



Commissioning with cosmic rays of the ALICE Muon Trigger System

Claudio Geuna

Journées de Rencontre Jeunes Chercheurs 2009





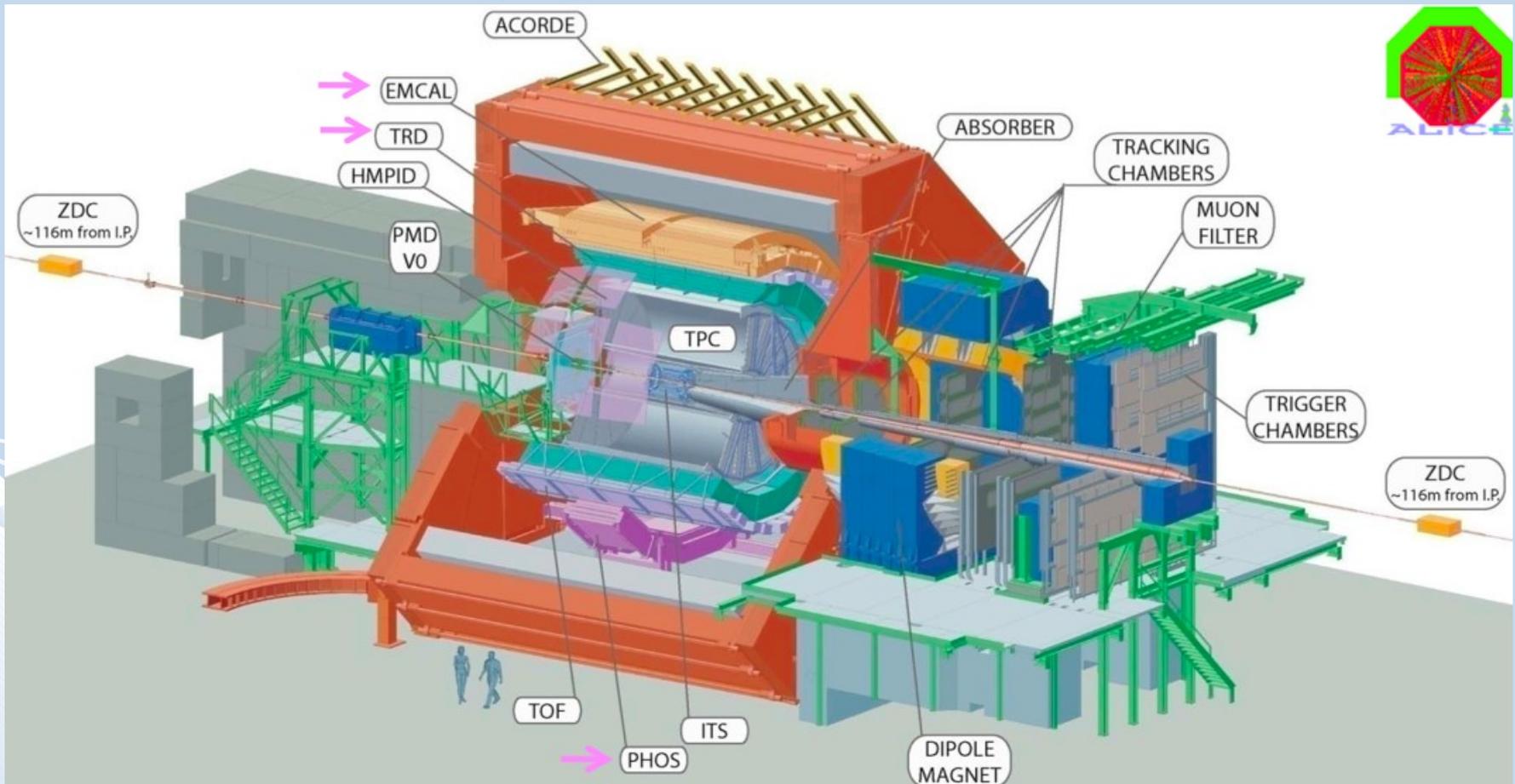
OUTLINE:

- 1) Overview of the trigger system of the ALICE Muon Spectrometer.**
- 2) Commissioning measurements and results.**



Alice experiment

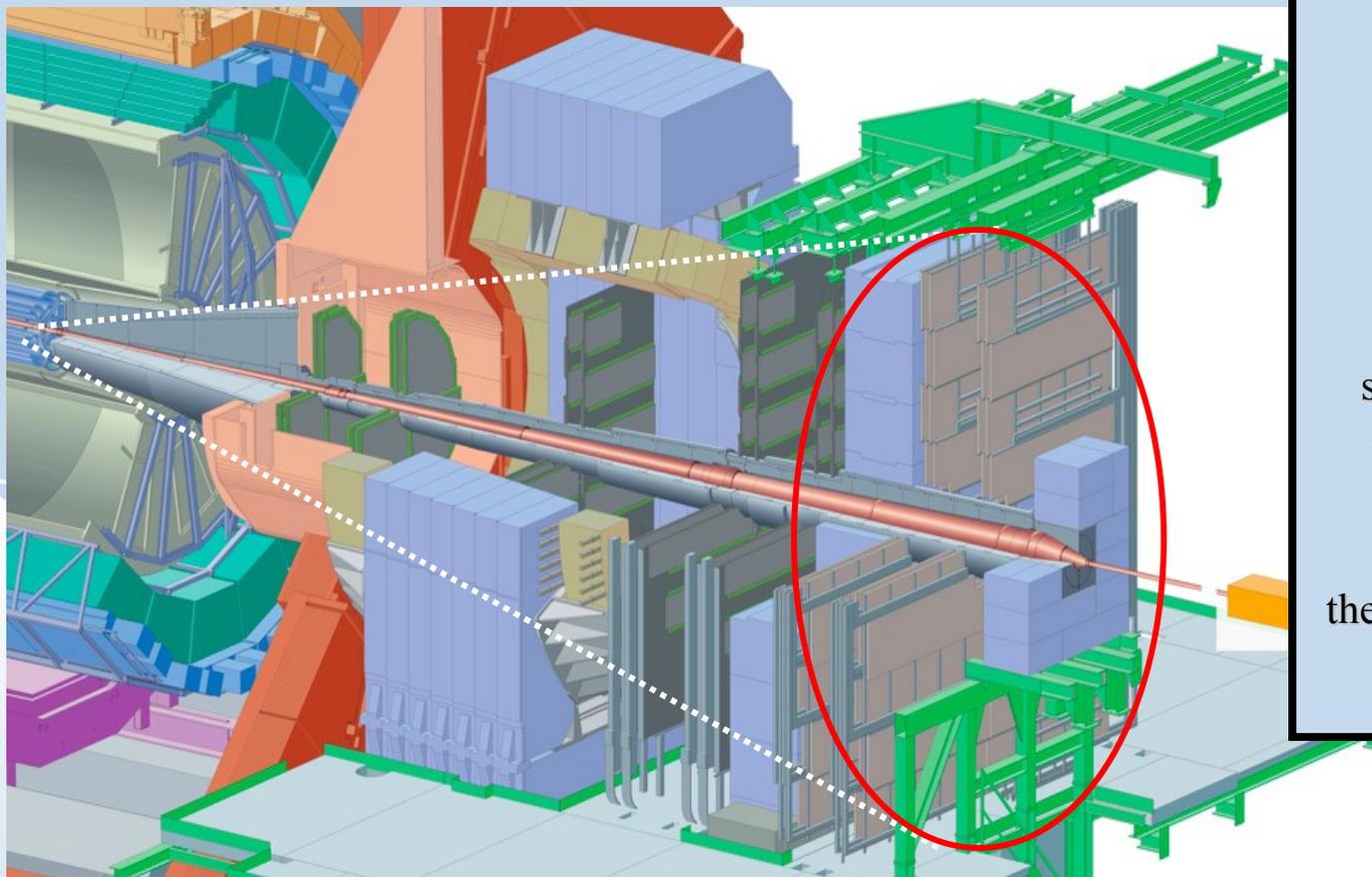
A Large Ion Collider Experiment



ALICE experiment @LHC is specifically dedicated to ultra-relativistic heavy-ion collisions



The Muon Spectrometer (I)



Forward muon spectrometer

Angular acceptance

$$2^{\circ} < \theta < 9^{\circ}$$

Pseudorapidity

$$-4 < \eta < -2.5$$

study of the production of open heavy flavour and heavy quarkonia (J/Ψ and Υ) through the muon (μ) decay channel



The Muon Spectrometer (II)

