# Using simulated quasar catalogs for the BAO in lyman- $\alpha$ analysis of eBOSS and DESI

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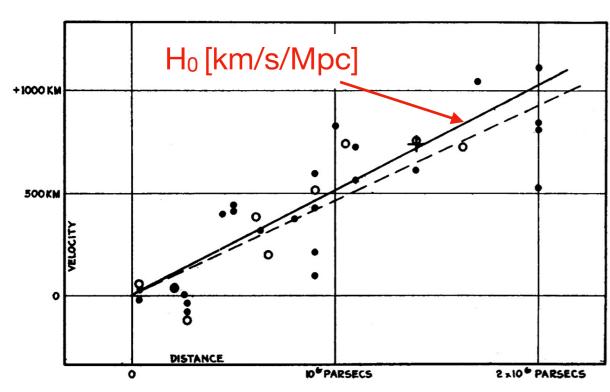
# Dark energy and the accelerated expansion of the universe

#### First measurements

#### 200g of dark energy sugar

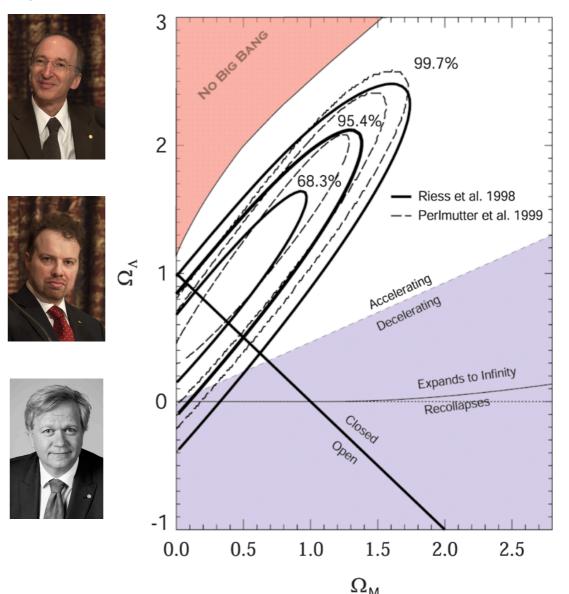
 1929 - expansion of the universe: Galaxies



