

Emergence of Flow in Relativistic Heavy Ion Collisions

François Gelis

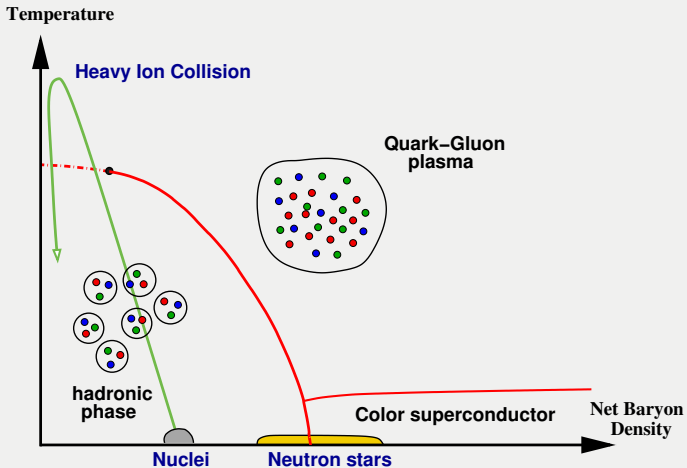
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Heavy Ion Collisions

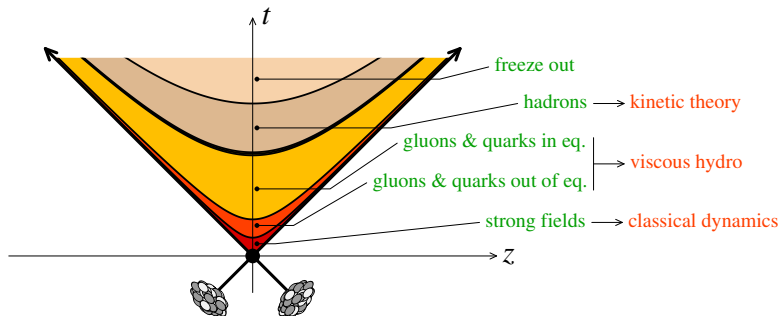
HEAVY ION COLLISIONS



EXPERIMENTAL FACILITIES : RHIC AND LHC



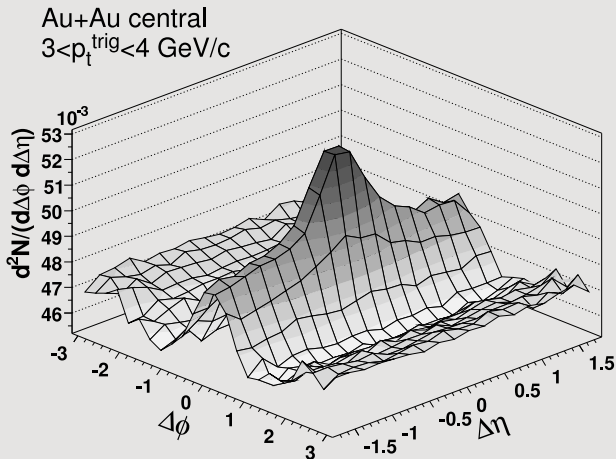
STAGES OF A NUCLEUS-NUCLEUS COLLISION



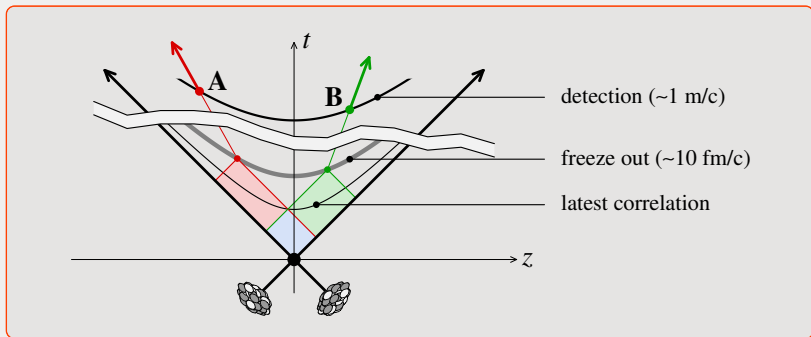
- Hydrodynamics successful at describing the bulk evolution
- In this talk : **Pre-hydrodynamical evolution**

EVIDENCE FOR HYDRODYNAMICAL EXPANSION (ONE OUT OF MANY...)

Two-particle correlations

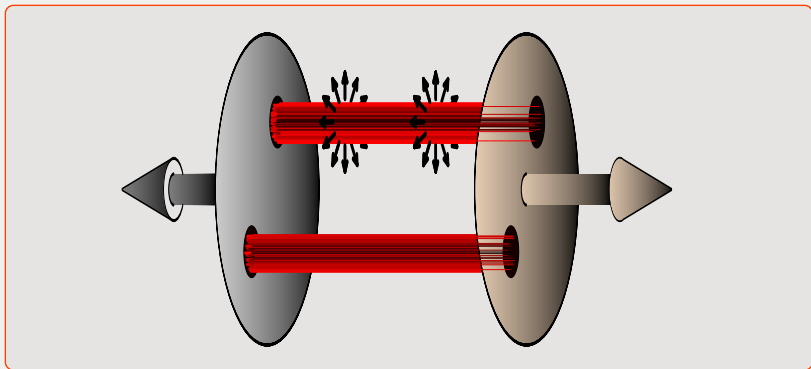


EARLY ORIGIN OF CORRELATIONS IN RAPIDITY



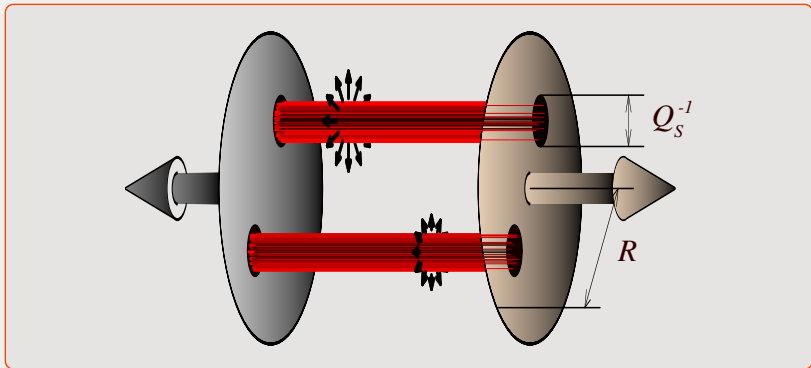
- $t_{\text{correlation}} \lesssim t_{\text{freeze out}} \times e^{-\frac{1}{2} |\Delta\eta|}$
- Correlations in azimuthal angle may be produced much later

