

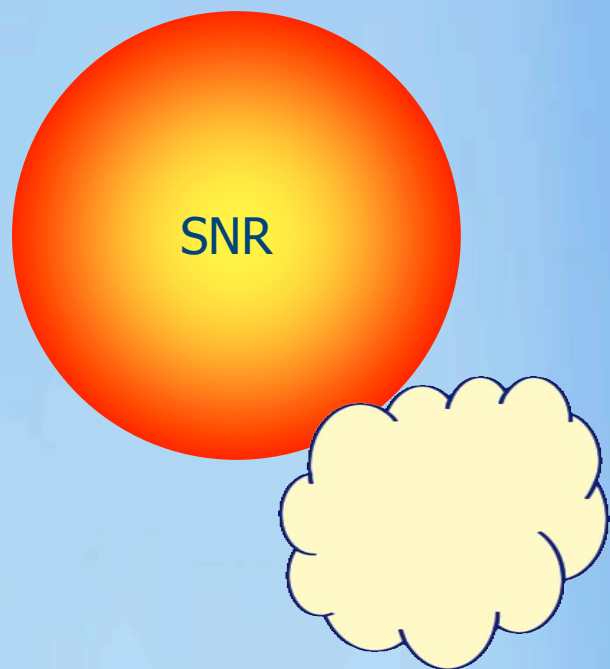
Where to search for Cosmic-Ray and Molecular Clouds interaction ?

Fabio Acero
& collaborators



Interaction of CRs with MCs

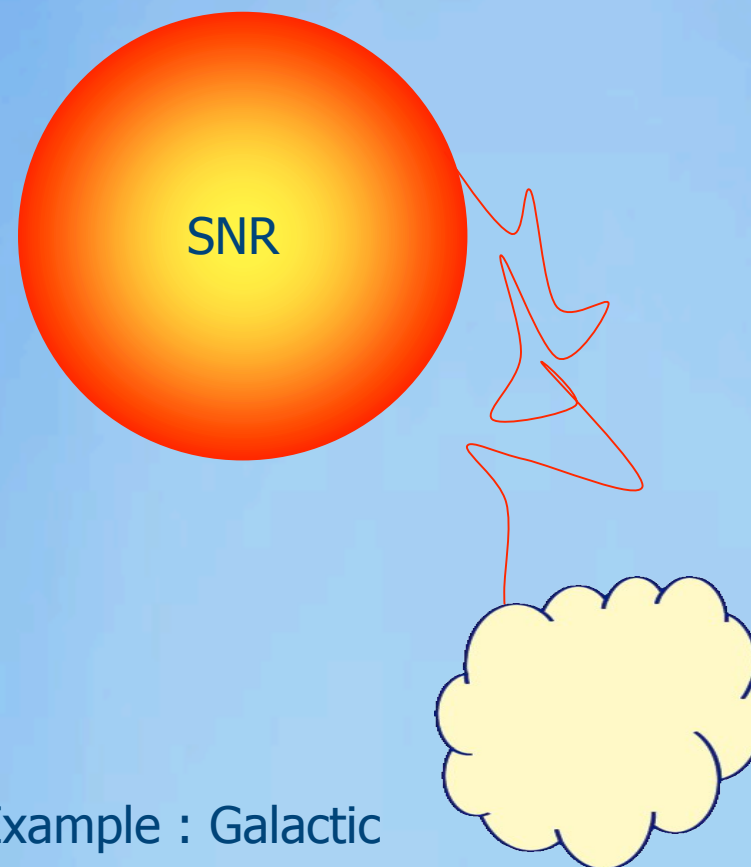
Direct interaction



Examples : IC443, W28, W44, ...

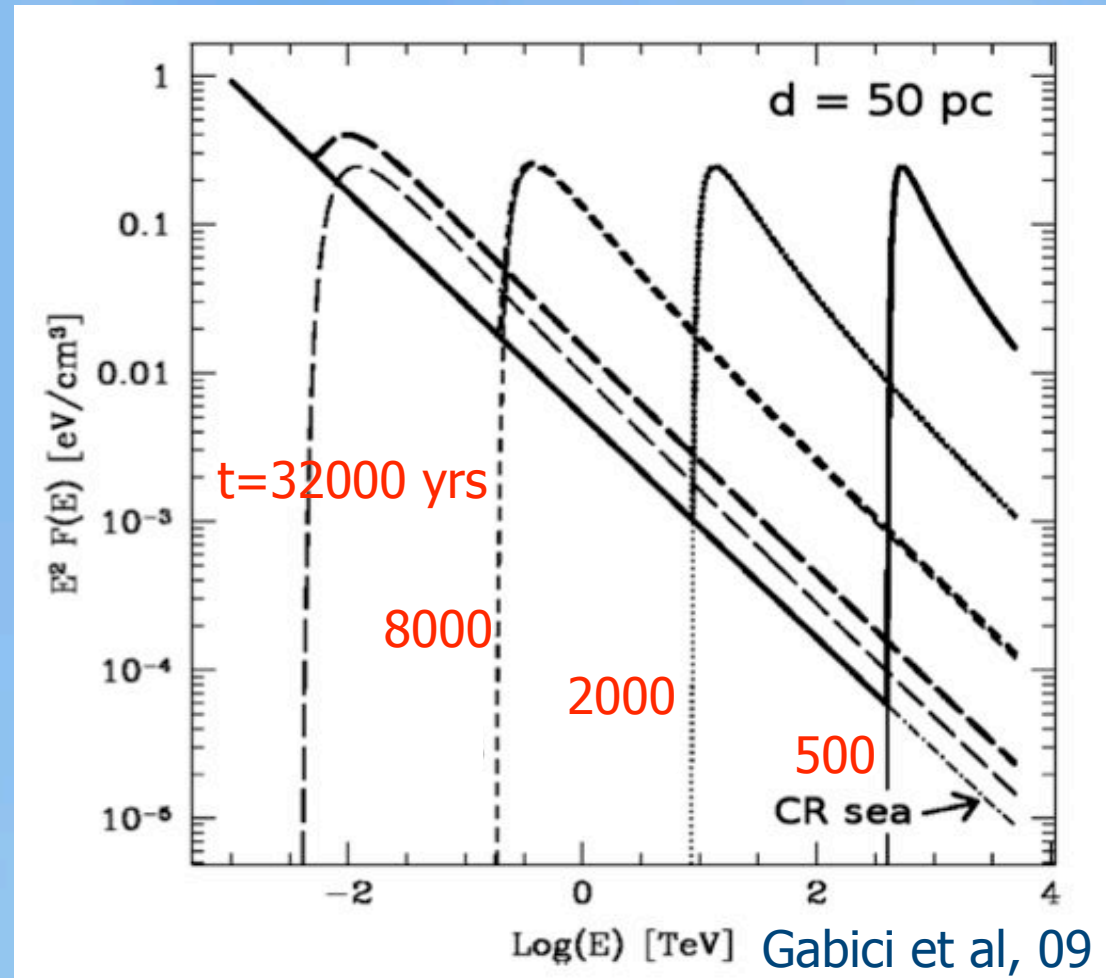
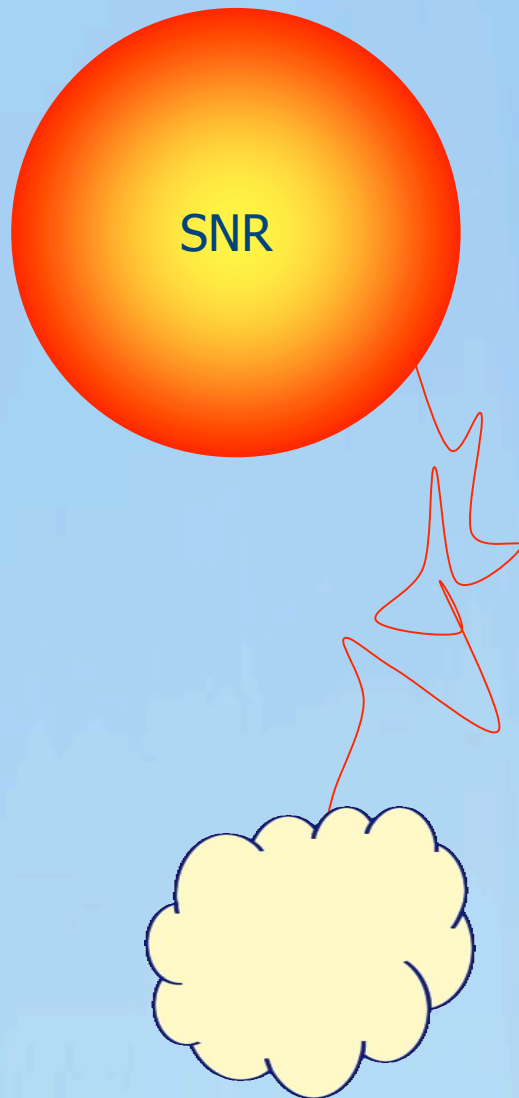
"Diffusive" interaction