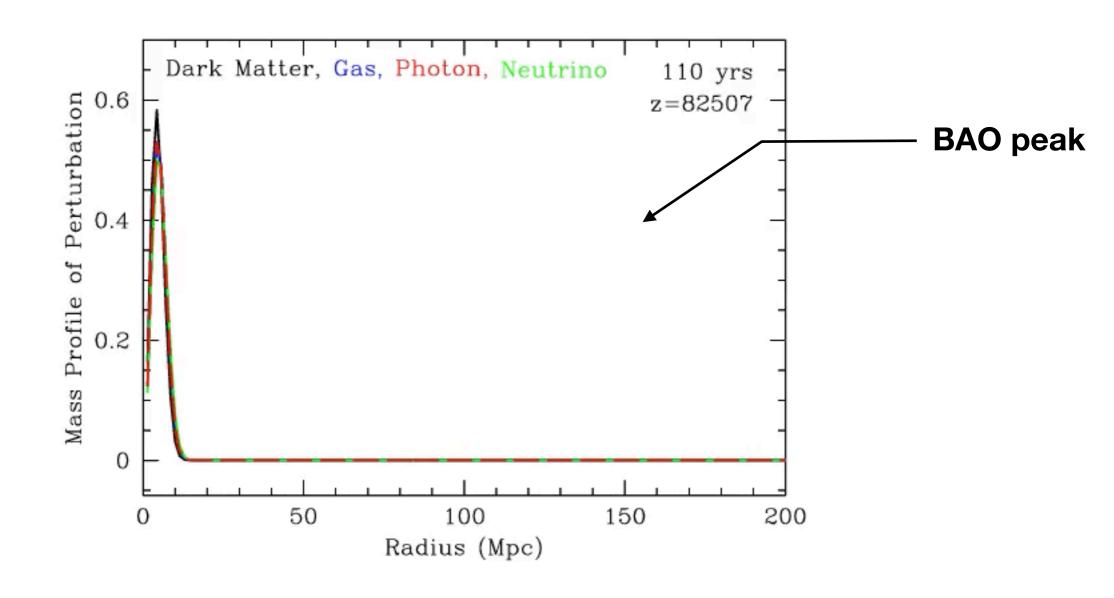


Original Hubble diagram

1998 - accelerated expansion: type la supernovae



# Baryon Acoustic Oscillation (BAO)



Standard ruler

200g salted butter plasma

$$z = \frac{\lambda_{obs}}{\lambda_{rf}} - 1$$

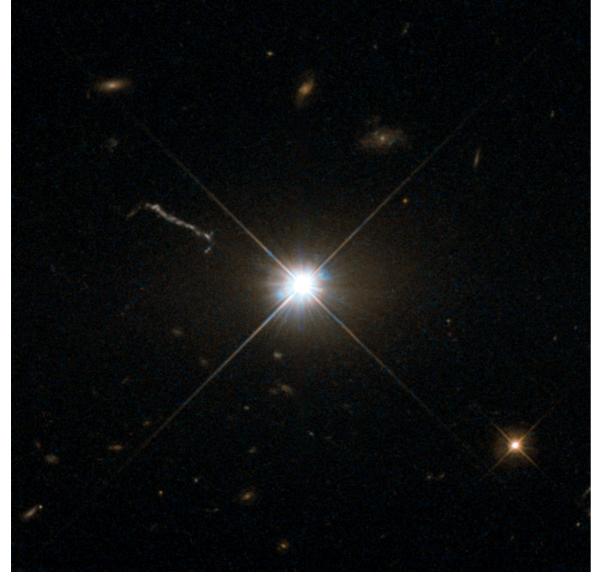
Eisenstein et al. 2007

### Quasars

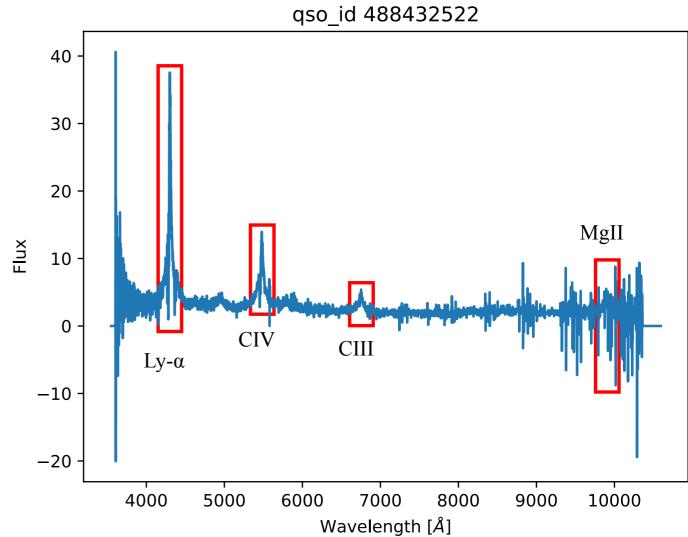
Very Luminous Active Galactic Nuclei (AGN)
Gaseous accretion disk



ESA/Hubble: artist's impression



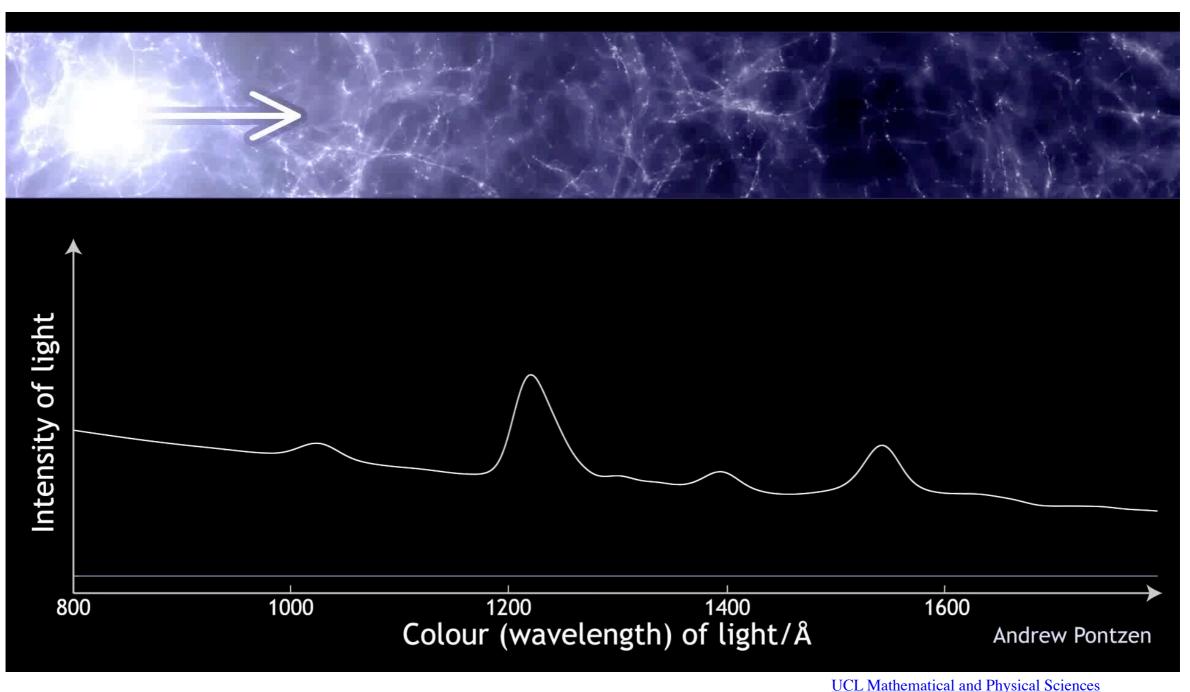
ESA/Hubble: Best image of bright quasar 3C 273



# Lyman-alpha (Ly- $\alpha$ ) forest

#### 250g of Neutral H clouds flour

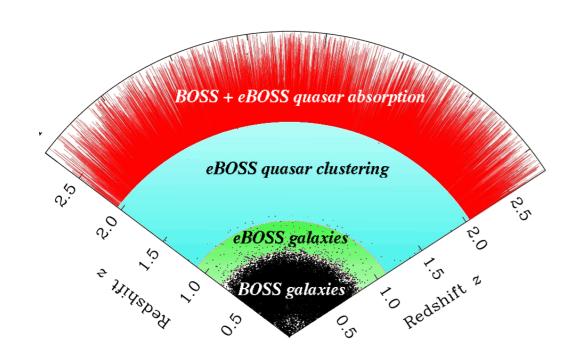
neutral H clouds on the line of sight (LOS) which absorb part of light from quasars -> ly-a forest



