Supersymmetry: Status 2018/19

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Supersymmetry addresses the following shortcomings of the SM:

- The Hierarchy Problem
- What is Dark Matter?
- Unification of gauge couplings (quarks and leptons fill complete SU(5) representations, but gauge couplings do not quite unify in the Standard Model without Supersymmetry)
- Possibly: the  $\sim$  3  $\sigma$  deviation of the measured anomalous magnetic moment of the muon  $a_{\mu}$  w.r.t. the Standard Model

### BUT: Up to now, no significant sign for Supersymmetry at the LHC

### Status of the MSSM after 36 $fb^{-1}$ at the LHC

Better not rely on interpretations of limits within simplified models (simplified decay cascades, typically 1 step), but

- consider versions of the "pMSSM" (differ by the number of independent parameters/soft Supersymmetry breaking terms at the weak scale)
- consider all (most relevant) searches by ATLAS and CMS simultaneously, interprete with realistic branching ratios and decay cascades within these models
- require the Higgs  $(H_{125})$  mass and couplings within present limits
  - $\rightarrow$  Need a large soft Supersymmetry breaking  $A_{top}$  and/or large soft Supersymmetry breaking stop mass terms
  - ightarrow large rad. corrections to a soft Supersymmetry breaking Higgs mass term
  - $\rightarrow$  a "little" finetuning problem
- Finally: derive lower limits on sparticle masses at 95% confidence level

### The Dark Matter Issue

Case of generic annihilation cross sections of higgsinos/winos/binos as LSP  $\chi_1^0$ :

• Higgsinos and winos annihilate too fast

 $\rightarrow$  Relic density too small (unless  $M_{\chi_1^0} \gtrsim 1 \text{ TeV} \rightarrow M_{Squark, Slepton} > 1 \text{ TeV}$ )

• Binos annihilate too slow  $ightarrow \Omega h^2$  too large

 $\rightarrow$  Binos as LSP/DM require enhanced annihilation cross sections:

- Large bino/wino/higgsino mixing requires tuning given the relatively small off-diagonal elements of the 4  $\times$  4 neutralino mass matrix
- coannihilation with charginos/sleptons/stops ( $\rightarrow M_{\chi_1^0} > 100$  GeV since near-degeneracy is required), or
- annihilation via  $H_{125}$  or heavy H/A-funnels

 $\rightarrow$  Higgsinos/winos as LSP lighter than 1 TeV require other sources for DM (their lower relic densities would allow to alleviate constraints from direct DM detection which are strong)

Or: unconventional cosmological evolution

### Recent pMSSM Scans

## GAMBIT collaboration (1705.07917), pMSSM7:

- Gaugino mass ratios motivated by GUT:  $M_3/lpha_s=M_2/lpha_2=M_1/lpha_1$
- Degenerate soft squark/slepton masses, but free  $A_t \neq A_b$ (the muon anomalous magnetic moment  $a_\mu$  cannot be fitted)
- Free soft Higgs masses  $M_{H_u}$ ,  $M_{H_d}$ ,  $\tan \beta \ ( \rightarrow \mu, B_\mu \text{ fixed by } M_Z, \tan \beta)$
- $\Omega h^2 \lesssim 0.1189$  (smaller  $\Omega h^2$  alleviates constraints from direct DM detection, but requires additional sources of dark matter)

## MasterCode (1710.11091), pMSSM11:

- Free gaugino masses
- Different soft squark/slepton masses for the first two/third generations, free  $A_t = A_b$
- Free  $\mu$ , tan  $\beta$ ,  $M_A$
- $\Omega h^2 = 0.1186 \pm 0.004$
- With or without fits to the muon anomalous magnetic moment  $a_{\mu}$

#### Limits on sparticle masses (Within $2\sigma$ of the "best fit point" to numerous search signal regions mainly from sparticle searches by ATLAS/CMS)

MasterCode: The parameters of the "best fit points" with or w/o  $a_{\mu}$  are completely different!

|                    | GAMBIT                   | MasterCode with $a_{\mu}$         | MasterCode w/o $a_{\mu}$ |
|--------------------|--------------------------|-----------------------------------|--------------------------|
| $M_{\chi_1^0}$     | $\gtrsim$ 60 GeV (H'ino) | 90 – 500 GeV (bino)               | > 90 GeV (H'ino)         |
| $M_{\chi_1^{\pm}}$ | $\gtrsim$ 90 GeV         | $\gtrsim$ 90 GeV                  | $\gtrsim$ 90 GeV         |
| $M_{\tilde{g}}$    | $\gtrsim 1,0$ TeV        | $\gtrsim 1,8$ TeV                 | $\gtrsim 1,0$ TeV        |
| $M_{\tilde{q}}$    | $\gtrsim 1,2$ TeV        | $\gtrsim 1,9$ TeV                 | $\gtrsim$ 800 GeV        |
| M <sub>ĩ</sub>     | $\gtrsim$ 0, 5 TeV       | $\sim$ 500 GeV/ $\gtrsim$ 1,0 TeV | $\gtrsim 500$ GeV        |
| $M_{	ilde{	au}}$   | $\gtrsim 1,3$ TeV        | $\gtrsim 110{ m GeV}$             | $\gtrsim 110$ GeV        |
| $M_{	ilde{\mu}}$   | $\gtrsim 1,3$ TeV        | 110 – 770 GeV                     | $\gtrsim 110$ GeV        |
| M <sub>A</sub>     | $\gtrsim 500$ GeV        | $\gtrsim 800 \; { m GeV}$         | $\gtrsim 800~{ m GeV}$   |

