

Quantum Gravity at the LHC

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motivation

- **standard model of particle physics**

theory of the strong, weak, and electro-magnetic interactions

very successful, up to $\sim \mathcal{O}(100)$ **GeV**

LHC to explore physics of the electroweak scale, Higgs particle

motivation

- standard model of particle physics
- what about gravity?

presently, no unified understanding of all fundamental forces

Planck mass $M_{\text{Pl}} \approx 10^{19} \text{GeV} \gg M_{\text{EW}}$

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- hierarchy problem

coupling of Higgs particle to gravity

one-loop level:

either no symmetry breaking

or Higgs particle mass of order M_{Pl}

or non-trivial cancelations at loop level (fine-tuning)

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- low-scale quantum gravity

what if the **fundamental** Planck scale M_D obeys

$$M_D \approx \mathcal{O}(M_{EW}) \approx \mathcal{O}(1\text{TeV}) \ll M_{Pl}$$

quantum gravity accessible at colliders

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quantum gravity accessible at colliders

- scenarios

large extra dimensions

(Arkani-Hamed, Dimopoulos, Dvali '98)

RS I and RS II

4d setup with many species, strong RG effects

(G Dvali, et. al.)

low scale quantum gravity

- **large extra dimensions**

(Arkani-Hamed, Dimopoulos, Dvali '98)

D=4+n compact spatial dimensions

fundamental Planck mass $M_D \approx \mathcal{O}(M_{EW}) \ll M_{\text{Planck}}$

low scale quantum gravity

- **large extra dimensions**

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D=4+n compact spatial dimensions

fundamental Planck mass $M_D \approx \mathcal{O}(M_{EW}) \ll M_{\text{Planck}}$

compact extra dimensions $M_{\text{Planck}}^2 \sim M_D^2 (M_D L)^n$

roughly $L \sim 10^{\frac{30}{n}-17} \text{cm} \left(\frac{1 \text{TeV}}{m_{EW}} \right)^{1+\frac{2}{n}}$

scale separation $1/L \ll M_D \ll M_{\text{Planck}}$

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- **quantum gravity scenarios**

string theory

effective theory for gravity

(Donoghue '94)

asymptotic safety for gravity

(Weinberg '79)

UV fixed points in $D = 4$ dimensions

(Reuter '96)

UV fixed points in $D > 4$ dimensions

(DL '03, Fischer, DL '06)

renormalisation group

- **RG scaling of gravitational coupling**

gravitational coupling $G(\mu) = Z_N(\mu)^{-1} \cdot G_N$

dimensionless coupling $g(\mu) = Z_N(\mu)^{-1} \cdot G_N \cdot \mu^{D-2}$

anomalous dimension $\eta_N = -\frac{d \ln Z_N}{d \ln \mu}$

RG running $\frac{dg(\mu)}{d \ln \mu} = (D - 2 + \eta_N) g(\mu)$

(DL '06, Niedermaier '06)