the components

Interaction vertex

Absorber

Dipole Magnet

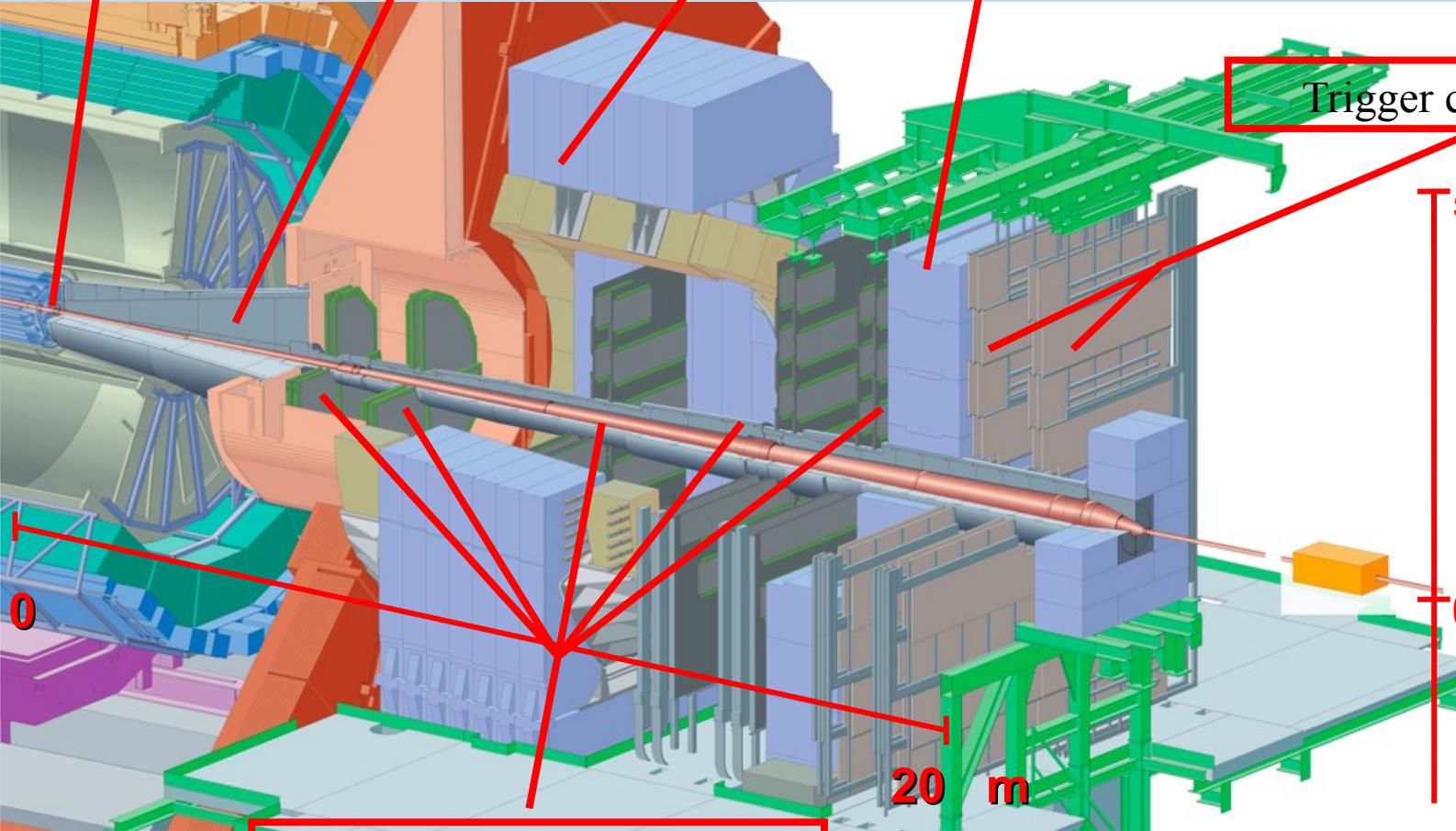
Muon filter

Trigger chambers

5 m

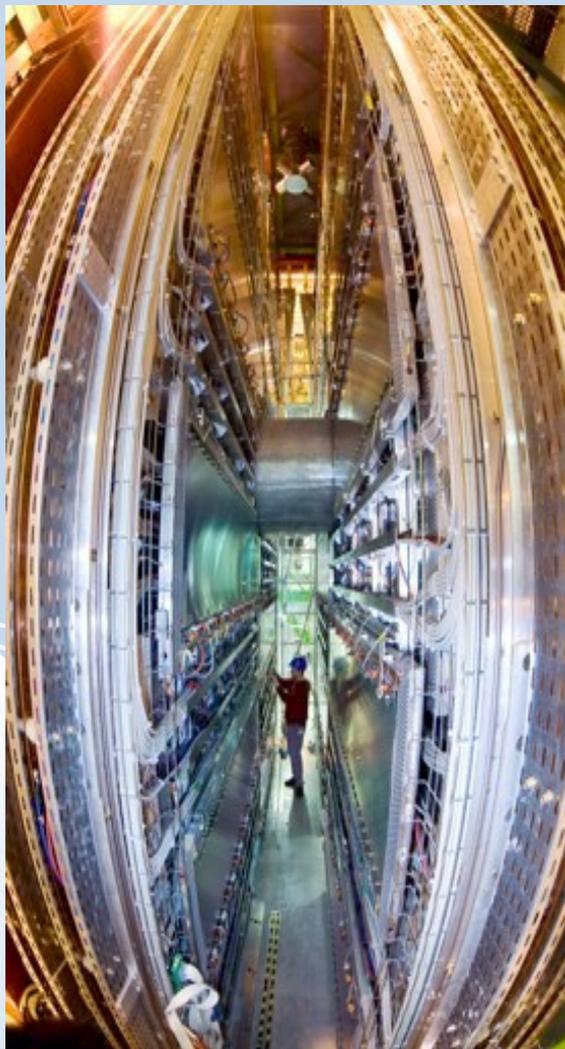
Tracking chambers

20 m

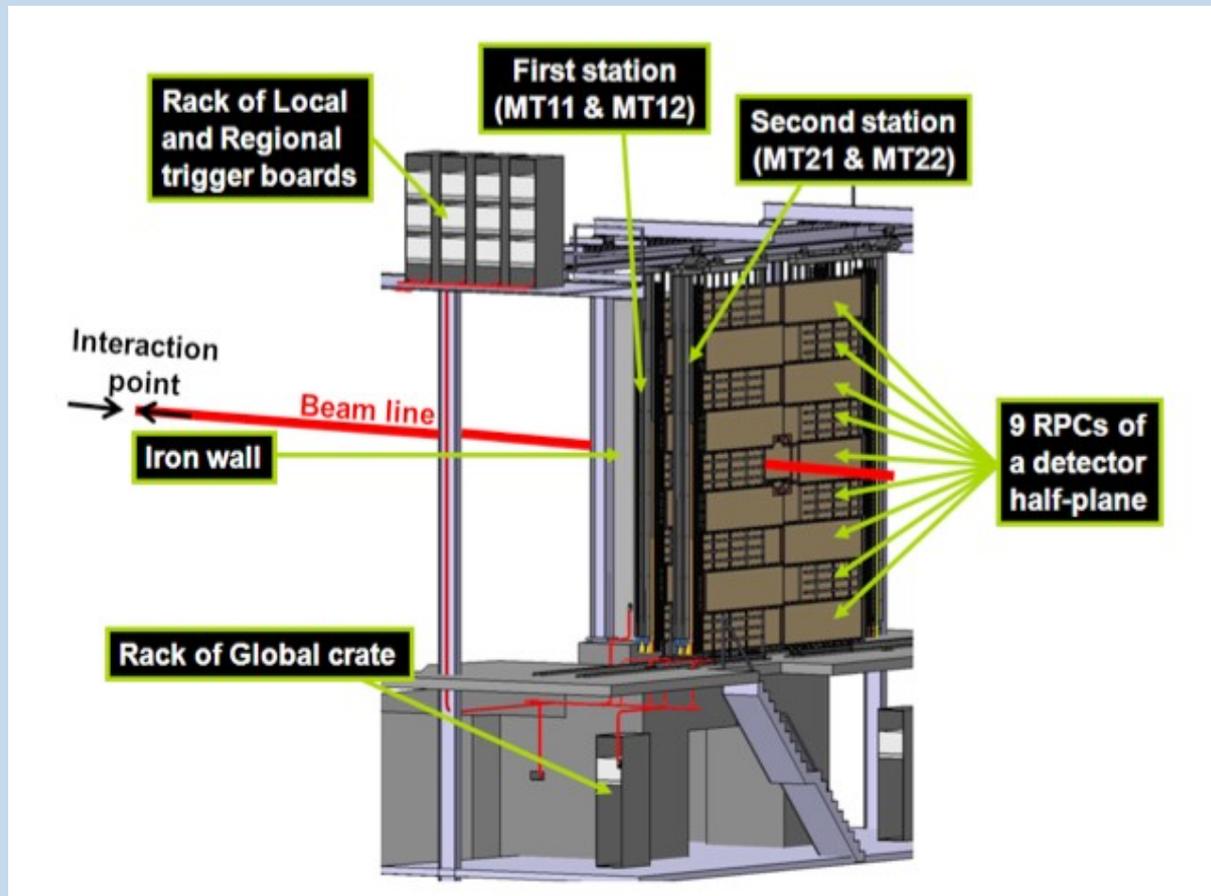




The trigger system: Muon Trigger Chambers



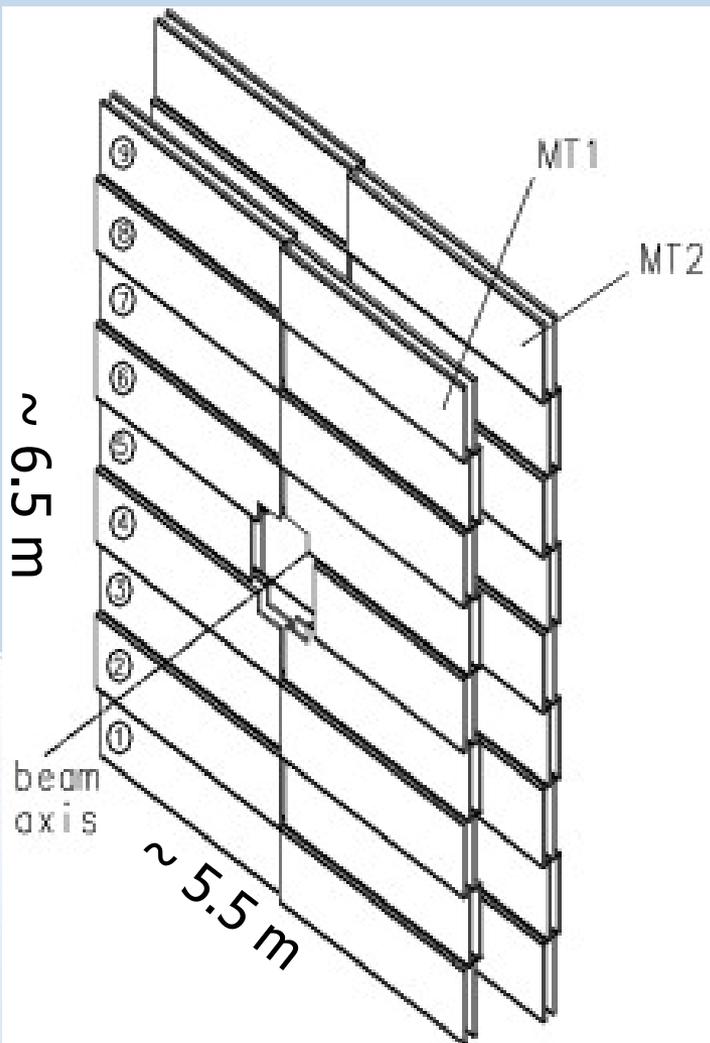
2 trigger stations (MT1 and MT2) located at about 16 m from the IP and 1 m apart from each other placed behind an iron muon filter



Layout of the trigger stations and the iron wall installed in the Alice cavern.



The trigger system: Muon Trigger Chambers



2 stations, of two planes each

total area: ~ 140 m²

72 RPCs of 3 different shapes

and dimensions :

~ 1.6 m² \div 2.1 m²

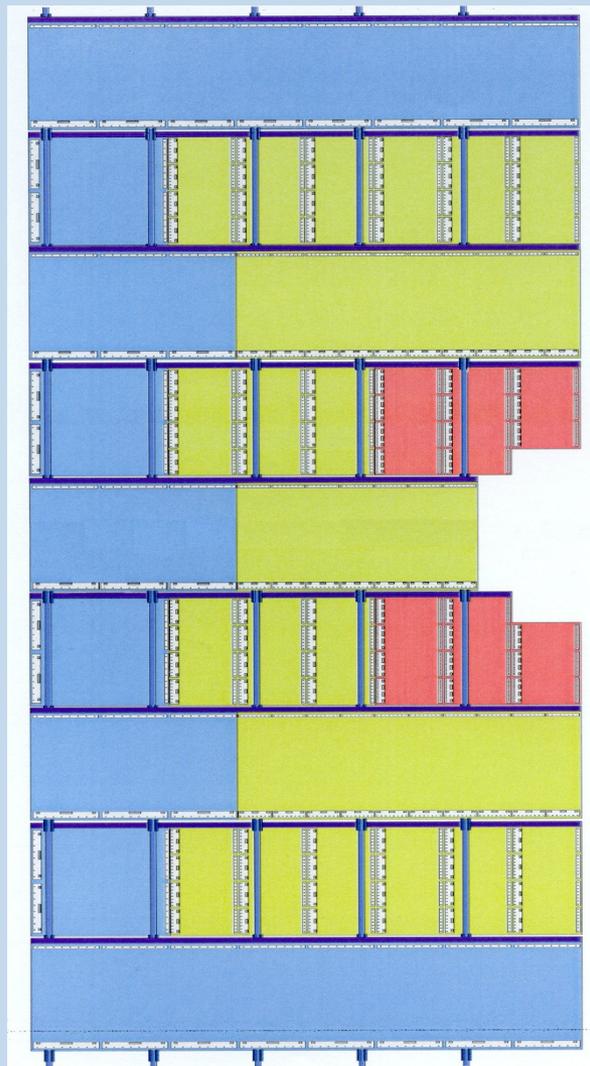
20992 strips and front-end channels

strip pitch:

10-45 mm ($\sim 1/2/4$ cm)

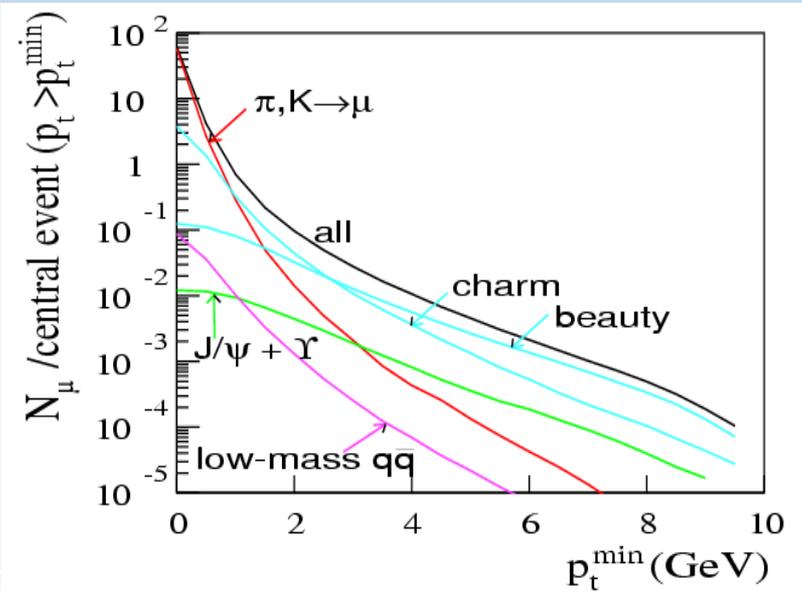
strip length:

170 \div 720 mm



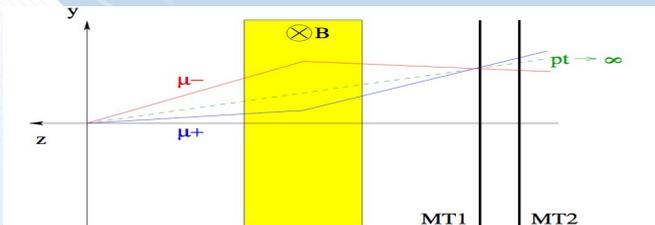


The principle of the trigger (I)



GOAL:

- Selection of (muon) tracks pointing to I.P. with p_t above 2 thresholds:
 - low $p_{t \text{ cut}}$ = 1 GeV/c
 - high $p_{t \text{ cut}}$ = 2 GeV/c
- Trigger signals for single μ , like-sign and unlike-sign μ pairs

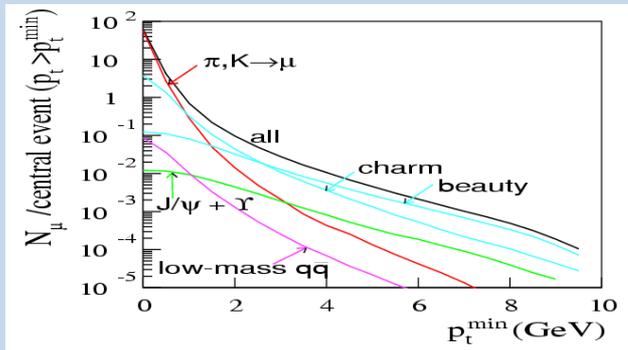


PRINCIPLE:

- cut on $p_t \Leftrightarrow$ cut on deviation between MT1 and MT2
 → select tracks in a road of a given width

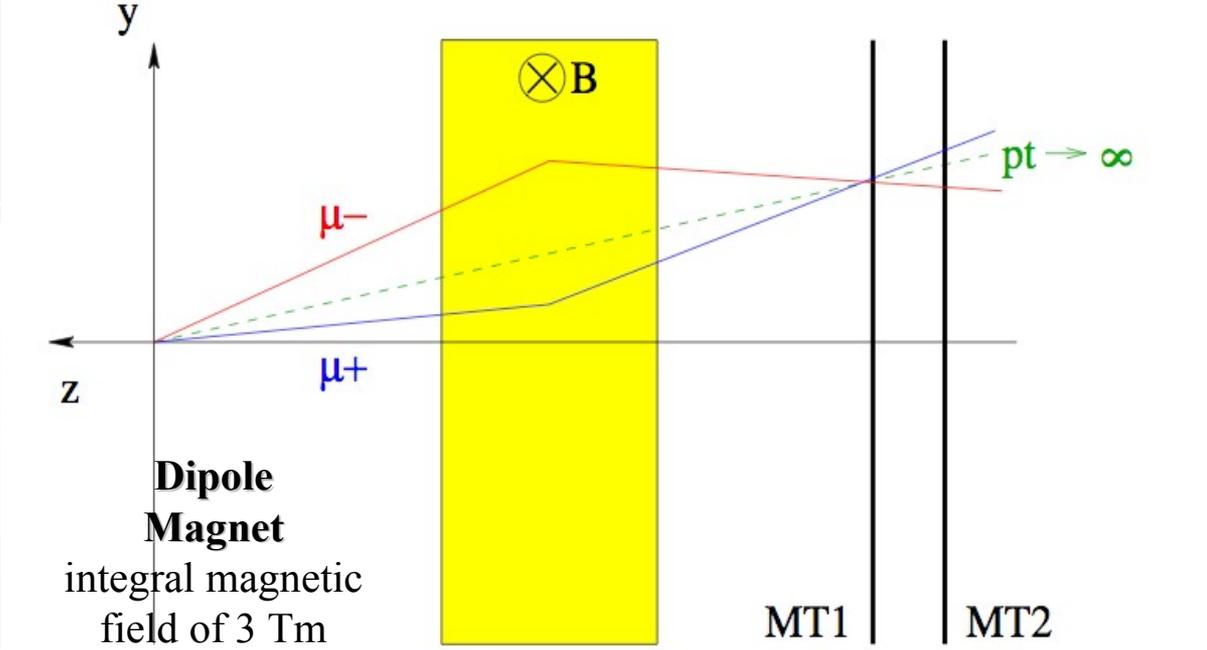


The principle of the trigger (II)



GOAL:

- Selection of (muon) tracks pointing to I.P. with p_t above 2 thresholds:
 - $low\ p_{t\ cut} = 1\ GeV/c$
 - $high\ p_{t\ cut} = 2\ GeV/c$
- Trigger signals for single μ , like-sign and unlike-sign μ pairs



