



Deblending galaxies with Variational Autoencoder: a multi-bands, multi-instruments analysis

(Arcelin, Doux, Roucelle, Aubourg)

Bastien Arcelin
05/06/2019

LSST France - Juin 2019

LSST Data

- Will look a lot like HSC data (hsc-release.mtk.nao.ac.jp/hscMap2/)
- For HSC 58% of the detected objects are identified as blended*
- Systematic in shear measurement



From Cyrille Doux
(LSST Europe, 14 Juin 2018)

* Bosch et al. (2017)

Goals and motivation

- Create a model for galaxy images from the data



Generative model

- Bayesian approach



Variational AutoEncoder

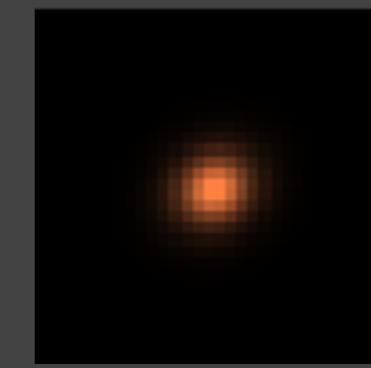
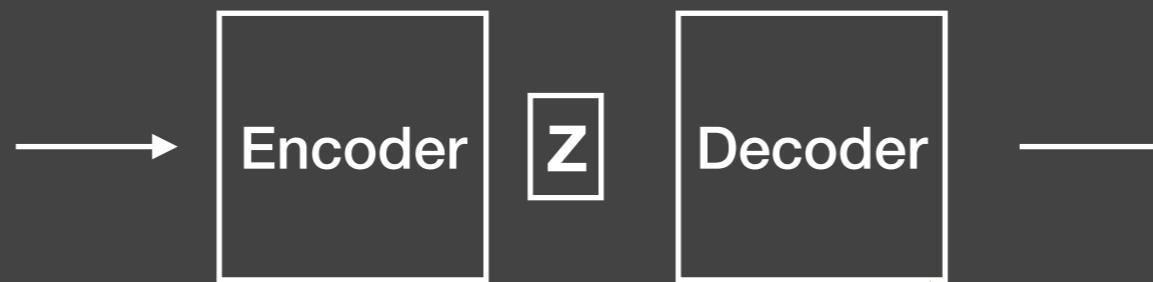
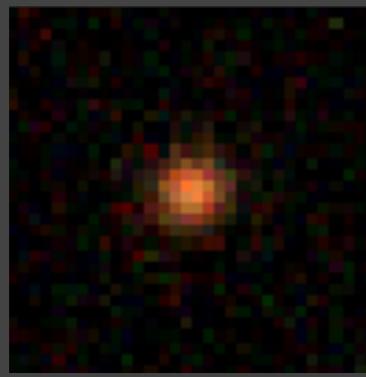
- Accurate reproduction of relevant weak lensing parameters



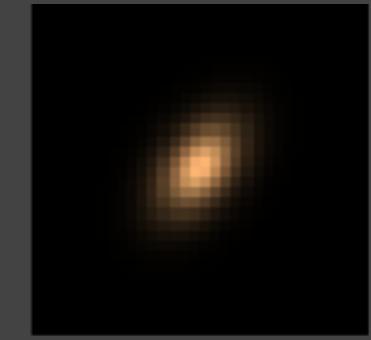
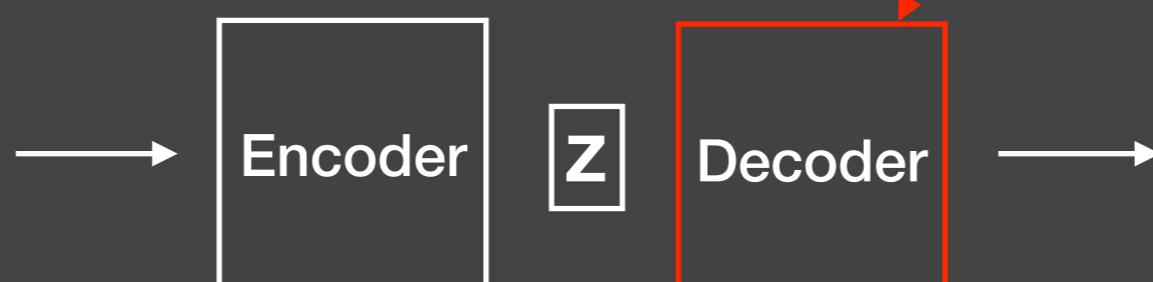
Shape and flux reproduction

Machine learning for deblending

Variational AutoEncoder (VAE) (Kingma+2014)

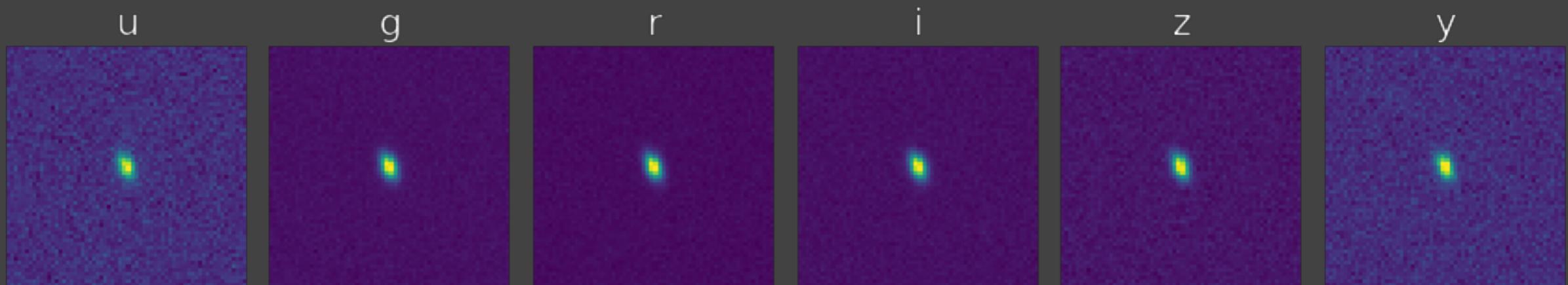


Deblender

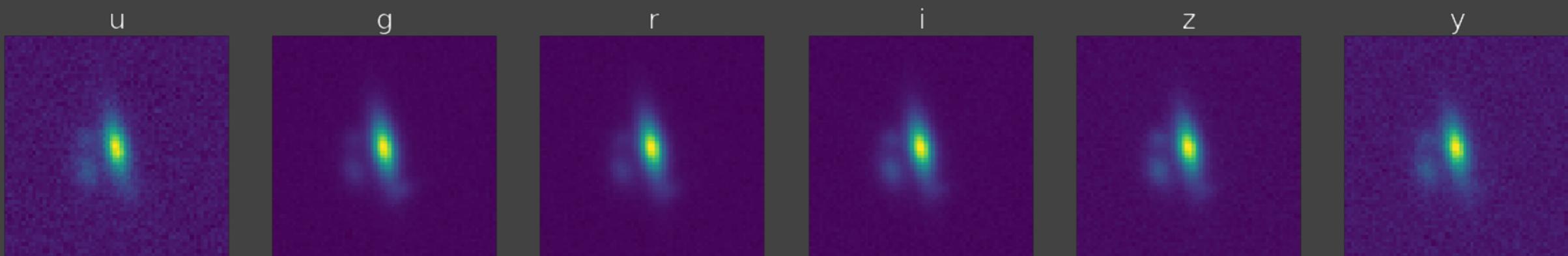


Training sample

- With GalSim from COSMOS catalog: 6 LSST pass-band filters
 - 200k galaxies for the VAE



