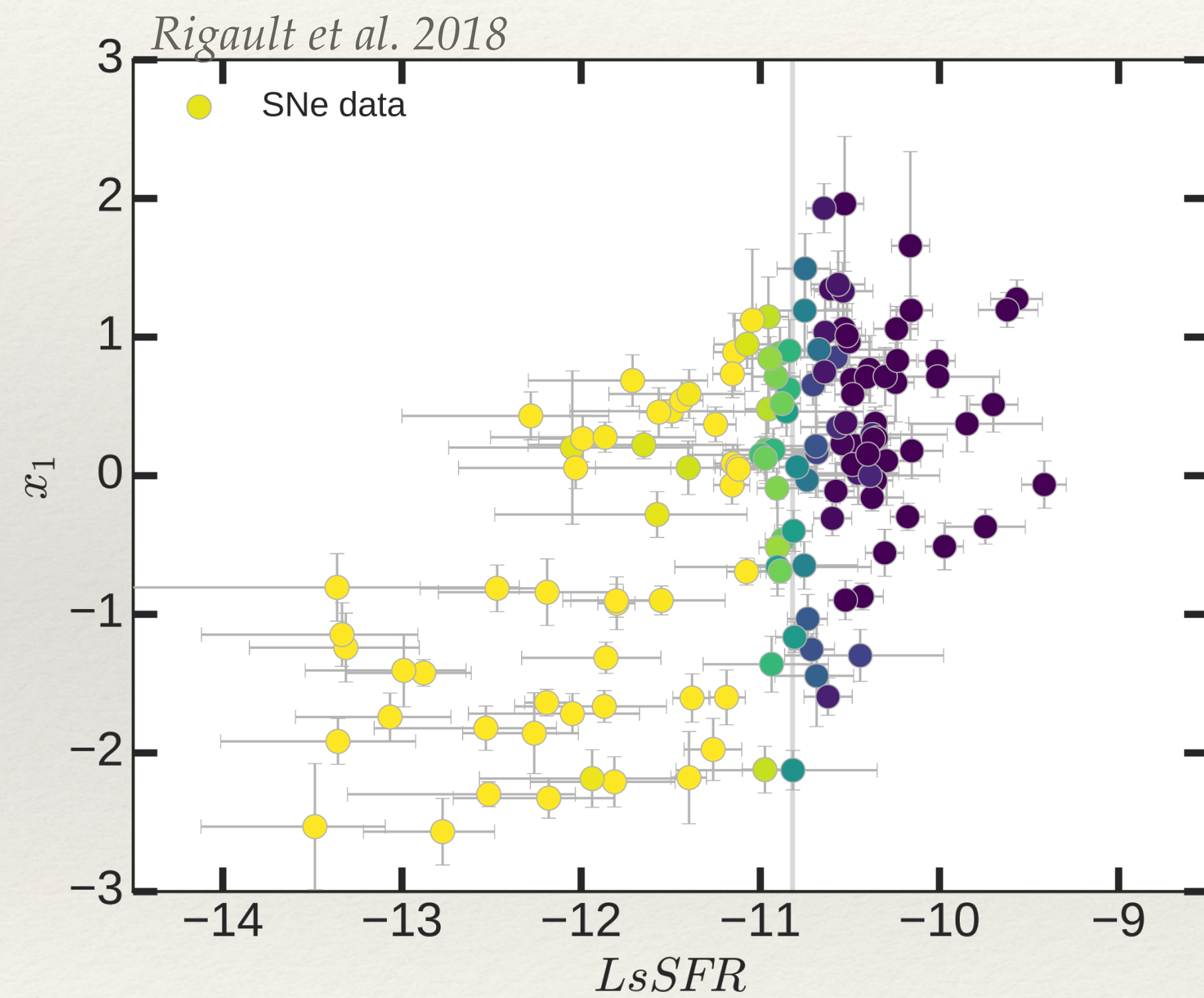
# Redshift Drift vs. Exotic Dark