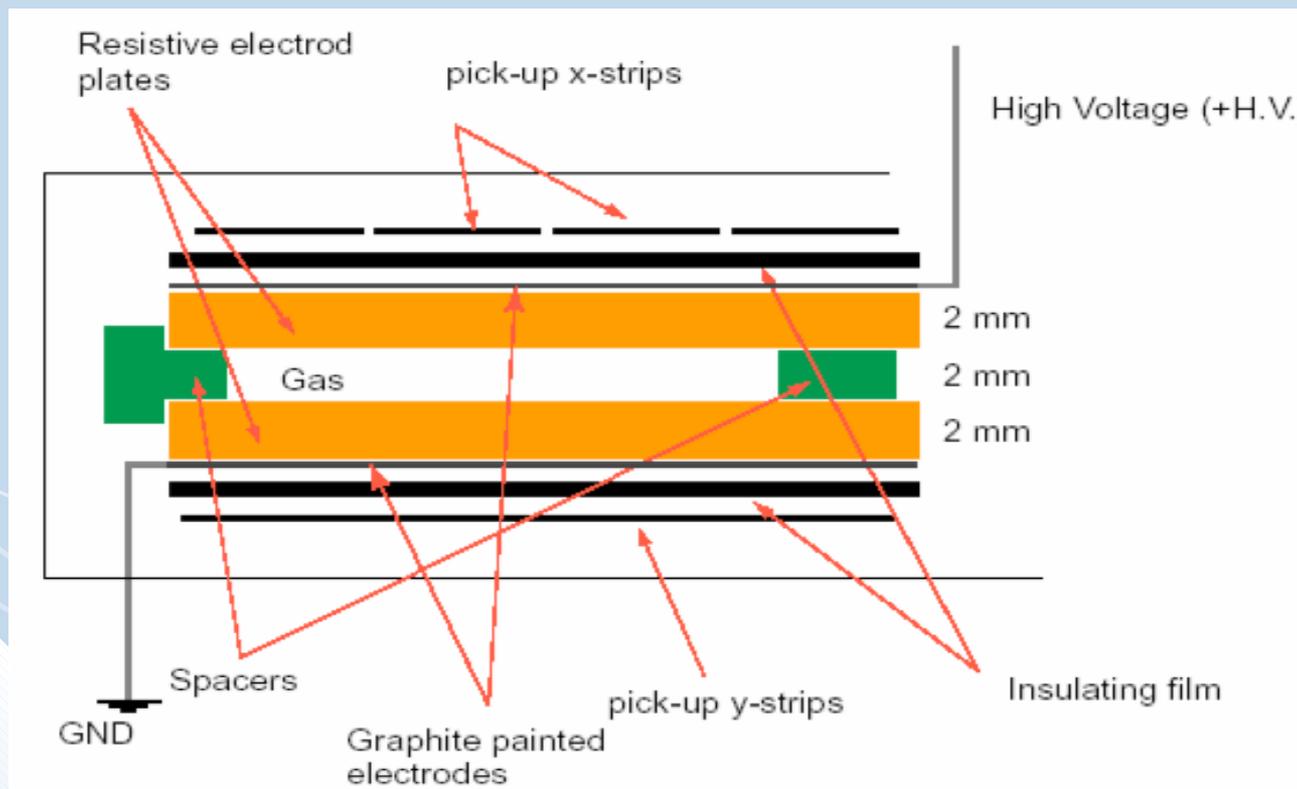
PRINCIPLE:

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Resistive Plate Chambers (I)

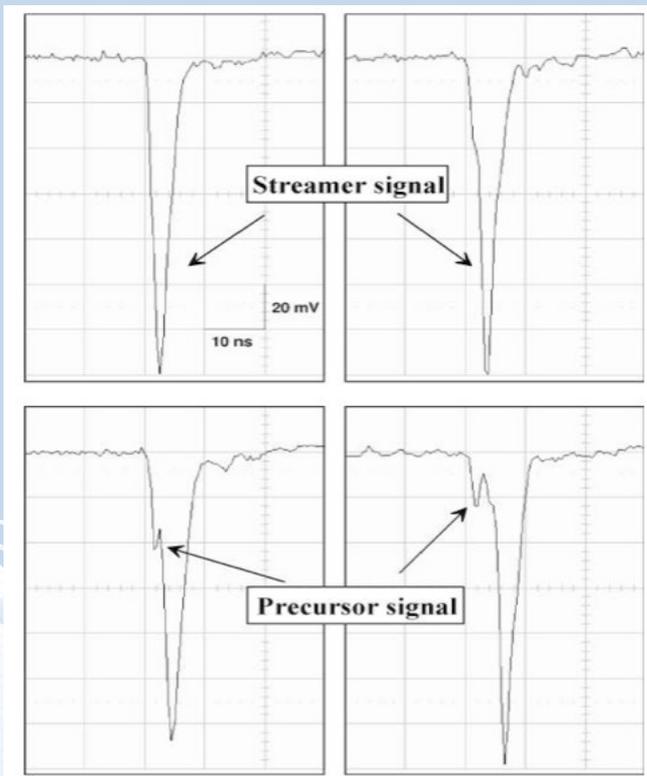
A Resistive Plate Chamber (RPC) is a planar geometry gaseous detector





Resistive Plate Chambers (II)

Gas mixtures: avalanche vs streamer



The **difference** between the streamer and avalanche modes lies in the **gas mixture (quenchers)** and **HV** applied between the two electrodes.

Main advantages of streamer and avalanche RPC operation

Operation mode	Advantages
Streamer	<i>Spatial resolution</i>
	<i>No amplification needed</i>
	<i>Lower noise rate</i>
Avalanche	<i>Time resolution</i>
	<i>Rate capability</i>
	<i>Slower ageing</i>

ALICE MuonTrigger RPCs

→ Typical avalanche operation voltages ~ 10 kV

→ Typical streamer operation voltages ~ 8 kV



Resistive Plate Chambers (III)

ALICE : both *A-A* and *p-p* data-taking

Requirements for A-A data taking:

- *Spatial resolution* ~ 1 cm
- *Occupancy* as low as possible (few % Pb-Pb) and *cluster-size* as close as possible to 1
- *Rate capability* ~ 3 Hz/cm² (Pb-Pb) and ~ 25 Hz/cm² (Ar-Ar)
- *Time resolution* ~ 2 ns

Requirements for p-p data taking:

- Expected *muon trigger rate* much lower than in A-A collisions
- *Rate capability* ~ 100 Hz/cm²
- Goal: *detector lifetime*



Resistive Plate Chambers (IV)

RPC performances

Rate capability up to ~ 100 Hz/cm² (p-p)



Time resolution ~ 2 ns



Low resistivity bakelite:

$$\rho = 2 \div 8 \cdot 10^9 \text{ } \Omega\text{cm}$$

Dual threshold

Front End Electronics

10 mV - 80 mV



Resistive Plate Chambers (IV)

RPC performances

Rate capability up to ~ 100 Hz/cm² (p-p) \longrightarrow

Time resolution ~ 2 ns \longrightarrow

Low resistivity bakelite:

$$\rho = 2 \div 8 \cdot 10^9 \text{ } \Omega\text{cm}$$

Dual threshold

Front End Electronics

10 mV - 80 mV

Requirements for A-A data taking:

Spatial resolution ~ 1 cm

Occupancy as low as possible ($< 1\%$ Pb-Pb) and *cluster-size* as close as possible to 1

Streamer mode



Resistive Plate Chambers (IV)

RPC performances

Rate capability up to ~ 100 Hz/cm² (p-p)



Low resistivity bakelite:

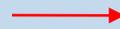
$$\rho = 2 \div 8 \cdot 10^9 \text{ } \Omega\text{cm}$$

Dual threshold

Front End Electronics

10 mV - 80 mV

Time resolution ~ 2 ns



Requirements for A-A data taking:

Spatial resolution ~ 1 cm

Occupancy as low as possible ($< 1\%$ Pb-Pb) and *cluster-size* as close as possible to 1



Streamer mode

Requirements for p-p data taking:

Expected *muon trigger rate* much lower than in A-A collisions

Goal: *detector lifetime*



“Highly saturated” avalanche mode



Resistive Plate Chambers (V)

The two gas mixtures in detail.....

The gas mixture used will be different for A-A and p-p collisions.

A-A collisions (a wet, low-gain streamer mixture will be used)

50.5% Ar 41.3% $C_2H_2F_4$ 7.2% C_4H_{10} 1% SF_6

p-p collisions (a wet, highly-saturated avalanche mixture will be used)

89.7% $C_2H_2F_4$ 10% C_4H_{10} 0.3% SF_6



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- 2) Commissioning measurements and results**



Muon Spectrometer commissioning with cosmic rays: goals

- **Global test of the Muon Spectrometer**



Muon Spectrometer commissioning with cosmic rays: goals

- Global test of the Muon Spectrometer
- **Test of the Trigger Chambers and of the Tracking Chambers separately**



Muon Trigger commissioning with cosmic rays: goals

- Global test of the Muon Spectrometer
- Test of the **Trigger Chambers** and of the Tracking Chambers separately



Detectors

Dark current and rate measurements

RPCs working point

Electronics

test of Front-End and trigger electronics

DAQ, DCS



Muon Trigger commissioning with cosmic rays: goals

- Global test of the Muon Spectrometer
- Test of the **Trigger Chambers** and of the Tracking Chambers separately



Detectors

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RPCs working point

Electronics

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DAQ, DCS

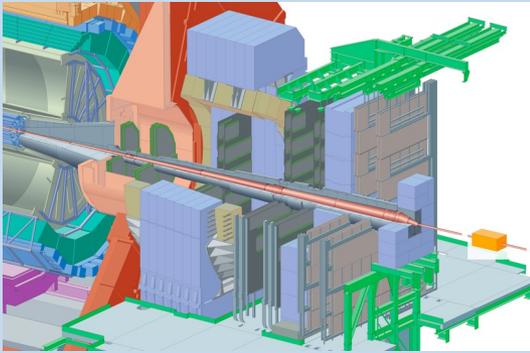


RPC working point in streamer

RPC working point in avalanche

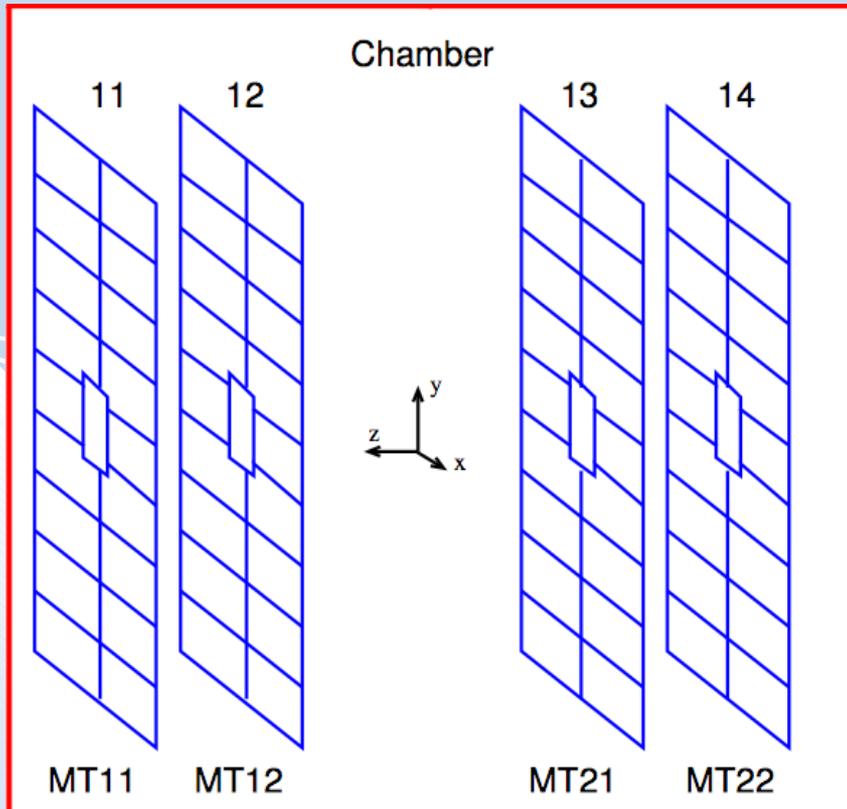


Trigger chamber efficiency (I)



Due to : *geometrical acceptance* and *projective geometry trigger* (the tracks triggered have to come from the Interaction Point IP)

the ALICE Muon Spectrometer is not designed to detect cosmic rays

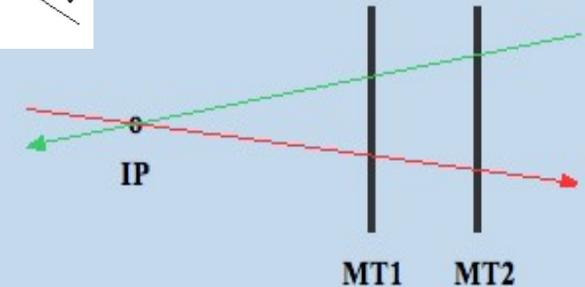
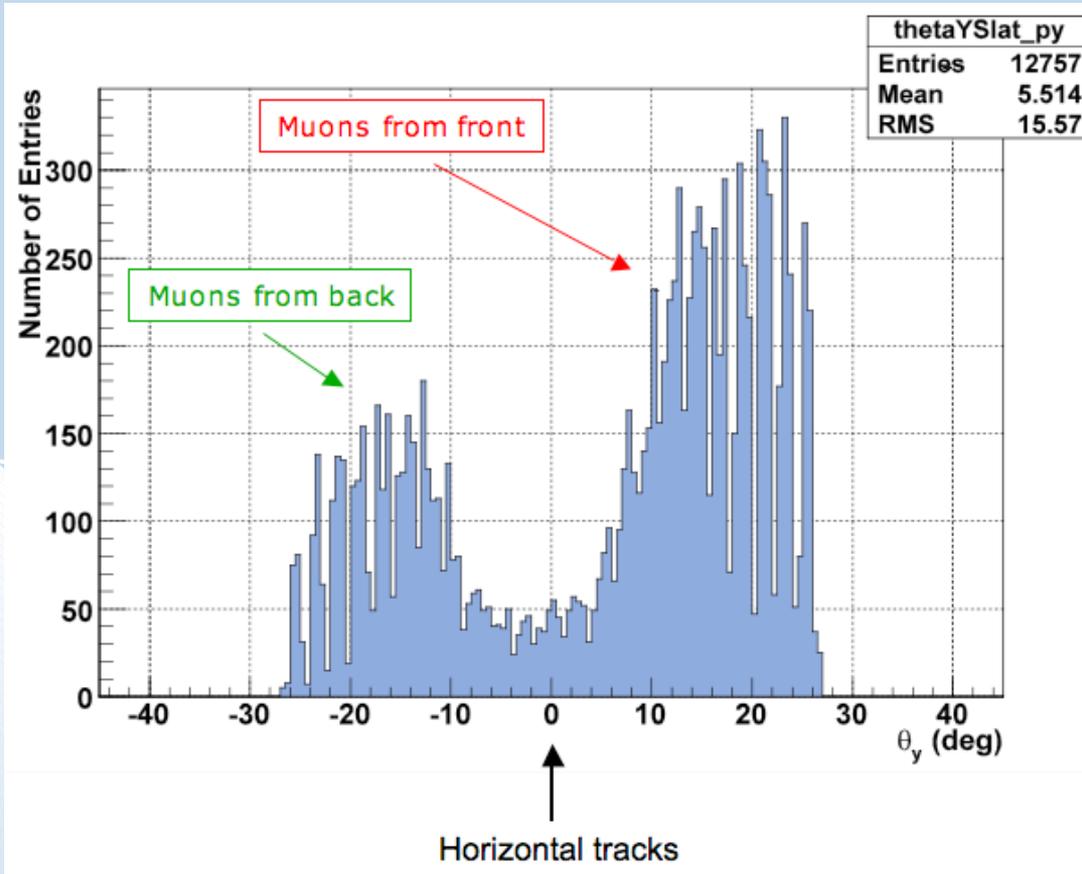
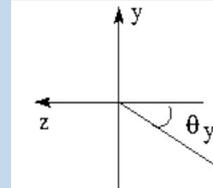