- 200k blends of galaxies for the deblender

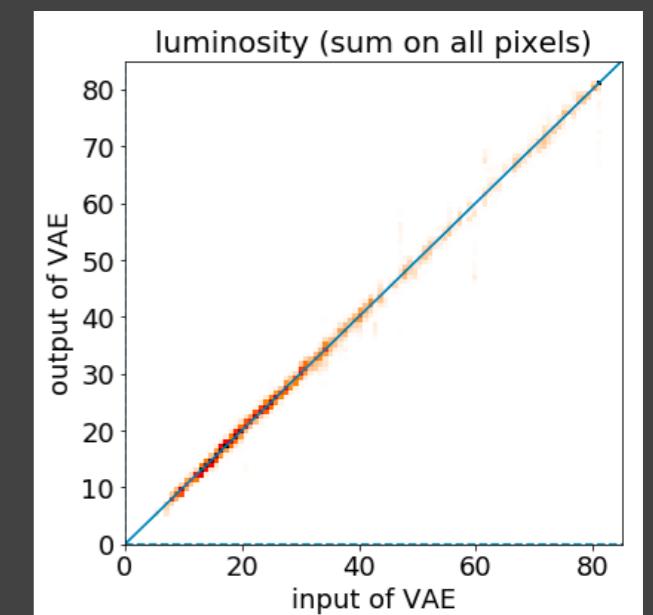
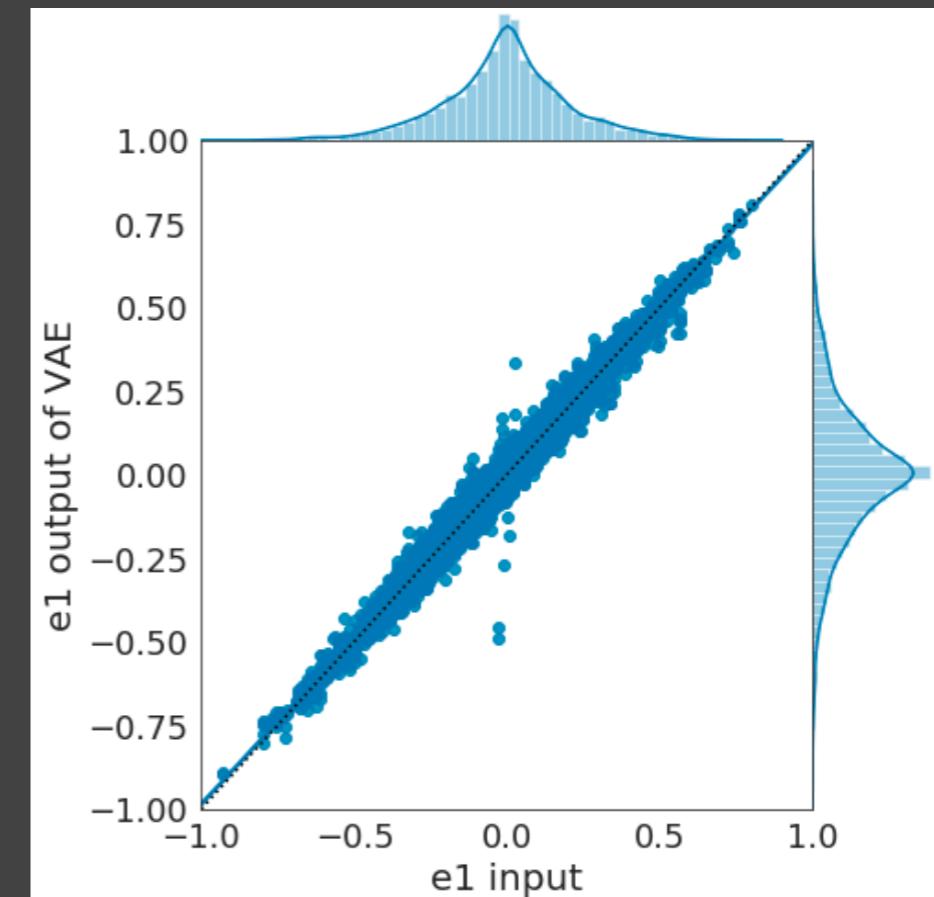
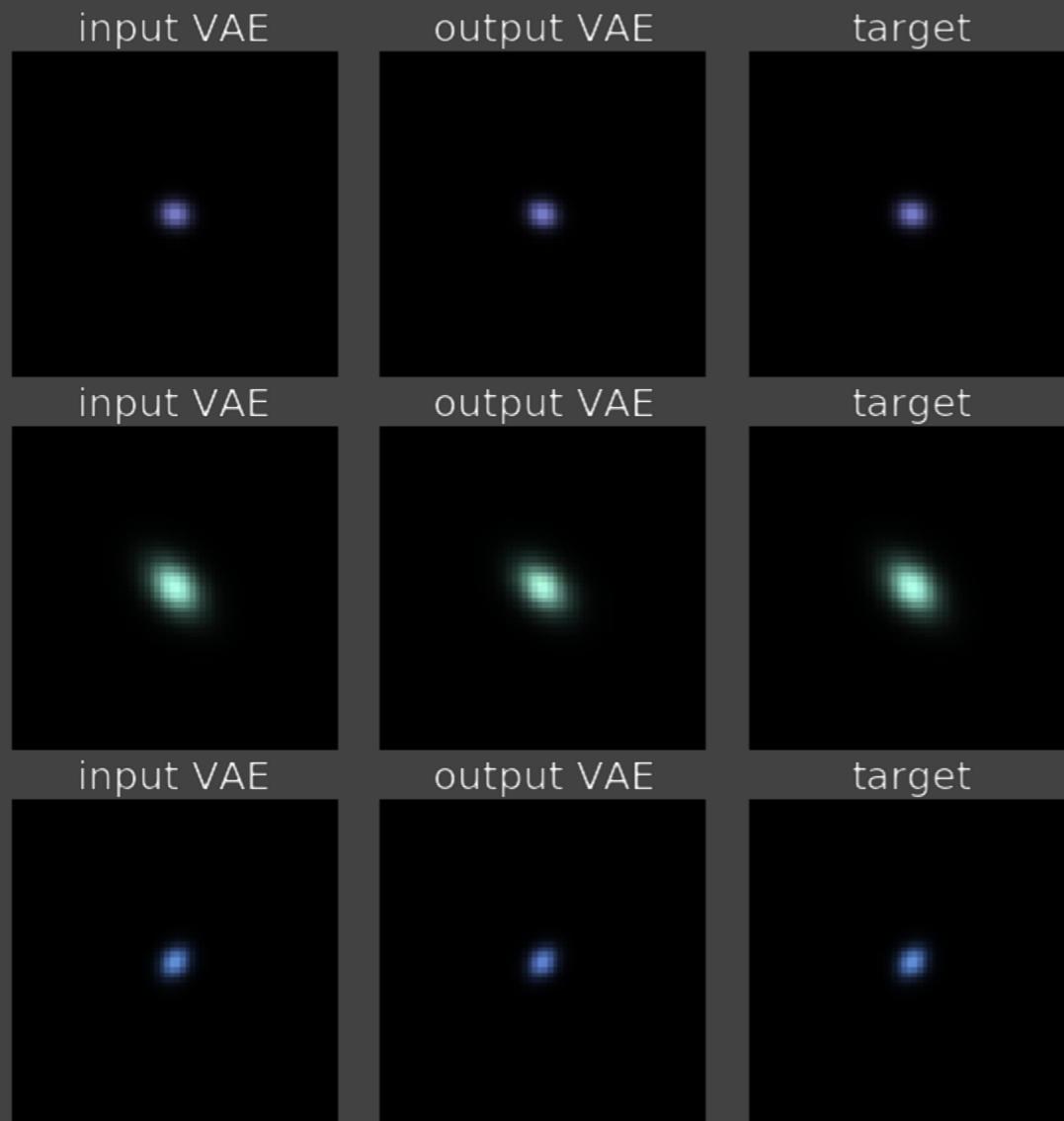


Training

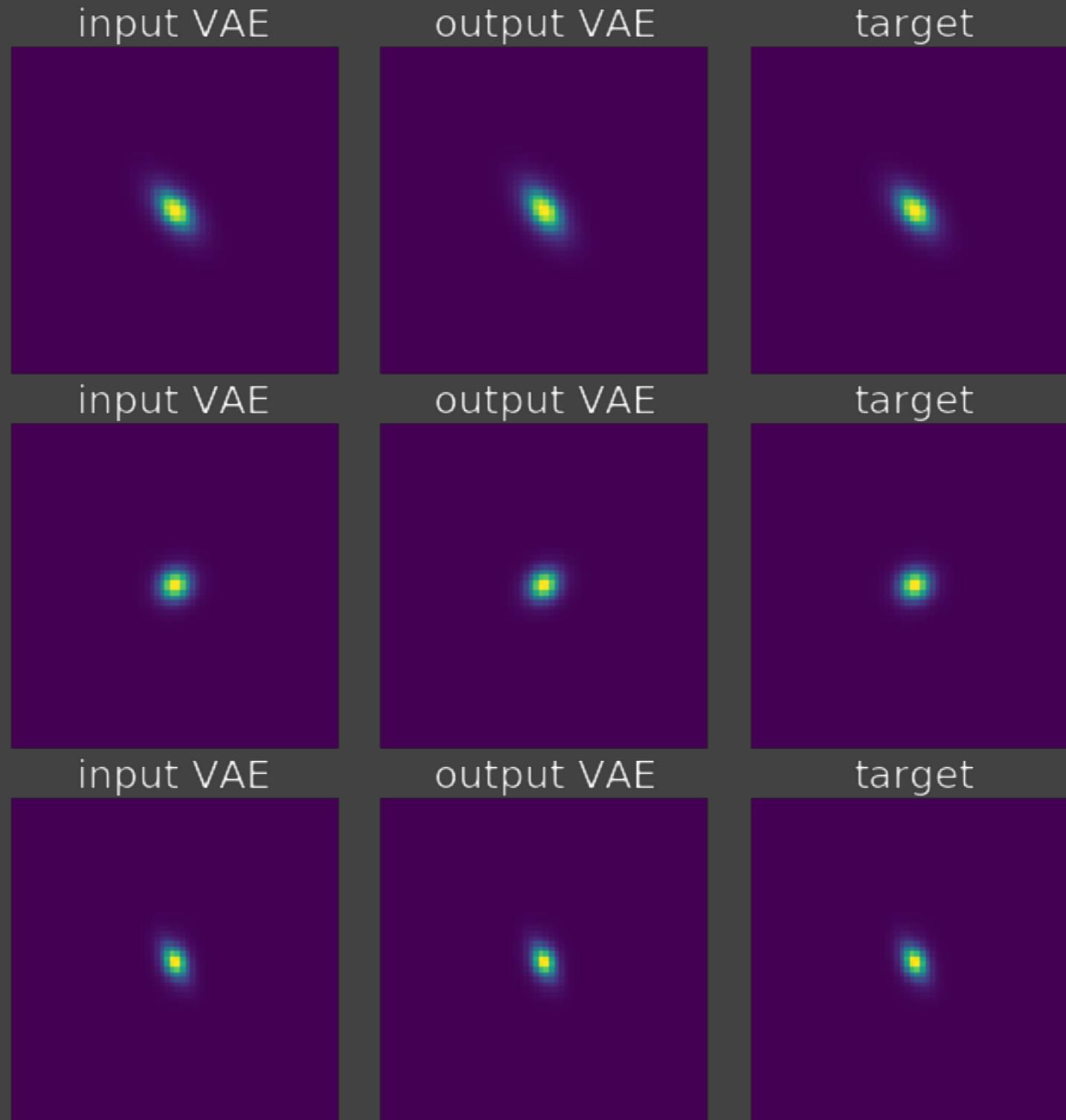
- 3 steps:
 - 1) Train VAE on noiseless images of single galaxy
 - 2) Train on noisy images of single galaxy
 - 3) Train on noisy images of blended galaxies
- 3 cases:
 - 1) Train on images composed of only ***R band-pass filter***
 - 2) Train on images composed of the ***6 LSST band-pass filters***
 - 3) Train on images composed of the ***6 LSST band-pass filters + the 4 Euclid band-pass filters***

