

Target Selection in DESI and Survey Validation

Christophe Yèche (CEA-Saclay)

The 5th joint meeting
eBOSS-DESI France

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Overview of Target Selection

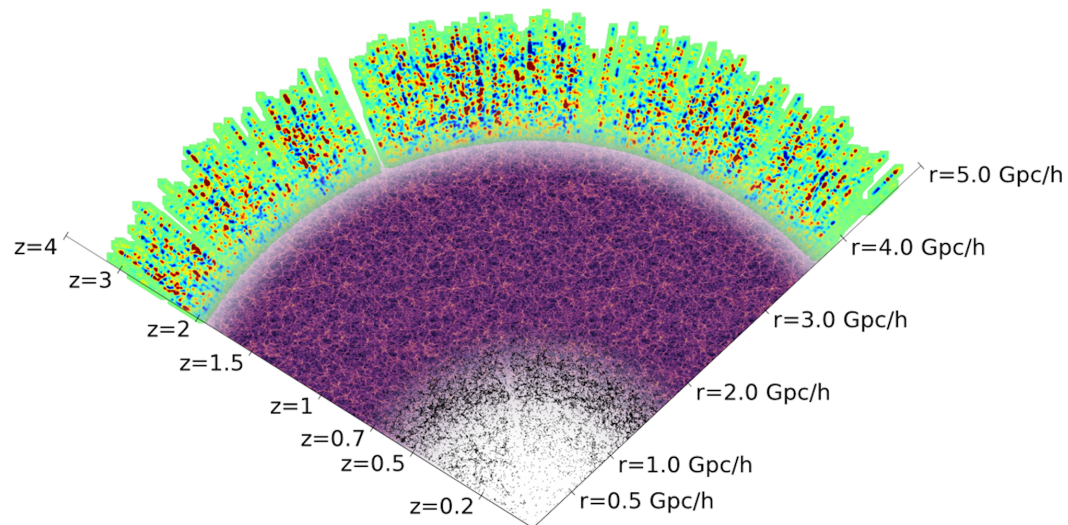


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Target Selection in DESI



Two surveys

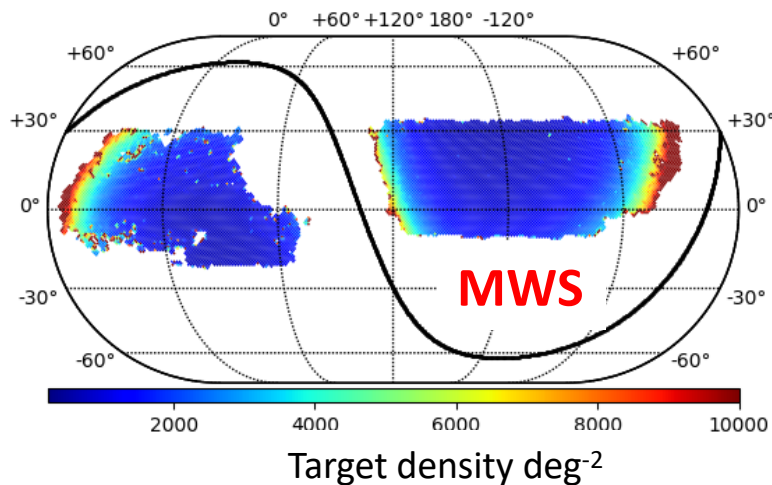
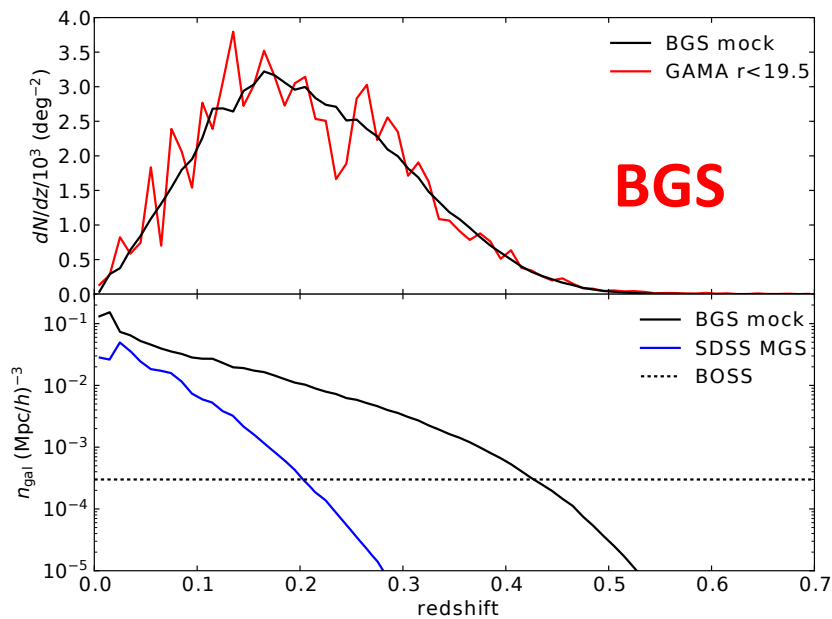
- Dark Time :
 - Dominated by ELGs
 - Bands optimized for ELG
- Bright Time :
 - ~4 nights/lunation
 - MWS/BGS share the observation time with the priority to BGS