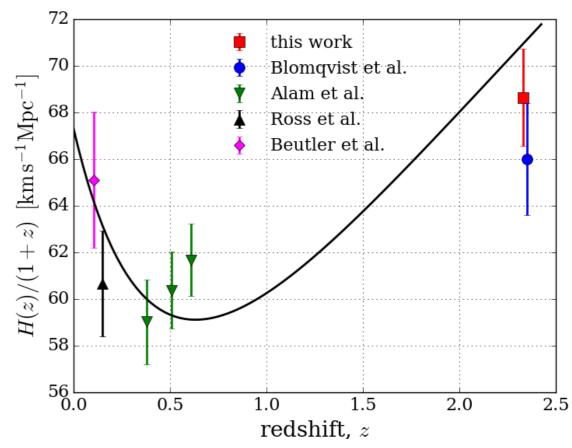
# eBOSS - extended Baryon Oscillation Spectroscopic Survey

#### Sloan Digital Sky Survey - NM, USA

- 2.5 m telescope
  - eBOSS: 200k + quasars in 2019
- 0.5 < z < 3
- 3D map of ⅓ of sky





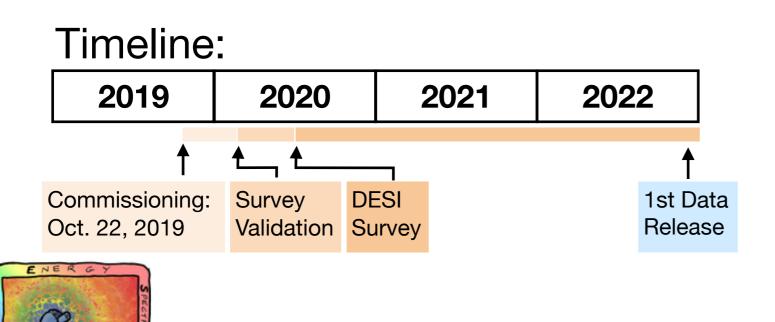


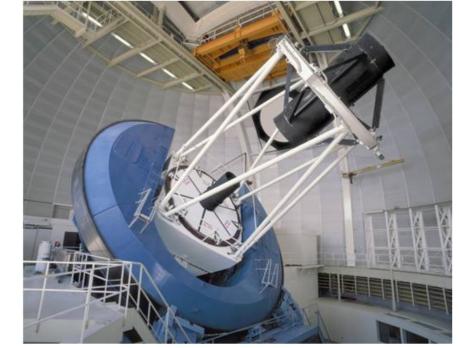
de Sainte Agathe et al. 2019

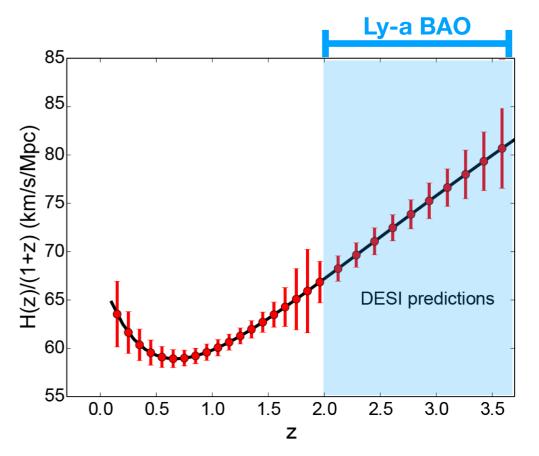
# DESI - Dark Energy Spectroscopic Instrument

Kitt Peak National Observatory - Az, USA

- 4 m Mayall telescope
- 2.5 million quasars to be observed
- 0.5 < z < 3.5
- 3D map of 1/3 of sky



