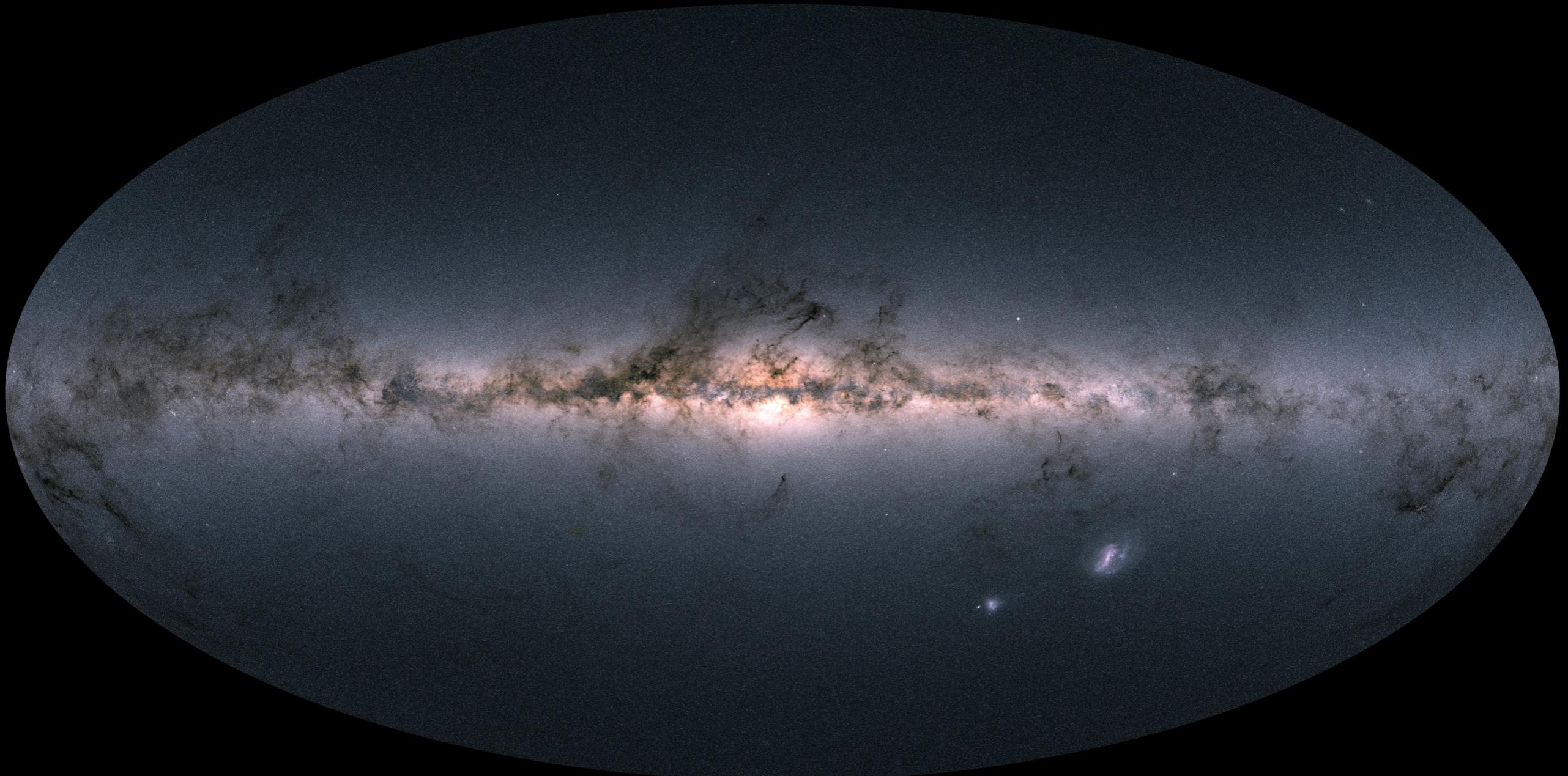




gaia

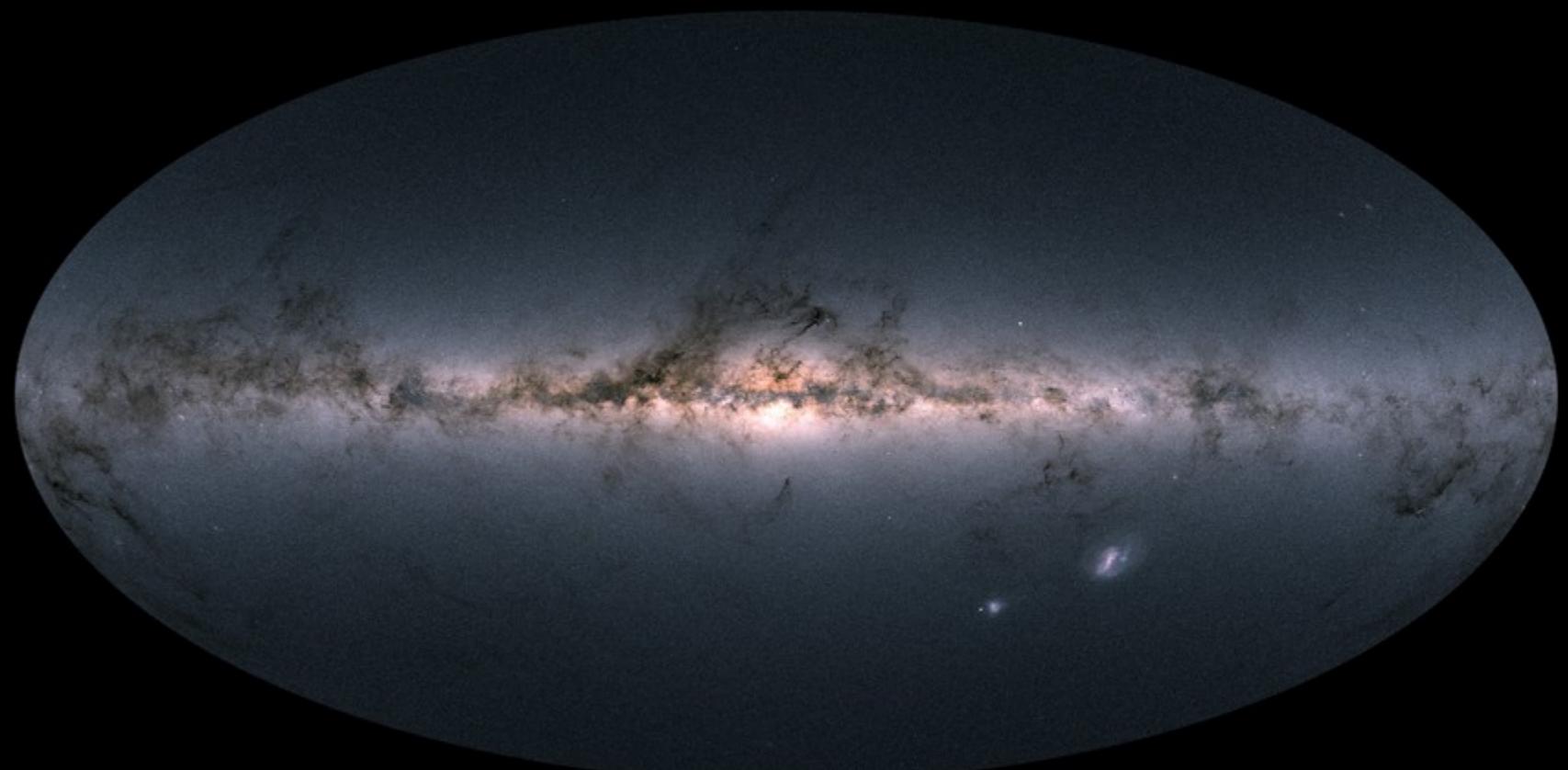
# Gaia DR2



Carine Babusiaux  
IPAG / OBSPM

# The second Gaia data release

- Gaia mission overview
- Gaia DR2 content and limitations
- A few (biaised!) applications



# The Gaia mission

ESA cornerstone mission

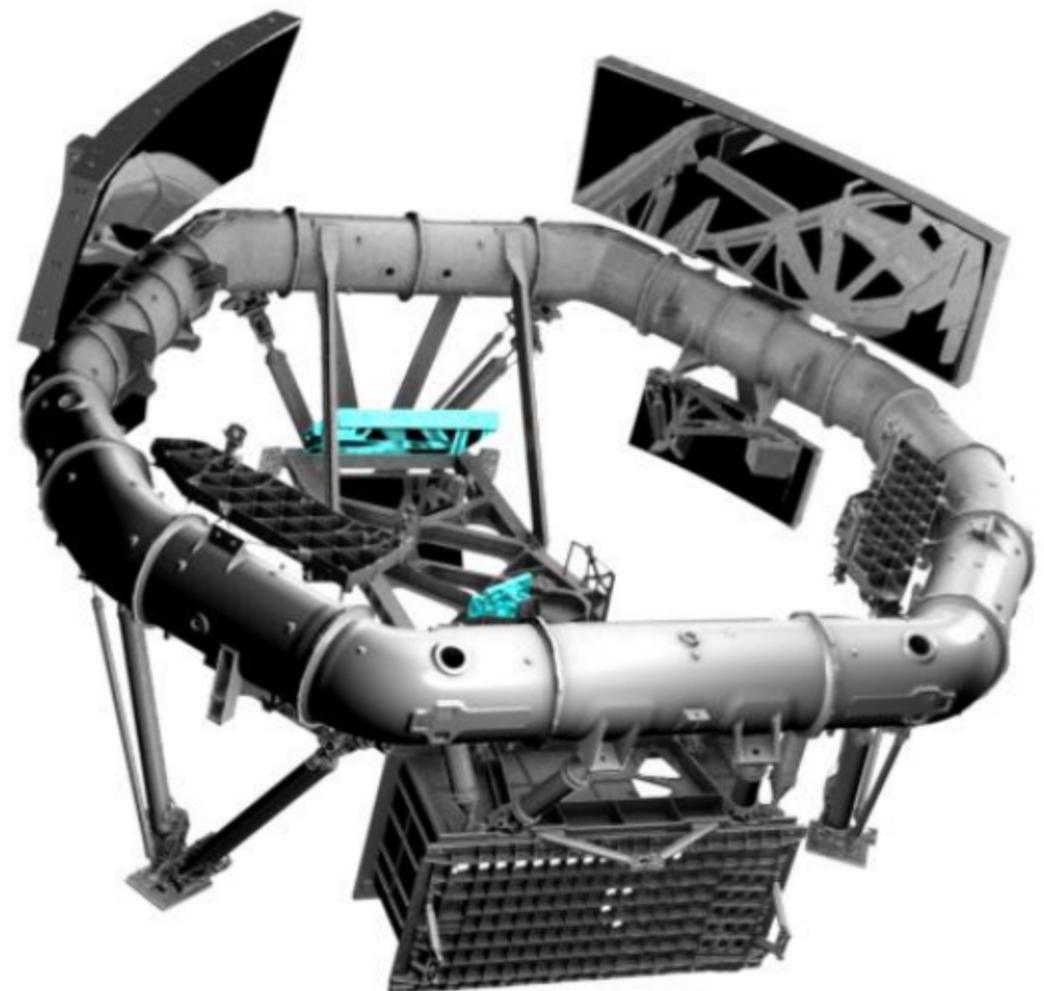
5 years (+) of mission

## 3 instruments

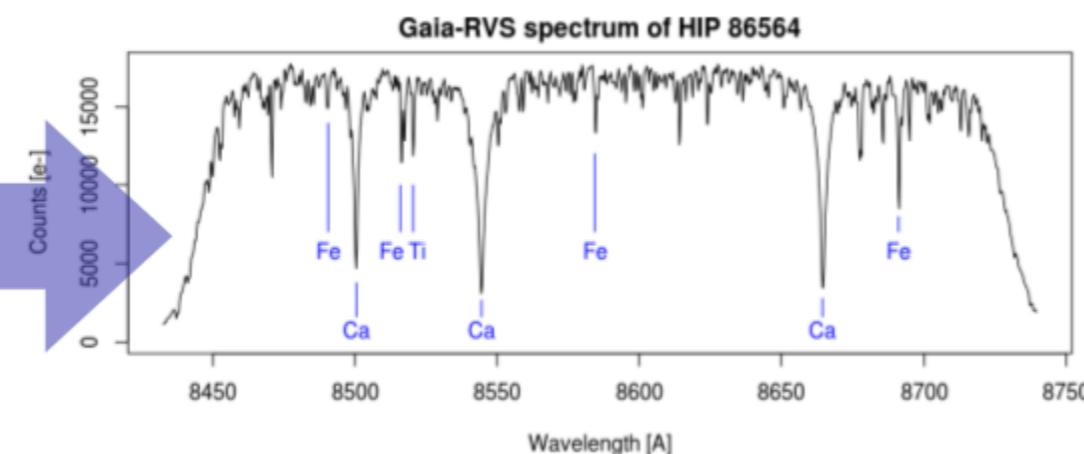
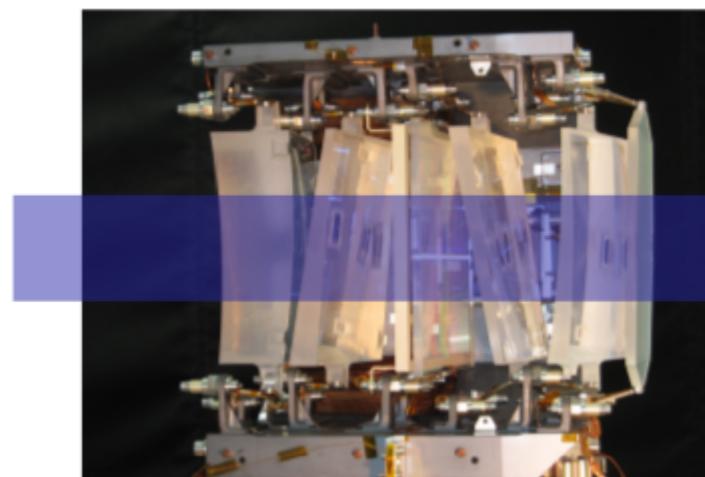
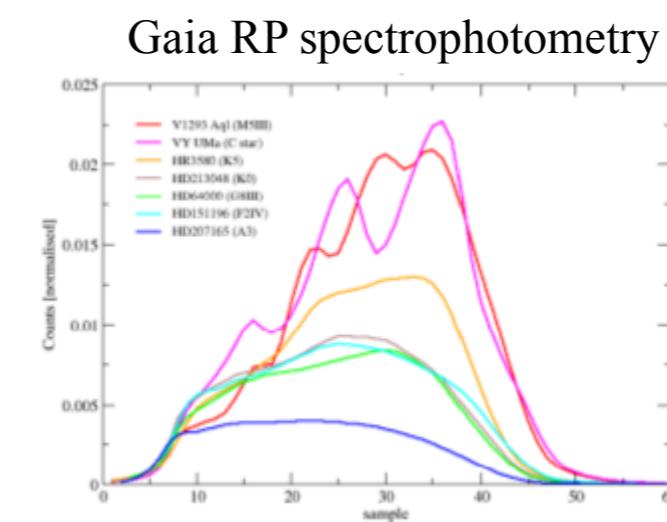
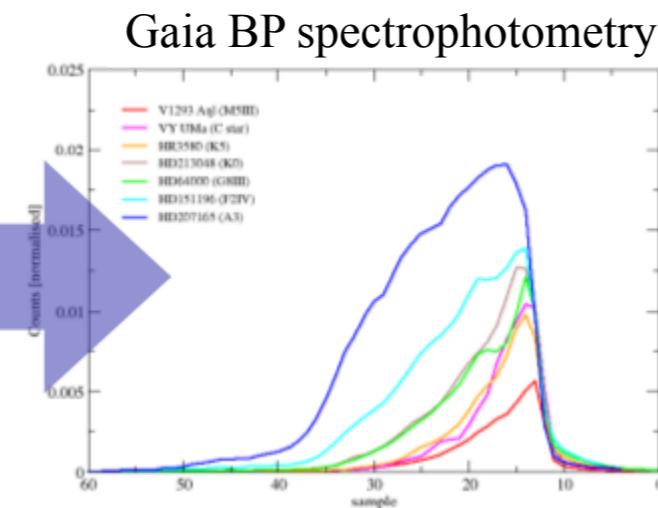
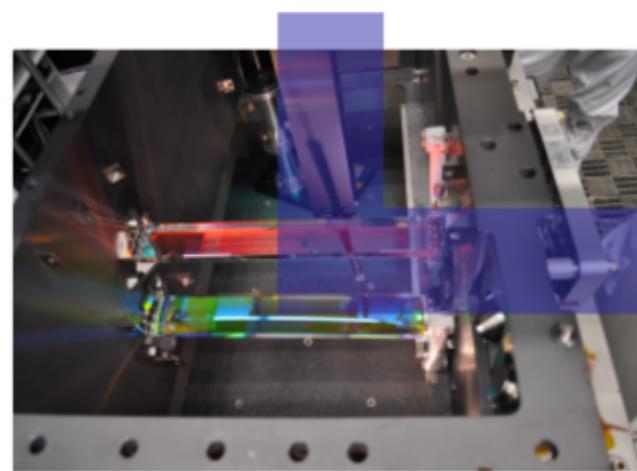
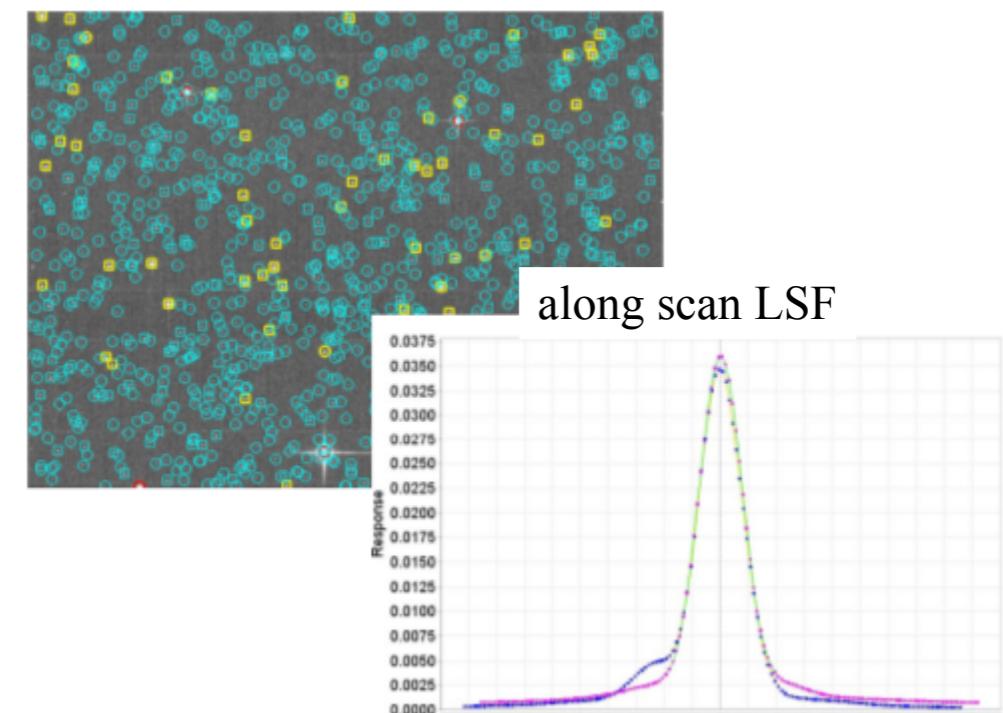
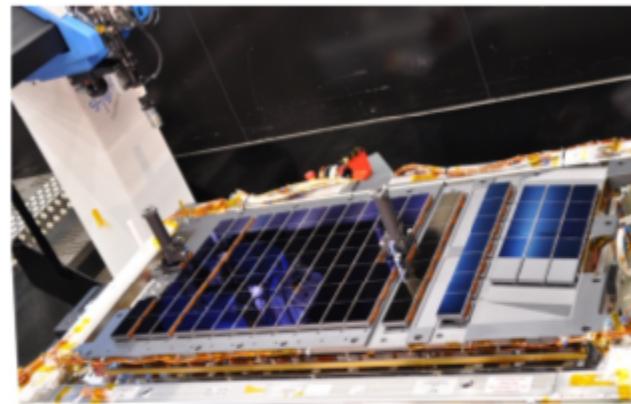
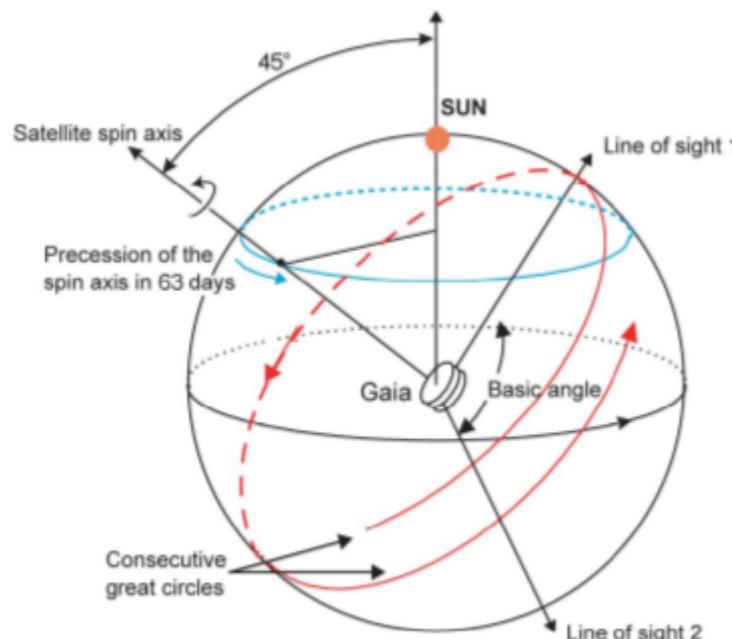
- Astrometry
- Spectrophotometry
- Spectroscopy (RVS)

