

## Workshop on Gas-filled Detectors and Systems (GDS): Rare-gas target handling and recycling systems

*Institut de Physique Nucléaire (IPN), Orsay (FRANCE), January 23-25, 2019*

# The JENSA windowless gas-jet target

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MICHIGAN STATE  
UNIVERSITY



**JENSA**

A logo featuring the word "JENSA" in large black letters, with a small atomic symbol containing a red and purple starburst above the letter "A".

# Thanks to the JENSA collaboration

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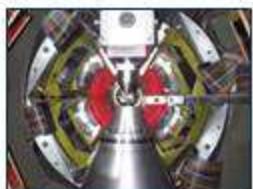


# Outline



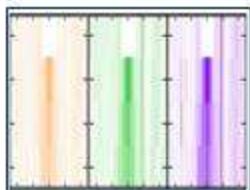
## Part 1

Jet Experiments in Nuclear Structure and Astrophysics



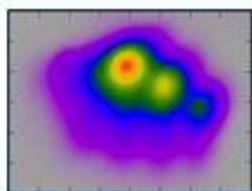
## Part 2

Overlap between jet and ion beam



## Part 3

Jet thickness measurements



## Part 4

Rare isotope beam experiments with JENSA



## Part 5

The future of JENSA

# **Part 1**

## **Jet Experiments in Nuclear Structure and Astrophysics**

# Astrophysical motivation

novae, X-ray bursts,  
supernovae, etc.

hydrogen and  
helium induced  
reactions

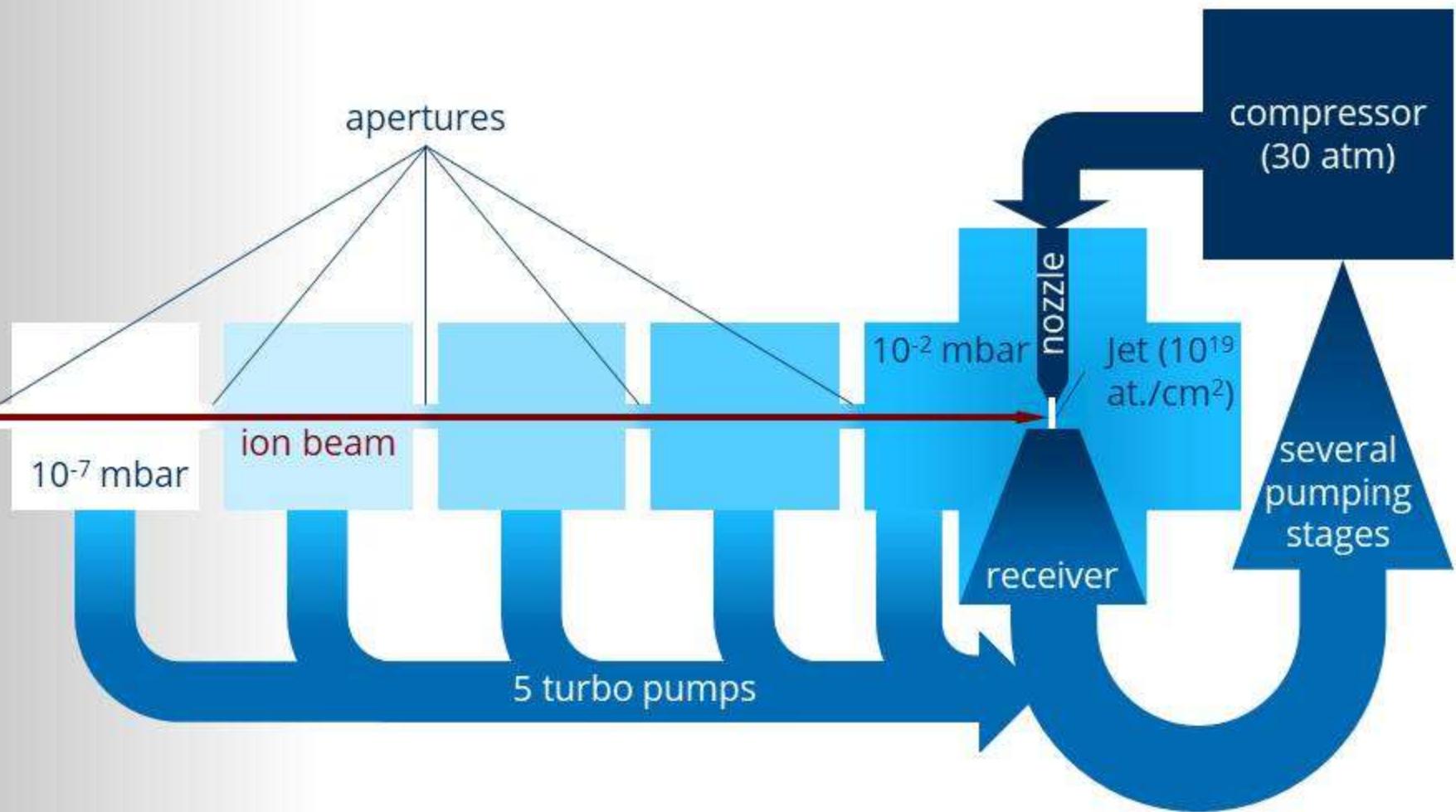
study in inverse  
kinematics with  
radioactive ion  
beams

measurements  
enabled by  
chemically pure,  
highly localized  
target with high  
density and low  
energy straggling



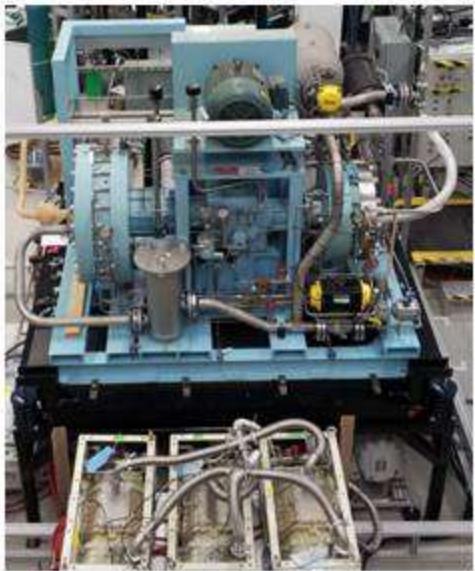
© David A. Hardy ([www.astroart.org](http://www.astroart.org))

# Recirculating gas system



# The JENSA gas jet target

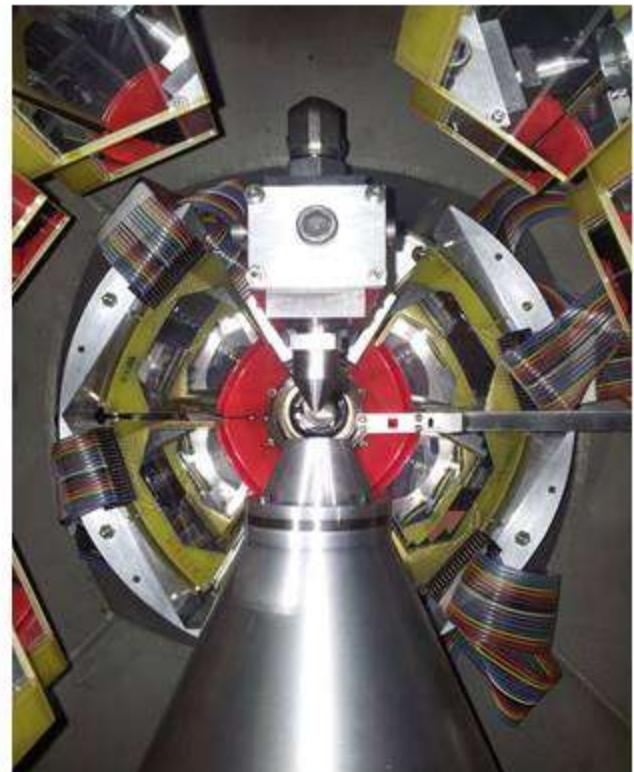
## Jet Experiments in Nuclear Structure and Astrophysics



