

A cosmic background image showing a colorful, abstract pattern of light and dark regions, resembling a nebula or a map of the universe. The colors range from deep blues and purples to bright yellows and oranges.

Cosmogenic activation

What is it?

How does it impact experiments?

How well we can model it?

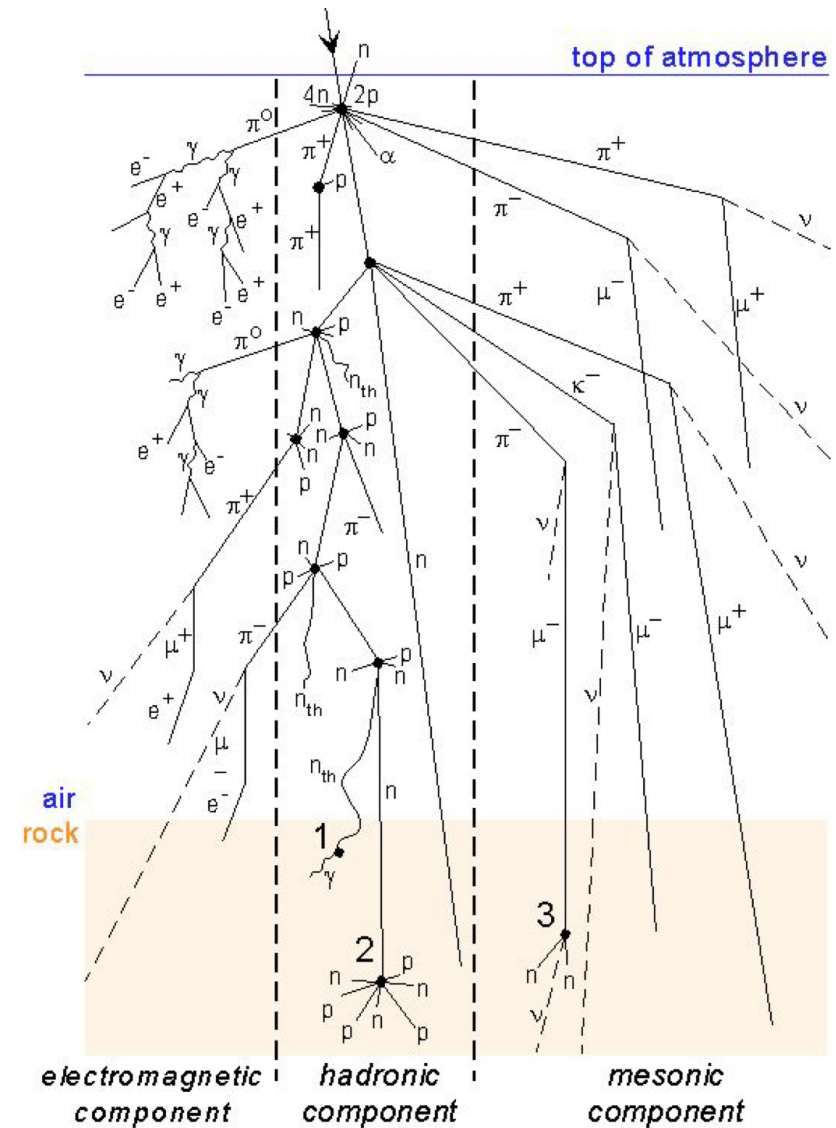
How to control it?

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.. with help from M. De Jesus & C. Augier

Cosmic Ray interactions

- Electromagnetic component stopped in atmosphere
- Hadronic (n) component penetrates a few m μ
$$I(mwe) \sim I_0 e^{-mwe/[1.5m]}$$

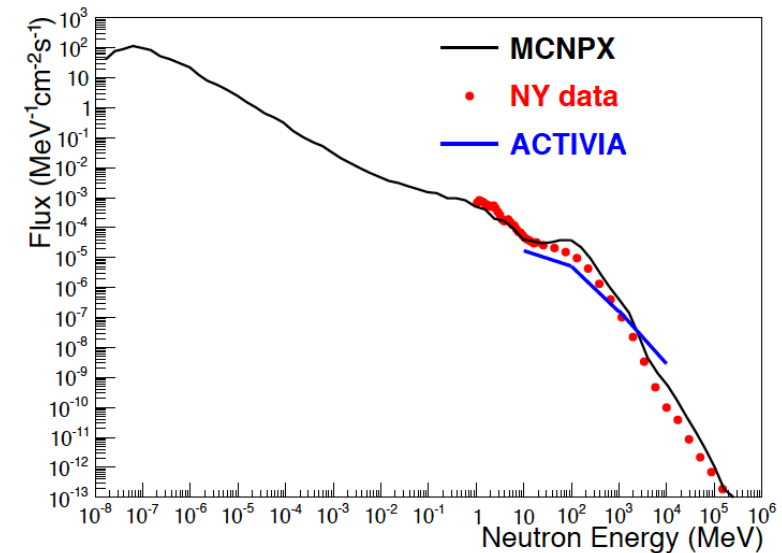
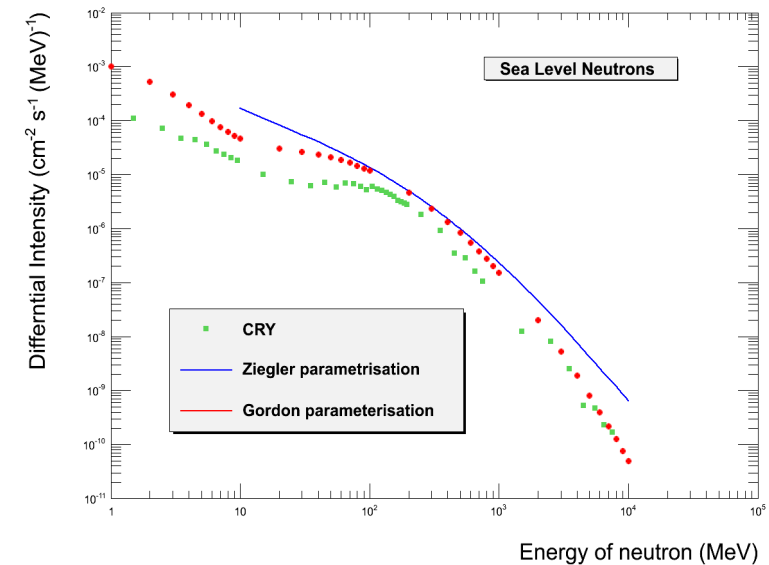
[³H in rock, arxiv:1607.08770]
- Deep underground, only high-energy muons can produce nuclear reactions
- Neutrinos? [HALO expt...]