Galaxy type	Redshift range	Bands used	Targets per deg ²	Exposures per deg ²	Good z 's per deg ²	Baseline sample
LRG	0.4–1.0	$g,r,z,W1$	480	610	430	6.0 M
ELG	0.6–1.6	g,r,z	2400	1870	1220	17.1 M
QSO (tracers)	< 2.1	$g,r,z,W1,W2$	170	170	120	1.7 M
QSO (Ly- α)	> 2.1	$g,r,z,W1,W2$	90	250	50	0.7 M
Total in dark time			3140	2900	1820	25.5 M
BGS	0.05–0.4	r	800	740	710	9.9 M
BGS–Faint	0.05–0.4	r	600	460	430	6.0 M
MWS	0.0	g,r (Gaia μ)	800+	720	720	10.1 M
Total in bright time			2200+	1920	1860	26.0 M



Selection of BGS galaxies and MWS stars

0.05 < z < 0.4

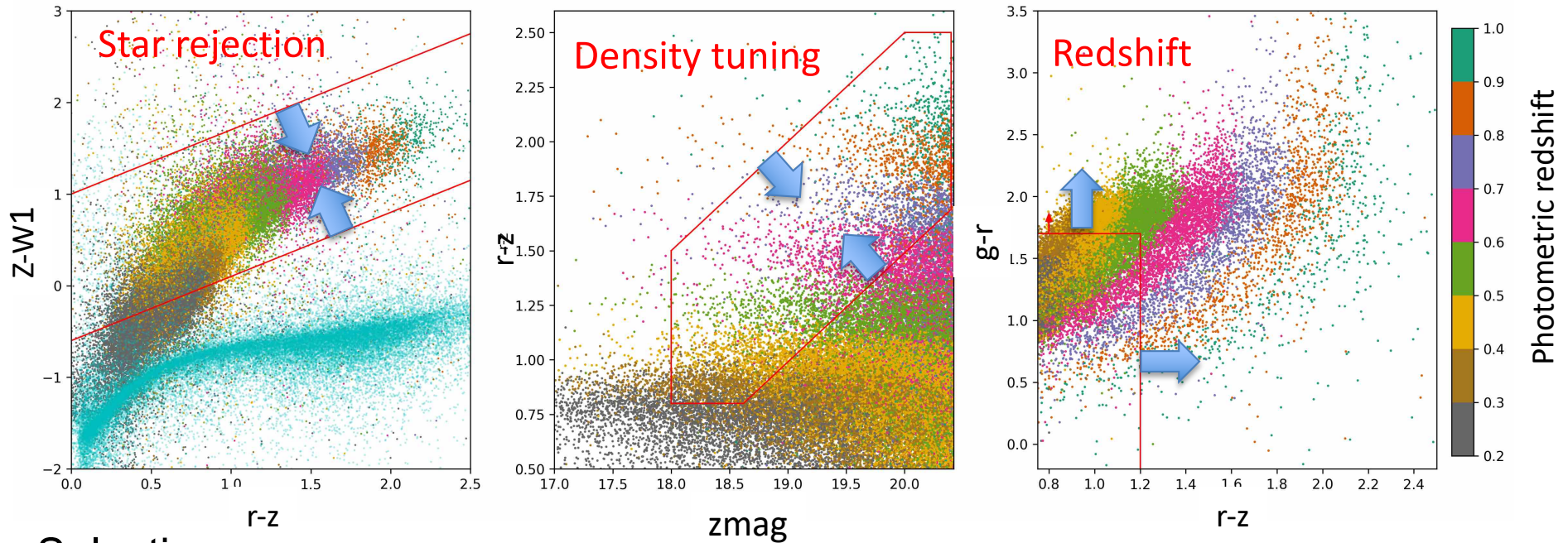


- **BGS Redshift distribution**
 - Median $z \sim 0.2$ ($z \sim 0.1$ for MGS)
 - 10 times SDSS density at $z \sim 0.3$
- **BGS Selection**
 - Morphology by comparing Gaia G and r band: $G-r > 0$ or $G=0$
 - $r < 19.5$ (priority 1)
 - $19.5 < r < 20$ (priority 2)
- **BGS Target density**
 - BGS-main: 800 deg^{-2}
 - BGS-faint: 600 deg^{-2}
- **Milky Way Stars**
 - Extra fibers used for MWS
 - Density: $\sim 700 \text{ deg}^{-2}$
 - Selection based on GAIA
 - 8 M of stars in $16 < r < 19$ range, down to $r \sim 20$ for fainter stars