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- **fixed points**

Gaussian: $g = 0$ **classical general relativity**

non-Gaussian: $\eta_N = 2 - D$ **strong quantum effects**

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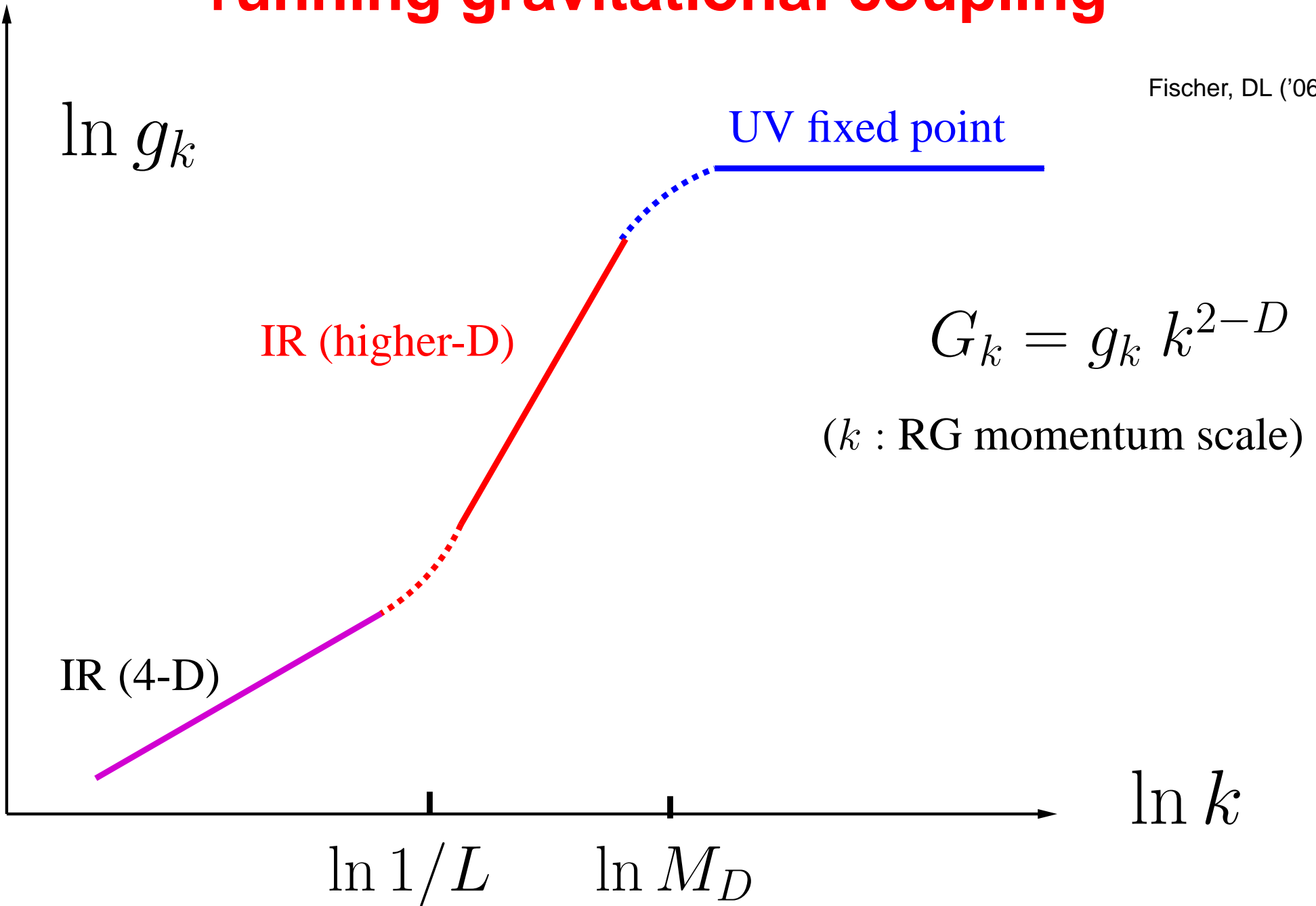
non-Gaussian: $\eta_N = 2 - D$ **strong quantum effects**

