Searches for SUSY in 2 and 3 lepton final states with ATLAS and SUSY global fits with GAMBIT

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THE UNIVERSITY of ADELAIDE February 4th 2019

Presented at CPPM

Outline

The physics program of ATLAS at the LHC

- New Physics (in particular SUSY) searches
- Status of Standard Model measurements
- Techniques we use for some searche
 - An introduction to Recursive Jigsaw Reconstruction (RJR)
- Application of RJR to searches for Electroweak SUSY with ATLAS data and recent results
- Constraints on EW-MSSM with GAMBIT
- Summary







 Hierarchy problem: Higgs mass subject to quadratically divergent loop corrections.
 → Incredible fine-tuning



 Grand unification: Standard Model coupling constants do not unify at high scales.
 → SM does not imply a Grand Unified Theory



 Dark matter: Cosmological data suggest presence of dark matter → No explanation within Standard Model







Analogies paint the picture





SM has a snowman's chance in hell

Give me a





Analogies paint the picture





We need SUSY!

- Fundamental symmetry between fermions and bosons introducing a set of new partner particles to the SM particles with half-spin difference.
- ✓ Opposite-sign loop corrections from SUSY particles. Quadratic divergencies cancel. → No (little) fine-tuning.
- ✓ If R-parity conserved: Lightest SUSY Particle (LSP) stable. → Natural candidate for dark matter.



✓ Unification of gauge couplings at $M_{GUT} \approx 10^{16} \text{ GeV}$





We need SUSY!

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- Higgs boson discovery and strong experimental bounds have put vanilla SUSY under pressure
- Within the MSSM stop and gluino masses enter at 1 and 2 loop level into the Higgs mass matrix, the Higgsino mass parameter µ at tree level
- → Search effort focus around "Natural SUSY" (e.g. <u>arXiv:1110.6926</u>) with relatively light gluinos, stops, higgsinos (remaining SUSY particles can be decoupled at high masses)





The ATLAS experiment



- Solenoidal magnetic field (2T) in the central region – momentum measurement
- Energy meas. down to ~1° to the beamline
- High resolution silicon detectors
- Granular EM and Had calorimetry
- Independent muon spectrometer
- Good coverage permits reconstruction of missing transverse momentum through object reconstruction







Data Samples – Run 2

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Exceptional LHC performance in 2015-2018 Improved luminosity and recording efficiency throughout the run.

Integrated ~150 fb⁻¹ at end of Run2 pp collisions



ي100 چ

98

96

94

92

90

Cumulative Recording Efficiency

TLAS Interna

2011 pp (s = 7 TeV

2012 pp (s = 8 TeV

2015 pp vs = 13 TeV

2017 pp

2018 pp

Q

s = 13 TeV

s = 13 TeV



LHC: More than nominal Luminosity



LHC design: $L = 1.0 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ Achieved (2016): $L = 1.4 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ Achieved (2018): $L = 2.1 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$

More lumi makes for a more challenging environment to extract results of interest





Data analysis

MC analysis







Select Objects/Data of Interest











Measuring the Standard Model with ever increasing accuracy









SM - Backgrounds to SUSY searches









HOWTO search for SUSY

If SUSY particles are accessible at LHC energies

If R-Parity is conserved (RPV and Long-lived SUSY scenarios would be a different talk entirely)

- Pair-production via strong/EW interaction
- Direct or cascade decays to stable LSP
- Potentially many high pT SM objects and large missing transverse momentum





Search strategy for LHC Run2:

- Looks for gluinos/squarks decaying to jet enriched final states due to large XS
- Then focus on EWKino searches as they become more accessible and sensitive





HOWTO search for SUSY



- ① Build signal regions (**SR**s) based on requirements on signal / background discriminating variables to target specific SUSY event topologies. Optimised for discovery & exclusion.
- 2 Determine Standard Model background in the SRs:







illustration by M-H Genest

SUSY, this is what we often "claim" we're searching for...





SUSY: Strong, 3rd gen and Electroweak Production



W

 $\frac{W}{q}$

Squark and Gluino mediated light jets





+ many more

3rd generation squarks



EWKino and slepton production











D. Alves et al J. Phys. G: Nucl. Part. Phys. 39 (2012) 105005

The way in which we design, and optimize, searches at LHC.....

.....not just an organising principle, this is what we search for!





Simplified Models











Data analysis

MC analysis





Data 2015 and 2016

tī(+EW) & single top

1800 200

p_TS [GeV]

10

200 400

600 800 1000 1200

Data / MC

SM Total

W+jets

Z+jets

Dibosor

Multi-ier

600 800 1000 1200 1400 1600



10

200

400 600 800 1000 1200 1400 1600 1800 2000

p^{CM}_{TS}[GeV]

Data / MC

1800 2000

p^{CM}_{TS}[GeV]





GeV

Events / 200

Data / MC

10⁸

10⁷

10⁶

10⁵

10

10

10

ATLAS Preliminary

vs=13 TeV, 36.1 fb

CBW for BJB-C1

400



ARC Centre of Excellence for Particle Physics at the Terascale



Plethora of observables used by SUSY searches to maximally exploit event information:

Reconstructed object multiplicities, momenta, energies, e.g. $N_{jet/b-tag/l/\gamma}$, p_T , $E_{T,miss}$, ...

Scale variables, e.g. $\mathbf{m}_{eff} = \Sigma p_T + E_{T,miss}$,

Angular variables, e.g. min $\Delta \Phi$ (jet, E_{T,miss}), ...

Mass variables, e.g. m_{ℓ} , $m_{T}^{b/\ell/j}$, $\Sigma m_{fat-jet}$, ...

Event shape variables, e.g. Aplanarity, ...

Hypothesis-based event variables e.g. m_{T2}, ...

More **complete** methods, e.g. new **recursive jigsaw reconstruction** [arxiv:1607.08307], ...





Missing Transverse Momentum



$$ec{E}_T^{miss} \equiv -\sum_i^{ ext{calo}} ec{E}_T^{\ i}$$

Infer presence of weakly interacting particles in LHC events by looking for missing transverse energy....may be composed of one or more objects, which may differ

We can learn more by using other information in an event to contextualize the missing transverse momentum \Rightarrow multiple weakly interacting particles?