Tests

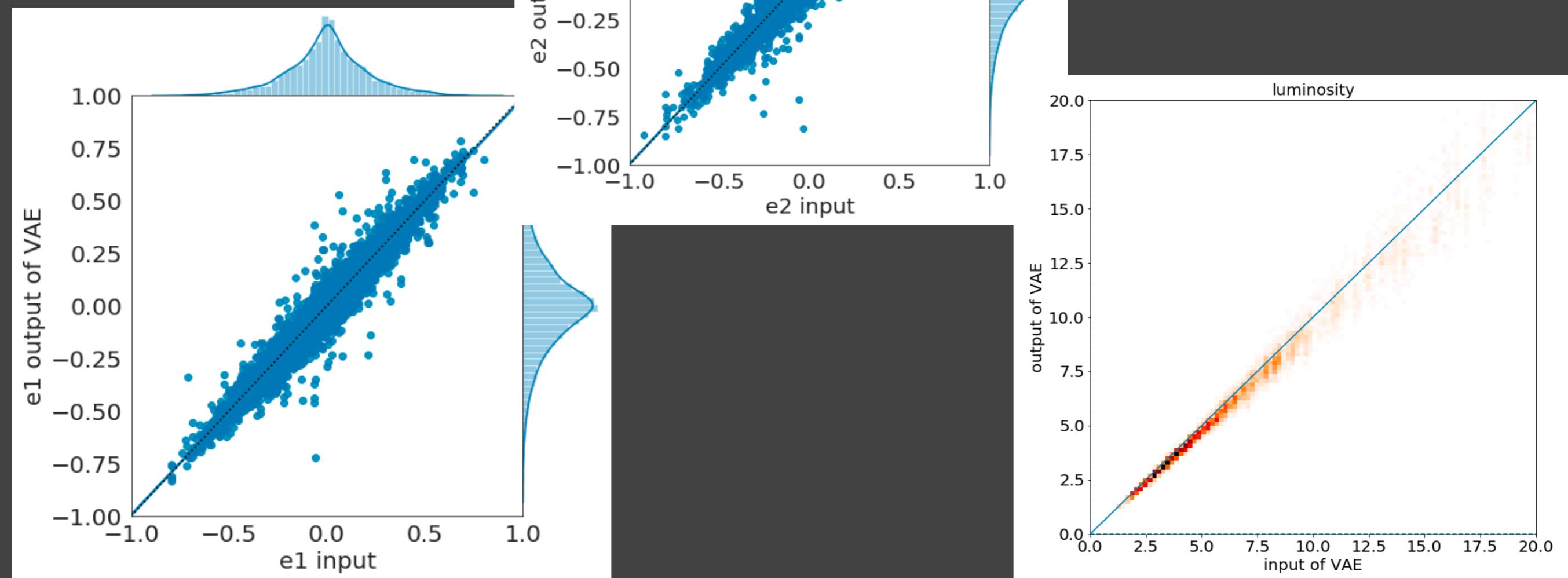
- Visual inspection
- Reproduction of shapes/fluxes



