

# I] Quantum Field Theory

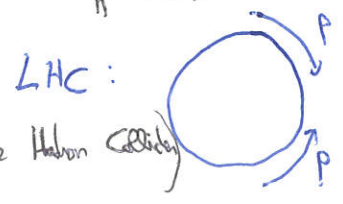
\*  $E \ll mc^2$  (at low energies) : relativity is unimportant  
 $\Rightarrow$  NRQM (Non Relativistic Quantum Mechanics) works perfectly!

eg.  $N_{particle} = C^t$  (Number of particle is a conserved quantity)

\*  $E \gtrsim mc^2$  (at high energies) : relativity is important (and things are more complicated)

$\Rightarrow$  Enough energy to pop out new particles of the vacuum.

• application :



Exp.  $\sim 13 \text{TeV}$  ( $m_p \sim 1 \text{GeV}$ )

processes such as :  $p + p \rightarrow p + p + \pi^0$

$p + p \rightarrow p + p + p + \bar{p}$

+ Higgs, ... (or more exotic particles)

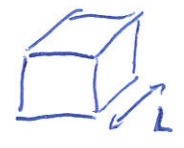
$\rightarrow$  many-body problem

$\rightarrow$  Need to construct a many-particle quantum theory to describe that.

## 1) Set-up QFT

a. Multi particle basis states (Fock space)

\* System in a box



The allowed values of  $\vec{k}$  for the wave function are discrete :

$$\vec{k} = \left( \frac{2\pi n_x}{L}, \frac{2\pi n_y}{L}, \frac{2\pi n_z}{L} \right)$$

• a state :  $|\vec{k}_1, \vec{k}_2, \dots, \vec{k}_{N_{TOT}}\rangle$  or  $|\dots\rangle = (a_{\vec{k}_1} + \dots + a_{\vec{k}_{N_{TOT}}}) |\dots\rangle$

BUT  $N_{TOT}$  is explicit

• occupation nb representation (without specifying  $N_{TOT}$ ) :  $|\dots, m(\vec{k}^i), m(\vec{k}^j), \dots\rangle$

\* def Number op.  $N(\vec{k}) |\dots\rangle = m(\vec{k}) |\dots\rangle$

nb of particles for each momentum

$\hookrightarrow$  count the occupation number for a given  $\vec{k}$

→  $H = \sum_{\vec{k}} \omega_{\vec{k}} N(\vec{k})$  , for 1k :  $H = \omega N$  , does it ring a bell ?

(2)

\* (Simple Harmonic Oscillator) SHO  $H_{SHO} = \omega \left( N + \frac{1}{2} \right)$   
 ↑  
 excitation level of the oscillator

⇒ (1-1 correspondence with an infinite system of H.O.)

$$H \approx \sum_{i=1}^{\infty} H_{HO}^i$$

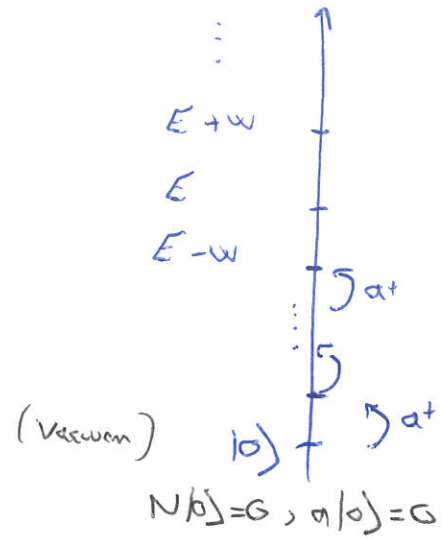
b. Simple Harmonic Oscillator

$$H = \frac{p^2}{2m} + \frac{1}{2} \omega^2 m X^2 = \frac{\omega}{2} (p^2 + q^2) \quad \text{with } [p, q] = -i$$

• def  $\left\{ \begin{array}{l} \text{lowering} \\ \text{raising} \end{array} \right. \varphi. \quad \left\{ \begin{array}{l} a = \frac{q + ip}{\sqrt{c}} \\ a^\dagger = \frac{q - ip}{\sqrt{c}} \end{array} \right. \quad \text{with } [a, a^\dagger] = 1 \rightarrow H = \omega \left( \underbrace{a^\dagger a}_{N(m) = m/m} + \frac{1}{2} \right)$