Recursive Jigsaw Reconstruction

New(ish) approach to reconstructing open final states



The strategy is to transform observable momenta iteratively *reference-frame to reference-frame*, traveling through each of the reference frames relevant to the topology

<u>**Recursive</u>**: At each step, specify only the relevant d.o.f.related to that transformation \Rightarrow apply a *Jigsaw Rule*.</u>

Repeat procedure recursively according to particular rules defined for each topology (the topology relevant to each reference frame)

Jigsaw: Each of these rules is factorizable/customizable/interchangeable like (strange) jigsaw puzzle pieces

Rather than obtaining one observable, get a *complete basis* of useful observables for each event





RJR technique

- Original method to reconstructing final states with weakly interacting particles.
- Transform observable momenta
 reference-frame to reference-frame
- Jigsaw rules: specify the unknown d.o.f. relevant to the transformation (customizable-interchangeable like jigsaw puzzle pieces)
- The procedure is repeated **recursively**, travelling through each of the reference frames relevant to the topology



• Rather than obtaining one observable, get a complete basis of useful variables diagonalized with physical observable: angles, energies, masses ...

PJ, C. Rogan, Phys. Rev. D96 112007 (2017) PJ, C. Rogan, M. Santoni, Phys. Rev. D95 035031 (2017) M. Santoni, "Probing Supersymmetry with Recursive Jigsaw Reconstruction", PhD Thesis Uni. Adelaide (Dec 2017) M. Santoni, JHEP 1805 058 (2018)





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SUSY searches with RJR



Scale Variables





where:

n : number of *visible ojects* considered as independent

m : number of *invisible ojects* considered as independent

 \mathcal{F} : frame under examination (can be PP ($\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0$) or P ($\tilde{\chi}_1^{\pm}$ or $\tilde{\chi}_2^0$))

Examples used in this analysis:

$$H_{1,1}^{PP} = (\ell_1 + \ell_2 + \ell_3)^{PP} . P() + (\tilde{\chi}_{1a}^0 + \tilde{\chi}_{1b}^0 + \nu_a)^{PP} . P()$$

$$HT_{4,1}^{PP} = \ell_1^{PP} . Pt() + \ell_2^{PP} . Pt() + jet_1^{PP} . Pt() + jet_2^{PP} . Pt() + (\tilde{\chi}_{1a}^0 + \tilde{\chi}_{1b}^0)^{PP} . Pt()$$

$$H_{2,1}^{P_a} = \ell_1^{P_a} . P() + \ell_2^{P_a} . P() + \tilde{\chi}_{1}^{0P_a} . P()$$

$$H_{2,1}^{P_b} = jet_1^{P_b} . P() + jet_2^{P_b} . P() + \tilde{\chi}_{1}^{0P_b} . P()$$





Compressed RJR



- In order to observe kinematic differences between signal and background we need an ISR system to give our sparticles a transverse kick: the response of the sparticle decay products is sensitive to the mass of the LSP
- In the limit of nearly degenerate parent sparticles \widetilde{p} and LSPs $\widetilde{\chi}$:

$$\vec{E}_T^{\text{miss}} \sim -\vec{p}_T^{\text{ISR}} \times \frac{m_{\tilde{\chi}}}{m_{\tilde{p}}}$$

Rather than relying on a clean mono-ISR signal we would like to be able to separate "ISR objects" from "sparticle objects"

- Accomplished with a simple decay view of the event
- CM: centre-of-mass system including all visible objects and MET
- ISR: radiation not coming from sparticle decays
- S: sparticle system
 - V: visible decay products
 - I: weakly interacting particles







Applications of RJR in ATLAS





General philosophy





M_P = Parent mass M_I = Invisible mass





Electroweak SUSY searches with RJR

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All North

600

 $m_{\widetilde{\chi}_{1}^{*}/\widetilde{\chi}_{2}^{0}}\left[\text{GeV}\right]$

700

500





The different shapes of these variables in the signal models as compared to the major backgrounds can be used in a more targeted way.

The interplay between the variables is also key - if we require one ratio to be large (for instance) it may make it **increasingly hard** for a complementary variable to have background events looking like signal events





Preselection for the 2lepton ISR tree



Similarly, where we require initial-state radiation, we need complementary variables to tease out sensitivity to a signal













3lepton Standard Tree Definitions



Region	$n_{ m leptons}$	$n_{ m jets}$ n	$n_{b-\mathrm{tag}}$	$p_{\rm T}^{\ell_1} [{\rm GeV}]$	$p_{\mathrm{T}}^{\ell_2}$ [GeV]	$p_{\mathrm{T}}^{\ell_3}$ [GeV]
CR3 <i>l</i> -VV	= 3	< 3	= 0	> 60	> 40	> 30
$VR3\ell$ - VV	= 3	< 3	= 0	> 60	> 40	> 30
SR3ℓ_High	= 3	< 3	= 0	> 60	> 60	> 40
$SR3\ell$ _Int	= 3	< 3	= 0	> 60	> 50	> 30
$SR3\ell_Low$	= 3	= 0	= 0	> 60	> 40	> 30
Region	$m_{\ell\ell} [{\rm GeV}]$	m_{T}^{W} [GeV]	$H_{3,1}^{\mathrm{PP}}$ [C	GeV] $\frac{p}{p_{\mathrm{T}}^{\mathrm{lab}}}$	$ \frac{lab}{T PP} - \frac{H^{PP}}{H^{PP}_{T 3,1}} - \frac{H^{P}_{T}}{H^{PP}_{T 3,1}} $	$\frac{P_{1,1}^{PP}}{I_{3,1}^{PP}} \qquad \frac{H_{1,1}^{Pb}}{H_{2,1}^{Pb}}$
$CR3\ell$ -VV	$\in (75, 105)$	$\in (0,70)$	>	250	< 0.2 >	0.75 –
$VR3\ell$ - VV	$\in (75, 105)$	$\in (70, 100)$	>	250	< 0.2 >	0.75 –
SR3ℓ_High	$\in (75, 105)$	▲ > 150	<u> </u>	550	< 0.2 >	0.75 h > 0.8
$SR3\ell_Int$	$\in (75, 105)$	> 130	>	450	< 0.15	> 0.8 > 0.75
$SR3\ell_Low$	$\in (75, 105)$	> 100	>	250	$\psi < 0.05 \psi >$	> 0.9



- Select events:
- with 3 high pT leptons
- l+l- pair at the Z-mass
- use RJ variables to define sensitive regions of phase space



Can leverage the behavior of the variables we design to target signals in a more natural way.