No sign for "dark spots" in the combined signal regions (light sparticles escaping detection)

→ Limits on squarks/gluinos depend strongly on assumptions (possible decay cascades)

## NMSSM

- Extra neutral CP-even and CP-odd scalars  $H_S$ ,  $A_S$  (not degenerate!) on top of the MSSM-like heavy ~degenerate SU(2) doublets H/A
- Extra singlino  $\tilde{S}$ on top of the MSSM-like charged/neutral bino/wino/higgsinos
- $H_S$ ,  $A_S$ ,  $\tilde{S}$  have small couplings to SM particles/MSSM sparticles, except to the Higgs sector from a coupling  $\lambda \tilde{H}_u \tilde{H}_d \tilde{H}_S$  in the superpotential (in terms of superfields)
- $\rightarrow$  Small direct production cross sections proportional to mixing angles<sup>2</sup> ~  $\lambda^2$ , but possible decay products of SM particles/MSSM sparticles
- $\rightarrow$  Still:  $H_S$ ,  $A_S$  decay into SM particles like  $H_{125}$  due to mixing

 $\rightarrow$  Searches for  $ggF \rightarrow H_S \rightarrow \gamma\gamma$  with  $M_{H_s} < 125$  GeV are sensitive to viable cross sections  $\times$  BR in the NMSSM:



From CMS-HIG-17-013 (13 TeV)



Possible Xsect  $\times$  BR in the NMSSM for 13 TeV using limits from 8 TeV, from 1512.04281

Searches for  $H_{125} \rightarrow A_5 A_5 / H_5 H_5$  are sensitive to BSM branching fractions of  $H_{125}$  allowed by indirect constraints from measured  $H_{125}$  couplings

Many possible final states, many recent and ongoing searches by ATLAS/CMS



NMSSM after 2017 LEP/LHC constraints

Significant improvement in the  $\mu\mu bb$ channel!

# If the singlino $\tilde{S}$ is the LSP (I)

A good DM candidate: a relic density  $\Omega h^2 \sim 0.119$  is possible even if  $\tilde{S}$  is very light (a few GeV) through annihilation via  $A_S$  funnel ( $\neq$  MSSM)

Constraints from direct detection experiments:

coloured region: NMSSM points allowed by constraints from LUX/PandaX-II  $\rightarrow$  Xsect possibly below the neutrino floor (black curve):



(from 1806.09478)

# If the singlino $\tilde{S}$ is the LSP (II)

Every NLSP (neutralino, chargino, slepton, stop...) will decay into  $\tilde{S} + H_{125}/H_S/A_S/Z, W, lepton, top...,$ the only available decay channels due to R-parity conservation



Notably if  $\tilde{S}$  is light (a few GeV) AND  $M_{NLSP} \approx M_{\tilde{S}} + M_X$ ,  $X = H_{125}/H_S/A_S/Z$ : Little energy is given to  $\tilde{S}$  in any decay  $NLSP_{heavy} \rightarrow X_{heavy} + \tilde{S}_{light}$  $\rightarrow$  Little  $E_T^{miss}$  from  $\tilde{S}$  in all Susy searches

 $\rightarrow$  Reduced lower limits on sparticle masses (A.Teixeira, U.E., 1406.7221, 1412.6394)

Re-analysis of sensitivities of the CMS squark search via jets and  $E_T^{miss}$  (1802.02110, after 36 fb<sup>-1</sup> at 13 TeV) by A. Titterton et al., 1807.10672:

MSSM with bino LSP Assume  $\tilde{q} \rightarrow q + bino$ 

NMSSM with  $\tilde{q} \rightarrow q + bino \rightarrow q + \tilde{S} + H_{125}$  $M_{bino} = M_{\tilde{S}} + M_{H_{125}} + 2 \text{ GeV}$ 



(red/black curves: expected/observed limits)

NMSSM: strong reduction of the lower limit on  $M_{Squark}$  for small  $M_{LSP} = M_{\tilde{S}}!$ 

Given the present absence of significant excesses:

What are the constraints on the NMSSM parameter space (masses and couplings)?

A simple but highly nontrivial question given the larger parameter space, and the possibilities of "dark spots": reduced sensitivities in case of quenched spectra, complicated decay cascades, ...

"Bottom up" approach: Start with constraints on neutralinos/charginos, the lower ends of heavier sparticle decay cascades:

Neutralino/chargino masses and couplings are needed for their simulations!