not due to CR sea



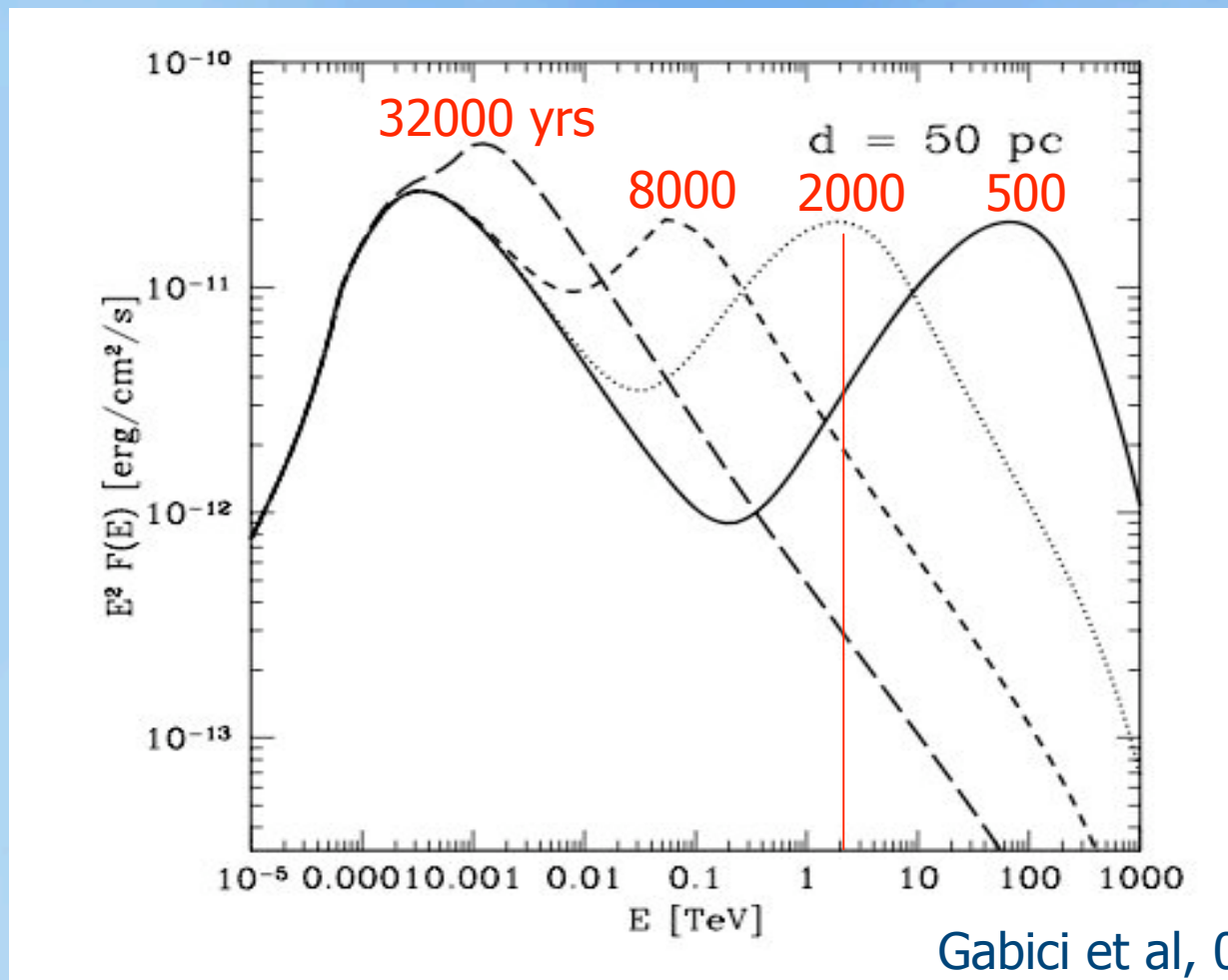
Example : Galactic center

Interaction of CRs with MCs



- ✓ Spectrum of Cosmic rays escaped from an SNR of age t

Interaction of CRs with MCs

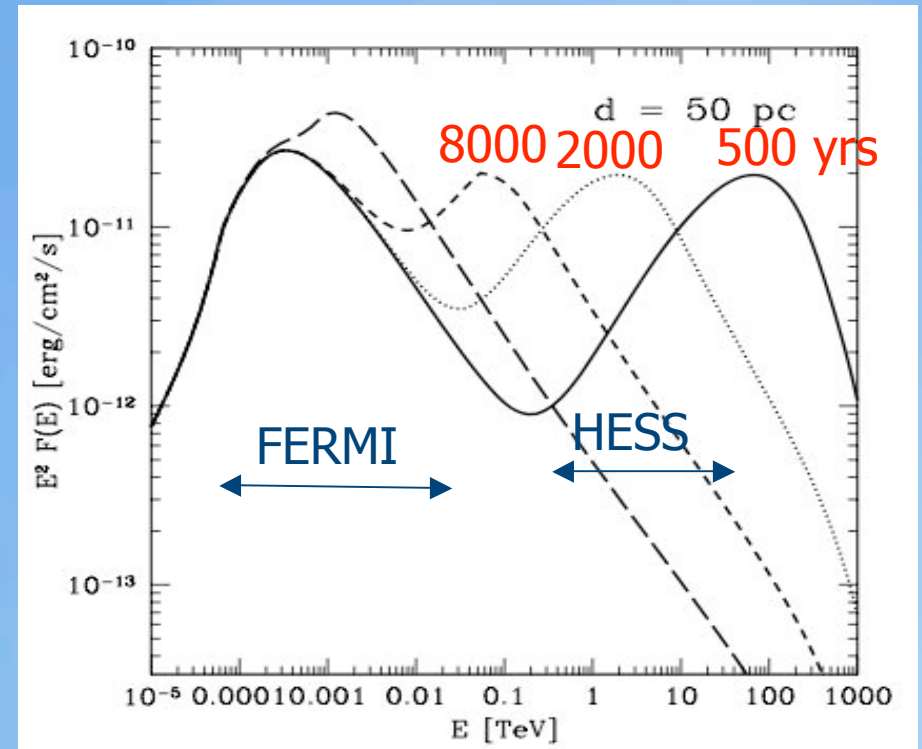


- ✓ Spectrum of the photons emitted by the MC

Molecular-Clouds/SNR

Little How-to

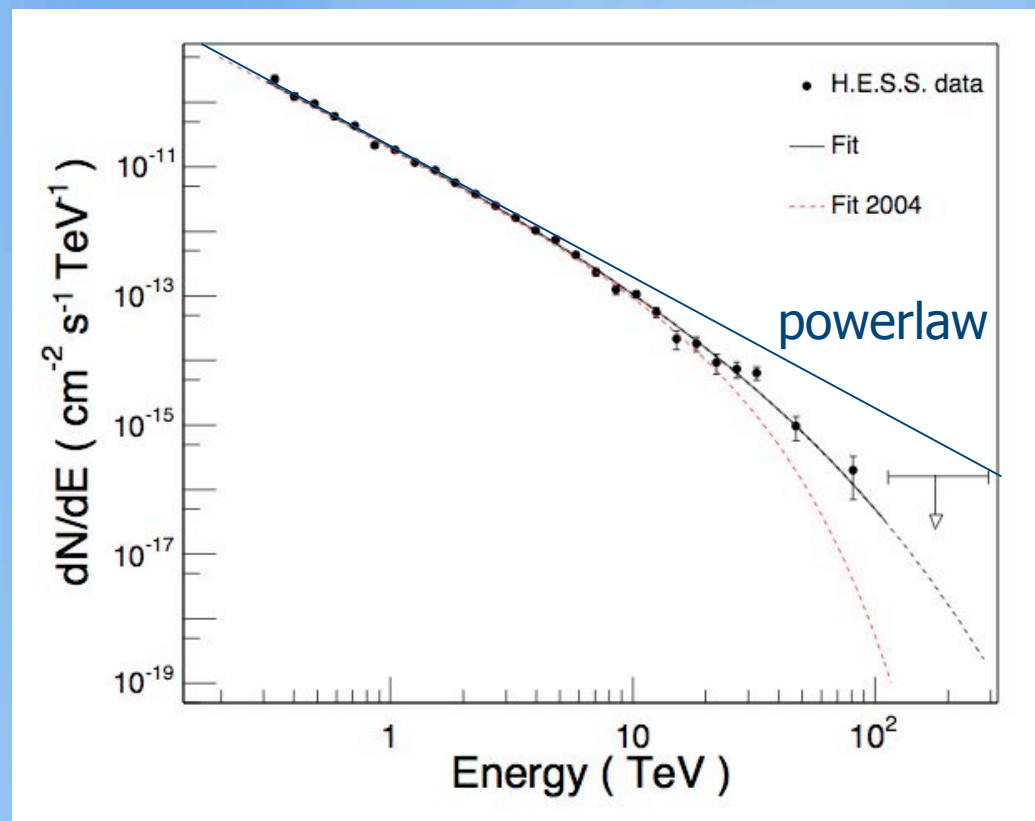
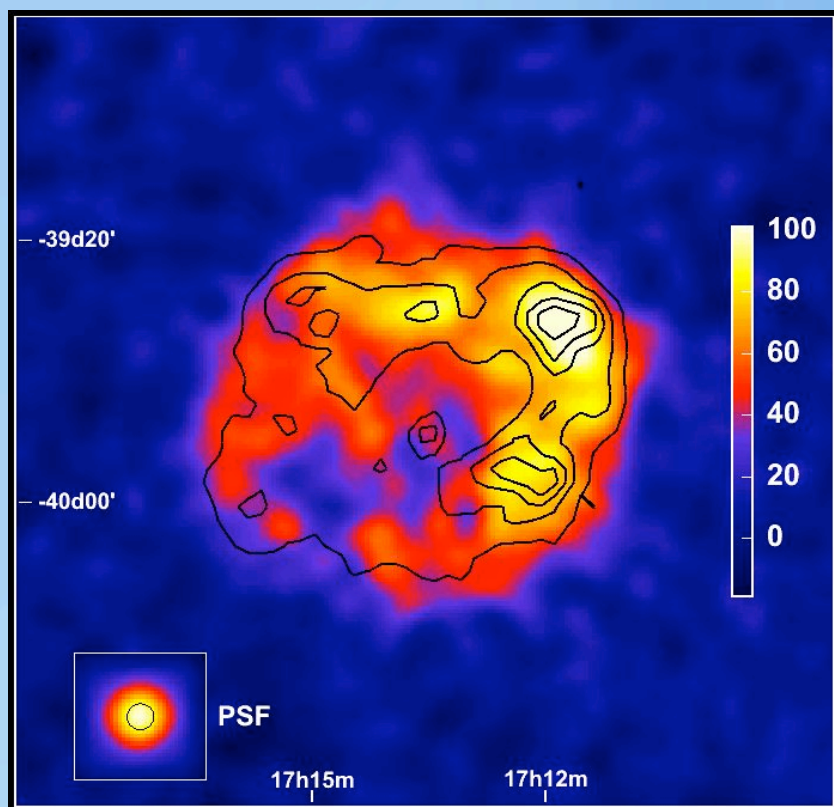
- ✓ Good candidates : dark sources
- ✓ CO counterpart at $d < 100$ pc from the SNR
- ✓ Young SNR (< 5000 yrs) for a HESS detection
- ✓ SNR with detected spectral cutoff is a plus
- ✓ Specific spectrum shape of the cloud hard spectrum with cutoff
- ✓ FERMI counterpart for MC