- Cosmic ray flux Φ_0
- Attenuation of flux in matter (mostly: n, μ): $\Phi_{\mu,n}(z)$
- Cross-section σ for production of isotope X on a target Y
- Production during a time t_{prod} : $\sigma \Phi (1 - \exp(-t_{\text{prod}}/\tau))$
- Decay of Y during a time t_{decay} : $\exp(-t_{\text{decay}}/\tau)$

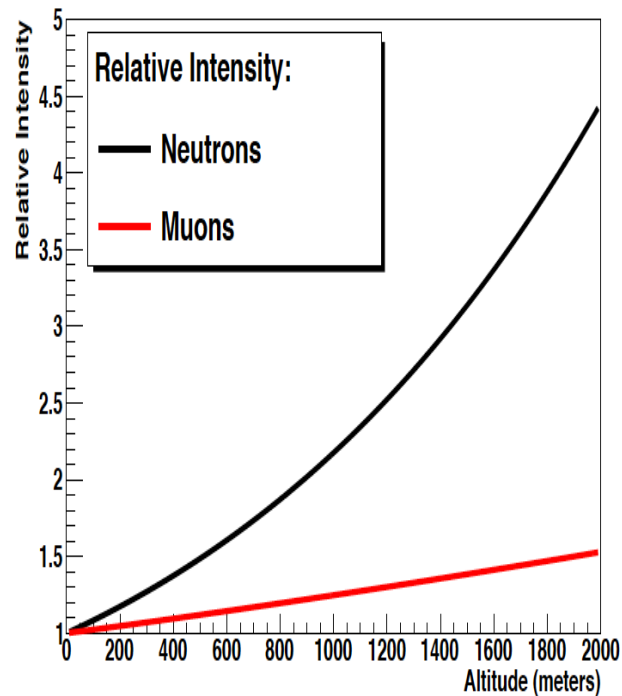
Uncertainties in neutron flux

- Parameterizations:
Gordon (New York data),
Ziegler
- Models: Activia, MCNPX,
COSMO, GEANT4,
FLUKA, ... (see
Kudryavtsev, LRT2017)



Attenuation of flux

■ Neutron flux calculator



Neutron Flux Calculation

seutest.com/cgi-bin/FluxCalculator.cgi

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Flux calculation

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Use this calculator to determine the relative neutron flux at a particular location. The output value is relative to the sea-level flux in New York City, New York, USA. This point has historically been the reference point for neutron flux measurements.

[Instructions for use are here \(opens a new window\)](#) and summarized at the bottom of this page.

Location

Latitude ☒ N degrees ☐ S degrees

Longitude ☐ E degrees ☒ W degrees

Elevation, Pressure or Depth - Enter a single value and check the appropriate box

Elevation ☒ ☒ feet ☐ meters

Station pressure ☐ ☒ mm Hg ☐ inches Hg ☐ millibar (hPa)

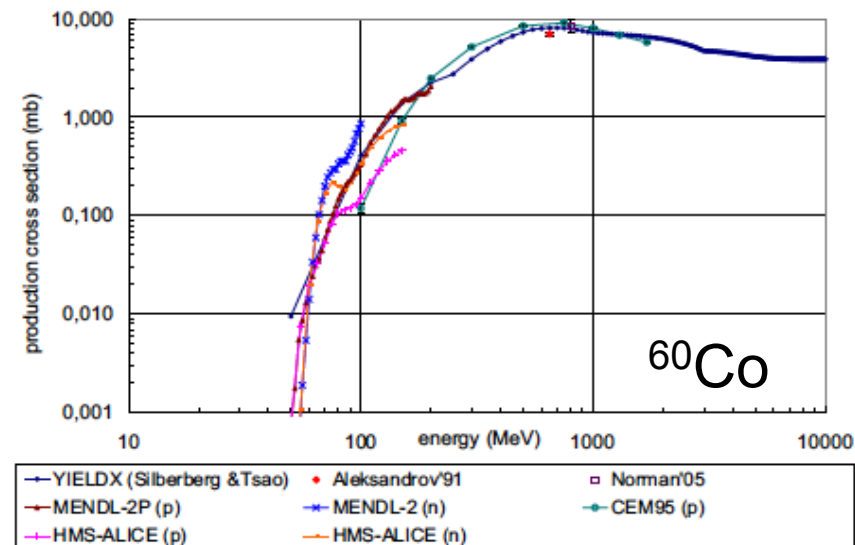
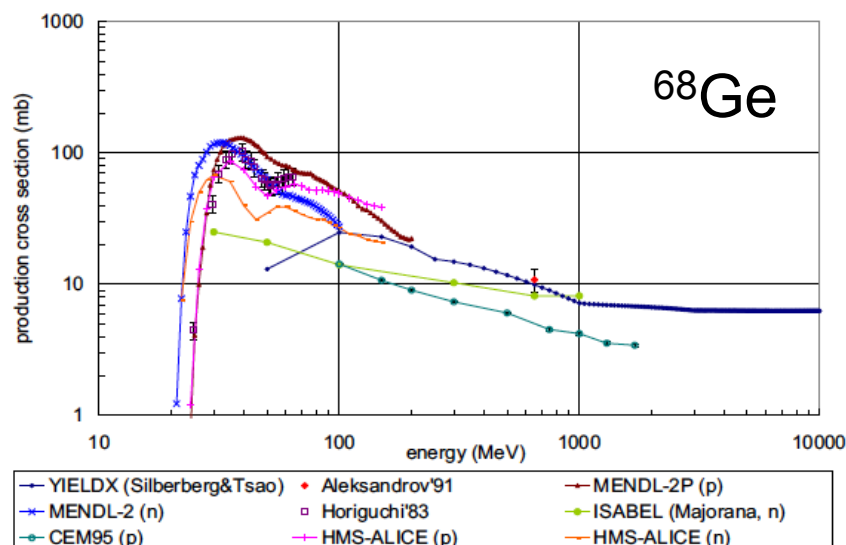
Atmospheric depth ☐ g/cm²