Selection of LRGs

0.4 < z < 1.0



- Selection

- z-W1 Vs r-z to veto stars
- No morphology criterion
- $r-z > 1.2$ or $g-r > 1.7$ for redshift range
- Sliding color magnitude cut to tune the density: $(z-17.18)/2 < r-z < (z-15.11)/2$
- Slightly different selections for DECaLS and MzLS+BASS

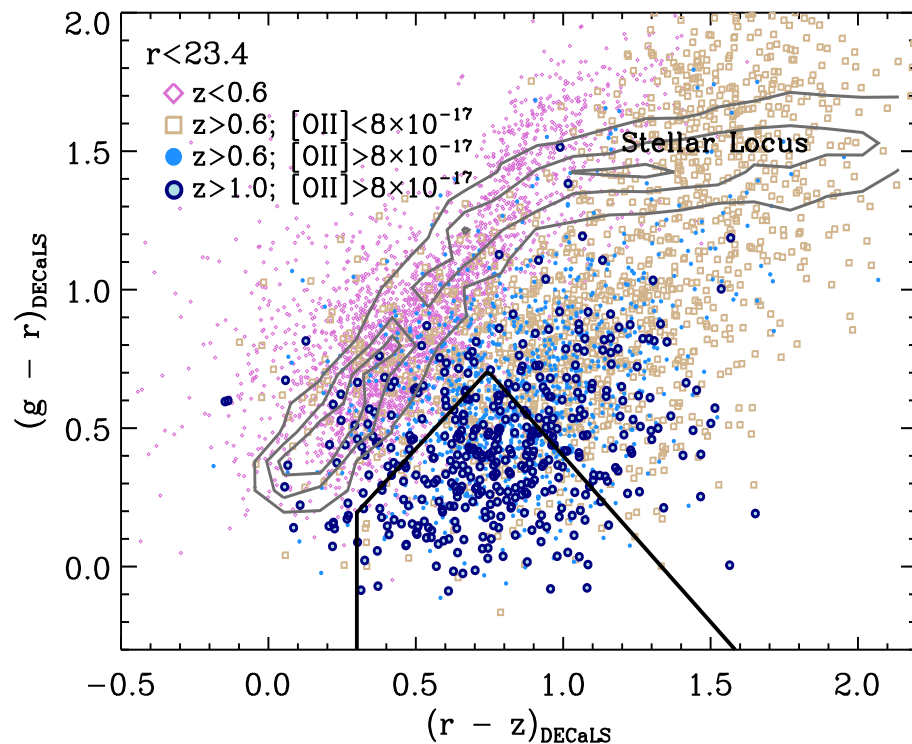
- Density

- One pass ($z < 20$) : 350 deg^{-2}
- Two passes ($20 < z < 20.4$): 130 deg^{-2}
- Good redshifts: 90% $\sim 430 \text{ deg}^{-2}$