UV fixed point implies weakly coupled gravity at **high energies**

$$\mu \rightarrow \infty : \quad G(\mu) \rightarrow g_* \mu^{2-D} \ll G_N$$

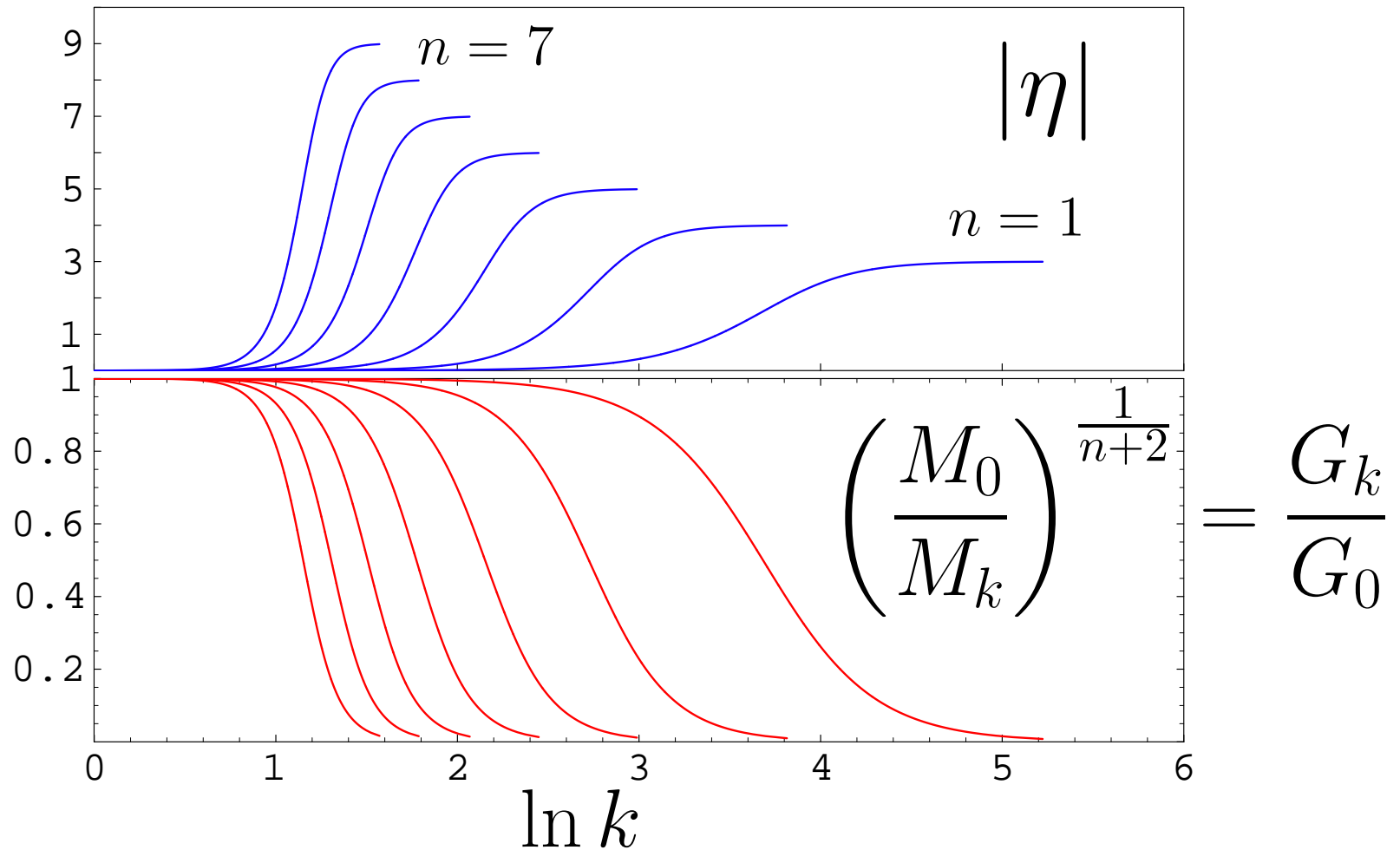
running gravitational coupling

Fischer, DL ('06)



RG running and anomalous dimension

DL ('03), Fischer, DL ('05,'06)



collider signatures of quantum gravity

- **real gravitons**

graviton production via $p p \rightarrow \text{jet} + G$

signature: missing energy

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lepton production $q\bar{q} \rightarrow \ell^+ \ell^-$ via graviton exchange

signature: deviations in SM reference processes

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- **mini-black holes**

black hole production and decay

signature: spectacular (many body final states)

gravitational Drell-Yan

- **effective theory** Giudice, Rattazzi, Wells ('98)

scattering amplitude for Drell-Yan lepton production

$$A = \mathcal{S}(s) \times T, \quad T = T^{\mu\nu} T_{\mu\nu} - \frac{1}{n+2} T_{\mu}^{\mu} T_{\nu}^{\nu}$$