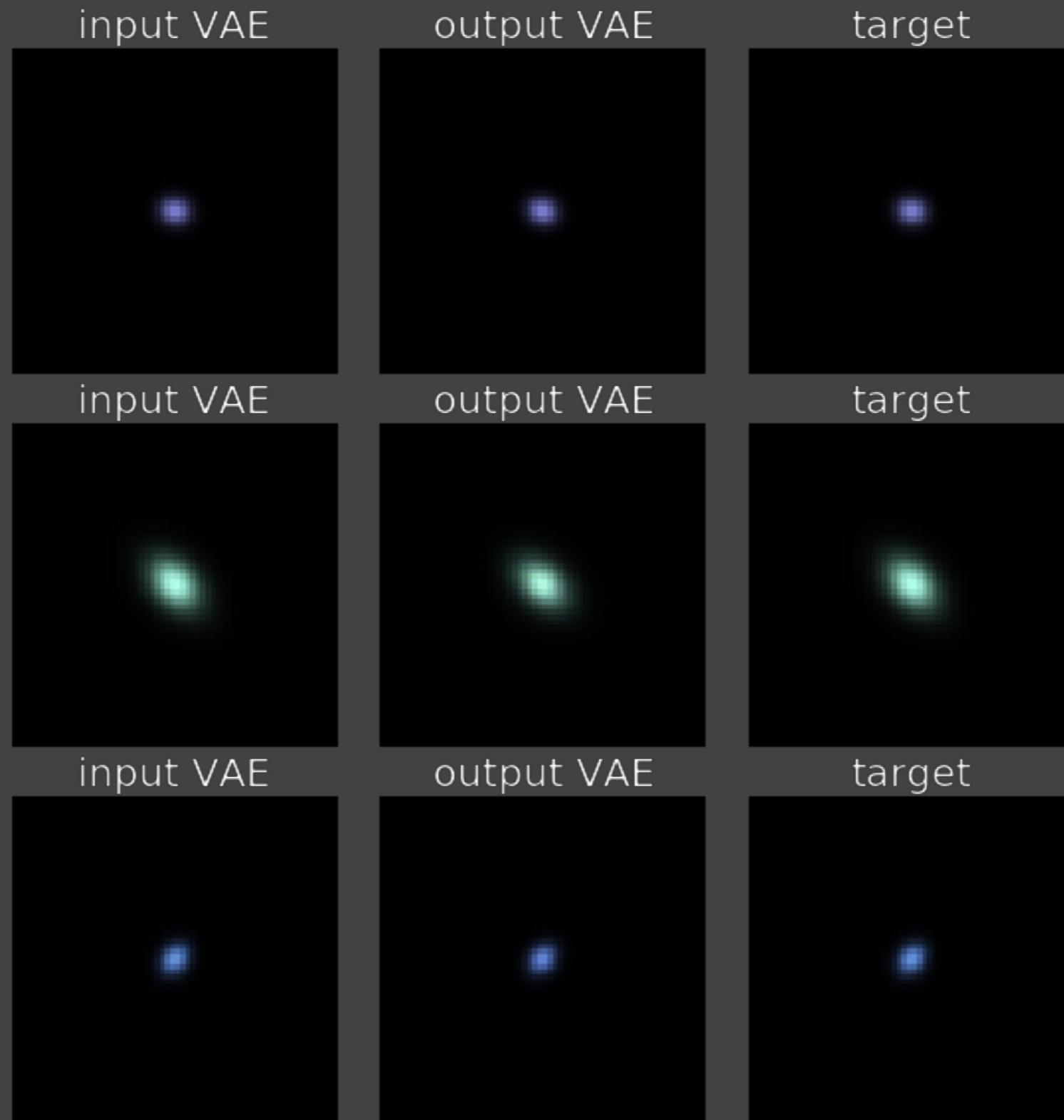
VAE: R band-pass filter



VAE: with LSST R band-pass filter



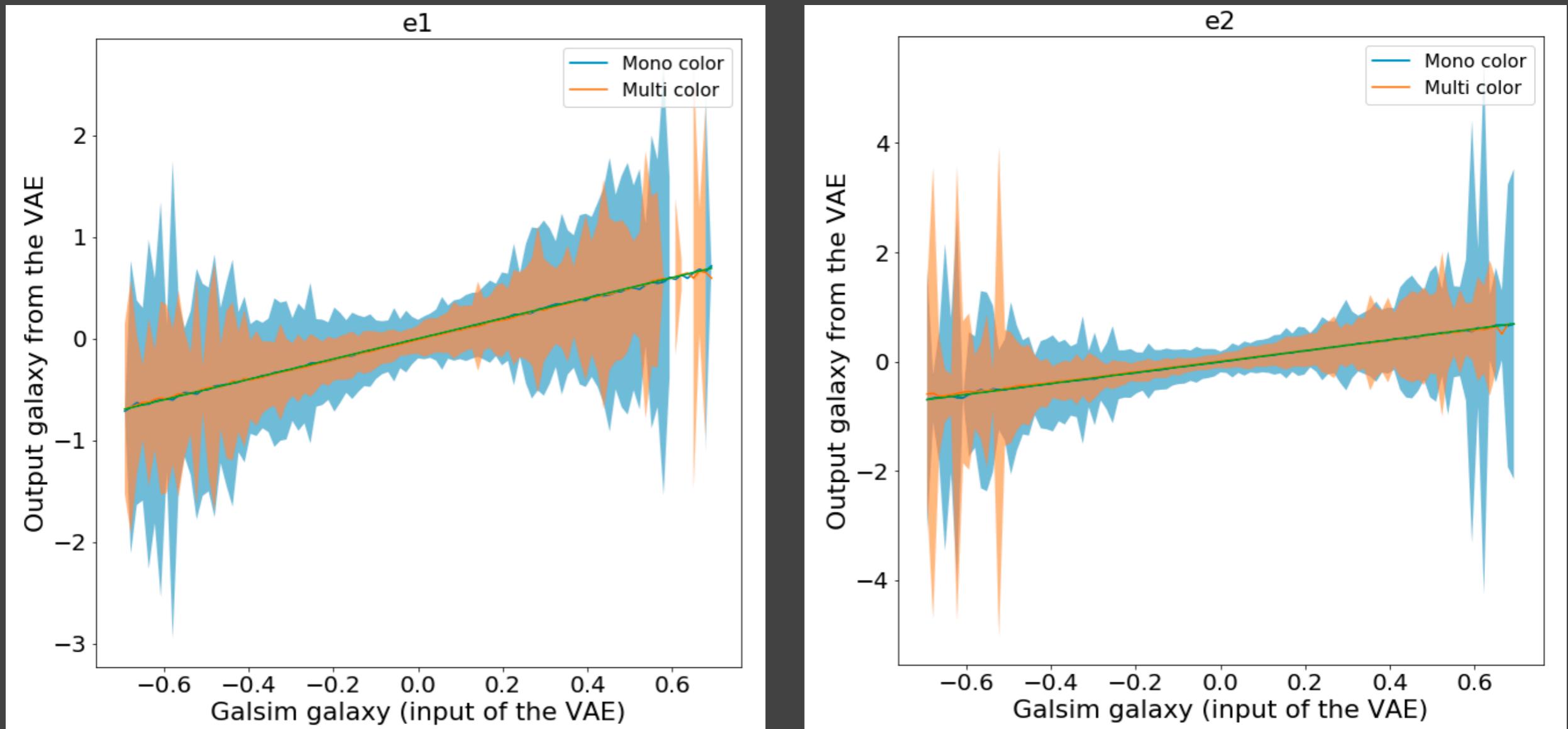
VAE: 6 LSST band-pass filters



VAE

mono-wavelength/multi-wavelengths comparison

- 6 passbands of LSST (comparison done on r band)