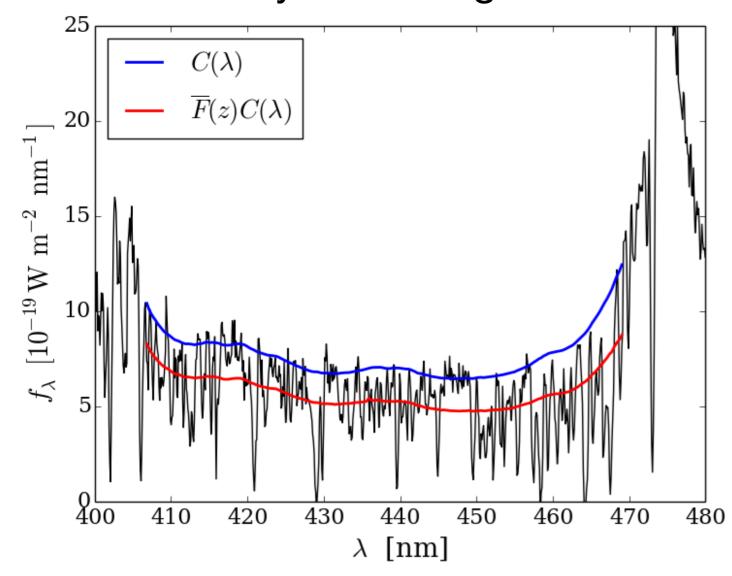


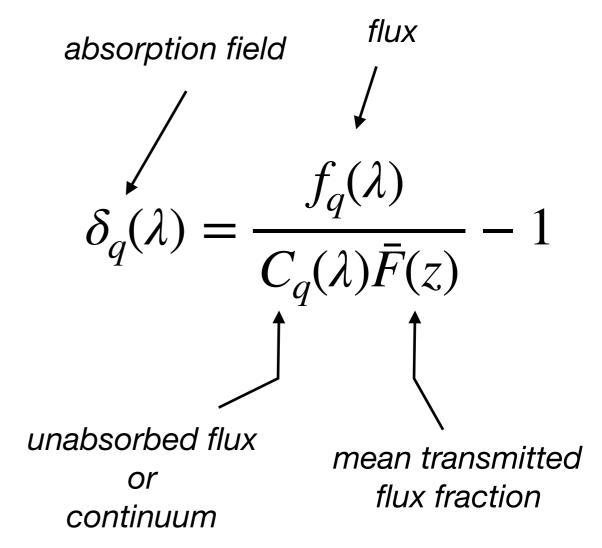


# Measuring the ly-a auto-correlation function:

Absorption field  $\delta$ 

#### In the lyman- $\alpha$ region:





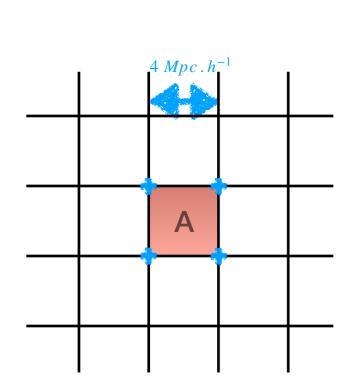
Bautista et al. 2017

# Measuring the ly-a auto-correlation function:

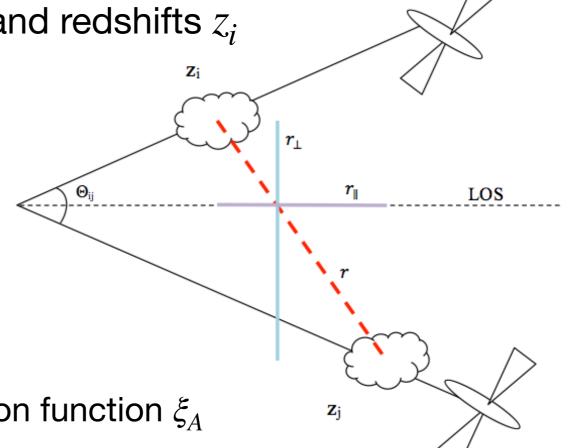
Angular separations  $\theta_{ij}$  and redshifts  $z_i$ 



Cosmology dependent: Planck 2016



Calculating the correlation function  $\xi_A$ 

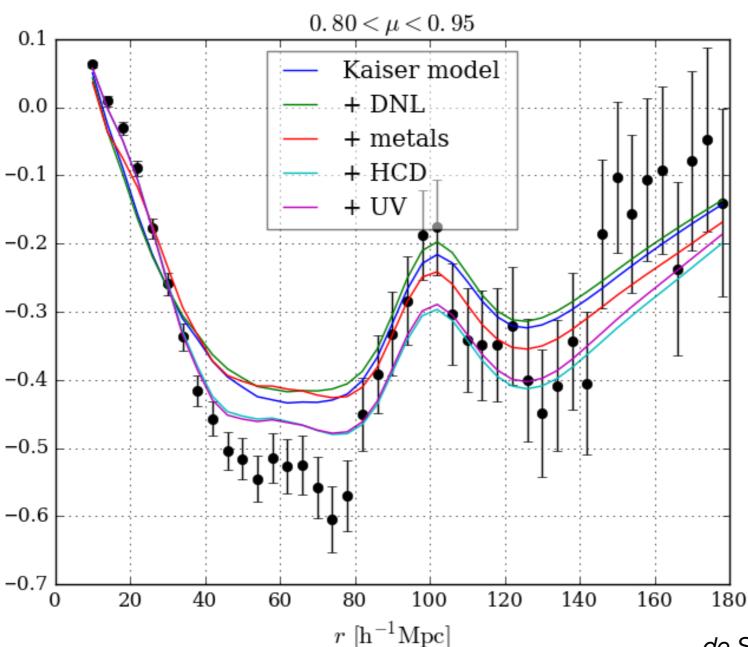


$$\xi_A = \langle \delta(x) \rangle \langle \delta(x+r) \rangle$$

# Measuring the ly-a auto-correlation function:

Model

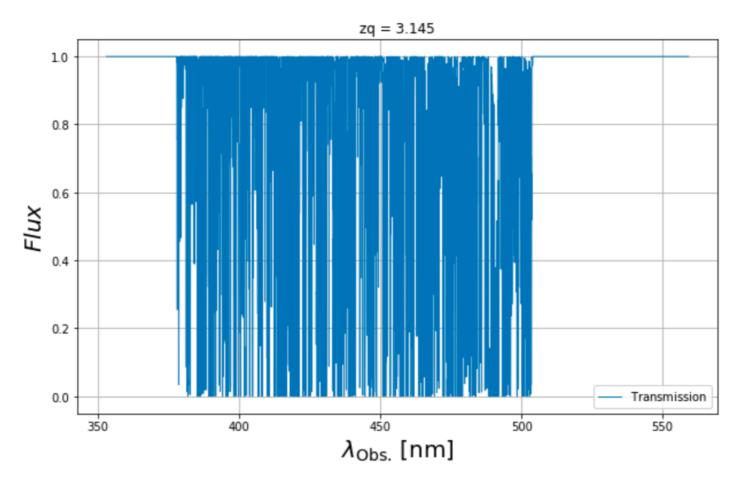
Decompose: 
$$\xi_A = \xi_{smooth} + \xi_{peak}$$