> 1 billion stars  $3 < G < 20.7$

~ 70 observations per source



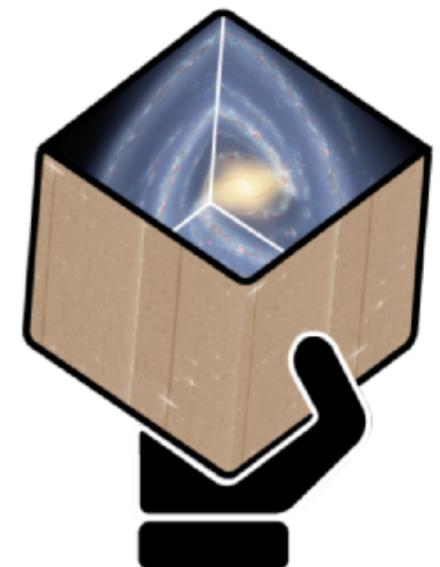
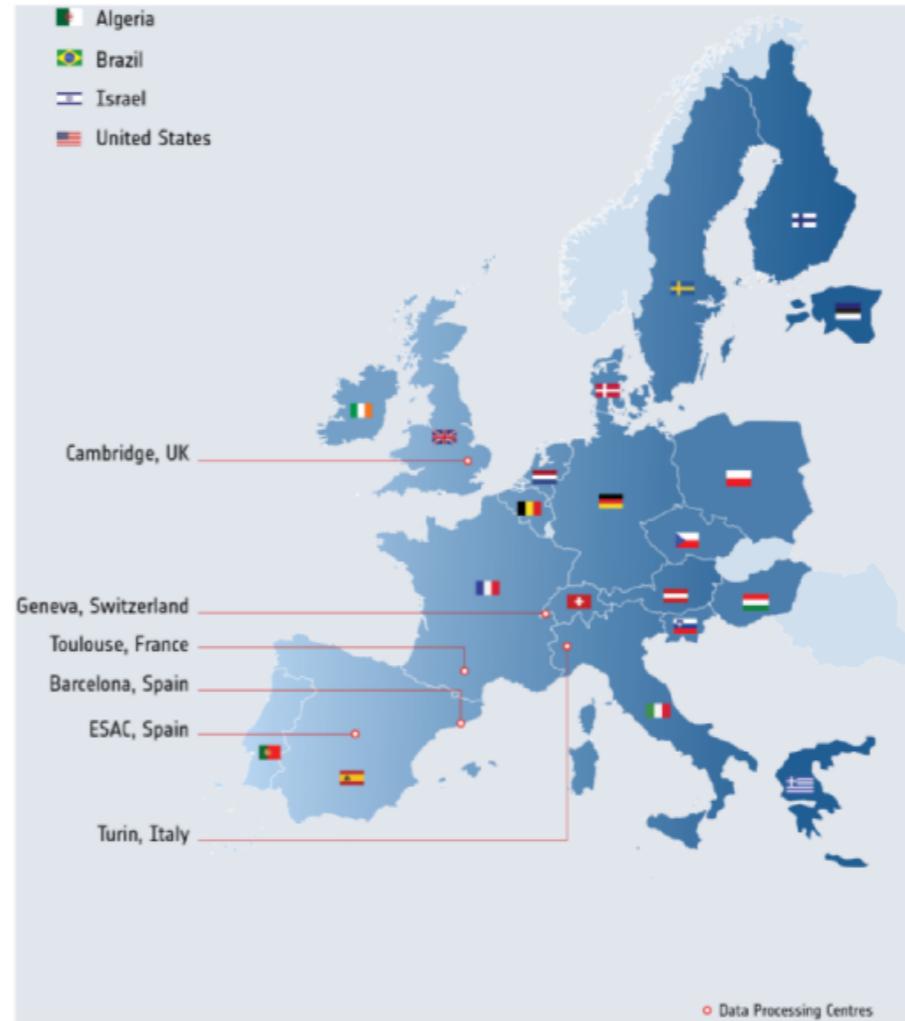
# Gaia instruments and measurements



Figures:  
ESA/Gaia/DPAC/Airbus DS

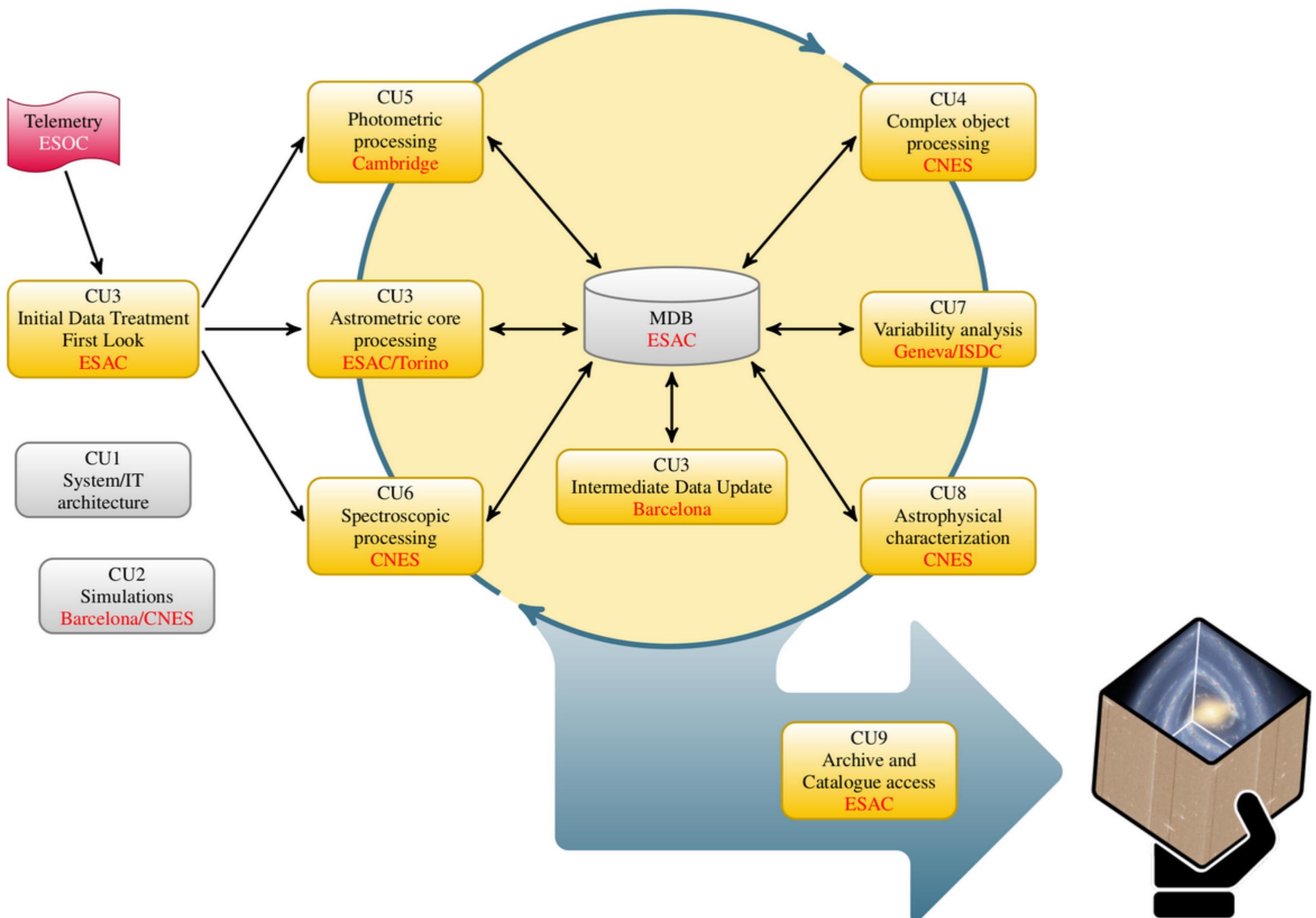
# Teamwork to deliver the promise of Gaia

- 17+ years of effort
  - 450 scientists and engineers
  - 160 institutes
  - 24 countries and ESA
  - 6 data processing centres



$\mu_s G$

# Gaia data processing



# The Gaia schedule

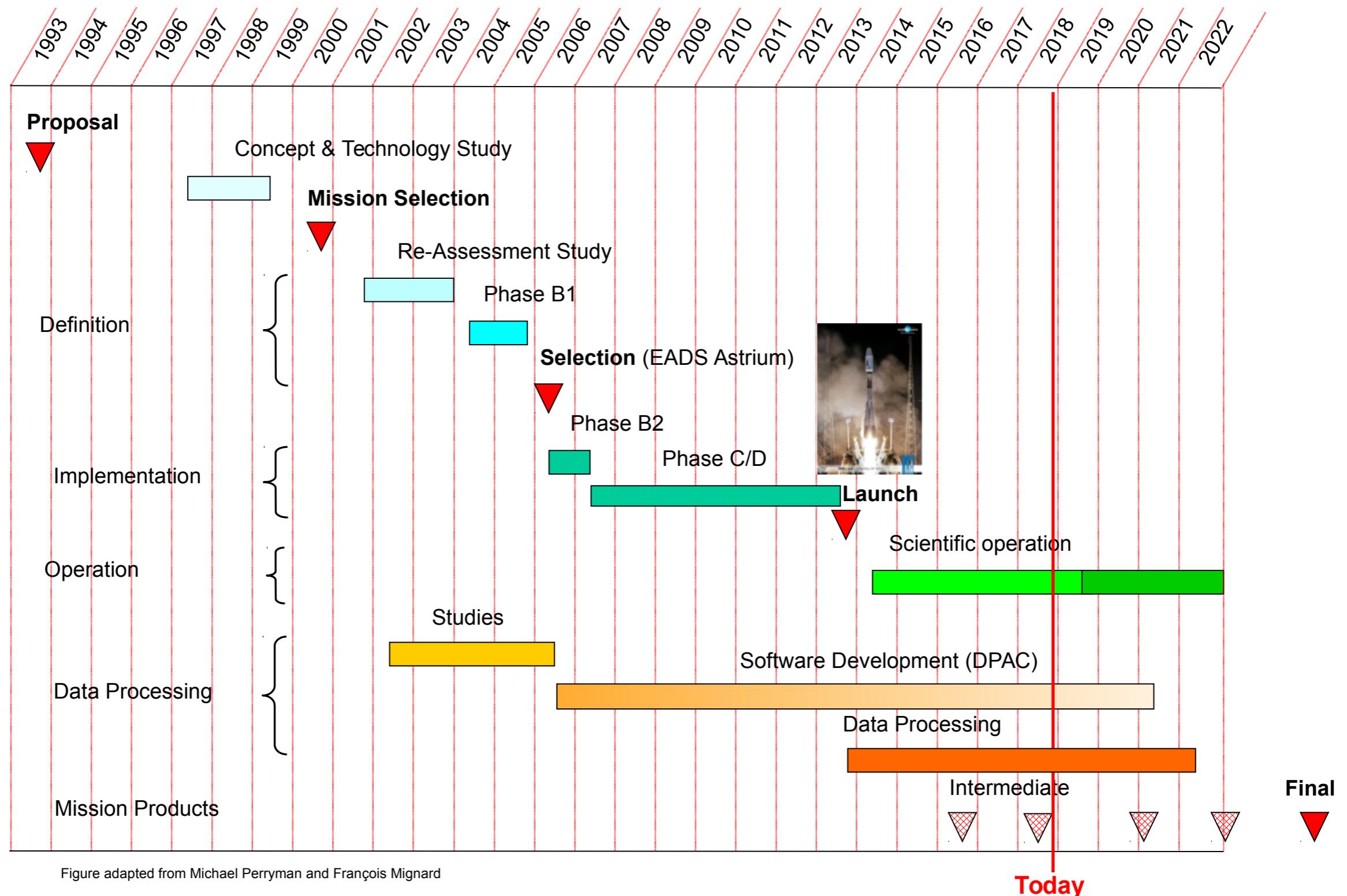


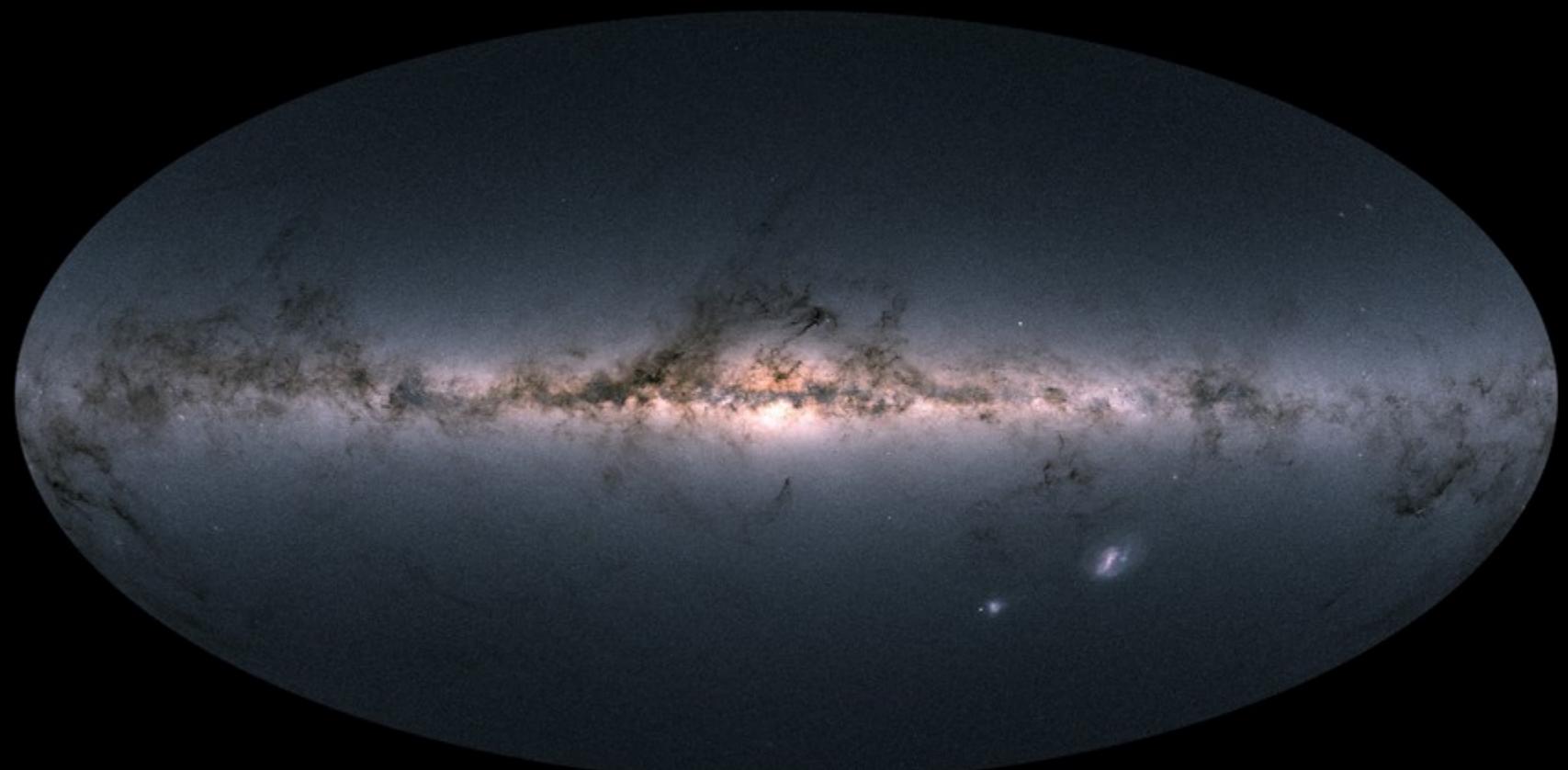
Figure adapted from Michael Perryman and François Mignard

Today



# The second Gaia data release

- Gaia mission overview
- Gaia DR2 content and limitations
- A few (biased!) applications

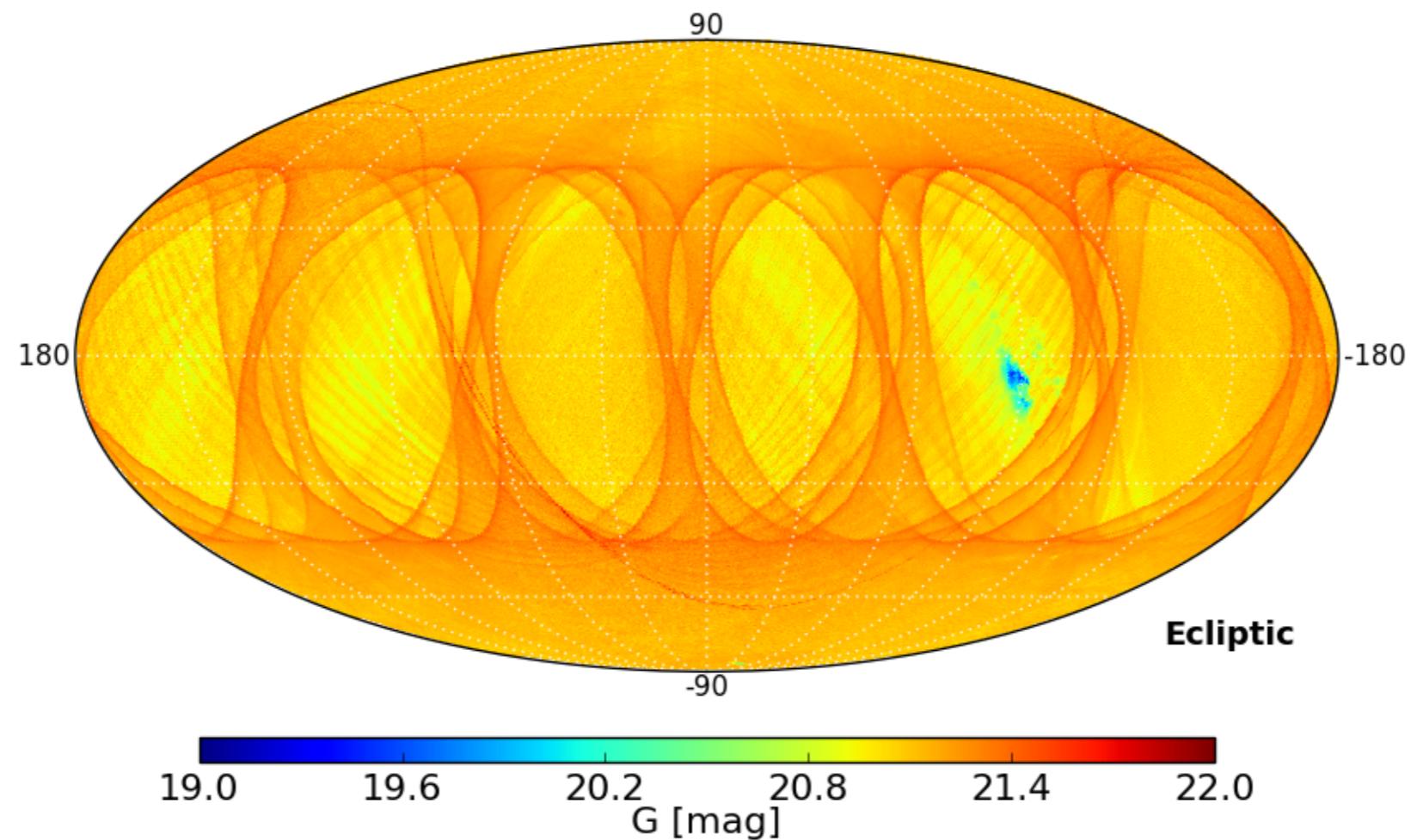


# Gaia DR2 content

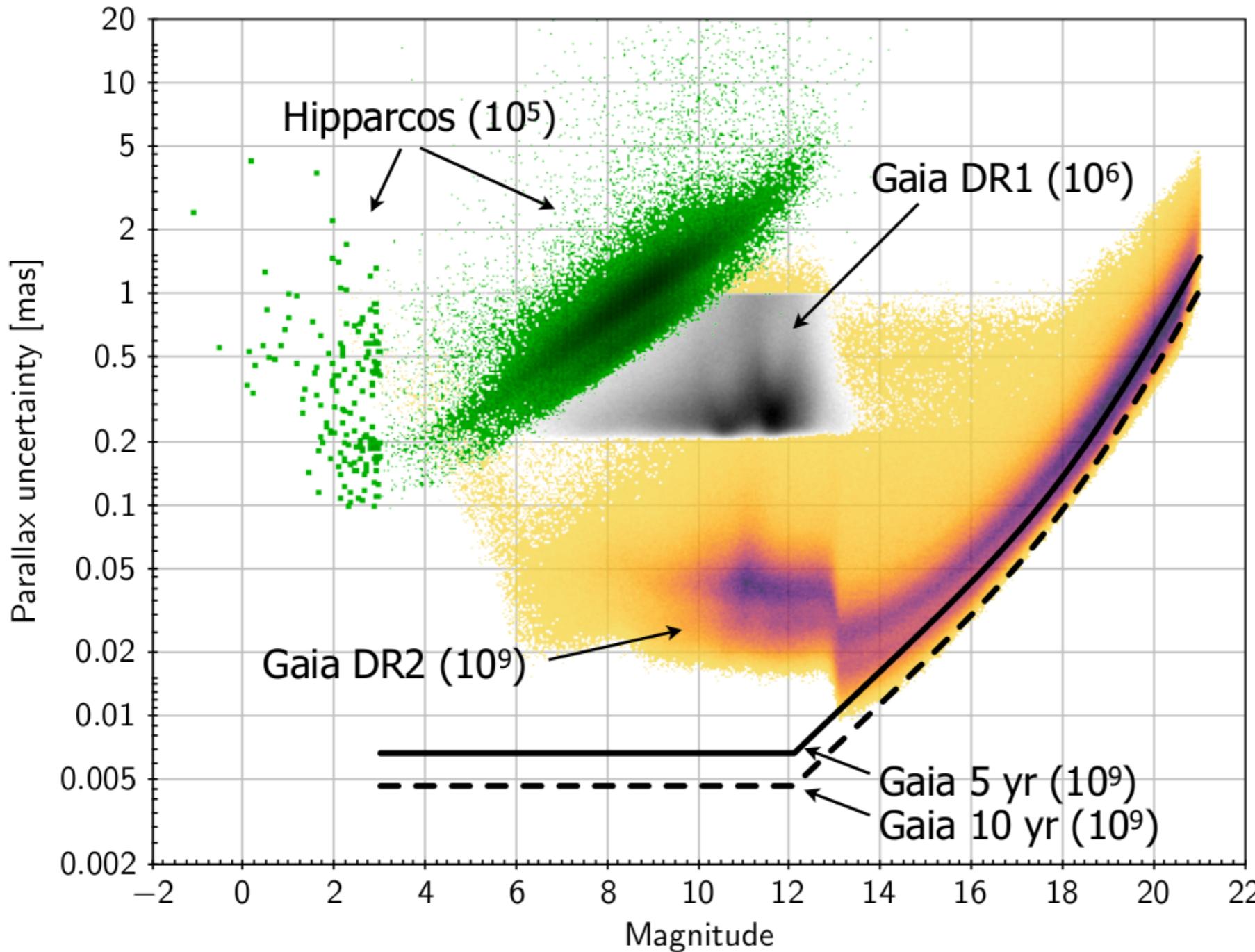
	<b>DR1</b> <b>(Sept 2016)</b>	<b>DR2</b> <b>(25 April 2018)</b>	<b>DR3</b> <b>(2021)</b>	<b>DR4</b> <b>(TBD)</b>
Parallaxes and proper motions	Hip/Tyc priors	Full Sample	++	++
Photometry	G	G, G <sub>BP</sub> , G <sub>RP</sub>	++	++
Variables	3 000	550 000	++	++
Radial velocities	-	RVs at G <sub>RVS</sub> <12	++	++
SSOs	-	pre-selected asteroids	New SSOs	++
Astrophysical parameters	-	for G < 17 : Teff, A <sub>G</sub> Radii and luminosities <i>from integrated phot</i>	Classification + parameters from BP/RP + RVS spectra	++
Systems	-	-	Non-single catalogue Extended objects	Exoplanet list
Spectra	-	-	Mean BP/RP spectra Mean RVS spectra	++
Epoch data	-	-	-	All epoch data

More DRs to be planned: mission extension !