Similar selection optimization performed for 2lepton regions







Region	$n_{ m lep}$	tons	$n_{ m jets}$ r	$p_{b- ext{tag}} = p_{ ext{T}}^{\ell_{ ext{J}}}$	[GeV]	$p_{\rm T}^{\ell_2}$ [GeV]	$p_{\mathrm{T}}^{\ell_3}$ [GeV]
CR3ℓ_ISR-V	VV	= 3	≥ 1	= 0	> 25	> 25	> 20
VR3ℓ_ISR-V	VV	= 3	≥ 1	= 0	> 25	> 25	> 20
$SR3\ell$ _ISR		$= 3 \in$	[1,3]	= 0	> 25	> 25	> 20
Region	$m_{\ell\ell} \; [{\rm GeV}]$	m_{T}^{W} [Ge	V] $\Delta \phi_{\rm ISR,I}^{\rm CM}$	$R_{\rm ISF}$	$p_{\mathrm{T\ ISR}}^{\mathrm{CM}}$ [6	GeV] $p_{T I}^{CM}$ [GeV]	$p_{\rm T}^{\rm CM}$ [GeV]
$CR3\ell$ _ISR-VV	$\in (75, 105)$	< 1	> 2.0	$\in (0.55, 1.0]$)	> 80 > 60	< 25
$VR3\ell$ _ISR-VV	$\in (75, 105)$	>	50 > 2.0	$\in (0.55, 1.0)$) >	> 80 > 60	> 25
$SR3\ell$ _ISR	$\in (75, 105)$	> 1	> 2.0	$\in (0.55, 1.0]$) >	100 > 80	< 25



Complementarity between the R_{ISR} variable and $P_{T ISR}$ studied in detail in: PJ, C. Rogan, M. Santoni, PRD 95 035013 (2017)




Control Regions – 3lepton











Control Regions – 2lepton









Validation Regions









Observable 2

Unblinded results



- Main background contribution is from VV (3I), VV and Z+jets (2I)
- Control and Validation Regions enriched in these processes demonstrate that the key backgrounds are well modeled
- Z+jets prediction from a dedicated photon template sample
- We see excesses, in 4 signal regions, all targeting the low mass splitting





Results – 2lepton



Signal region	SR2ℓ_High	SR2ℓ_Int	SR2ℓ_Low	SR2ℓ_ISR
Total observed events	0	1	19	11
Total background events	1.9 ± 0.8	2.4 ± 0.9	8.4 ± 5.8	$2.7^{+2.8}_{-2.7}$
Other Fit output, $Wt + t\bar{t}$ Fit output, VV Z+jets	$\begin{array}{c} 0.02 \pm 0.01 \\ 0.00 \pm 0.00 \\ 1.8 \pm 0.7 \\ 0.07 \substack{+0.78 \\ -0.07} \end{array}$	$\begin{array}{c} 0.05^{+0.12}_{-0.05} \\ 0.00 \pm 0.00 \\ 2.4 \pm 0.8 \\ 0.00^{+0.74}_{-0.00} \end{array}$	$\begin{array}{c} 0.02^{+1.07}_{-0.02} \\ 0.57 \pm 0.20 \\ 1.5 \pm 0.9 \\ 6.3 \pm 5.8 \end{array}$	$\begin{array}{c} 0.06\substack{+0.33\\-0.06}\\ 0.28\substack{+0.34\\-0.28}\\ 2.3\pm1.1\\ 0.10\substack{+2.58\\-0.10}\end{array}$
Fit input, $Wt + t\bar{t}$ Fit input, VV	0.00 1.9	0.00 2.6	0.63 1.6	0.28 2.4







Results – 3 lepton



Signal region	SR3ℓ_High	SR3ℓ_Int	SR3ℓ_Low	SR3ℓ_ISR
Total observed events	2	1	20	12
Total background events	1.1 ± 0.5	2.3 ± 0.5	10 ± 2	3.9 ± 1.0
Other Triboson Fit output, VV	$\begin{array}{c} 0.03^{+0.07}_{-0.03} \\ 0.19 \pm 0.07 \\ 0.83 \pm 0.39 \end{array}$	0.04 ± 0.02 0.32 ± 0.06 1.9 ± 0.5	$\begin{array}{c} 0.02 \substack{+0.34 \\ -0.02} \\ 0.25 \pm 0.03 \\ 10 \pm 2 \end{array}$	$\begin{array}{c} 0.06^{+0.19}_{-0.06} \\ 0.08 \pm 0.04 \\ 3.8 \pm 1.0 \end{array}$
Fit input, VV	0.76	1.8	9.2	3.4







Results – ISR Signal Regions





We see different yields in data compared to our prediction in the ISR SRs, most prominently in the 3 lepton region (lower plots).





3lepton – ISR Signal Region









Results – Low Mass Signal Regions





Similarly, there are excess events in data compared to our prediction in the Low mass SRs. The upper right distribution *was not used* in the event selection.





3lepton – Low Mass Signal Region







0.9



Statistical interpretation



Signal region	$\langle \epsilon \sigma \rangle_{\rm obs}^{95}$ [fb]	$S_{\rm obs}^{95}$	S ⁹⁵ _{exp}	$p_0(Z)$
SR3ℓ_ISR	0.42	15.3	$6.9^{+3.1}_{-2.2}$	0.001 (3.02)
SR2ℓ_ISR	0.43	15.4	$9.7^{+3.6}_{-2.5}$	0.02 (1.99)
SR3ℓ_Low	0.53	19.1	$9.5_{-1.8}^{+4.2}$	0.016 (2.13)
SR2ℓ_Low	0.66	23.7	$16.1_{-4.3}^{+6.3}$	0.08 (1.39)
SR3ℓ_Int	0.09	3.3	$4.4^{+2.5}_{-1.5}$	0.50 (0.00)
SR2ℓ_Int	0.09	3.3	$4.6^{+2.6}_{-1.5}$	0.50 (0.00)
SR3ℓ_High	0.14	5.0	$3.9^{+2.2}_{-1.3}$	0.23 (0.73)
SR2ℓ_High	0.09	3.2	$4.0^{+2.3}_{-1.2}$	0.50 (0.00)

To remain as conservative as possible, and to avoid model dependent statements, *we do not combine the significances*



Excesses of 3.0σ , 2.0σ , 2.1σ and 1.4σ in the four regions targeting moderately compressed EWK SUSY.