# Simulations - Mocks

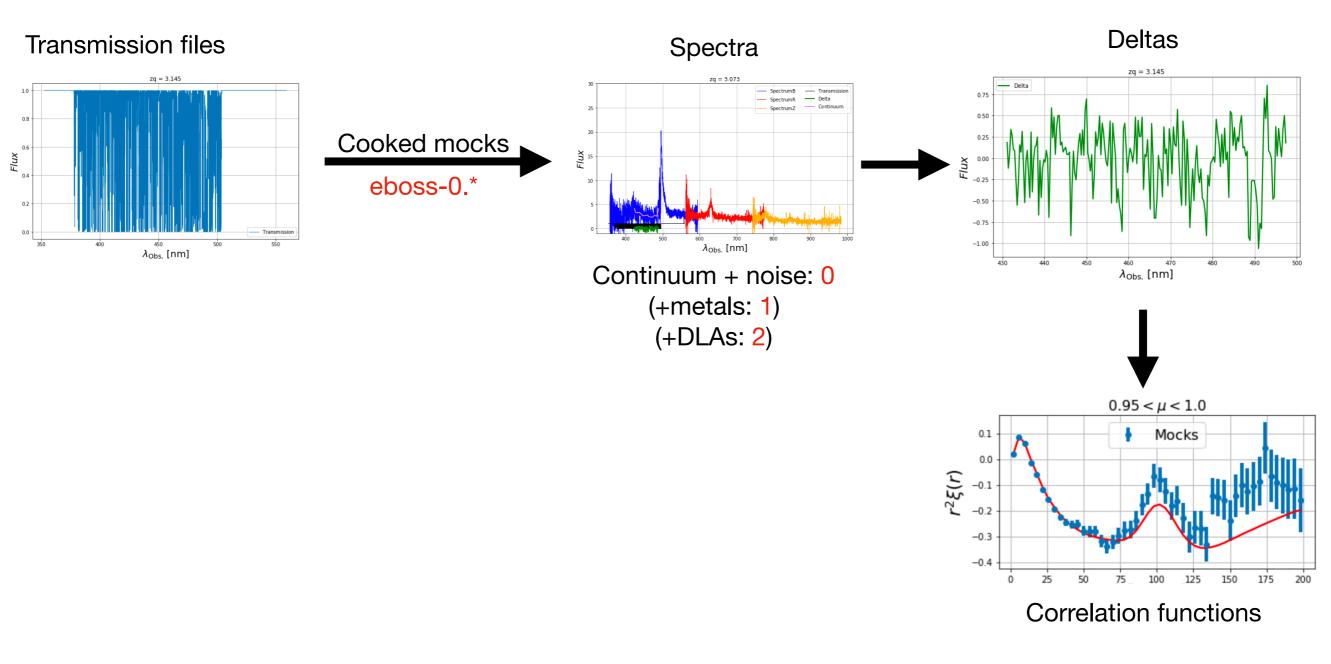
#### Concept

- 2 different sets developed by teams in Saclay and London
- Focus on Saclay mocks
  - ullet Random gaussian fields  $\delta_k$  tuned for correct  $\xi_{1D}$  and  $\xi_{3D}$



