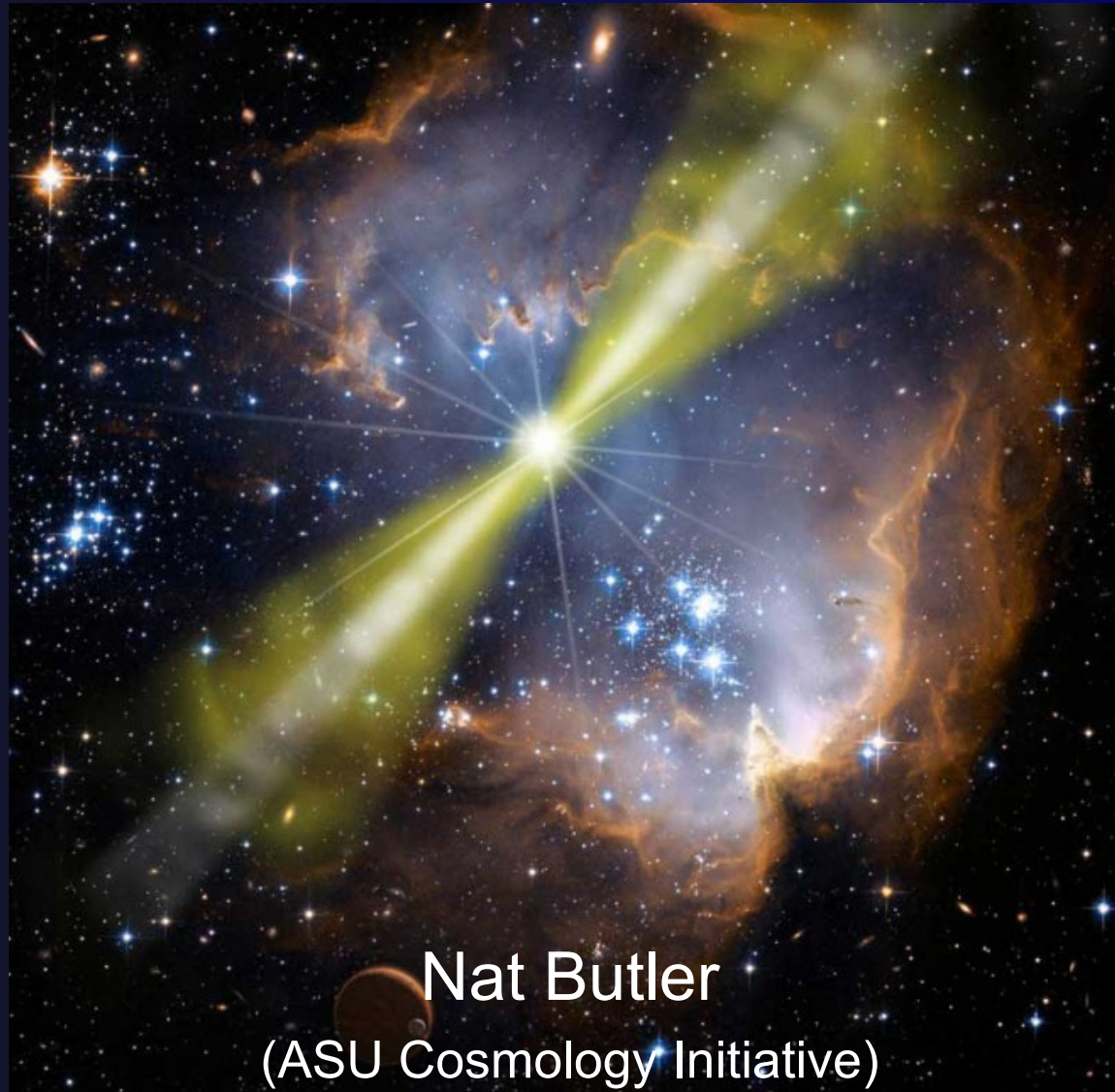


Chasing Astrophysical Transients

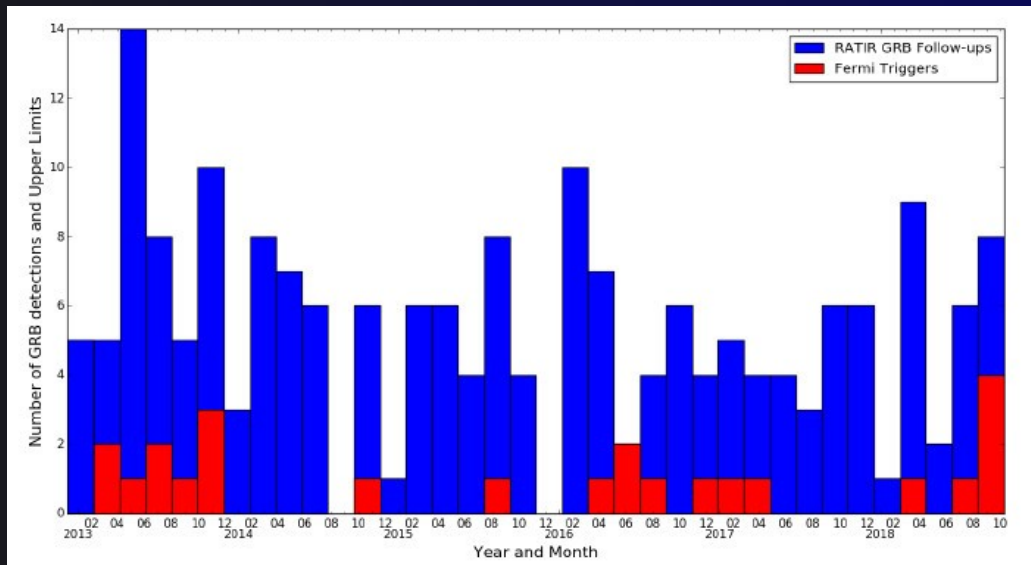


Nat Butler
(ASU Cosmology Initiative)