Solar Modulation

Solar modulation % (0 is minimum flux / active sun, 100 is maximum flux / quiet sun)

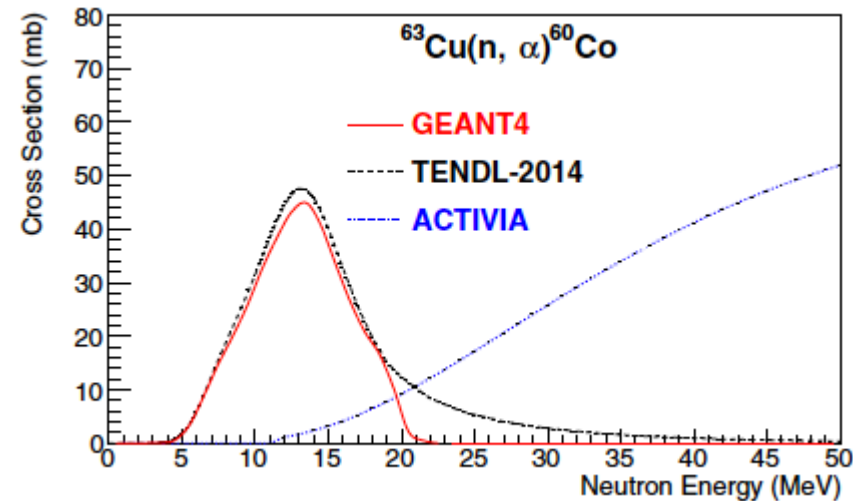
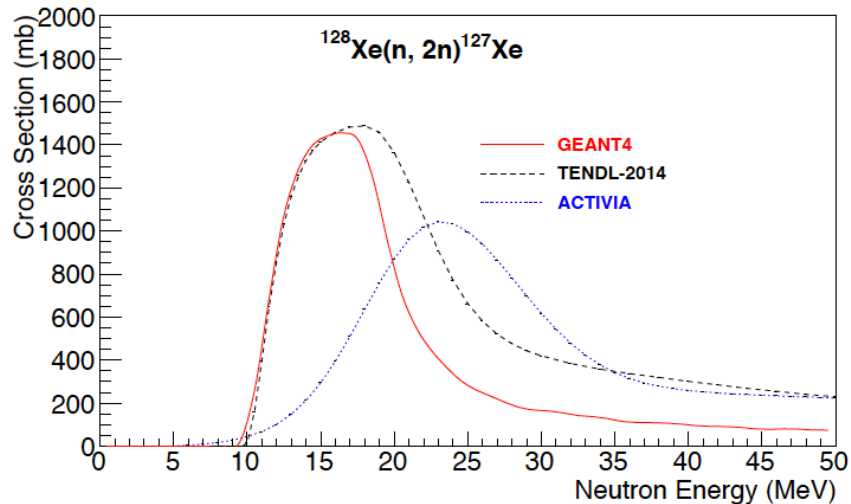
Relative flux (NYC, NY, USA = 1.00)

Production Cross-Sections

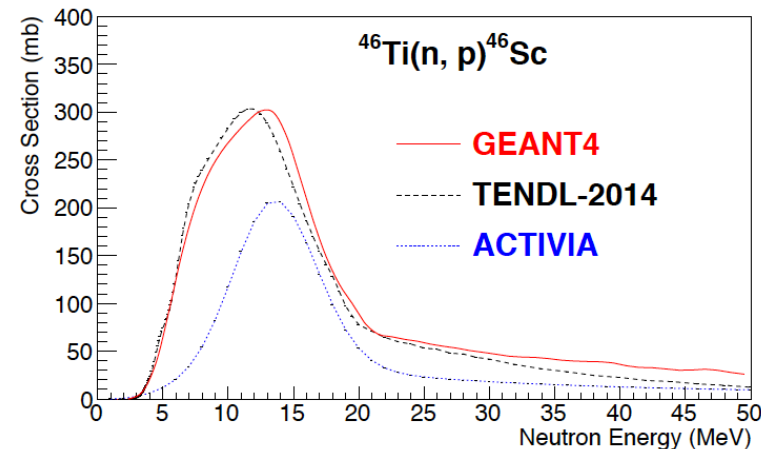


- Production cross-sections (excitation functions) of ^{68}Ge and ^{60}Co in natural Ge (Cebrian et al. Journal of Physics, Conference series, **39** (2006) 344; Cebrian et al. Astroparticle Physics, **33** (2010) 316; Review by Cebrian at LRT2013).
- Cross-sections:
 - Silberberg, Tsao and Barghouty. Astrophys. J., **501** (1998) 911; YIELDX.
 - MENDL-2/2P libraries: <https://www-nds.iaea.org/publications/iaea-nds/iaea-nds-0136.htm>.