**Transmission field example** 

### Analysis of the Mocks



Mocks

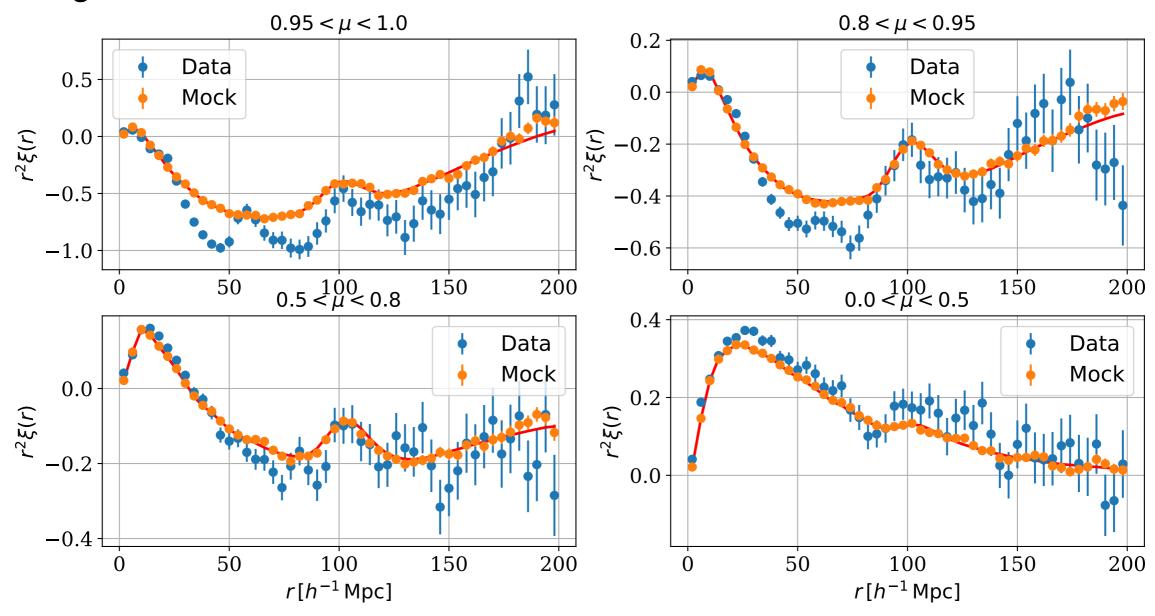
Quickquasars

Picca

# BAO analysis on mocks

Divide the sky in 4 wedges of angle  $\mu = \frac{r_{\parallel}}{r}$ 

#### **Along the LOS**

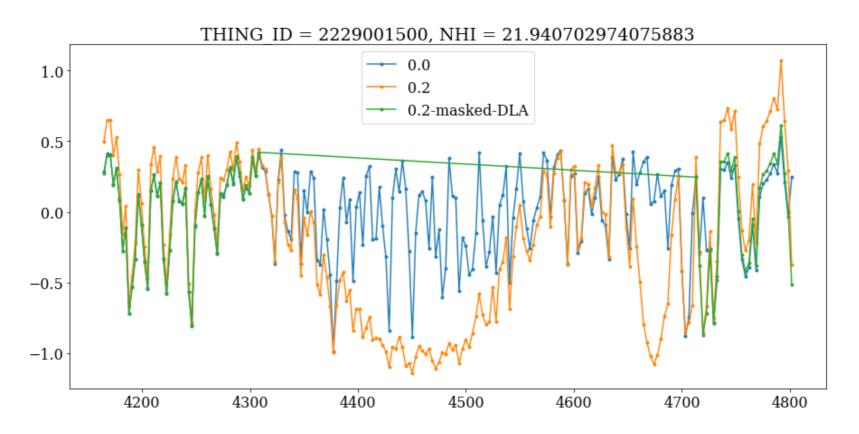


Across the LOS

# Damped Ly-a absorbers (DLAs)

Regions along the LOS of the quasar with high concentration of neutral hydrogen gas i.e.  $n_{HI} \ge 2.10^{20} \ atoms/cm^2$ 

- $^{ riangle}$  Skew estimation of  $\delta$  field for specific wavelength
- Run finder algorithm to find and mask DLAs in the forest