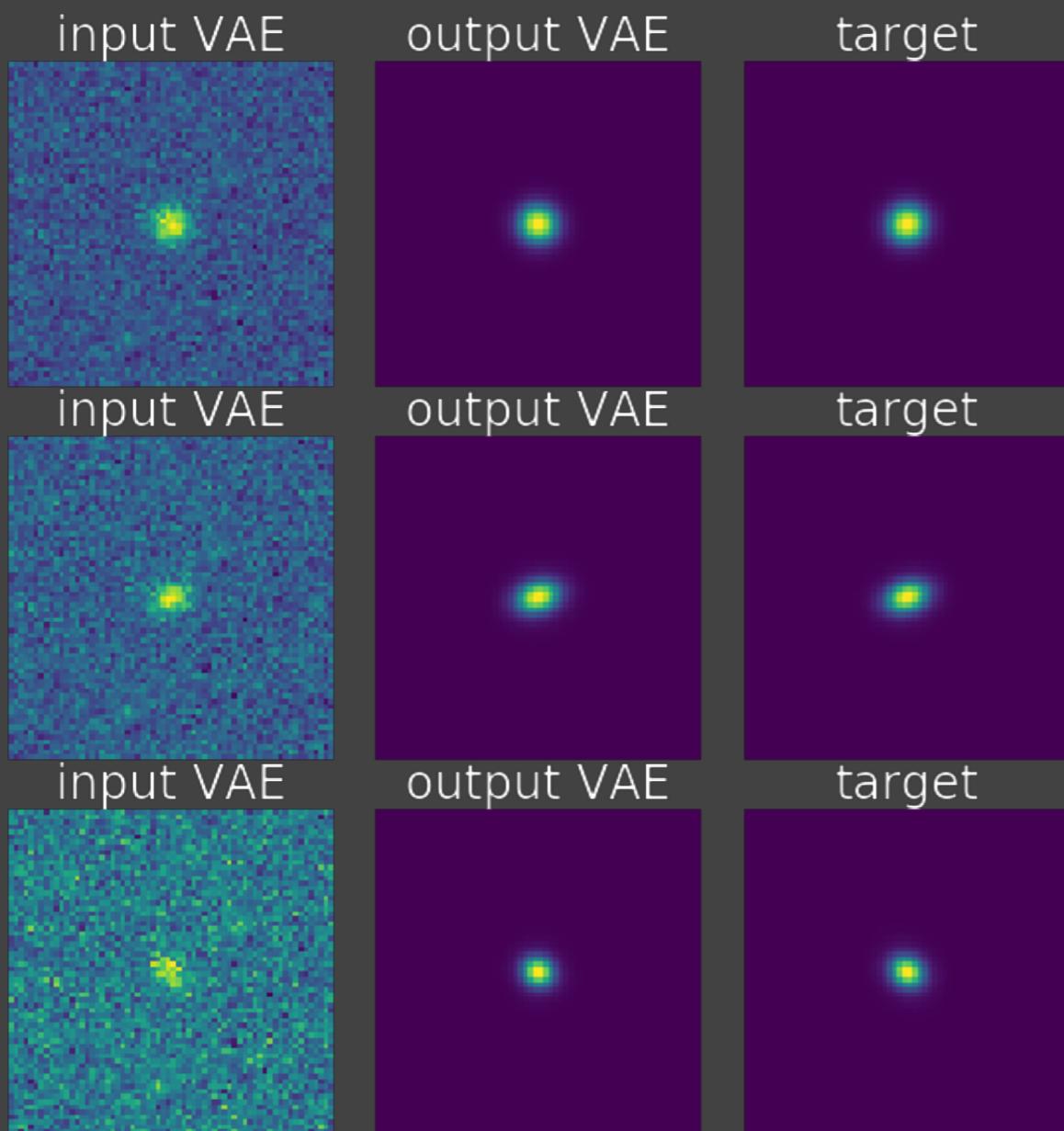


→ VAE gives better results using 6 bands rather than one

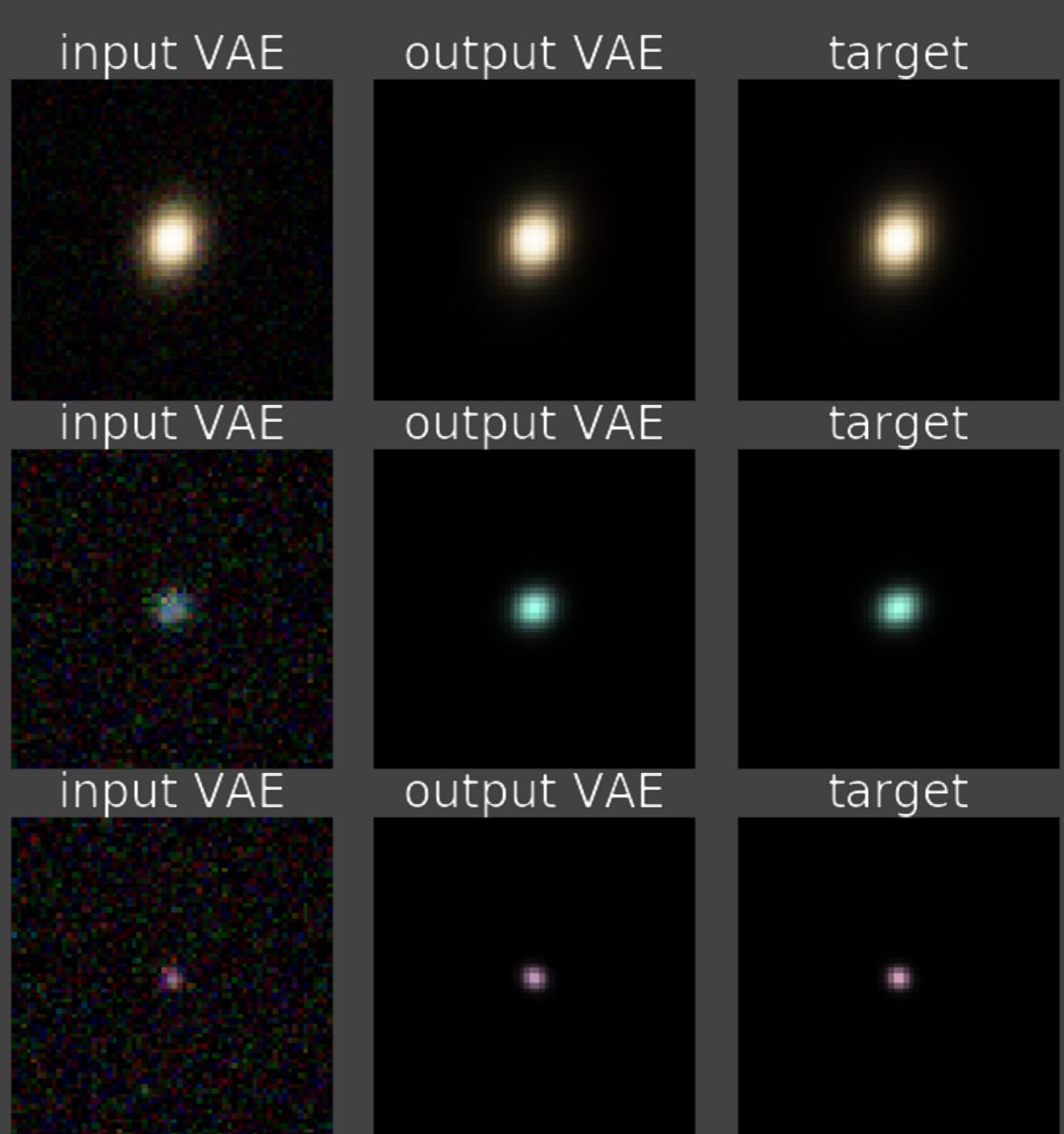
VAE: noisy images

- Poissonian noise added

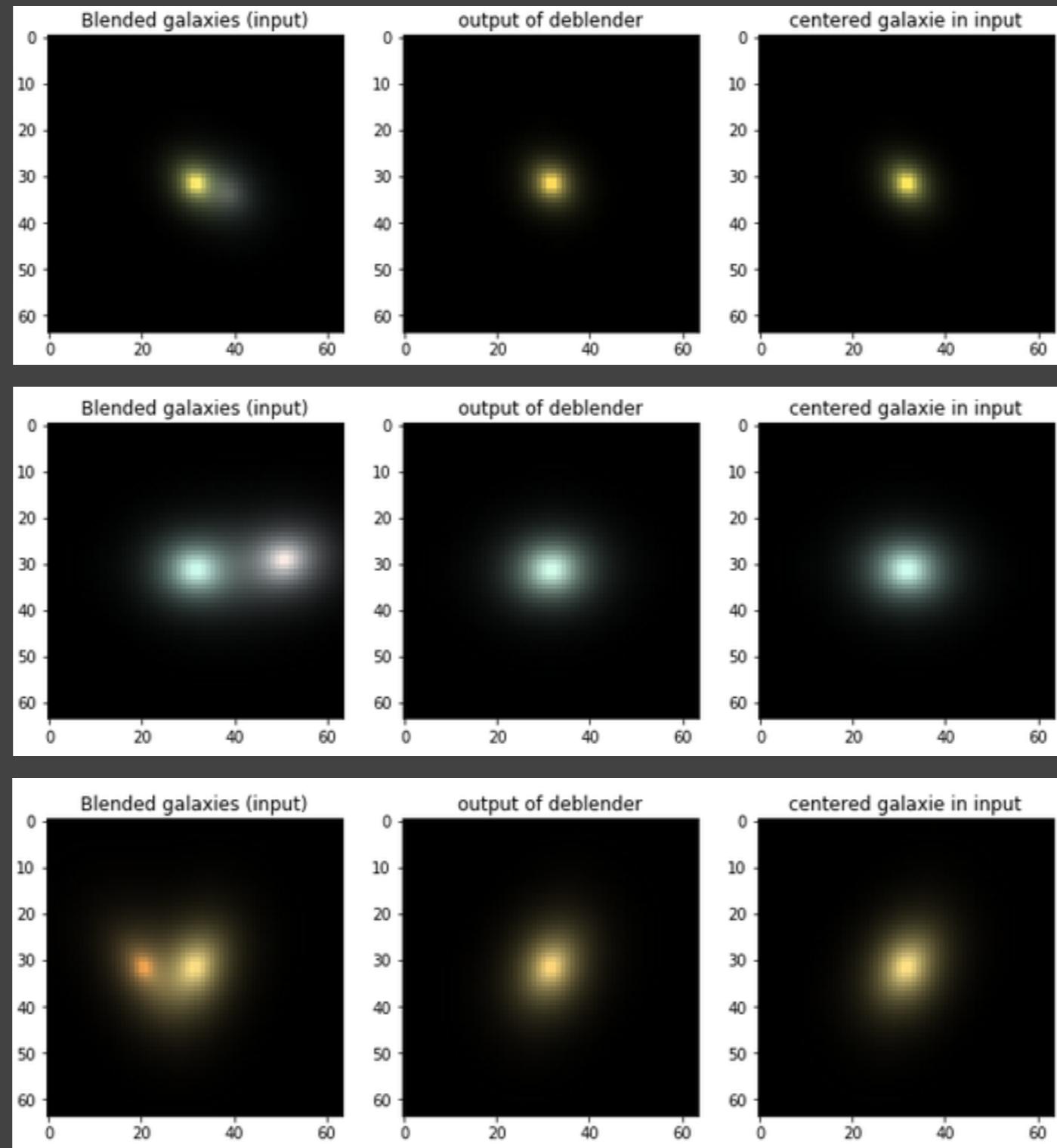