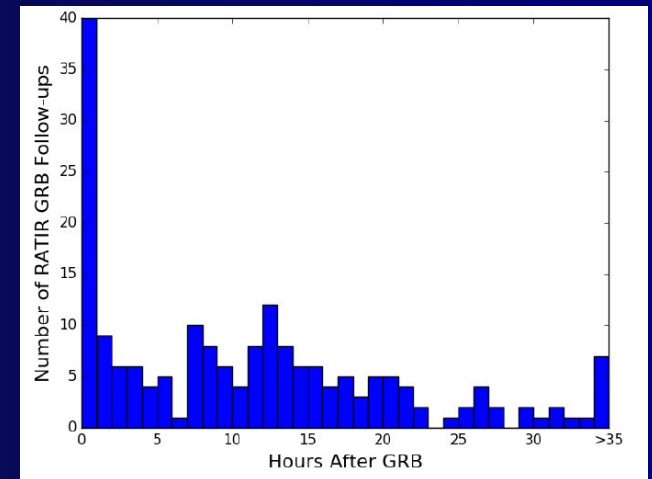
Outline

Lesson from Swift & RATIR for SVOM & Colibri

* High-z GRBs * Jetting constraints * GW followup



Past ~6 years (2.7 GRBs/mo)



THE REIONIZATION AND TRANSIENTS
INFRARED/OPTICAL PROJECT

RATIR



<http://www.ratir.org>

Nathaniel Butler

Chris Klein (w/ Josh Bloom)

Alejandro Farah

Jesus Gonzalez

Alexander Kuttyrev (w/ Harvey Moseley, Neil Gehrels)

Ori Fox

Dave Rapchun

Alan Watson (w/ Michael Richer)

(PI & Project Manager; ASU)

(Software Lead; UC Berkeley)

(Mechanical Design; UNAM)

(Optical Design; UNAM)

(IR Dewar Lead; Goddard)

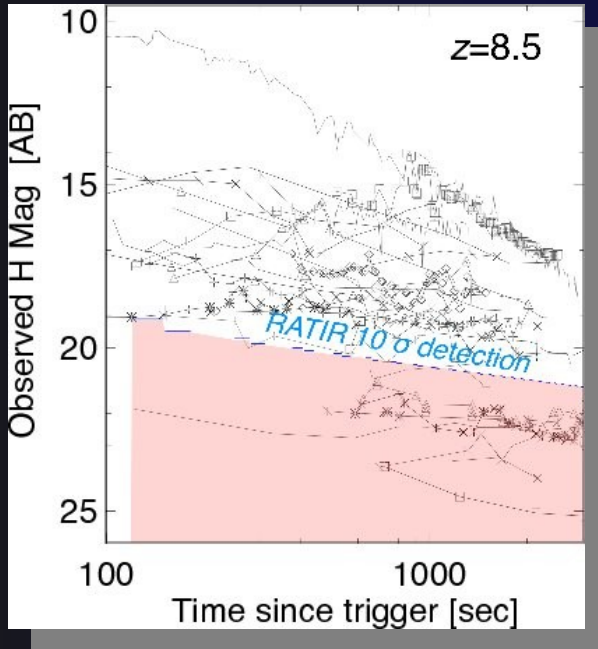
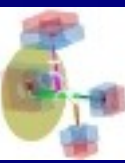
(NIR Detectors, Goddard)

(NIR Dewar, Goddard)

(Telescope Automation; UNAM)

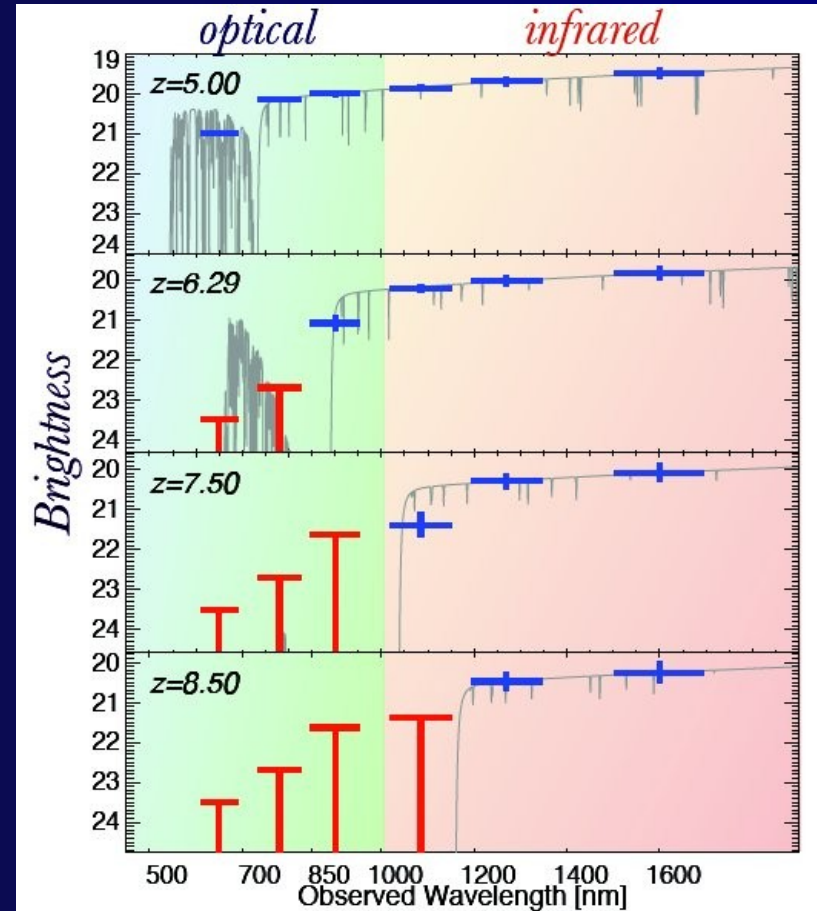
William Lee, David Clark, Carlos Roman, Leonid Georgiev (UNAM), J. Xavier Prochaska, Enrico Ramirez-Ruiz (UCSC/Lick), Gennadiy Lotkin, Frederick Robinson, Matthew Samuel, Leroy Sparr (Goddard), ...