Example of DLAs in 1 mock quasar

### DLA finder: performances

200

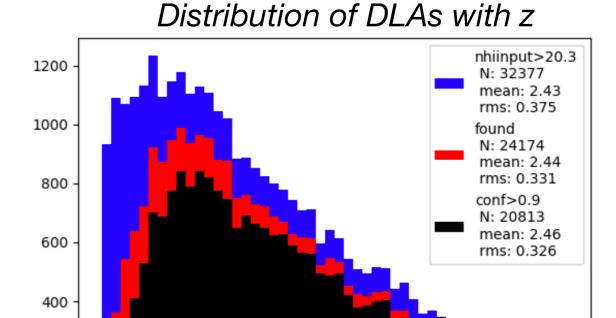
Found =  $\Delta z \leq 0.006$ 

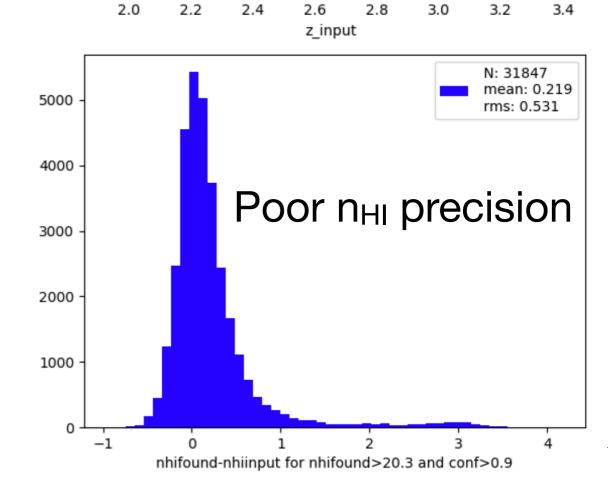
Efficiency: how many DLAs are found

▶ 60%

Purity: if a DLA is found, does it exist

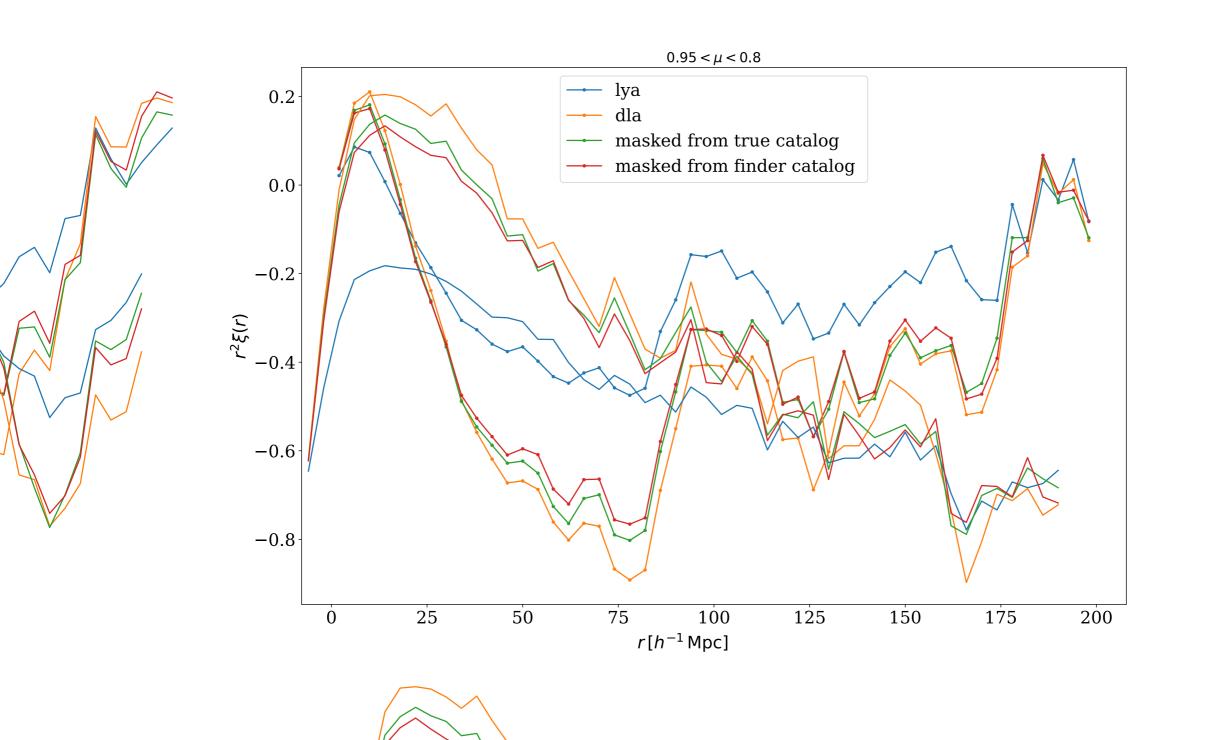
▶ 90%





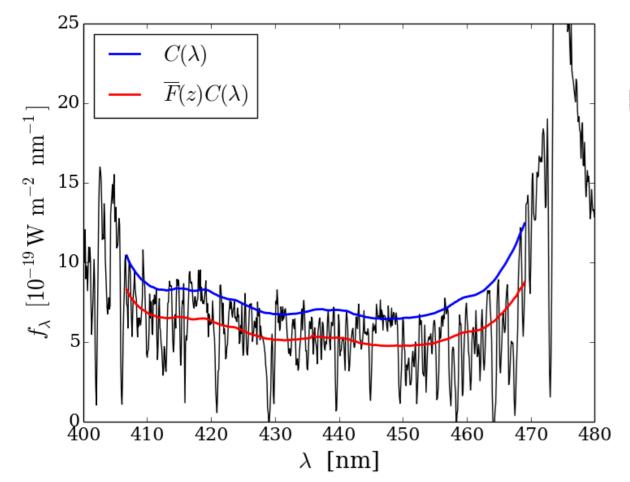
# Damped Ly-a absorbers (DLAs)