This is the largest excess seen at any LHC experiment in a search for Supersymmetry





Statistical interpretation



Signal region	SR2ℓ_Low	SR2ℓ_ISR
ee	9 (4.5±3.9)	3 (1.2±1.2)
μμ	10 (3.9±2.6)	8 (1.5±1.5)
Signal region	SR3ℓ_Low	SR3ℓ_ISR
eee	6 (3.5±0.7)	3 (1.1±0.3)
ееµ	$6(2.0\pm0.4)$	$3(0.9\pm0.3)$
$\mu\mu\mu$	$7(2.7\pm0.6)$	$4(1.5\pm0.4)$
μμе	$1(1.9\pm0.4)$	2 (0.4±0.1)

The four signal regions with excesses were studied in terms of their flavour composition - looks as expected. MANY other cross-checks done....

Improved limits at high mass compared to previous analysis.....with weaker limits at low mass due to excesses observed.







Statistical interpretation





Analysis with the best reach in Electroweak searches with intermediate W/Z bosons.

In the process of updating this work with full Run 2 data: 36 ->150 fb⁻¹ (~4x more data)

Largest excess ($\geq 3\sigma$) in any SUSY search!





ATLAS – 4 lepton







Hints in some EWK SUSY channels would suggest we should see excesses in similar phase space.

Region	$N(e,\mu)$	$N(\tau_{\rm had-vis})$	$p_{\mathrm{T}}\left(au_{\mathrm{had-vis}} ight)$	Z boson	Selection	Target
SR0A	≥ 4	= 0	> 20 GeV	veto	$m_{\rm eff} > 600 {\rm GeV}$	General
SR0B	≥ 4	= 0	> 20 GeV	veto	$m_{\rm eff} > 1100 {\rm GeV}$	RPV <i>LLĒ</i> 12k
SR0C	≥ 4	= 0	> 20 GeV	require 1st & 2nd	$\begin{array}{l} E_{\mathrm{T}}^{\mathrm{miss}} > 50 \mathrm{GeV} \\ E_{\mathrm{T}}^{\mathrm{miss}} > 100 \mathrm{GeV} \end{array}$	higgsino GGM
SR0D	≥ 4	= 0	> 20 GeV	require 1st & 2nd		higgsino GGM
SR1	= 3	≥ 1	> 30 GeV	veto	$m_{\rm eff} > 700 { m GeV}$	RPV <i>LLĒi</i> 33
SR2	= 2	≥ 2	> 30 GeV	veto	$m_{\rm eff} > 650 { m GeV}$	RPV <i>LLĒi</i> 33

arXiv:1804.03602, Phys. Rev. D 98, 032009 (2018)







New Physics interpretation



Sample	SR0A	SR0B	SROC	SR0D	SR1	SR2
Observed	13	2	47	10	8	2
SM Total	10.2 ± 2.1	1.31 ± 0.24	37 ± 9	4.1 ± 0.7	4.9 ± 1.6	2.3 ± 0.8
ZZ tīZ	2.7 ± 0.7 2.5 ± 0.6	0.33 ± 0.10 0.47 ± 0.13	28 ± 9 3.2 + 0.4	0.84 ± 0.34 1.62 ± 0.23	0.35 ± 0.09 0.54 ± 0.11	0.33 ± 0.08 0.31 ± 0.08
Higgs VVV	1.2 ± 1.2 0.79 ± 0.17	0.13 ± 0.13 0.22 ± 0.05	0.9 ± 0.8 2.7 ± 0.6	0.28 ± 0.25 0.64 ± 0.14	0.5 ± 0.5 0.18 ± 0.04	0.32 ± 0.32 0.20 ± 0.06
Reducible Other	2.4 ± 1.4 0.53 ± 0.06	$\begin{array}{c} 0.000^{+0.005}_{-0.000}\\ 0.165 \pm 0.018 \end{array}$	$\begin{array}{c} 0.9^{+1.4}_{-0.9} \\ 0.85 \pm 0.19 \end{array}$	$\begin{array}{c} 0.23^{+0.38}_{-0.23} \\ 0.45 \pm 0.10 \end{array}$	3.1 ± 1.5 0.181 ± 0.022	1.1 ± 0.7 0.055 ± 0.012
$\langle \epsilon \sigma \rangle_{obs}^{95}$ fb	0.32	0.14	0.87	0.36	0.28	0.13
S ⁹⁵ _{obs}	12	4.9	31	13	10	4.6
S 95 exp	$9.3^{+3.6}_{-2.3}$	$3.9^{+1.6}_{-0.8}$	23^{+8}_{-5}	$6.1^{+2.1}_{-1.3}$	$6.5^{+3.5}_{-1.3}$	$4.7^{+2.0}_{-1.3}$
CL_b $p_{s=0}$	0.76 0.23	0.74 0.25	0.83 0.15	0.99 0.011	0.86 0.13	0.47 0.61
Ζ	0.75	0.69	1.0	2.3	1.2	0





2.3 σ deviation from SM in 4lepton EWKino search in region sensitive to \approx 200GeV

Still to be updated with 4x more data!