INTERPRETATION IN TERMS OF FLOW



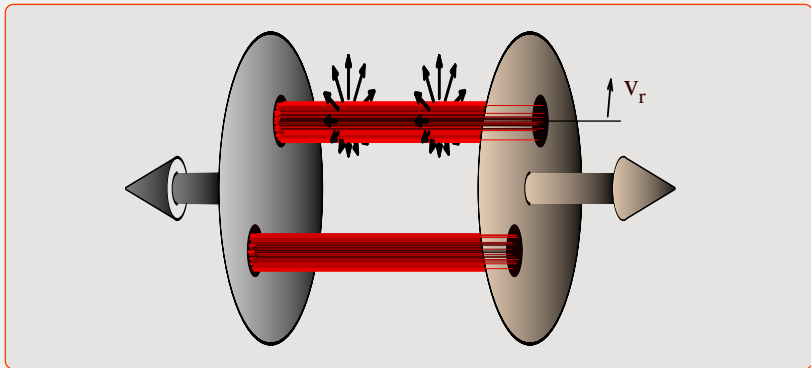
- Post-collision color fields organized in “flux tubes”
- Coherent production in a each tube, Isotropic on average

INTERPRETATION IN TERMS OF FLOW



- Post-collision color fields organized in “flux tubes”
- Coherent production in a each tube, Isotropic on average
- No correlations between different tubes

INTERPRETATION IN TERMS OF FLOW



- Post-collision color fields organized in “flux tubes”
- Coherent production in a each tube, Isotropic on average
- No correlations between different tubes
- Radial collective expansion \Rightarrow angular collimation

Equations of hydrodynamics (conservation laws)

$$\partial_{\mu} T^{\mu\nu} = 0 \quad , \quad \partial_{\mu} J_B^{\mu} = 0$$

Assumptions and inputs

i. Near equilibrium form of $T^{\mu\nu}$:

$$T^{\mu\nu} = \underbrace{(p + \epsilon) v^{\mu} v^{\nu} - p g^{\mu\nu}}_{\text{ideal hydro}} \oplus \underbrace{(\eta, \zeta) \partial v}_{\text{viscous terms}} \oplus \dots$$

ii. Equation of State: $p = f(\epsilon)$

iii. Transport coefficients: η, ζ, \dots

iv. Initial condition

Equations of hydrodynamics (conservation laws)

What makes it work so well?

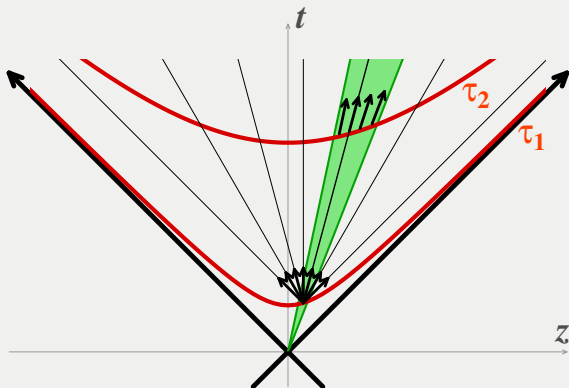
- Pressure tensor not too far from isotropy
- System not too far from equilibrium
- Low viscosity (compared to entropy density)

Not easy to get in Quantum Chromodynamics...

- ii. Equation of State: $p = f(\epsilon)$
- iii. Transport coefficients: η, ζ, \dots
- iv. Initial condition

COMPETITION BETWEEN EXPANSION AND INTERACTIONS

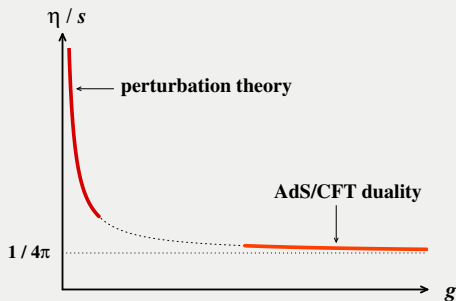
- Very different from isotropization in a box
- Sustained interactions are needed for isotropy to persist despite the expansion

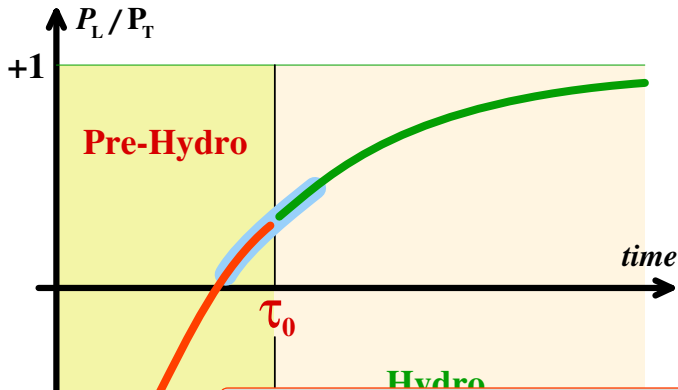


SHEAR VISCOSITY AT WEAK AND STRONG COUPLING (IN EQUILIBRIUM)