Energy

Rigault et al. 2018





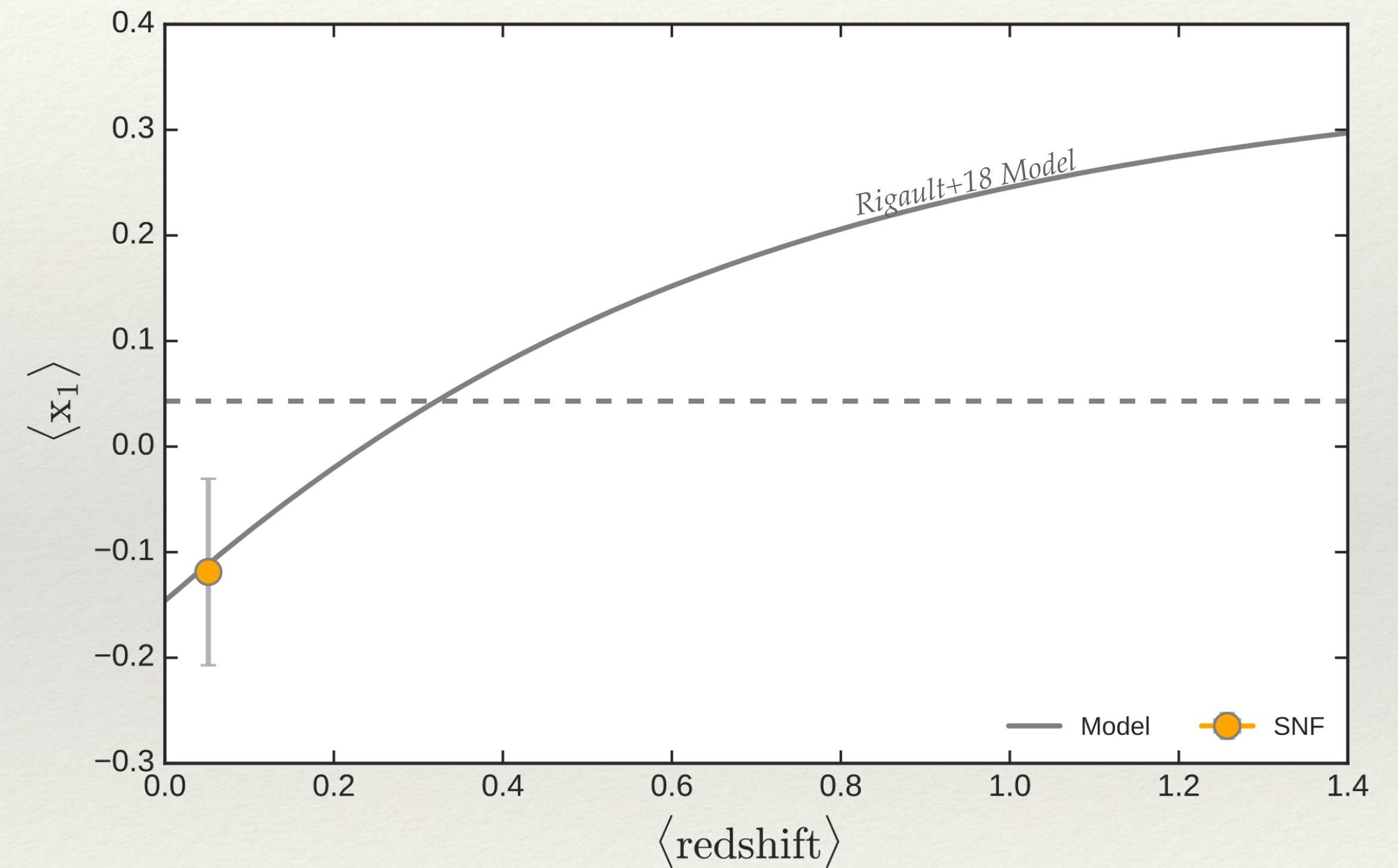