diaphragm compressor



target chamber and pumps



nozzle, receiver and Si detectors

# Attention should be paid to ...

- Overlap between gas-jet target and ion beam



- Distribution of the jet density

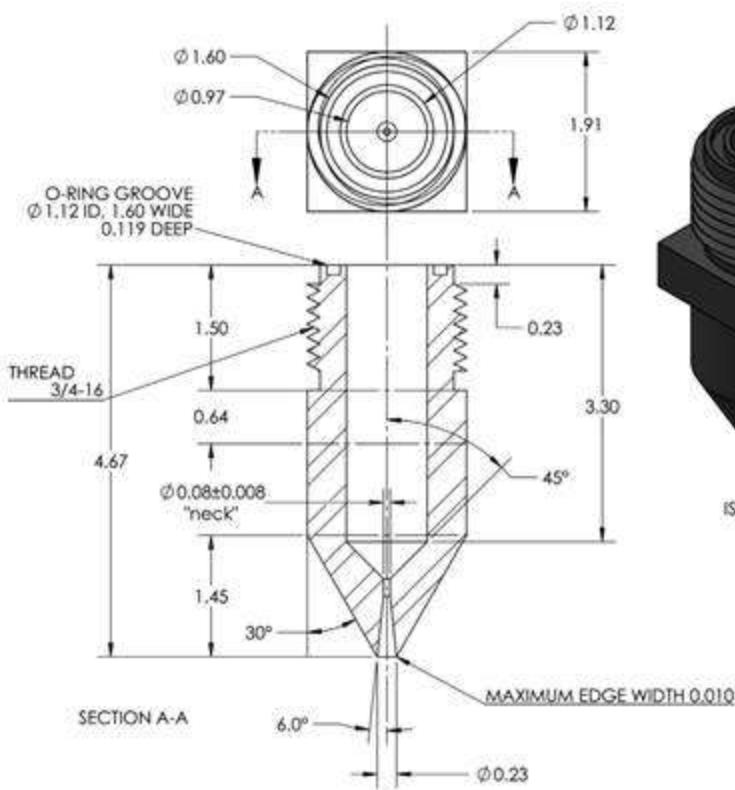
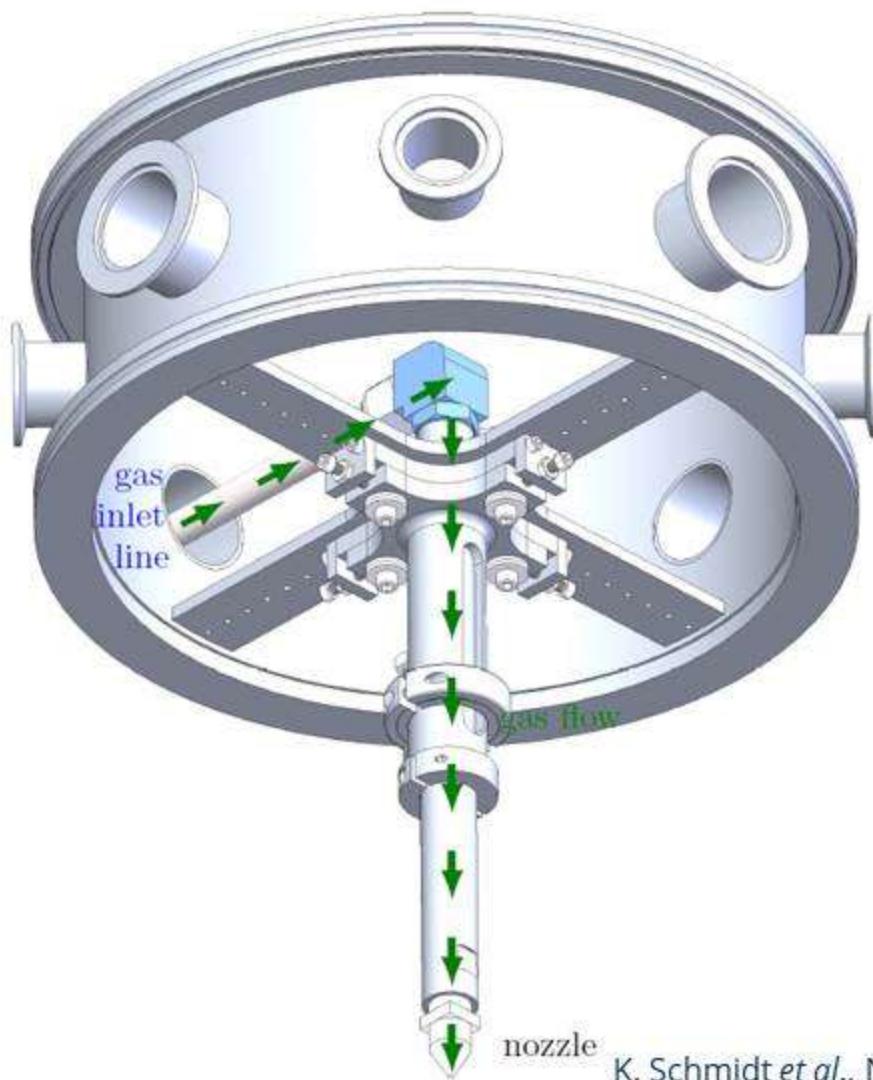
of  $S_{34}$  (35–41) is displayed in Table 2. The value quoted for (35) is the published value. There are indications (C. Rolfs, private communication) that this value includes some measurements with a gas-jet target that used an incorrect value for the target thickness. This apparently increases the extracted value of  $S_{34}$  to 0.40 keV-barns.

B.W. Filippone, Ann. Rev. Part. Sci. **36**, 717 (1986)