# DR2 completeness



# DR2 astrometry



Typical parallax precision:

- $G=15$ : 0.03 mas
- $G=17$ : 0.1 mas
- $G=20$ : 0.7 mas

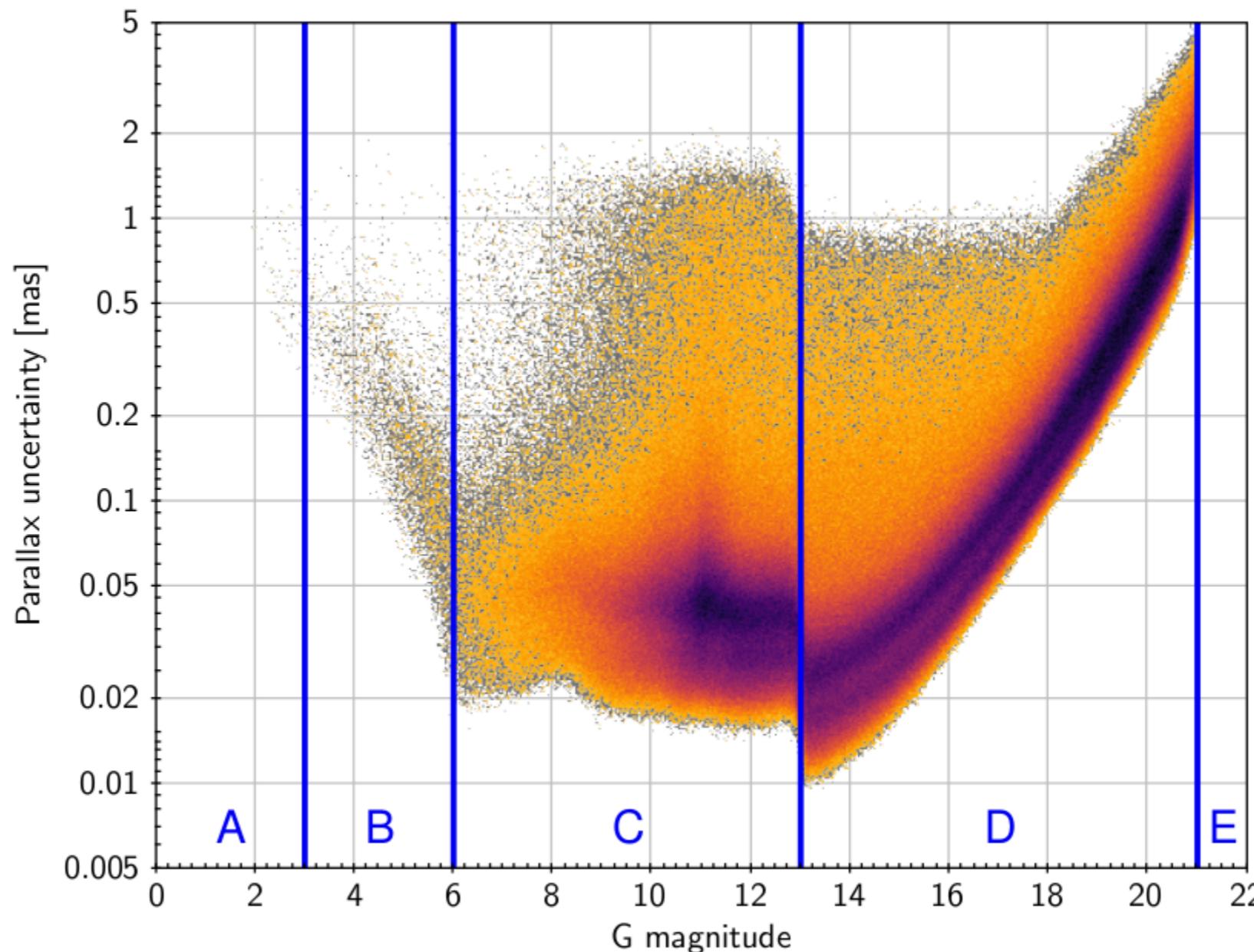
Still single-star solution

Systematics below 0.1 mas

Global zero point  $\sim -0.03$  mas

Spatial correlations at  $\sim 1$  and  $\sim 20$  degree scales

# DR2 astrometry : formal uncertainties



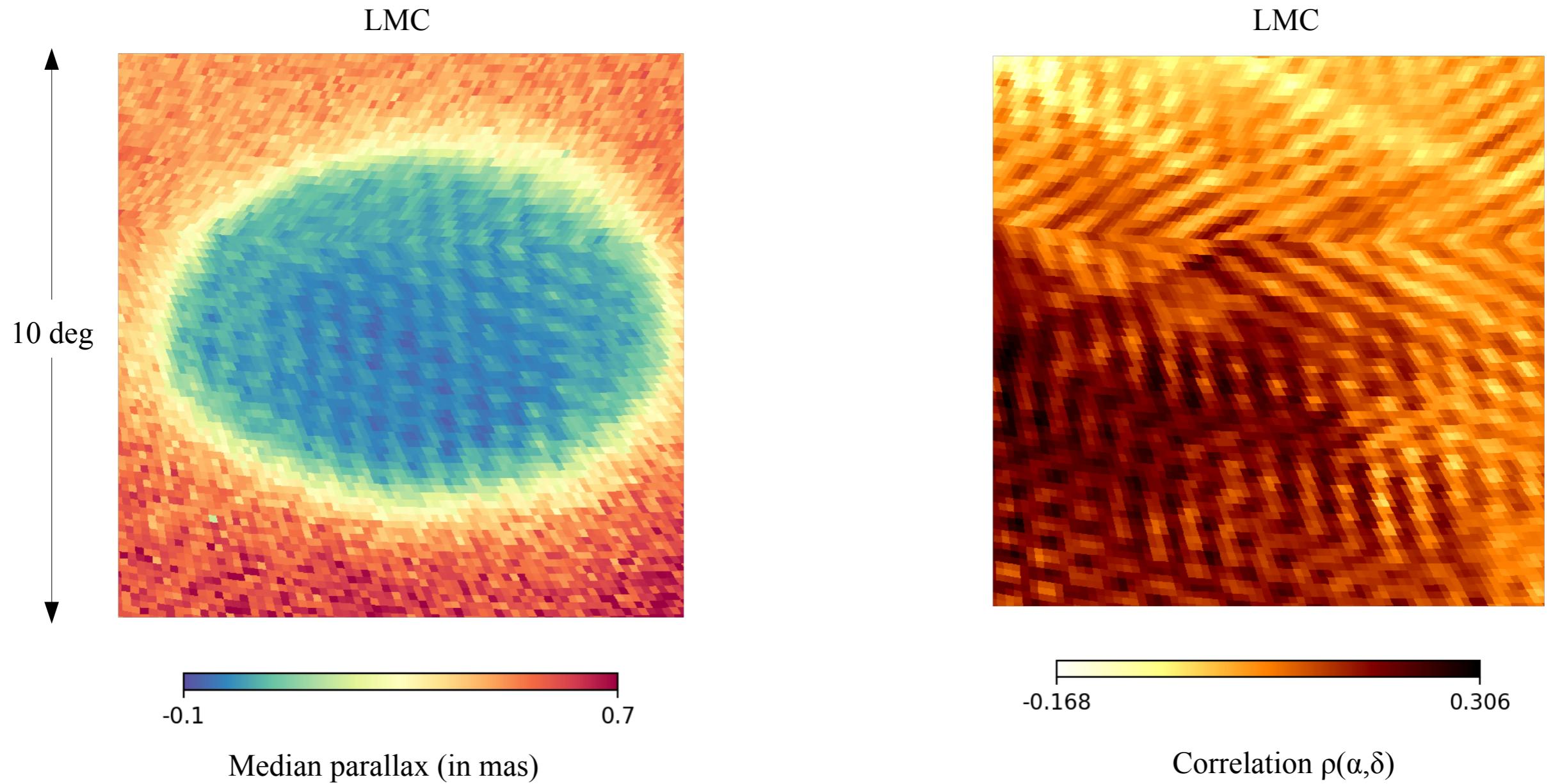
*Regimes of G :*

- A: Too bright**
- B: Partly saturated (unreliable)**
- C: Detector and calibration limited**
- D: Photon limited**
- E: Too faint (not published)**

Provided: **formal uncertainties**  
estimated from internal consistency of the measurements  
do not represent the total error

# DR2 astrometry : small scale systematics

Quasi-periodic patterns imprinted by the Gaia scanning law

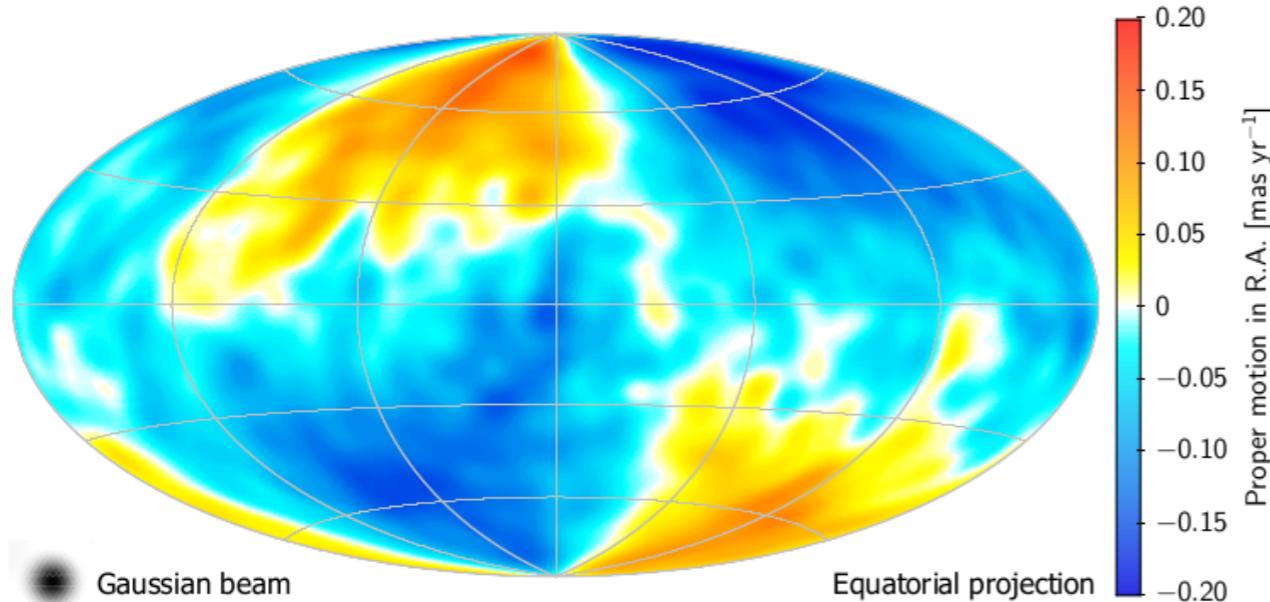


Characteristic period  $\sim 0.6$  deg, RMS variation  $\sim 0.03$  mas

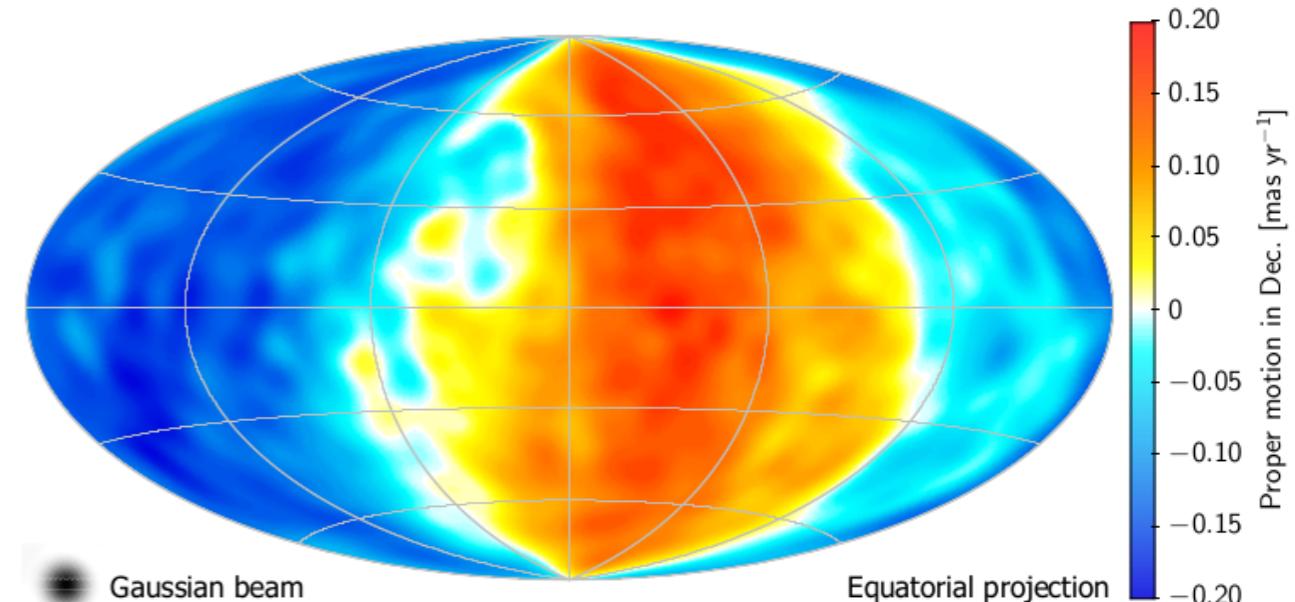
# DR2 astrometry : bright stars PM systematics

Proper motion difference versus Hipparcos

R.A.



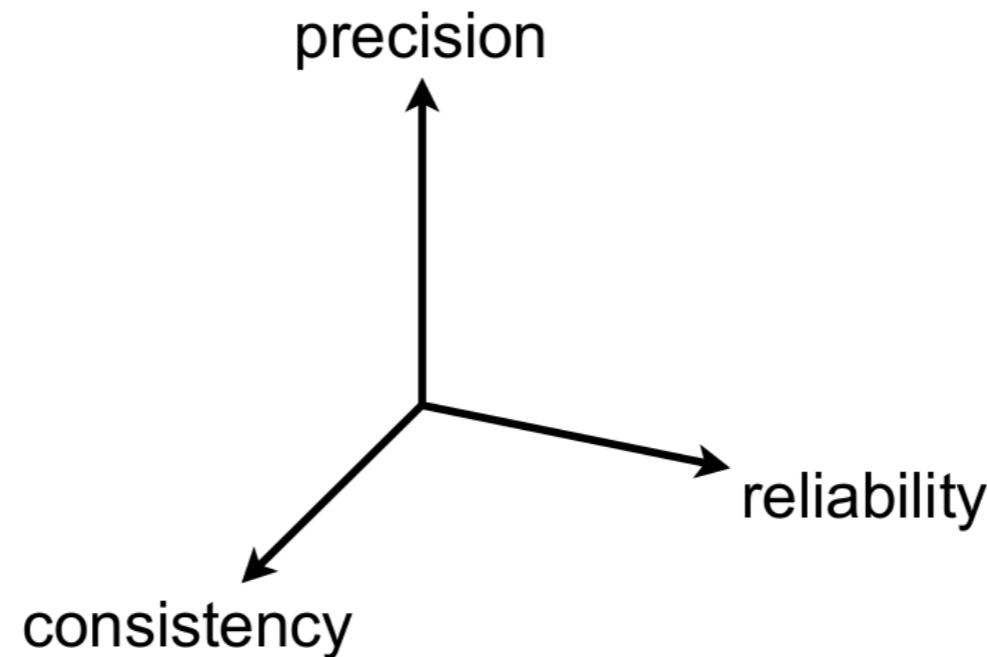
Dec.



Global rotation  $\sim 0.15$  mas yr<sup>-1</sup>

# DR2 astrometry : quality indicators

---

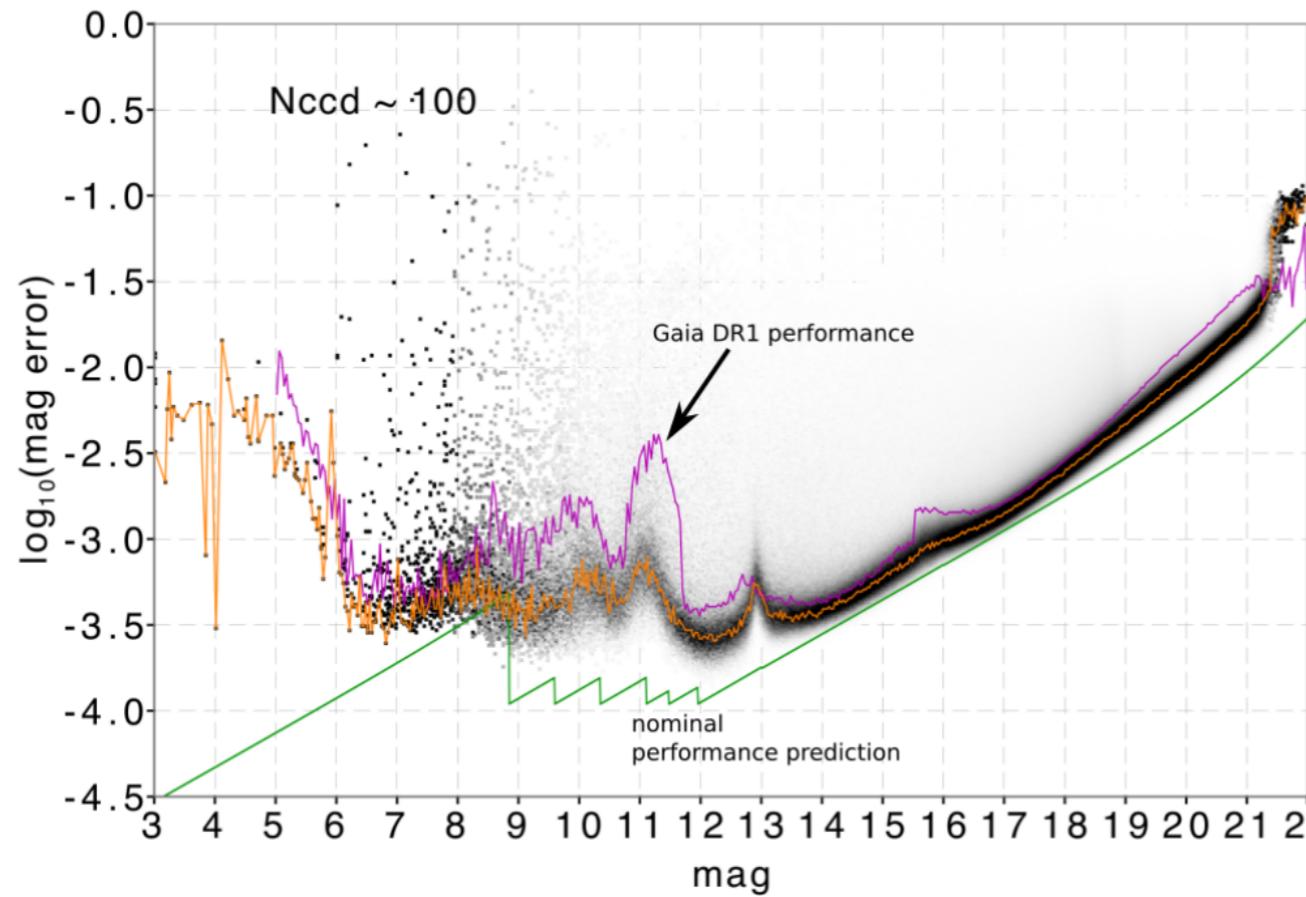


**Precision:** `parallax_error`, `pmra_error`, `pmdec_error` (internal!)

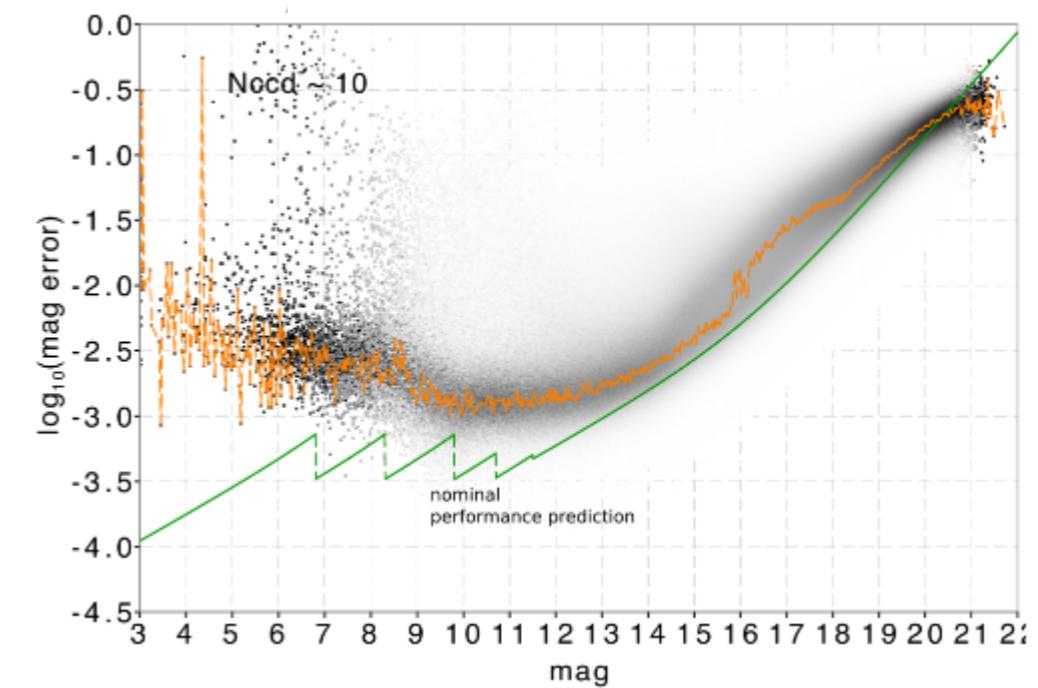
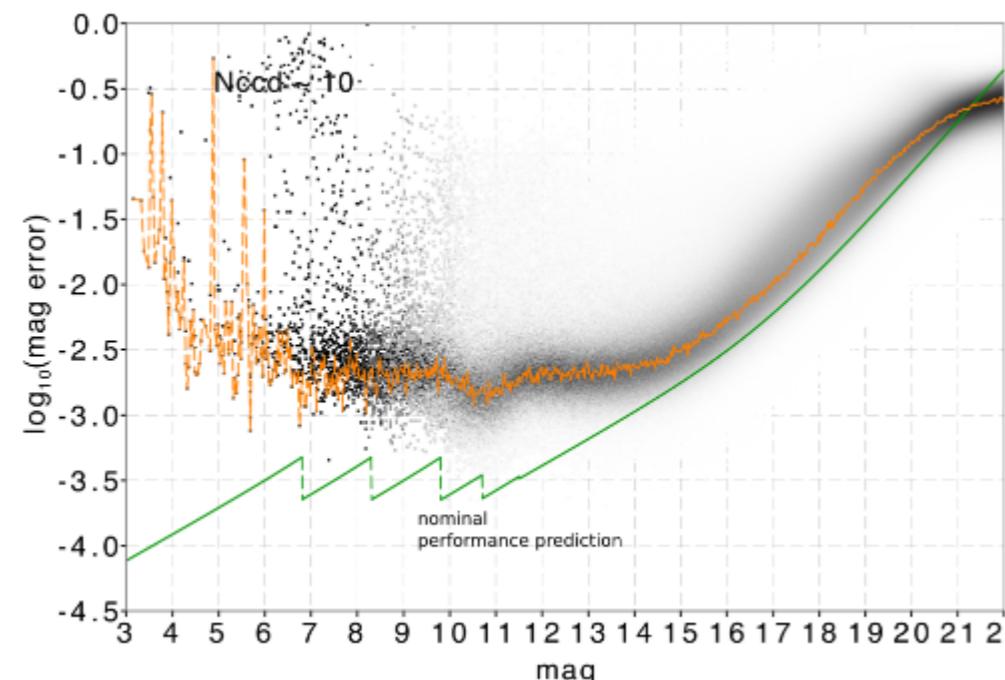
**Reliability:** `visibility_periods_used`