Production Cross-Section at Low Energy

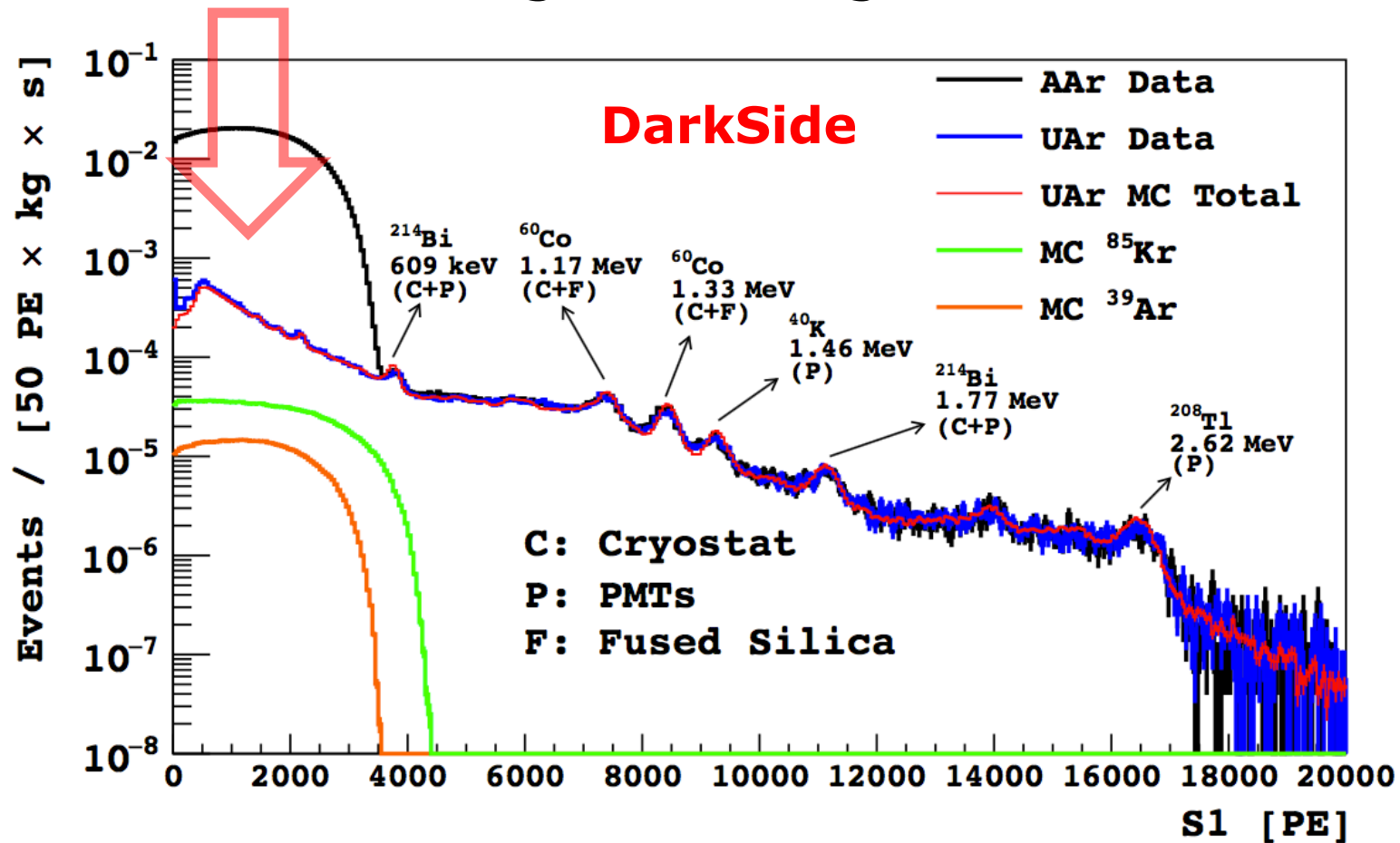


- Evaluated Nuclear Data File – ENDF (used in GEANT4, G4NDL).
- Experimental Nuclear Reaction Data, EXFOR.
- TENDL (-2015) – output from TALYS.
- From Zhang et al. *Astroparticle Physics*, **84** (2016) 62.
- ACTIVIA (Silberberg and Tsao): J. J. Back and Y. A. Ramachers, *Nucl. Instr. and Meth. A* **586** (2008) 286.



Impact in Argon

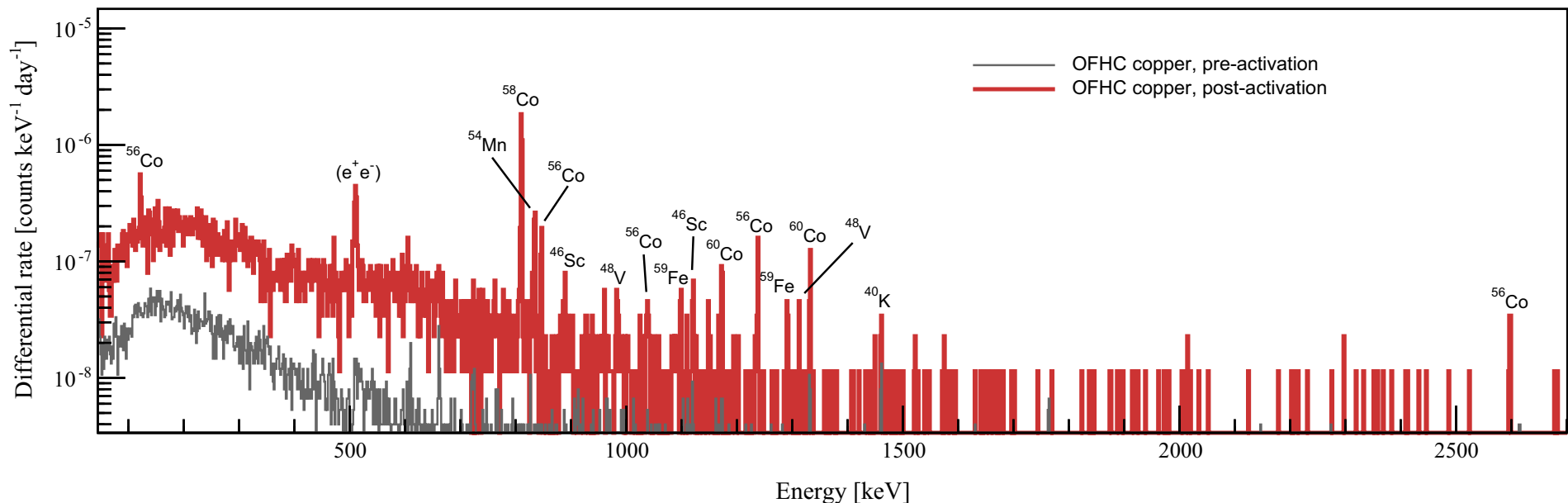
- Huge impact on low-energy background of removal of cosmogenic background of ^{39}Ar