To study the *global features* of the system, the nearly horizontal tracks triggered are very useful for our aims



Trigger chamber efficiency (II)

angular distribution of detected cosmic muons



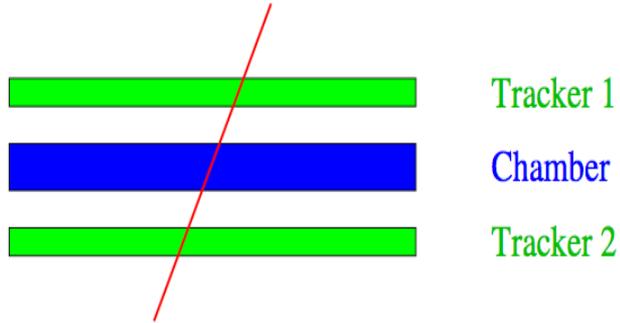
- θ_y *angular distribution* reflects the cosmic muon zenithal angle distribution
- *Distribution not symmetric:* lower efficiency for muons from the back due to difference in timing between the two stations



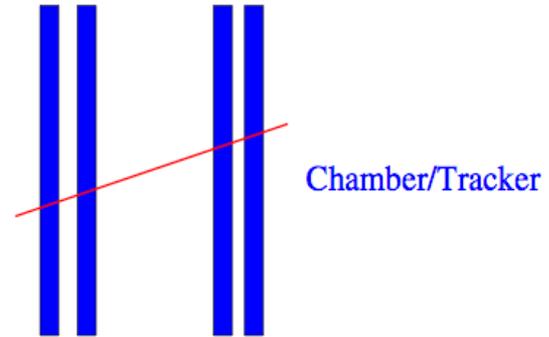
Trigger chamber efficiency (III)

Experimental set-up

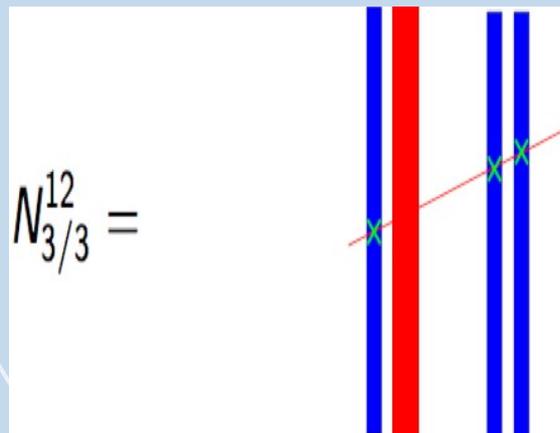
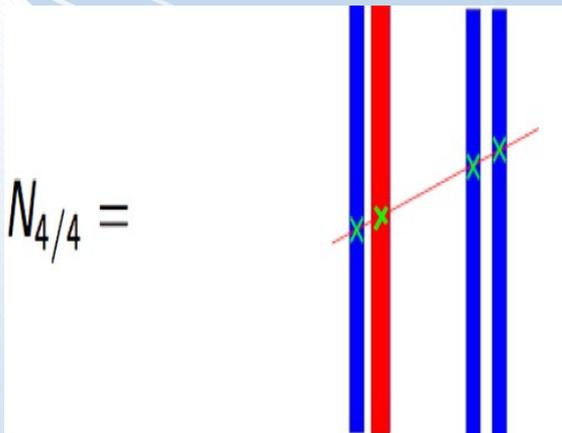
In laboratory:



Muon Spectrometer:



The trigger algorithm searches for hits in at least 3 out of 4 trigger chambers. We define:



The *efficiency* for plane *a* is given by

$$\epsilon_a = \frac{N_{4/4}}{N_{3/3}^a + N_{4/4}}$$



Trigger chamber efficiency (IV)

Selection of *tracks* in data analysis for efficiency evaluation

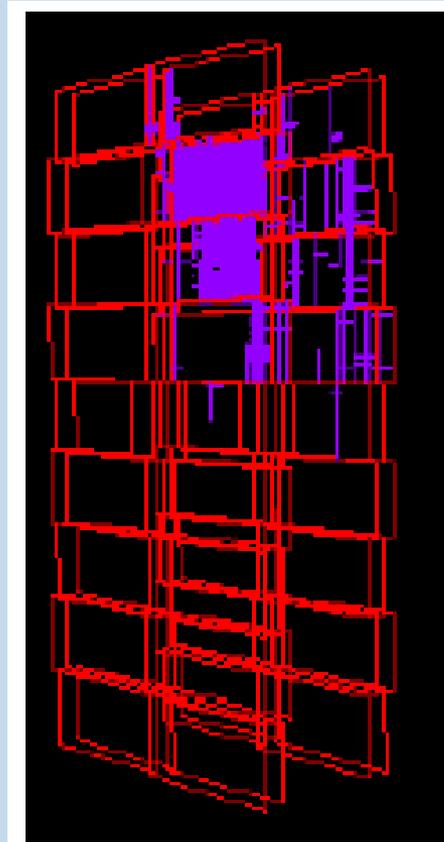
- 1) ~ 60% shower
- 2) ~ 40% single tracks



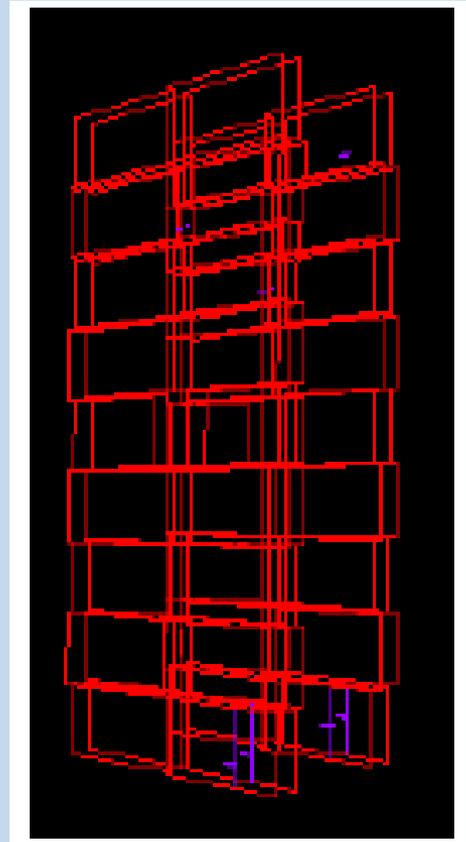
used in efficiency calculation

Total trigger rate :

- ~ 0.18 Hz (Streamer)
- ~ 0.23 Hz (Avalanche)



Event display of a cosmic shower



Event display of a cosmic muon



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- Tests in Streamer mode (March-April 2009)

- Tests in Avalanche mode (August-September 2009)



RPC status: efficiency vs. H.V.

RPC working point in **streamer**:

Cosmics run (Mar-Apr 2009):

The goal is to find the working voltage by H.V. scan from a “nominal” H.V. value (corresponding to H.V. = 0 V in the plot)

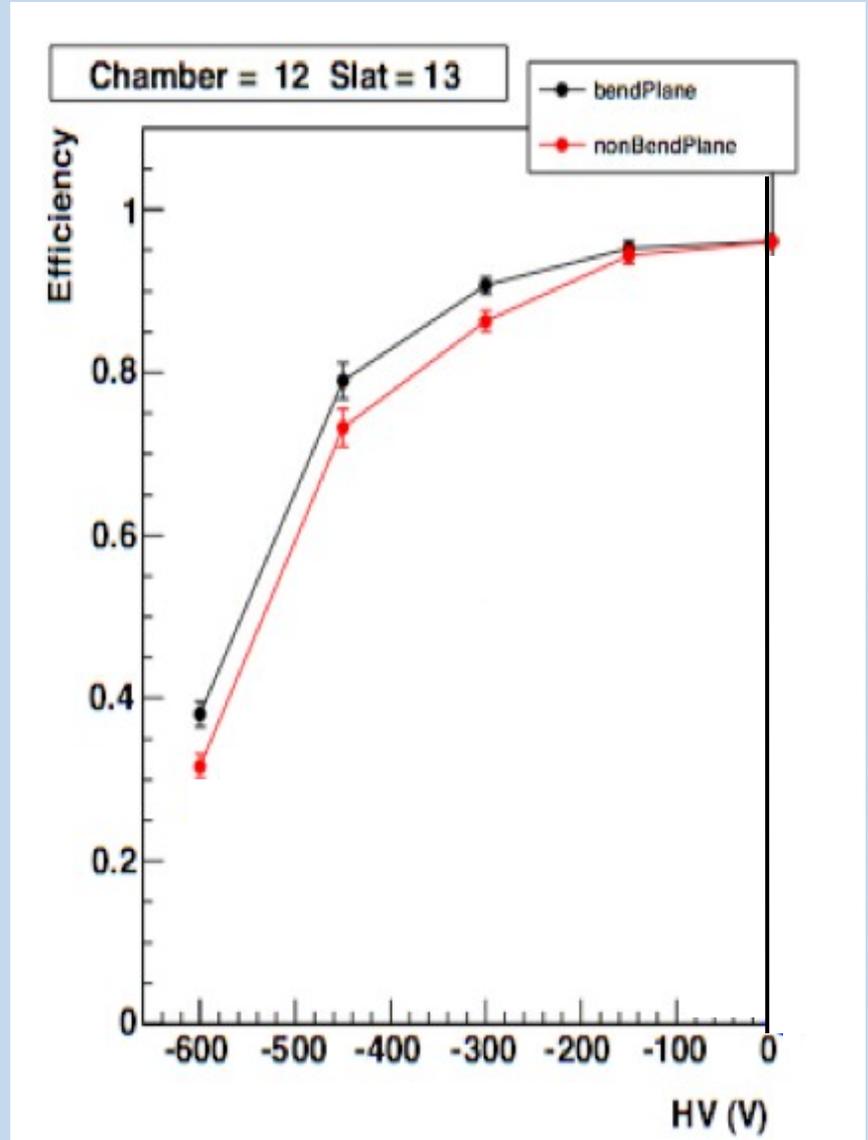
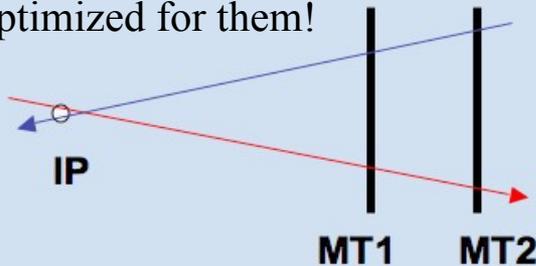
determined during a period of preliminary tests.

—————→

CAVEAT

Difficult to determine the absolute value of RPC efficiency due to specific cosmic run conditions:

- 1) Low statistics
- 2) Systematic effects : cosmics from the direction opposite to IP. The timing between the two trigger stations is not optimized for them!

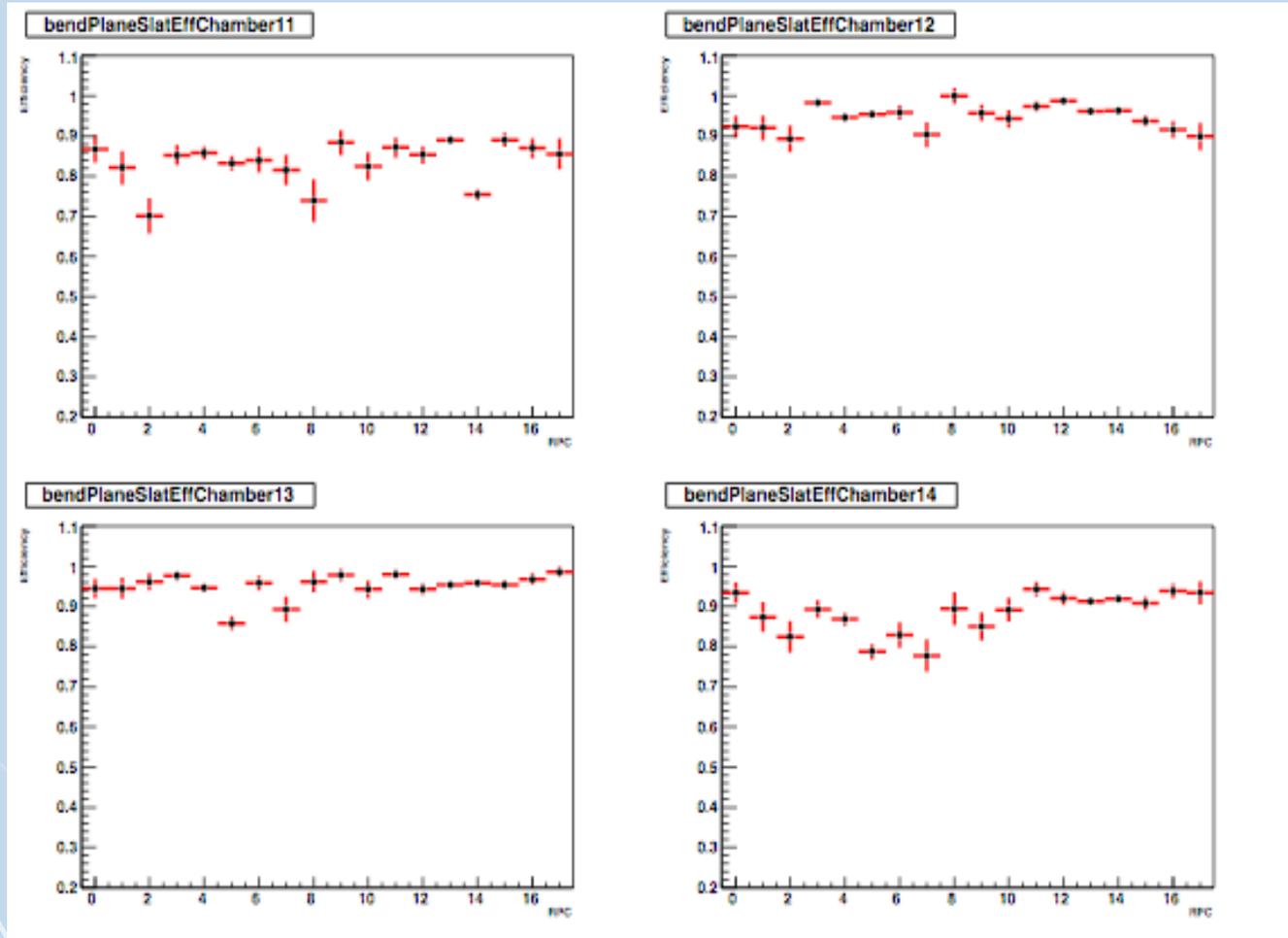




“Efficiency” in the bending plane (per slat) for all the 4 stations (streamer mixture)