**Consistency:** Renormalised Unit Weight Error (see Gaia DR2 Known issue page)

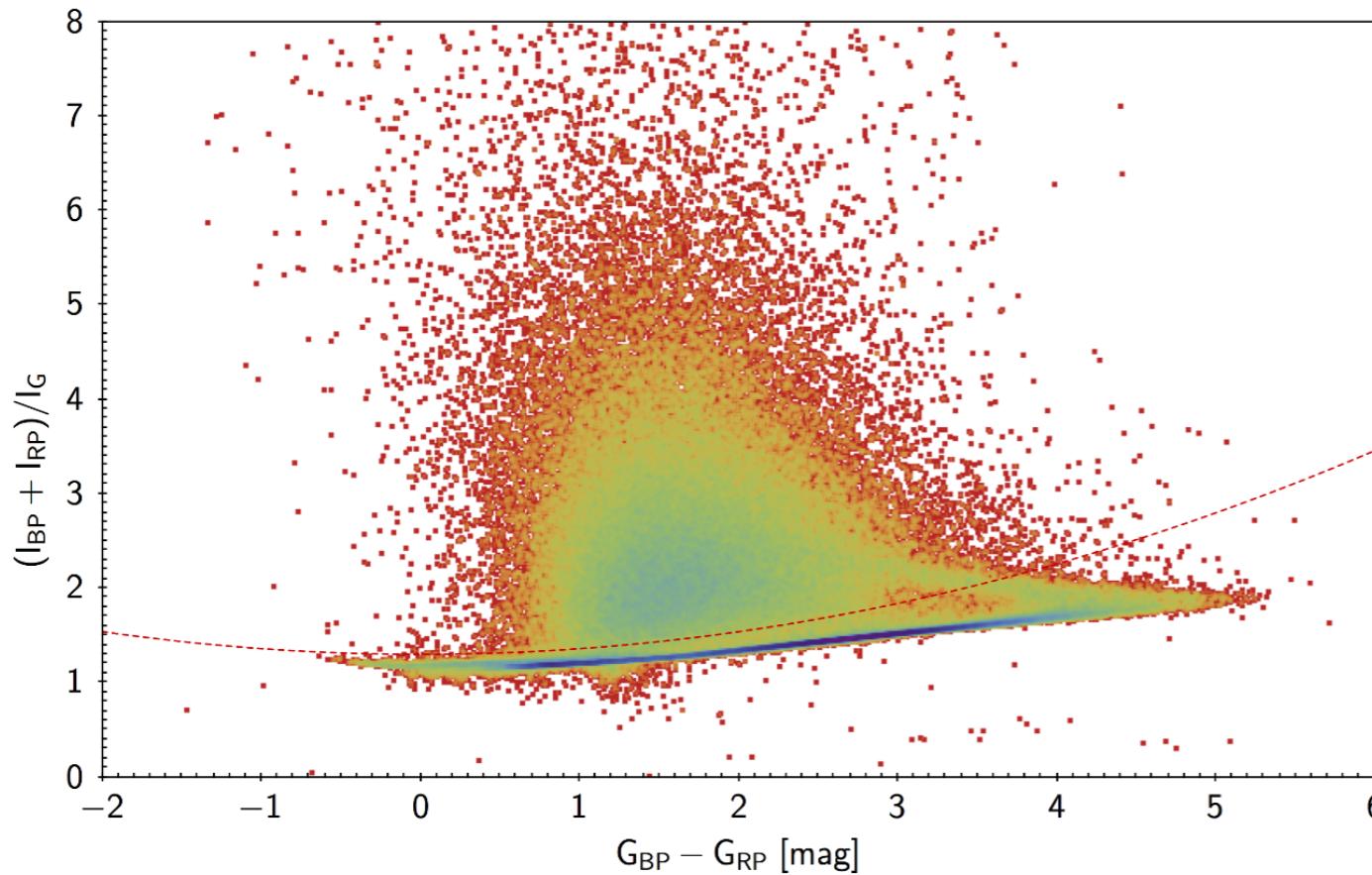
# DR2 photometry



~ 10 mmag level systematics  
BP & RP integrated photometry  
 $G=20$ :  $\sigma_G \sim 0.02$  mag,  $\sigma_{XP} \sim 0.2$  mag  
Passbands provided  
! No deblending  
! G band system different from DR1

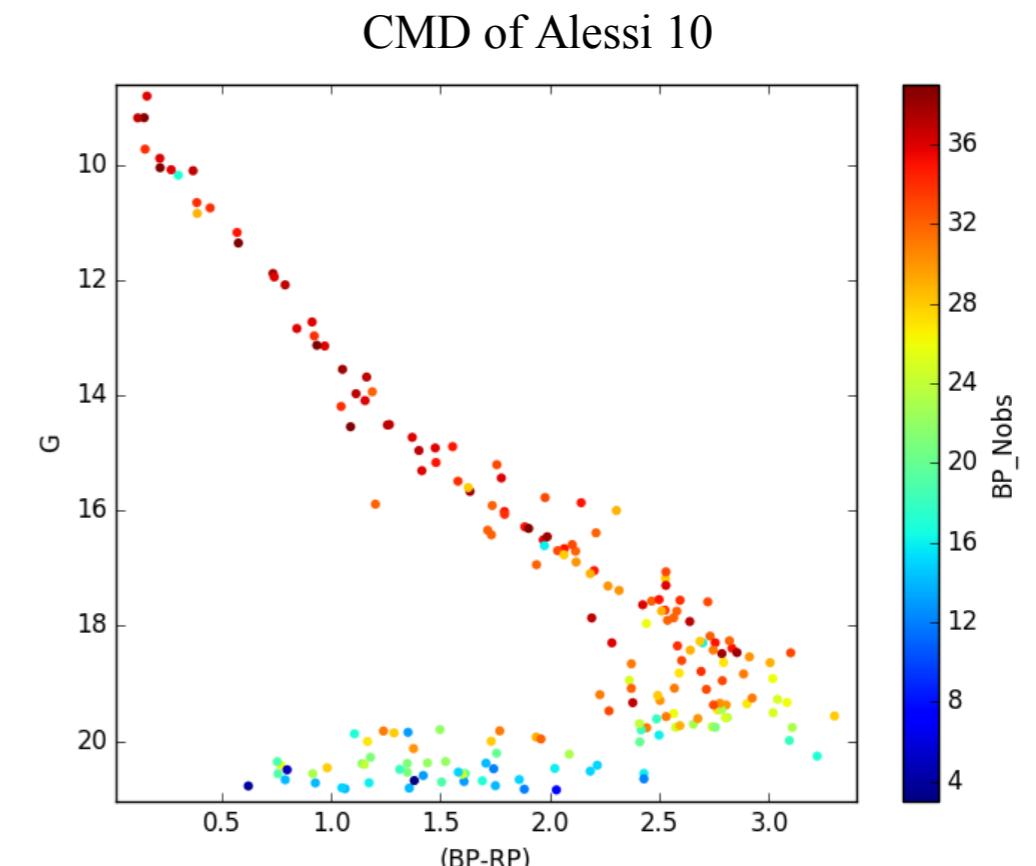


# DR2 photometry : main issues



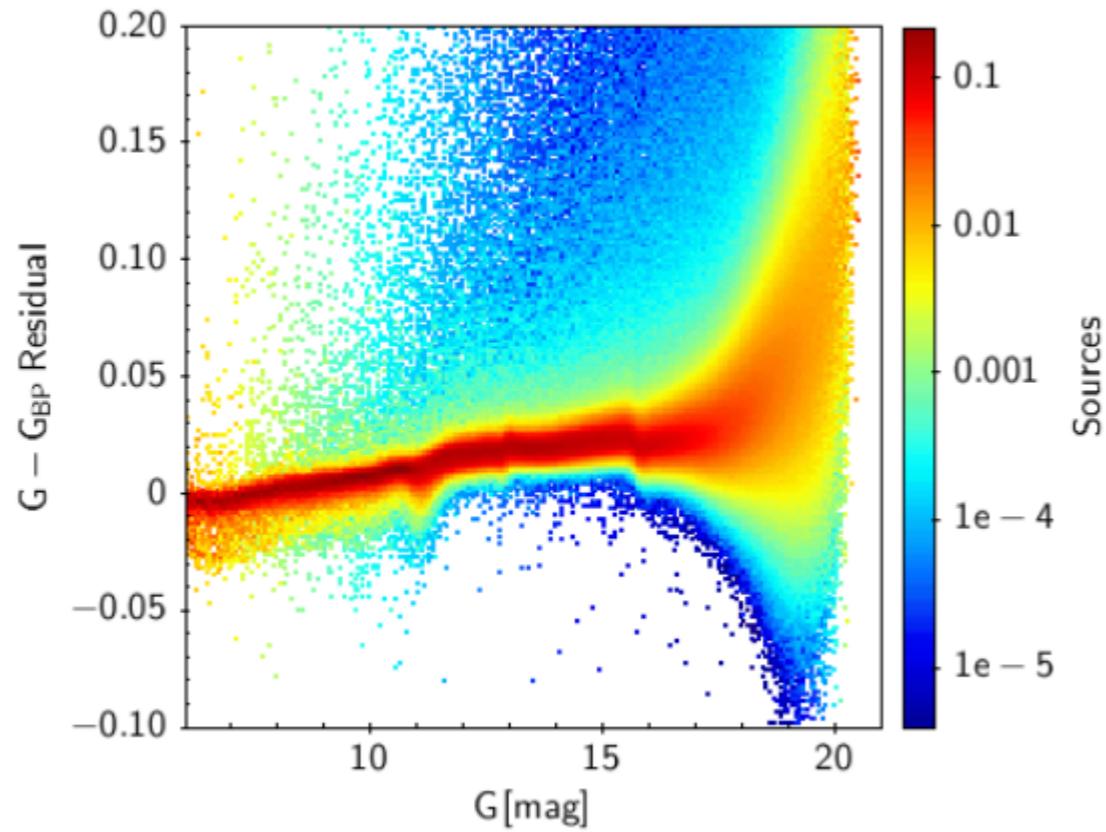
BP/RP excess flux  
← crowded spectrophotometry

BP background under-estimation



# DR2 photometry : calibrations

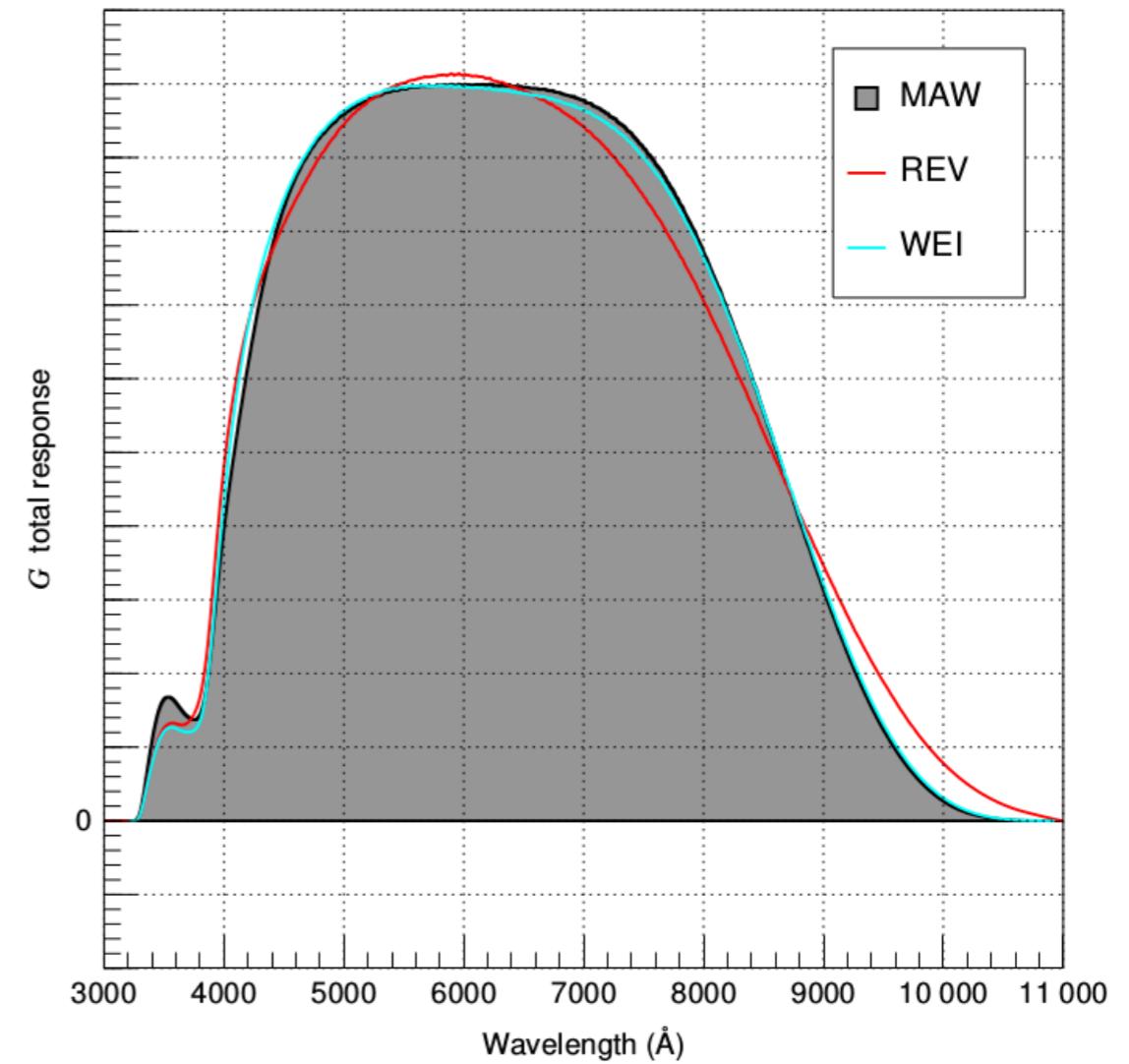
G fading of  $\sim 3$  mmag / mag



*Arenou et al. 2018*

*Weiler 2018*

New calibrations of the Gaia photometric bands

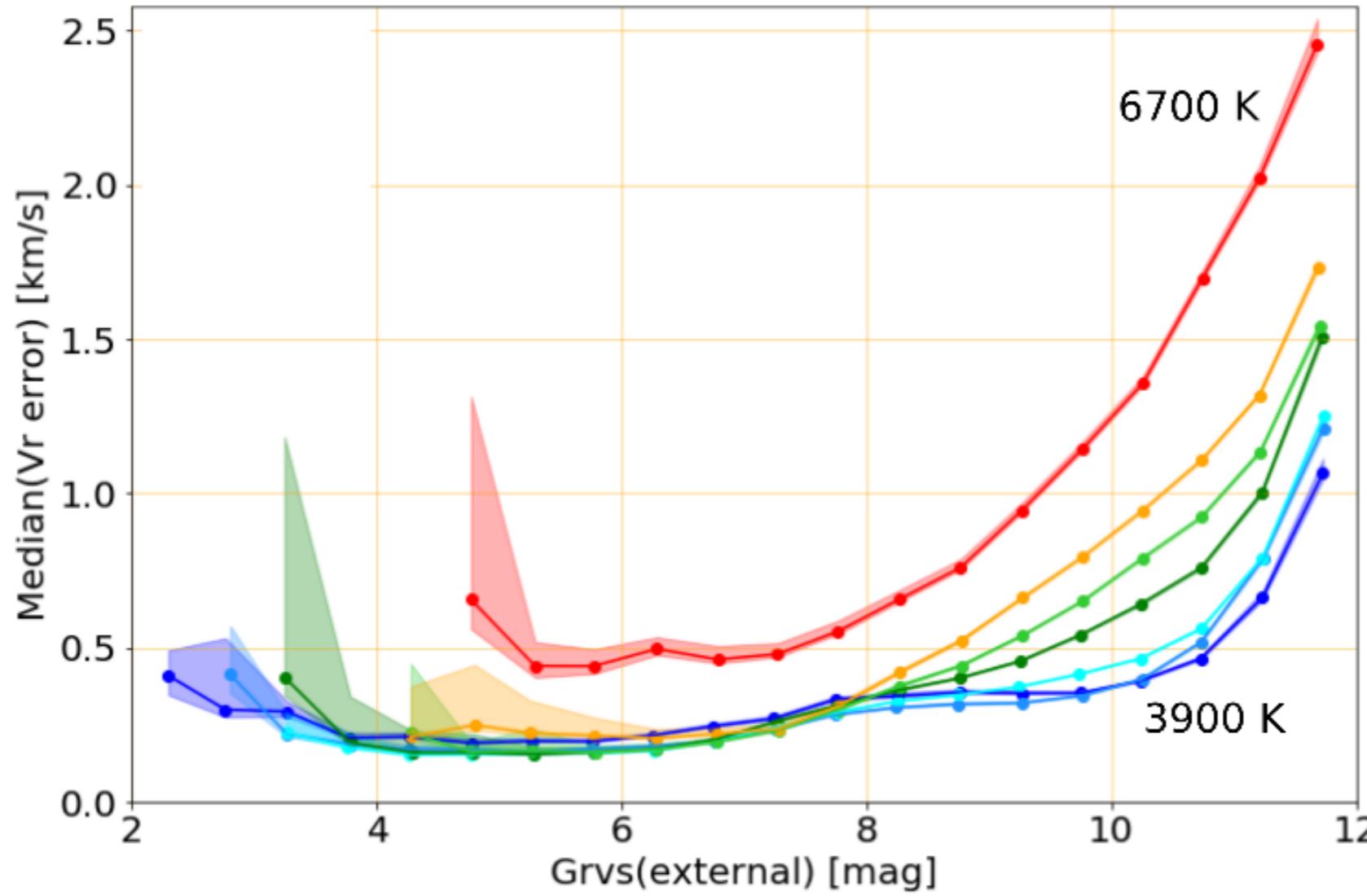


*Evans et al. 2018*

*Weiler 2018*

*Maiz Apellaniz & Weiler 2018*

# DR2 radial velocities

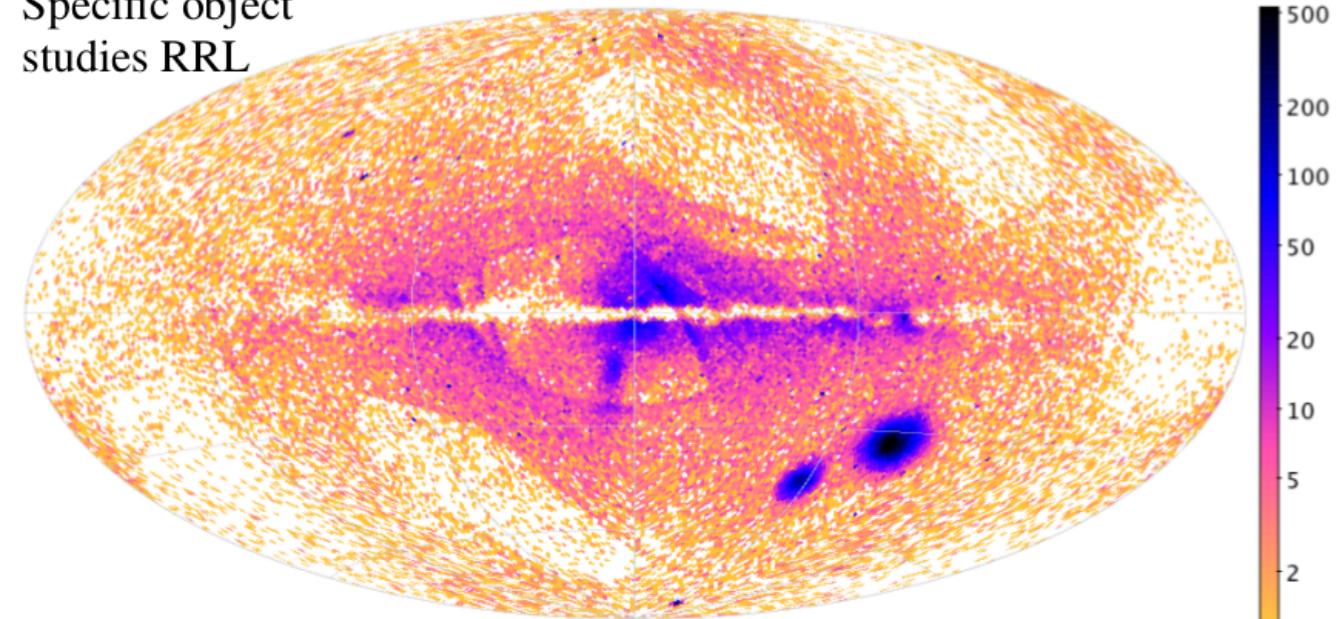


More than 7 million Vr

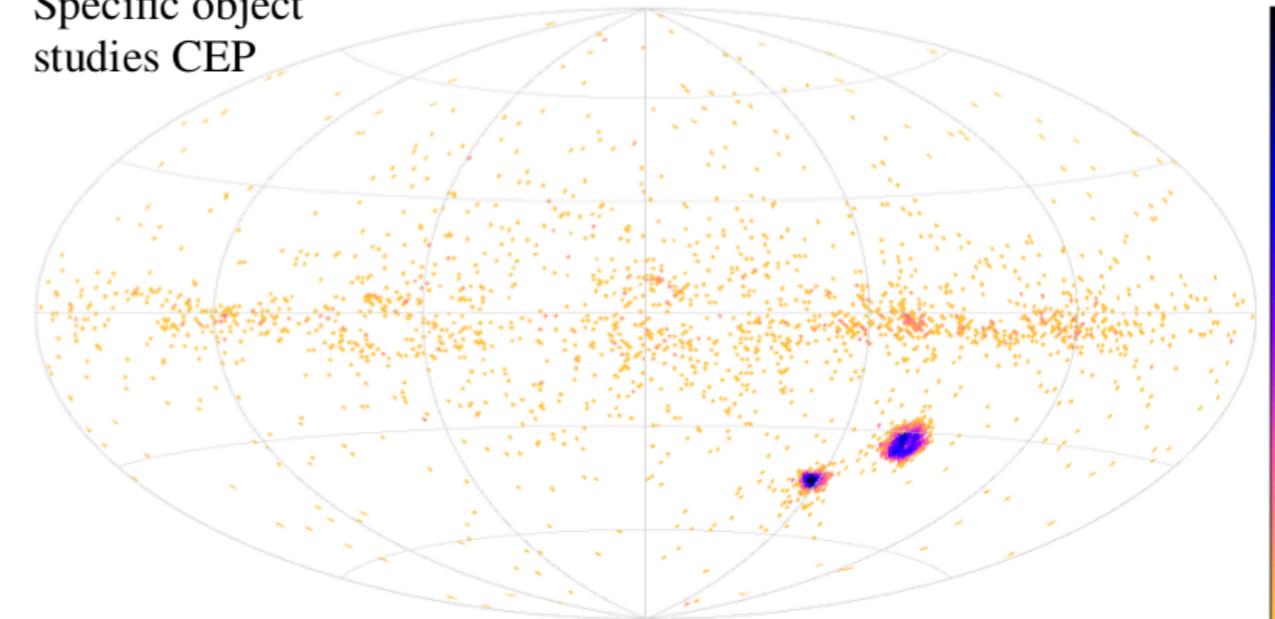
- Only for  $\sim 3550 < \text{Teff} < 6900 \text{ K}$
- Zero point at  $\sim 0.1 \text{ km/s}$
- Systematics as a function of magnitude