RATIR GRB Sensitivity

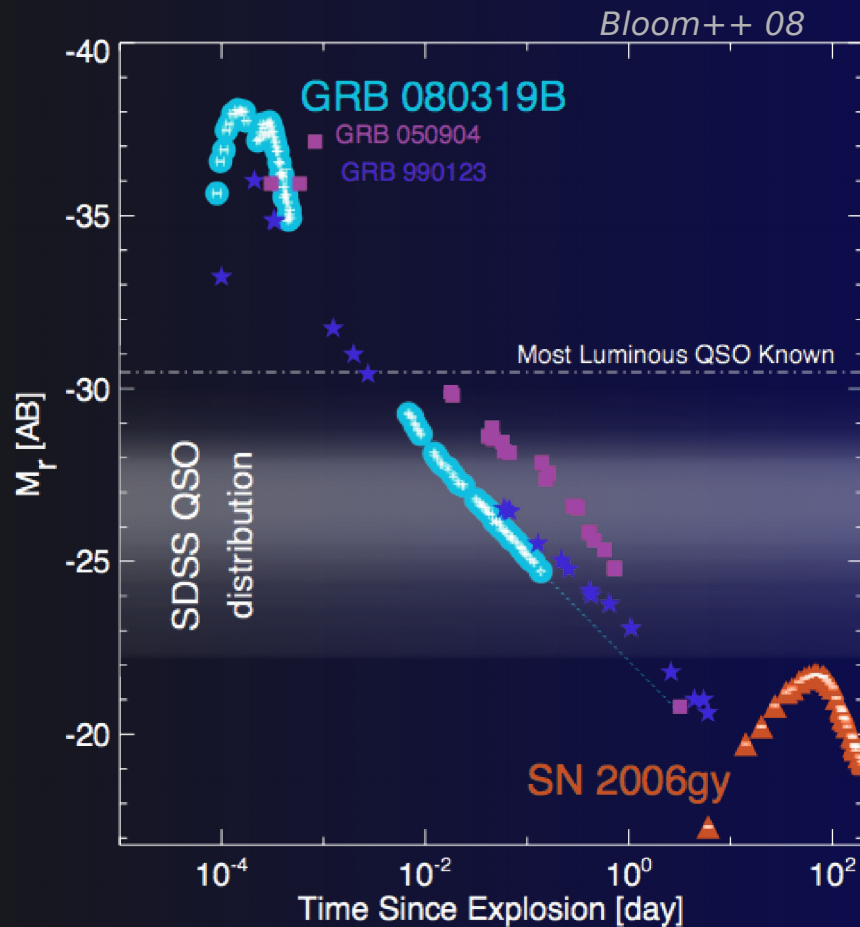


Detects 83% of Swift events,
even at $z=8.5$

Will observe 12 Swift events
per year within 10 minutes
(actual 9/yr within 1 hr).



The Promise of GRBs: Extremely Distant & Piercing



1000x brighter than
brightest QSO

Median (Swift) $z=2$

10% $z>4$ (Detectable $z>10$)

GRBs as Cosmological Lighthouses



I. DIRECT

Standard Candles?

High-Energy Observables
+ Redshift
→ Cosmology

II. INDIRECT

Backlights at Edge of Universe

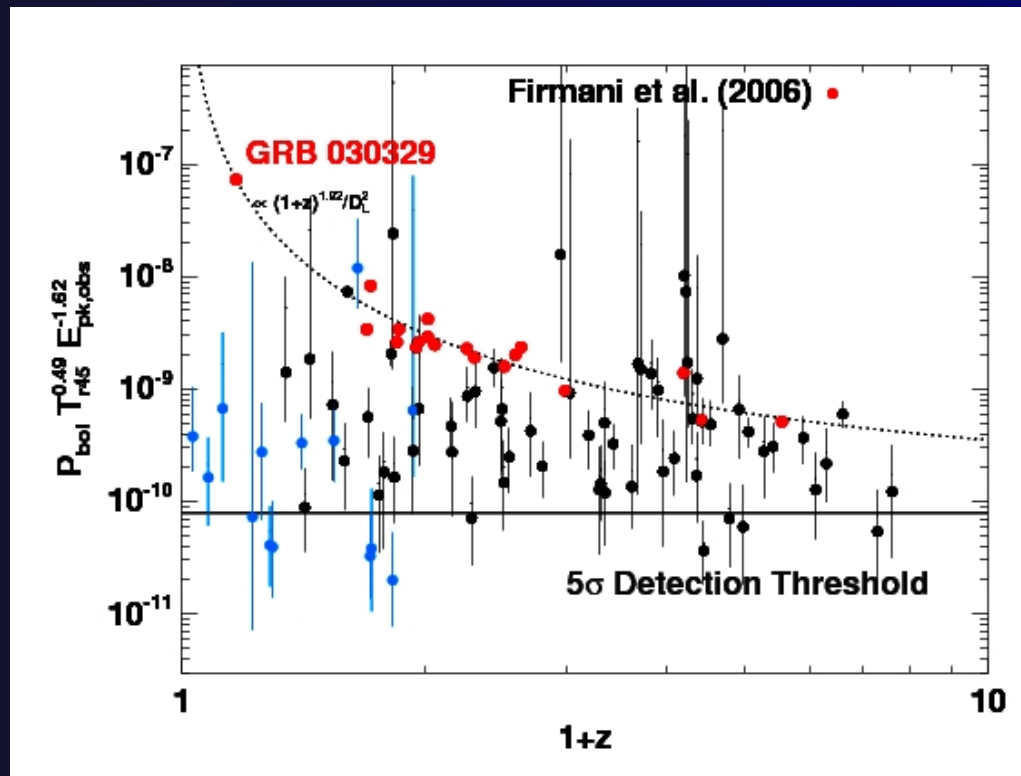
Optical/IR/X-ray Absorption
→ Gas/Dust in Distant Galaxies, Inter-Galactic
Medium