Effects on the correlation function



### Continuum estimation

$$\delta_q(\lambda) = \frac{f_q(\lambda)}{C_q(\lambda)\bar{F}(z)} - 1$$



Bautista et al. 2017

No way of knowing TRUE continuum

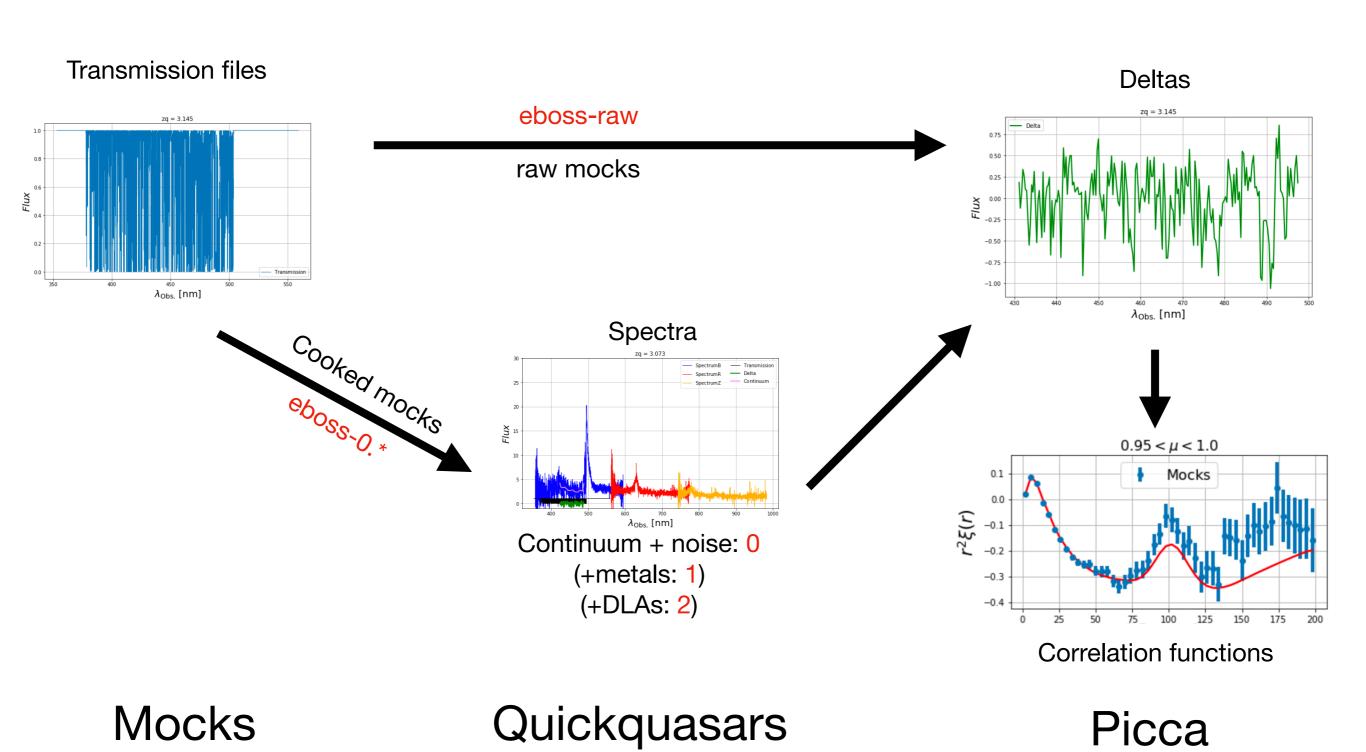
needs to be estimated

$$C_q(\lambda) = C(\lambda_{RF}) (a_q + b_q log(\lambda))$$

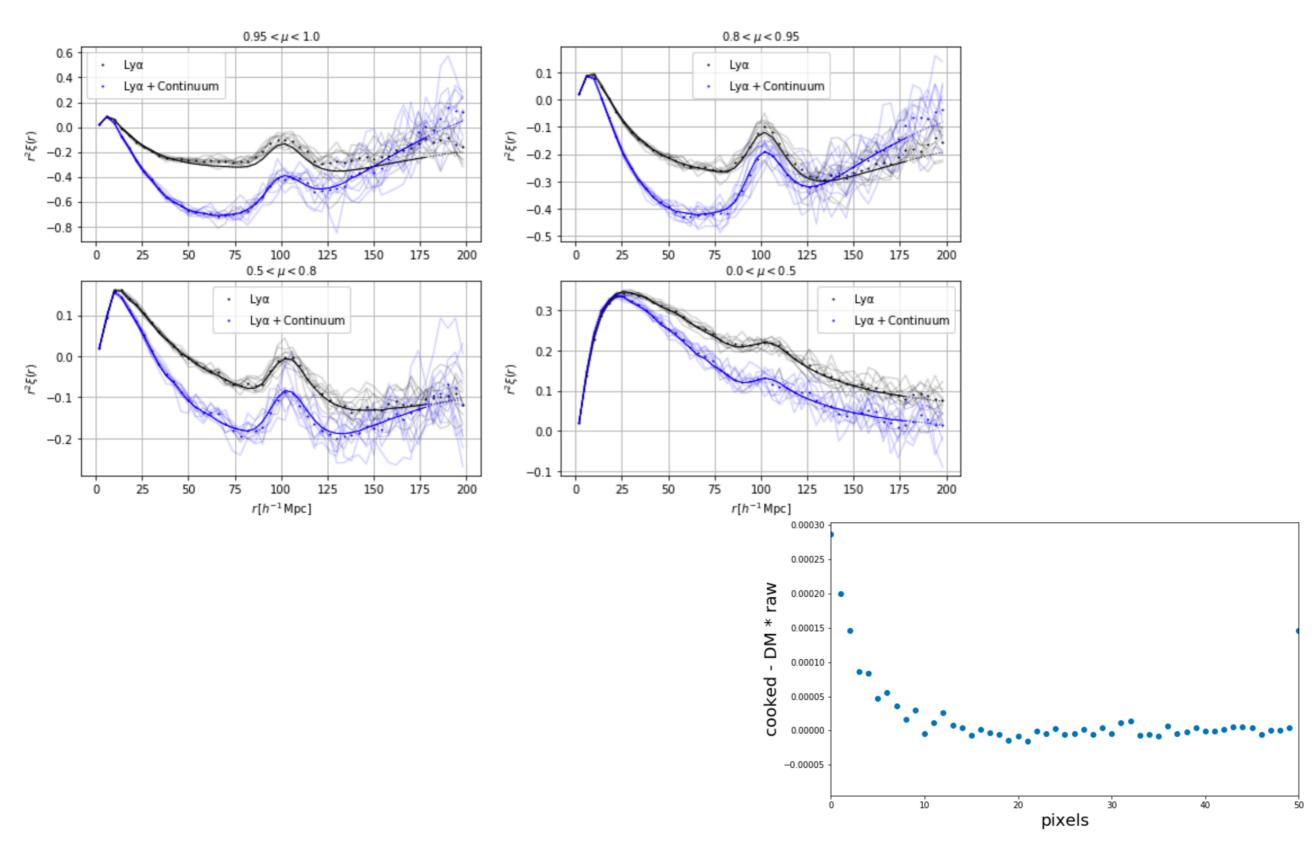
Introduces additional correlation

distortion matrix

# Analysis of the Mocks



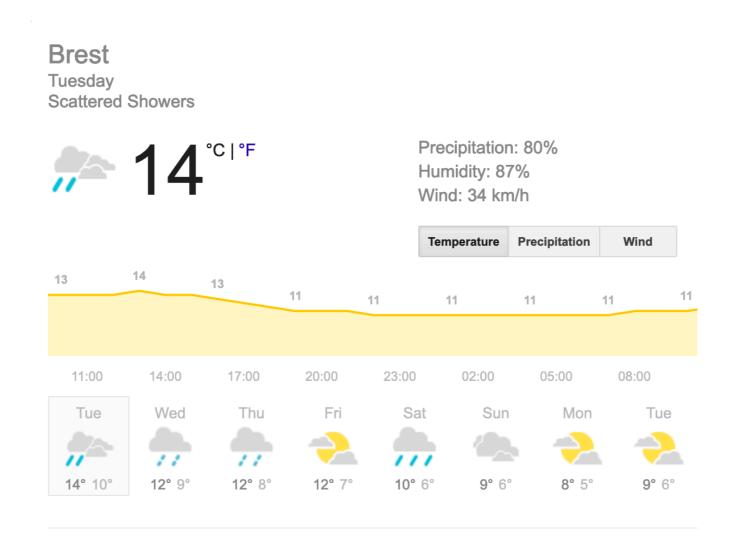
# Comparison between raw and cooked mocks: auto correlation



#### Conclusion

#### Mix all the ingredients together

#### 10cl of Bretagne rain water



#### Thunderstorms Warning (Yellow)

Finistere

9 hours ago - Météo-France

Des phénomènes habituels dans la région mais occasionnellement et localement dangereux sont prévus, (ex ...

### Conclusion

#### Bake at 210° for 35 minutes



### Conclusion

Wash it down with some Chuchen