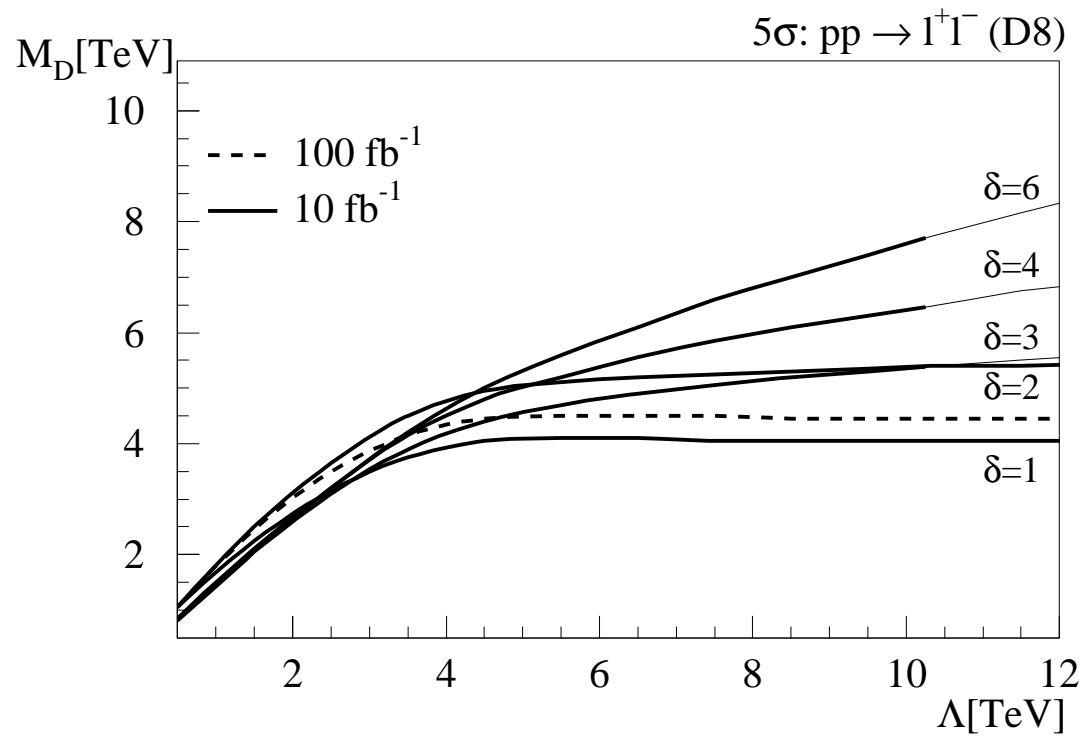
$$\mathcal{S}(s) = \frac{1}{M_D^{n+2}} \int_0^{\infty} dm_{\text{kk}} m_{\text{kk}}^{n-1} \frac{1}{s + m_{\text{kk}}^2}$$

UV divergent for $n \geq 2$.