Check of the status of the system

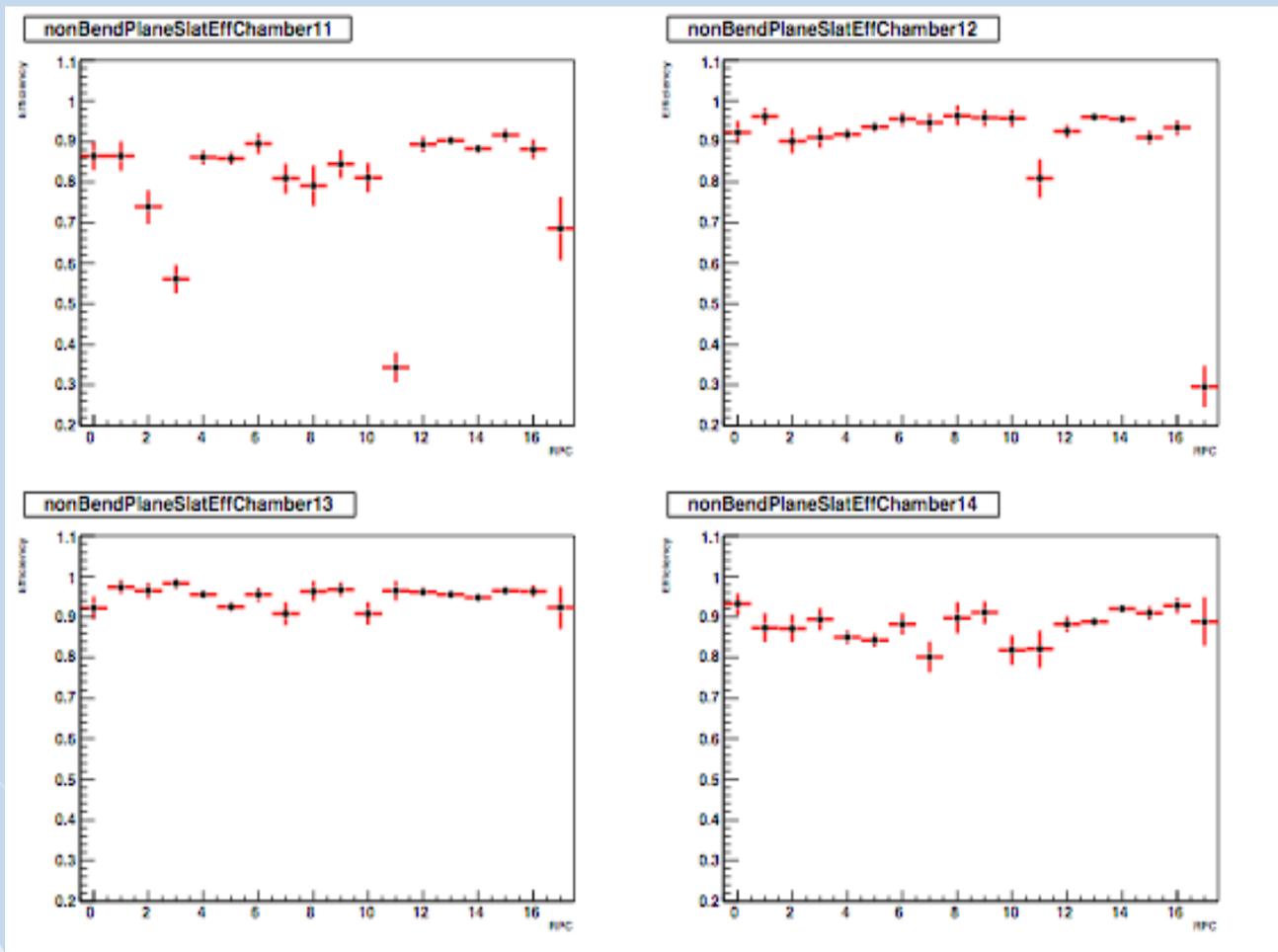




“Efficiency” in the non bending plane (per slat) for all the 4 stations (streamer mixture)



Check of the status of the system

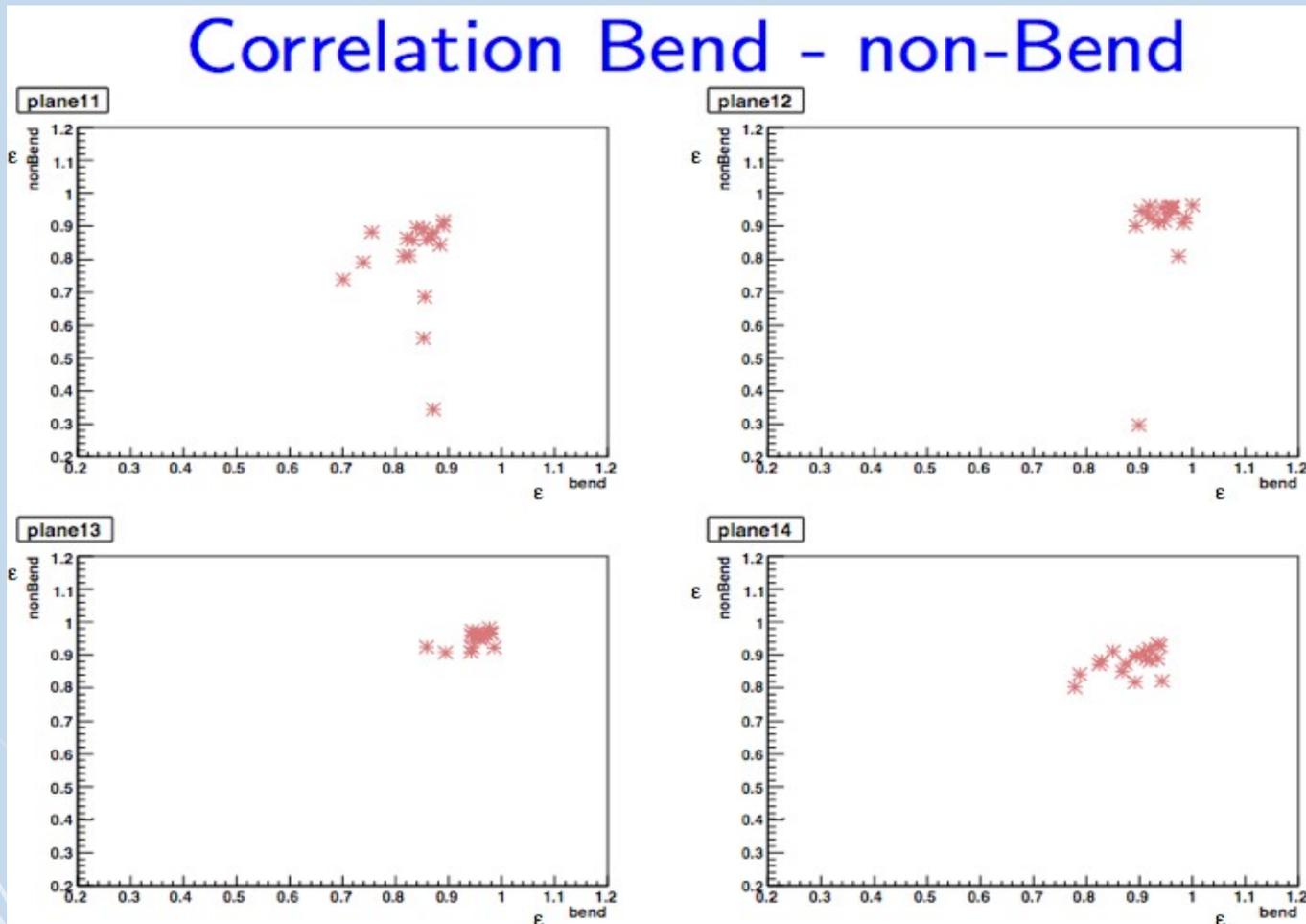




Correlation between bending and non-bending efficiency (streamer mixture)



Hardware issues can be more clearly understood considering this graphic





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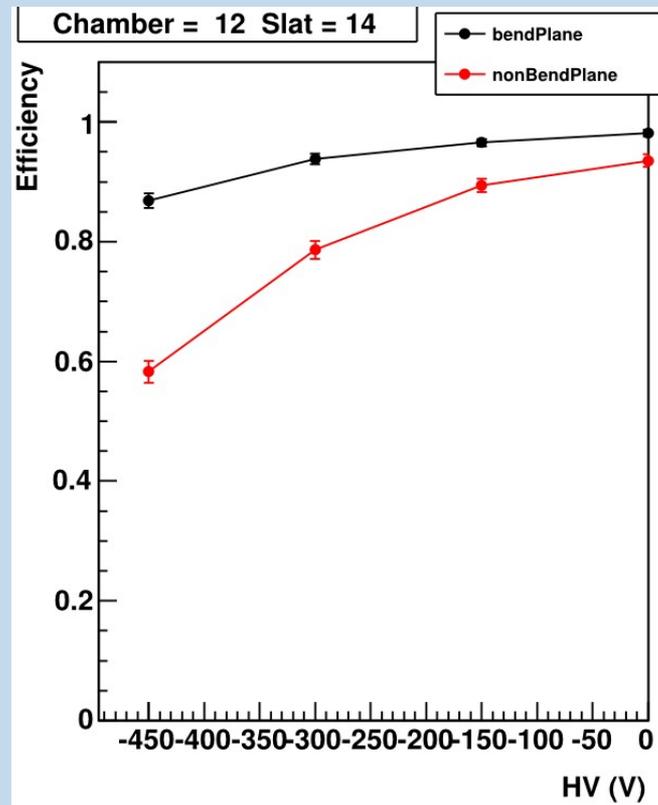
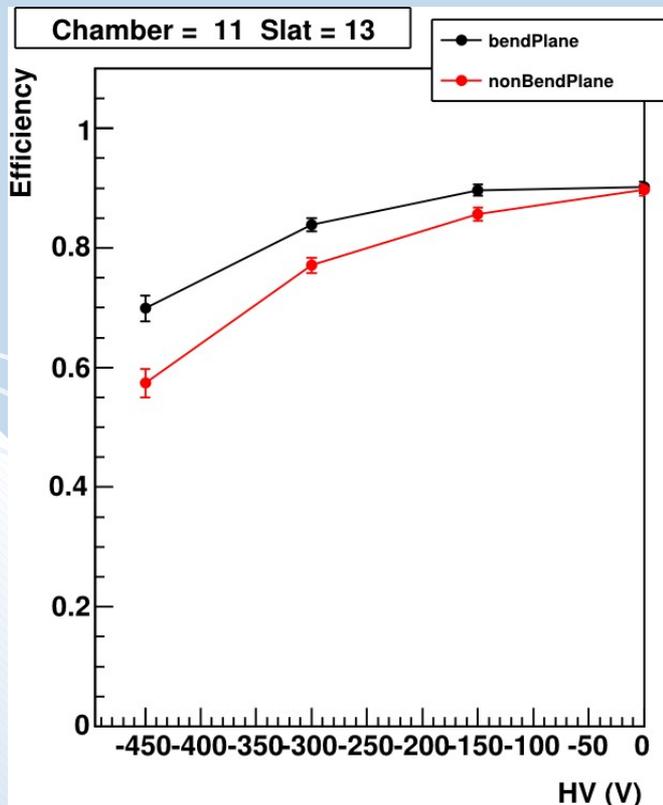


RPC status: efficiency vs. H.V.

RPC working point in **avalanche**:

Cosmic run (Ago-Sep 2009):

The goal is to find the working voltage by H.V. scan from the “nominal” H.V. value estimated with preliminary test .