APOD 11/26/07

Standard Candles?

Essentially **3 Observables**: Duration, Flux, Hardness

Assume Cosmology, Observables \rightarrow Distance (or z)



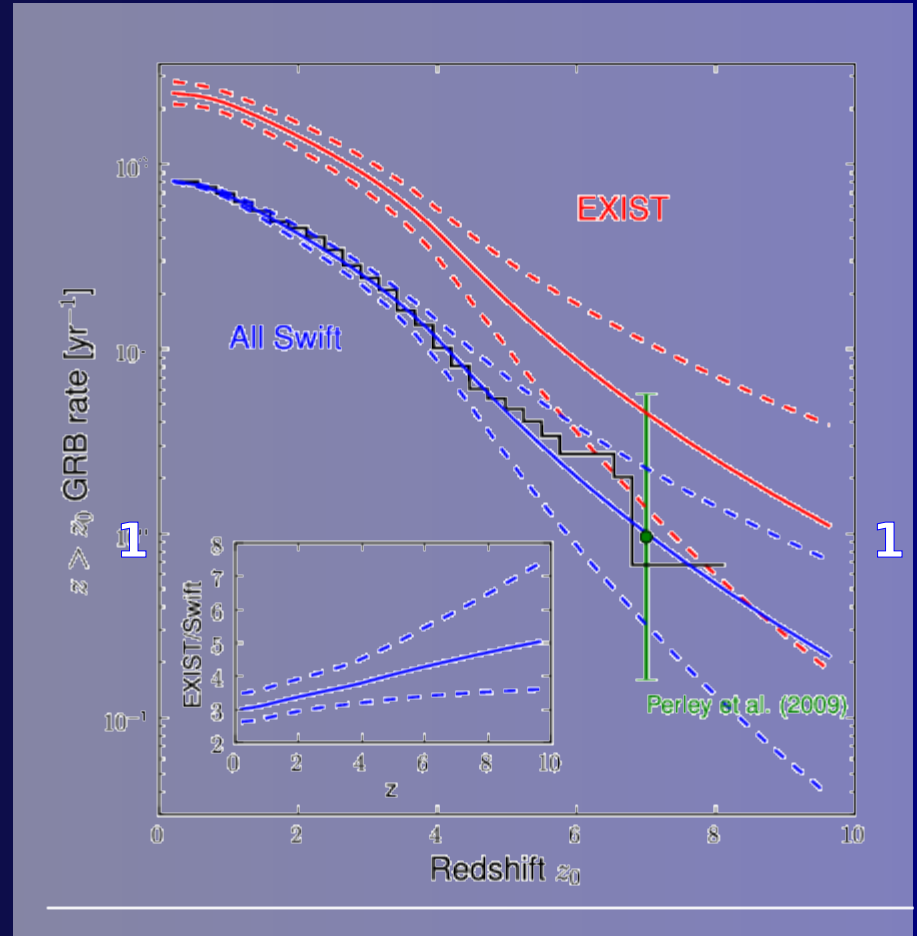
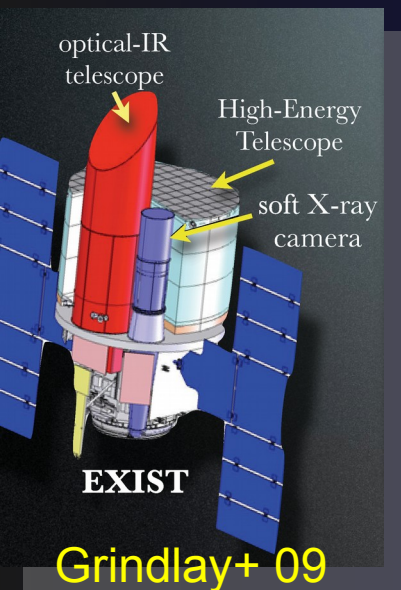
Pernicious Statistics! (*Butler+ 07,09,10*)