Good candidate : RXJ 1713

Good Example

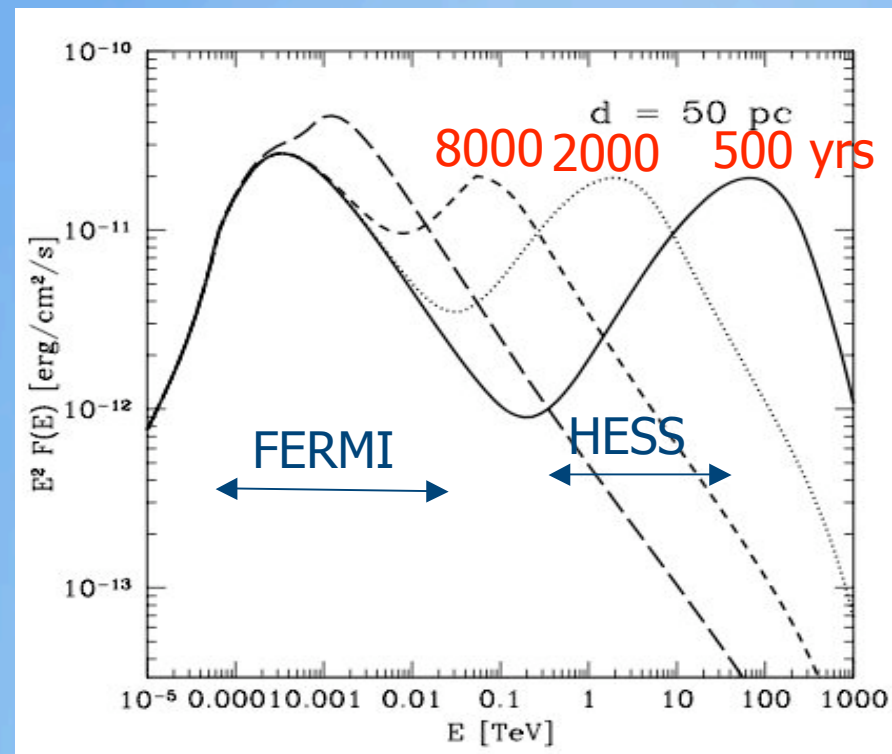
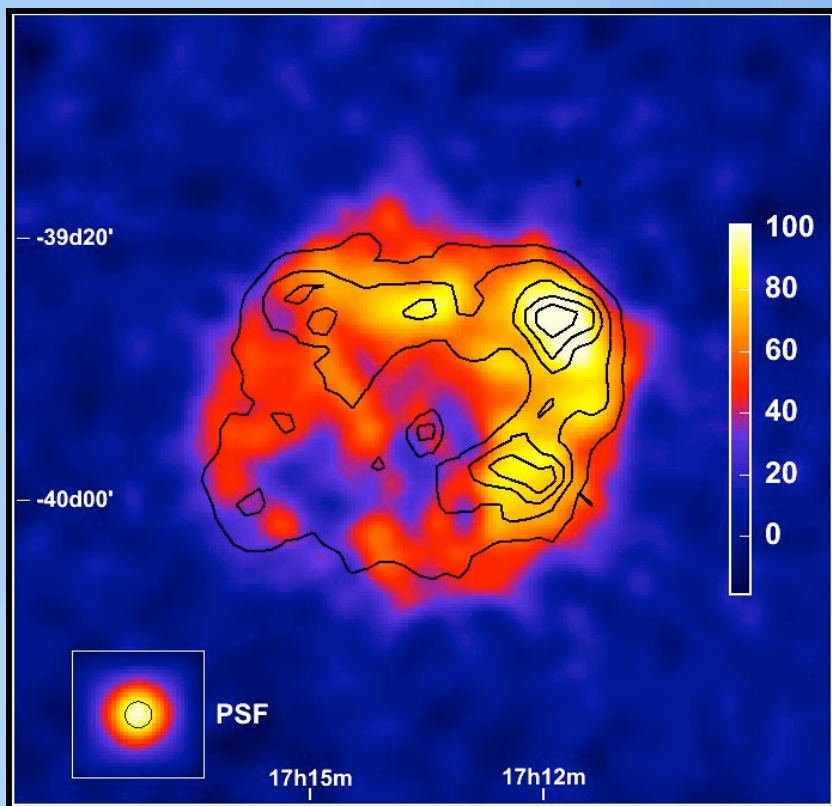
RX J1713



Aharonian et al 07

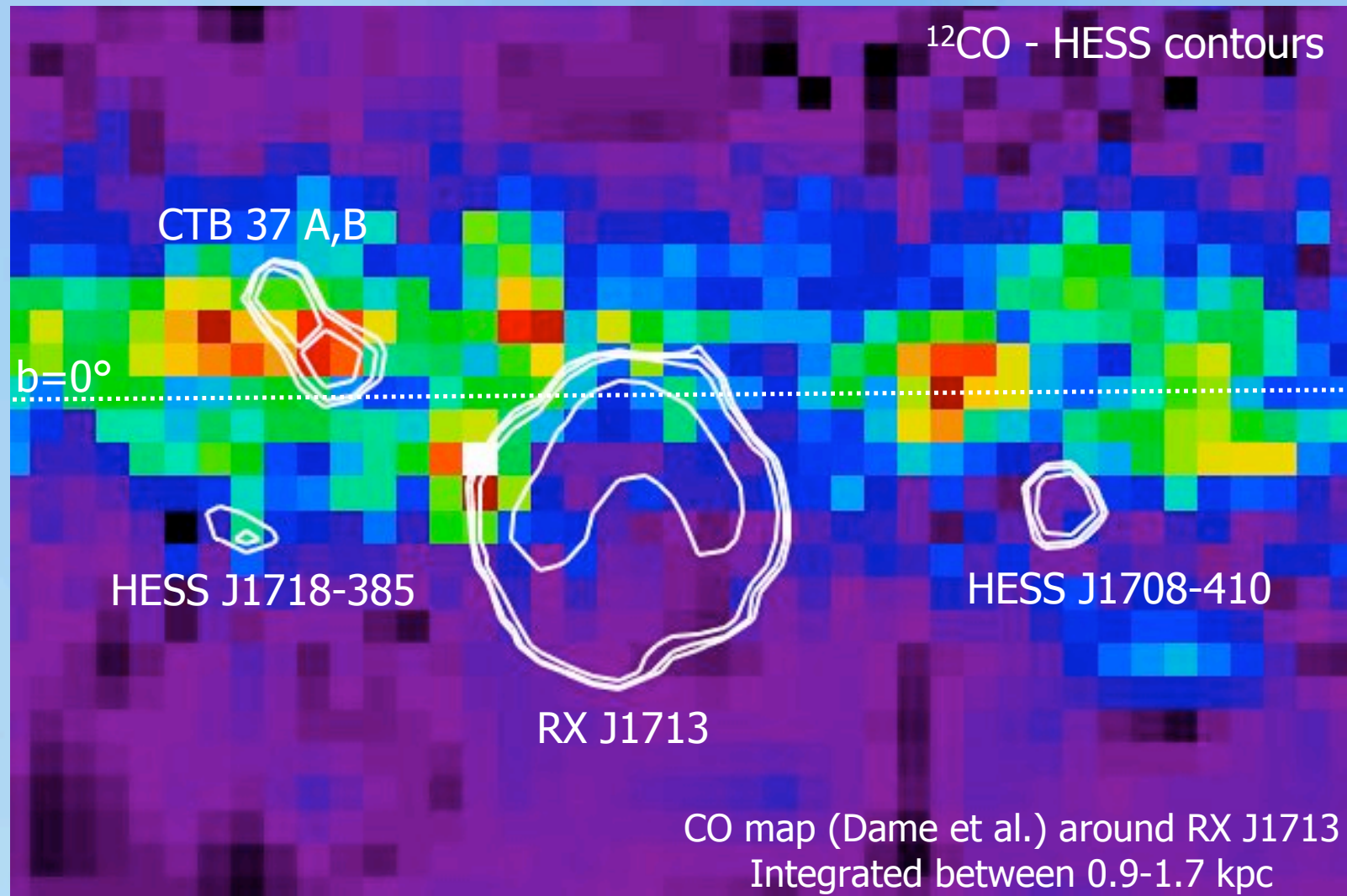
Good Example

RX J1713



Molecular clouds ?