# DR2 variables

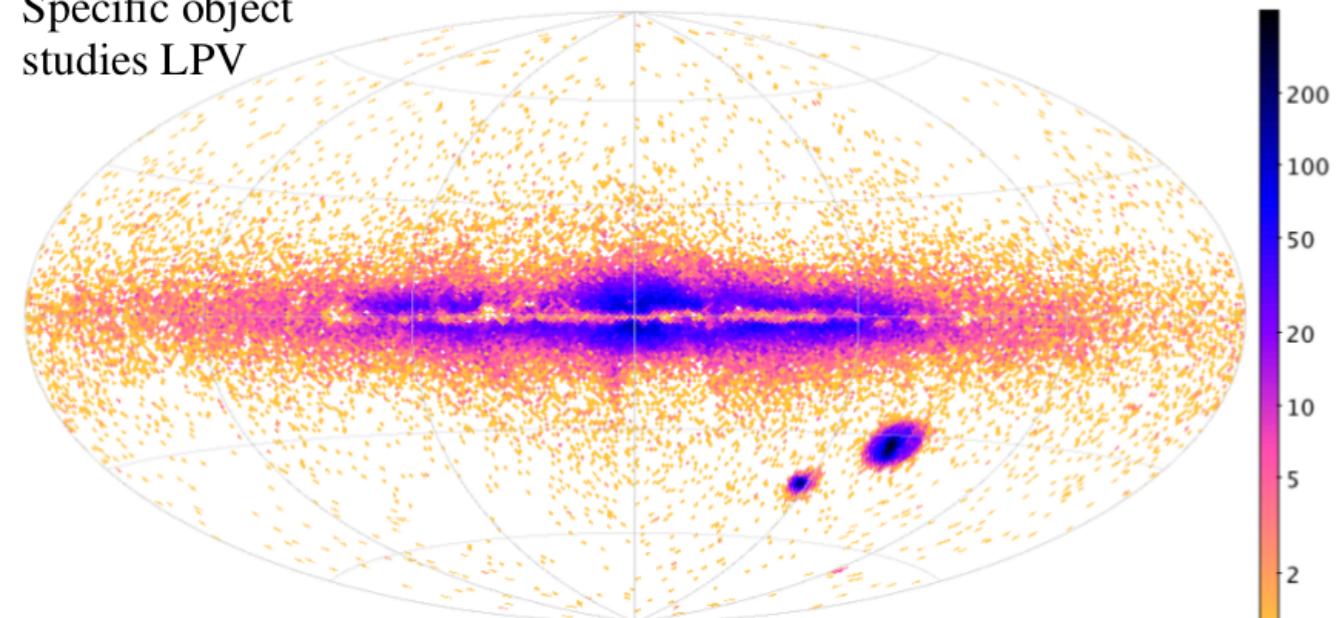
Specific object studies RRL



Specific object studies CEP

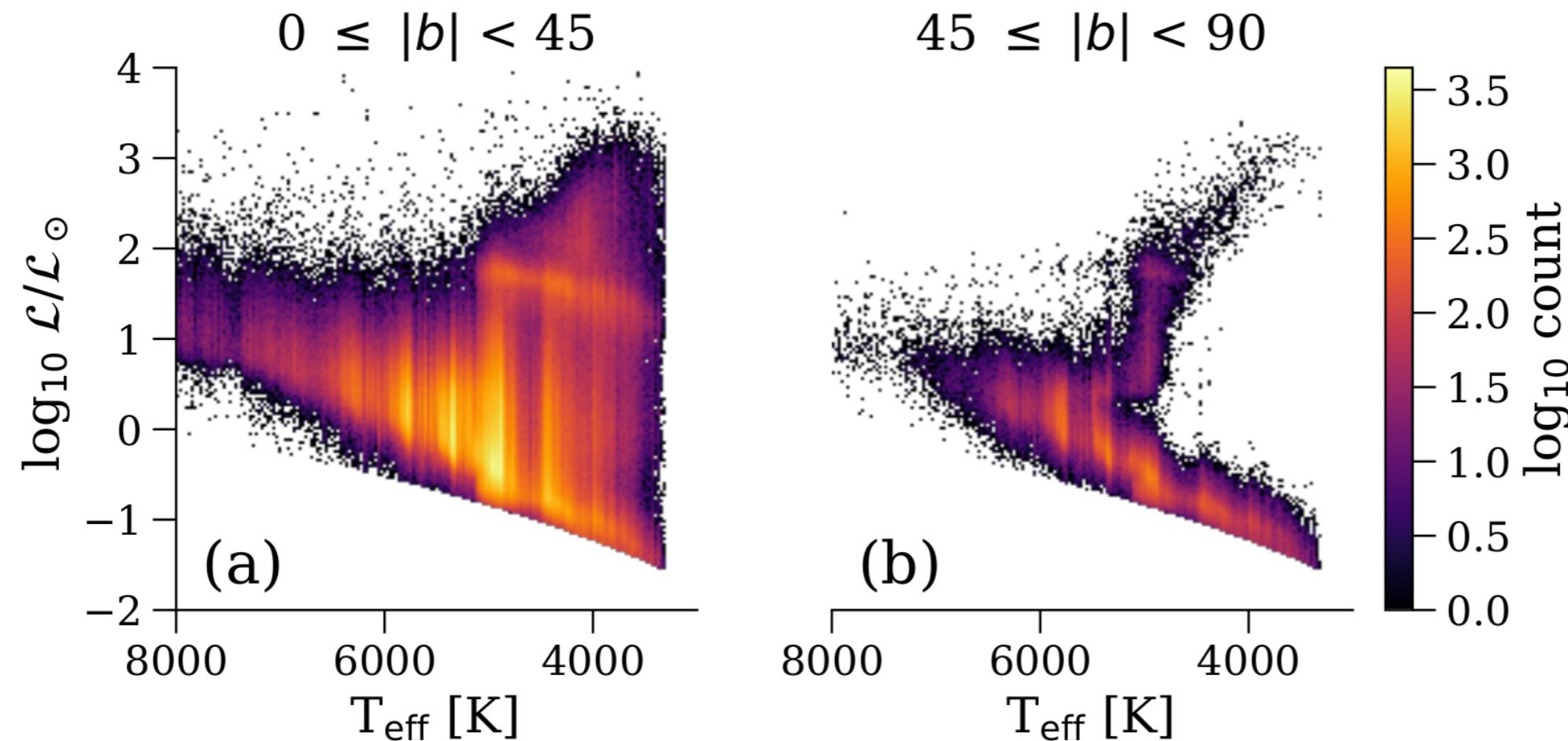


Specific object studies LPV



- 551 thousand variables identified
  - ▶ many more to come in future
- Subset classified by variability type
  - ▶ based on 2+ transits
- Overlapping subset studied in detail
  - ▶ based on 12+ transits

# DR2 astrophysical parameters



$T_{\text{eff}}$ ,  $A_G$ ,  $E(G_{\text{BP}} - G_{\text{RP}})$ , luminosity, radius based only on  $G$ ,  $G_{\text{BP}}$ ,  $G_{\text{RP}}$  and parallax  
→ strong degeneracies

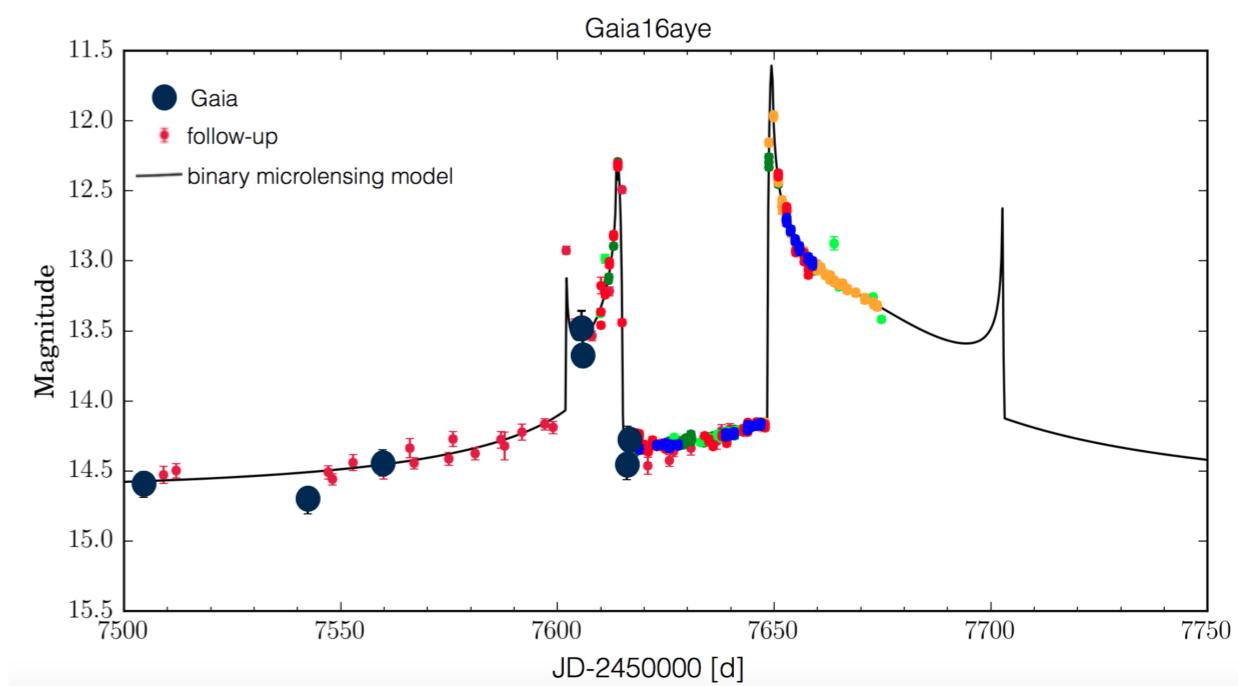
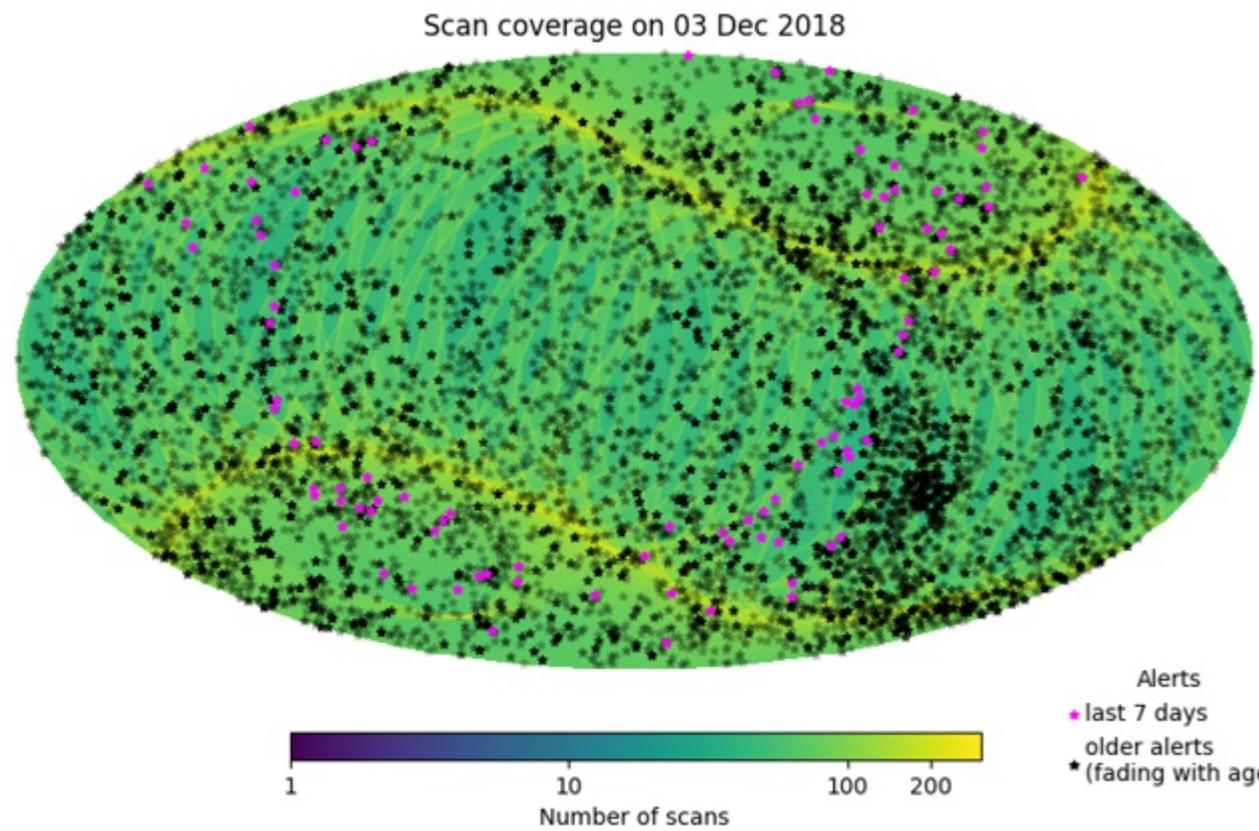
# Lots of documentation to read !

---

- Preliminary solution with associated limitations  
single star solution, calibration issues, ...
- Complex selection function  
scanning law, on-board limitations, pre & post processing filtering
- Uncertainties and covariance matrix
- Several systematics  
parallax zero point, small and large scale systematics,...  
photometry in crowded fields, faint G<sub>BP</sub>, G magnitude drift,...
- Many artefacts  
stars closer than proxima, excess flux

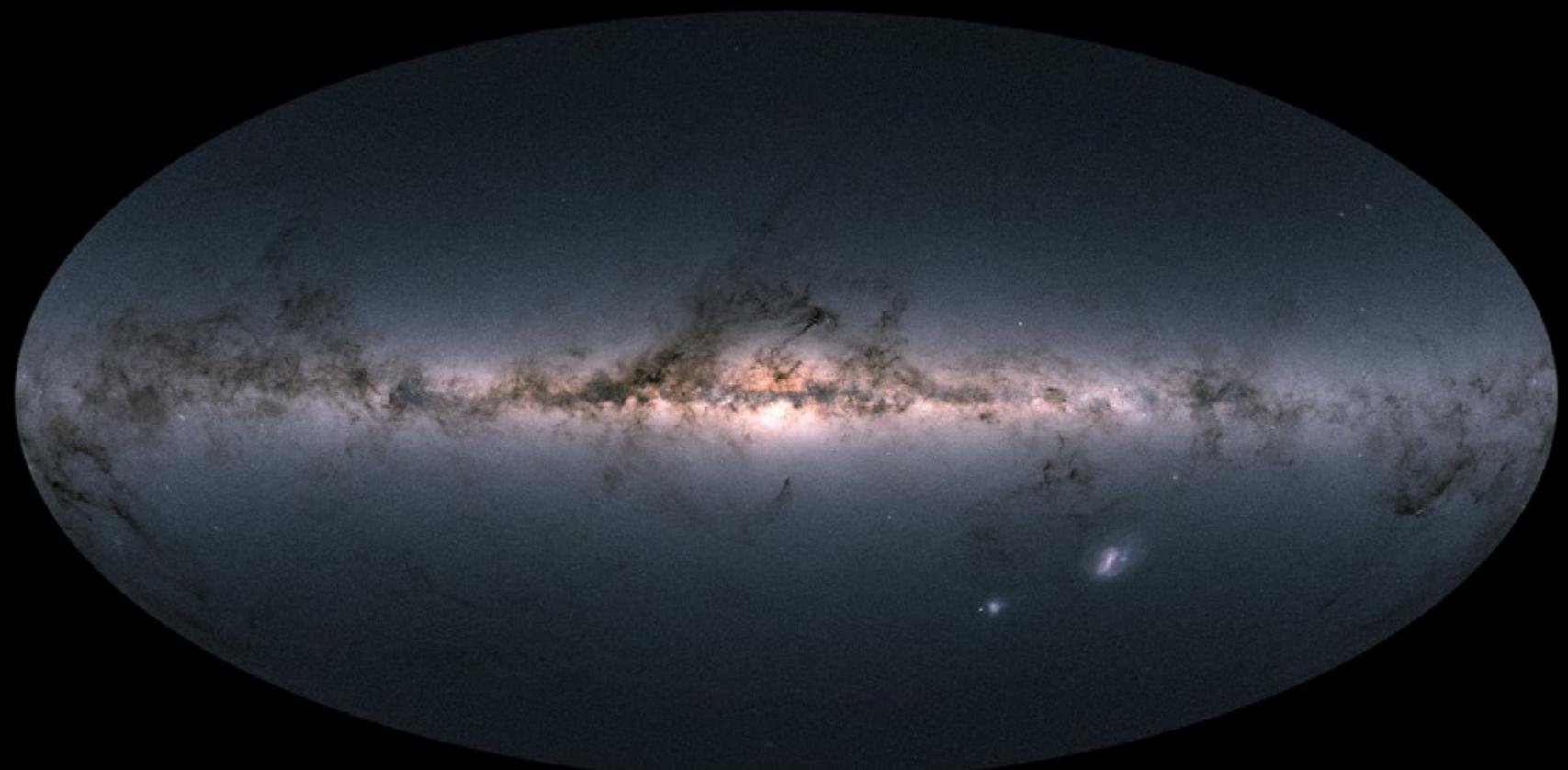
# Gaia Photometric Science Alerts

started publishing alerts in July 2014



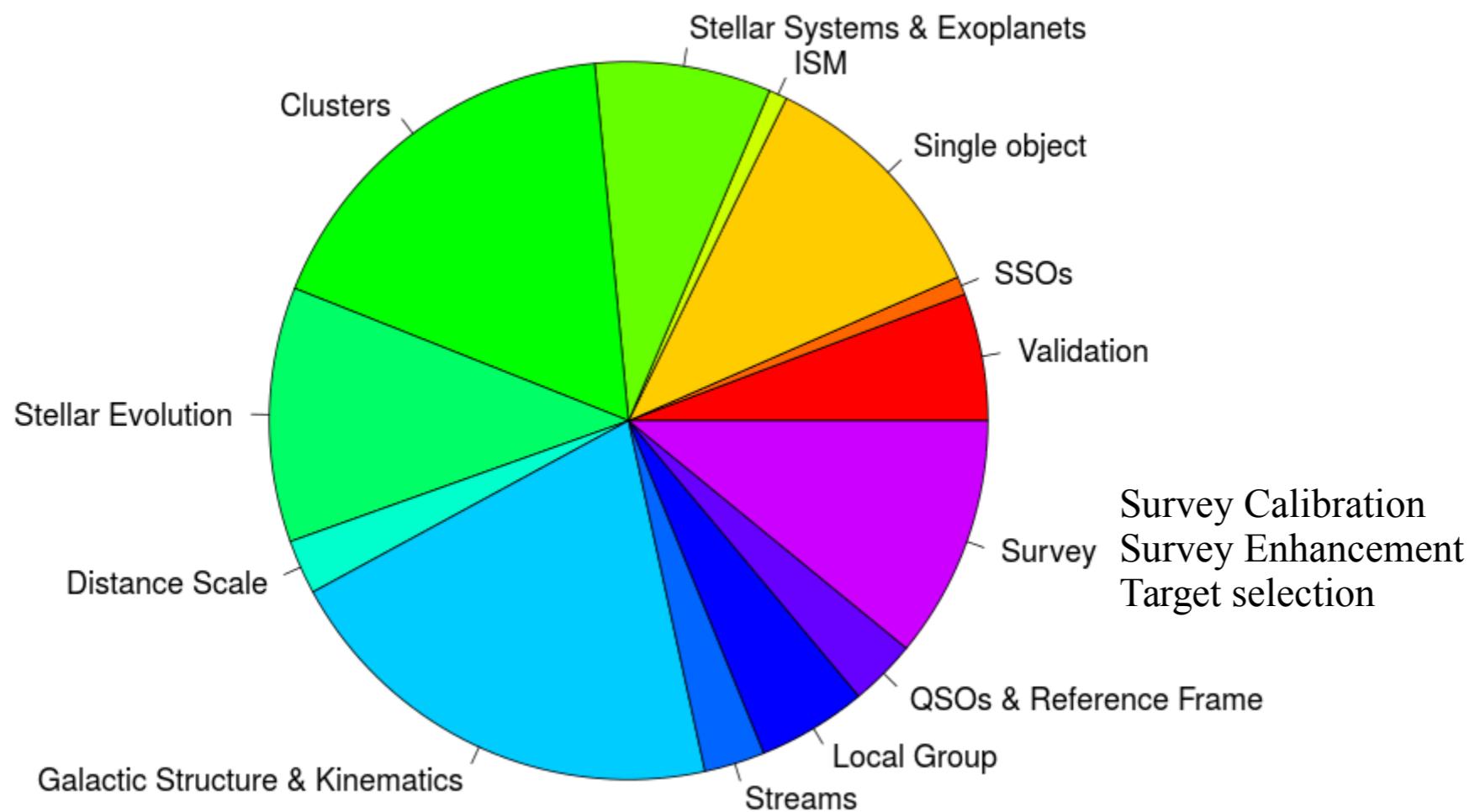
# The second Gaia data release

- Gaia mission overview
- Gaia DR2 content and limitations
- A few (biaised!) applications