Selection of ELGs

$0.6 < z < 1.6$



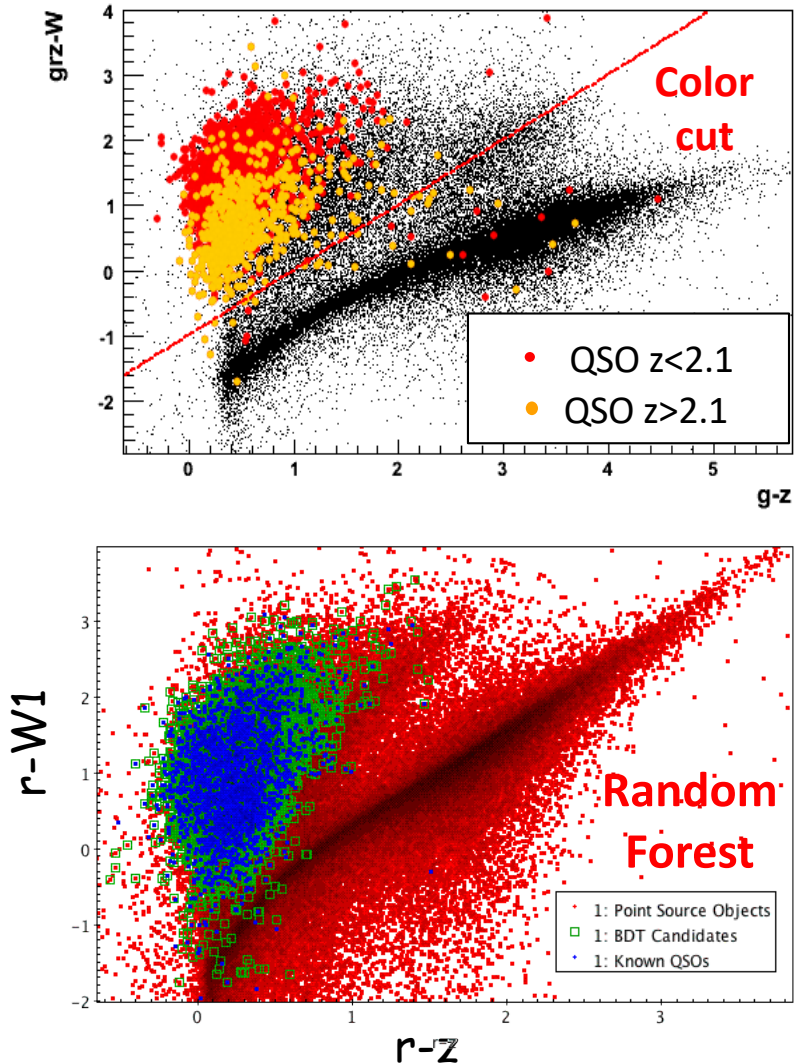
From DEEP2, validated with COSMOS

- Redshift distribution
 - Median $z \sim 1.0$
 - $nP(k) > 1$ up to $z \sim 1.3$
- Selection
 - g, r, z optical bands
 - $g < \sim 23.5$ to reach density goal
 - No morphology criterion
- Different selection in North/South
 - $g < 23.6 \rightarrow g < 23.4$ and color box
- Target density goal
 - 2400 deg^{-2}
- 2nd generation selection
 - Ongoing work to develop 2nd generation selections (Number Density Modeling and Random Forest)
 - Pilot surveys with MMT



Selection of QSOs

$z < 2.1$ (mean $z \sim 1.5$) \rightarrow QSO clustering
 $z > 2.1$ (mean $z \sim 2.5$) \rightarrow Ly α clustering