Weak coupling result [Arnold, Moore, Yaffe (2000)]

$$\frac{\eta}{s} \approx \frac{5.1}{g^4 \ln\left(\frac{2.4}{g}\right)}$$

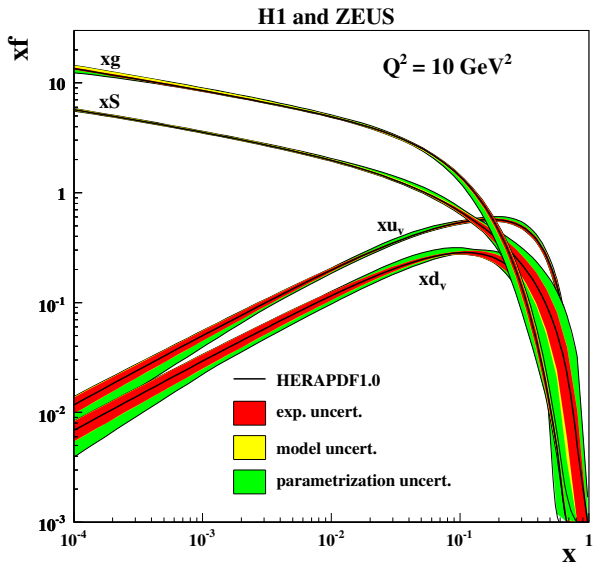




GOAL : smooth matching to Hydrodynamics

- The pre-hydro model should bring the system to a situation that hydrodynamics can handle
- Pre-hydro and hydro should agree over some range of time \Rightarrow no τ_0 dependence
- Description as close as possible to QCD

QCD description

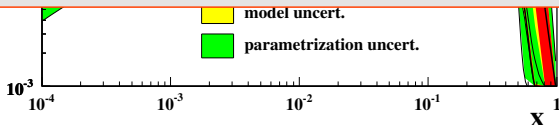
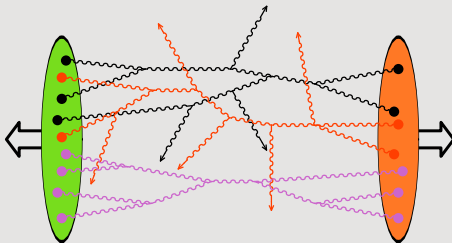


PARTON DISTRIBUTIONS IN A NUCLEON: $x \sim p_T/E_{\text{coll}}$

H1 and ZEUS

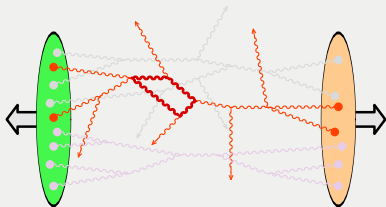


$x \sim p_T/E_{\text{coll}} \ll 1$: dense, multi-parton interactions



SCALE SEPARATION IN THE DENSE REGIME

$$\mathcal{S} = \underbrace{-\frac{1}{4} \int F_{\mu\nu} F^{\mu\nu}}_{\text{slow gluons}} + \underbrace{\int (J_1^\mu + J_2^\mu) A_\mu}_{\text{fast sources}}$$



In the dense regime: $J^\mu \sim g^{-1}$, $A^\mu \sim g^{-1}$, $f_k \sim g^{-2}$

$$T^{\mu\nu} \sim \frac{1}{g^2} \left[c_0 + c_1 g^2 + c_2 g^4 + \dots \right]$$

$c_0 \equiv$ tree level, $c_1 \equiv$ one loop, etc...

SHEAR VISCOSITY IN THE DENSE REGIME

From kinetic theory :

$$\frac{\eta}{s} \sim \frac{\text{mean free path}}{\text{de Broglie wavelength}}$$

- **(de Broglie wavelength)**⁻¹ $\sim Q$
- **(mean free path)**⁻¹ $\sim \underbrace{g^4 Q^{-2}}_{\text{cross section}} \times \underbrace{\int_{\mathbf{k}} f_{\mathbf{k}}}_{\text{density}} \underbrace{(1 + f_{\mathbf{k}})}_{\text{Bose enhancement}}$

If $g \ll 1$ but $f_{\mathbf{k}} \sim g^{-2}$ (**weakly coupled, but strongly interacting**)

$$\frac{\eta}{s} \sim g^0$$

LEADING ORDER

- Leading Order = sum of all tree diagrams

Expressible in terms of **classical solutions of Yang-Mills equations** :

$$[\mathcal{D}_\mu, \mathcal{F}^{\mu\nu}] = J_1^\nu + J_2^\nu$$

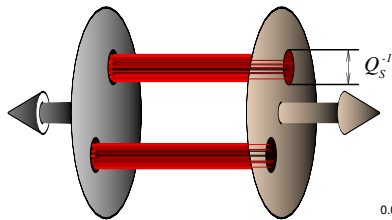
- Initial condition : $\lim_{x^0 \rightarrow -\infty} \mathcal{A}^\mu(x) = 0$

Components of the energy-momentum tensor

$$T_{LO}^{00} = \frac{1}{2} [\underbrace{\mathbf{E}^2 + \mathbf{B}^2}_{\text{class. fields}}] \quad T_{LO}^{0i} = [\mathbf{E} \times \mathbf{B}]^i$$

$$T_{LO}^{ij} = \frac{\delta^{ij}}{2} [\mathbf{E}^2 + \mathbf{B}^2] - [\mathbf{E}^i \mathbf{E}^j + \mathbf{B}^i \mathbf{B}^j]$$