Impact: ^{60}Co activation in Cu

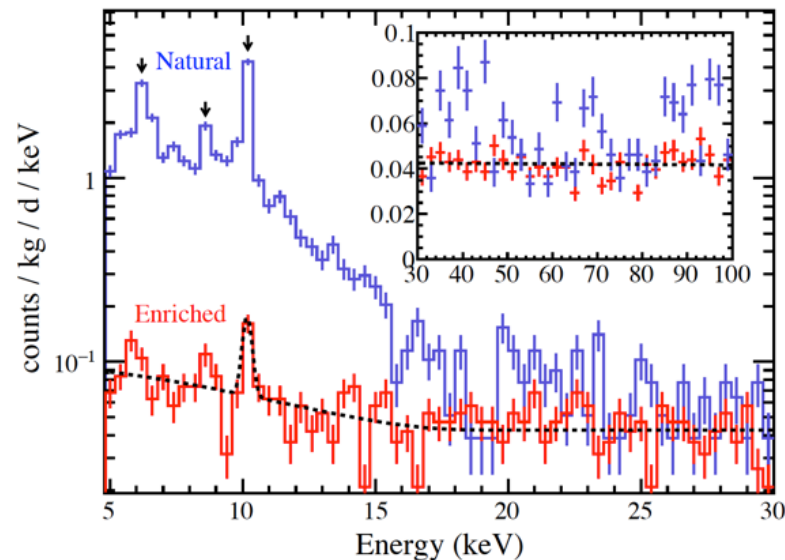
Ex. of measurement: Baudis et al., EPJ C 75 (2015) 485

- Exposure of 10.35 kg Cu (OFHC) to **cosmic rays** 3470 m above sea level (x11) for 345 days
- ^{60}Co activation (at saturation): $340^{+82}_{-60} \mu\text{Bq/kg}$ (but others: x3)
- Not dominant γ background wrt U/Th + ^{40}K [see eg LUX: AP 62 (2015) 33; Majorana NIMA 828 (2016) 22]... but with efforts



Impact: ^3H in Ge

- ^3H is dangerous for low-energy searches:
low end point (18 keV) large half-life (12 y)
- Expected to be the main background for SuperCDMS
SNOLAB Ge-HV search for low-mass WIMPs



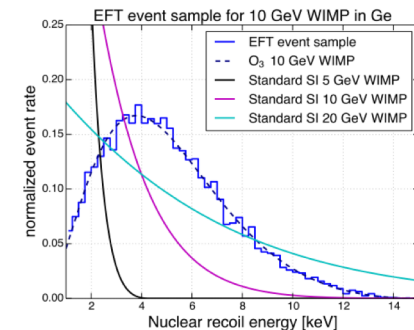
Phys. Rev. Lett. 118, 161801 (2017)

The reduction in the low energy events can be seen between our enriched and natural detectors.

The natural detectors spent significant surface time in Los Alamos and production was not closely tracked.