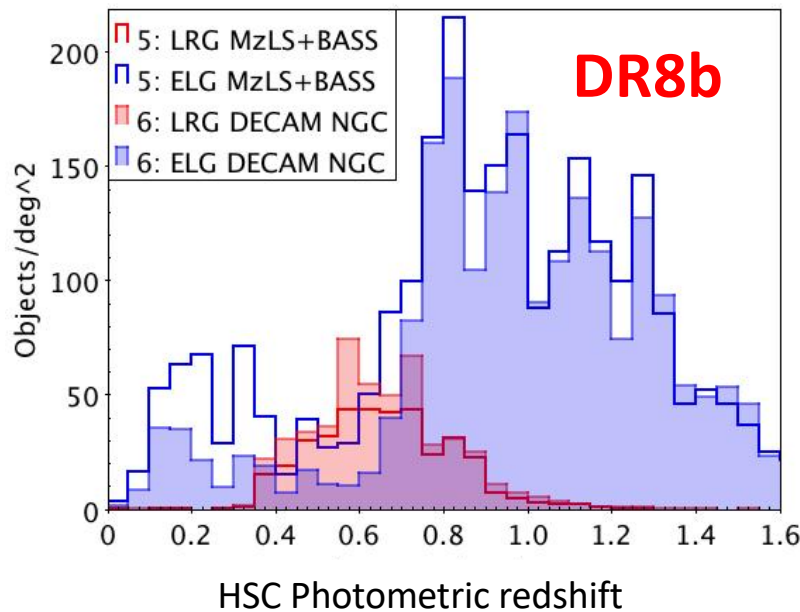
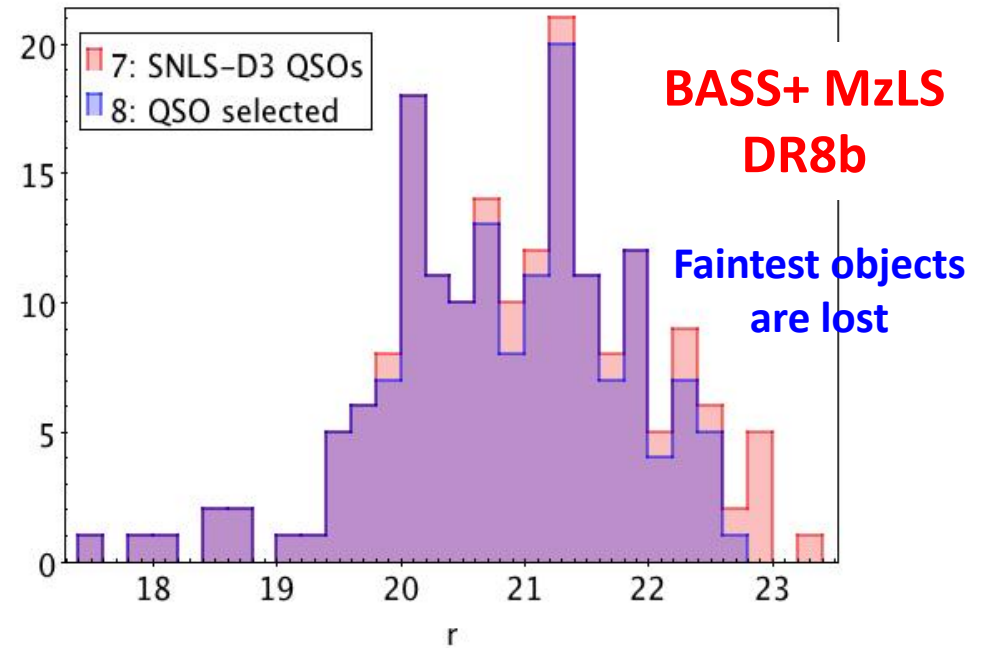
- Common main selection
 - Point source morphology
 - grz optical bands and W_1W_2 NIR bands
 - $r < 22.7$
- Two approaches
 - 1st generation: Color cut selection
 - 2nd generation: Machine learning approach with Random Forest
- Target density goal
 - 260 deg^{-2}
 - Up to 4 passes for SNR of Ly- α QSO
- Success rate
 - $z < 2.1$: 120 deg^{-2}
 - $z > 2.1$: 50 deg^{-2}



Validations with DECaLS and BASS+MzLS

- Validation of QSOs

- Use SLNS D3 field (1 deg²) with QSOs selected by their variability (210 deg⁻²)
- Achieve FDR goals : ~170 QSOs deg⁻² with (>50) with z>2.1 but with a density a very low statistic...



- Validation of LRGs and ELGs

- Photometric redshifts from Hyper Suprime-Cam Subaru (HSC)
- Encouraging results with color cut selection for DECaLS
- More low-z galaxies for BASS+MzLS



Preparation of Survey Validation



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Goals of Survey Validation (SV)

- **Quality of the spectra**
 - Exposure times: achieve expected S/N
 - Calibration: test that calibration satisfies spectrophotometry requirements
 - Pipeline performances: Achieve redshift accuracy
- **Validation of Target Selection (discussed in next slides)**
 - Final target efficiency → Target densities
 - Final $n(z)$ for each tracer
 - Finalization of the TS
- **Observation constraints**
 - 3-4 months at 50% efficiency
 - 100 / 60 fields per month for (LRG,ELG,QSO) / BGS
 - 4x times the nominal exposure time



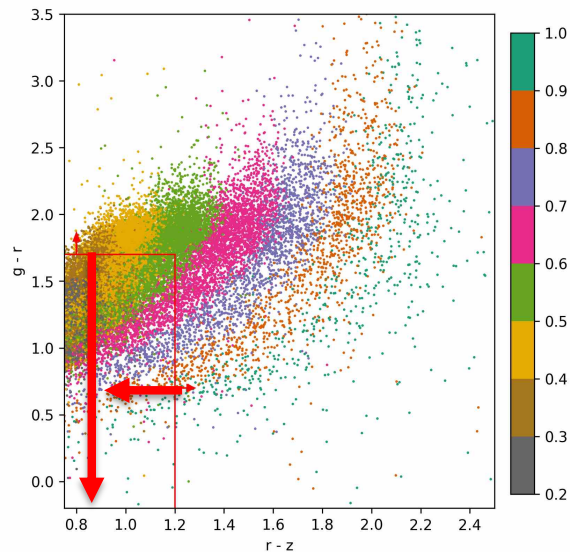
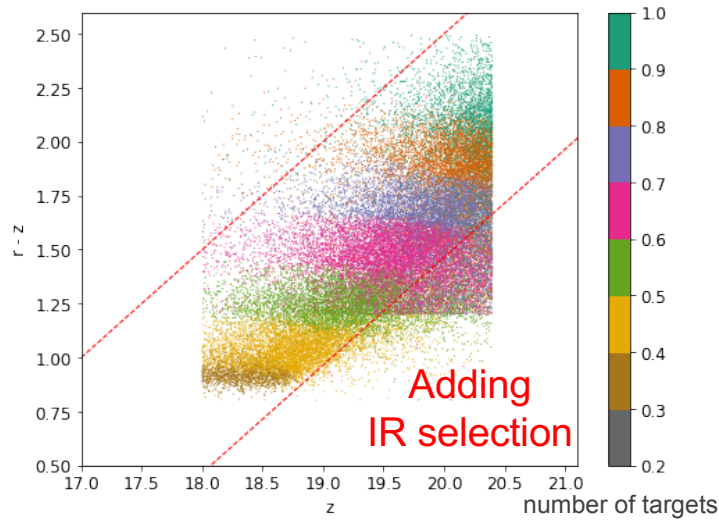
Preparation of Target Selection for SV