Require – as promised by Supersymmetry – a viable DM relic density consistent with constraints from direct DM detection

At the LHC, neutralinos/charginos are produced via  $W^{\pm *} \rightarrow \chi_i^0 + \chi_j^{\pm}$ (or  $Z^* \rightarrow \chi_i^{\pm} + \chi_j^{\mp}, \chi_i^0 + \chi_j^0$ )

Searches by ATLAS/CMS: Trileptons from Z + W (or *bb* from  $H_{125}$ ):



Results are typically interpreted for wino-like  $\chi_2^0 + \chi_1^{\pm}$ : Largest cross sections  $\rightarrow$  strongest constraints

But: Higgsinos have only half the cross section (even adding  $\chi^0_2$ ,  $\chi^0_3$ )  $\rightarrow$  weaker constraints

For limits on the NMSSM singlino-higgsino sector (with C. Hugonie, 1806.09478): Scan the parameter space with singlino LSP, require a viable relic density consistent with constraints from direct DM detection, apply bounds from the CMS trilepton search in 1801.03957 (the strongest ones) Comparison of limits the in the  $M_{\chi_1^0}/M_{\chi_1^\pm} \sim M_{\chi_2^0}$  plane:

CMS, assuming wino-like  $\chi_2^0$  and  $\chi_1^{\pm}$ : NMSSM, singlino LSP and higgsino-like



Blue: Excluded iff the bino mass satisfies  $M_1 > 300$  GeV as motivated by the GUT relation  $M_1 \approx M_{Gluino}/6$  and  $M_{Gluino} \gtrsim 1.8$  TeV  $\rightarrow$  no bino/higgsino mixing

 $\rightarrow$  Substantial reduction of limits!

Allowed regions in the plane  $M_{\tilde{\chi}_1^{\pm}} - M_{\tilde{\chi}_1^{0}}$  in the constrained NMSSM: universal soft susy breaking terms at the GUT scale, but non-universal soft Higgs mass terms (allows to estimate the necessary amount of finetuning):



→ Relatively low finetuning for  $M_{\tilde{\chi}_1^0} \sim M_Z/2$ ,  $M_{\tilde{\chi}_1^0} \sim M_{H125}/2$  or  $M_{\tilde{\chi}_1^0} \sim M_{\tilde{\chi}_1^\pm}$ where s-channel annihilation or co-annihilation is possible Otherwise: s-channel annihilation via  $A_S$  with  $M_{\tilde{\chi}_1^0} \sim M_{A_S}/2$ → Many regions with relatively low fine-tuning  $\approx 100$  remain to be tested

### Dark Spots for neutralino/chargino searches:

- Mixed bino higgsino NLSP  $\chi_2^0$ : reduces production cross section further
- $\chi^0_{2,3}$  cascade decays via light  $H_S$  or  $A_S$  (escape searches for  $H_{125}$  via  $b\bar{b}$ )
- Light staus  $\tilde{\tau}$  as NLSP: Hardly constrained by the LHC (limits from LEP),  $\rightarrow$  less "Trileptons" in the final state

### Hints for Excesses?

Recent searches for neutralinos/charginos by ATLAS/CMS:

Some local  $2-3\sigma$  excesses in some bins/some signal regions (e.g. ATLAS 1806.02293 in SR2 $\ell_{low/ISR}$  and SR3 $\ell_{low/ISR}$ )  $\rightarrow$  If interpreted in terms of simplified models: No significant deviations from observed w.r.t. expected limits

#### GAMBIT collaboration, 1809.02097:

combine 4 ATLAS and 4 CMS electroweakino searches after 39 fb<sup>-1</sup> ( $\approx$  10 signal regions each, up to  $\sim$  40 bins), compare to simulations within a pMSSM electroweakino sector (bino, wino, higgsinos), allowing for cascade decays  $\rightarrow$  local 3,2  $\sigma$  excess for  $M_{\chi_1^0} \sim 50$  GeV,  $M_{\chi_1^\pm} \sim 150$  GeV via contributions from  $\chi_2^0$ ,  $\chi_3^0$  and  $\chi_2^\pm$  multi-W/Z cascade decays ( $M_{\chi_1^0} \sim 8 - 155$  GeV,  $M_{\chi_1^\pm} \sim 104 - 259$  GeV within 95% CL)

But: Missing covariance matrices for stat. analysis including more search results

Viable relic density possible for bino-like LSP via  $Z/H_{125}$  funnels

#### $\rightarrow$ To confirm or to rule out!

# Conclusions:

- Due to  $M_{Higgs} \sim 125$  GeV the MSSM has a "little" finetuning problem of at least  $\mathcal{O}(1\%)$ , of  $\mathcal{O}(1\%)$  with (grand) unified soft Supersymmetry breaking terms, somewhat less in the NMSSM
- Of course: even with  $M_{Squark}$ ,  $M_{Gluino} > 1 2$  TeV Supersymmetry still solves the "BIG" hierarchy problem
- To derive definite constraints on the high dimensional parameter space is a challenging task, notably in the NMSSM ( $\rightarrow$  dark spots)
- But a MUST for the future unless significant excesses appear