GAMBIT: The Global And Modular BSM Inference Tool

gambit.hepforge.org

EPJC **77** (2017) 784

arXiv:1705.07908

- Extensive model database not just SUSY
- Extensive observable/data libraries
- Many statistical and scanning options (Bayesian & frequentist)
- Fast LHC likelihood calculator
- Massively parallel
- Fully open-source

Members of:ATLAS, Belle-II, CMS, CTA,
Fermi-LAT, DARWIN, IceCube,
LHCb, SHiP, XENONAuthors of:DarkSUSY, DDCalc, Diver,
FlexibleSUSY, gamlike, GM2Calc,
Ica lat_pulike_PolyChord_Pivot

IsaJet, nulike, PolyChord, Rivet, SOFTSUSY, SuperIso, SUSY-AI, WIMPSim

- Fast definition of new datasets and theories
- Plug and play scanning, physics and likelihood packages



Collaborators:

Peter Athron, Csaba Balázs, Ankit Beniwal, Florian Bernlochner, Sanjay Bloor, Torsten Bringmann, Andy Buckley, Eliel Camargo-Molina, Marcin Chrząszc, Jan Conrad, Jonathan Cornell, Matthias Danninger, Tom Edwards, Joakim Edsjö, Ben Farmer, Andrew Fowlie, Tomás Gonzalo, Will Handley, Sebastian Hoof, Selim Hotinli, Felix Kahlhoefer, Suraj Krishnamurthy, Anders Kvellestad, Julia Harz, Paul Jackson, Tong Li, Greg Martinez, Nazila Mahmoudi, James McKay, Are Raklev, Janina Renk, Chris Rogan, Roberto Ruiz de Austri, Patrick Stoecker, Roberto Trotta, Pat Scott, Nicola Serra, Daniel Steiner, Puwen Sun, Aaron Vincent, Christoph Weniger, Sebastian Wild, Martin White, Yang Zhang



40+ participants in 10 Experiments & 14 major theory codes





GAMBIT code structure







Model independent LHC limits

- Custom parallelised Pythia MC + custom detector sim
- Can generate 20,000 events on 12 cores in < 5 s (we use a lot more than that for recent papers)</p>
- Then apply Poisson likelihood with nuisance parameters for systematics
- Combine analyses using best expected exclusion, unless public covariance matrix is available (CMS are being very helpful on this)





GAMBIT status

- GAMBIT was released as an open source public tool in 2017
- First physics studies include GUTscale SUSY models, the MSSM7, axion models and Higgs portal dark matter models
- See gambit.hepforge.org for more info, all samples are available via Zenodo

Eur. Phys. J. C manuscript No. (will be inserted by the editor)

GAMBIT: The Global and Modular Beyond-the-Standard-Model Inference Tool

The GAMBIT Collaboration: First Author^{a,1}, Second Author^{b,2} ¹ First Address, Street, City, Country ²Second Address, Street, City, Country

Received: date / Accepted: date





When supercomputers go over to the dark side

Pespice obdies on data and pienty of triedness, we still don't know what dark matter is: marten white and Pat Scott describe how a new software tool called GAMBIT will test how novel theories stack up when confronted with real data

The more meaning the strike mass the more than the the strike mass the strike

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Eur. Phys. J. C manuscript No. (will be inserted by the editor)

CoEPP-MN-18-7

Combined collider constraints on neutralinos and charginos

The GAMBIT Collaboration: Peter Athron^{1,2}, Csaba Balázs^{1,2}, Andy Buckley³, Jonathan M. Cornell⁴, Matthias Danninger⁵, Ben Farmer⁶, Andrew Fowlie^{1,2,9}, Tomás E. Gonzalo¹⁰, Julia Harz¹¹, Paul Jackson^{2,12}, Rose Kudzman-Blais⁵, Anders Kvellestad^{6,10,a}, Gregory D. Martinez¹³, Andreas Petridis^{2,12}, Are Raklev¹⁰, Christopher Rogan¹⁴, Pat Scott⁶, Abhishek Sharma^{2,12}, Martin White^{2,12,b}, Yang Zhang^{1,2}

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Abstract Searches for supersymmetric electroweakinos have entered a crucial phase, as the integrated luminosity of the Large Hadron Collider is now high enough to compensate for their weak production cross-sections. Working in a framework where the neutralinos and charginos are the only light sparticles in the Minimal Supersymmetric Standard Model, we use GAMBIT to perform a

relic density can be obtained through the Higgs-funnel and Z-funnel mechanisms, even assuming that all other sparticles are decoupled. All samples, GAMBIT input files and best-fit models from this study are available on <code>Zenodo</code>.

Contents

The GAMBIT collaboration have recently performed the most comprehensive study of the MSSM electroweakino sector to date

 We focussed on collider constraints from LHC and LEP, but also looked at the implications for dark matter (precise implications depend on the mass scale of the sparticles that we decoupled)





What if we assume there is no SUSY?

- We have the option of "capping" the LHC likelihood in our scan results, to prevent potential signals from providing a better fit to the data than the SM
- This amounts to testing the exclusion power of the included LHC searches
- We find *no general constraint* on the MSSM EW sector from the LHC in this case, and we also explain why (the searches are overoptimised on specific simplified SUSY models)





Uncapping the likelihood

- If we allow for the presence of a signal, our results get more interesting
- A particular mass scale is picked out by a series of anomalies in ATLAS and CMS searches
- All electroweakinos are light, and we either have:

Bino < winos < higgsinos Or Bino < higgsinos < winos









 Contribution from each analysis to the 1σ, 2σ and 3σ best-fit regions

 $\ln \mathcal{L}(s+b) - \ln \mathcal{L}(b)$

- Blue: better than background-only
 Red: worse than background-only
- Most important contributions to best-fit region:
 - · ATLAS_4lep
 - · ATLAS_RJ_3lep
 - · ATLAS_MultiLep_2lep_jet
 - ATLAS_MultiLep_3lep
 - · CMS_MultiLep_3lep











https://project-hl-lhc-industry.web.cern.ch/content/project-schedule



- LHC Run 1 and Run 2 have been completed.
 - Data analyses in full progress.
- LS2 (Long Shutdown 2019-2020) is about to start.
 - ATLAS Phase I Upgrade.
- LS3 (Long Shutdown 2024-2026) after Run 3.
 - LHC Major Upgrade to HL-LHC (High Luminosity LHC).
 - ATLAS Phase II Upgrade.





ATLAS Phase I Upgrade for LHC Run 3

- Run 3 LHC expectations: $L = 3x10^{24} \text{ cm}^{-2} \text{s}^{-1}$ (int. $L = 300 \text{ fb}^{-1}$) at $\sqrt{\text{s}} = 14 \text{ TeV}$.
- Need better trigger capabilities (efficiency and fake rejection).
- Maintain same acceptance and p_{τ} thresholds with higher pile-up.