gravitational Drell-Yan

- effective theory + Monte Carlo simulations

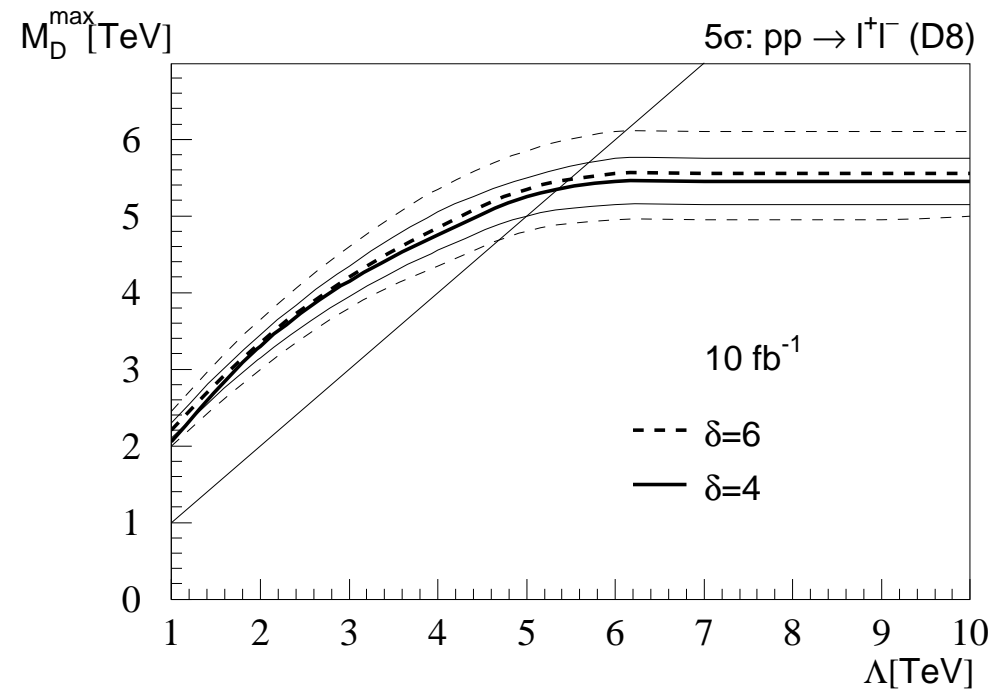
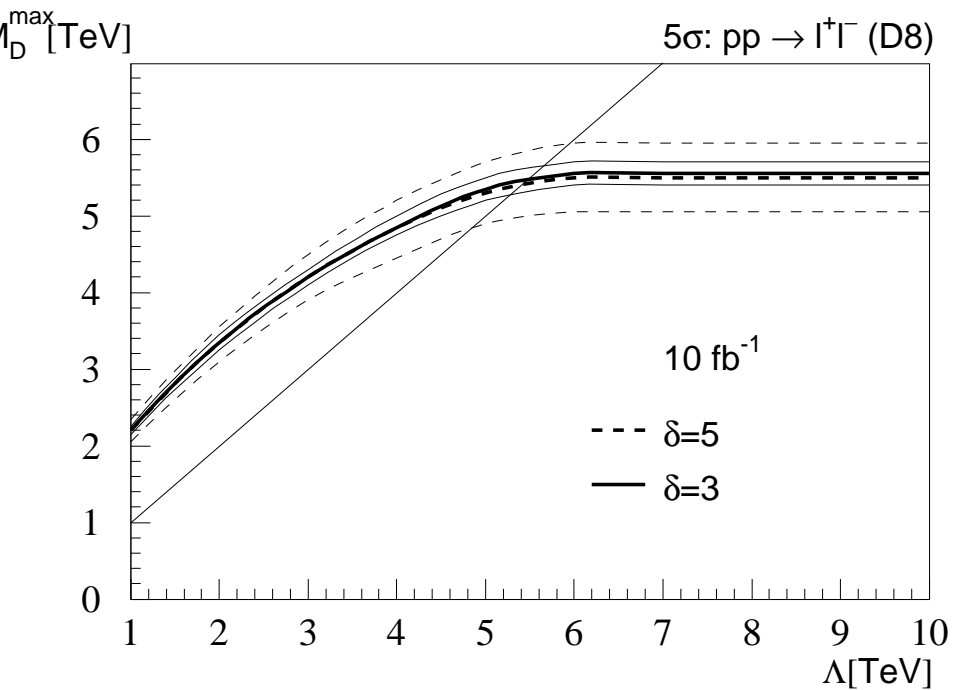
Giudice, Plehn, Strumia ('04)



gravitational Drell-Yan

- renormalisation group + Monte Carlo simulation

DL, Plehn ('07)



mini-black holes

Dimopoulos, Landsberg ('01)
Giddings, Thomas ('01)

- **classical Schwarzschild black holes**

metric

$$ds^2 = -f(r) dt^2 + f^{-1}(r) dr^2 + r^2 d\Omega_{d-2}^2, \quad f = 1 - \frac{G_N M}{r^{d-3}}$$

classical Schwarzschild radius

$$r_{\text{cl}} = (G_N M)^{1/(d-3)}$$

- **production cross section**

semi-classical

$$\hat{\sigma} = F \times \pi r_{\text{cl}}^2 (M = \sqrt{s}) \times \theta(\sqrt{s} - M_{\text{min}})$$

form factor F

mini-black holes

- **RG improved black holes**

Falls, DL, Raghuraman (ERG '08, and to appear)

running gravitational coupling

$$G_N \rightarrow G(r), \quad f(r) \rightarrow f_{\text{imp}}(r) = 1 - \frac{G(r) M}{r^{d-3}}$$