• ↳ They generate a ladder of states

$$\begin{aligned} H |E\rangle &= E |E\rangle \\ H a^\dagger |E\rangle &= (E + \omega) a^\dagger |E\rangle \\ H a |E\rangle &= (E - \omega) a |E\rangle \end{aligned}$$



c. Operator formalism for Fock space

→ We apply this formalism  $\left( \begin{array}{l} \text{annihilation} \\ \text{creation} \end{array} \right. \varphi. \quad \left( a_{\vec{k}} \text{ \& } a_{\vec{k}}^\dagger \right)$  for each momentum :

1 particle :  $|\vec{k}\rangle = a_{\vec{k}}^\dagger |0\rangle$   
 2 = :  $|\vec{k}, \vec{k}'\rangle = a_{\vec{k}}^\dagger a_{\vec{k}'}^\dagger |0\rangle$

$$H = \sum_{\vec{k}} \omega_{\vec{k}} a_{\vec{k}}^\dagger a_{\vec{k}}$$

- We set-up the operator formalism for a multi particle theory, based on SHO.
- Any observables may be written in terms of  $a, a^\dagger$ , very unfull.
- Need to construct the dynamics

Classical Particle Mechanics:

state of a system described by coordinate  $q_a(t)$

dynamics:  $L(q_1, \dots, q_n, \dot{q}_1, \dots, \dot{q}_n, t) = T - V$  potential  $E$   
kinetic  $E$   
 $S = \int_{t_1}^{t_2} L(t) dt$

general Hamilton's principle:  $(E-L eq)$

$$\frac{\partial L}{\partial q_a} = \dot{p}_a \quad \text{with} \quad p_a \equiv \frac{\partial L}{\partial \dot{q}_a}$$

Quantum Particle Mechanics:

deriving  $q_a(t)$

$$\begin{cases} q_a(t) \\ p_a(t) \end{cases} \rightarrow \begin{cases} \hat{q}_a(t) \\ \hat{p}_a(t) \end{cases} \text{ operator-valued functions}$$

↖ work in Heisenberg picture

expecting:  $[\hat{p}_a(t), \hat{q}_b(t)] = -i\delta_{ab}$

$$[\hat{p}_a, \hat{q}_b] = [\hat{p}_a, \hat{q}_b] = 0$$

Schrödinger picture:

$1 > g(t)$   
 $q_a$  small  $g \neq 0$

$$i\hbar \frac{d}{dt} |K(t)\rangle = H |K(t)\rangle \Rightarrow |K(t)\rangle = e^{-iH(t-t_0)} |K(t_0)\rangle$$

Heisenberg picture:

$1 > \text{and } L$   
 $Op. = f(t)$

$$i\hbar \frac{d}{dt} O_H(t) = [O_H(t), H] \quad \text{and} \quad |K(t)\rangle = e^{-iH(t-t_0)} |K(t_0)\rangle \rightarrow Op. \text{ satisfy Heisenberg eq.}$$

Classical Field Theory: (or: Classical electrodynamics)

observables ( $ex: E$ ) are defined at each point in space-time  $\phi_a(x,t)$

$$L(t) = \int d^3x \mathcal{L}(\phi_a(x), \partial_\mu \phi_a(x))$$

kinetically

$$S = \int_{t_1}^{t_2} dt \int d^3x \mathcal{L}(x,t)$$

com:  $\frac{\delta \mathcal{L}}{\delta \phi_a} = \partial_\mu \pi_a^\mu$  with  $\pi_a^\mu = \frac{\delta \mathcal{L}}{\delta(\partial_\mu \phi_a)}$

or:  $\mathcal{L} = \frac{1}{2} [\partial_\mu \phi_a \partial^\mu \phi_a - m^2 \phi_a^2]$  com:  $(\partial_\mu \partial^\mu + m^2) \phi_a = 0$

for massless field

QFT: To quantize an classical theory we do exactly what we did

to quantize classical particle mechanics

$$\begin{cases} \phi_a(x) \\ \pi_a^\mu(x) \end{cases} \rightarrow \begin{cases} \hat{\phi}_a \\ \hat{\pi}_a^\mu \end{cases} \text{ op. valued functions}$$

$$[\hat{\phi}_a(x_1), \hat{\pi}_b^\mu(x_2)] = i\delta_{ab} \delta^3(x_1 - x_2)$$

$$[\hat{\phi}_a, \hat{\phi}_b] = [\hat{\pi}_a^\mu, \hat{\pi}_b^\nu] = 0$$

example: scalar field solution of the K-G eq:

$$\phi(x) = \int d^3k [a_k e^{-ik \cdot x} + a_k^\dagger e^{ik \cdot x}] \quad \text{with} \quad [a_k, a_{k'}^\dagger] = \delta^{(3)}(k - k')$$

commutators of creat/annihilat op.

