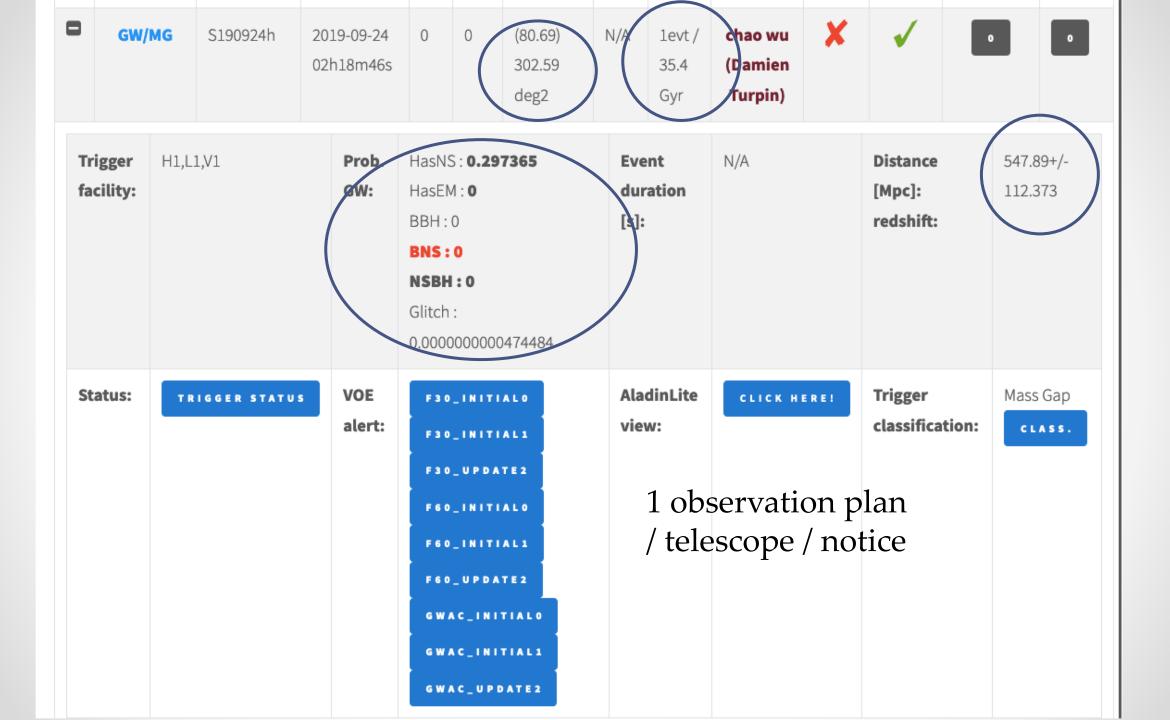
## GW triggers informations

What I have to care of ?

S. Antier, J.G Ducoin, D. Corre, N. Leroy

	gwalert   윤 28   육 0   & Ajouter un thème		6 i ¢	Q Recherche		@	☆	÷
	new GCN notice	Mardi 24 septembre —	Nev	v notice				
	ID : S190924h vith status Initial EventPage: https://gracedb.ligo.org/superevents/S190924h/view/ Observation program launched				¢	€ ₽	\$	•••
51	05 GW alert 							
	GW NAME: 5170724h Trigger Time: 2019-09-24T02:18:46.84665 Instruments: H1,L1,V1 EventPage: https://gracedb.ligo.org/superevents/S190924h/view/ Search: CBC HasRemnant: 0.0 Delay since alert: 0:45:13.872291	54		servation pla send to CSC		oduo	ced	
	I	Dimanche 29 septembre						
	incoming-webhook APPLI 8 h 36							
	new GCN notice							
	ID : S190828j with stitus Update EventPage: https://gracedb.lige.org/Superevents/S190828j/view/ Observation program launched							
	GW alert							
	GW NAME : S190828j Trigger Time: 2019-08-28T06:34:05.75647 Instruments: H1,L1,V1 EventPage: https://gracedb.ligo.org/superevents/S190828j/view/ Search: CBC HasRemnant: 0.0 Delay since alert: 32 days, 0:01:45.539436	2						
	0 Envover un message à #gwalert						@	3