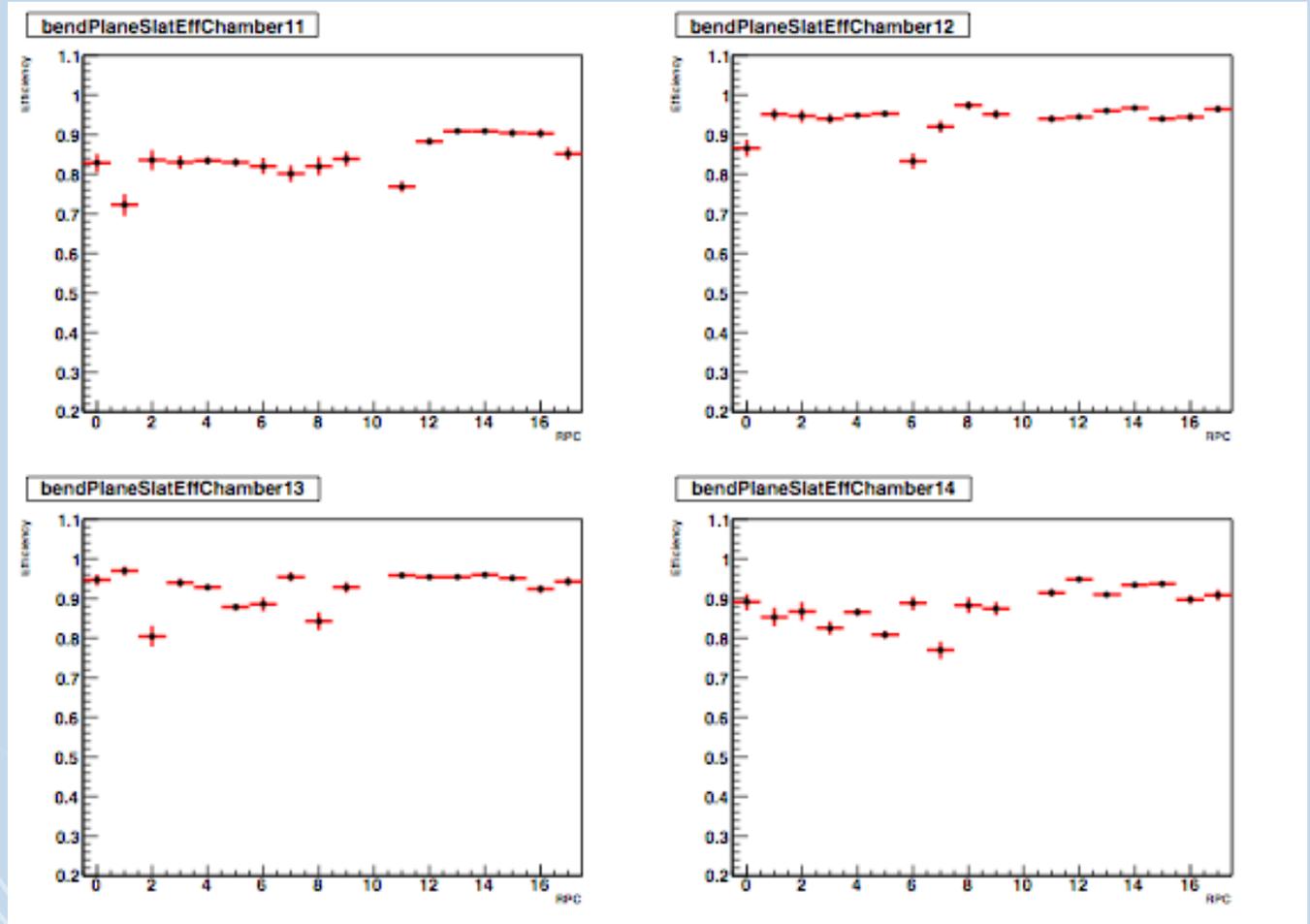




“Efficiency” in the bending plane (per slat) for all the 4 stations (avalanche mixture)



Check of the status of the system

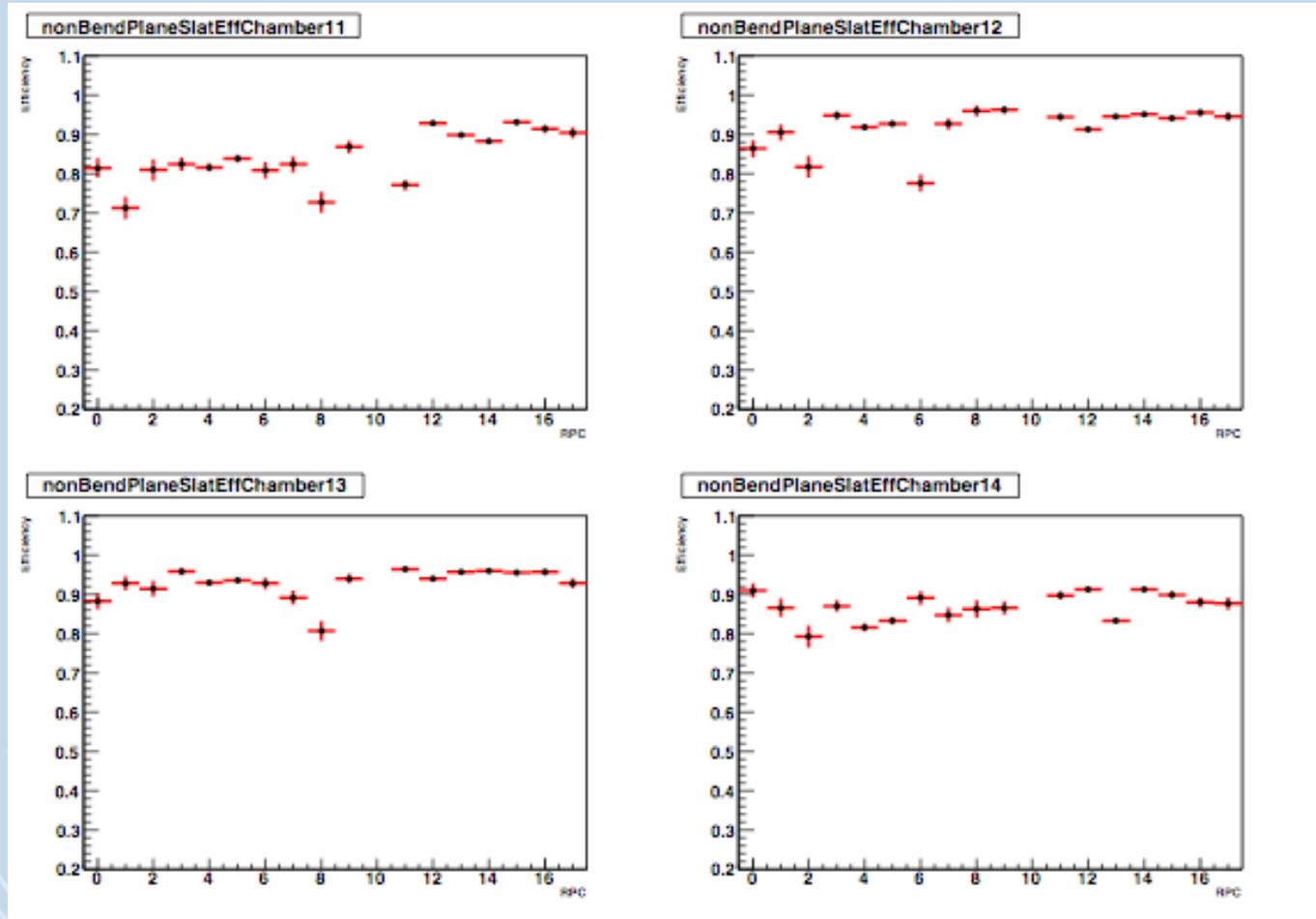




“Efficiency” in the non bending plane (per slat) for all the 4 stations (avalanche mixture)



Check of the status of the system





Conclusions

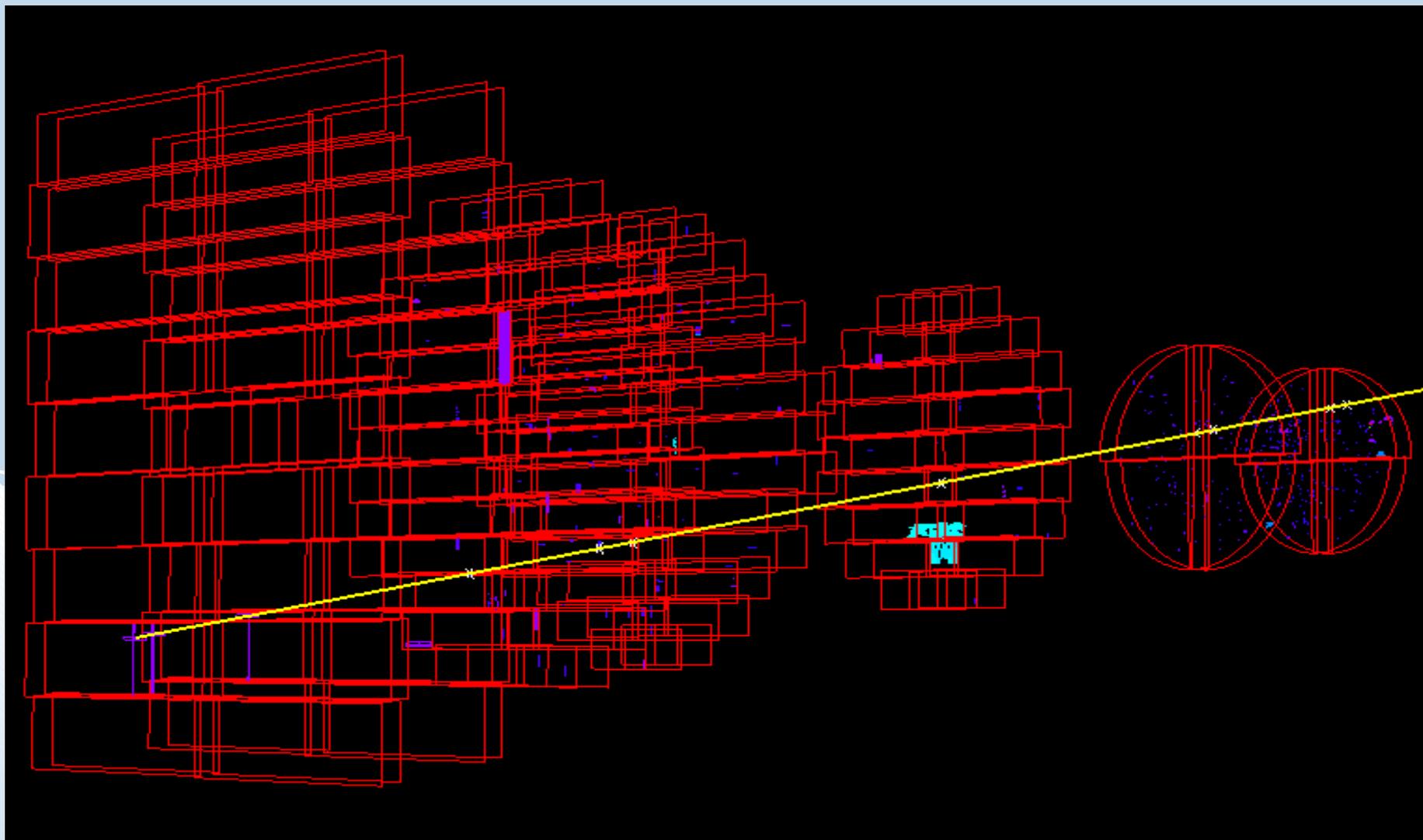
First and successful long-term test of the full detector in **avalanche mode**

MTR and **MTK** were stable and permanently operational, all along the cosmic run

The run has allowed to *test the muon spectrometer in a configuration very close to the final one* (almost all the detection and read-out elements active)