Fraction at High-z

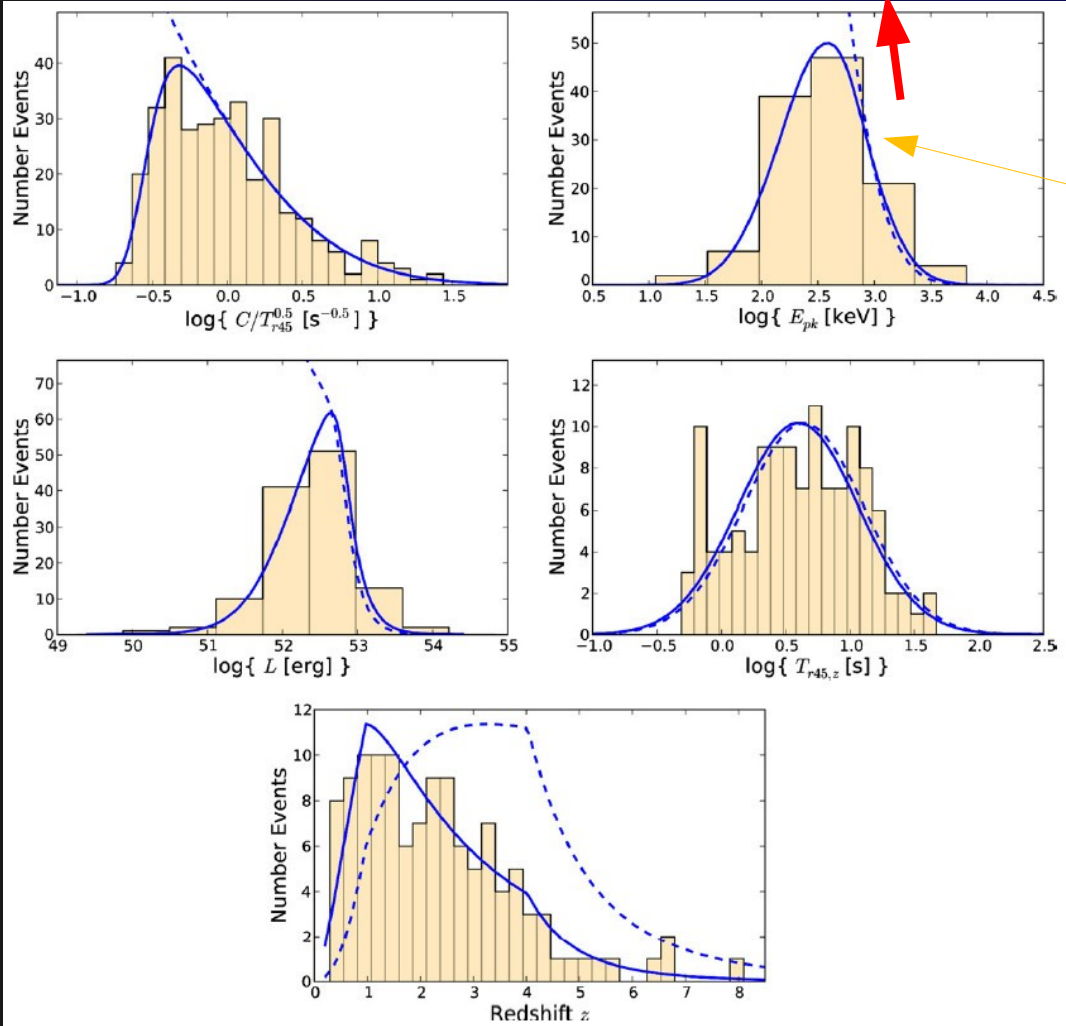
5 ± 3 Swift GRBs/yr @ $z > 5$

1 ± 1 Swift GRB/yr @ $z > 7$

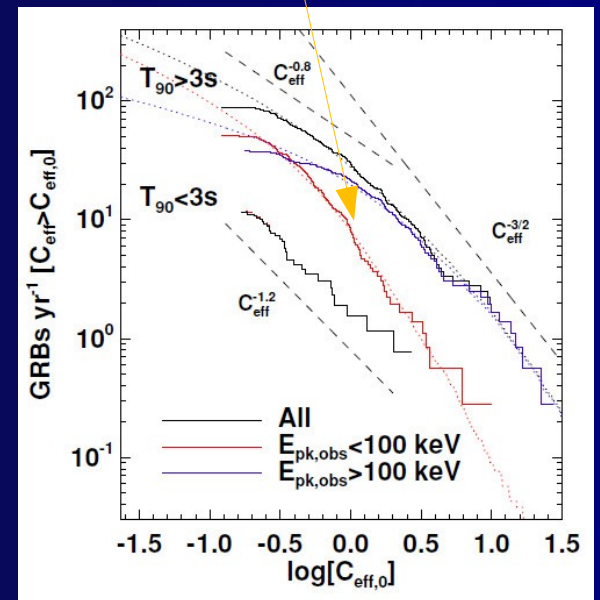
(see, also, Perley...Butler+ 09,
Fynbo+ 09)