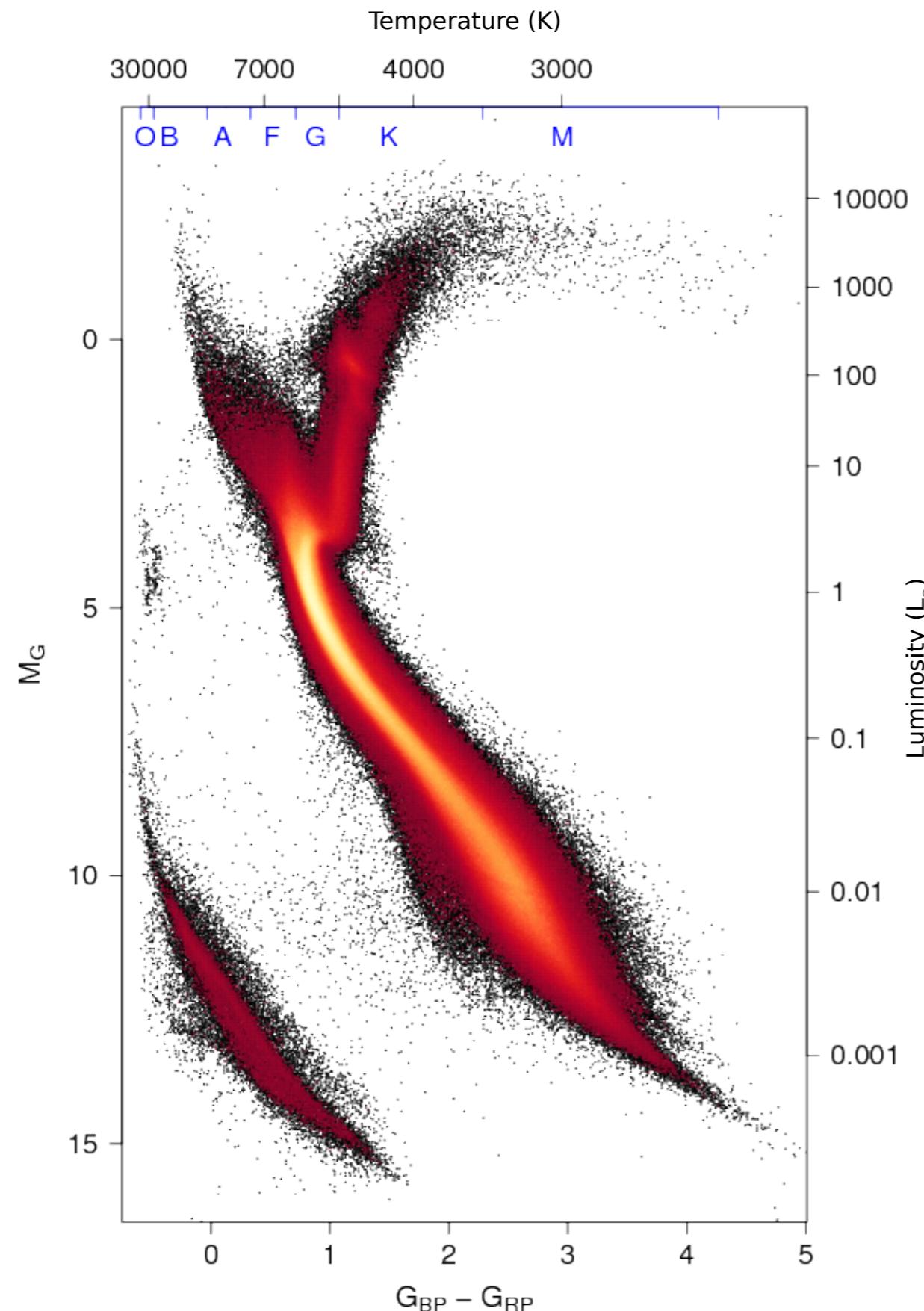


# Gaia DR2 Science applications

> 410 papers on astro-ph using Gaia DR2 ( $\sim 2/\text{jours}$ )

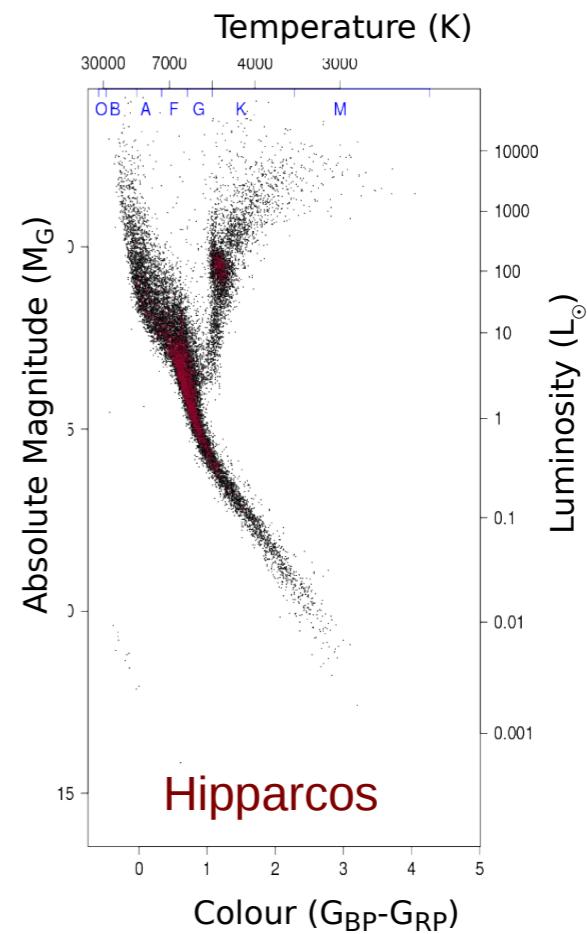


# Gaia DR2 H-R diagram



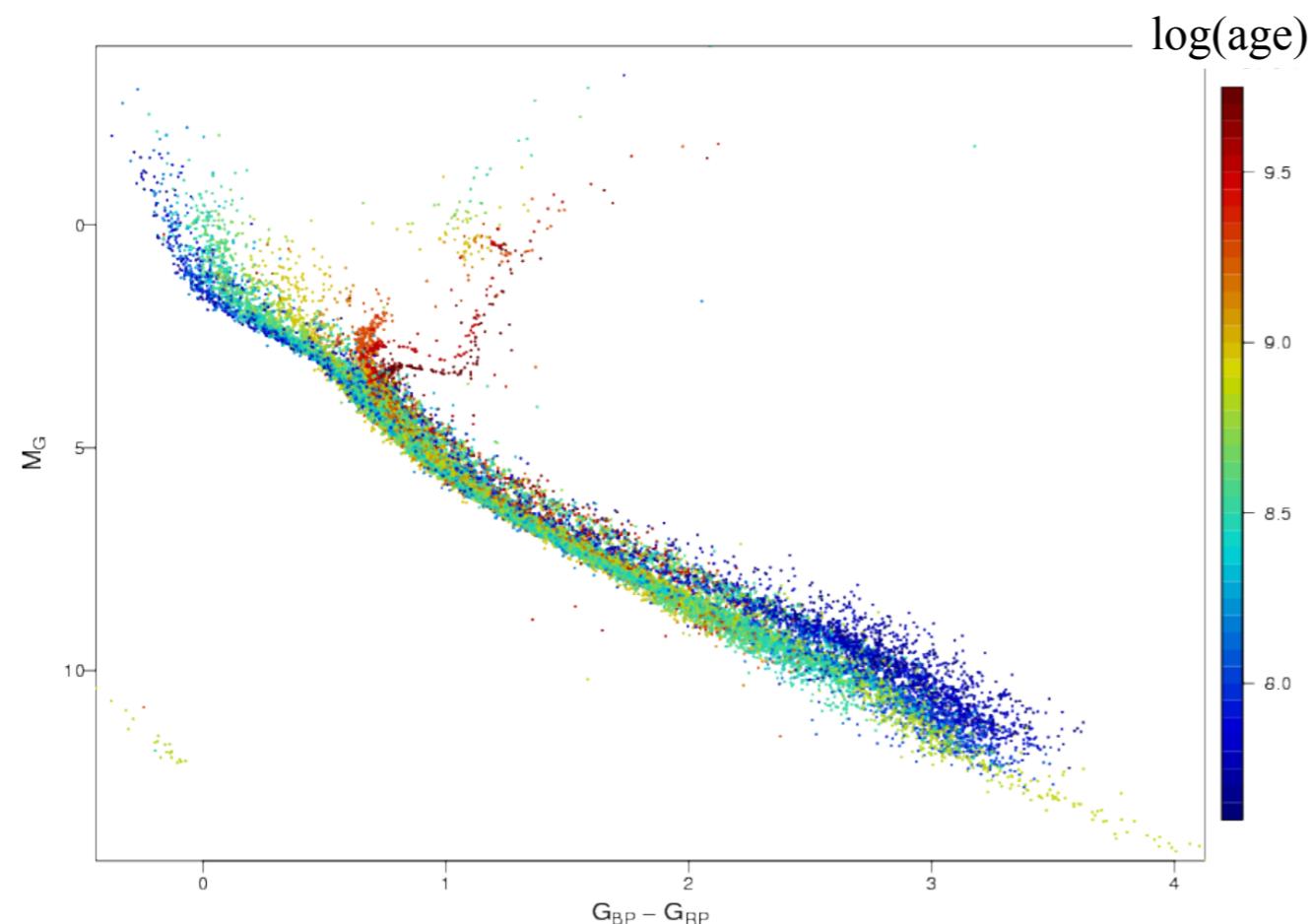
$\sigma_\pi/\pi < 10\%$ ,  $E(B-V) < 0.015$