LO : STRONG PRESSURE ANISOTROPY AT ALL TIMES

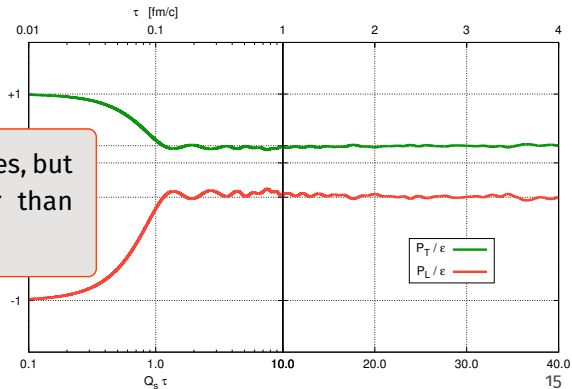


At $\tau = 0^+$

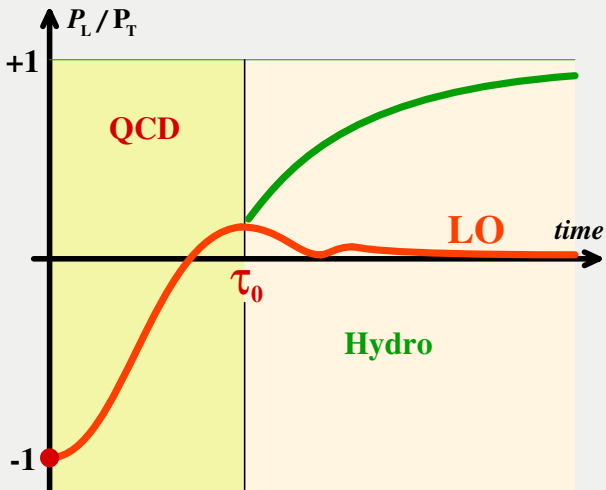
$$\mathbf{E} \parallel \mathbf{B} \parallel \hat{z}$$

$$P_T = \epsilon, P_L = -\epsilon$$

P_L rises to positive values, but remains much smaller than P_T (free streaming)

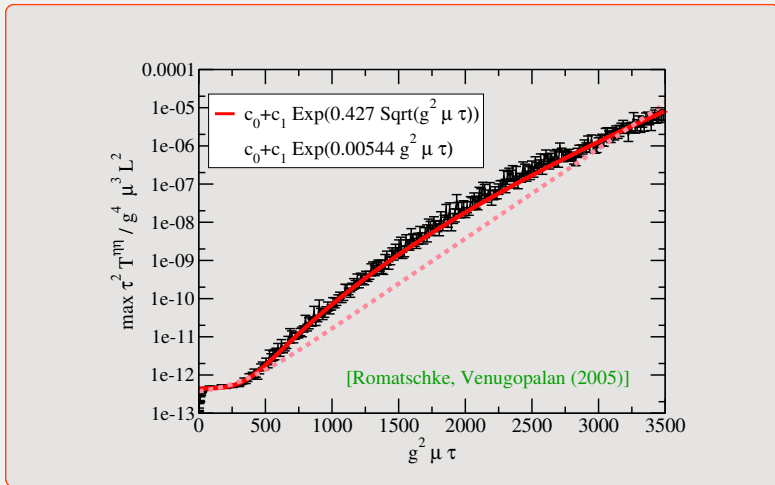


LO : UNSATISFACTORY MATCHING TO HYDRODYNAMICS

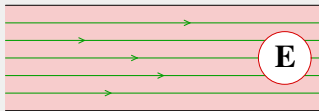


Instabilities, Resummation

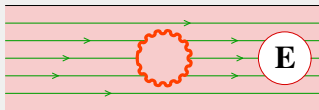
INSTABILITY OF CLASSICAL SOLUTIONS



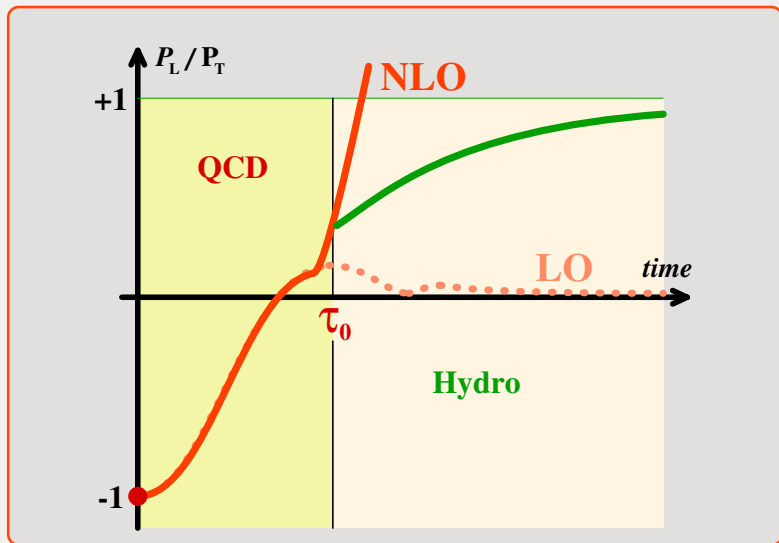
- LO = longitudinal chromo-E and chromo-B fields



- NLO = gluon loop embedded in this field



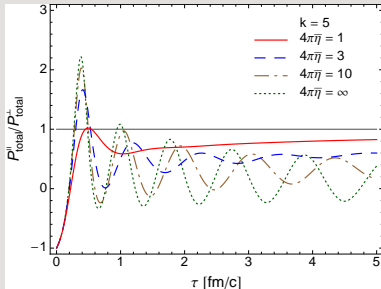
- instability \sim imaginary part of the loop \sim gluon pair production
- BUT : at NLO, no feedback of the produced gluons on the LO field!