^{12}CO



CRs would diffuse preferentially along the magnetic fields lines

A new shell type SNR

Unidentified sources

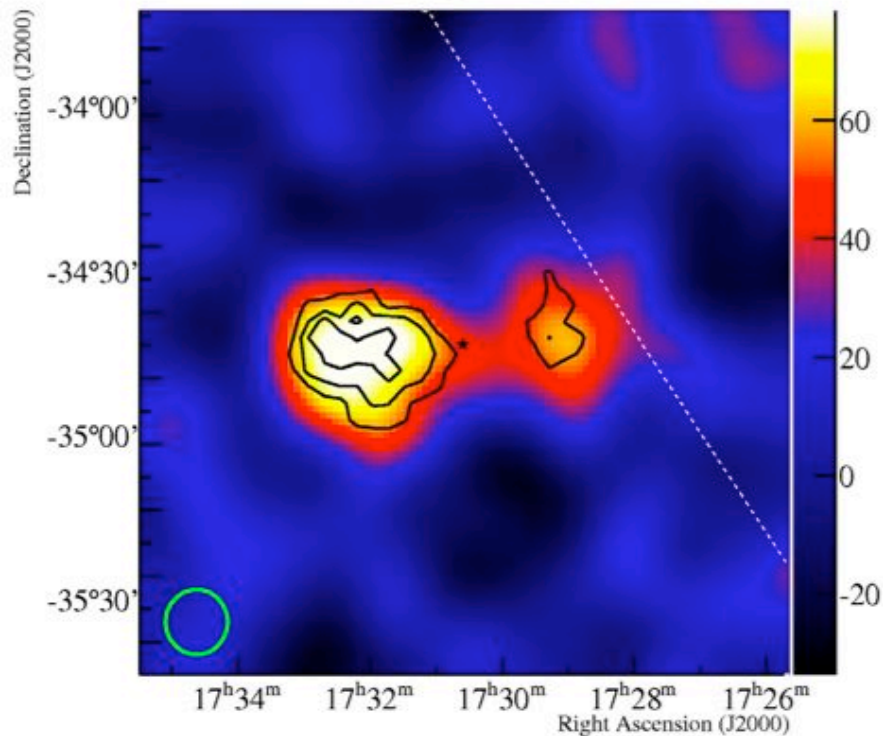
HESS J1731-347

A&A 477, 353–363 (2008)
DOI: 10.1051/0004-6361:20078516
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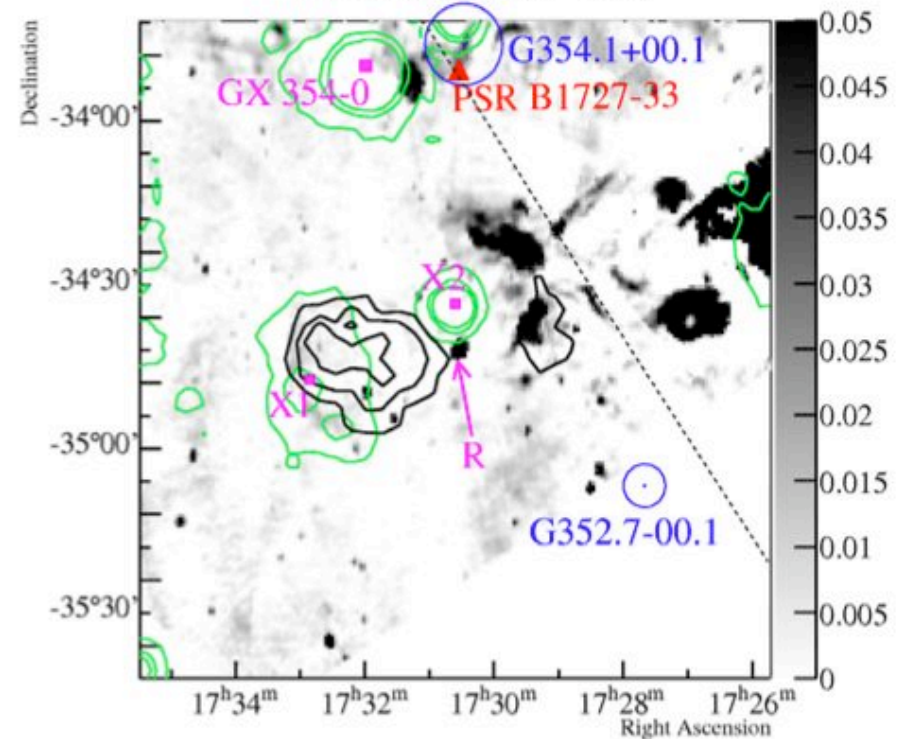
Astronomy
&
Astrophysics

HESS very-high-energy gamma-ray sources without identified counterparts

HESS J1731-347



HESS J1731-347



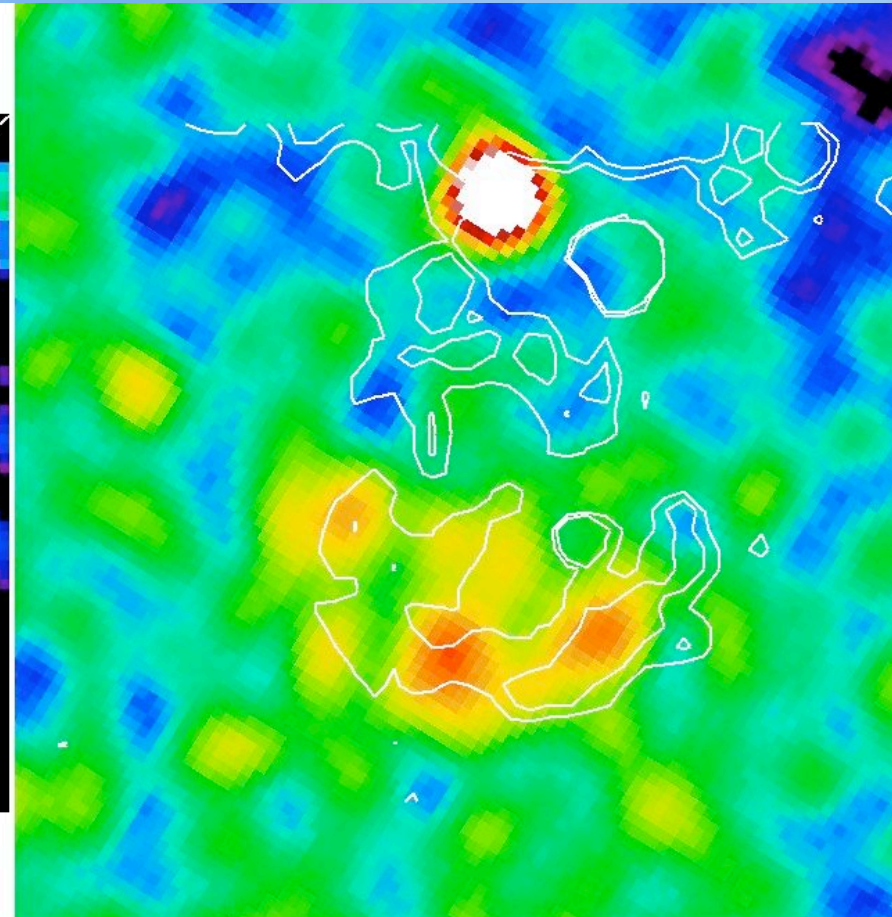
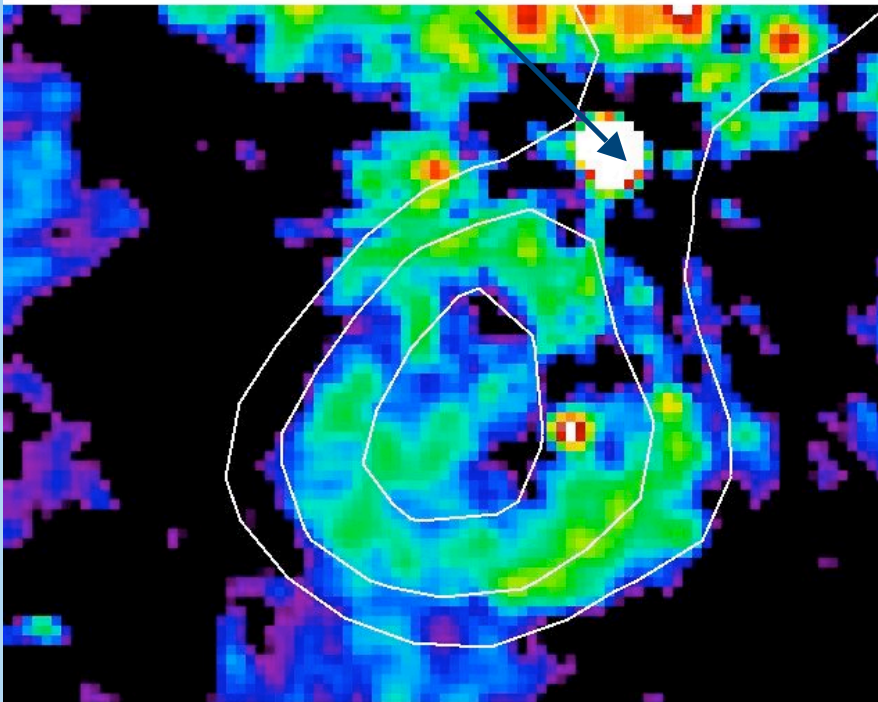
HESS J1731-347