~ 4 000 000 stars, < 2 kpc

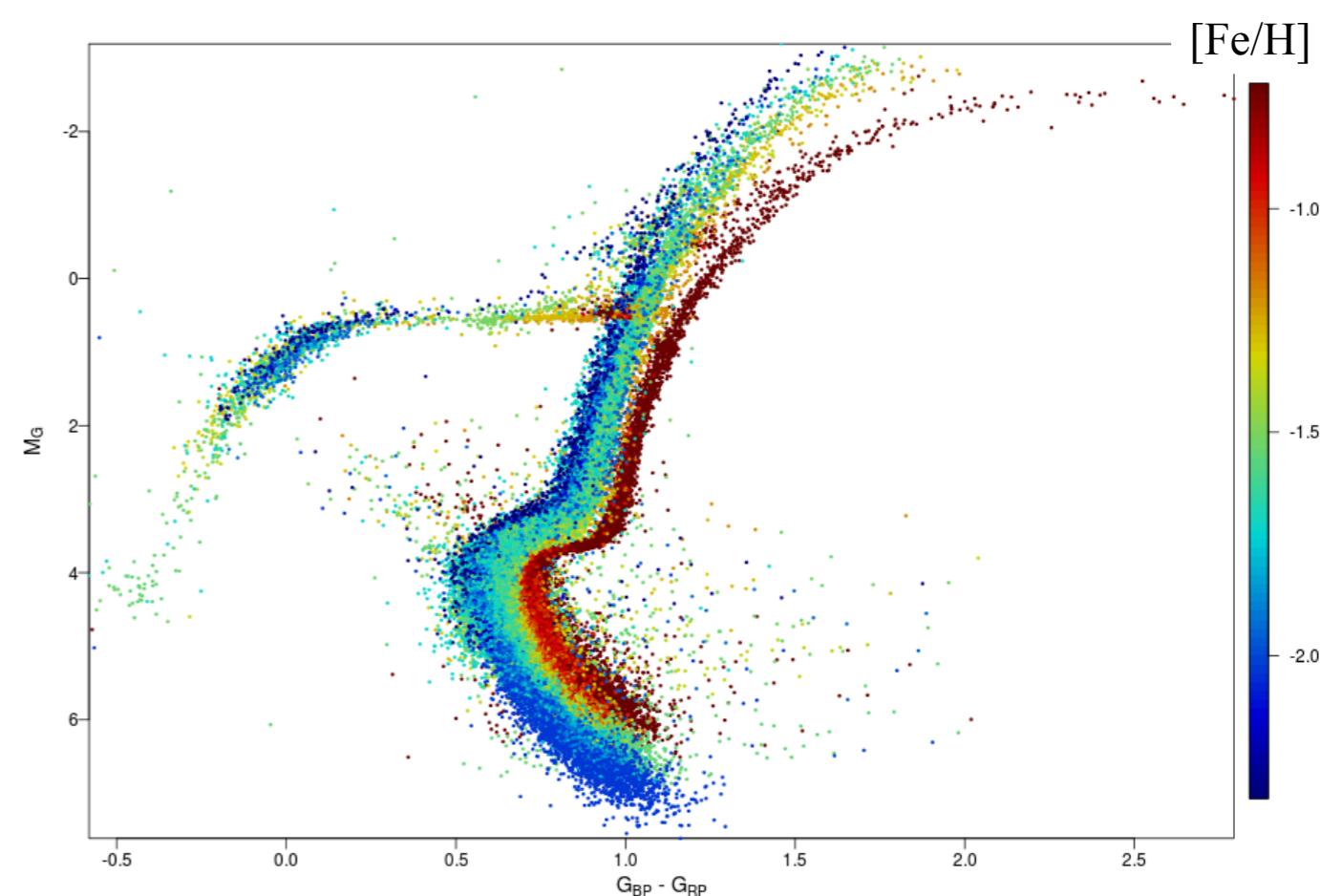


# Clusters → empirical isochrones

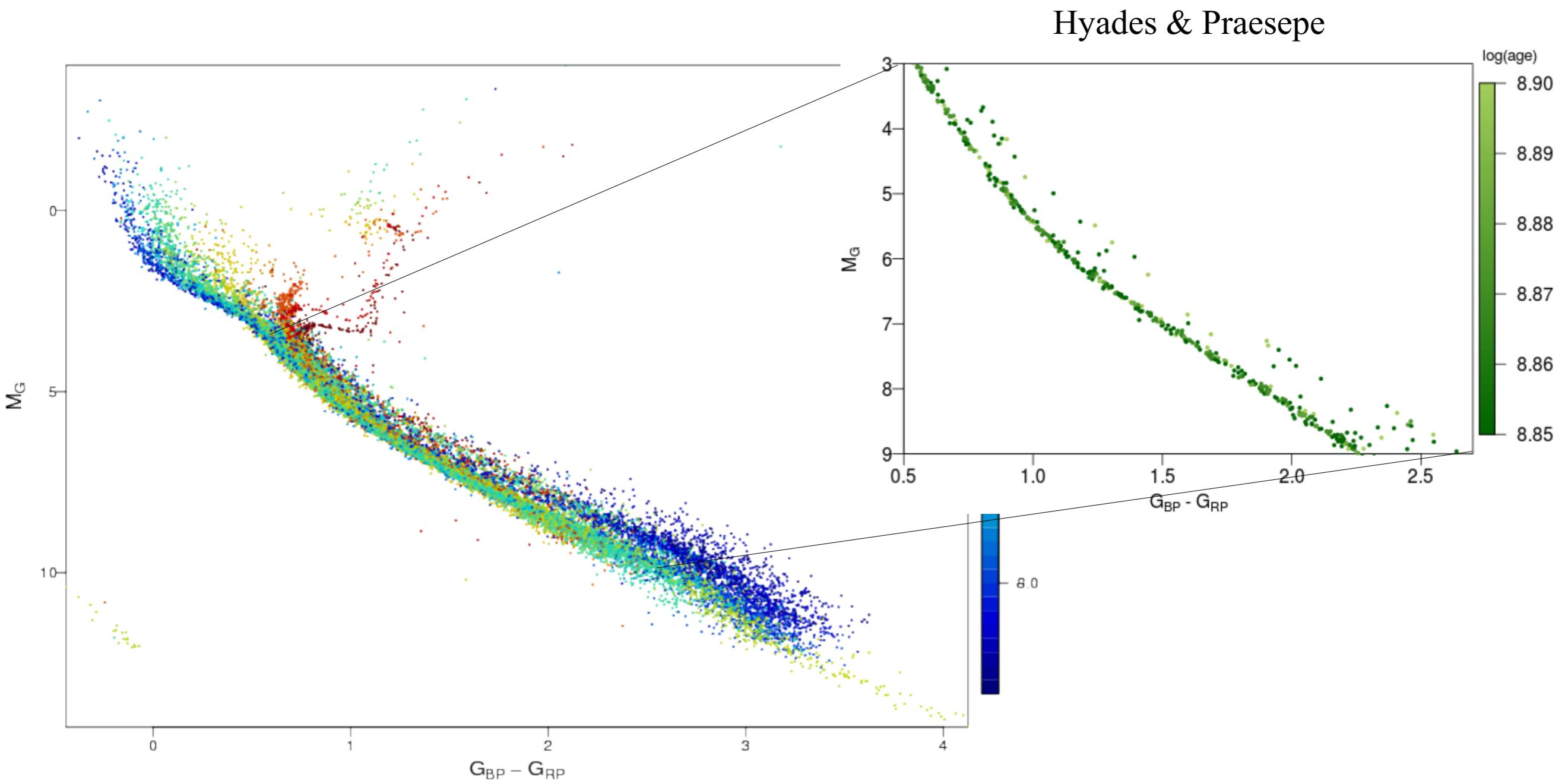
32 open clusters



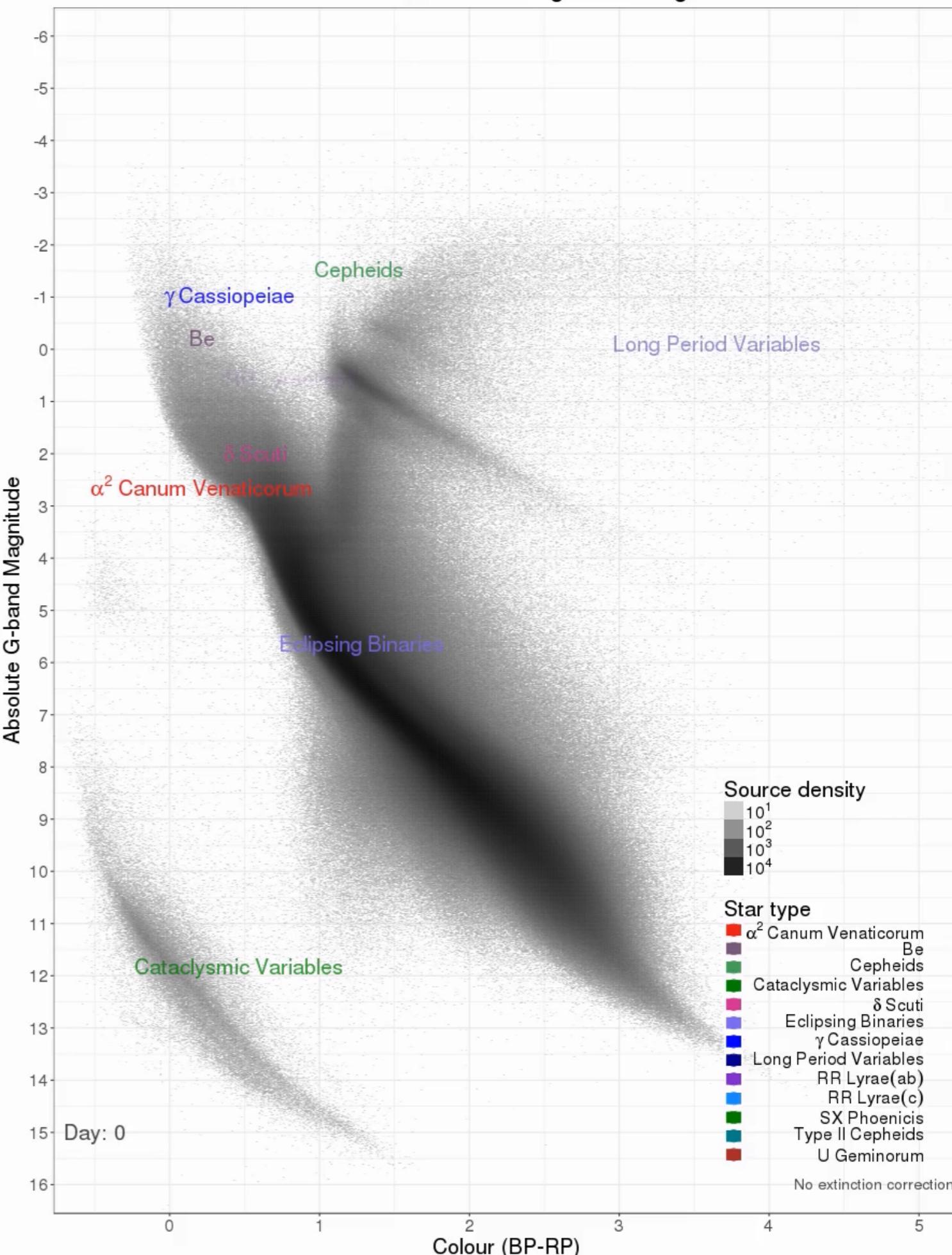
14 globular clusters



# Clusters → empirical isochrones



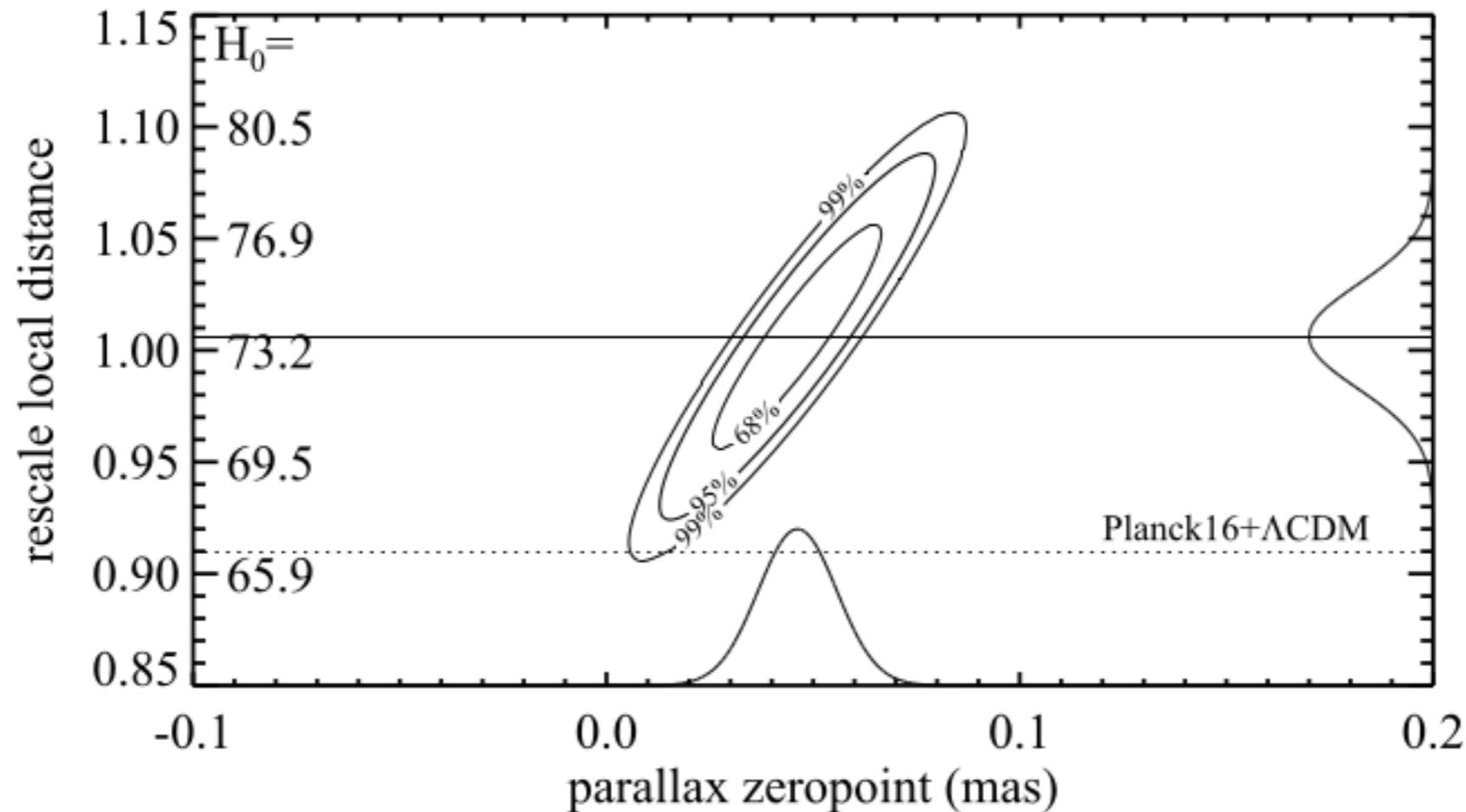
# Variability in the HRD



Gaia Collaboration, Eyer et al. 2018

crédit ESA/Gaia/DPAC/CU7

# Cepheids distance scale

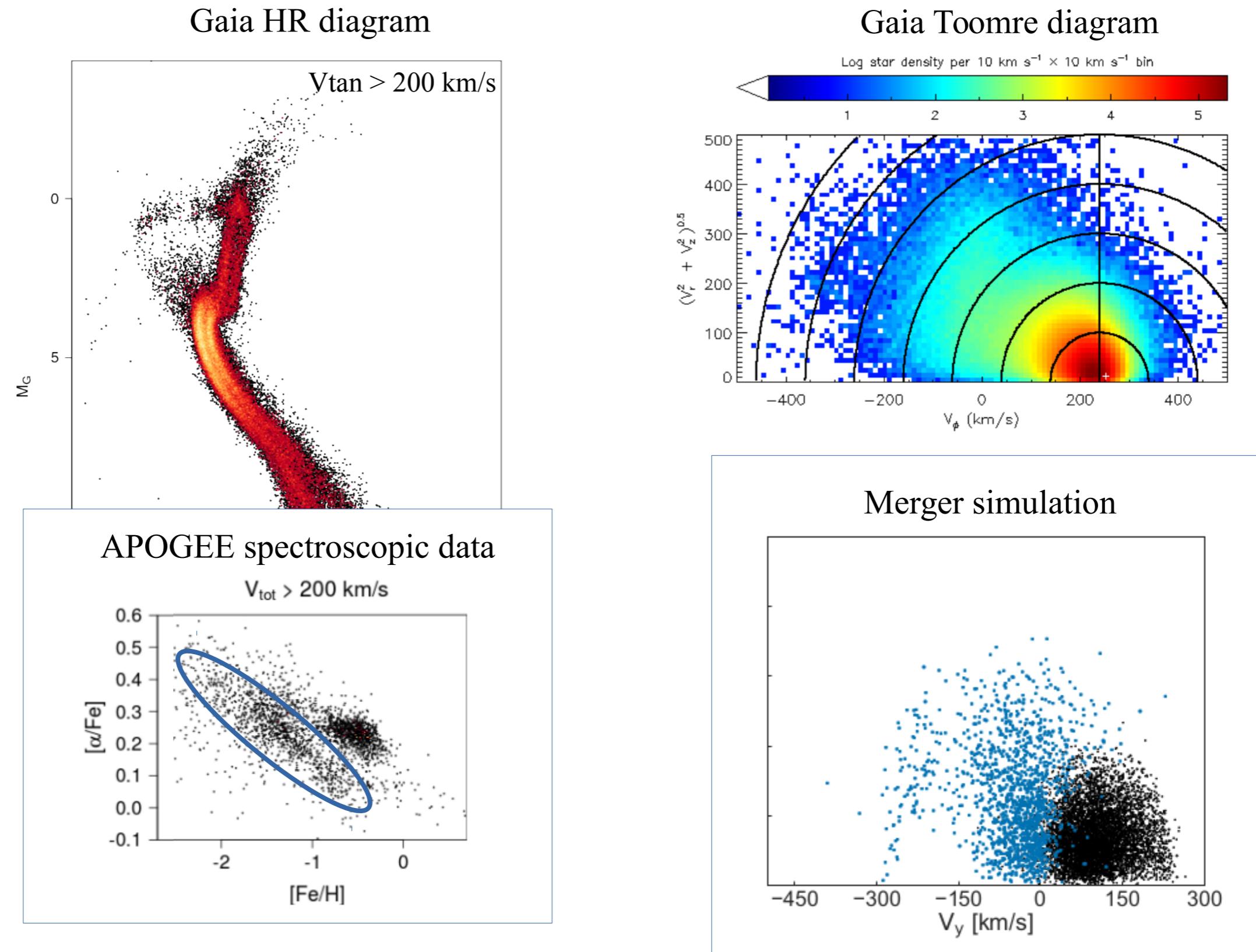


*Reiss et al. 2018*

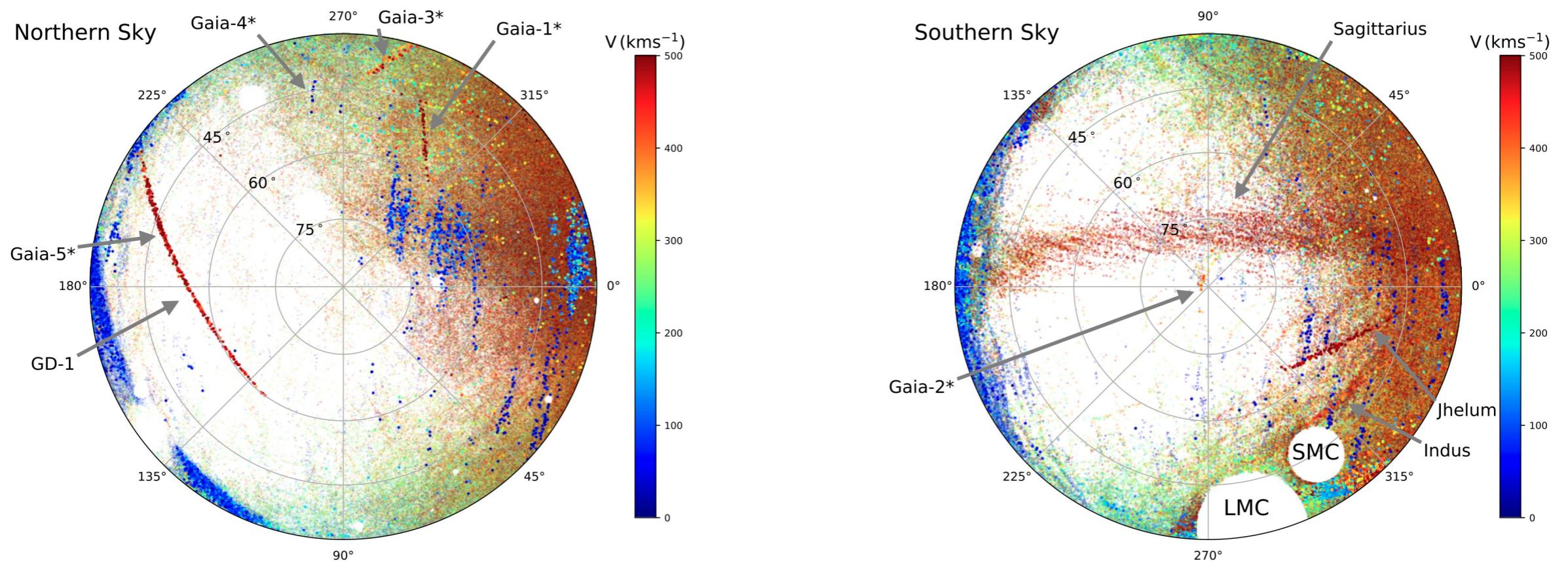
See also Shanks et al. 2018 on this  $H_0$  debate

+ Groenewegen et al. 2018 on the Cepheids distance scale with Gaia DR2

# Gaia-Enceladus the last big merger shaping both the halo and the thick disk?



# Stellar streams in the halo

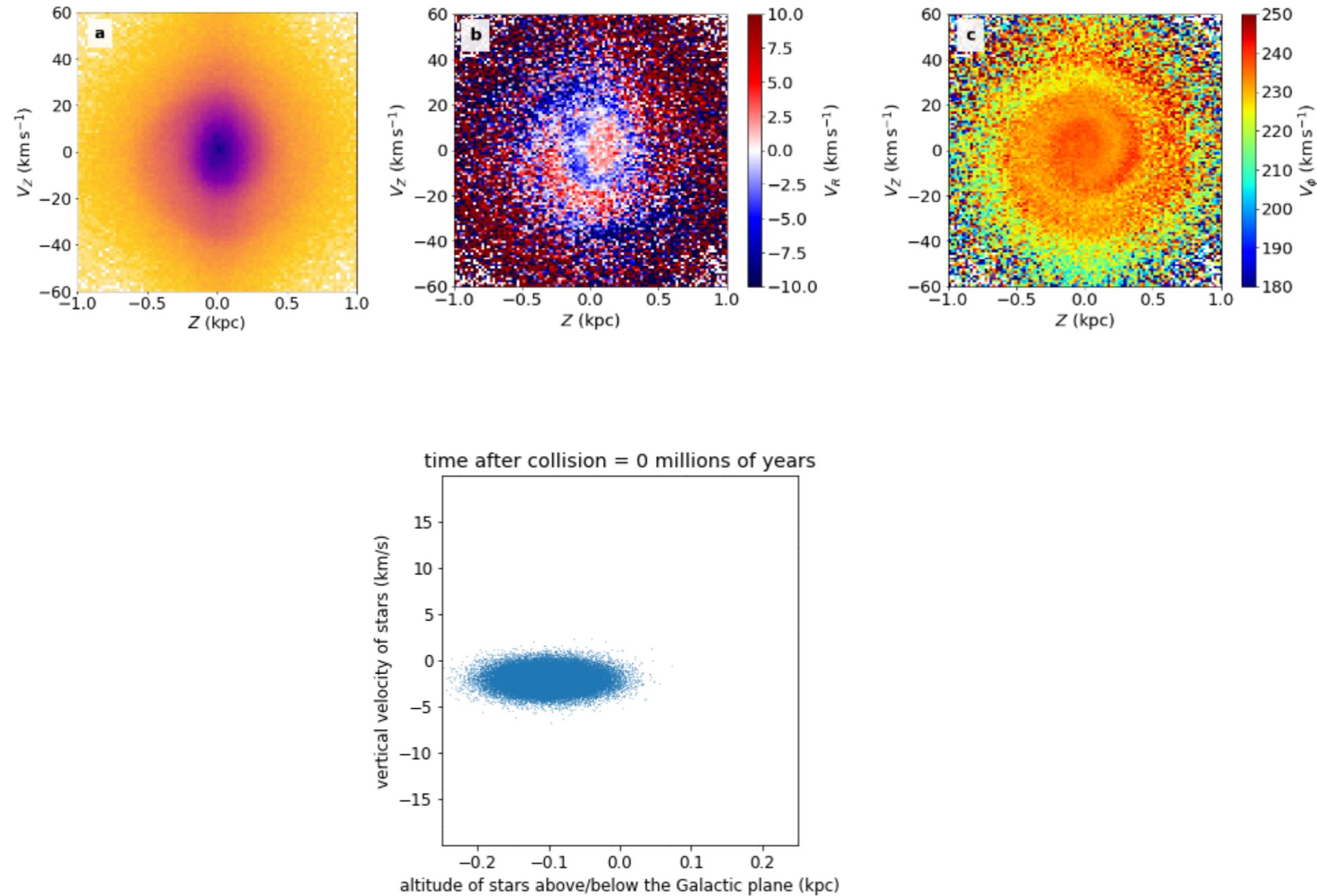


Malhan, Ibata, Martin 2018

Evans *et al.* : implications of Gaia-Enceladus / Sausage for the DM

Necib *et al.* :  $42 \pm 25\%$  of the accreted DM is in kinematic substructure

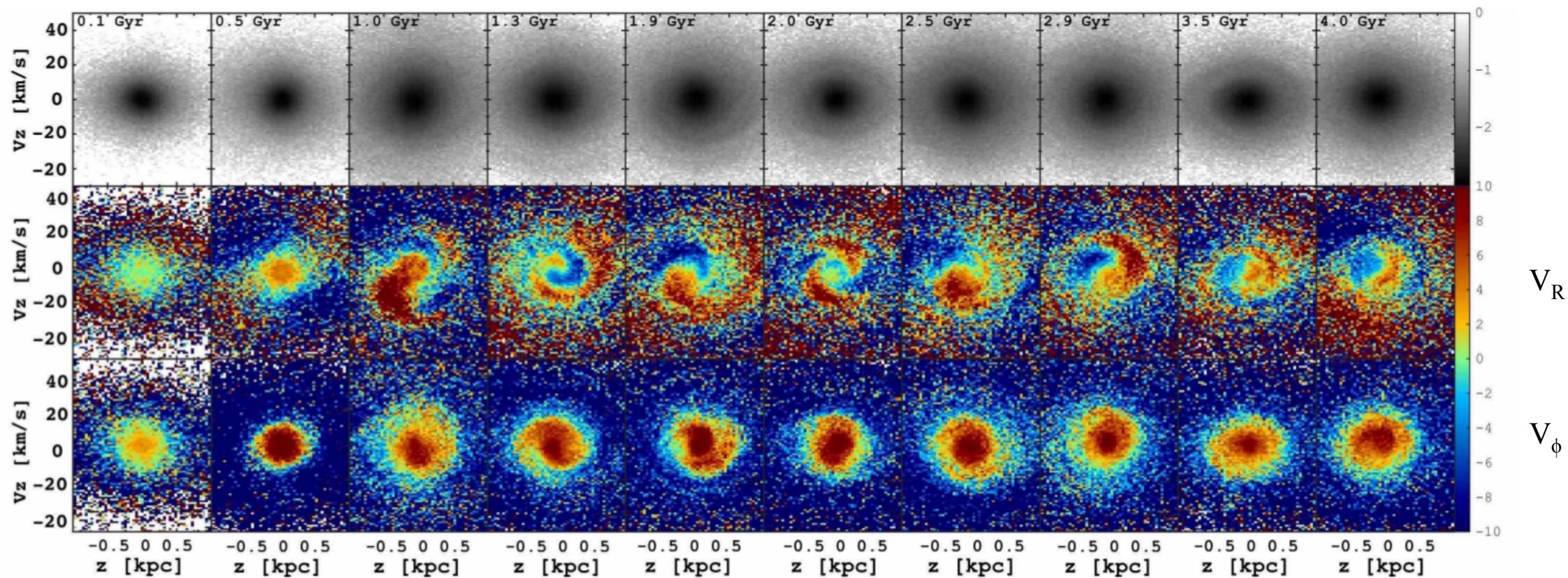
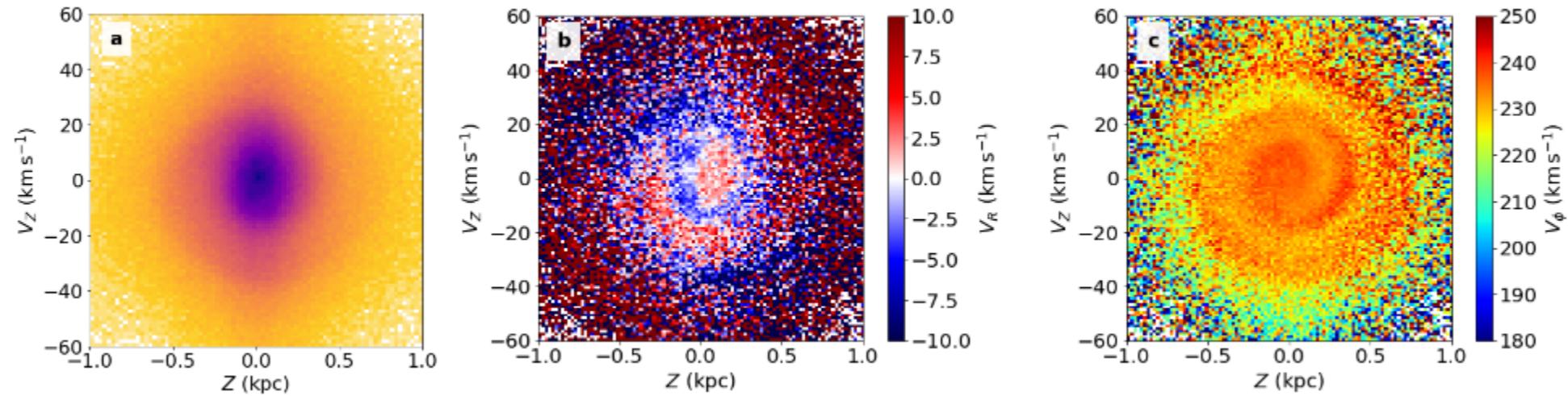
# The perturbed Milky Way disk