Eclairs Soft Response → More GRBs

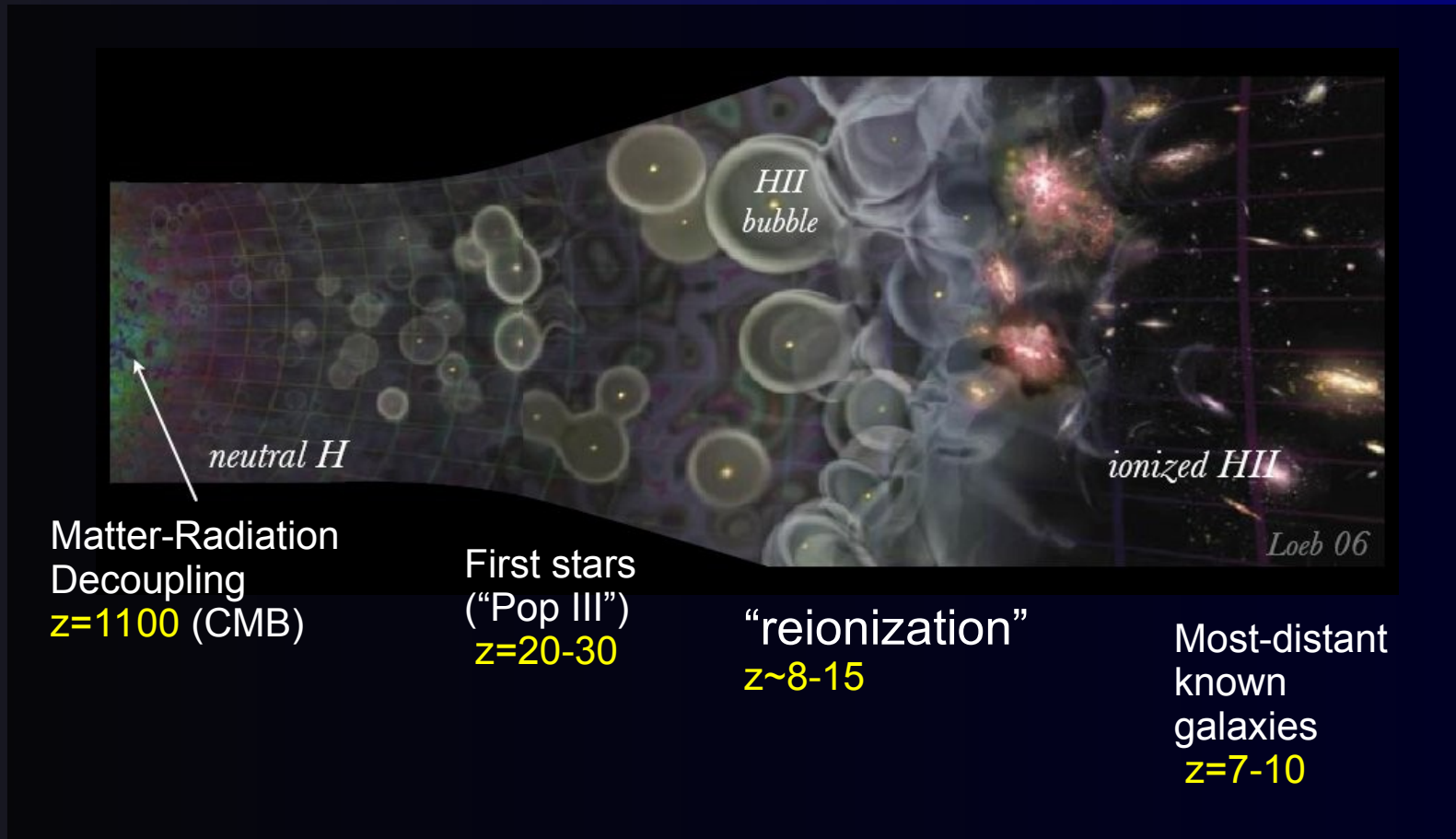


Soft GRB Population!

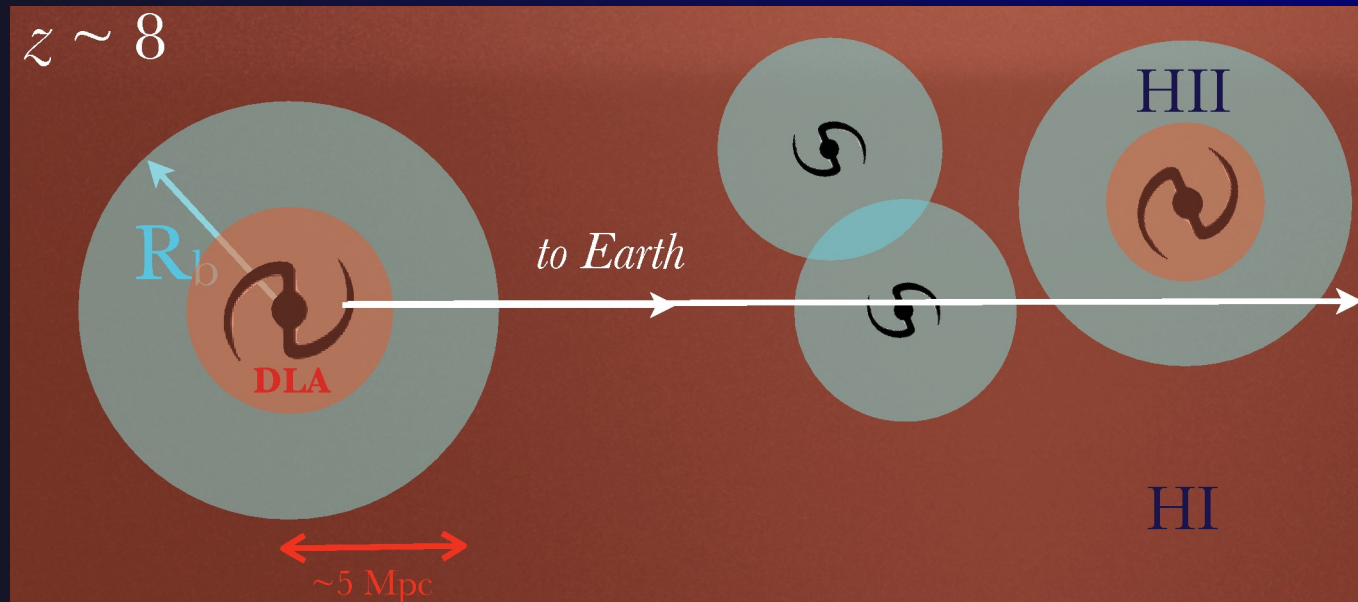


SWIFT/BAT Rates, Butler et al. 2010

Detect High-z GRBs → Epoch of Reionization



Epoch of Reionization



Directly Infer neutral fraction of Hydrogen $x_{\text{H I}}$ versus redshift
(Miralda-Escudé 98, McQuinn+ 08,
Mesinger & Furlanetto 08)