ROSAT

ATCA Radio + HESS contours

ROSAT + ATCA radio contours

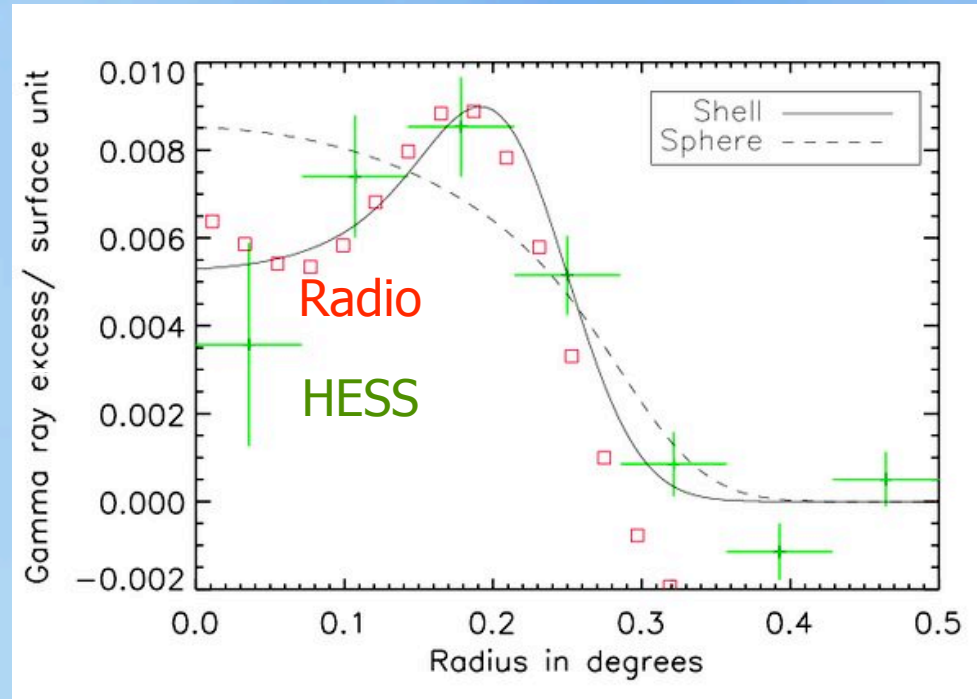
HII region at 3.2 ± 0.8 kpc



HESS J1731-347

HESS

Radial profiles



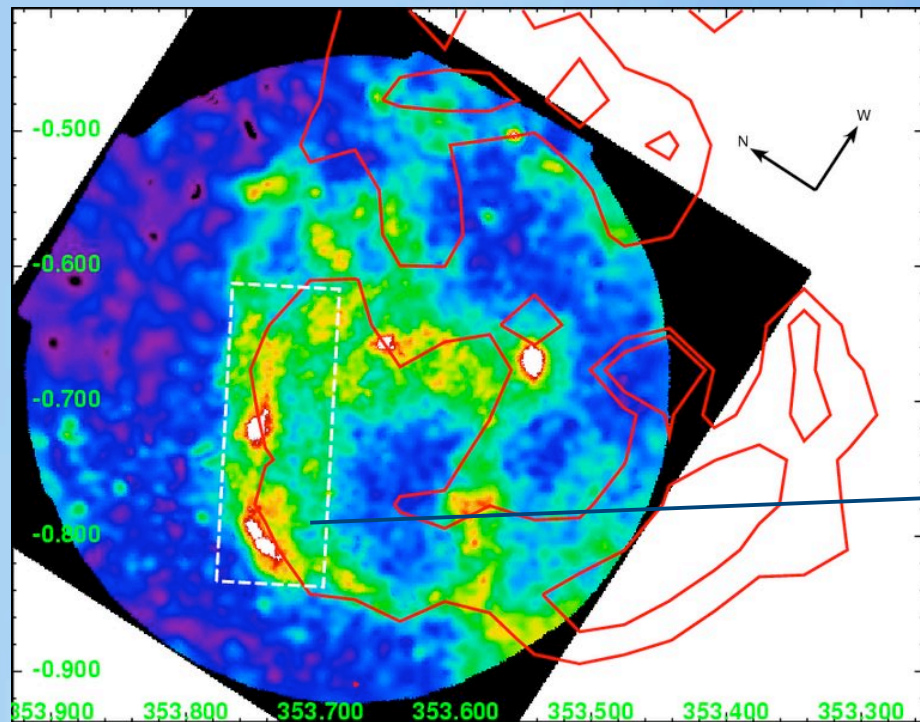
Acero et al '09, arXiv:0907.0642

- *Spatially coincident with radio shell
- *Indication of shell like morphology
- Not statistically significant (yet)

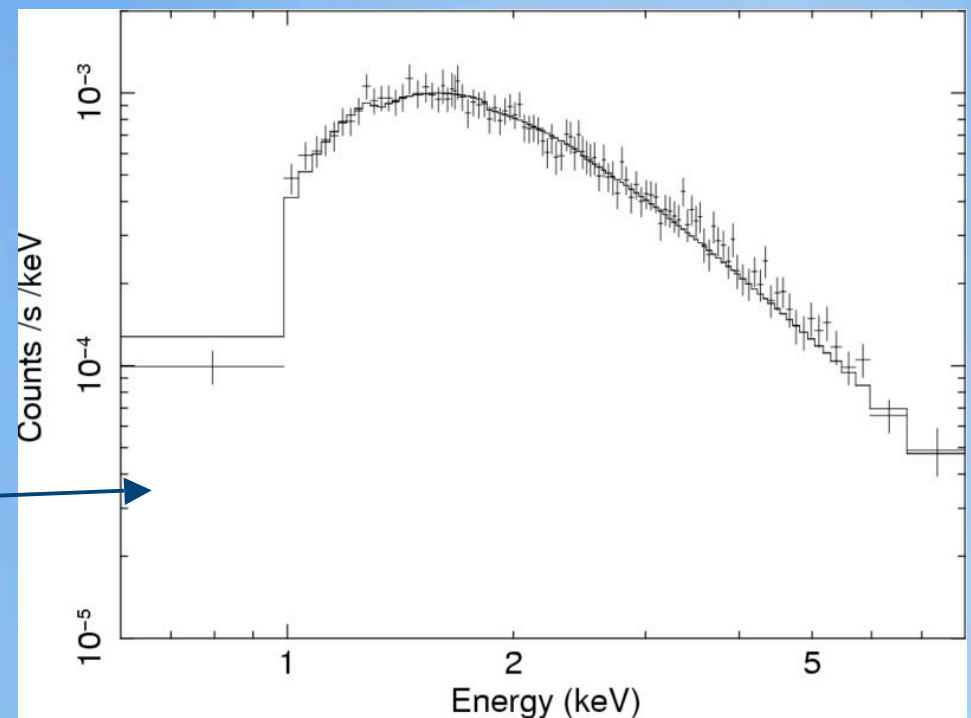
HESS J1731-347

Non thermal shell !