# Part 2

## Overlap between jet and ion beam

# Precise alignment of the nozzle



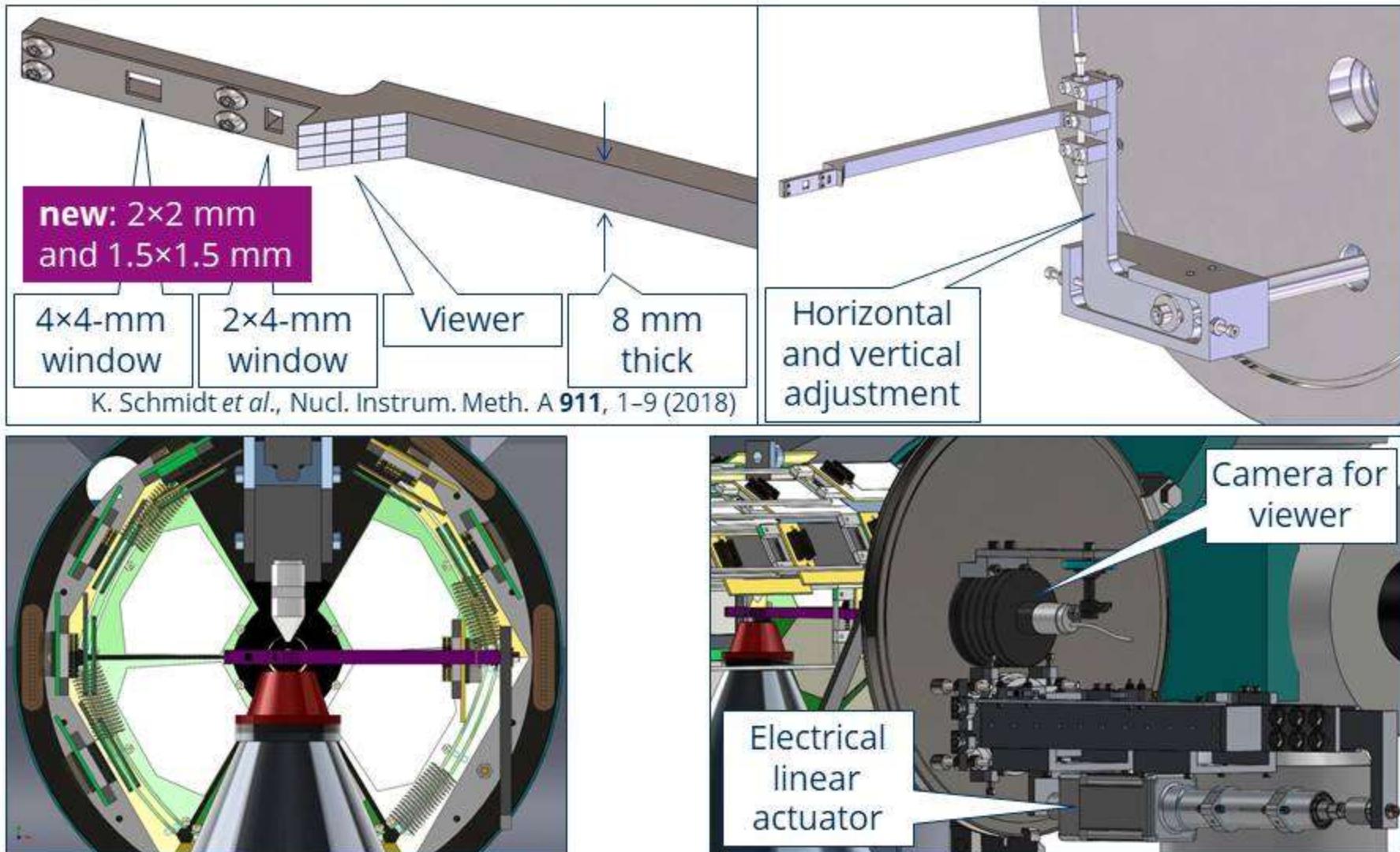
Nozzle for a supersonic jet

minimum inner diameter

0.8 mm

K. Schmidt *et al.*, Nucl. Instrum. Meth. A **911**, 1-9 (2018)

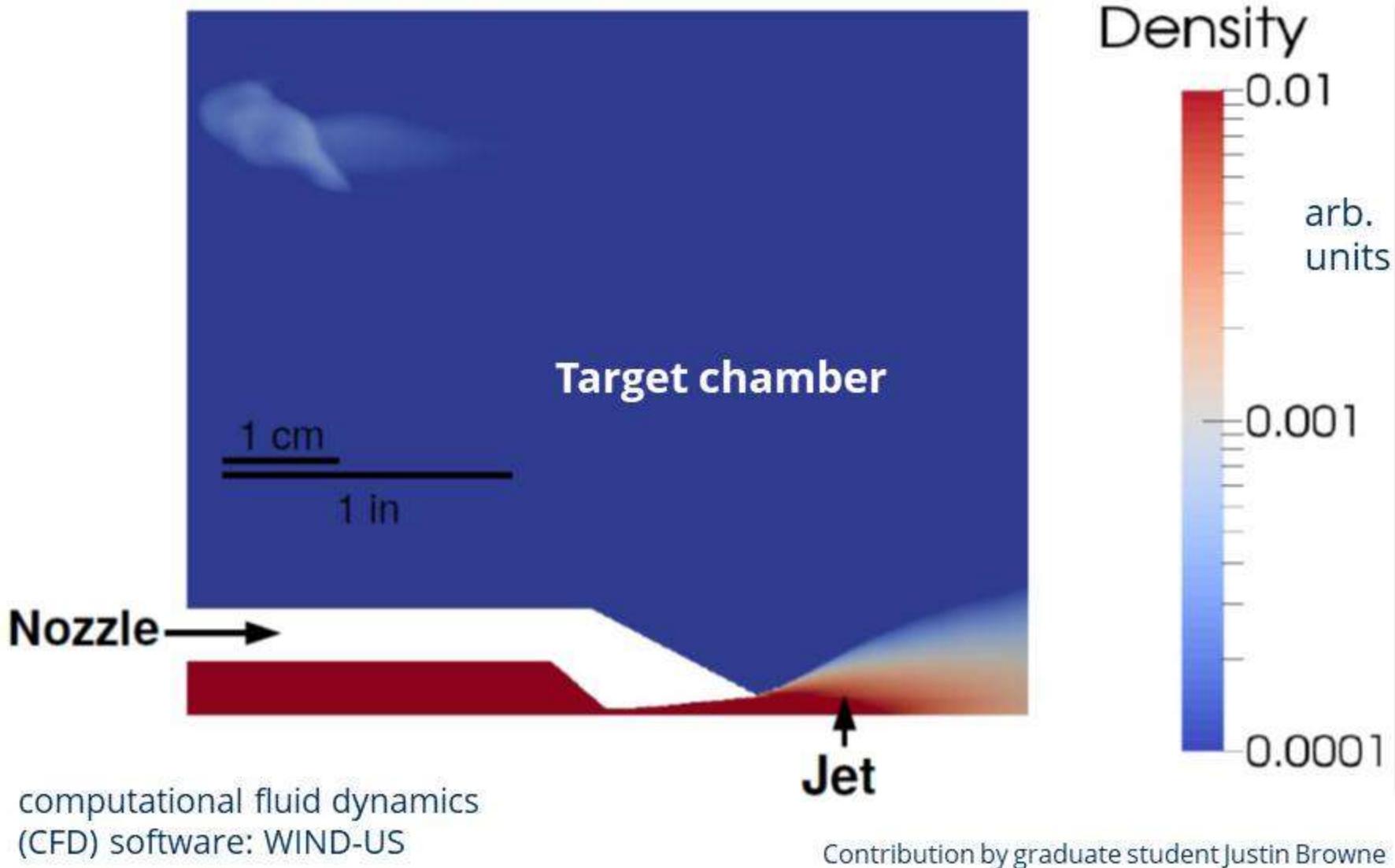
# Beam viewer for transmission check



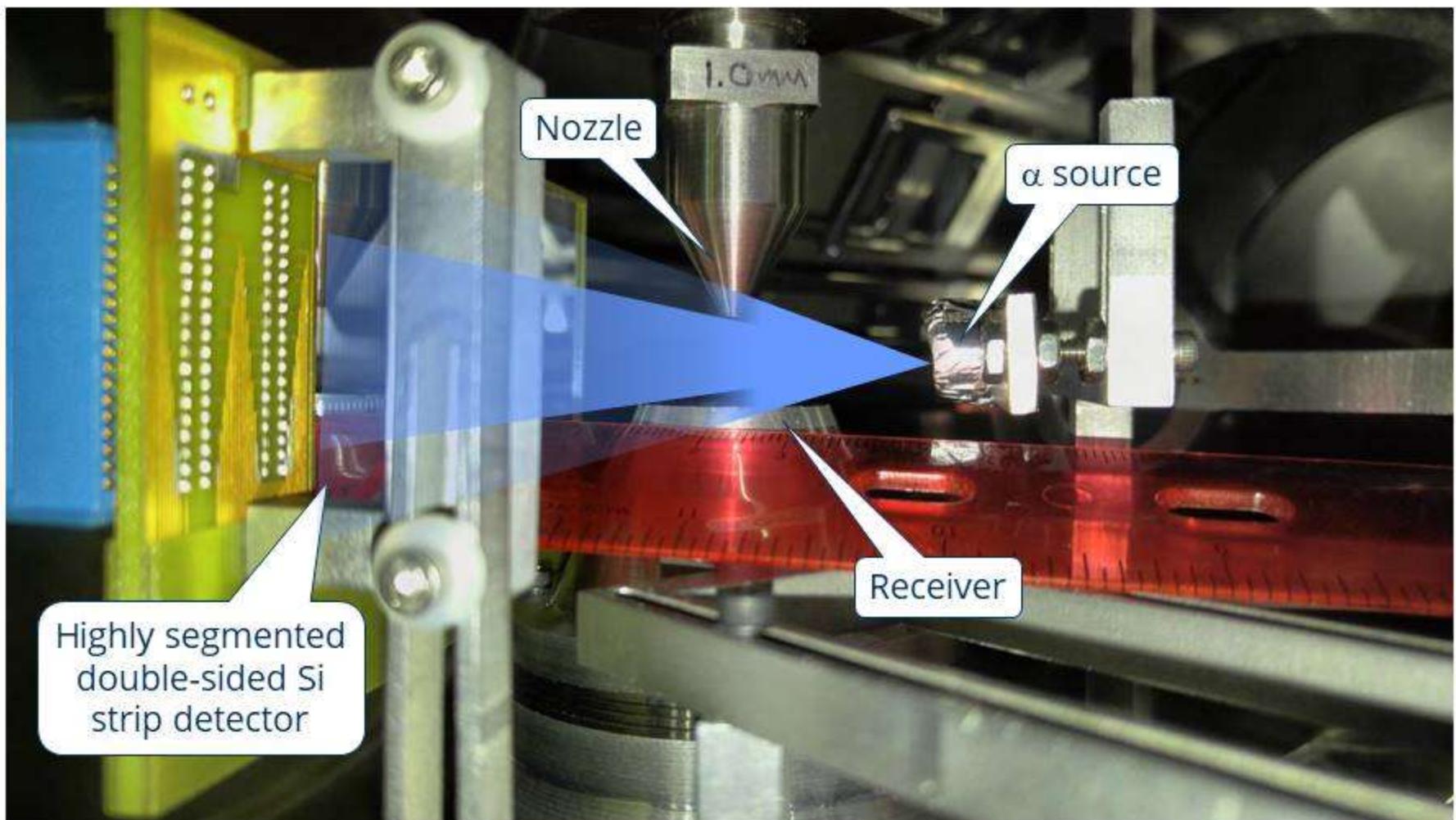
# Part 3

## Jet thickness measurements

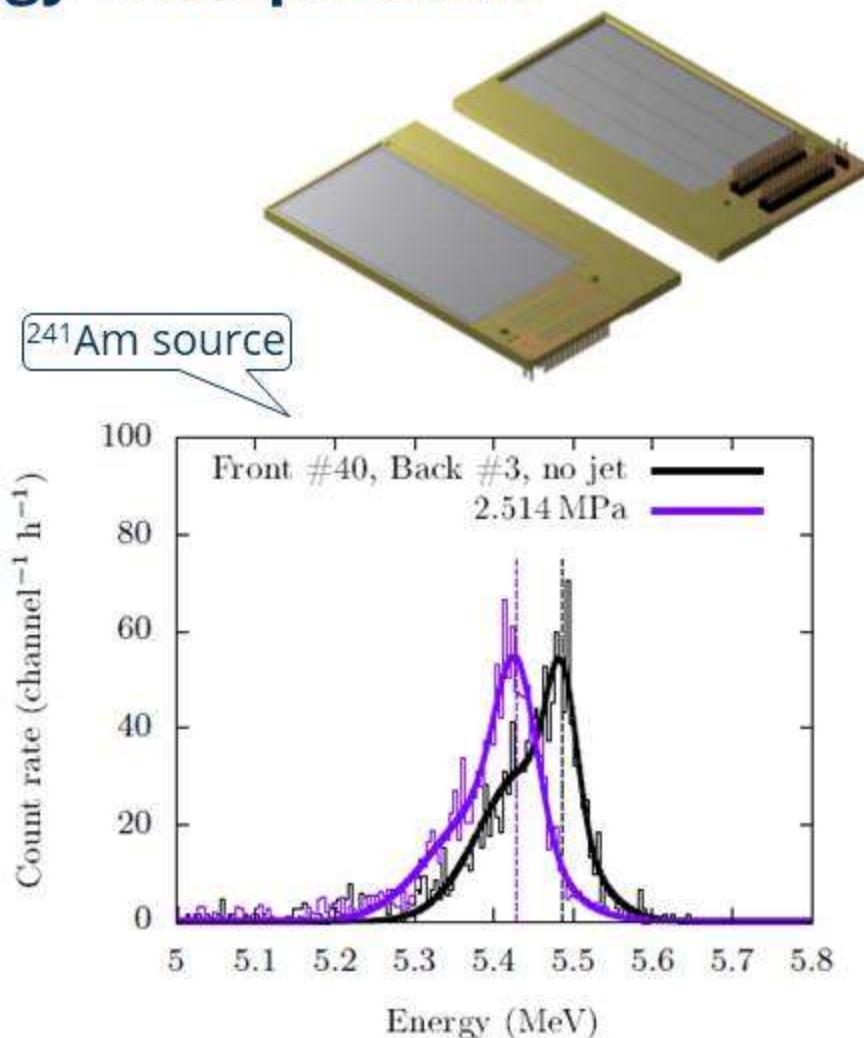
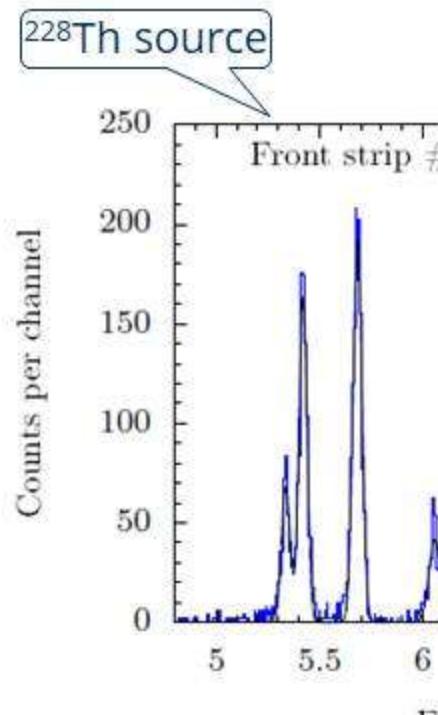
# CFD simulation of the jet