Only R pass-band filter images



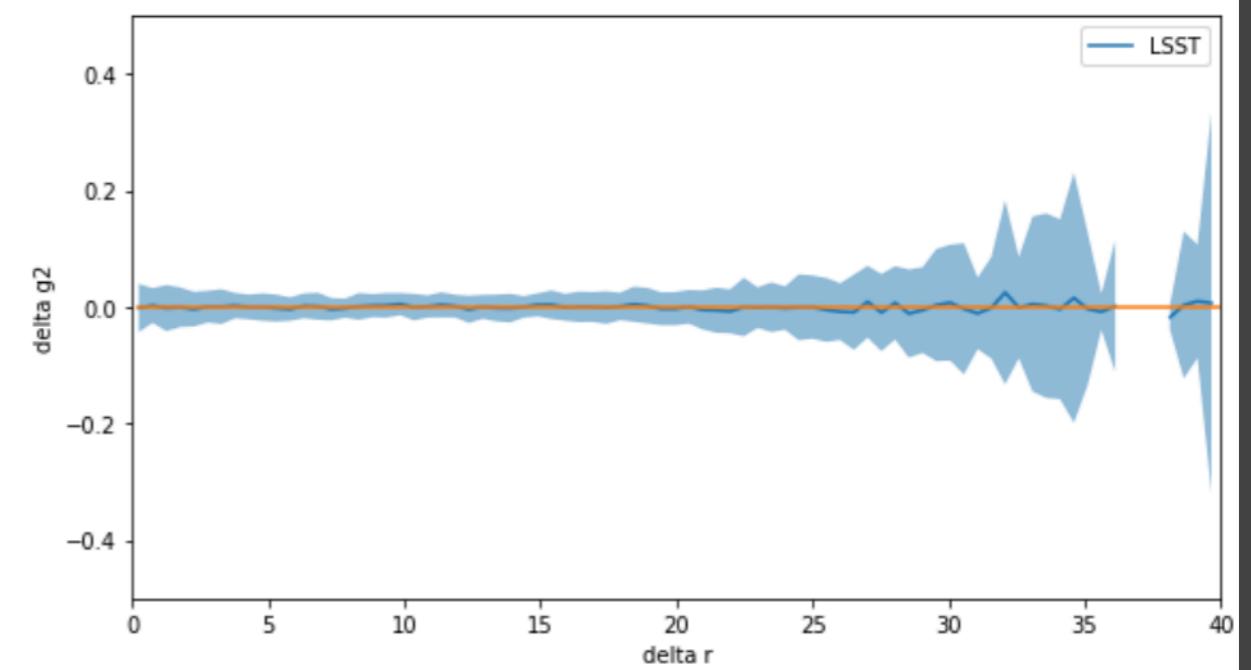
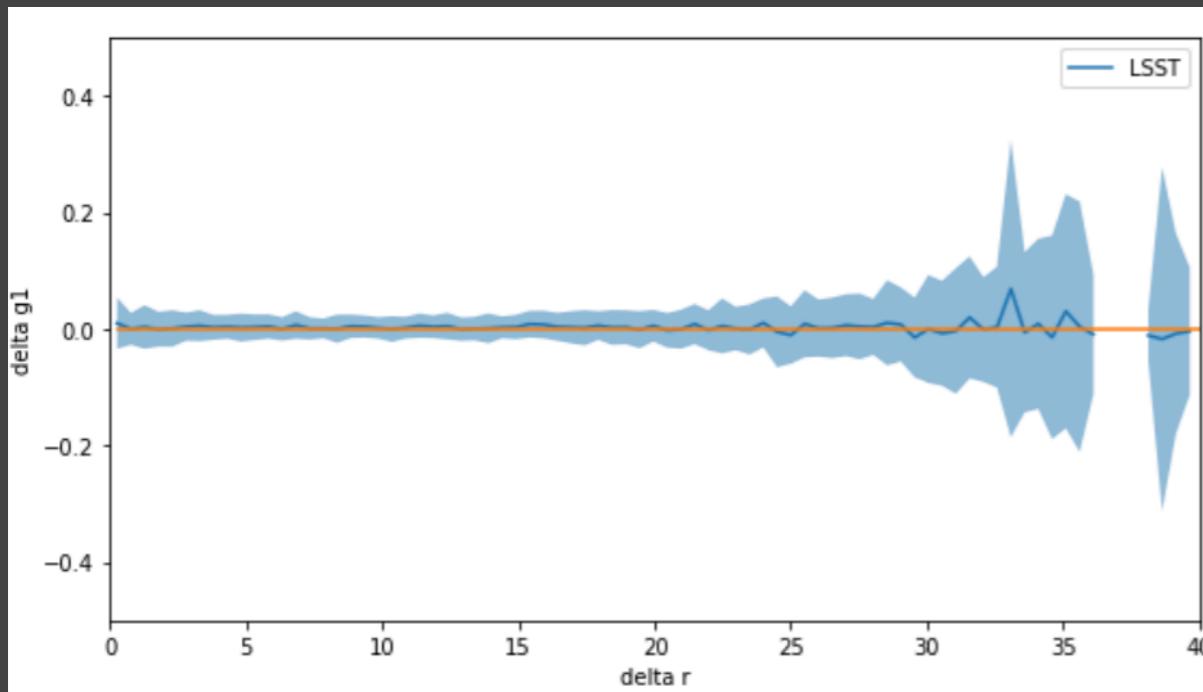
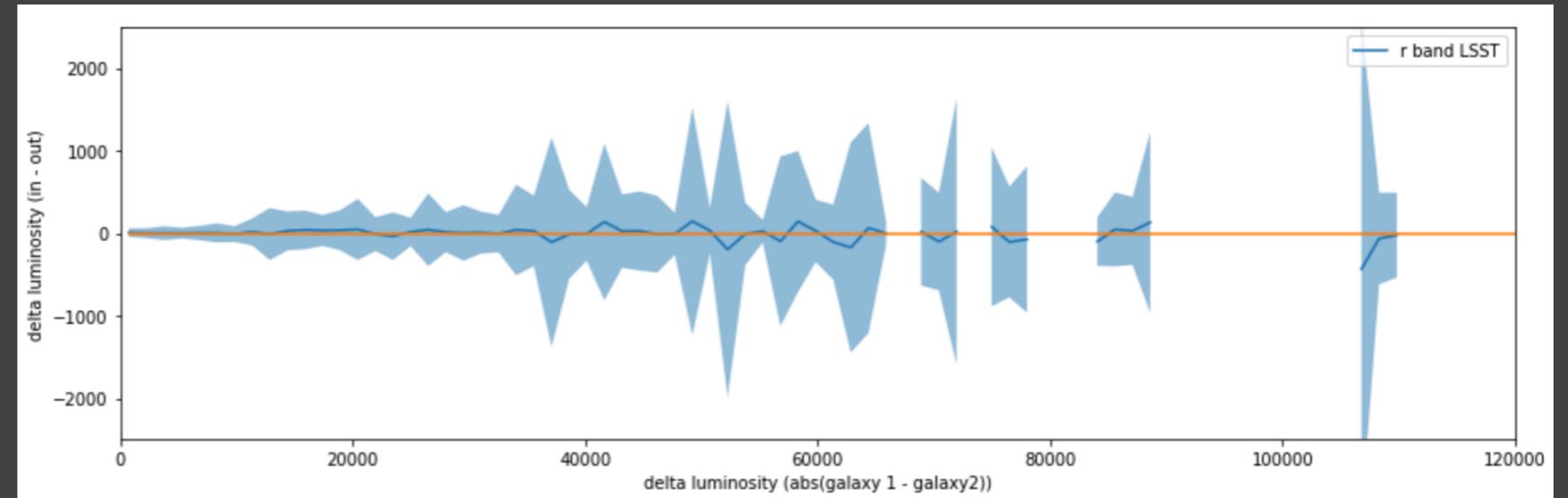
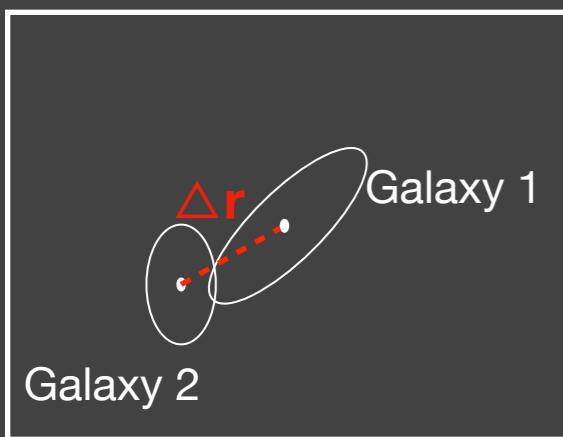
6 LSST pass-band filters images