improved Schwarzschild radius r_s from

$$f'_{\text{imp}}(r_s) = 0$$

critical black hole mass M_c from

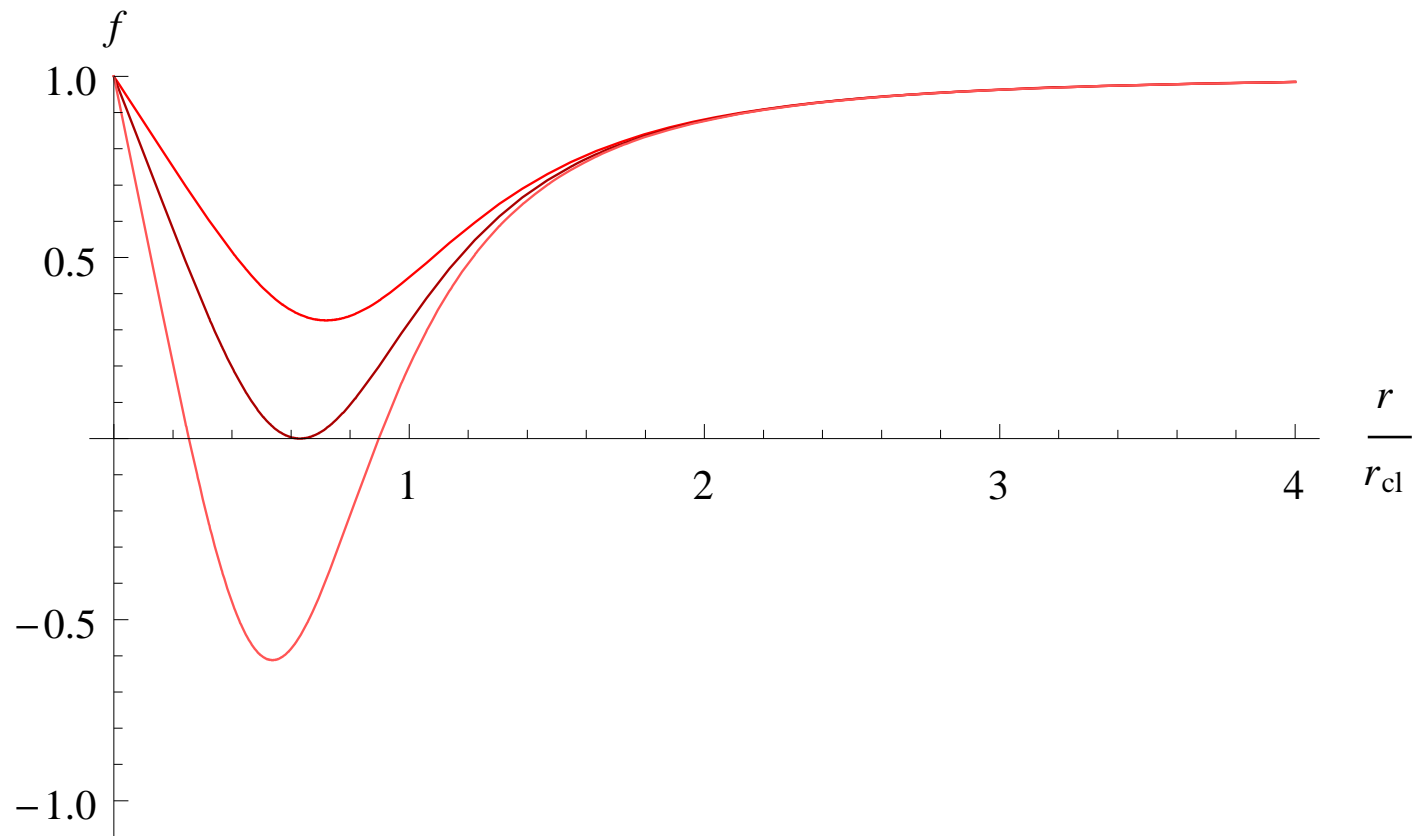
$$(d - 3)M_c = r \partial_r G(r)|_{r=r_c(M_c)}$$

mini-black holes

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metric, dependence on M (**D=6**)