Trigger and Data Acquisition System
 Fast Tracker (FTK), Topological Triggers,
 High level Trigger, Common Readout, ...

Inner Detector TRT, SCT, Pixels

- Software
- Optical Readout





Summary

• The search for new physics at the LHC continues

- There is considerable scope to improve techniques and methods used to perform measurements and searches with the LHC data.
- We have demonstrated this with one such method called *Recursive Jigsaw Reconstruction (RJR)*
- We used RJR to execute a search in 3lepton and 2lepton+jets final states that yields one of the largest excesses of any new physics search performed by ATLAS (if real it may well be beyond 5σ with the full Run2 data).
- These are included in a SUSY EWK-fit with GAMBIT, yielding results that point to a potential scale of new physics





CHEP 2019 – International Conference on Computing in High-Energy and Nuclear Physics



4-8 November, 2019 Adelaide, Australia

Backup slides

Questions we hear a lot from astrophysicists

"We keep hearing that the lightest neutralino is a good dark matter candidate"

"You've spent almost a decade not seeing supersymmetry at the LHC"

"What are the LHC constraints on lightest neutralino dark matter?"

The lightest neutralino is a natural dark matter candidate, and is the subject of most studies

illustration by M-H Genest

How the MSSM might appear...

Name	Spin	P_{R}	Gauge Eigenstates	Mass Eigenstates
Higgs bosons	0	+1	H^0_u H^0_d H^+_u H^d	h^0 H^0 A^0 H^{\pm}
squarks	0	-1	$\widetilde{u}_L \ \widetilde{u}_R \ \widetilde{d}_L \ \widetilde{d}_R$ $\widetilde{s}_L \ \widetilde{s}_R \ \widetilde{b}_L \ \widetilde{b}_R$ $\widetilde{t} \ \overline{b}_R$	(same) (same) \widetilde{t}_1 \widetilde{t}_2 \widetilde{b}_1 \widetilde{b}_2
sleptons	0	D -1	$\widetilde{e}_{L} \widetilde{e}_{R} \widetilde{\nu}_{e}$ $\widetilde{\mu}_{L} \widetilde{\mu}_{R} \widetilde{\nu}_{\mu}$ $\widetilde{\tau}_{L} \widetilde{\tau}_{R} \widetilde{\nu}_{\tau}$	(same) (same) $\widetilde{\tau}_1 \widetilde{\tau}_2 \widetilde{\nu}_{\tau}$
neutralinos	1/2	-1	$\widetilde{B}^0 \hspace{0.2cm} \widetilde{W}^0 \hspace{0.2cm} \widetilde{H}^0_u \hspace{0.2cm} \widetilde{H}^0_d$	$\widetilde{\chi}^0_1 \hspace{0.2cm} \widetilde{\chi}^0_2 \hspace{0.2cm} \widetilde{\chi}^0_3 \hspace{0.2cm} \widetilde{\chi}^0_4$
charginos	1/2	-1	\widetilde{W}^{\pm} \widetilde{H}^+_u \widetilde{H}^d	$\widetilde{\chi}_1^{\pm}$ $\widetilde{\chi}_2^{\pm}$
gluino	1/2	-1	\widetilde{g}	(same)

Parameters

15.4

19.1

23.7

16.1

0.43

0.53

0.66

SR2ℓ ISR

SR3ℓ_Low

SR2ℓ_Low

Ge,

0

Events

Data / Bkg

0.02 (1.99)

0.016 (2.13)

0.08(1.39)

Largely unique selection of events compared to earlier analysis on same dataset

0.1 0.2 0.3 0.4 0.5 0.6 0.7

0

Exclusions for high mass reach 600 GeV and low mass points cannot be excluded due to excesses

Signal Region	$\langle \epsilon \sigma \rangle_{ m obs}^{95} [{ m fb}]$	$S_{\rm obs}^{95}$	S_{exp}^{95}	$p(s=0)\left(Z\right)$
SR3ℓ_ISR	0.42	15.3	$6.9^{+3.1}_{-2.2}$	0.001 (3.02)
SR2ℓ_ISR	0.43	15.4	$9.7^{+\overline{3.6}}_{-2.5}$	0.02 (1.99)
SR3ℓ_Low	0.53	19.1	$9.5_{-1.8}^{+4.2}$	0.016 (2.13)
SR2ℓ_Low	0.66	23.7	$16.1_{-4.3}^{+6.3}$	0.08 (1.39)

0.8 0.9

 $H_{11}^{P_b}/H_{21}^{P_b}$



2lepton Standard Tree Definitions

Region	$n_{ m leptons}$	$n_{ m jets}$	$n_{b-\mathrm{tag}}$	$p_{\mathrm{T}}^{\ell_1,\ell_2}$ [GeV]	$p_{\mathrm{T}}^{j_1,j_2}$ [GeV]	$m_{\ell\ell} \; [{\rm GeV}]$	$m_{jj} [{ m GeV}]$	m_{T}^{W} [GeV]
$CR2\ell$ -VV	$\in [3,4]$	≥ 2	=0	> 25	> 30	$\in (80, 100)$	> 20	$\in (70, 100)$
								if $n_{\text{leptons}} = 3$
$CR2\ell$ -Top	= 2	≥ 2	=1	> 25	> 30	$\in (80, 100)$	$\in (40, 250)$	
$VR2\ell$ - VV	= 2	≥ 2	=0	> 25	> 30	$\in (80, 100)$	$\in (40,70)$	_
							or $\in (90, 500)$	—
$VR2\ell$ -Top	=2	≥ 2	=1	> 25	> 30	$\in (20, 80)$	$\in (40, 250)$	_
						or > 100		_
$\mathrm{VR}2\ell_{-}\mathrm{High}{-}\mathrm{Zjets}$	= 2	≥ 2	= 0	> 25	> 30	$\in (80, 100)$	$\in (0, 60)$	_
							or $\in (100, 180)$	_
$VR2\ell_Low-Zjets$	=2	= 2	= 0	> 25	> 30	$\in (80, 100)$	$\in (0, 60)$	—
							or $\in (100, 180)$	—
$\mathrm{SR}2\ell_{-}\mathrm{High}$	= 2	≥ 2	= 0	> 25	> 30	$\in (80, 100)$	$\in (60, 100)$	_
$SR2\ell$ _Int	=2	≥ 2	= 0	> 25	> 30	$\in (80, 100)$	$\in (60, 100)$	_
$\mathrm{SR}2\ell_{-}\mathrm{Low}$	= 2	= 2	= 0	> 25	> 30	$\in (80, 100)$	$\in (70, 90)$	_