# Setup for jet thickness study

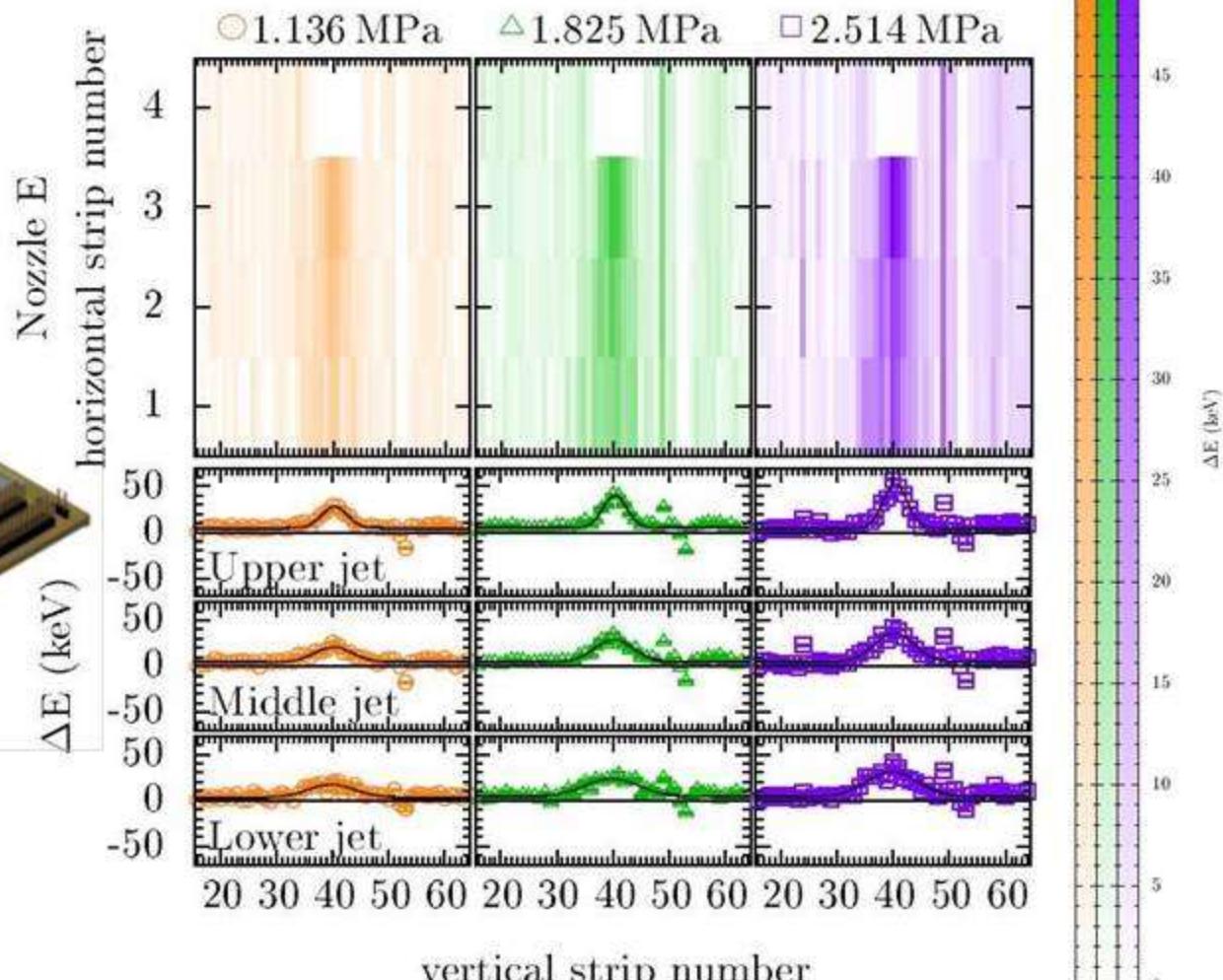
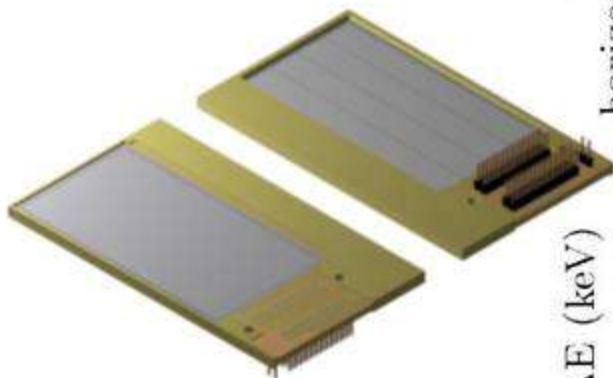


# Energy calibration and energy-loss spectrum



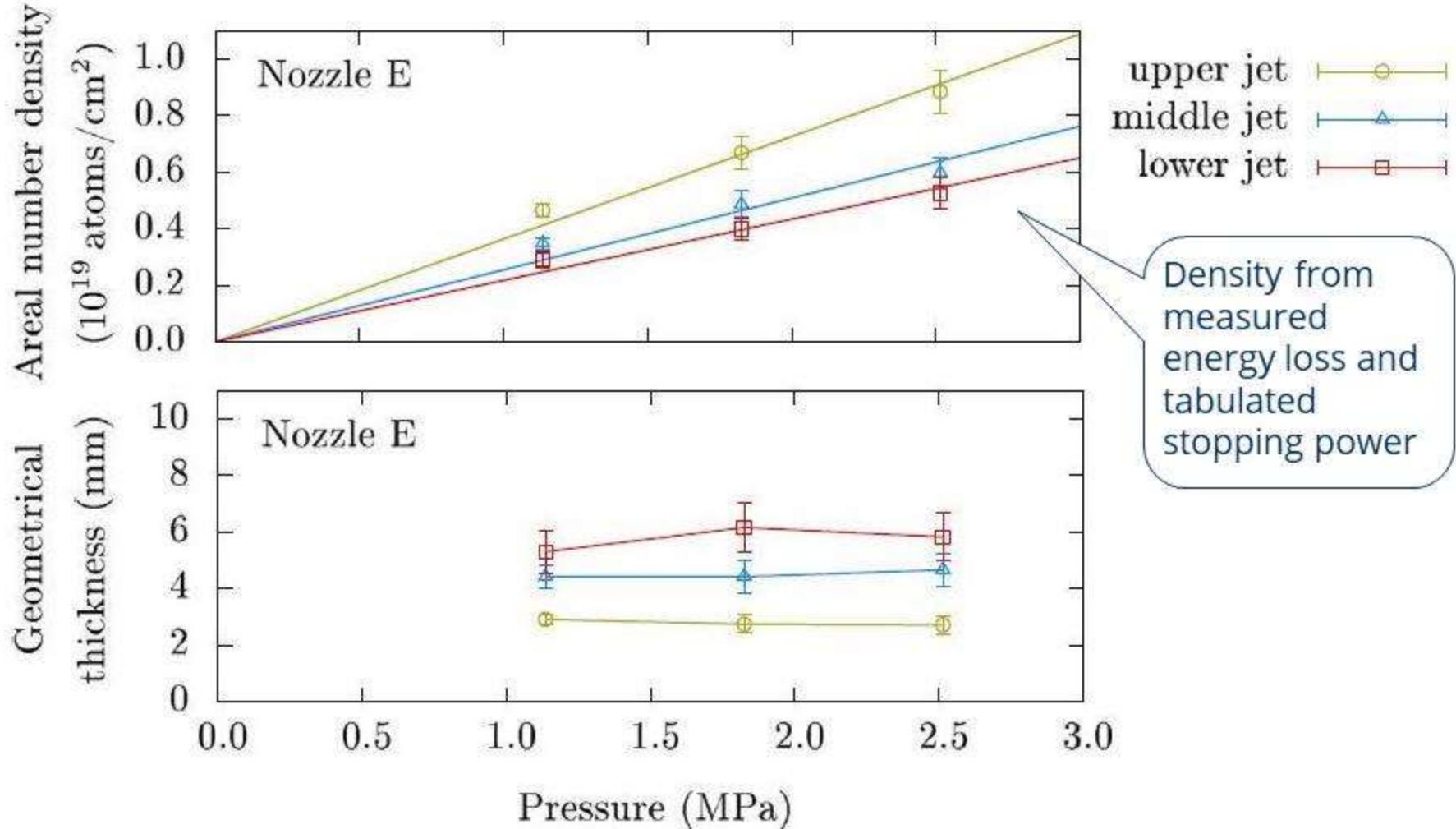
K. Schmidt *et al.*, Nucl. Instrum. Meth. A 911, 1-9 (2018)