- **Definition of the supersets**
 - Relax the cuts
 - Alternative selections (LRGs: Optical/IR or QSO Color cuts/Random Forest)
 - Tune the redshift distributions
 - Optimize the redshift overlap: BGS/LRG, LRG/ELG
 - ~Double fiber budget compared to nominal selection
- **Subsampling**
 - If the density of superset too high → Subsampling according categories
 - Mandatory for LRGs, possible for ELGs, not required for QSOs
 - Subsampling managed by MTL (Merged Target List)



SV superset for LRGs

$0.4 < z < 1.0$

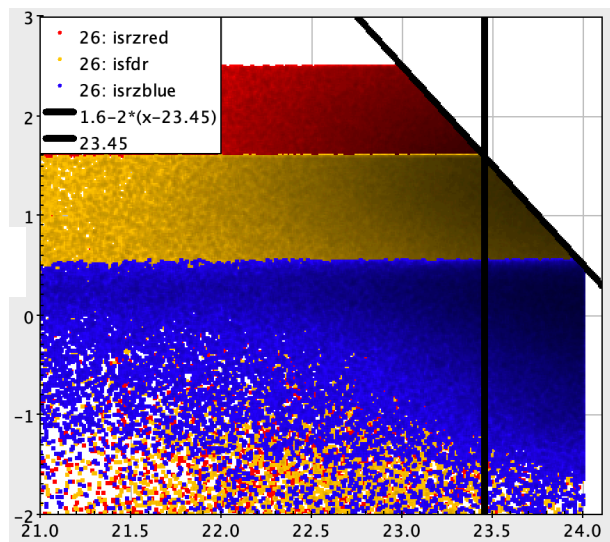
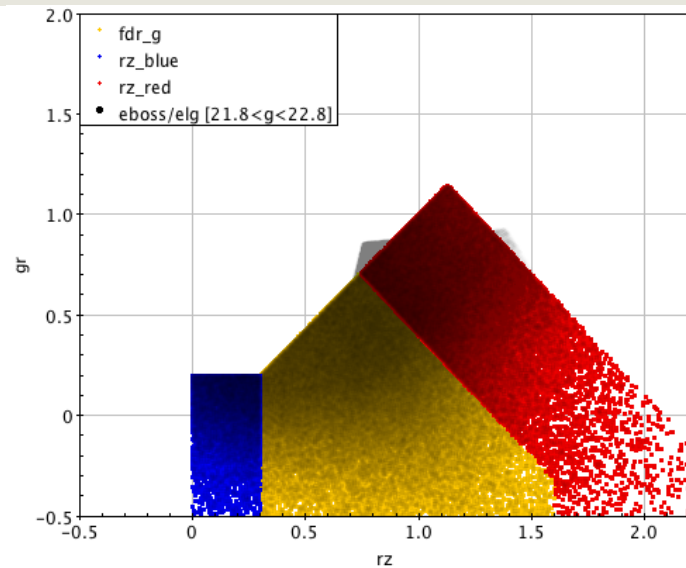


- Optical/IR selection
 - Sliding mag. cut :
(W1 - 17.74)/0.4 < r-W1
- + Low-z extension
 - Bright LRGs (or $z < 19.4$)
- + High-z extension
 - Faint LRGs (or $z < 20.8$)
- + Relaxed cuts
 - Stellar rejection + Sliding cuts
 - + $g-r < 1.45$
- + more relaxed cuts
 - No $g-r$ cut and $r-z > 0.65$
- Final budget
 - Five bits are defined
 - Optical+IR: 600 deg^{-2}
 - All bits: 2300 deg^{-2}