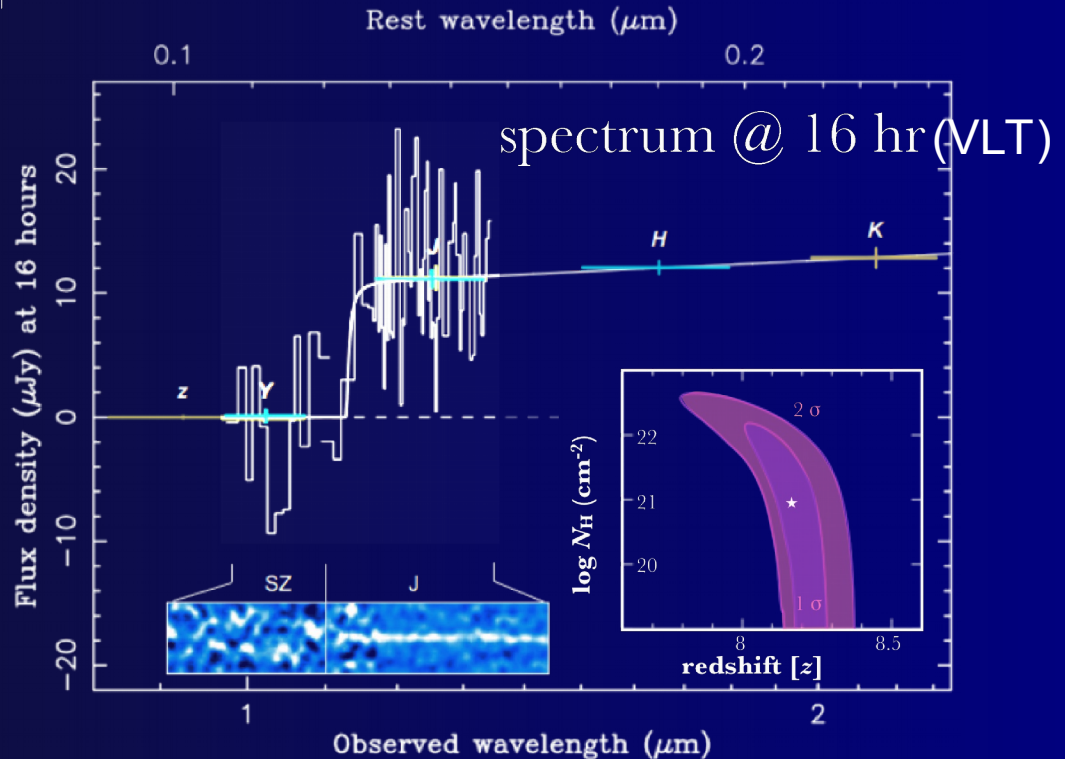
GRB 090423 ($z \approx 8.3$)

Large $N_{\text{H I}}$ obscures ionization state!

(*Tanvir+ 09, also Salvaterra+ 09*)

But 10-20% of GRBs still have sufficiently low $N_{\text{H I}}$.

(*Chen+ 08*)



RATIR Constraints on Jetting

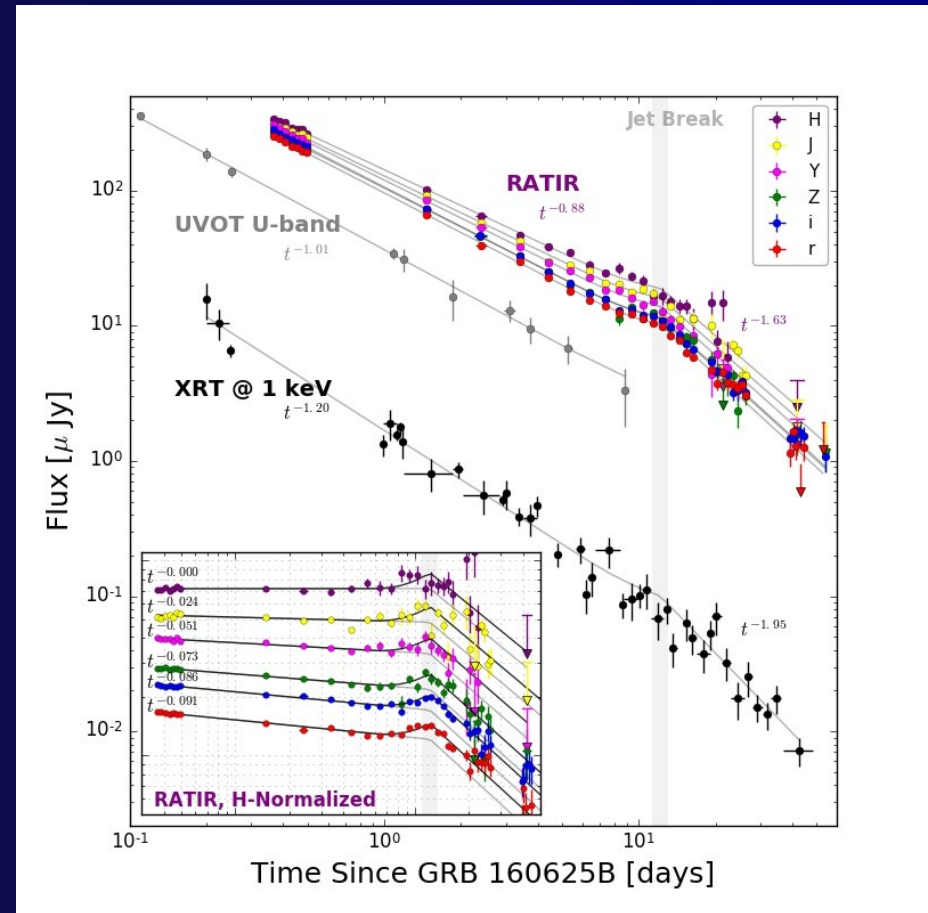
GRB 160625B (Troja et al. 2018; Nature):