# Energy loss profiles



K. Schmidt *et al.*, Nucl. Instrum. Meth. A **911**, 1–9 (2018)

# $10^{19}$ atoms/cm<sup>2</sup> in a 4-mm He jet

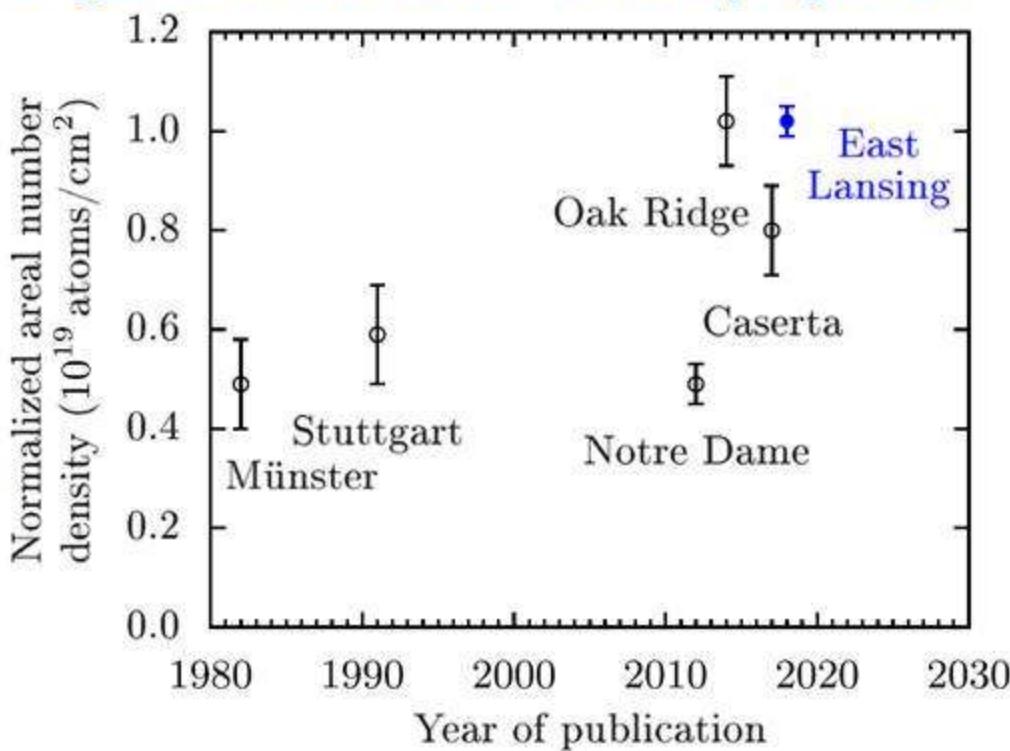


K. Schmidt *et al.*, Nucl. Instrum. Meth. A **911**, 1-9 (2018)

# Supersonic helium-jet targets in nuclear astrophysics

JENSA is the world's densest helium-jet target used in nuclear astrophysics studies

Normalized to  
2.859 MPa  
input pressure



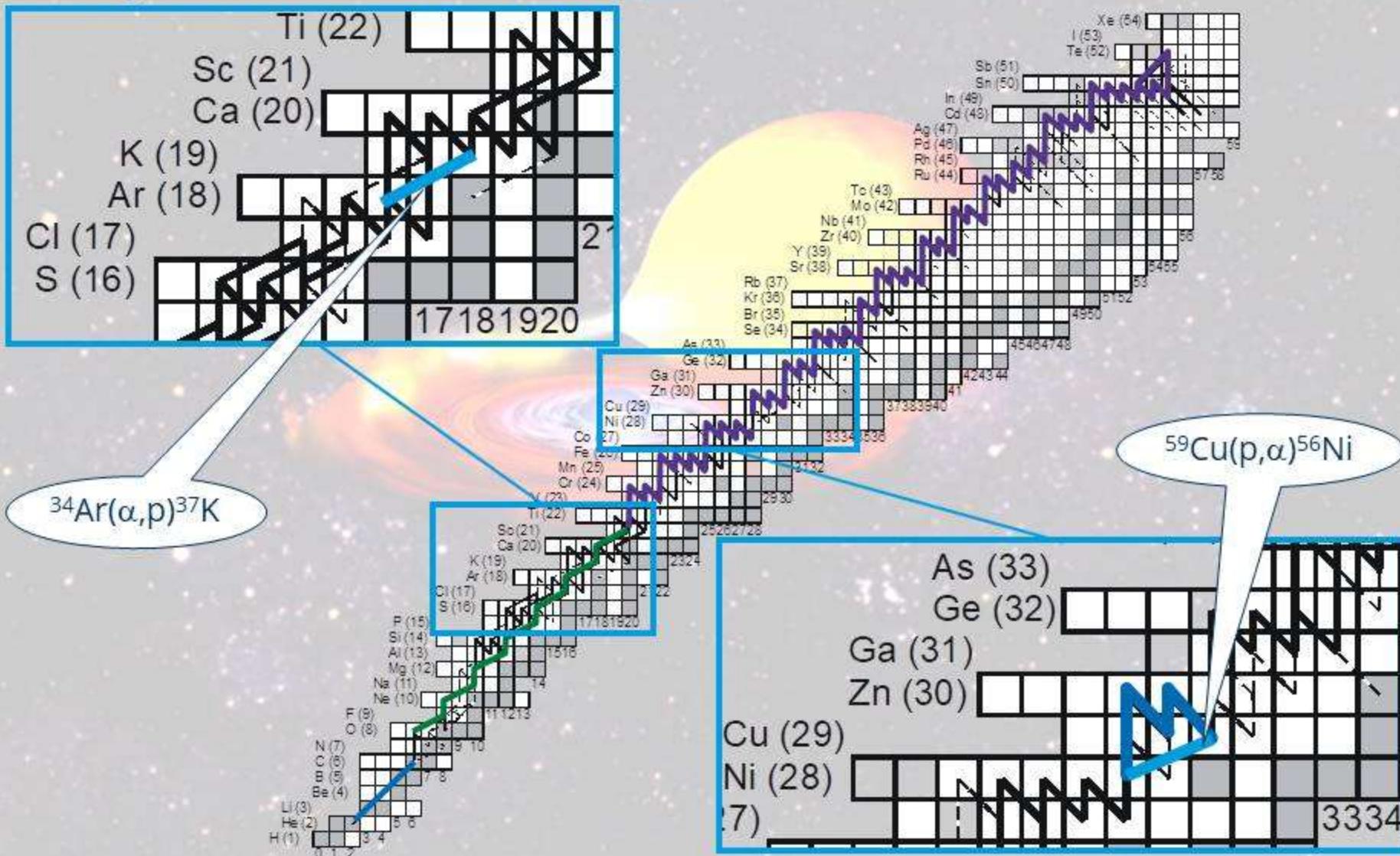
Location	Year	Input pressure (kPa)	${}^4\text{He}$ jet density ( $10^{18}$ at./cm $^2$ )	${}^4\text{He}$ jet FWHM (mm)	Distance from nozzle (mm)
Münster	1982	200	$0.34 \pm 0.06$	$2.5 \pm 0.2$	1 to 5
Stuttgart	1991	38	$0.078 \pm 0.013$	$2.6 \pm 0.2$	$\sim 1.5$
Notre Dame	2012	150	$0.259 \pm 0.021$	$2.2 \pm 0.2$	$\sim 4$
Oak Ridge	2014	2859	$10.2 \pm 0.9$	$5.1 \pm 0.3$	$\sim 1$
Caserta	2017	700	$1.97 \pm 0.21$	Not reported	$\sim 5.5$
East Lansing	2018	2515	$9.0 \pm 0.3$	$2.03 \pm 0.09$	$\lesssim 4$