MAJORANA

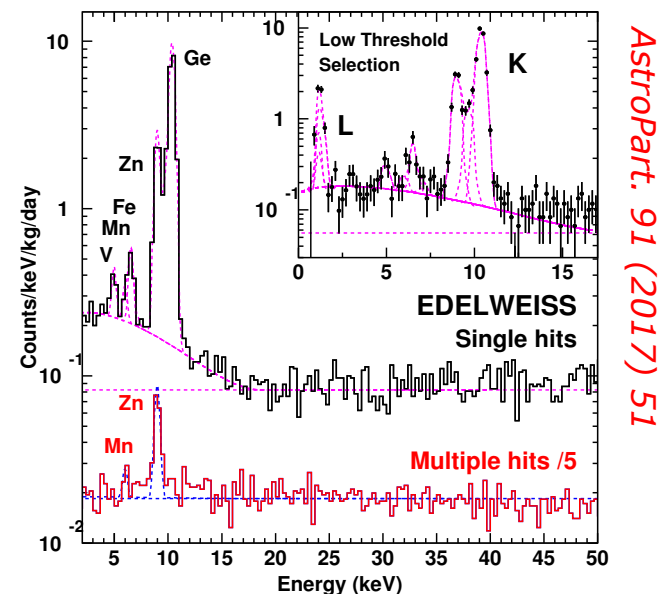
Compared to
5-10-20 GeV
WIMPs



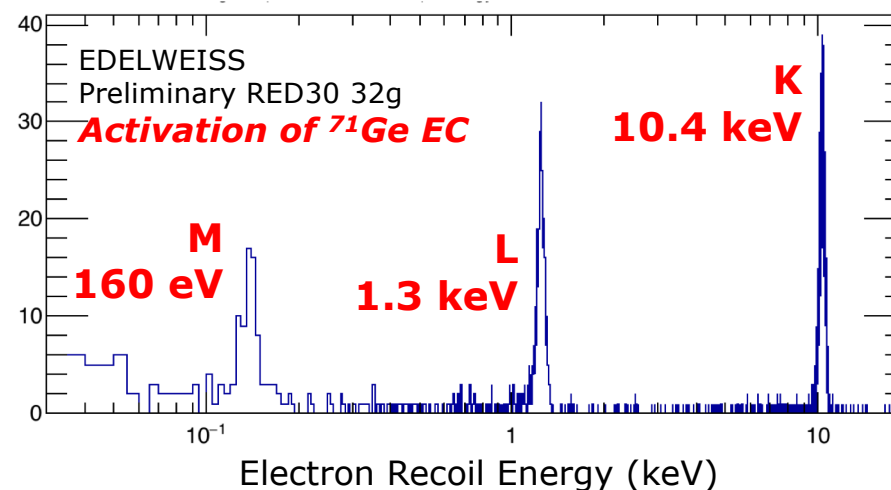
EDELWEISS
measurement:
 82 ± 21 atom/day/kg

Good Impact? Bulk Calibration...

- Cosmogenic activation is uniform over the entire volume:
useful for low-energy calibration of massive detectors and fiducial volume measurements



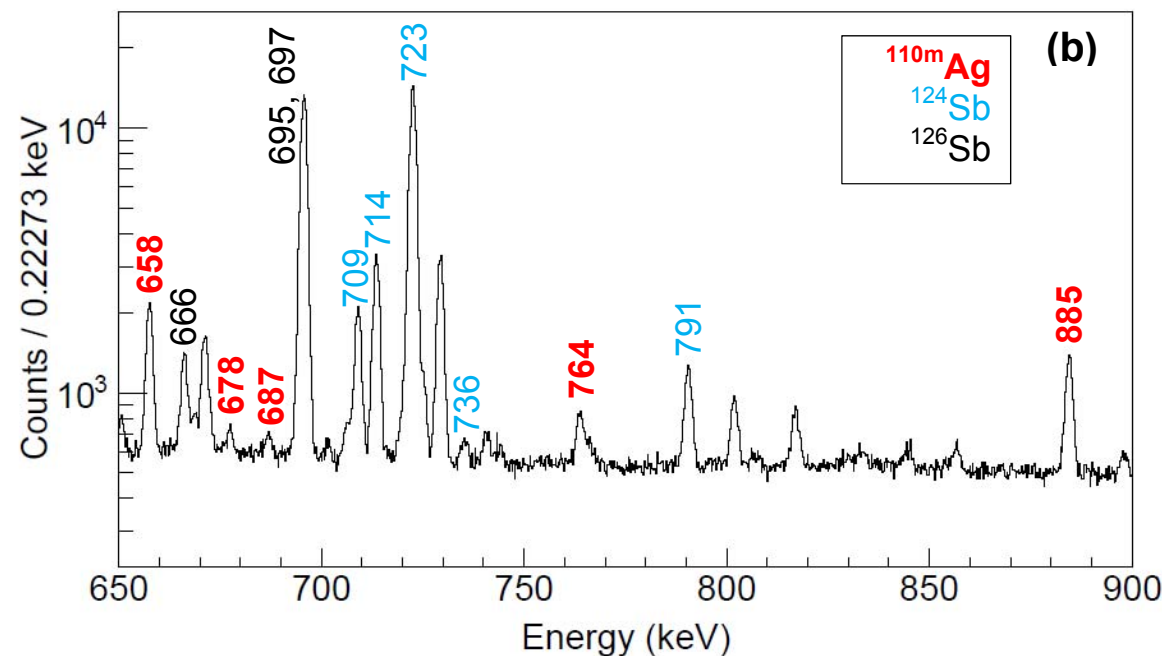
- However thermal neutron activation (here: ⁷¹Ge) is more selective & cleaner