SV supersets for ELGs

$0.6 < z < 1.6$

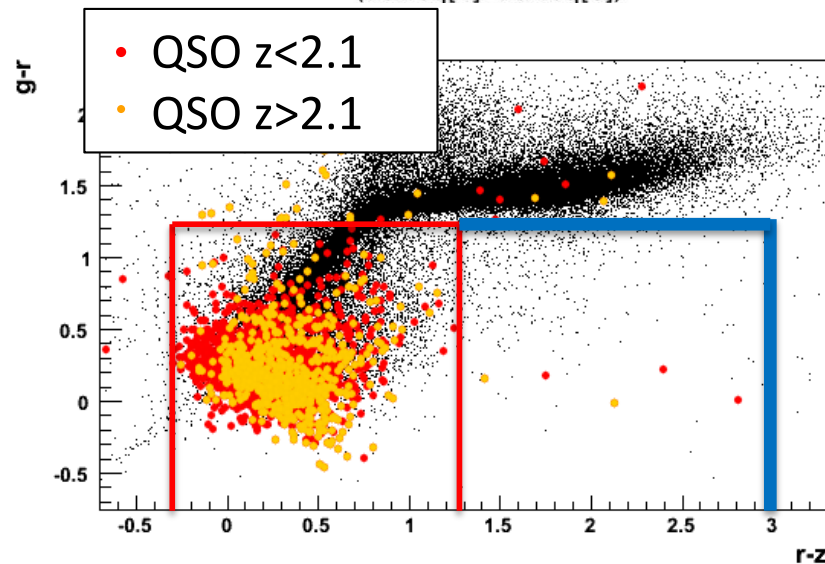
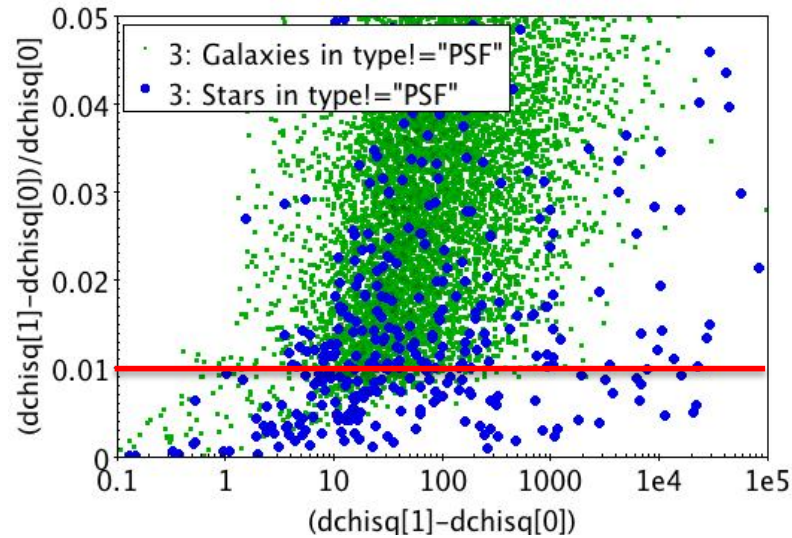


- Faint objects
 - Magnitude limit: $23.45 < g < 23.65$
- Red region
 - ELGs with similar redshift
 - But lower OII flux...
- Blue region
 - ELGs with very high redshift and reasonable OII flux
 - But close to stellar locus and low- z galaxy region
- Three bits
 - Combination of faint / red / blue bits
 - Density: With sliding color cut on mag. to reject faint objects in red region $\rightarrow \sim 4500 \text{ deg}^{-2}$



SV supersets for QSOs

$z < 2.1$ (mean $z \sim 1.5$) \rightarrow QSO clustering
 $z > 2.1$ (mean $z \sim 2.5$) \rightarrow Ly α clustering



- Relax PSF definition
 - Recover objects with $(dchisq[1]-dchisq[0])/dchisq[0] < 0.01$
- Select fainter objects
 - Magnitude limit $r < 22.7 \rightarrow r < 23.0$
- Relax color cut selection
 - Enlarge grz and WISE color boxes
 - Reduce star stripe
- Relax random forest
 - Tune RF probability
- Add High-z faint QSOs
 - $22.7 < r < 23.5$ (optimization of f_{NL})
- Three bits
 - Color Cut / RF / high-z faint QSO
 - Density: $400/400/70 \rightarrow 600 \text{ deg}^{-2}$
 - Completeness: $\sim 95\%$



Observation in SV

- LRG and QSO observations
 - 10 fields (with 8x exposure time): faint LRGs and high-z QSOs
 - 10 fields (with 4x exposure time): bright LRGs and low-z QSOs
 - 10 fields (with 4x exposure time): all QSOs
- ELG observations
 - 25 isolated fields with 4x exposure time
 - Density: $\sim 4000 \text{ deg}^{-2}$
- Schedule
 - From Feb./March 2020 to June 2020
 - 55 specific fields with increased exposure time (2-3 months)
 - Final TS selection \rightarrow 1% survey (1 month)

SV and science is starting at DESI in less than one year!