K. Schmidt *et al.*, Nucl. Instrum. Meth. A **911**, 1–9 (2018)

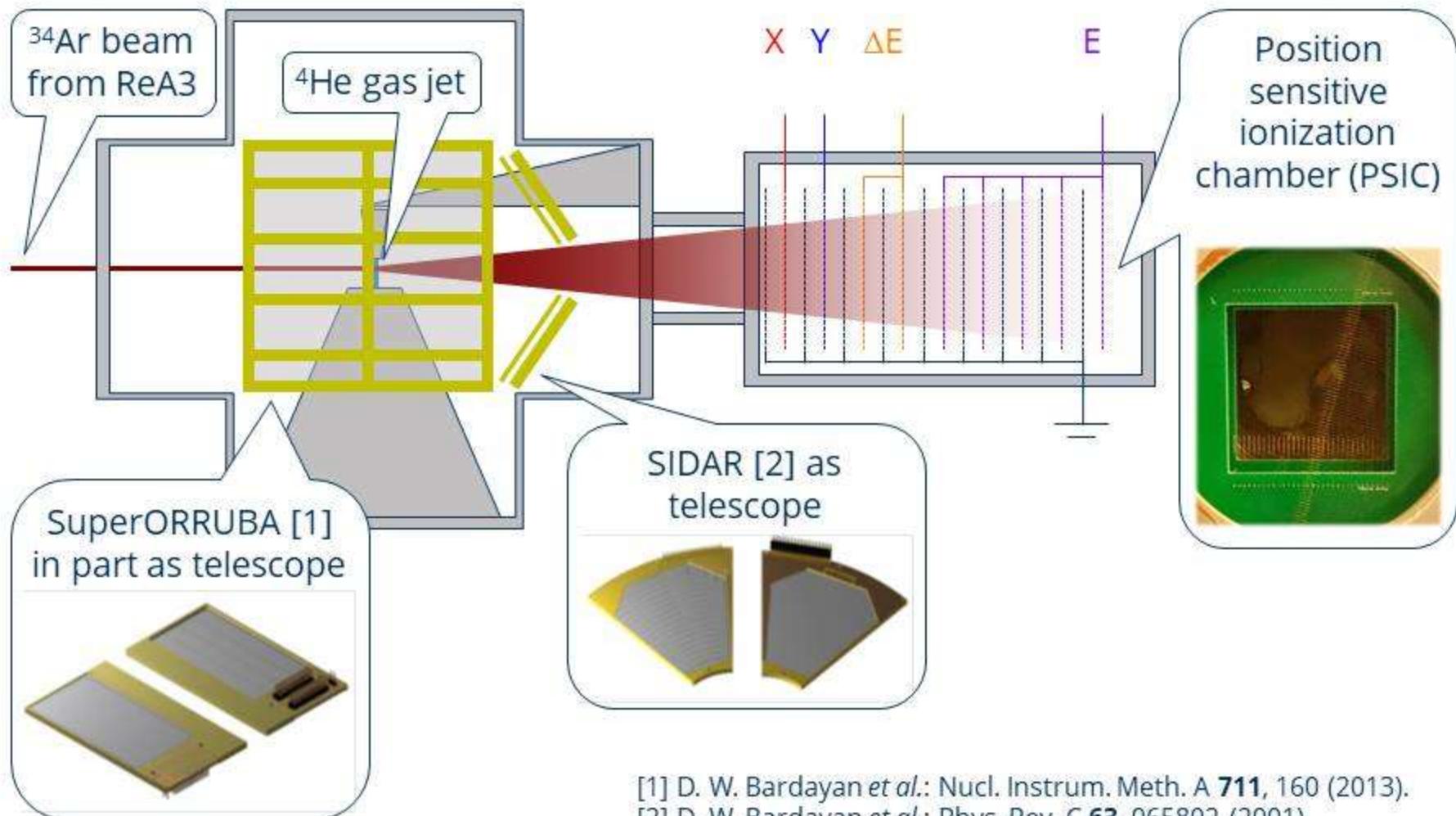
# Part 4

## Rare isotope beam experiments with JENSA

# X-ray burst reaction studies



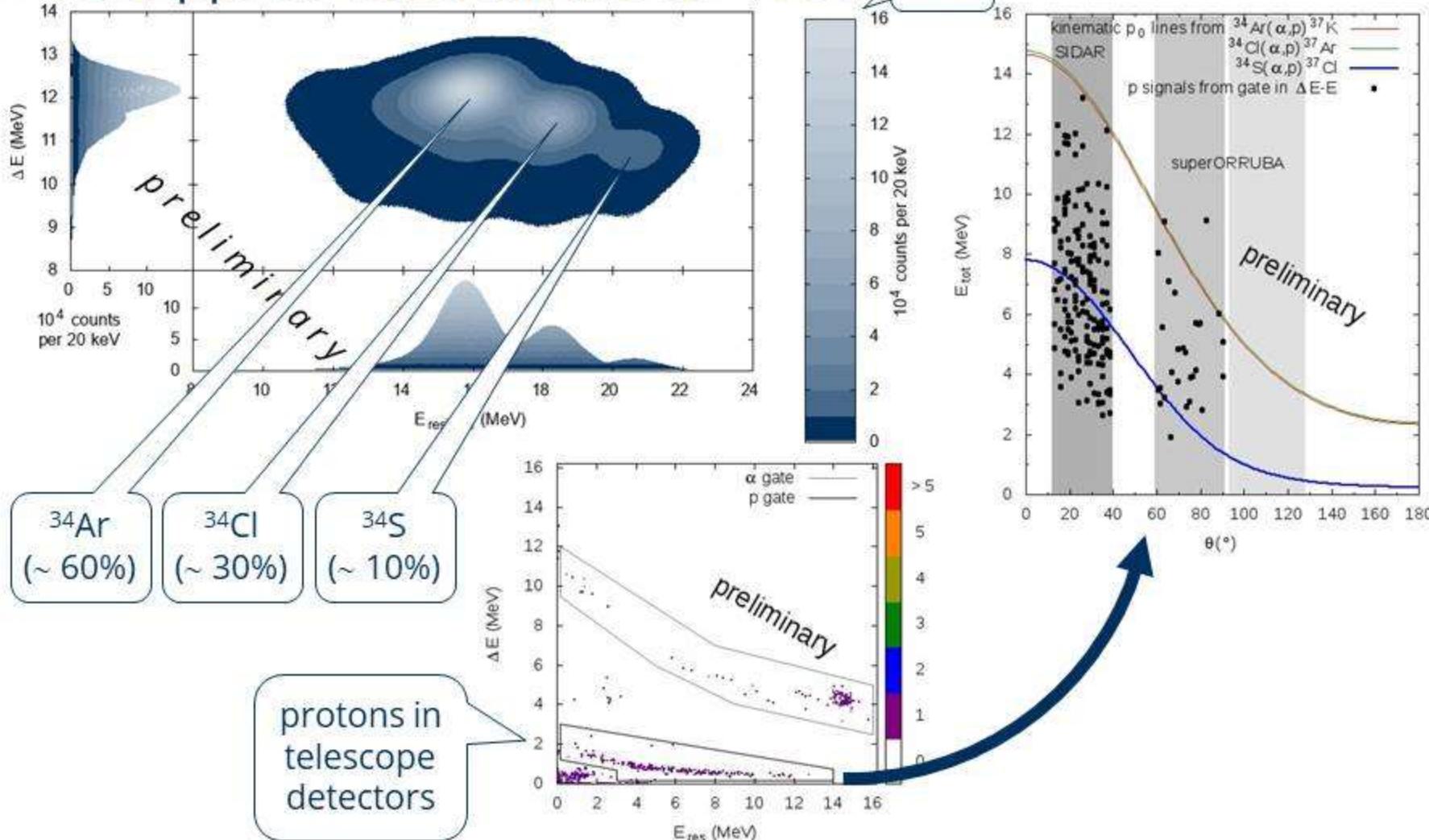
# Setup to study $^{34}\text{Ar}(\alpha, \text{p})^{37}\text{K}$



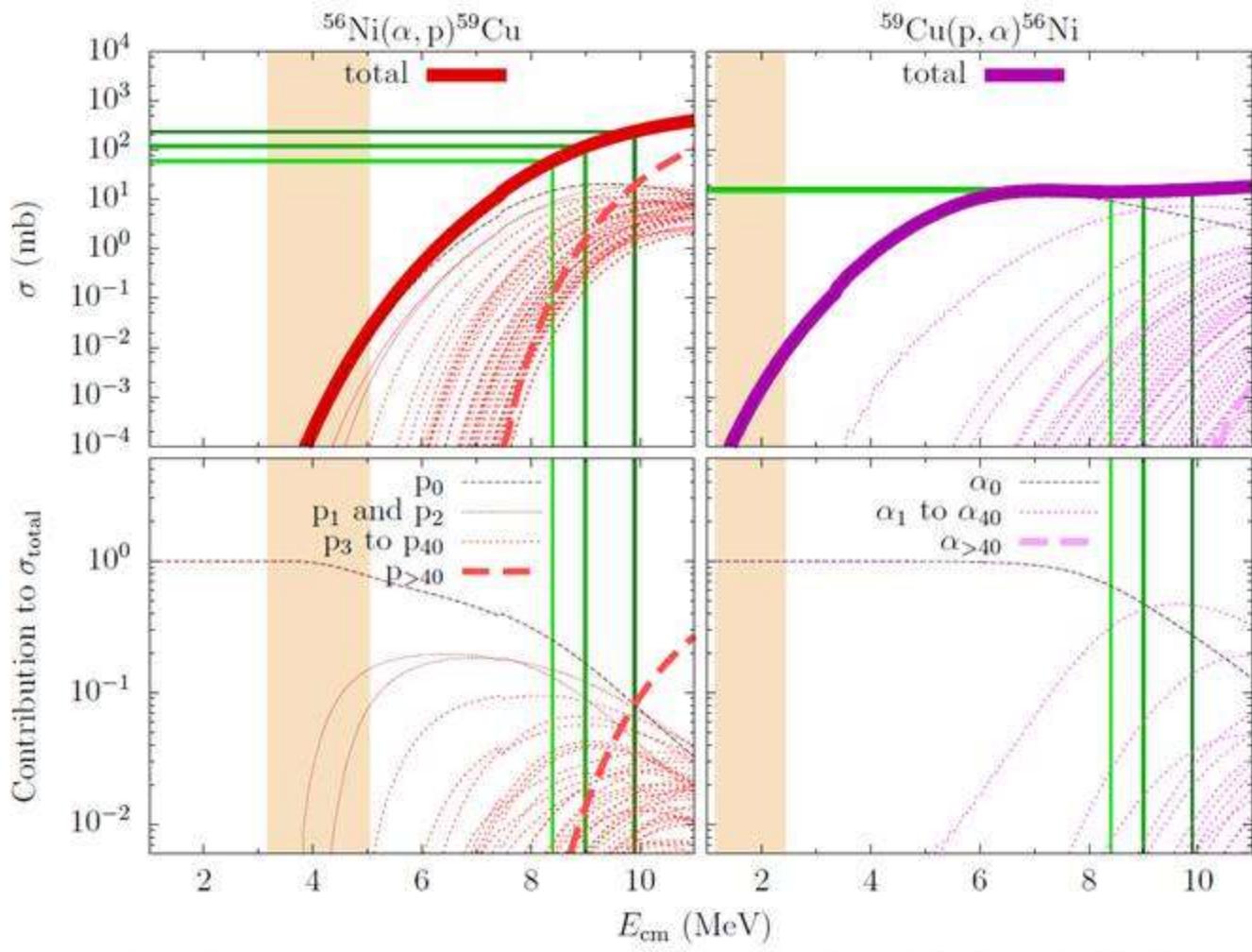
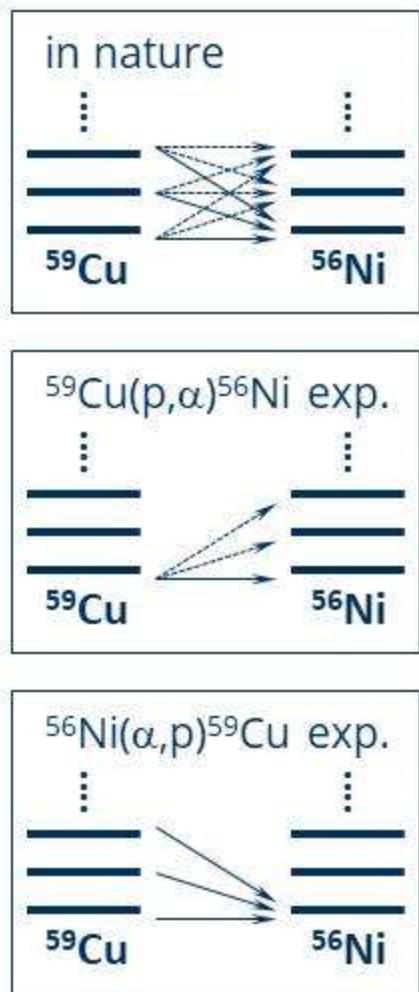
- [1] D. W. Bardayan *et al.*: Nucl. Instrum. Meth. A **711**, 160 (2013).  
[2] D. W. Bardayan *et al.*: Phys. Rev. C **63**, 065802 (2001).