Production yield measurement in TeO_2

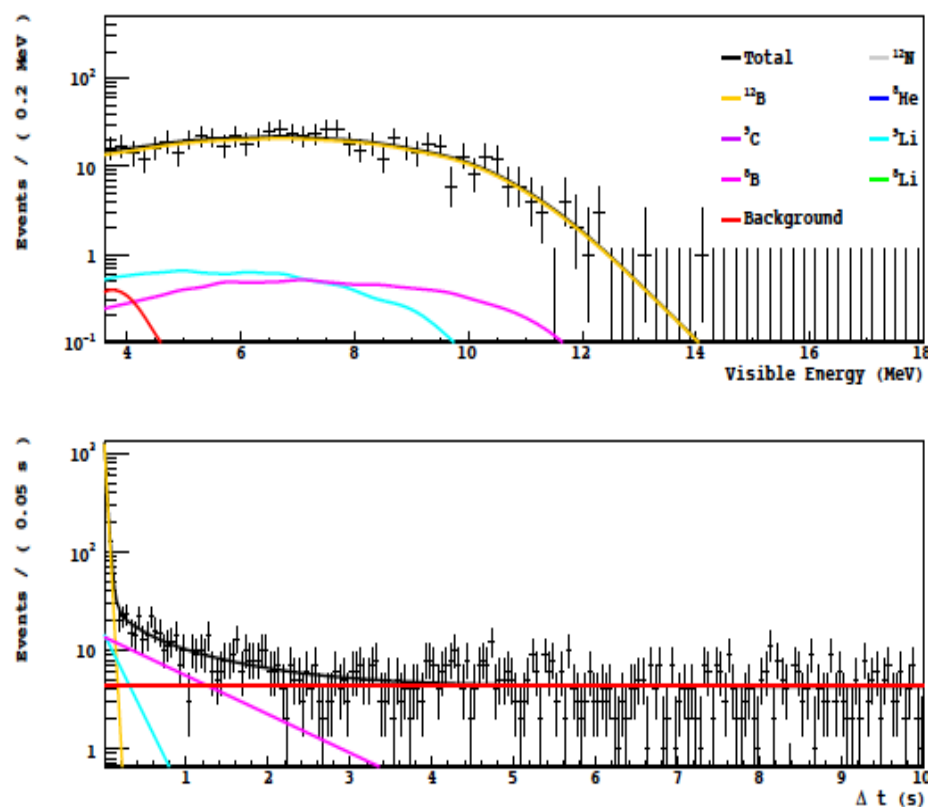
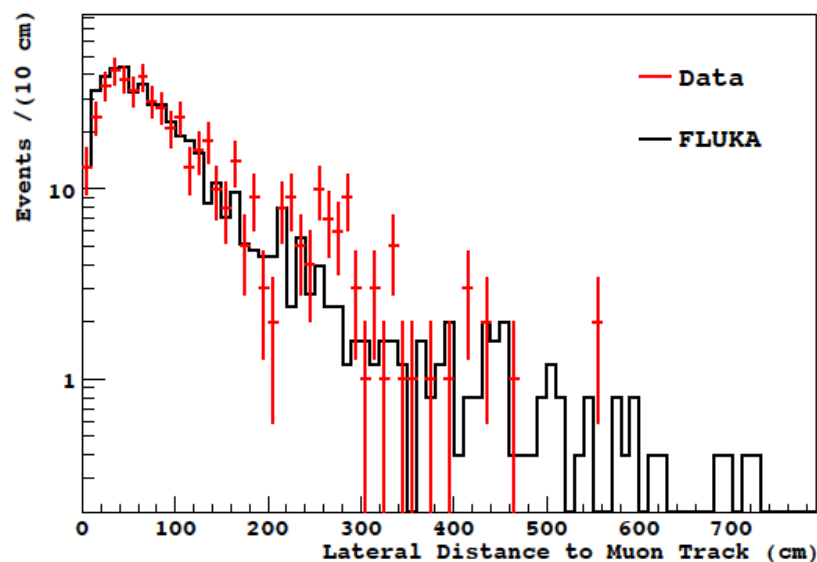
CUORE, PRC 92 (2015) 024620

- Replace cosmic rays with ($\rightarrow 800$ MeV) n and ($\rightarrow 28$ GeV) p **beams**
- Measure activation of TeO_2 powder with Ge
- Check that $T_{1/2} > 1\text{y}$ isotopes with $Q > 2.5$ MeV have negligible impact on backgrounds in region of interest for $2\beta 0\nu$



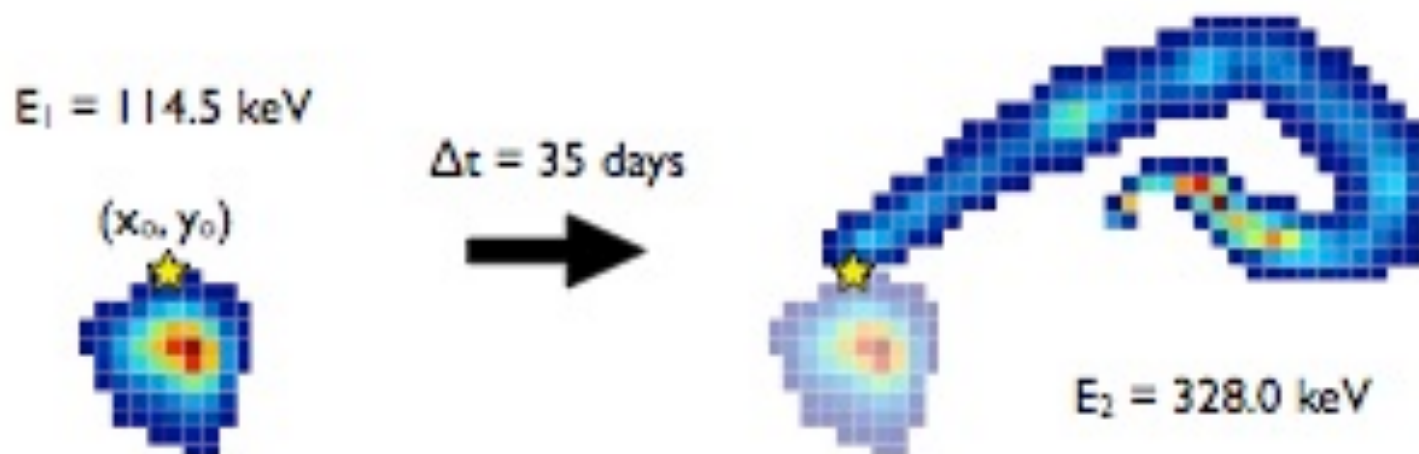
Underground Production Measurement

- Kamland, Borexino: measure cosmogenic isotope production correlated with muon track ($\langle E_\mu \rangle \sim 270$ GeV)
- Ex: Borexino, JCAP 1308 (2013) 049, arXiv:1304.7381



Production yield measurement in Si

- ^{32}Si β^- decay may be an important bkg for CDMS SNOLAB Si-HV
- Origin: spallation of Ar in atmosphere
- Decay cascade:
 $^{32}\text{Si} \beta^- \Rightarrow ^{32}\text{P} \quad Q = 227 \text{ keV} \quad T_{1/2} = 150 \text{ y}$
 $^{32}\text{P} \beta^- \Rightarrow ^{32}\text{S} \quad Q = 1.7 \text{ MeV} \quad T_{1/2} = 14 \text{ d}$
- DAMIC CCD: from observed 13 Si/P pairs with <7 pixel separation and <70 days apart, deduce a production rate of 80^{+110}_{-65} atoms/kg/day @95%CL. Reconstruction efficiency: 49%.