XMM + ATCA radio contours



MOS instrument spectrum



Acero et al '09

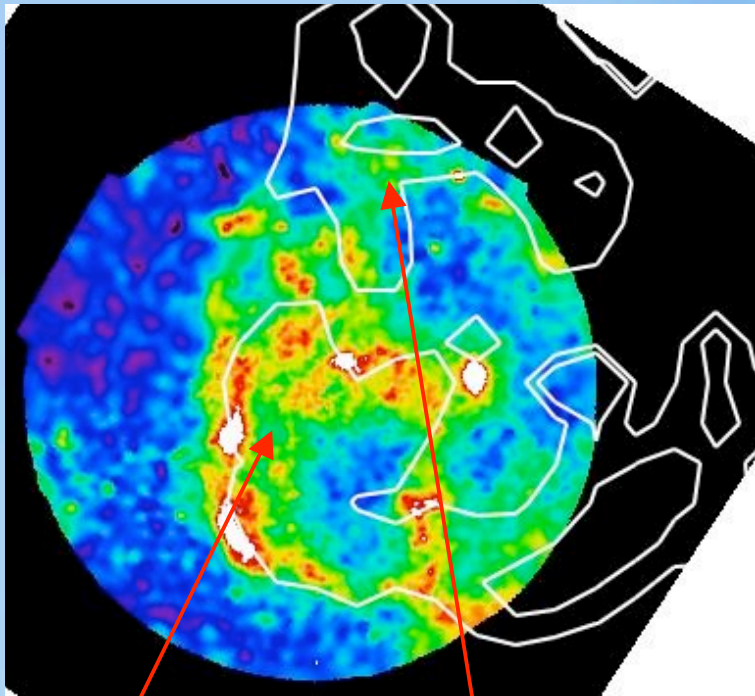
SNR's filament well fitted by an absorbed powerlaw