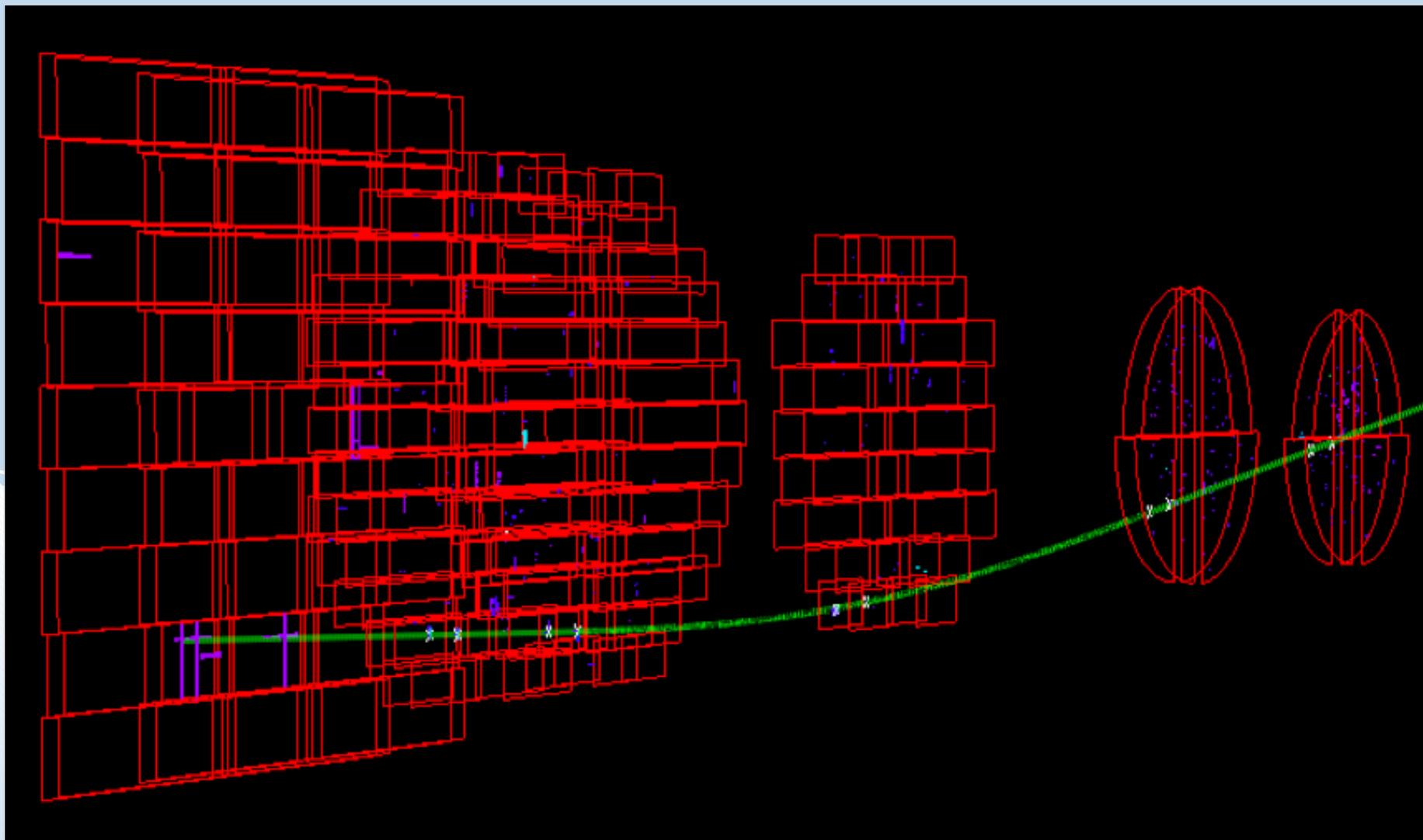


A muon track in the ALICE Muon Spectrometer with **dipole magnet OFF**



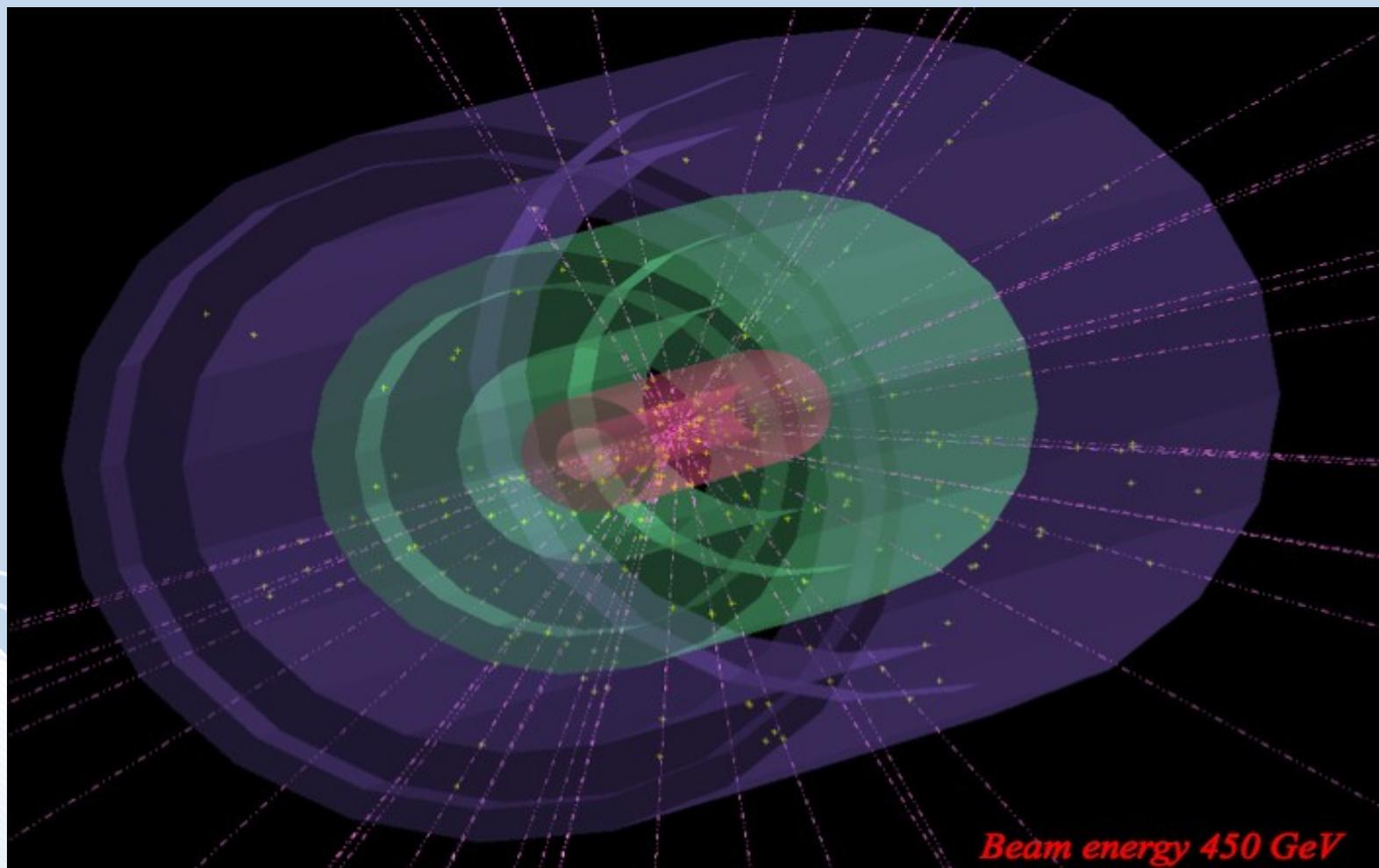


A muon track in the ALICE Muon Spectrometer with **dipole magnet ON**

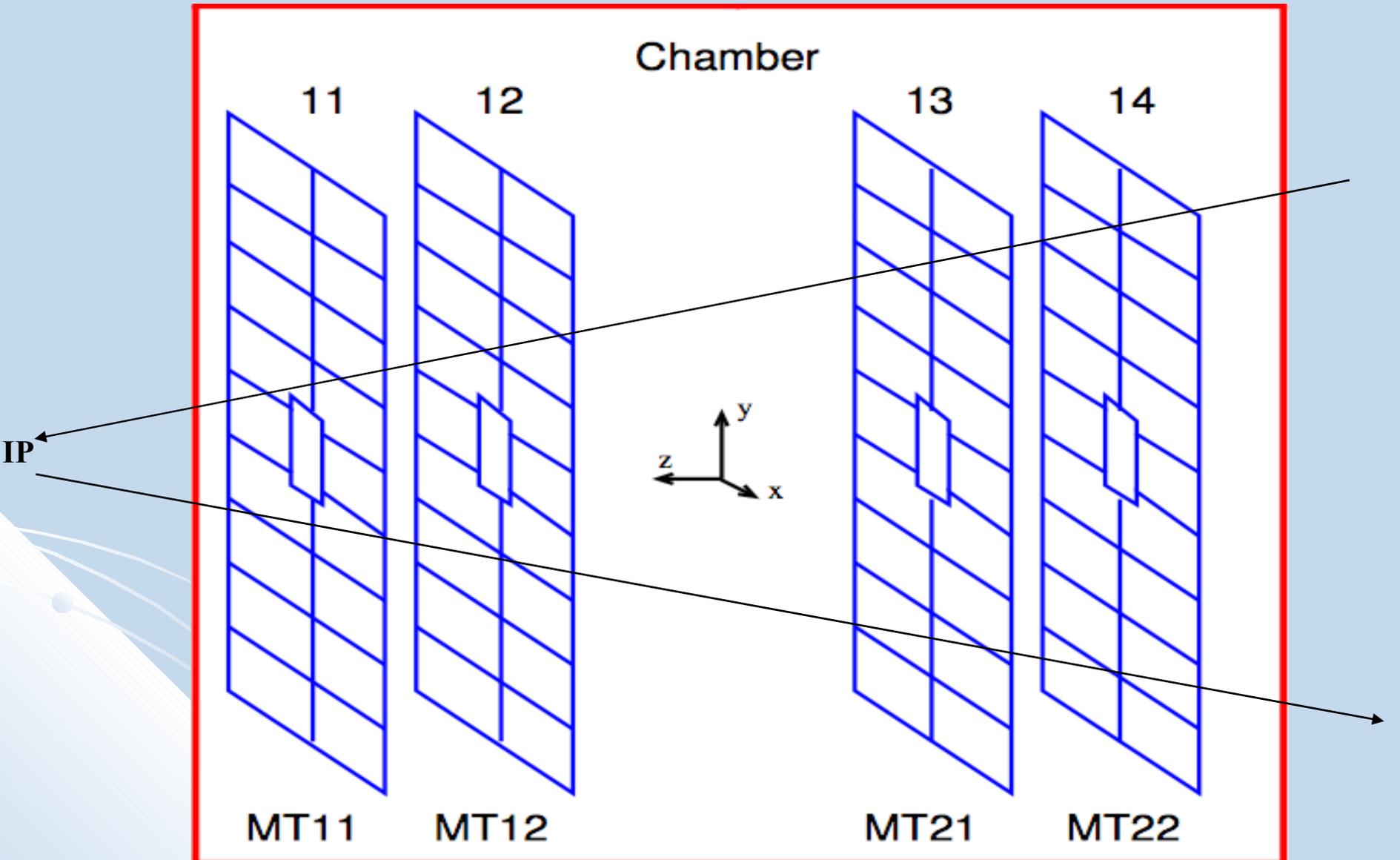




LHC restart: ALICE first event !!!



A handful of tracks have been reconstructed pointing back to a unique vertex



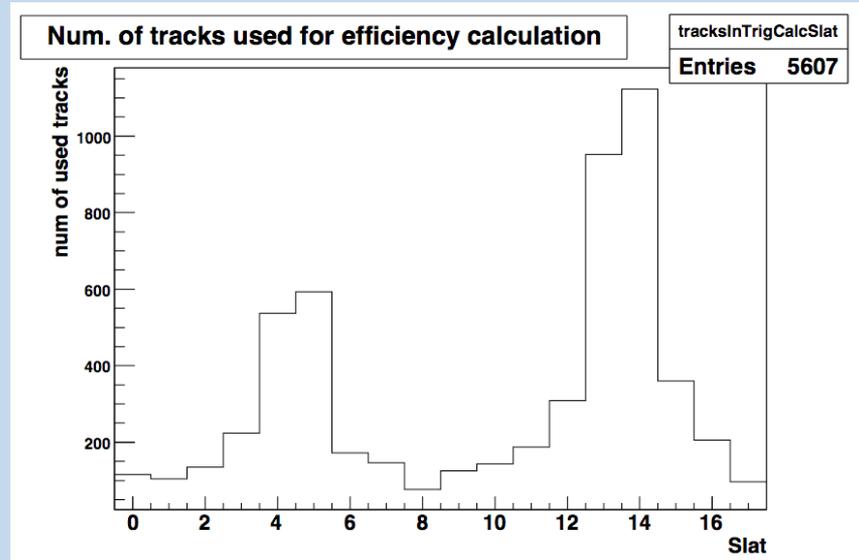


Commissioning with cosmic rays of the ALICE Muon Trigger System

05 9.Out	234	225	209	193	177	155	133	16	38	60	76	92	108	117	04 9.In
	LC7L9B1	LC8L9B1	LC9L9B1	LC4L9B1	LC3L9B1	LC2L9B1	LC1L9B1	RC1L9B1	RC2L9B1	RC3L9B1	RC4L9B1	RC5L9B1	RC6L9B1	RC7L9B1	
06 8.Out	233	224	208	192	176	154	132	15	37	59	75	91	107	116	03 8.In
	LC7L8B1	LC8L8B2	LC9L8B2	LC4L8B2	LC3L8B2	LC2L8B2	LC1L8B2	RC1L8B2	RC2L8B2	RC3L8B2	RC4L8B2	RC5L8B2	RC6L8B2	RC7L8B1	
07 7.Out	232	222	206	190	174	152	130	13	35	57	73	89	105	115	02 7.In
	LC7L7B1	LC8L7B2	LC9L7B2	LC4L7B2	LC3L7B2	LC2L7B2	LC1L7B1	RC1L7B1	RC2L7B2	RC3L7B2	RC4L7B2	RC5L7B2	RC6L7B2	RC7L7B1	
08 6.Out	231	220	204	188	172	150	128	11	33	55	71	87	103	114	01 6.In
	LC7L6B1	LC8L6B2	LC9L6B2	LC4L6B2	LC3L6B4	LC2L6B4	LC1L6B3	RC1L6B3	RC2L6B4	RC3L6B4	RC4L6B2	RC5L6B2	RC6L6B2	RC7L6B1	
09 5.Out	230	219	203	187	171	149	127	10	32	54	70	86	102	113	00 5.In
	LC7L5B1	LC8L5B2	LC9L5B2	LC4L5B2	LC3L5B3	LC2L5B3	LC1L5B2	RC1L5B2	RC2L5B3	RC3L5B3	RC4L5B2	RC5L5B2	RC6L5B2	RC7L5B1	
10 4.Out	229	218	202	186	170	148	126	9	31	53	69	85	101	112	17 4.In
	LC7L4B1	LC8L4B2	LC9L4B2	LC4L4B2	LC3L4B4	LC2L4B4	LC1L4B3	RC1L4B3	RC2L4B4	RC3L4B4	RC4L4B2	RC5L4B2	RC6L4B2	RC7L4B1	
11 3.Out	228	217	201	185	166	144	124	8	29	51	67	83	99	111	16 3.In
	LC7L3B1	LC8L3B2	LC9L3B2	LC4L3B2	LC3L3B3	LC2L3B3	LC1L3B2	RC1L3B2	RC2L3B2	RC3L3B2	RC4L3B2	RC5L3B2	RC6L3B2	RC7L3B1	
12 2.Out	227	216	198	182	160	138	122	5	21	43	65	81	97	110	15 2.In
	LC7L2B1	LC8L2B2	LC9L2B2	LC4L2B2	LC3L2B2	LC2L2B2	LC1L2B2	RC1L2B2	RC2L2B2	RC3L2B2	RC4L2B2	RC5L2B2	RC6L2B2	RC7L2B1	
13 1.Out	226	215	199	183	162	140	124	7	23	45	66	82	98	109	14 1.In
	LC7L1B1	LC8L1B2	LC9L1B2	LC4L1B2	LC3L1B2	LC2L1B2	LC1L1B2	RC1L1B2	RC2L1B2	RC3L1B2	RC4L1B2	RC5L1B2	RC6L1B2	RC7L1B1	