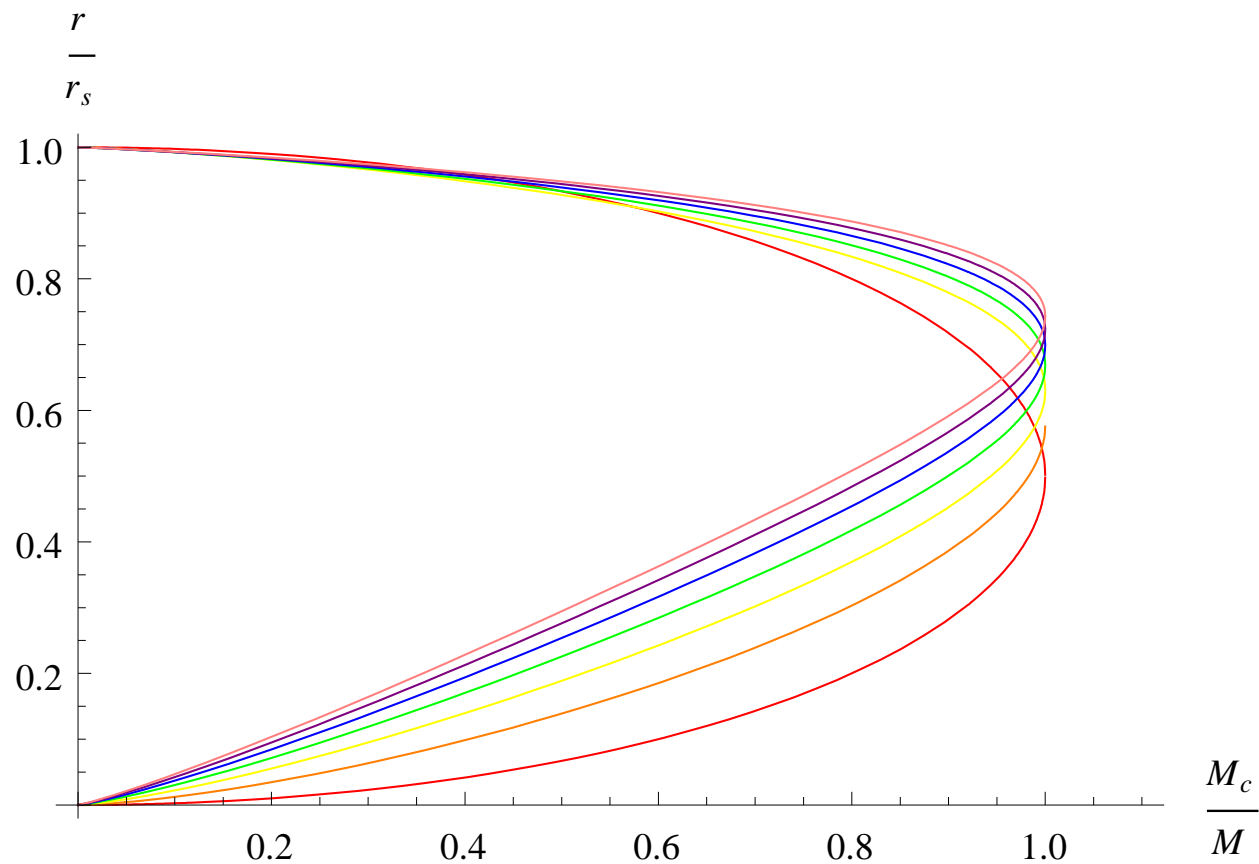


mini-black holes

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improved Schwarzschild radii, various dimension



BH production at the LHC

- semi-classical vs renormalisation group

Hiller, DL (to appear)

production cross section at the LHC $pp \rightarrow$ **final state**

$$\sigma = \sum_{i,j} \int_0^1 dx_1 \int_0^1 dx_2 f_i(x_1) f_j(x_2) \hat{\sigma}(q_i q_j \rightarrow \mathbf{final\ state})$$

parton distribution functions from **CTEQ61**
evaluated at $Q^2 = M_{\text{BH}}^2$.

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elastic BH production $pp \rightarrow$ **BH**