Region	$H_{4,1}^{\mathrm{PP}}$ [GeV]	$H_{1,1}^{\mathrm{PP}}$ [GeV]	$\frac{p_{\mathrm{T}\ \mathrm{PP}}^{\mathrm{lab}}}{p_{\mathrm{T}\ \mathrm{PP}}^{\mathrm{lab}}+H_{\mathrm{T}\ 4,1}^{\mathrm{PP}}}$	$\frac{\min(H_{1,1}^{\mathrm{Pa}}, H_{1,1}^{\mathrm{Pb}})}{\min(H_{2,1}^{\mathrm{Pa}}, H_{2,1}^{\mathrm{Pb}})}$	$\frac{H_{1,1}^{\mathrm{PP}}}{H_{4,1}^{\mathrm{PP}}}$	$\Delta \phi_{ m V}^{ m P}$	${ m min}\Delta\phi(j_1/j_2,ec{p}_{ m T}^{ m miss})$
CR2ℓ-VV	> 200	_	< 0.05	> 0.2	_	$\in (0.3, 2.8)$	_
$CR2\ell$ -Top	> 400	—	< 0.05	> 0.5	—	$\in (0.3, 2.8)$	_
VR2ℓ-VV	> 400	> 250	< 0.05	$\in (0.4, 0.8)$	_	$\in (0.3, 2.8)$	_
VR2ℓ-Top	> 400	—	< 0.05	> 0.5	—	$\in (0.3, 2.8)$	_
$VR2\ell$ _High-Zjets	> 600	—	< 0.05	> 0.4	—	$\in (0.3, 2.8)$	_
$VR2\ell_Low-Zjets$	> 400	—	< 0.05	_	$\in (0.35, 0.60)$	_	_
SR2ℓ_High	> 800	_	< 0.05	> 0.8	_	$\in (0.3, 2.8)$	_
$\mathrm{SR}2\ell_{-}\mathrm{Int}$	> 600	_	< 0.05	> 0.8	_	$\in (0.6, 2.6)$	_
$SR2\ell_Low$	> 400	_	< 0.05	_	$\in (0.35, 0.60)$	_	> 2.4







Region	$n_{ m leptons}$	$N_{ m jet}^{ m ISR}$	$N_{ m jet}^{ m S}$	$n_{ m jets}$	$n_{b-\mathrm{tag}}$	$p_{\mathrm{T}}^{\ell_1,\ell_2}$ [GeV]	$p_{\mathrm{T}}^{j_{1},j_{2}}$ [GeV]
CR2ℓ_ISR-VV	$\in [3,4]$	≥ 1	≥ 2	> 2	= 0	> 25	> 30
$CR2\ell_ISR$ -Top	=2	≥ 1	$\stackrel{-}{=} 2$	$\in [3, 4]$	= 1	> 25	> 30
VR2ℓ_ISR-VV	$\in [3,4]$	≥ 1	≥ 2	≥ 3	= 0	> 25	> 20
$VR2\ell_{ISR}$ -Top	=2	≥ 1	= 2	$\in [3, 4]$	= 1	> 25	> 30
$VR2\ell_{ISR}$ -Zjets	= 2	≥ 1	≥ 1	$\in [3, 5]$	= 0	> 25	> 30
$SR2\ell_{-}ISR$	= 2	≥ 1	= 2	$\in [3,4]$	= 0	> 25	> 30

Region	$m_Z [{ m GeV}]$	$m_J [{ m GeV}]$	$\Delta \phi^{\rm CM}_{\rm ISR,I}$	$R_{\rm ISR}$	$p_{\rm T\ ISR}^{\rm CM}$ [GeV]	$p_{\mathrm{T~I}}^{\mathrm{CM}}$ [GeV]	$p_{\rm T}^{\rm CM}~[{\rm GeV}]$
CR2ℓ_ISR-VV	$\in (80, 100)$	> 20	> 2.0	$\in (0.0, 0.5)$	> 50	> 50	< 30
$CR2\ell_{ISR}$ -Top	$\in (50, 200)$	$\in (50, 200)$	> 2.8	$\in (0.4, 0.75)$	> 180	> 100	< 20
VR2ℓ_ISR-VV	$\in (20, 80)$	> 20	> 2.0	$\in (0.0, 1.0)$	> 70	> 70	< 30
	or > 100						
$VR2\ell$ _ISR-Top	$\in (50, 200)$	$\in (50, 200)$	> 2.8	$\in (0.4, 0.75)$	> 180	> 100	> 20
$VR2\ell_{ISR}$ -Zjets	$\in (80, 100)$	< 50 or > 110	—	_	> 180	> 100	< 20
$SR2\ell$ _ISR	$\in (80, 100)$	$\in (50, 110)$	> 2.8	$\in (0.4, 0.75)$	> 180	> 100	< 20





Overlap Plots



 $(RJR \cap CA) / (RJR \cup CA) [\%]$

 $(RJR \cap CA) / (RJR \cup CA) [\%]$













GAMBIT collaboration performed a global electroweak fit using available collider and direct DM constraints