# Beam composition and proton signals

~3000 pps at 1.625 MeV/u for 108 h



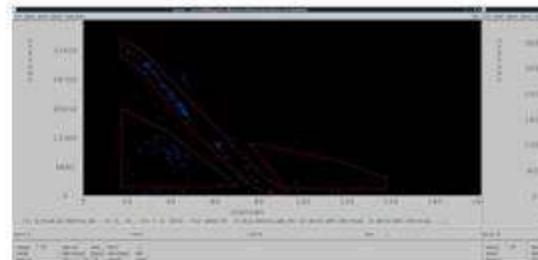
# The $^{56}\text{Ni}(\alpha, p)^{59}\text{Cu}$ reaction



forward and reverse cross sections are directly related via detailed balance

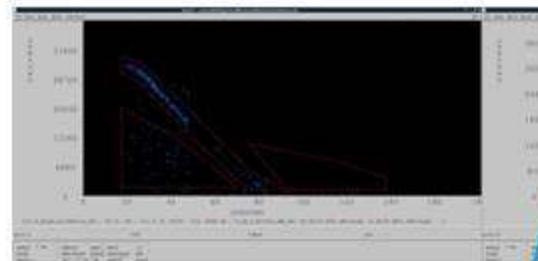
# Preliminary results from the online analysis (Dec. 2018)

2.65 MeV/u



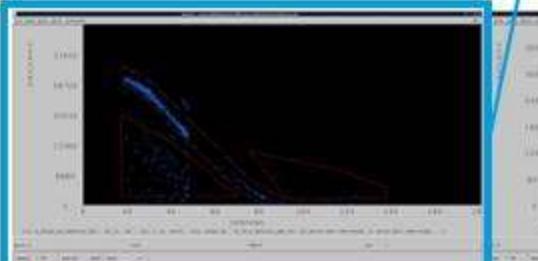
In the gates in the IC spectra, there 58 and 5 protons.

2.40 MeV/u

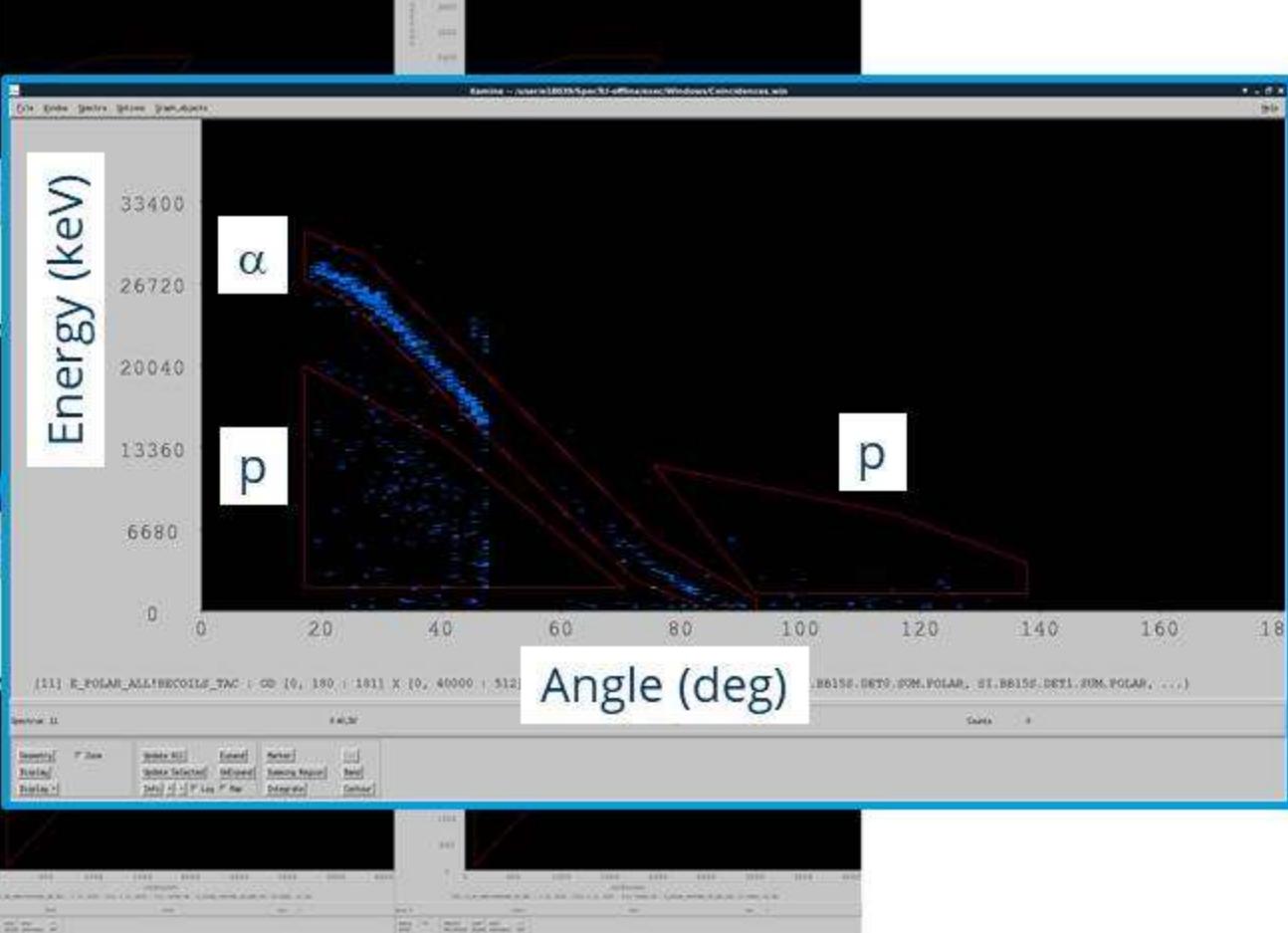


In the gates in the IC spectra, there 67 and 6 protons.

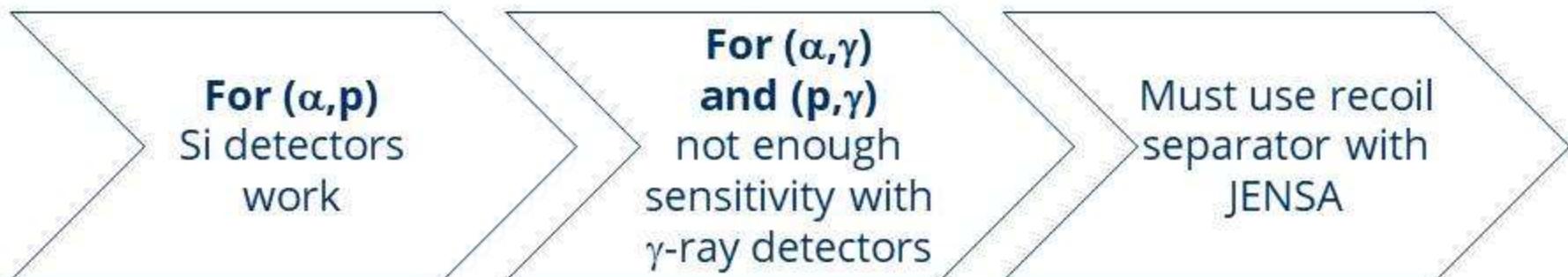
2.25 MeV/u



In the gates in the IC spectra, there 131 and 6 protons.