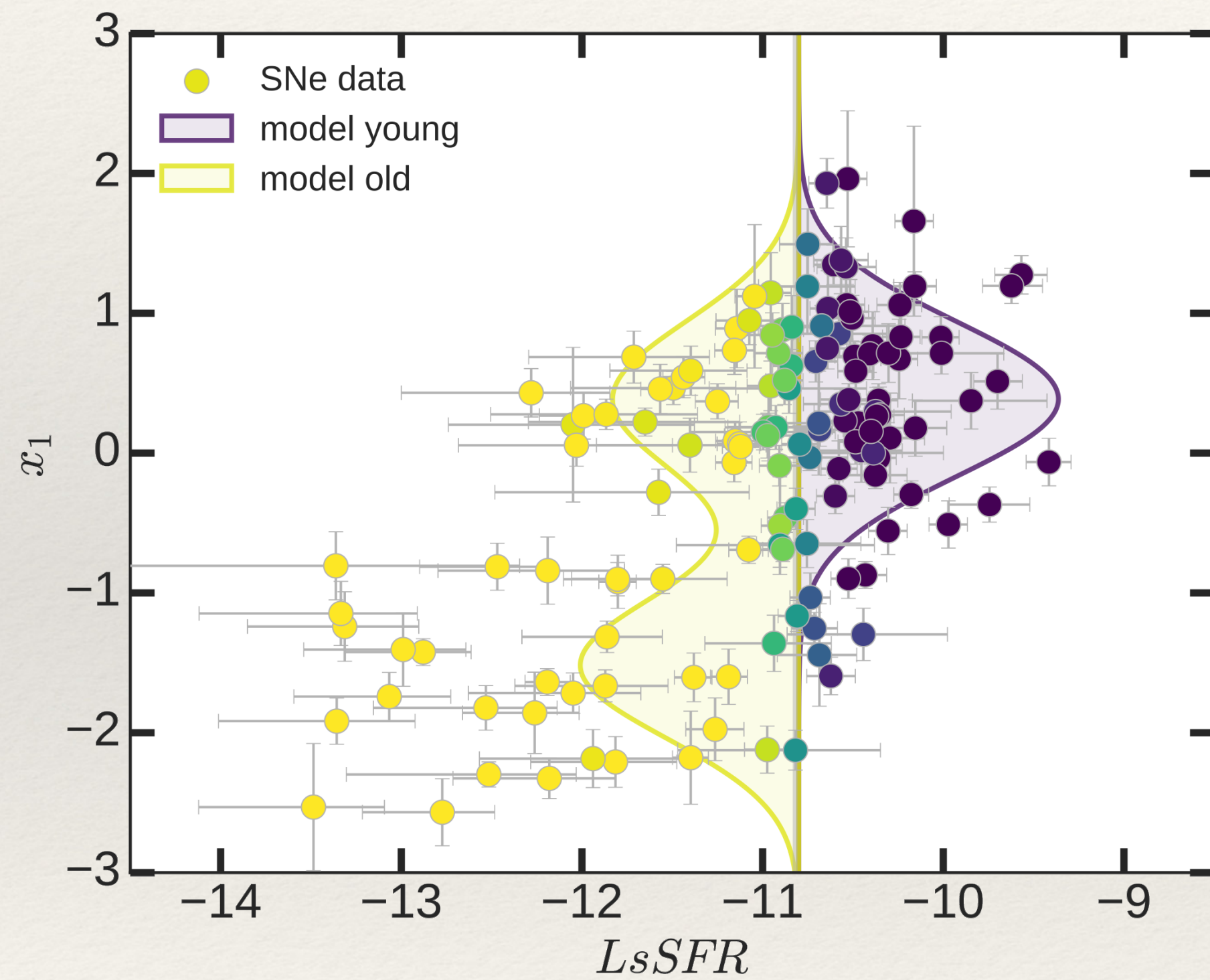
# Testing Redshift Drift model