Perturbation by the Sgr dwarf galaxy

*Antoja et al. 2018*

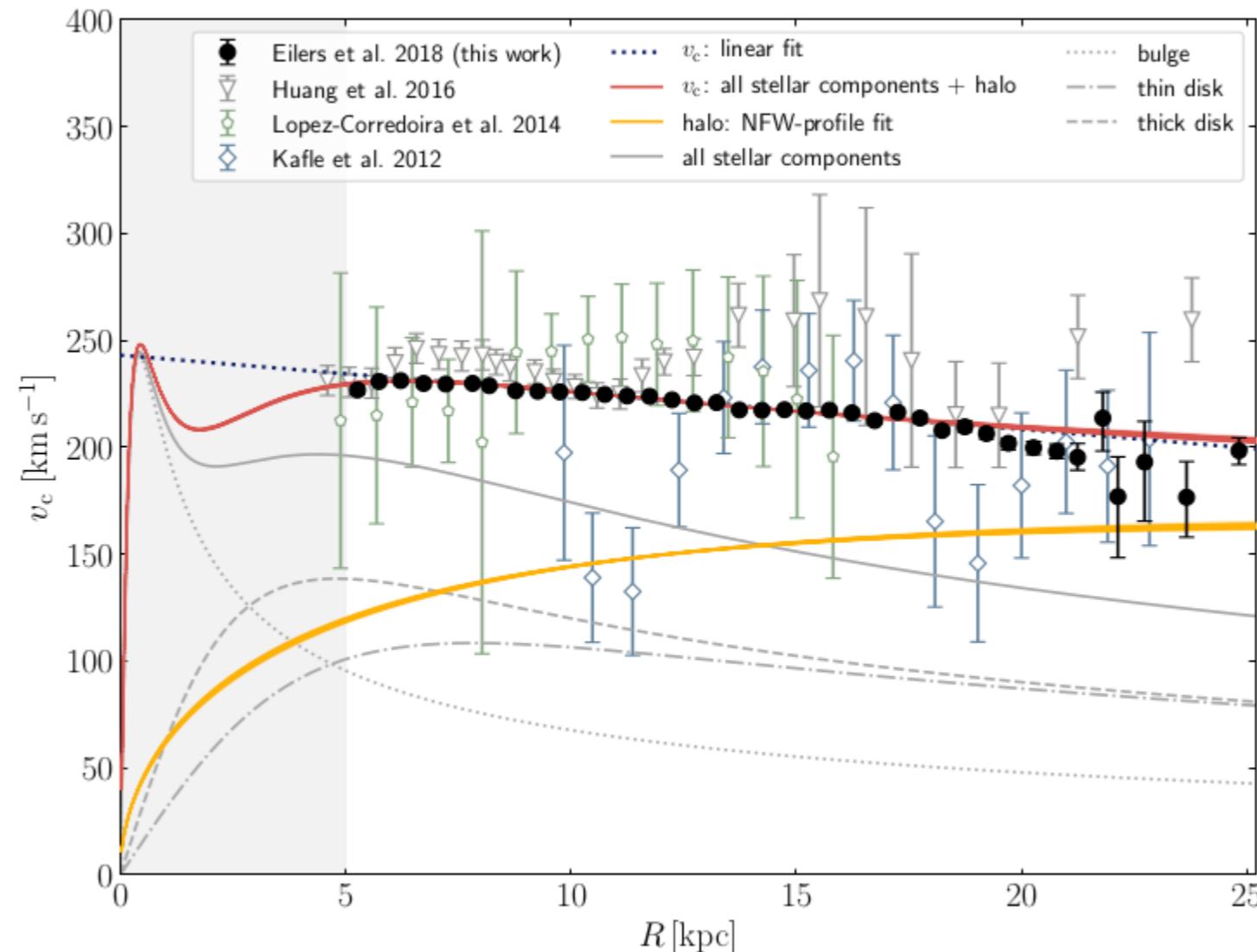
# The perturbed Milky Way disk



Perturbation by the bar buckling  
Khoperskov et al. 2018

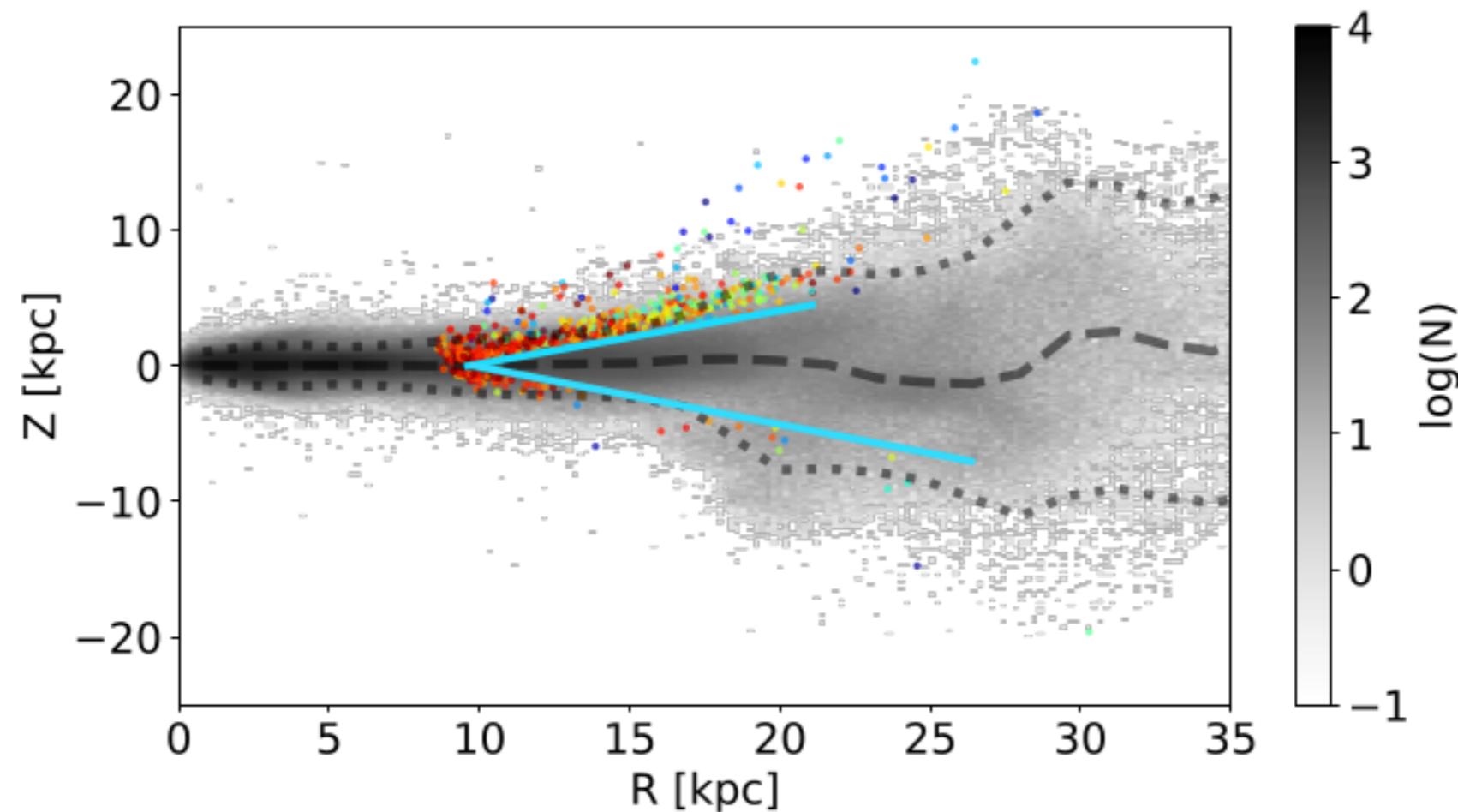
long-lived phase-space spirals

# The Milky Way circular velocity curve

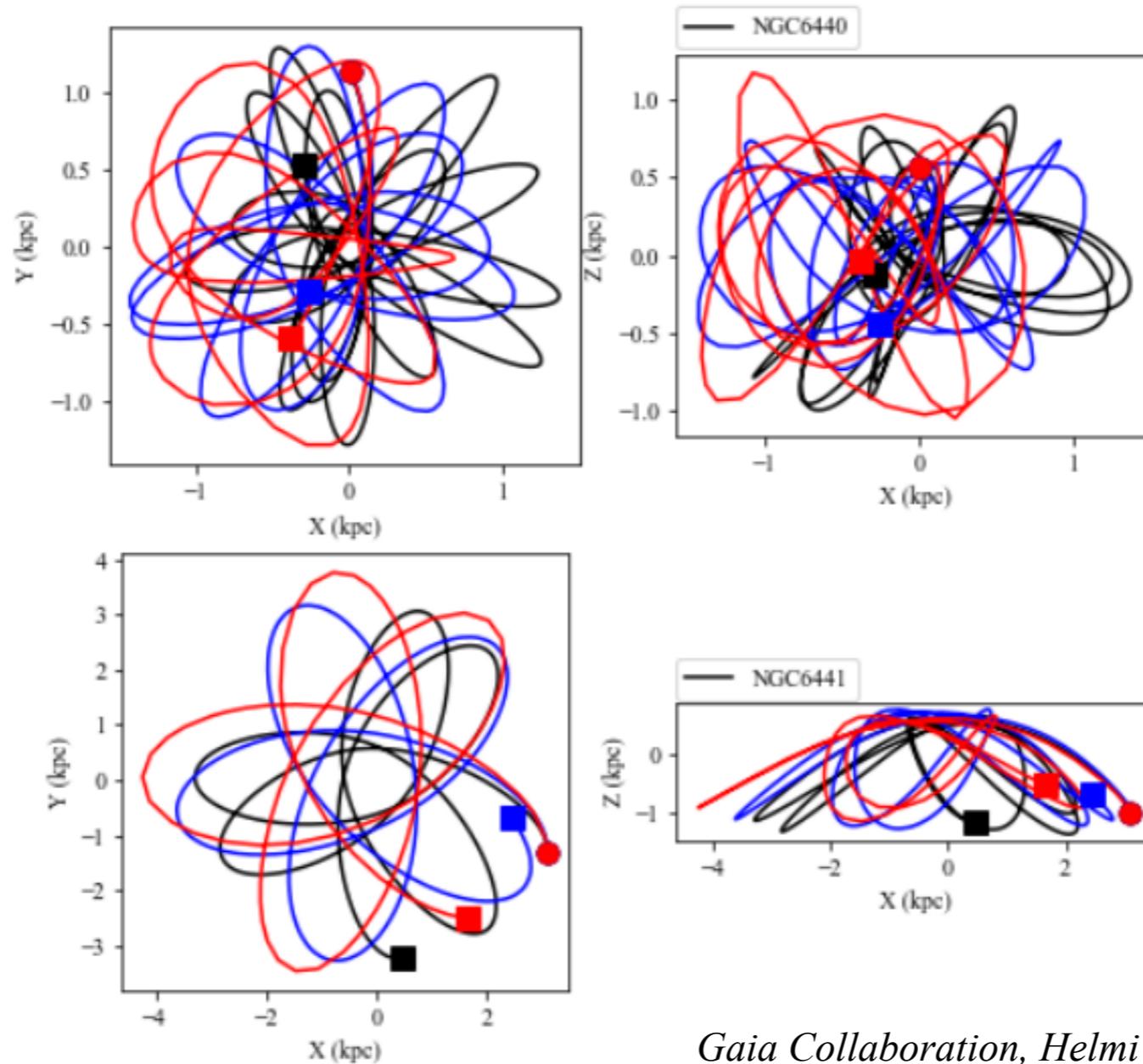


$$\rho_{\text{DM}}(R_\odot) = 0.30 \pm 0.03 \text{ GeV cm}^{-3}$$

# The disc flare due to the Sgr dSph repeated passage ?



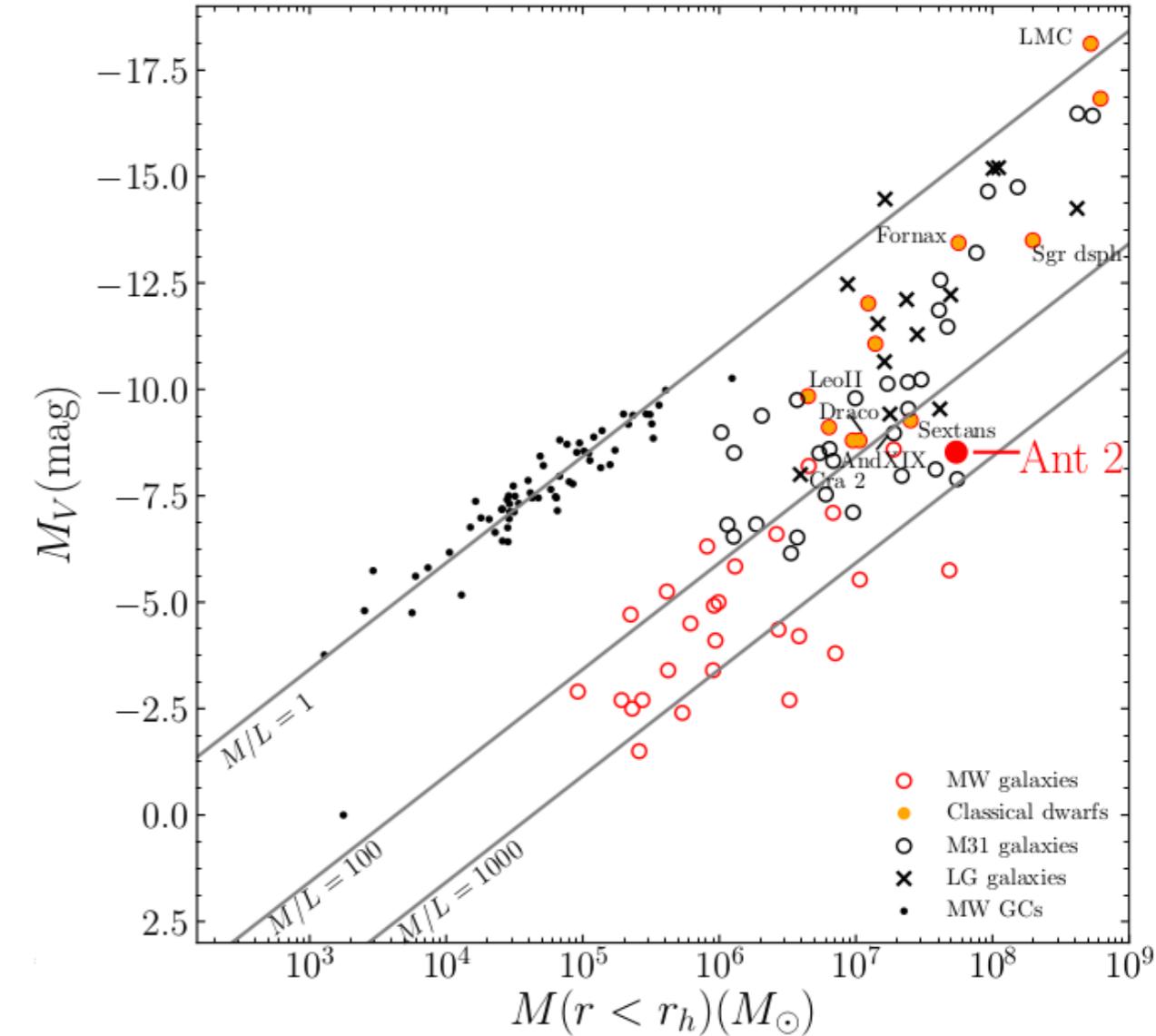
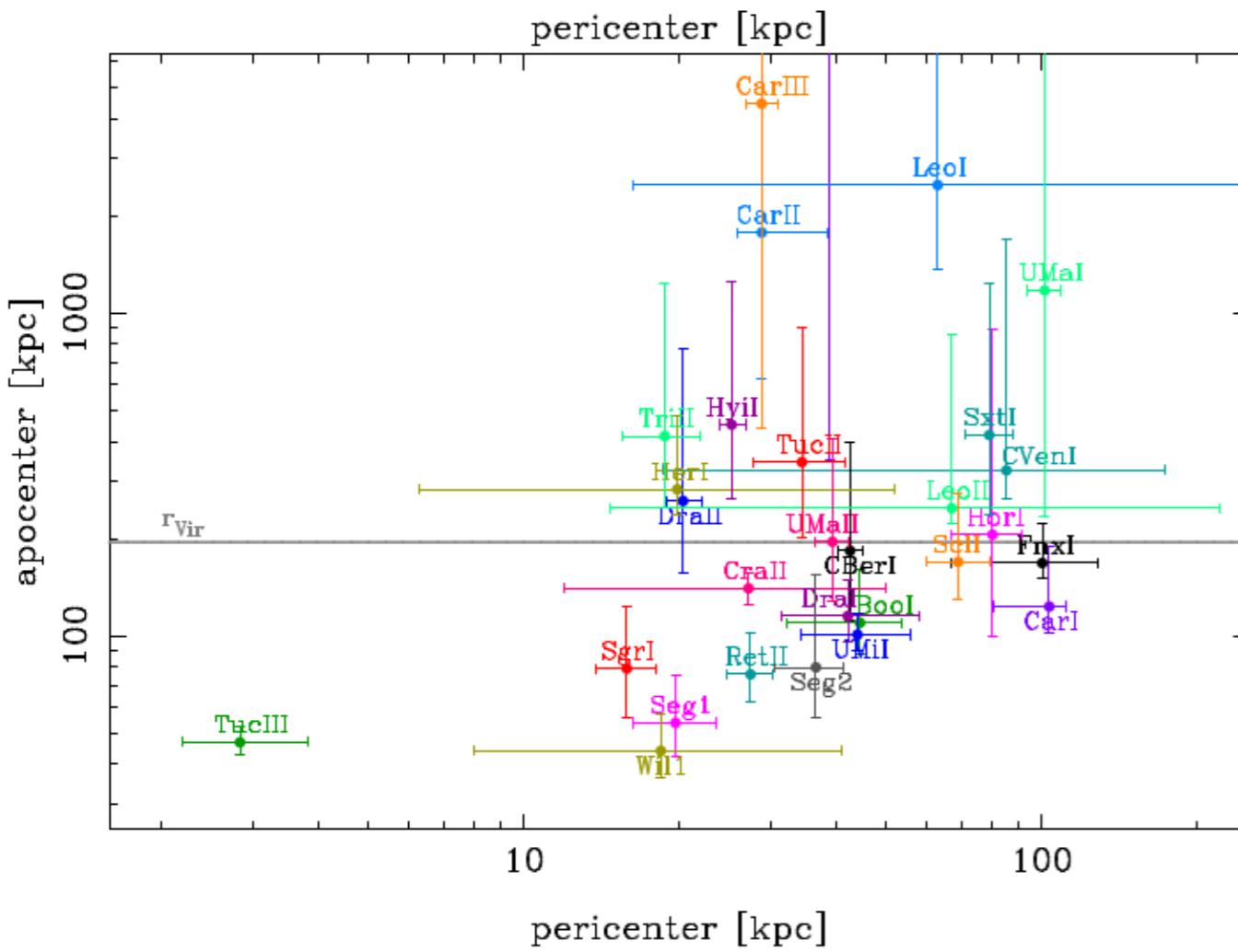
# Globular Clusters



*Gaia Collaboration, Helmi et al. 2018*

Used to derive the mass of the MW + DM:  
Watkins et al. 2018, Posti & Helmi 2018

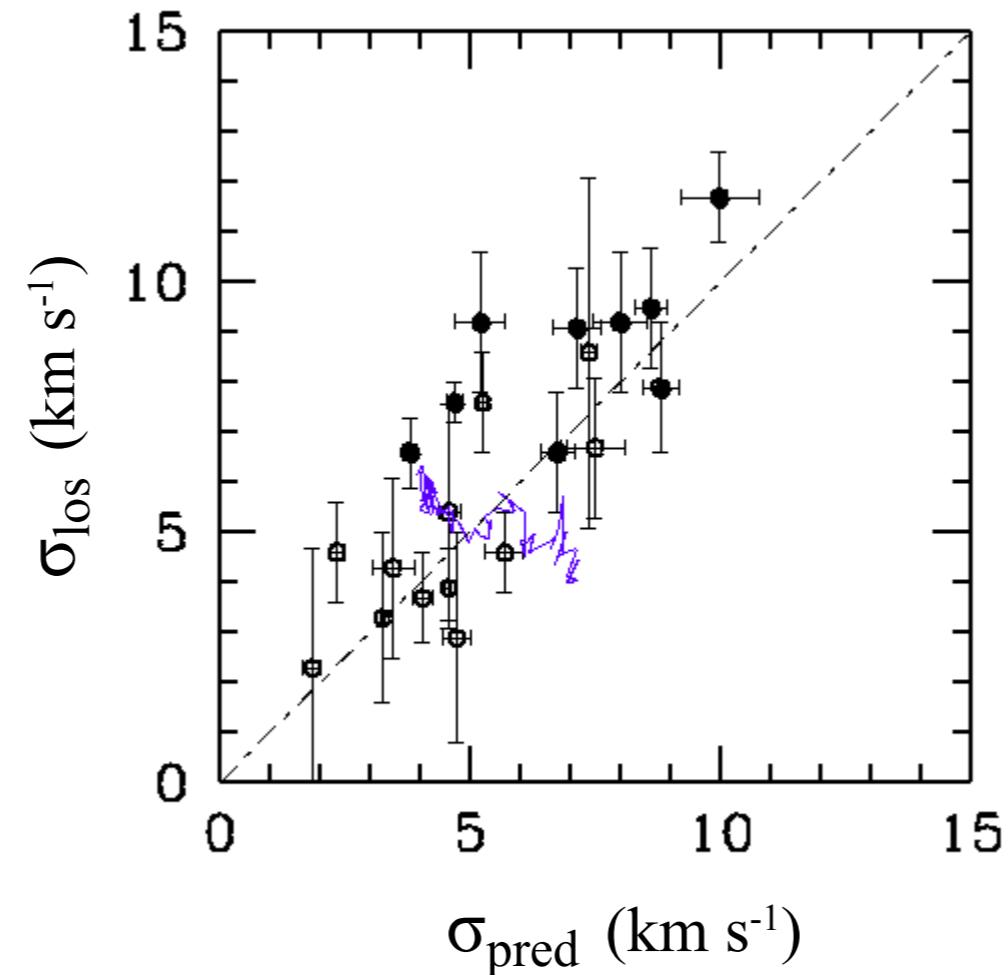
# Dwarf spheroidals



Orbit determinations :  
*Gaia Collaboration, Helmi et al. 2018*  
*Fritz et al. 2018*

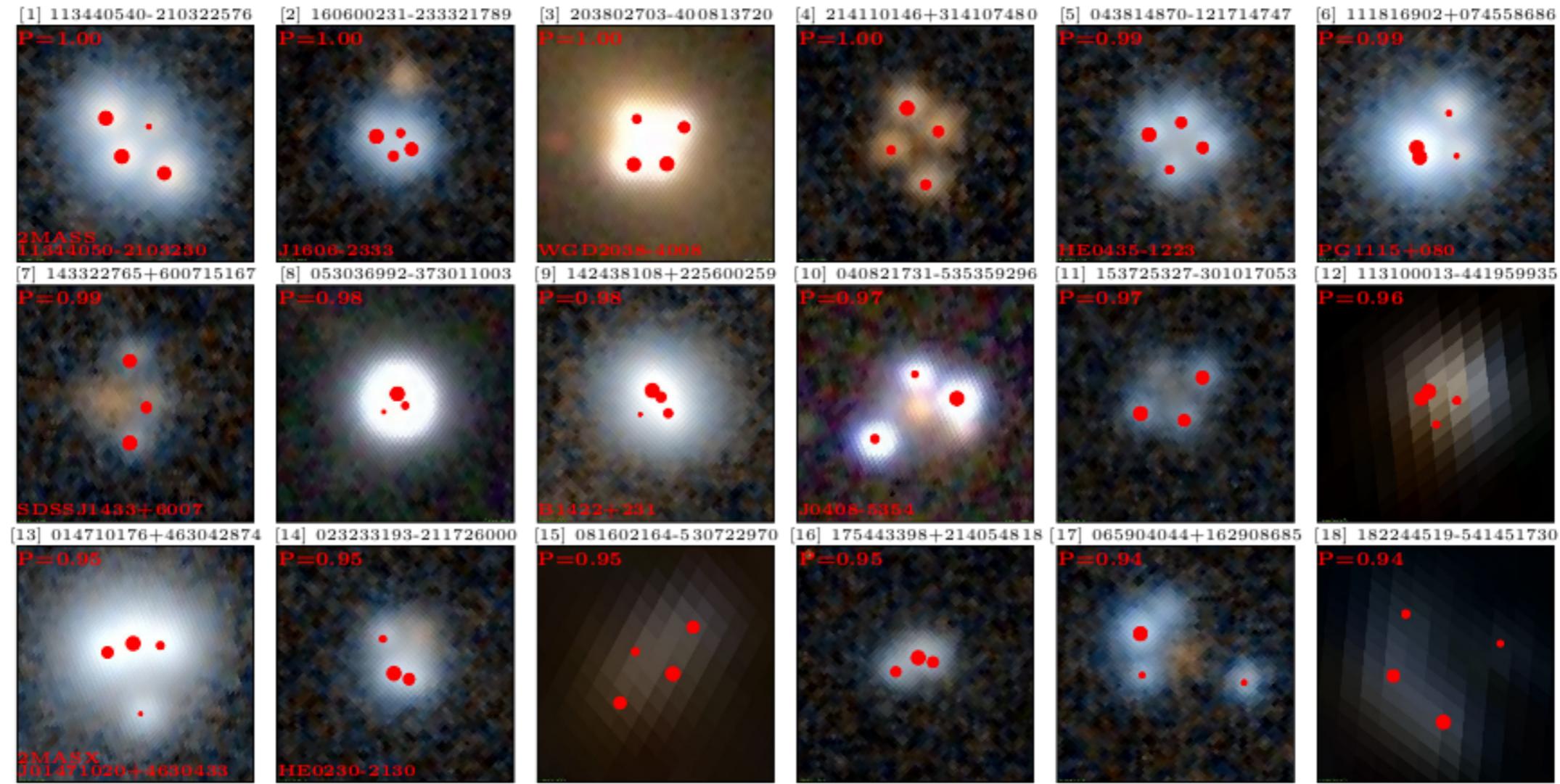
*Torrealba et al. 2018*  
Discovery of Antlia 2

# No more need for DM in dSph ?



Velocity dispersion predicted by a model where dSph are at their first passage

# Gravitational lenses



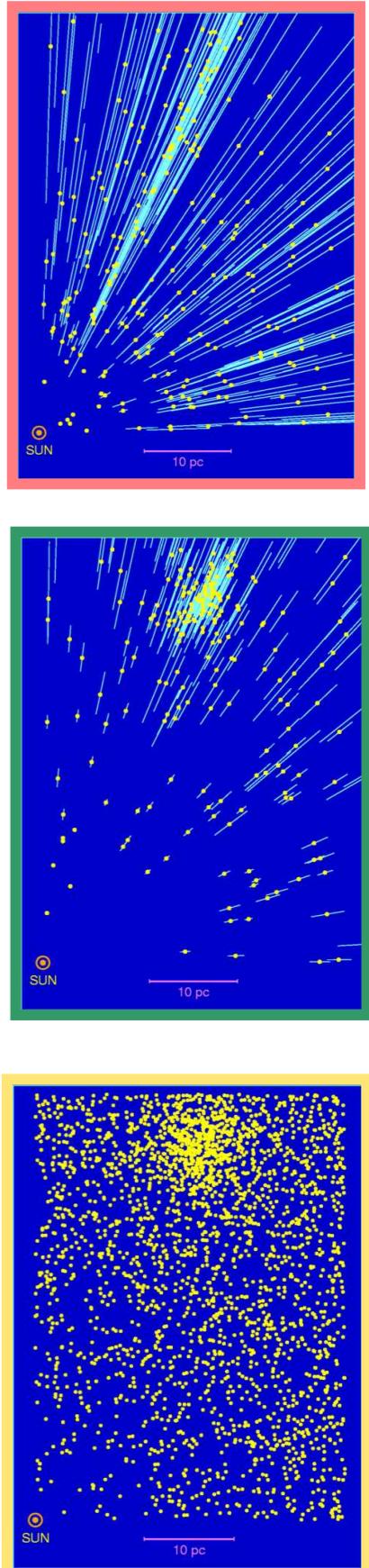
Search for gravitational lens systems in Gaia DR2

*Delchambre et al. 2018*

# More than yesterday, less then tomorrow...

Hyades

	<b>DR2</b> <b>(25 April 2018)</b>	<b>DR3</b> <b>(2021)</b>	<b>DR4</b> <b>(TBD)</b>
Parallaxes and proper motions	Full Sample	++	++
Photometry	G, G <sub>BP</sub> , G <sub>RP</sub>	++	++
Variables	550 000	++	++
Radial velocities	RVs at G <sub>RVS</sub> <12	++	++
SSOs	pre-selected asteroids	New SSOs	++
Astrophysical parameters	for G < 17 : Teff, A <sub>G</sub> Radii and luminosities <i>from integrated phot</i>	Classification + parameters from BP/RP + RVS spectra	++
Systems	-	Non-single catalogue Extended Objects	Exoplanet list
Spectra	-	Mean BP/RP spectra Mean RVS spectra	++
Epoch data	-	-	All epoch data



1960

1990 (Hipparcos)

2022 (Gaia)