Color flux tube model : [Ryblewski, Florkowski (2013)]

$$\underbrace{(p^\mu \partial_\mu + g F^{\mu\nu} p_\nu \partial_p^\mu)}_{\text{Lorentz force}} G = \underbrace{\frac{dN}{d\Gamma}}_{\text{Schwinger}} + \underbrace{C_p[G]}_{\text{collisions}}$$

$$\partial_\mu F^{\mu\nu} = j^\nu \quad (\text{feedback})$$



- Field converted into particles by instability
- Nearly constant P_L/P_T
- Ratio depends on $\bar{\eta} \equiv \eta/s$

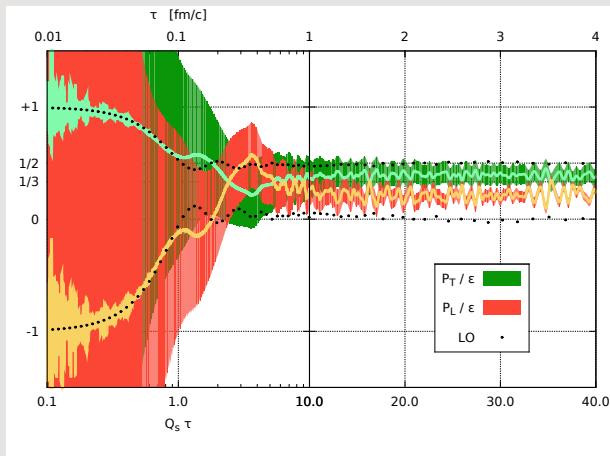
QCD EVOLUTION AT NLO + RESUMMATION OF SECULAR TERMS

$$\begin{aligned}\mathcal{O}_{\text{secular}} &\equiv \underbrace{\mathcal{O}_{\text{LO}} + \mathcal{O}_{\text{NLO}} + \text{subset of higher orders}}_{\text{terms that have the fastest growth in time}} \\ &= \int [\mathbf{D}\boldsymbol{\alpha}(\mathbf{u})] \exp \left[-\frac{1}{2\hbar} \int_{\mathbf{u},\mathbf{v}} \boldsymbol{\alpha}(\mathbf{u}) \boldsymbol{\sigma}^{-1}(\mathbf{u},\mathbf{v}) \boldsymbol{\alpha}(\mathbf{v}) \right] \mathcal{O}_{\text{LO}} [\mathcal{A}_{\text{ini}} + \boldsymbol{\alpha}]\end{aligned}$$

- In this resummation, the observable is as an average over classical field evolutions with fluctuating initial conditions
- Roughly speaking:
 - the secular resummation promotes a classical initial state to a quantum coherent state
 - fluctuations of $\mathcal{A}_{\text{in}} \sim \text{zero point fluctuations}$
- The precise form of the variance ($\hbar \boldsymbol{\sigma}$) is obtained from an NLO (analytical) calculation

QCD EVOLUTION AT NLO + RESUMMATION OF SECULAR TERMS

[Epelbaum, FG (2013)]



HOW CRUCIAL WAS THE INITIAL CONDITION ?

QCD: coherent initial state

$$A = \mathcal{A}_{LO} + \int_{\mathbf{p}} c_{\mathbf{p}} \alpha_{\mathbf{p}} \quad \langle c_{\mathbf{p}} c_{\mathbf{p}'} \rangle \sim \frac{1}{2} \delta_{\mathbf{p}\mathbf{p}'}$$

Occupation number :

$$\langle \tilde{A} \tilde{A}^* \rangle_{\tau=0^+} = \underbrace{\tilde{\mathcal{A}}_{LO} \tilde{\mathcal{A}}_{LO}^*}_{\sim \delta(\mathbf{p}_z) f(\mathbf{p}_\perp)} + \frac{1}{2}$$

HOW CRUCIAL WAS THE INITIAL CONDITION ?

Incoherent initial state:

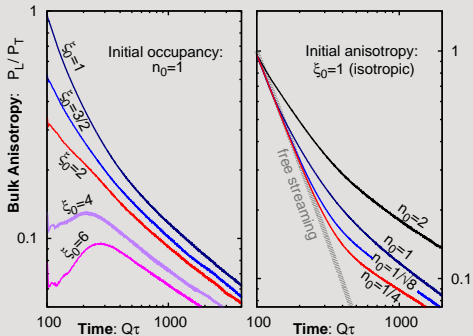
$$\bar{A} = \int_{\mathbf{p}} c_{\mathbf{p}} \alpha_{\mathbf{p}} \quad \langle c_{\mathbf{p}} c_{\mathbf{p}'} \rangle \sim \delta_{\mathbf{p}\mathbf{p}'} \left[\frac{1}{2} + f_0(\mathbf{p}) \right]$$

$\frac{1}{2}$ \iff zero point fluctuations

$f_0(\mathbf{p})$ \iff initial particle distribution ($\sim g^{-2}$)

If $f_0(\mathbf{p}) \gg 1$, approximate $\frac{1}{2} + f_0 \rightarrow f_0$?

How τ Incoherent initial state:



- No dependence on the coupling (can be scaled out)
- $P_T/P_L \sim \tau^{-2/3}$

Classical approximation in Kinetic Theory

Dyson-Schwinger
equations

→

Kadanoff-Baym
equations

→

Boltzmann :
 $p^\mu \partial_\mu f = C_p[f]$

- Collision term:

$$C_p[f] = \frac{i}{2} \left[\Sigma_{11}(p) + \left(\frac{1}{2} + f(p)\right) \left(\Sigma_{21}(p) - \Sigma_{12}(p) \right) \right]$$



$$\begin{aligned} \Rightarrow C_p[f] = & \frac{g^4}{4E_p} \int_{\mathbf{k}} \int_{\mathbf{p}'} \int_{\mathbf{k}'} (2\pi)^4 \delta(P + K - P' - K') \\ & \times \left[f(\mathbf{p}') f(\mathbf{k}') (1 + f(\mathbf{p})) (1 + f(\mathbf{k})) \right. \\ & \left. - f(\mathbf{p}) f(\mathbf{k}) (1 + f(\mathbf{p}')) (1 + f(\mathbf{k}')) \right] \end{aligned}$$