Extra Slides



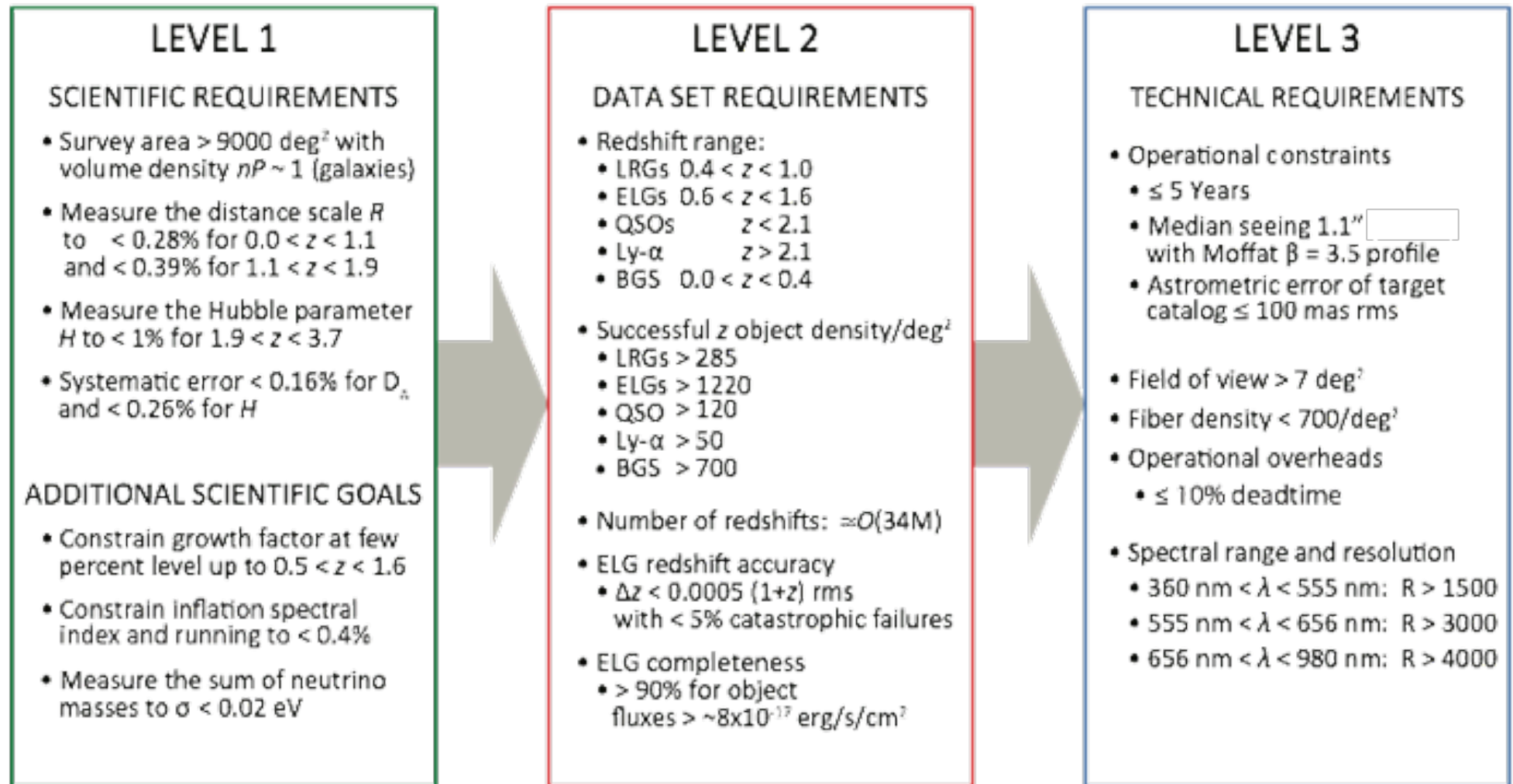
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Science Requirements

SCIENTIFIC OBJECTIVE: PERFORM A STAGE-IV BAO SPECTROSCOPIC SURVEY



DESI-0844v4



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