# Capture reaction studies with JENSA

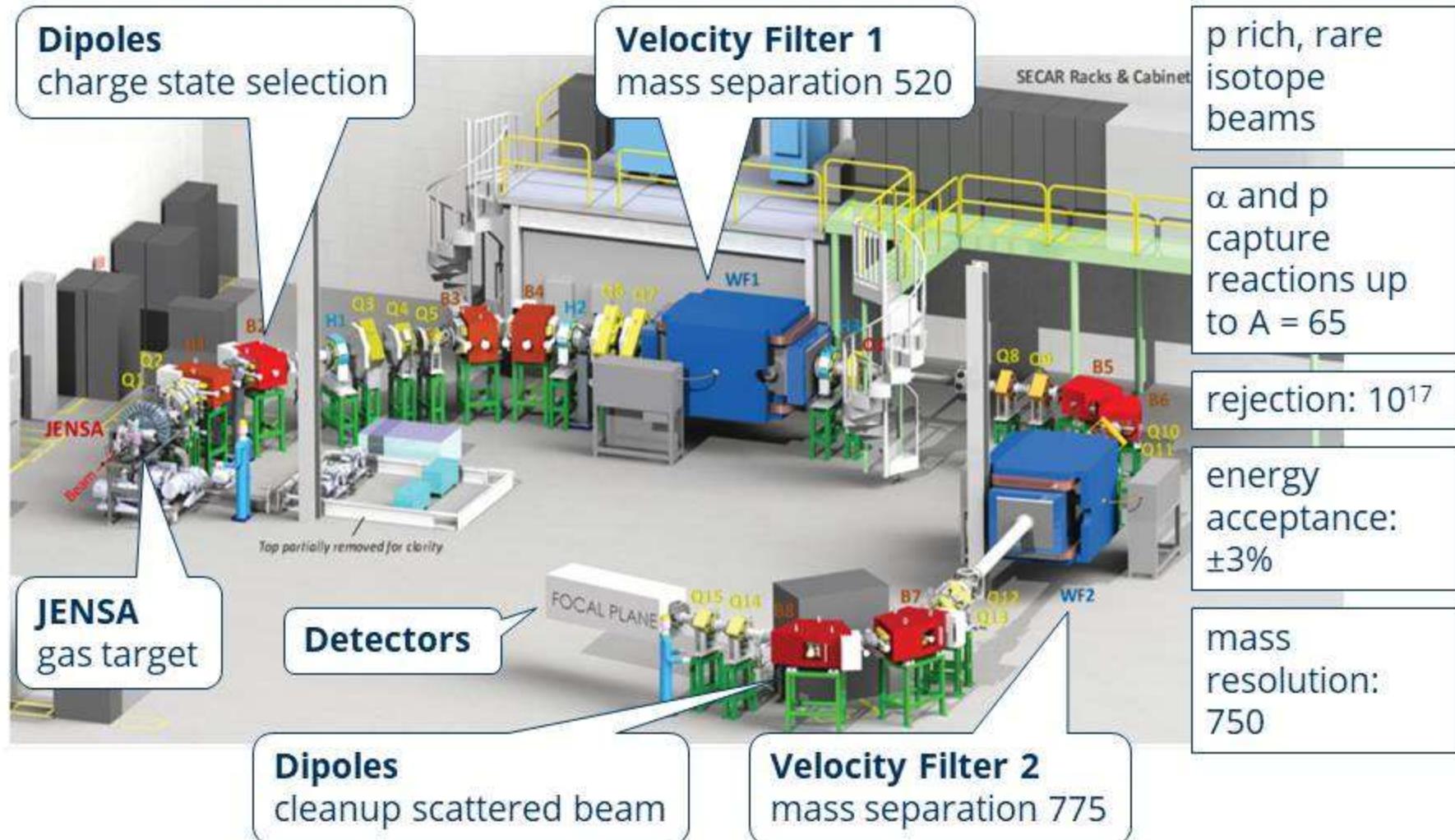


# Part 5

## The future of JENSA

# SECAR

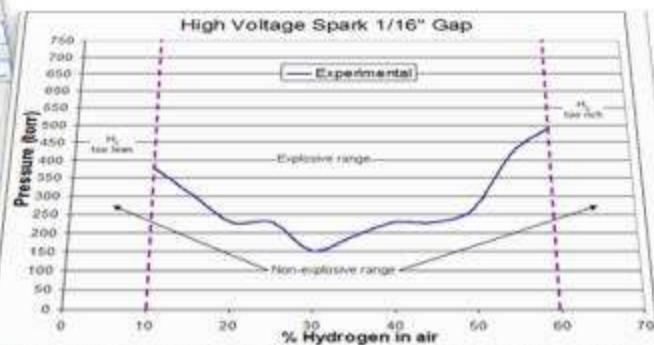
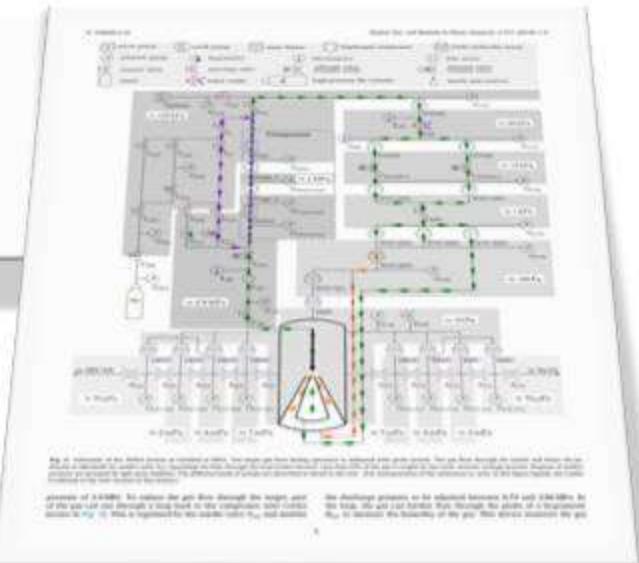
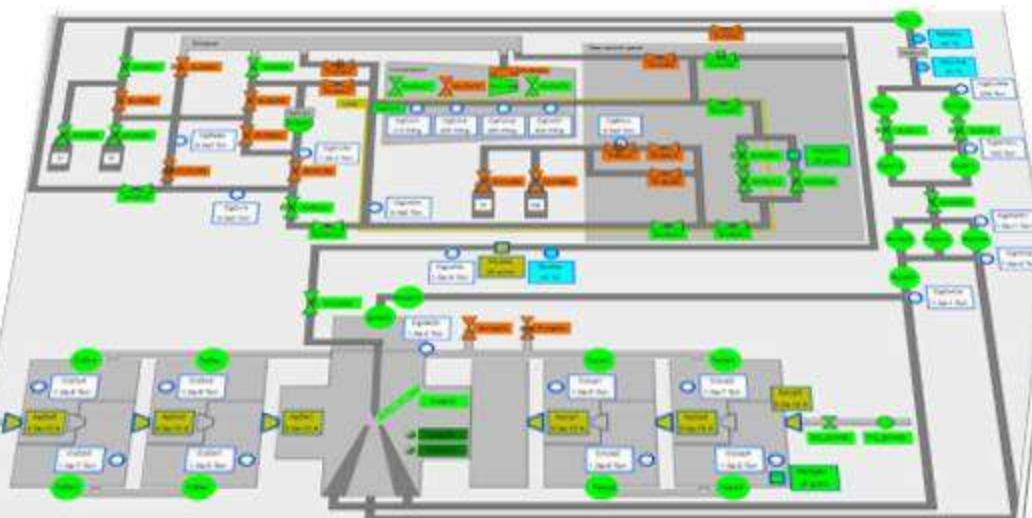
## Recoil Separator for Capture in Reactions



## JENSA is currently upgraded for $^3\text{He}$ and hydrogen

Facility for Rare Isotope Beams  
National Superconducting Cyclotron Laboratory  
*JENSA Operation with Hydrogen Activity Hazard Document*

# **JENSA Operation with Hydrogen Activity Hazard Document**

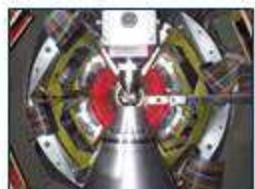


**Figure 5: Pressure and volumetric ratio conditions necessary to ignite a hydrogen-air mixture.** The figure was taken from an article published by The American Ceramic Society, "Materials Innovations in an Emerging Hydrogen Economy," Volume 202.

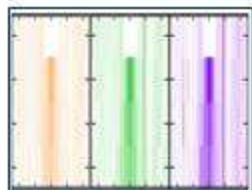
# Summary



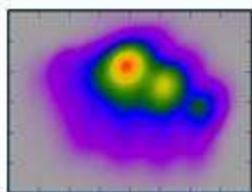
**JENSA** is able to study reactions important for nuclear astrophysics



**JENSA** has been successfully commissioned



**JENSA** is the world's densest gas-jet target in nuclear astrophysics with helium densities of  $10^{19}$  atoms/cm $^2$



**JENSA** enabled first direct reaction studies of  $^{34}\text{Ar}(\alpha,\text{p})^{37}\text{K}$  and  $^{56}\text{Ni}(\alpha,\text{p})^{59}\text{Cu}$



**JENSA** will be the main target for SECAR enabling  $(\alpha,\gamma)$  and  $(\text{p},\gamma)$  capture reaction measurements