Deblender : preliminary results

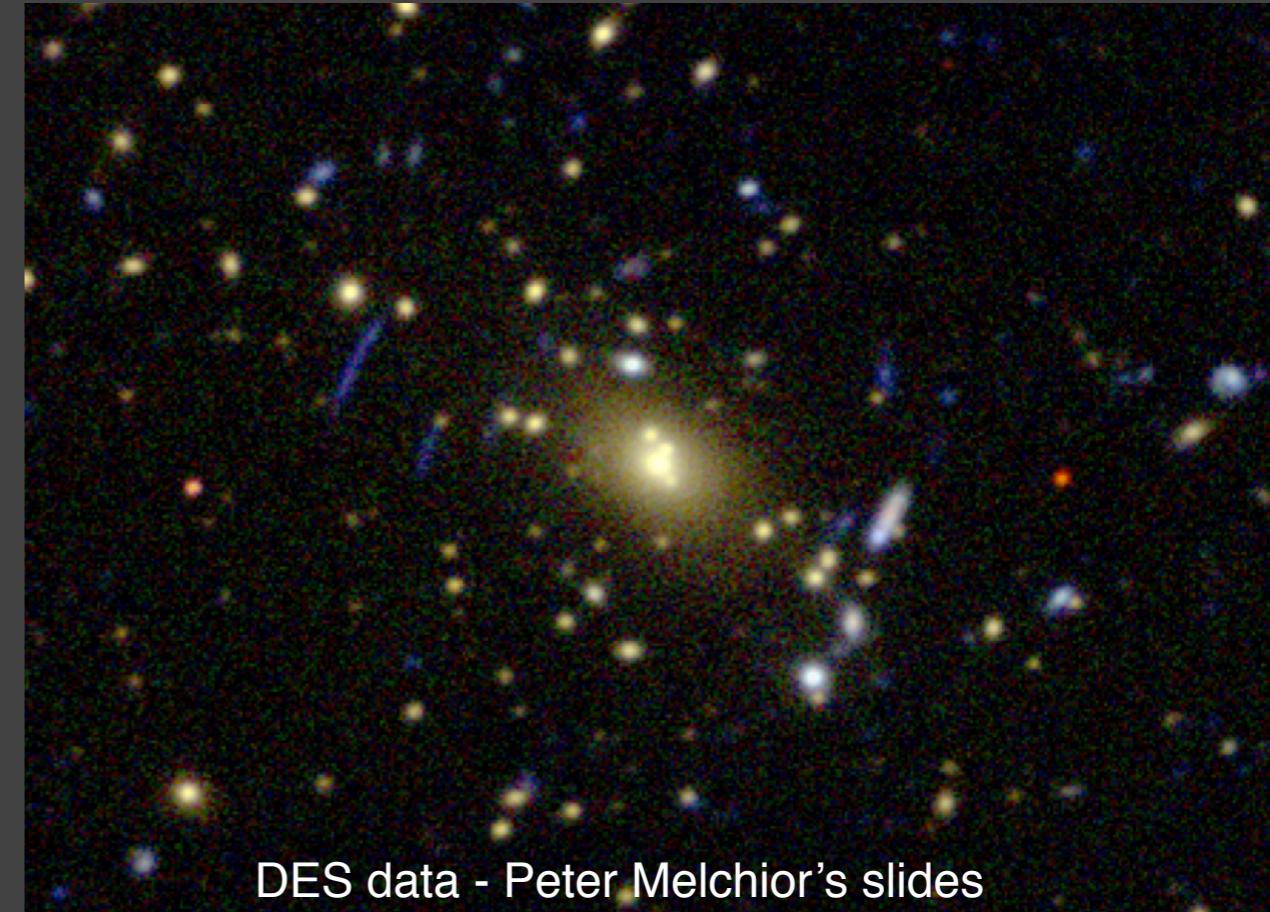


Deblender : preliminary results shape and luminosity reproduction



LSST+Euclid Data

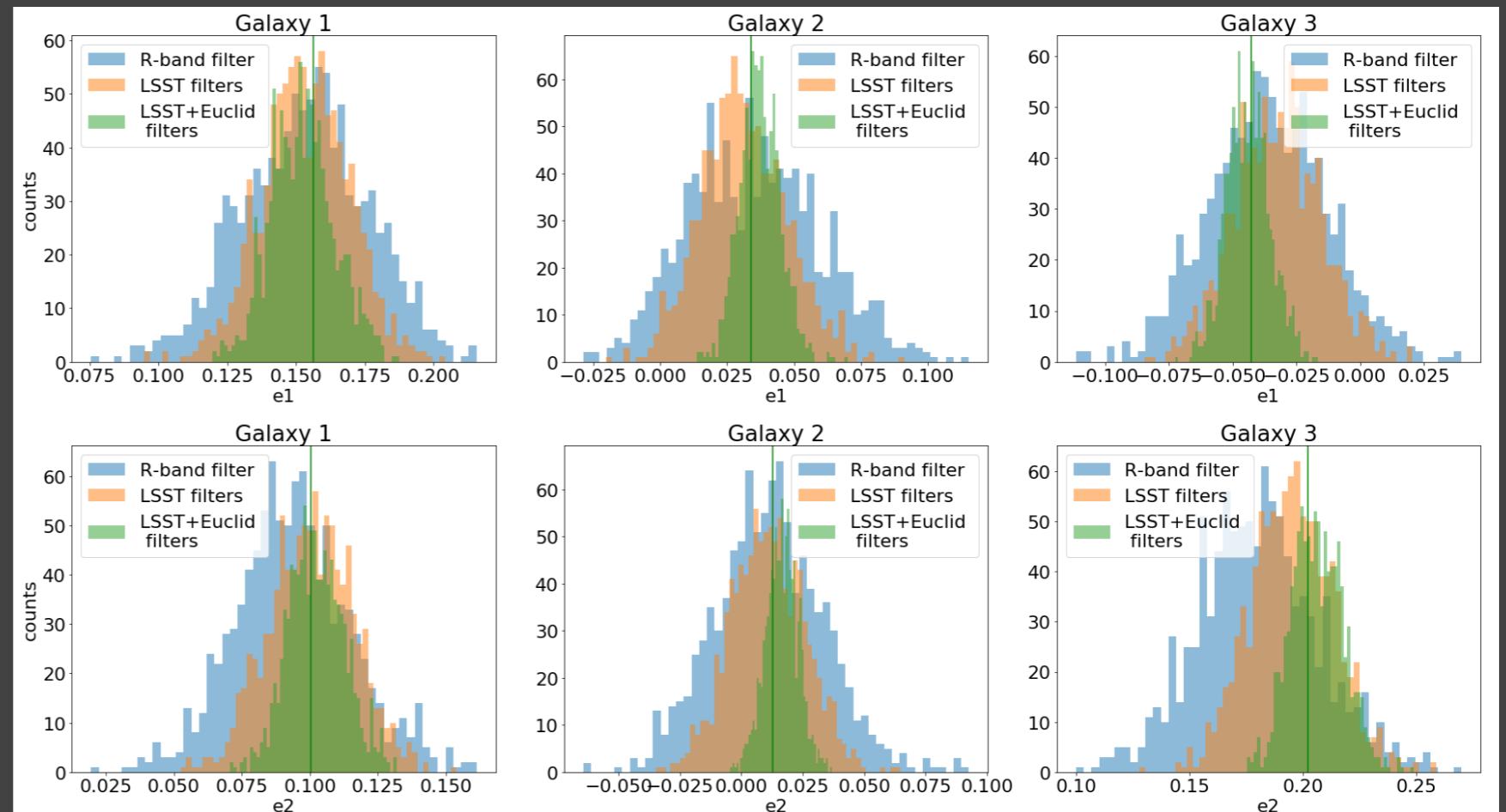
- Why using Euclid data:
 - ✓ Adding infrared bands (x3)
 - ✓ Adding a wide optical band
 - ✓ Different PSF



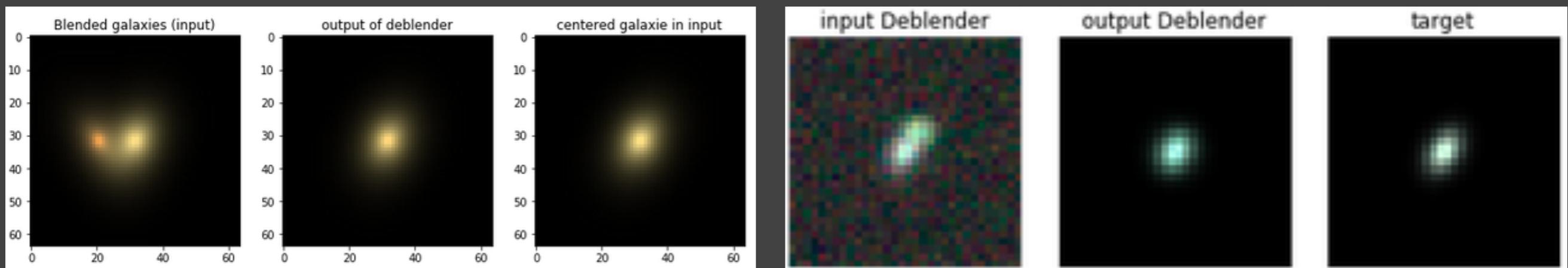
LSST + Euclid : preliminary results

Figure of merit: shape reproduction

- VAE



- Deblender



Where are we ?

- 3 steps:

- 1) Train VAE on noiseless images of single galaxy
- 2) Train on noisy images of single galaxy
- 3) Train on noisy images of blended galaxies



Tests :

- Unbiased reproduction of weak lensing parameters
- architectures, latent space, normalization...

- 3 cases:

- 1) Train on images composed of only ***R band-pass filter***
- 2) Train on images composed of the ***6 LSST band-pass filters***
- 3) Train on images composed of the ***6 LSST band-pass filters + the 4 Euclid band-pass filters***

Conclusion

- VAE :
 - Can reproduce accurately galaxies
 - Provides prior on single galaxies
- **Using several passbands → improvement of the precision**
- **Using several instruments (LSST+Euclid - different PSF) → improvement of the precision**
- Deblender :
 - Reproduction of the shape and of the luminosity.

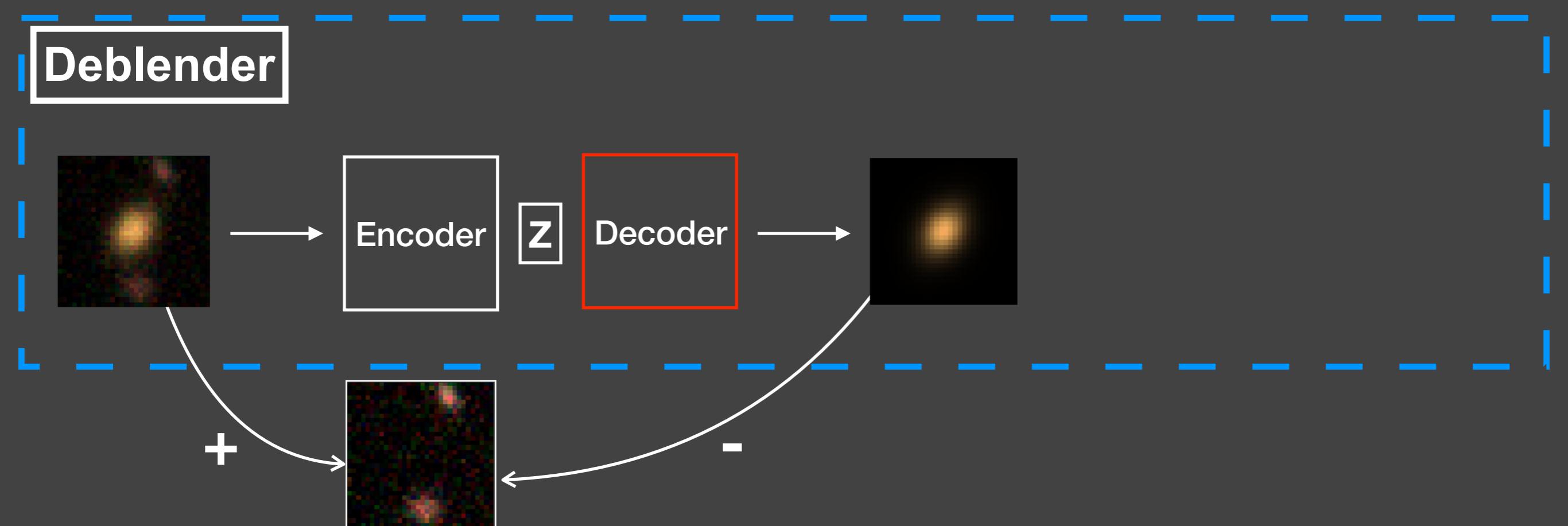
Paper in preparation: Arcelin, Doux et al. (2019). Deblending galaxies with variational Autoencoder: a multi-bands, multi-instruments analysis. (*in prep*)

What's next ?

- **Next:**
 - Train on real data (or DC2 ?)
 - More complex architecture for networks
 - Iterative deblender
 - Work on latent space: have generative model

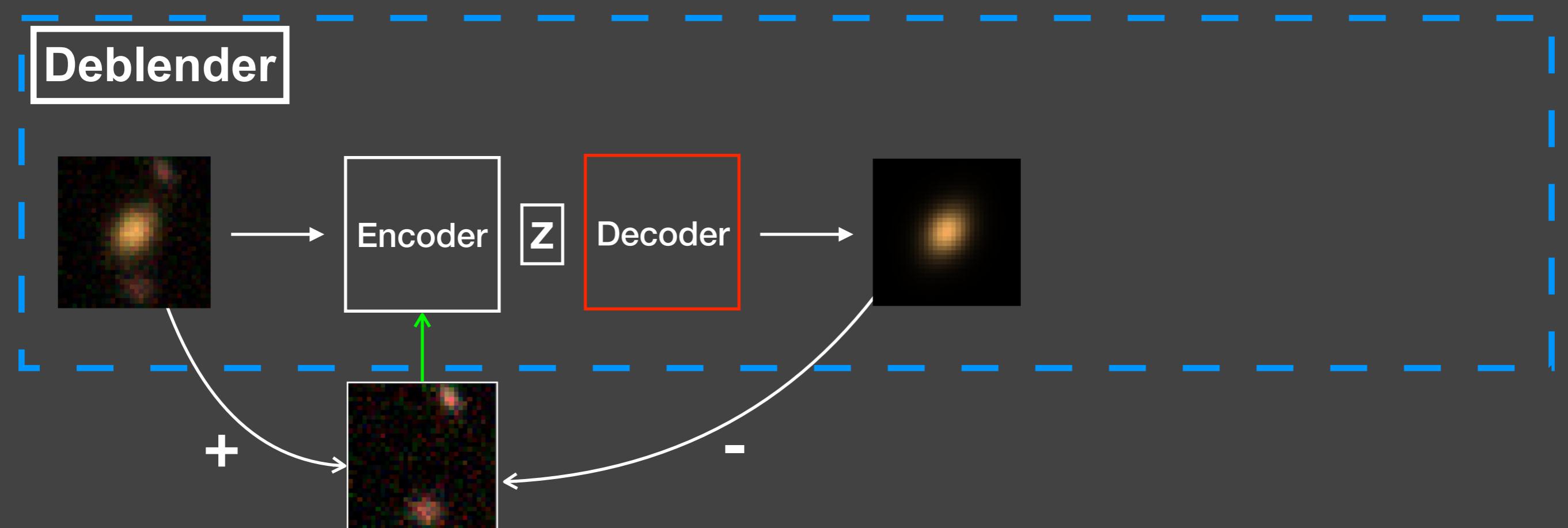
What's next ?