## Content of alert received

## Notice Contents

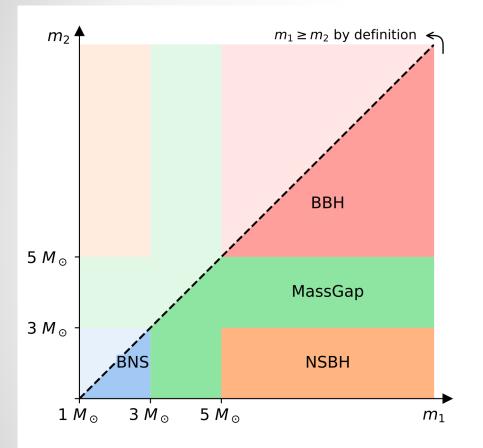
The table below is a representation of the contents of a LIGO/Virgo GCN Notice.

Root					
IVORN	<pre>ivo://gwnet/LVC#[{T,M}]SYYMMDDabc-{1,2,3}- {PreLiminary,Initial,Update,Retraction}</pre>				
Role	{observation,test}				
Who					
Date	Time sent (UTC, ISO-8601), e.g. 2018-11-01T22:34:49				
Author	LIGO Scientific Collaboration and Virgo Collaboration				
WhereWhen	Time of signal (UTC, ISO-8601), e.g. 2018-11-01T22:22:46.6544				
What					
GraceID	GraceDB ID: [{T, M}]SYYMMDDabc. Example: MS181101abc				
Packet Type	GCN Notice type: {Preliminary, Initial, Update, Retraction}				
Notice Type	Numerical equivalent of GCN Notice type: {150, 151, 152, 164}				
FAR	Estimated false alarm rate in Hz				
Sky Map	<pre>Versioned URL of HEALPix FITS sky localization file in the format https://gracedb.ligo.org/api/superevents /[{T,M}]SYYMMDDabc/files /{bayestar,LALInference,cWB}.fits.gz,[0-8].Example: https://gracedb.ligo.org/api/superevents/S190901ap/files /bayestar.fits.gz,0</pre>				

 https://emfollow.docs.ligo.org/userguide /content.html

Group	СВС	Burst			
Pipeline	{Gstlal,MBTAOnline,PyCBC,S PIIR}	{cWB,oLIB}			
CentralFreq		Central frequency in Hz			
Duration	N/A	Duration of burst in s			
Fluence		Gravitational-wave fluence in erg $\rm cm^{-2}$			
BNS, NSBH, BBH, MassGap, Noise	Probability that the source is a <u>BNS</u> , <u>NSBH</u> , <u>BBH</u> , or <u>MassGap</u> merger, or <u>terrestrial</u> (i.e, noise) respectively				
HasNS, HasRemnant	Probability, under the assumption that the source is not noise, that at least one of the compact objects was a neutron star, and that the system ejected a non-zero amount of neutron star matter, respectively.	N/A			

## Properties of the candidate



Terrestrial : a chance background fluctuation or a glitch

**Properties**: Probabilities that the source has each of the following properties, *assuming that it is not noise* (e.g., assuming that it is a BNS, NSBH, BBH, or MassGap merger):

- HasNS: The mass of one or more of the binary's two companion compact objects is consistent with a neutron star.
- HasRemnant: A non-zero amount of neutron star material remained outside the final remnant compact object (a necessary but not sufficient condition to produce certain kinds of electromagnetic emission such as a short GRB or a kilonova).

All of the quantities in the Classification and Properties sections are model dependent to some extent: the Classification section takes into consideration prior knowledge of astrophysical compact binary merger rates from previous LIGO/Virgo observations, and both the Classification and Properties sections depend on details of neutron star physics (e.g. maximum NS mass, equation of state). See the earlier subsection of the Data Analysis section for implementation details.