Control of exposure

Example: MAJORANA, NIMA 779 (2015) 52

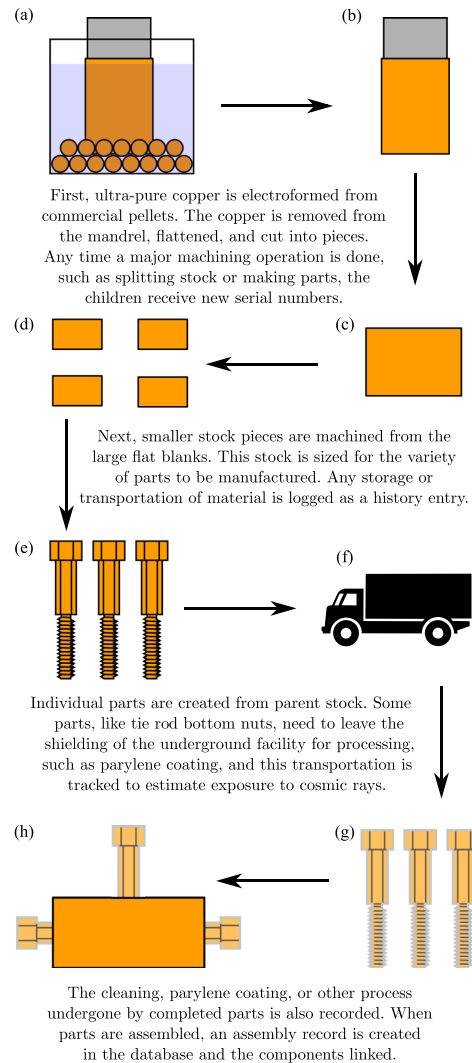
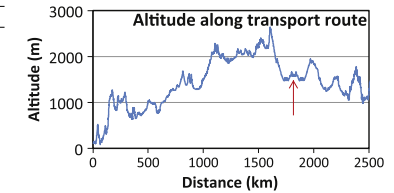
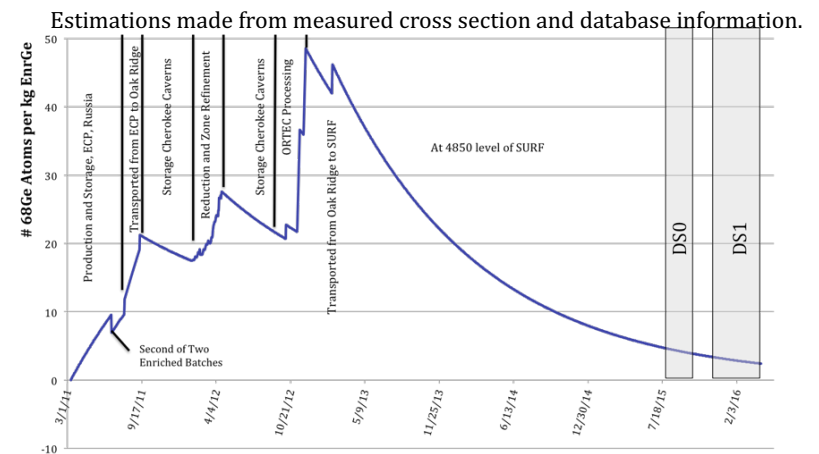


Table 1
FedEx Tracking information (upper left table) and the inferred set of transportation and storage record entries for the PTDB (lower table). All dates occur in the calendar year 2012. The figure shown in the upper right presents the altitude variation experienced by the part during shipment, with a notable day long stop in Denver, CO located at red arrow.

FedEx Tracking data (reverse chronological order)		
Date, Time	Activity	Location
Jun 13, 12:06	Delivered	Lead, SD
Jun 13, 10:37	Out for delivery	Rapid City, SD
Jun 13, 8:50	At local facility	Rapid City, SD
Jun 13, 5:00	At local facility	Rapid City, SD
Jun 12, 21:10	In transit	Denver, CO
Jun 11, 8:32	In transit	Boise, ID
Jun 9, 2:50	In transit	Hermiston, OR
Jun 8, 20:45	Left FedEx origin	Pasco, WA
Jun 8, 14:23	Picked up	Richland, WA



Parts Tracking Database records information (chronological order)			Exposure calculation
Record Type	Start - End (Date, Time) - (Date, Time)	Movement or Location (City, State)	Sea-level Equivalent Exposure (hours)
Transport	Jun 8, 14:23 PDT - Jun 8, 14:43 PDT	Richland, WA → Pasco, WA	0.4
Storage	Jun 8, 14:43 PDT - Jun 8, 20:45 PDT	Pasco, WA	4.6
Transport	Jun 8, 20:45 PDT - Jun 8, 21:30 PDT	Pasco, WA → Hermiston, OR	0.9
Storage	Jun 8, 21:30 PDT - Jun 9, 2:50 PDT	Hermiston, OR	6.2
Transport	Jun 9, 2:50 PDT - Jun 9, 8:03 MDT	Hermiston, OR → Boise, ID	11.2
Storage	Jun 9, 8:03 MDT - Jun 11, 8:32 MDT	Boise, ID	114.9
Transport	Jun 11, 8:32 MDT - Jun 11, 20:42 MDT	Boise, ID → Denver, CO	95.2
Storage	Jun 11, 20:42 MDT - Jun 12, 21:10 MDT	Denver, CO	133.5
Transport	Jun 12, 21:10 MDT - Jun 13, 3:12 MDT	Denver, CO → Rapid City, SD	29.8
Storage	Jun 13, 3:12 MDT - Jun 13, 10:37 MDT	Rapid City, SD	21.1
Transport	Jun 13, 10:37 MDT - Jun 13, 11:34 MDT	Rapid City, SD → Lead, SD	2.9
Storage	Jun 13, 11:34 MDT - Jun 13, 12:06 MDT	Lead, SD	2.7



- Production in underground lab
- Storage in underground lab
- Adapt production process to reduce exposure
- Shallow labs? (remove hadronic component)
- Measurements to further tests present codes