# Testing Redshift Drift model

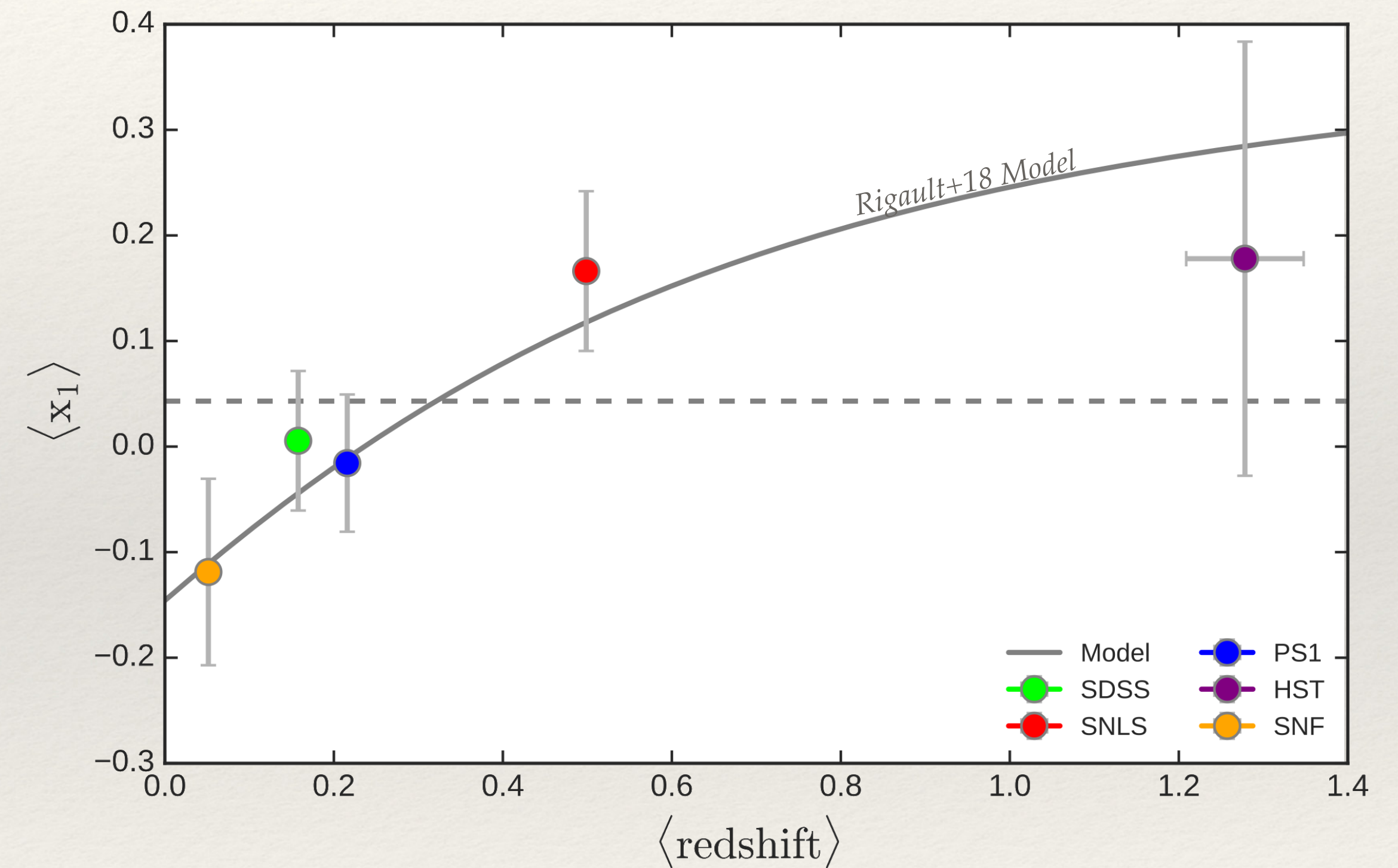