- Iterations by subtracting the output of the deblender to the blended image



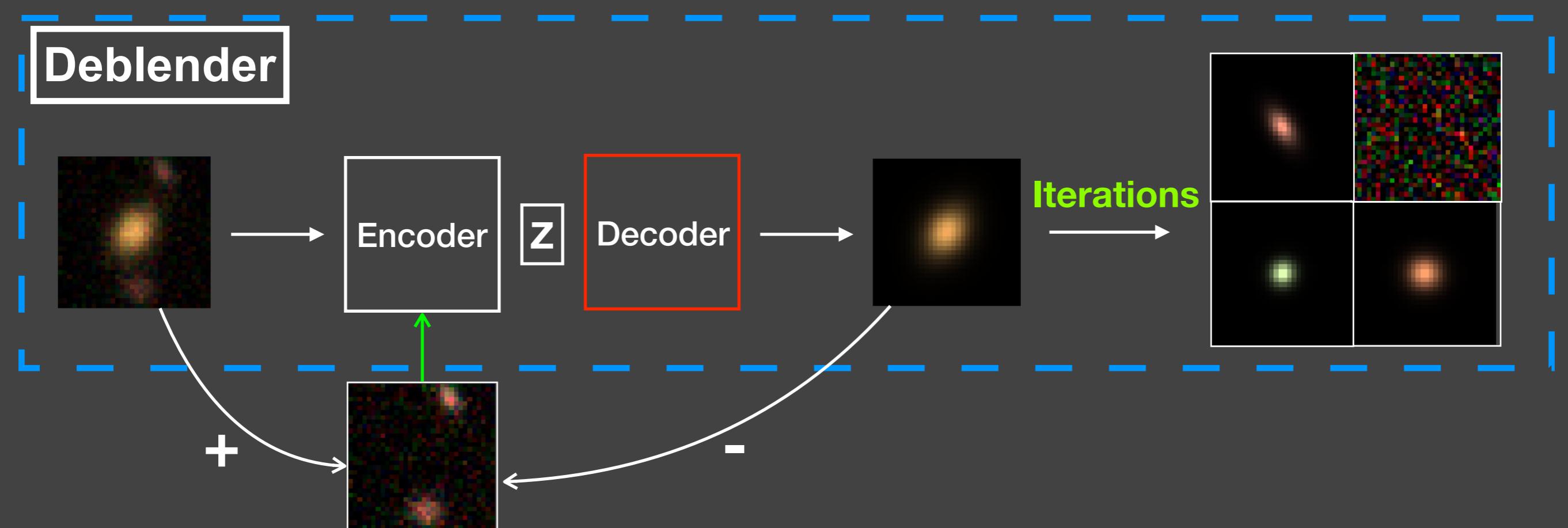
What's next ?

- Iterations by subtracting the output of the deblender to the blended image

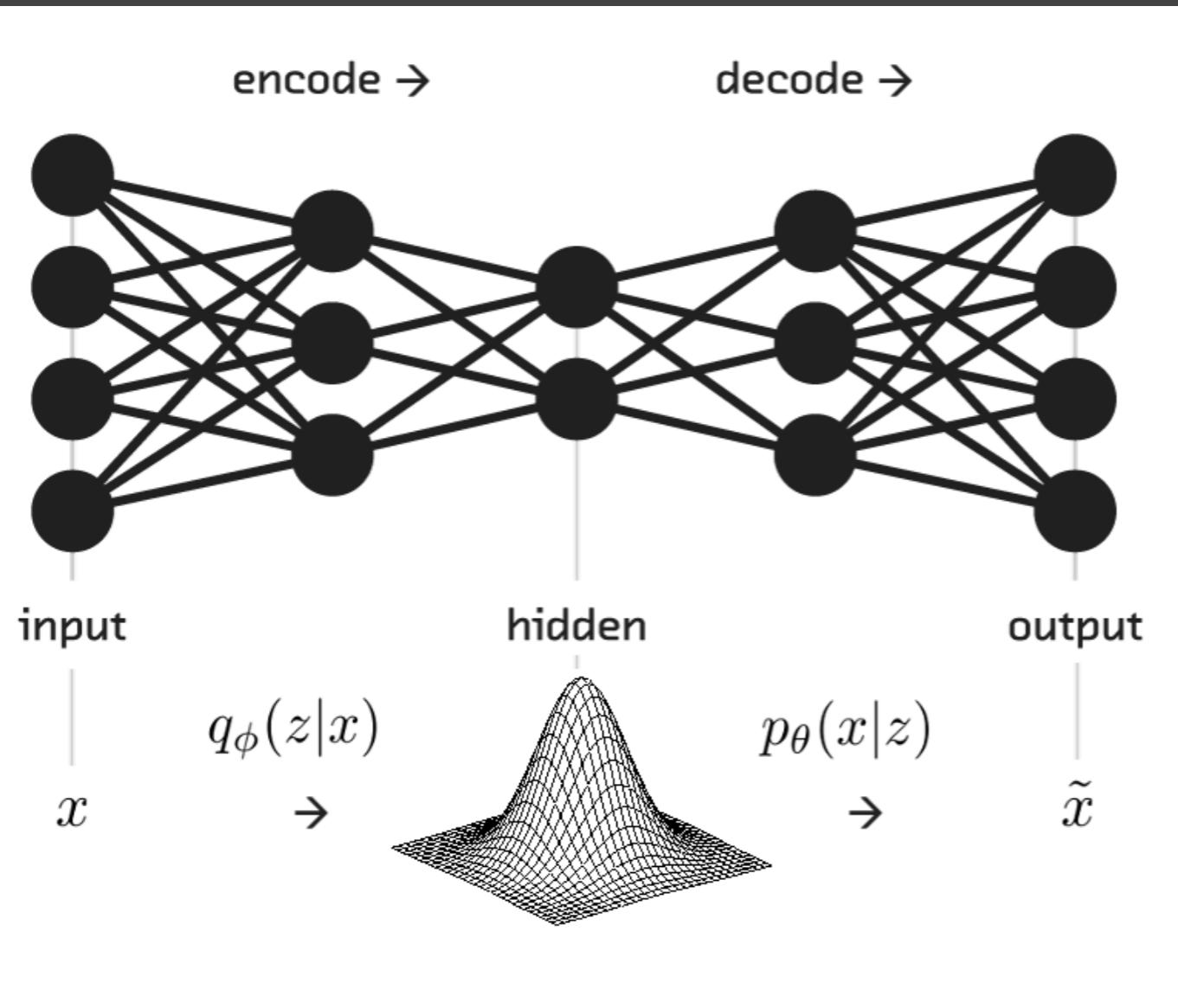


What's next ?

- Iterations by subtracting the output of the deblender to the blended image

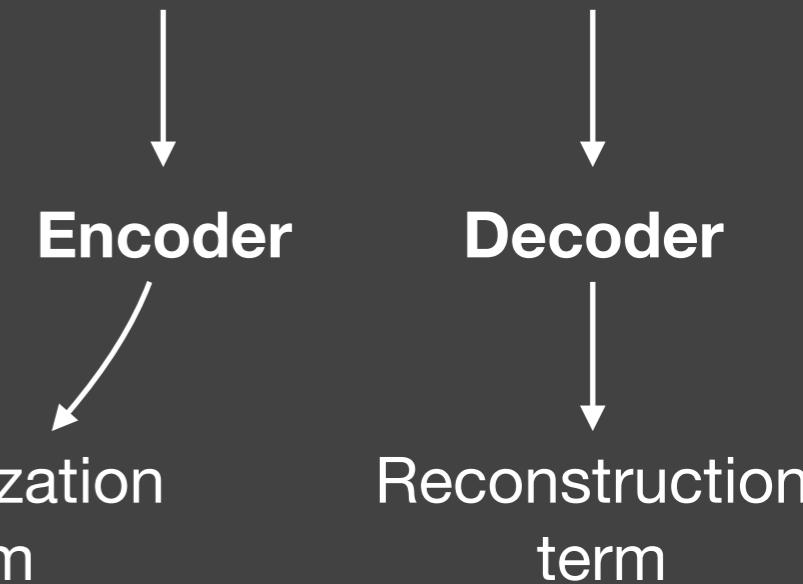


Variational autoencoder (Kingma+2014)



Loss function :

$$ELBO = - D_{KL} \left(q_\phi(z|x) \parallel p(z) \right) + \mathbb{E}(\log p_\theta(x|z))$$



- <http://blog.fastforwardlabs.com>

What's next ?

- Work on latent space: generate galaxies with VAE