D Weak classical approximation:

$$C_p[f] = \frac{g^4}{4E_p} \int_{\mathbf{k}} \int_{\mathbf{p}'} \int_{\mathbf{k}'} (2\pi)^4 \delta(\mathbf{P} + \mathbf{K} - \mathbf{P}' - \mathbf{K}') \\ \times \left[\left(\frac{1}{2} + f(\mathbf{p}')\right) \left(\frac{1}{2} + f(\mathbf{k}')\right) \left(\frac{1}{2} + f(\mathbf{p}) + \frac{1}{2} + f(\mathbf{k})\right) \right. \\ \left. - \left(\frac{1}{2} + f(\mathbf{p})\right) \left(\frac{1}{2} + f(\mathbf{k})\right) \left(\frac{1}{2} + f(\mathbf{p}') + \frac{1}{2} + f(\mathbf{k}')\right) \right]$$

(f^3 and f^2 correct, but spurious f^1 terms)

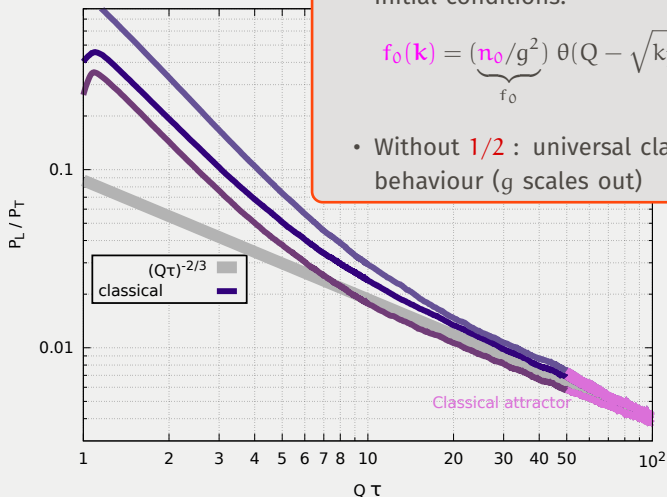
Strong classical approximation: drop the $\frac{1}{2}$ (f^3 correct)

$$\times \left[f(\mathbf{p}') f(\mathbf{k}') (1 + f(\mathbf{p})) (1 + f(\mathbf{k})) \right. \\ \left. - f(\mathbf{p}) f(\mathbf{k}) (1 + f(\mathbf{p}')) (1 + f(\mathbf{k}')) \right]$$

\mathbf{K}'

ISOTROPIZATION IN A LONGITUDINALLY EXPANDING SYSTEM

[Epelbaum, FG, Jeon, Moore, Wu (2015)]



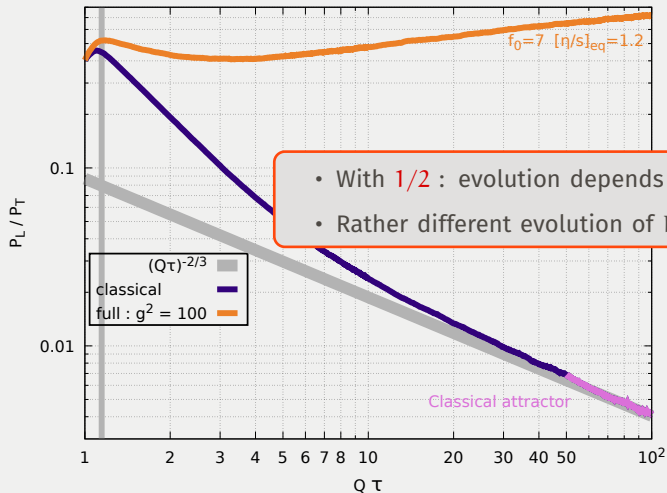
- Initial conditions:

$$f_0(\mathbf{k}) = \underbrace{(n_0/g^2)}_{f_0} \theta(Q - \sqrt{k_{\perp}^2 + \xi_0 k_z^2})$$

- Without $1/2$: universal classical behaviour (g scales out)

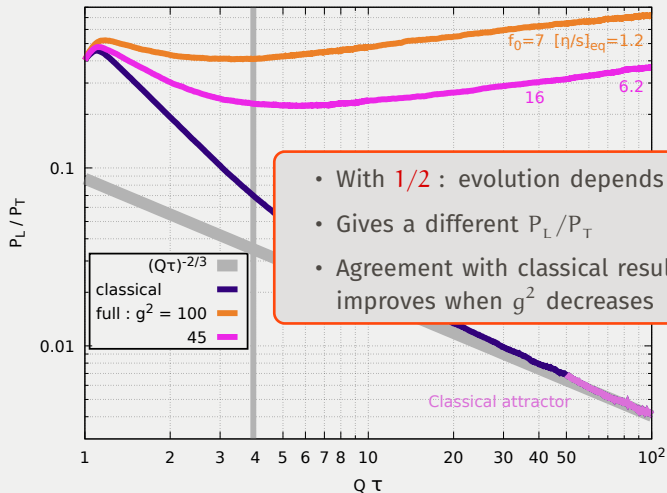
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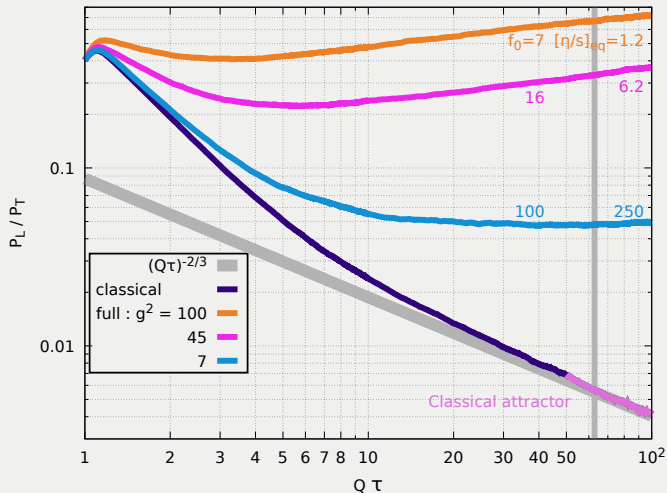
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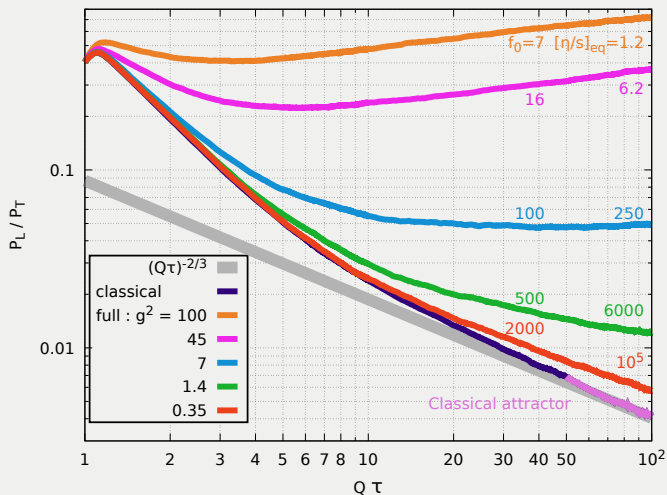
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WHY IS THE ZERO-POINT $1/2$ IMPORTANT ?