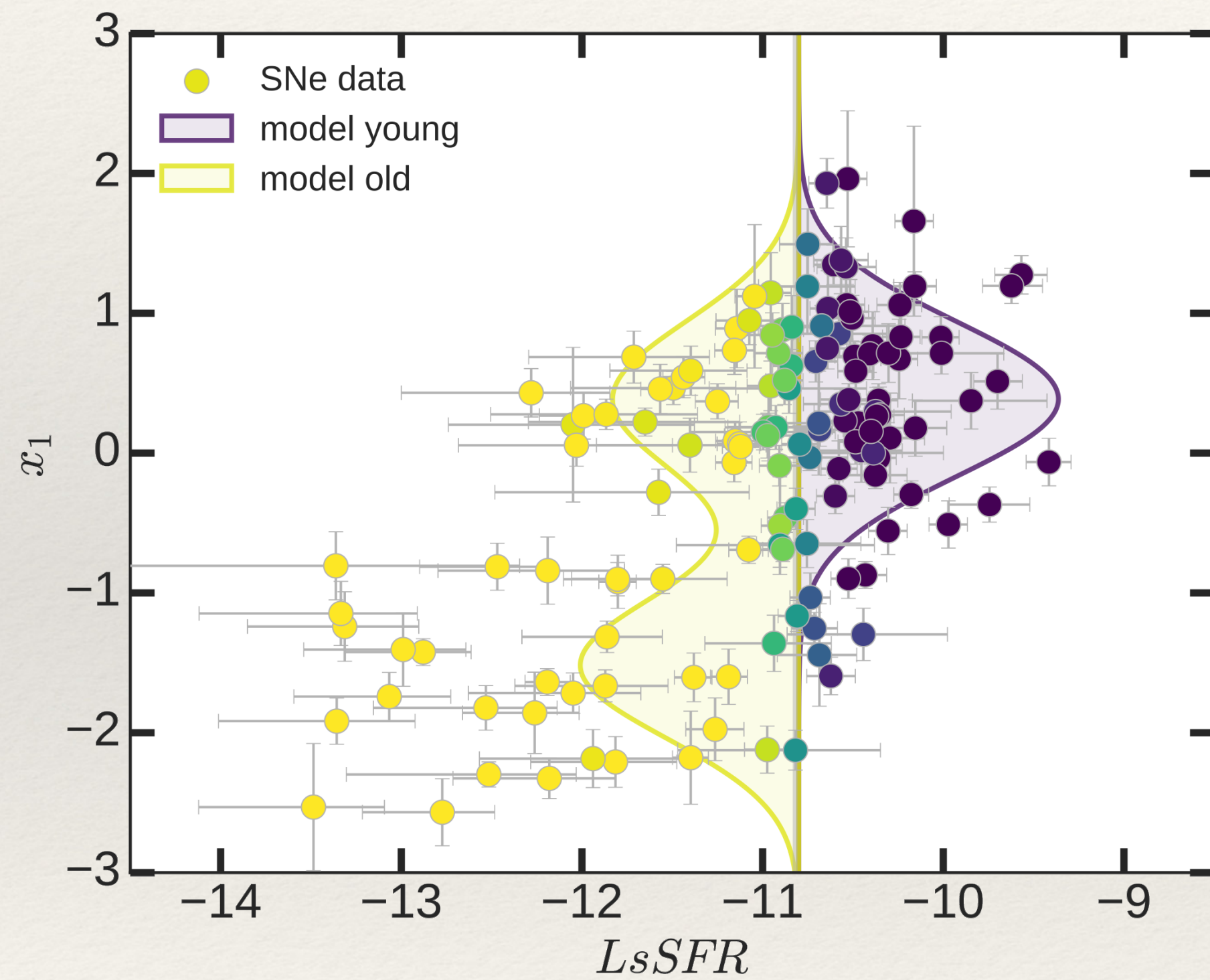
Nicolas et al. in prep