X-ray synchrotron emission \rightarrow e^- accelerated to ~ 10 TeV

12CO study

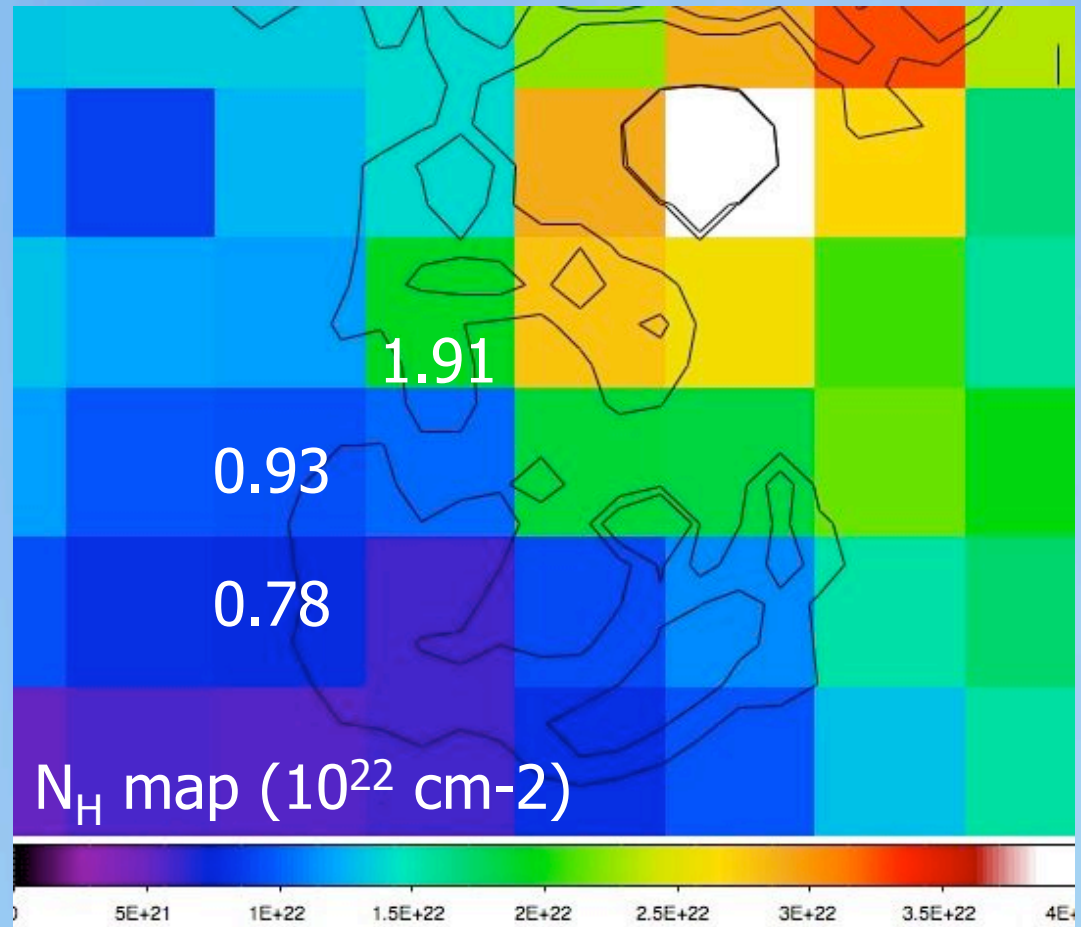
Spectra

XMM
ATCA contours



1.05 (1.01-1.11) 1.96 (1.75-2.20)
 $N_H(10^{22} \text{ cm}^{-2})$

12CO
ATCA contours



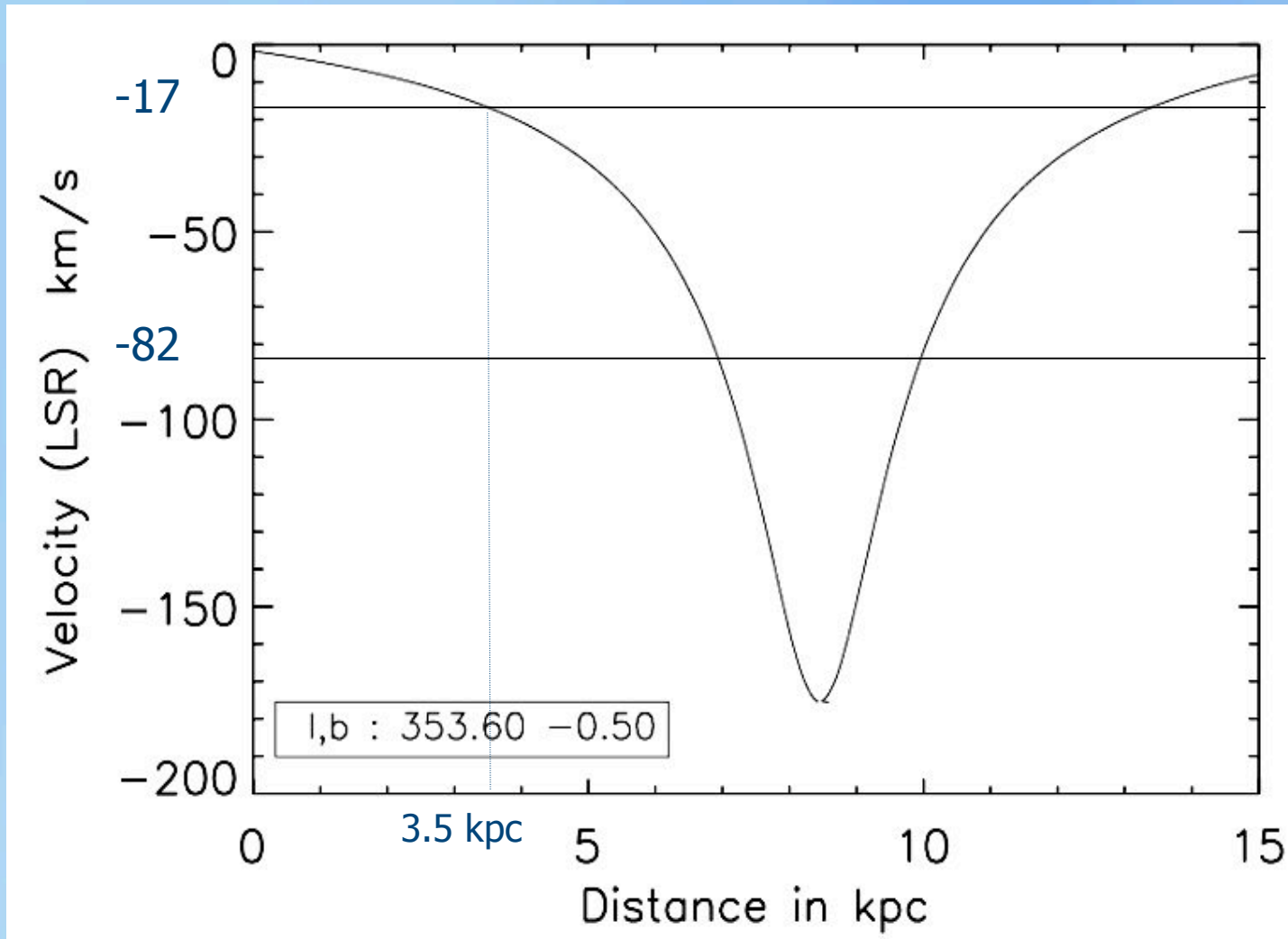
N_H map (10^{22} cm^{-2})

12 CO Integrated between 0->-17 km/s

12CO study Distance

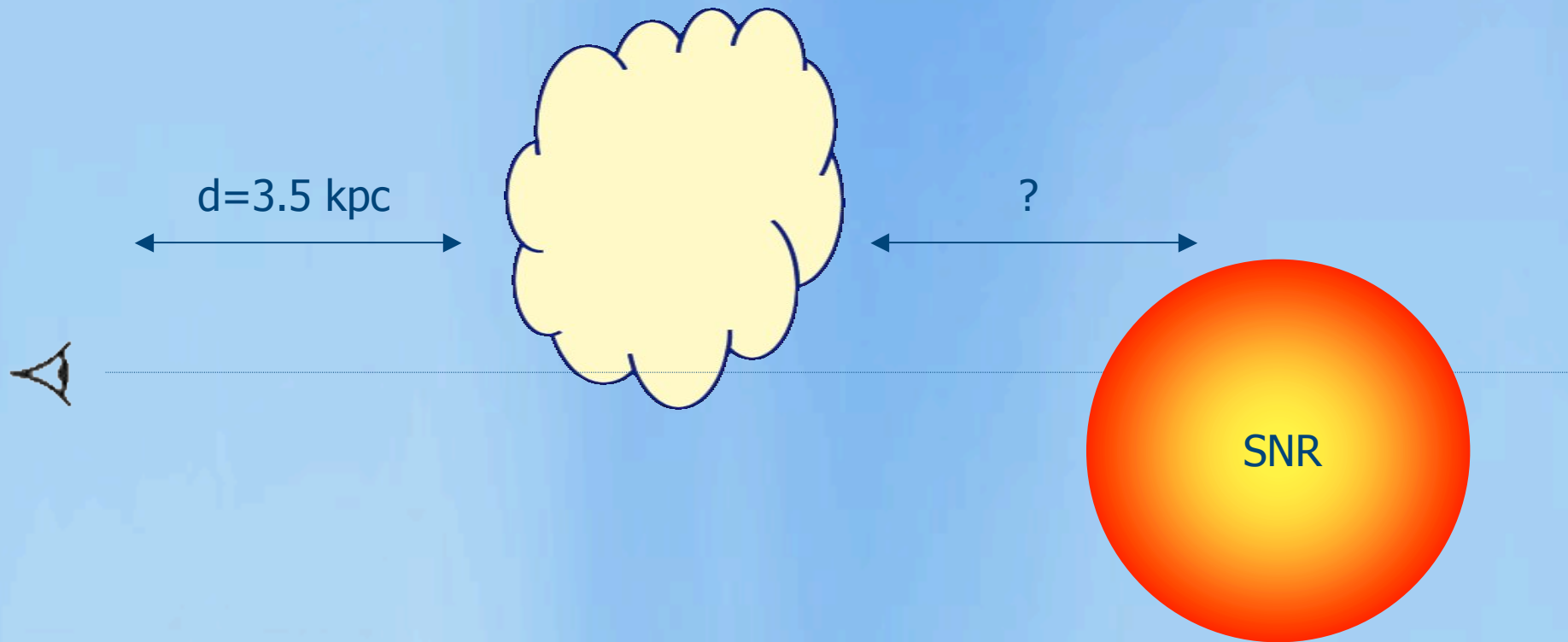
**Lower Limit
of 3.5 kpc**

Errors of +/- 0.5 kpc
using different
Galactic rotation
Models.



Galactic Rotation : Fich et al. 89

Cartoon



J1731 : Conclusion

- ✓ Gamma-ray Morphology suggesting a shell like remnant
-> Recent observations will hopefully confirm this
- ✓ Shell structure seen in radio and X-rays
- ✓ First shell SNR discovered based on Gamma observations !!!
- ✓ J1731 could join the club of RX J1713, Vela Jr, RCW 86 and SN 1006
- ✓ No cutoff in spectrum (Aharonian 08)
-> Recent observations will test the spectrum with better statistics
- ✓ Molecular Cloud in the vicinity

-> Direct/Diffusive/no interaction

Conclusion

- ✓ Looking for a methodology to detect emission from MC
 - > Look around young SNR
 - > Cutoff in the spectrum's SNR is a proof of escaping CRs
 - > Specific spectrum shape
- ✓ RX J1713 -> Good example but nothing seen so far
- ✓ New interesting candidate -> HESS J1731
 - > MC in the vicinity

Thank you !