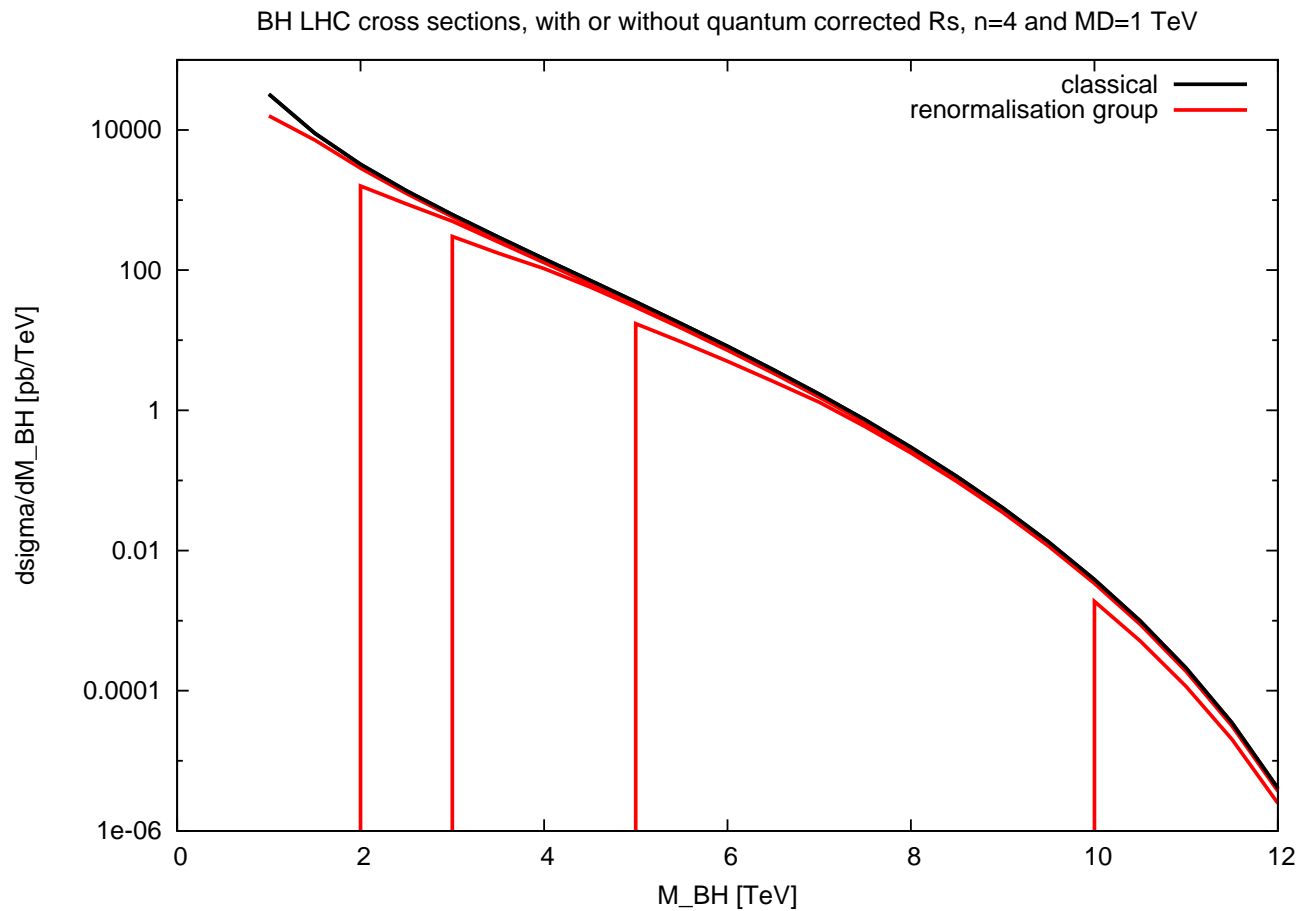
$$\frac{d\sigma}{dM} = \frac{2M}{s} \sum_{i,j} \int_{M^2/s}^1 \frac{dx}{x} f_i\left(\frac{M^2}{xs}\right) f_j(x) \hat{\sigma}(q_i q_j \rightarrow \text{BH})|_{\hat{s}=M^2}.$$

BH production at the LHC

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$n = 4$ extra dimensions



conclusions

- **low scale quantum gravity**

exciting idea, signatures at particle colliders

many scenarios, eg. ADD, RS I, RS II, $N = 10^{32}, \dots$

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renormalisation group: non-trivial UV fixed point

asymptotic safety in four and higher dimensions

Reuter (1996), Souma (1999)

Lauscher, Reuter (2001), Reuter, Saueressig (2001)

Forgacs, Niedermayer (2002), Niedermayer (2002)

DL (2003), Percacci, Perini (2003)

Bonanno, Reuter (2004), Percacci (2004)

Bonanno (2005), Lauscher, Reuter (2005)

Percacci (2005), Fischer, DL (2006)

Codello, Percacci (2006)

Codello, Percacci, Rahmede (2007), DL (2008), \dots

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many scenarios, eg. ADD, RS I, RS II, $N = 10^{32}, \dots$

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- **signatures of quantum gravity at the LHC**

if Planck scale of order TeV

QG effects accessible, insensitive to UV cutoff

implications for eg. di-lepton production, mini-black holes

conclusions

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- **challenges for theory and experiment**

determine the Planck scale

differentiate amongst quantum gravity scenarios

distinguish QG from other new physics signatures

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thanks!