# Testing Redshift Drift model

Nicolas et al. in prep

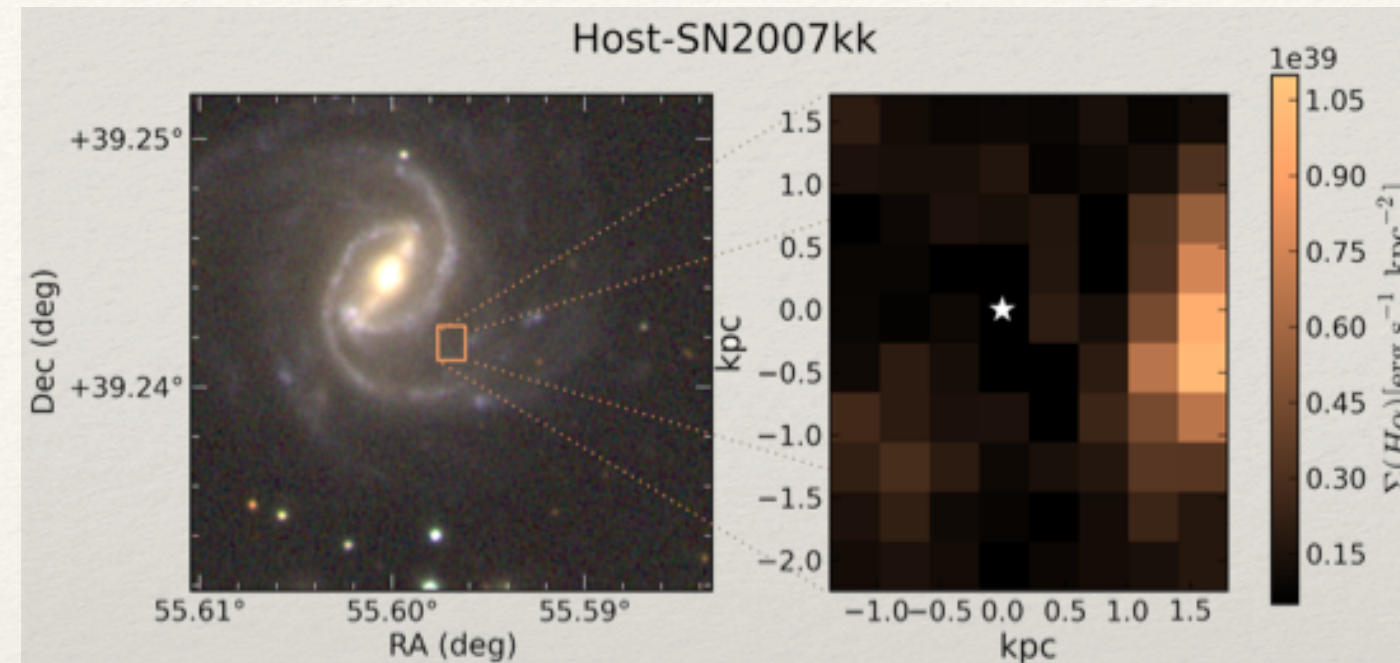




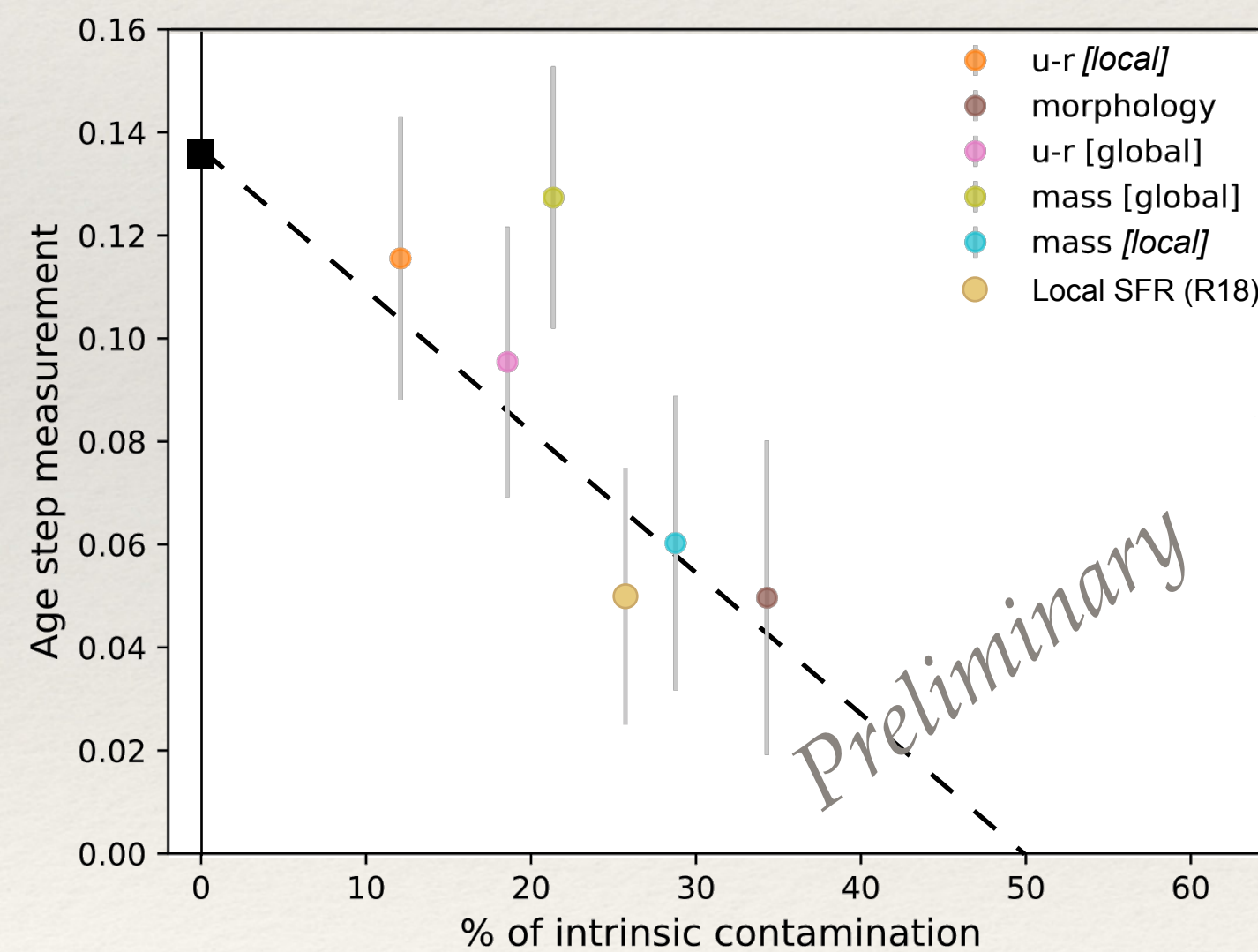
# How to measure environment at high-z ?