- The $1/2$'s ensure that the terms f^3 and f^2 are correct
- The quadratic terms are important in anisotropic systems

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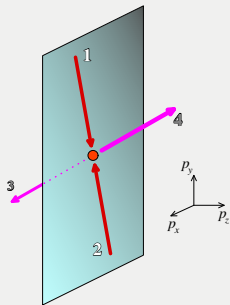
- The 1/2's ensure that the terms f^3 and f^2 are correct
- The quadratic terms are important in anisotropic systems

- No 1/2 \implies no f^2 terms in Boltzmann eq. :

$$\partial_t f_4 \sim g^4 \int_{123} \dots [f_1 f_2 (f_3 + f_4) - f_3 f_4 (f_1 + f_2)] \\ + \dots [f_1 f_2 - f_3 f_4]$$

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$$\partial_t f_4 \sim g^4 \int_{123} \dots \left[\cancel{f_1 f_2 (f_3 + f_4)} - \cancel{f_3 f_4 (f_1 + f_2)} \right] + \dots \left[\cancel{f_1 f_2} - \cancel{f_3 f_4} \right]$$

- When the distribution is very anisotropic, trying to produce the particle 4 at large angle results in $f_3 \approx f_4 \approx 0 \implies$ nothing left
- Cubic terms \Leftrightarrow stimulated emission : ineffective to produce particles in empty regions of phase-space
- Leading term: $f_1 f_2$

More insights from kinetic theory

Boltzmann in the relaxation time approximation

$$\left(\partial_\tau - \frac{p_z}{\tau}\right) f(\tau, \mathbf{p}) = -\frac{f - f_{\text{eq}}}{\tau_R}$$

$\tau_R \equiv$ relaxation time

$f_{\text{eq}} \equiv$ local equilibrium dist

- $\tau_R = \infty$: no collisions
- $\tau_R \sim \epsilon^{-1/4}$: conformal; rate scales as inverse temperature

- Define moments :

$$L_n \equiv \int_{\mathbf{p}} \mathbf{p}^2 P_{2n}(p_z/p) f(\tau, \mathbf{p}) \quad , \quad g_n \equiv \tau \partial_\tau \ln L_n$$

$$L_0 = \epsilon = P_L + 2P_T, \quad L_1 = P_L - P_T$$

Boltzmann \Leftrightarrow **coupled equations for** L_n

$$\partial_\tau L_0 = -\frac{a_0 L_0 + c_0 L_1}{\tau}$$

$$\partial_\tau L_n = -\frac{a_n L_n + c_n L_{n+1} + b_n L_{n-1}}{\tau} - \frac{L_n}{\tau_R} \quad (n \geq 1)$$

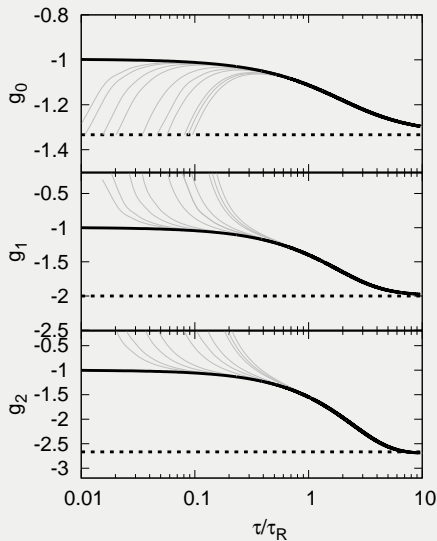
$a_n, b_n, c_n =$ pure numbers, known explicitly (depend only on the free streaming part of Boltzmann eq.)