↳ sum over all momenta of creat/ann op.

$$\rightarrow H = \int d^3k \omega_k [a_k^\dagger a_k + \frac{1}{\epsilon} \delta^{(3)}(0)]$$

∞ but only ΔE has a physical meaning (we renormalize with E(ground state) = ∞)

\* quantizing classical field theory → forced upon us a particle interpretation of the field

$$\phi(\vec{x}, 0)|0\rangle = \int \frac{d^3k}{2\omega_k} e^{-i\vec{k} \cdot \vec{x}} |k\rangle$$

field op act on vacuum

picks out a linear combination of momentum eigenstates

- metaphor with hammer on vacuum
- creation of particle at  $\vec{x}$
- when it act on n particles it produces  $n+1$  ←  $n-1$ .

## 2) S-matrix, perturbation theory and Feynman diagrams

a. The interacting picture

- \* do for only free fields
- \* We now consider a more interesting case when fields interact between each other

\* free fields → interacting fields

\* We would like (as before)  $\phi = \phi(a, at)$  because we know how it acts on the states of the theory

\* But w.r to Heisenberg eq. are no longer plane waves but much more complicated

\* Trick called interaction picture:

$$H = H_0 + H_I \quad \text{def: } | \psi(t) \rangle_I \equiv e^{iH_0 t} | \psi(t) \rangle_S \quad \left( \text{if } H_I = 0 \Rightarrow I = S \right)$$

(free fields)      (contains interaction)

$$\text{E.O.M of fields: } i \frac{d}{dt} \phi_I(t) = [\phi_I(t), H_0]$$

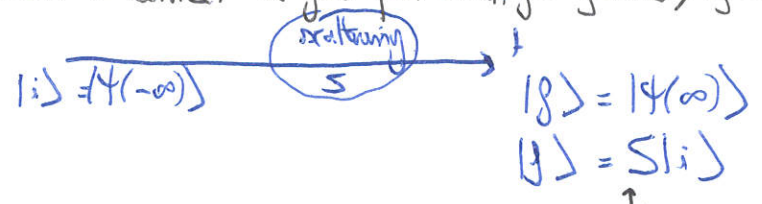
Great! because fields in an I.P. evolve just like free field in the H.P.

→ we can use all the free fields results we obtained so far!

\* all the complications have been relegated to the E.O.M of states :

$$i \frac{d}{dt} |\psi(t)\rangle_{\mathcal{I}} = H_{\mathcal{I}}(t) |\psi(t)\rangle_{\mathcal{I}}$$

\* we want to connect the far past with far future, after collision has taken place :



↑ scattering op.

\* def: S-matrix element:  $\langle f | S | i \rangle$  : amplitude to find the system in some given state  $|f\rangle$  starting from  $|i\rangle$ .

we integrate both sides

$$|\psi(t)\rangle = |i\rangle - i \int_{-\infty}^t dt_1 H_{\mathcal{I}}(t_1) |\psi(t_1)\rangle$$

we don't know that so we replace by

$$|\psi(t)\rangle = |i\rangle - i \int_{-\infty}^t dt_1 H_{\mathcal{I}}(t_1) |i\rangle + (-i)^2 \int_{-\infty}^t dt_2 H_{\mathcal{I}}(t_2) H_{\mathcal{I}}(t_1) |\psi(t_1)\rangle$$

by iterations ...  $\infty$  times taking the limit  $t \rightarrow \infty$  :

$$S = \sum_{n=0}^{\infty} (-i)^n \int_{-\infty}^{\infty} dt_1 \int_{-\infty}^{t_1} dt_2 \dots \int_{-\infty}^{t_{n-1}} dt_n H_{\mathcal{I}}(t_1) \dots H_{\mathcal{I}}(t_n)$$

Dyson's formula :  $S = T \left( e^{-i \int_{-\infty}^{\infty} dt H_{\mathcal{I}}(t)} \right)$   
 ordering product

b. Perturbativity, Feynman amplitudes

- time dep. of  $g$  is trivial (free field E.O.M)