		Best e	xpected SRs		All SRs; neglect correlations			
Analysis	Local signif. (σ)		EWMSSM fit (σ)	#SRs	Local signif. (σ)		$\begin{array}{c} \text{EWMSSM} \\ \text{fit} \ (\sigma) \end{array}$	#SRs
Higgs invisible width	0.9	0.3	0.2	1	0.9	0.3	0.2	1
Z invisible width	0	1.3	1.3	1	0	1.3	1.3	1
ATLAS_4b	0.7	0	0	1	2.1	0	0	2^*
ATLAS_4lep	2.3	2.0	0	1	2.5	1.0	0	4
ATLAS_MultiLep_2lep_	_0jet 0.9	0.3	0.1	1	1.3	0	0	6
ATLAS_MultiLep_2lep_	jet 0	0	0.5	1	0.8	0.5	0.3	3
ATLAS_MultiLep_3lep	1.8	1.6	0.6	1	1.2	0.4	0.3	11
ATLAS_RJ_2lep_2jet	0	0.3	0.5	1	1.5	1.8	1.5	4
ATLAS_RJ_3lep	2.8	2.4	1.0	1	3.5	2.6	0.5	4
CMS_1lep_2b	0.9	0.3	0.3	1	0	0	0	2
CMS_2lep_soft	0.4	0.2	0.2	12	0.4	0.2	0.2	12
CMS_20Slep	0	0.4	0.6	7	0	0.4	0.6	7
CMS_MultiLep_2SSlep	0.2	0	0	1	0.2	0	0	2
CMS_MultiLep_3lep	0	0	0.5	1	0	0	0	6
Combined	3.5	1.5	0.3	31	4.2	1.3	0	65





Our best-fit point has neutralino masses of $(m_{\tilde{\chi}_1^0}, m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_3^0}, m_{\tilde{\chi}_4^0}) \approx (49.4, 141.6, 270.3, 290.2) \text{ GeV}$, and chargino masses of $(m_{\tilde{\chi}_1^\pm}, m_{\tilde{\chi}_2^\pm}) \approx (142.1, 293.9) \text{ GeV}$. We find a local significance of 3.5σ for this excess. If there is indeed a supersymmetric signal resembling these properties the ATLAS and CMS experiments should be sensitive to it using the full LHC Run 2 dataset.

* GAMBIT: The Global and Modular Beyond-the-Standard-Model Inference Tool, Eur. Phys. J. C 77 (2017) 784, [arXiv:1705.07908].





Detector Performance Highlights







Trigger Performance Highlights



- Major challenge in 2016: Maintain trigger performance in fierce luminosity & pile-up conditions
- Main physics triggers for SUSY searches: Generic E_{T.miss}, jet, lepton triggers

Di-electron triggers

HLT_e17_lhvloose_nod0

100

120

Data

60

Z→ ee MC

80





1.4

1.2

0.8

0.6

0.4

0.2

0

20

40

ATLAS Internal

Data 2016, \sqrt{s} = 13 TeV, 33.5 fb⁻¹

Trigger Efficiency











Inclusive Olepton search - backgrounds

• Dominant backgrounds estimated in 4 CRs for each SR → extrapolation to VRs/SRs with transfer factors (TFs)







Example for gluino pair production with decays to jets and E_T^{miss}

Fit components estimate the total background. This is compared to the observed yield in the various signal region

Signal Region	RJR-G1a	RJR-G1b	RJR-G2a	RJR-G2b	RJR-G3a	RJR-G3b	RJR-G4					
	MC expected events											
Diboson	3.06	1.54	2.91	1.34	0.80	0.37	0.24					
Z/γ^*+ jets	28.56	13.03	28.01	9.41	8.56	2.90	2.05					
W+jets	13.99	6.40	14.66	4.98	4.45	1.71	0.99					
$t\bar{t}(+\mathrm{EW}) + \mathrm{single top}$	6.04	1.96	6.50	1.99	2.74	1.32	0.97					
		Fit	ted background	events								
Diboson	3.1 ± 0.6	1.5 ± 0.4	2.9 ± 0.8	1.34 ± 0.34	0.8 ± 0.24	0.37 ± 0.22	0.24 ± 0.13					
Z/γ^*+ jets	23.9 ± 3.0	10.9 ± 1.5	23.6 ± 2.8	7.9 ± 1.1	7.22 ± 1.0	2.5 ± 0.6	1.73 ± 0.33					
W+jets	11.4 ± 1.7	5.2 ± 0.8	11.7 ± 2.1	4.0 ± 0.7	3.5 ± 0.7	1.4 ± 0.6	0.79 ± 0.27					
$t\bar{t}(+\mathrm{EW}) + \mathrm{single top}$	4.8 ± 2.1	1.6 ± 1.1	5.6 ± 2.8	1.7 ± 1.0	2.4 ± 1.1	$1.14^{+1.20}_{-1.14}$	$0.83\substack{+1.19 \\ -0.83}$					
Michael Jeo	0.21 _ 0.21		0.0 ± 0.0	0.21 _ 0.21								
Total bkg	43 ± 4	19.2 ± 2.2	44 ± 4	15.2 ± 1.7	13.9 ± 1.6	5.3 ± 1.4	3.6 ± 1.3					
Observed	38	16	48	15	19	11	6					
$\sqrt{\frac{95}{\text{obs}}}$	0.00	0.20	0.00	0.20	0110	0.00	0.20					
$S_{\rm obs}^{95}$	13.9	9.4	20.1	10.0	14.5	13.6	9.1					
$S_{exp}^{\overline{95}}$	$16.2^{+6.6}_{-4.9}$	$10.7^{+4.1}_{-2.8}$	$17.4^{+5.4}_{-5.5}$	$9.5^{+4.2}_{-2.4}$	$9.7^{+4.0}_{-2.1}$	$10.2^{+3.8}_{-1.4}$	$7.6^{+2.3}_{-1.7}$					
p_0 (Z)	$0.50 \ (0.00)$	$0.50 \ (0.00)$	$0.23 \ (0.73)$	$0.50 \ (0.00)$	$0.09 \ (1.36)$	0.12(1.15)	0.18 (0.90)					





Signal region – Olepton search results!



q

 \tilde{Y}_1^0





 \tilde{q}







Good improvement on the parameter space we're now probing wrt run1 ©

Still no clear signal $\ensuremath{\mathfrak{S}}$



Phys. Rev. D 97, 112001 (2018)





Signal region – Olepton results!









RJR technique – used in ATLAS



1400

m_ã [GeV]



Yields in agreement with expectation, for the most part....some modest excess at the highest mass

Also has excellent performance in signal models that the analysis wasn't optimised to probe.









- Between the two analysis approaches we select non-overlapping events

- If we were to see an excess in both, or one and not the other, we get a very powerful piece of information instantly