Free streaming fixed point ($\tau_R = \infty$)

- All the g_n behave as τ^{-1} , with fixed ratios
- $L_1/L_0 \rightarrow -\frac{1}{2}$, i.e. $P_L/P_T \rightarrow 0$

Interacting fixed point ($\tau_R \sim \epsilon^{-1/4}$)

- $g_0 \rightarrow -4/3$, $g_1 \rightarrow -2$
- Locally isotropic distribution



- Universal attractor
- $\tau \lesssim \tau_R$: trajectories first approach free streaming fixed point
- $\tau \gtrsim \tau_R$: trajectories go to the local equilibrium fixed point

Summary

- Strong fields: short mean free path despite weak coupling
- LO: no pressure isotropization, NLO: secular instabilities
- Beyond NLO: **Semi-classical approximation**
 - **Weak classical approximation:**
non-renormalizable, sensitive to UV cutoff
 - **Strong classical approximation:**
underestimates large angle scatterings
poor unless η/s very large
- **Kinetic theory:** avoids all these difficulties (but does not cope well with screening effects at long distance)
- Beyond semi-classical in QFT: **Two-PI formalism**
(Luttinger-Ward functional)

ANALOGUE IN QUANTUM MECHANICS

- Consider the Liouville–von Neumann equation :

$$i \hbar \frac{\partial \hat{\rho}_\tau}{\partial \tau} = [\hat{H}, \hat{\rho}_\tau]$$

- Introduce the Wigner transforms :

$$W_\tau(\mathbf{x}, \mathbf{p}) \equiv \int d\mathbf{s} e^{i\mathbf{p}\cdot\mathbf{s}} \langle \mathbf{x} + \frac{\mathbf{s}}{2} | \hat{\rho}_\tau | \mathbf{x} - \frac{\mathbf{s}}{2} \rangle$$

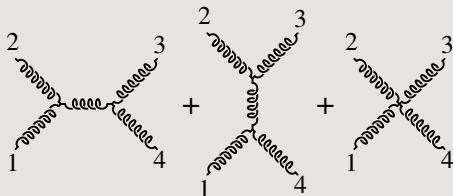
$$\mathcal{H}(\mathbf{x}, \mathbf{p}) \equiv \int d\mathbf{s} e^{i\mathbf{p}\cdot\mathbf{s}} \langle \mathbf{x} + \frac{\mathbf{s}}{2} | \hat{H} | \mathbf{x} - \frac{\mathbf{s}}{2} \rangle$$

- LvN equation is equivalent to Moyal-Groenewold equation

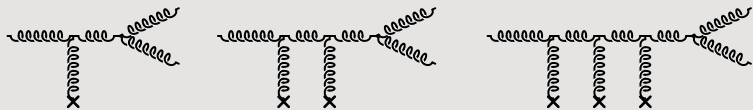
$$\begin{aligned} \frac{\partial W_\tau}{\partial \tau} &= \mathcal{H}(\mathbf{x}, \mathbf{p}) \frac{2}{i \hbar} \sin \left(\frac{i \hbar}{2} \left(\overleftarrow{\partial}_{\mathbf{p}} \overrightarrow{\partial}_{\mathbf{x}} - \overleftarrow{\partial}_{\mathbf{x}} \overrightarrow{\partial}_{\mathbf{p}} \right) \right) W_\tau(\mathbf{x}, \mathbf{p}) \\ &= \underbrace{\{\mathcal{H}, W_\tau\}}_{\text{Poisson bracket}} + \underbrace{\mathcal{O}(\hbar^2)}_{\text{deviation from classical dynamics}} \end{aligned}$$

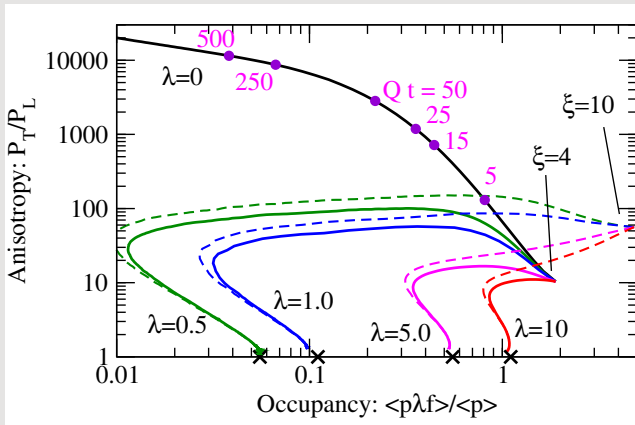
KINETIC THEORY FOR GLUONS [Kurkela, Zhu (2015)]

$2 \rightarrow 2$



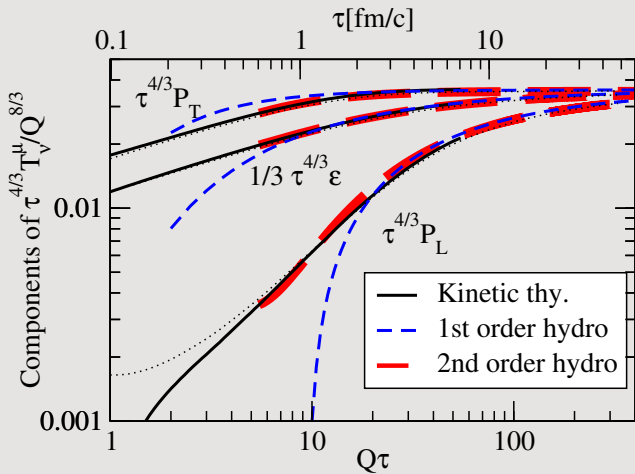
$1 \rightarrow 2, 2 \rightarrow 1$ + Landau-Pomeranchuk-Migdal resummation





For $\lambda = 0.5$, the **Strong CSA** breaks down at $Q\tau \approx 2$, while simple estimates suggested that it would be valid up to $Q\tau \approx \alpha_s^{-3/2} \approx 350$

Consistent with hydrodynamics before full isotropization



Energy-momentum tensor

$$\begin{aligned} T^{\mu\nu} = & \nabla^\mu \varphi \nabla^\nu \varphi - g^{\mu\nu} \mathcal{L} + \left[\nabla_x^\mu \nabla_y^\nu G_{xy} \right]_{x=y} \\ & + \frac{1}{2} g^{\mu\nu} \left\{ V''(\varphi_x) G_{xx} - \left[\nabla_\alpha^x \nabla_y^\alpha G_{xy} \right]_{x=y} \right\} \\ & - g^{\mu\nu} \frac{\delta \Phi}{\delta \sqrt{-g}} \end{aligned}$$

ISOTROPIZATION IN A FIXED BOX