- states is complicated !  $\rightarrow$  is going to be taken into account perturbatively

\* let's try to get an idea of how perturbation theory is going to work in I.P. :

\* let's take  $H_{\mathcal{I}} = g \psi^\dagger \psi \phi$   
 complex field      scalar field

$$\phi(x) = \int d^3k [ a_{\mathbf{k}} e^{-ikx} + a_{\mathbf{k}}^\dagger e^{ikx} ]$$

annihilate      create      meson

$$\psi(x) = \dots b_{\mathbf{k}} \dots c_{\mathbf{k}}^\dagger \dots$$

$$\psi^\dagger(x) = \dots c_{\mathbf{k}} \dots b_{\mathbf{k}}^\dagger \dots$$

annihilate      create anti-meson      (if spin 1/2 ~ fermion)  
 annihilation      create meson

\* small  $g \gg g^2 \gg g^3 \dots$  ( $H_I(x_1) \gg H_I(x_1)H_I(x_2) \gg \dots$ )

⑥

perturbation theory: imposing a cut makes sense

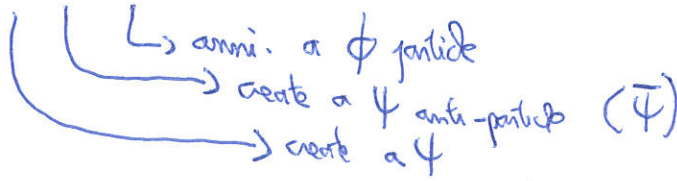
\* first order in  $g$ :

$$\langle g | H_I(x_1) | i \rangle = \langle g | g \begin{matrix} \psi^\dagger(x_1) \psi(x_1) \phi(x_1) \\ \left( \begin{smallmatrix} c \\ b^\dagger \end{smallmatrix} \right) \left( \begin{smallmatrix} b \\ c^\dagger \end{smallmatrix} \right) \left( \begin{smallmatrix} a \\ a^\dagger \end{smallmatrix} \right) \end{matrix} | i \rangle$$

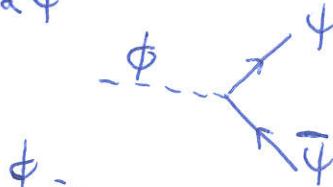
$\sim$  Feynman Amplitude  
 set of rules called Feynman rules

o int. doesn't conserve  $N_{TOT}$  (we made it!) and it contributes to a number of processes.

ex:  $b^\dagger c^\dagger a$ :

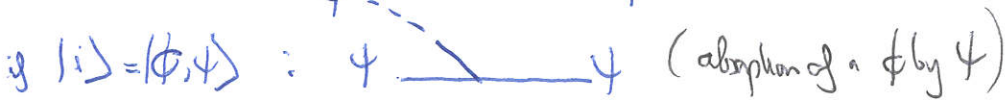


if  $|i\rangle = |\phi\rangle$ :



this is a number!

ex:  $c^\dagger c a$ :

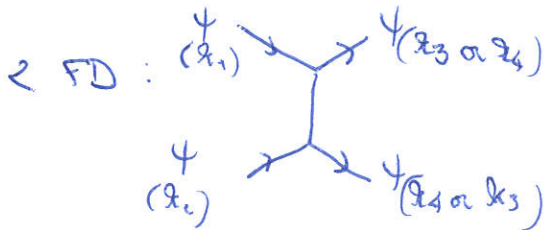


\* second order in  $g$ : H acts twice on the  $|i\rangle$

$$\langle g | H_I(x_1) H_I(x_2) | i \rangle = g^2 \langle g | \psi^\dagger(x_1) \psi(x_1) \phi(x_1) \psi^\dagger(x_2) \psi(x_2) \phi(x_2) | i \rangle$$

Let's consider the scattering process:  $\psi + \psi \rightarrow \psi + \psi$

$$\begin{cases} |i\rangle = |\vec{p}_1(N), \vec{p}_2(N)\rangle = b^\dagger(p_1) b^\dagger(p_2) |0\rangle \\ |f\rangle = \vec{p}_3 \leftarrow \leftarrow \leftarrow (p_3) - (p_4) \leftarrow \end{cases}$$



$\longleftrightarrow$  S. Amplitude in  $g^2$ :

$$\mathcal{A} = g^2 \left[ \frac{1}{(p_3 - p_1)^2 - m^2} + \frac{1}{(p_4 - p_2)^2 - m^2} \right]$$