Polarization detection by Master

Late-time, extraordinary high-cadence jet break

→ best observations to date of a jet break

→ the jet is very narrow (few degrees) and may have edge structure



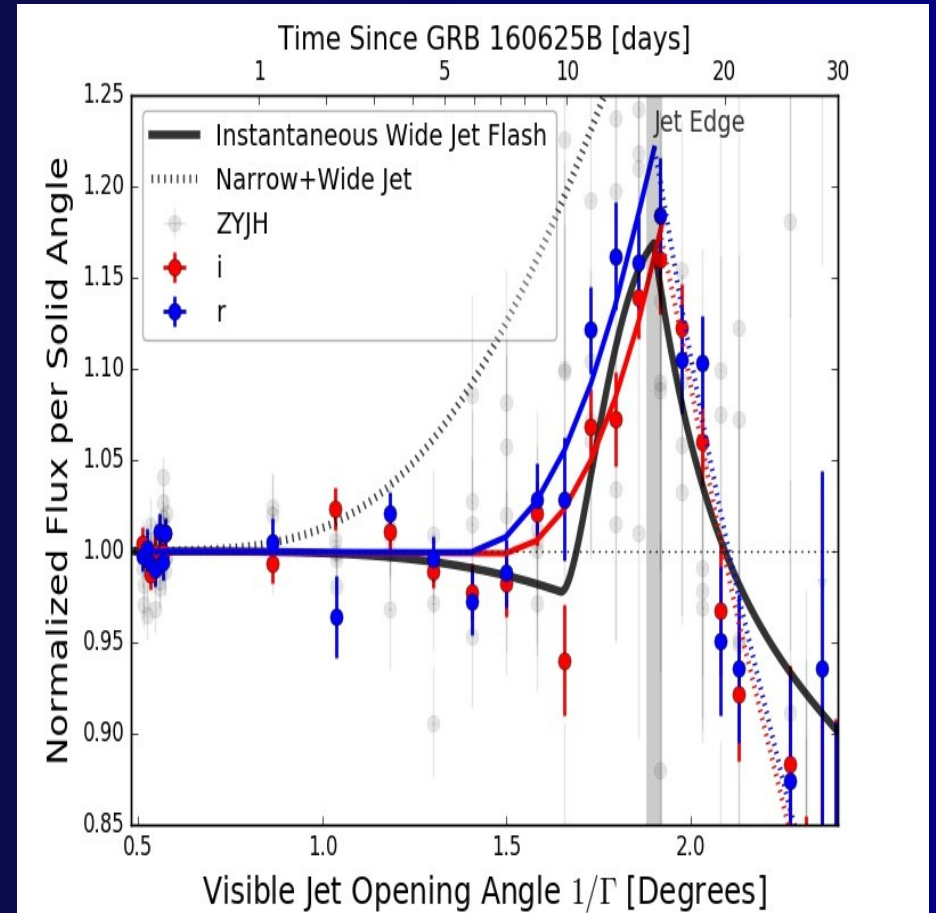
Strausbaugh et al. 2018

RATIR Constraints on Jetting

GRB 160625B (Troja et al. 2018; Nature):

Zoom-in lightcurve bump:

→ possible edge structure inferred from bump at time of “jet break”



Strausbaugh et al. 2018

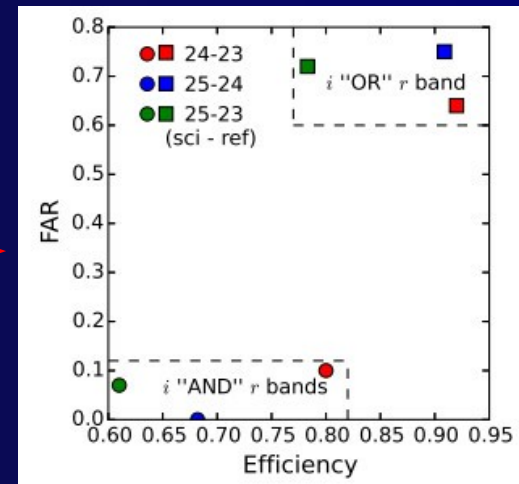
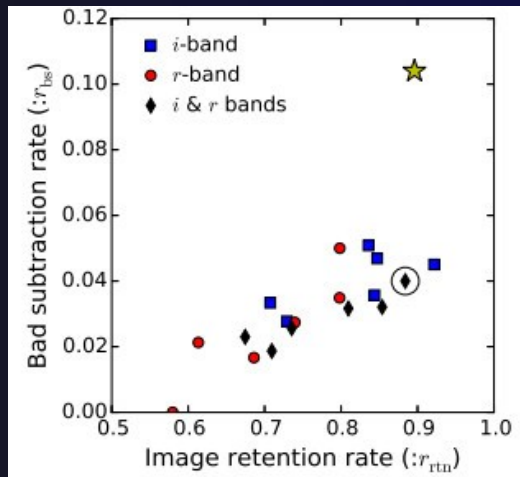
$$\theta_{\text{jet}} = 3.27 \left(\frac{t_{\text{jet}}}{\text{days}} \right)^{3/8} \left(\frac{1+z}{2} \right)^{-3/8} \left(\frac{E_{\text{iso}}}{10^{53} \text{ erg}} \right)^{-1/8} \left(\frac{\eta}{0.2} \right)^{1/8} \left(\frac{n}{0.1 \text{ cm}^{-3}} \right)^{1/8} = 2.28 \left(\frac{t_{\text{jet}}}{12.6 \text{ days}} \right)^{3/8} \text{ degrees}$$

GW Followup with RATIR

“Reionization And Transients InfraRed camera”

GW Followup from 2015+ described in Golkhou et al. 2018.

We target the known galaxies and do image subtraction.



Consider seeing, depth of image pairs, avoid bad subtractions.

Use multiple bands to drive down FAR.