Briday et al. in prep

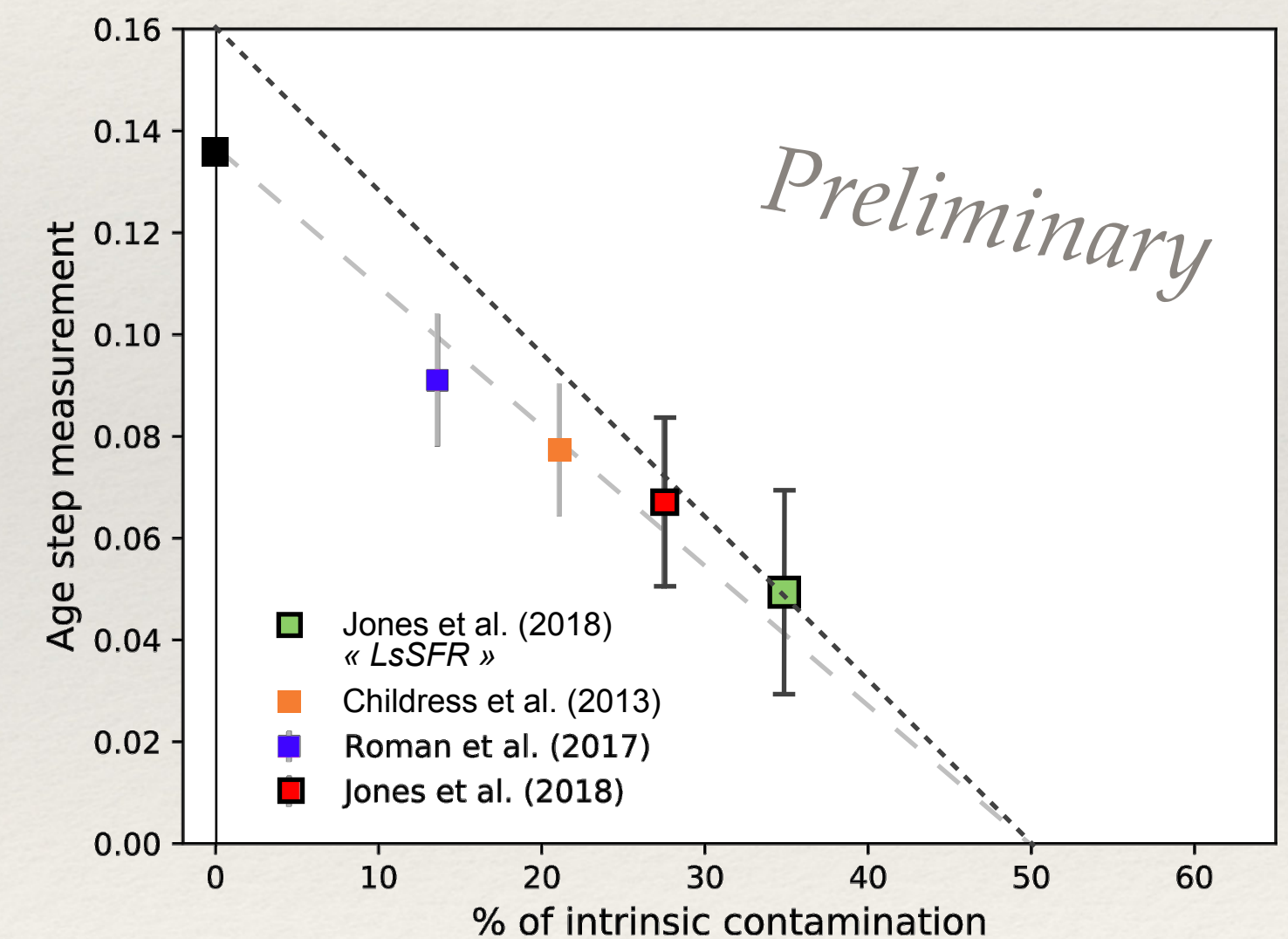
— On SNf data —



— On Public data —



Misclassification = Tracer Inaccuracy  
e.g. *Spiral* → *Young Progenitor*





# Host

Dark Energy ( $w_0, w_a$ )

Hubble Constant

Peculiar Velocities

USNAC ERC group is actively working on it.