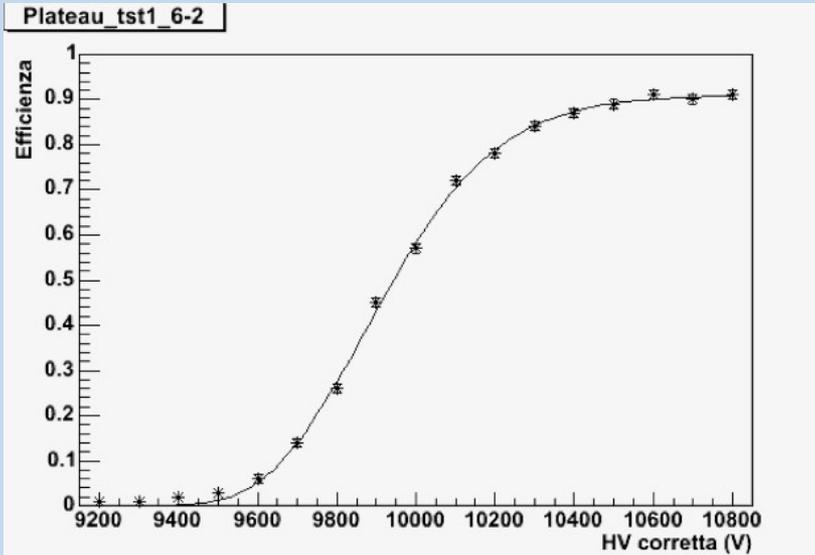
□ = RPC

□ = Board

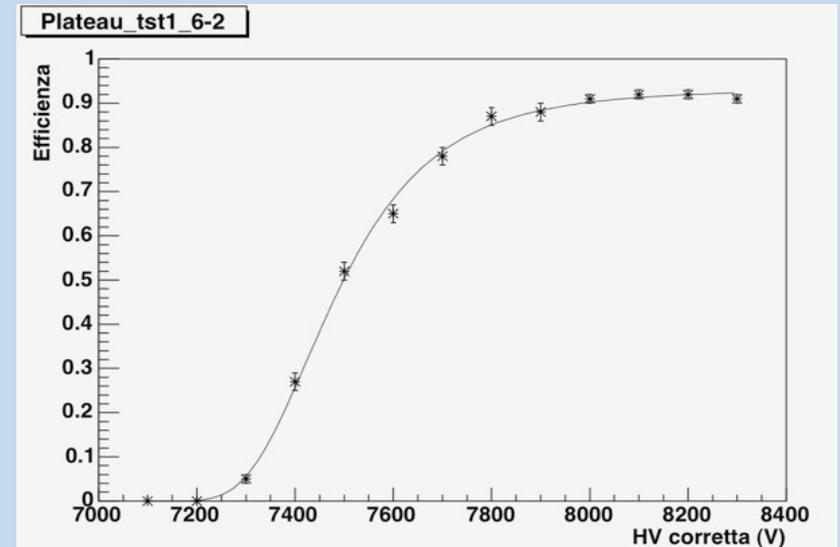
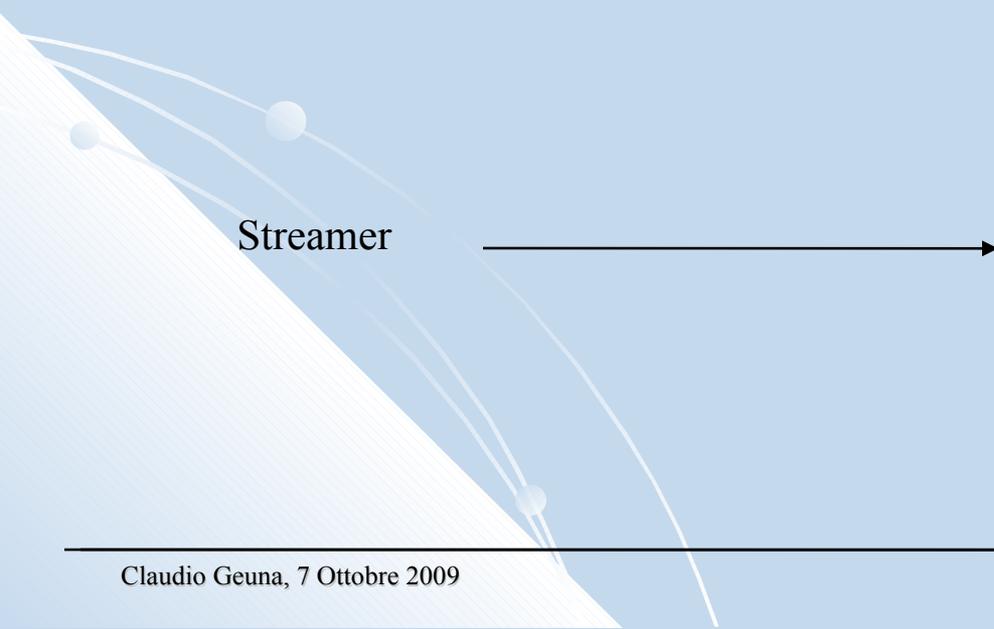




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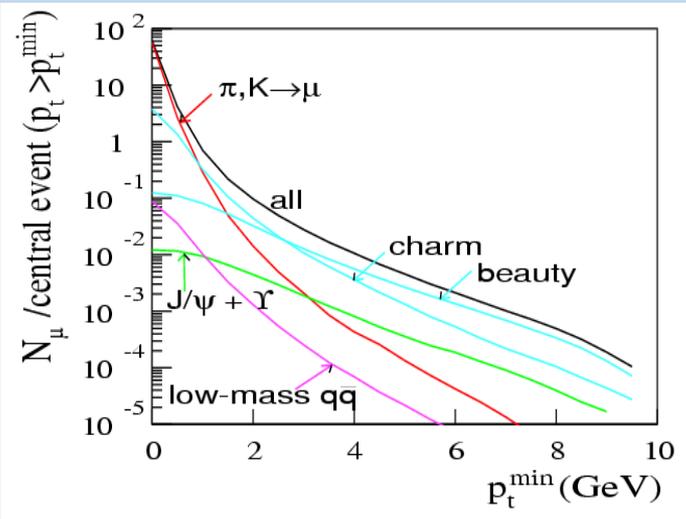


Avalanche





The principle of the trigger (I)



GOAL:

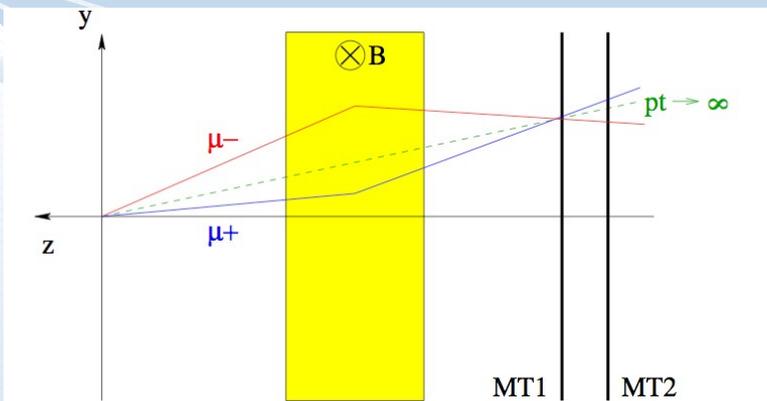
- Selection of (muon) tracks pointing to I.P. with p_t above 2 thresholds:

$$\text{low } p_{t \text{ cut}} = 1 \text{ GeV/c}$$

$$\text{high } p_{t \text{ cut}} = 2 \text{ GeV/c}$$

- Trigger signals for single μ , like-sign and unlike-sign μ pairs

8 strips in the vertical direction and 1 in the horizontal direction.



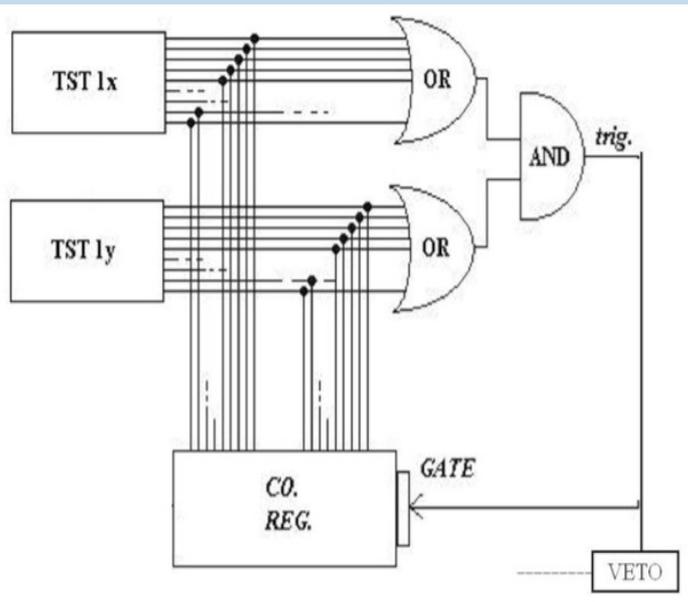
PRINCIPLE:

cut on $p_t \Leftrightarrow$ cut on deviation between MT1 and MT2
 \rightarrow select tracks in a road of a given width

Noise measurements - The Autotrigger method

The noise of the detectors is quantified by the dark counting rate, i.e. the counting rate of the detectors with no beam or irradiation, when the hits are only due to cosmic rays and intrinsic noise.

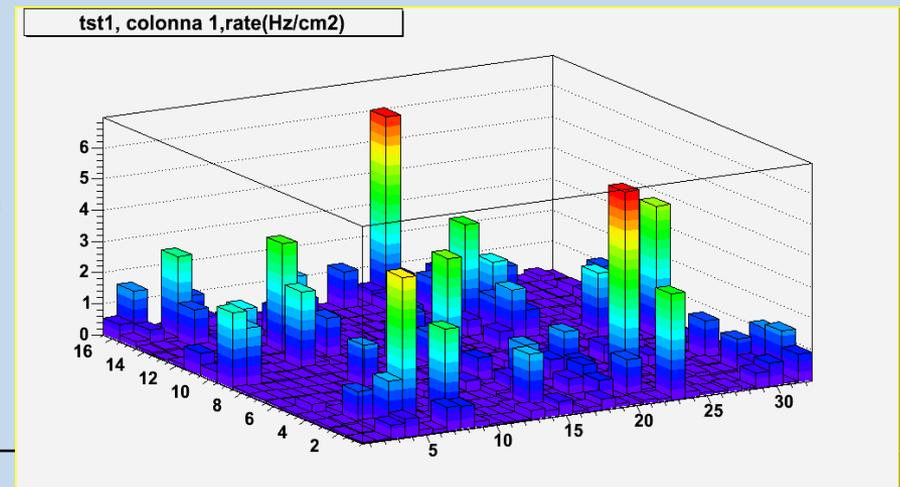
The counting rate is measured locally with the *autotrigger method* : the trigger is given by the detector itself, selecting events with at least one hit on both strip planes. The logical scheme of the electronic chain for the autotrigger measurements is shown below.



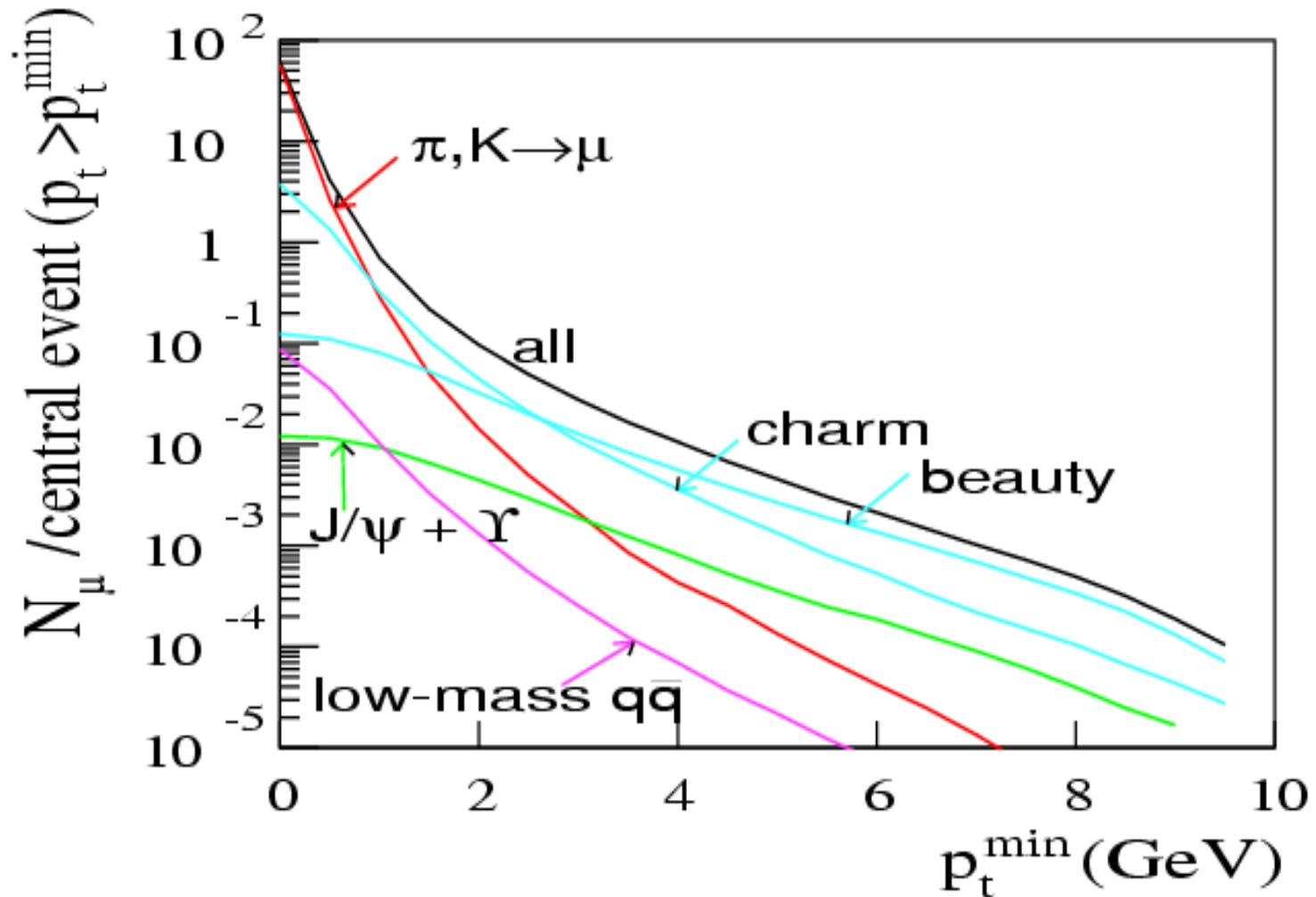
Logical scheme of the electronic chain for the autotrigger measurement.

The detector surface can be divided in **cells** defined by the crossing of strips in the two direction. Such a method provides the noise map of the detectors, which makes the detection of noisy spots possible.

Here there is the noise map obtained with gap 1210 (left) at the voltage of 8300 V (streamer mixture)



Taglio in impulso trasverso



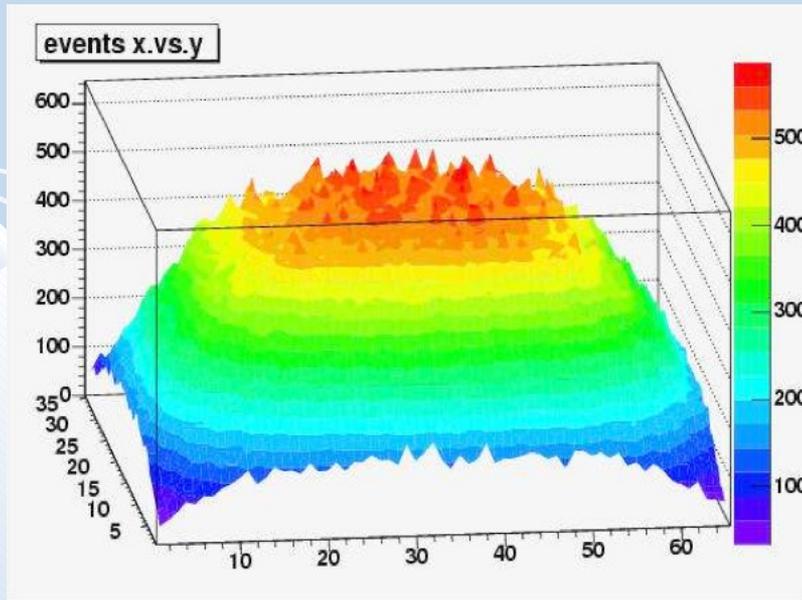
Efficiency maps (I)

To evaluate better the uniformity of the detectors , and to detect any imperfection, though small, **efficiency maps** are measure at two voltage value.

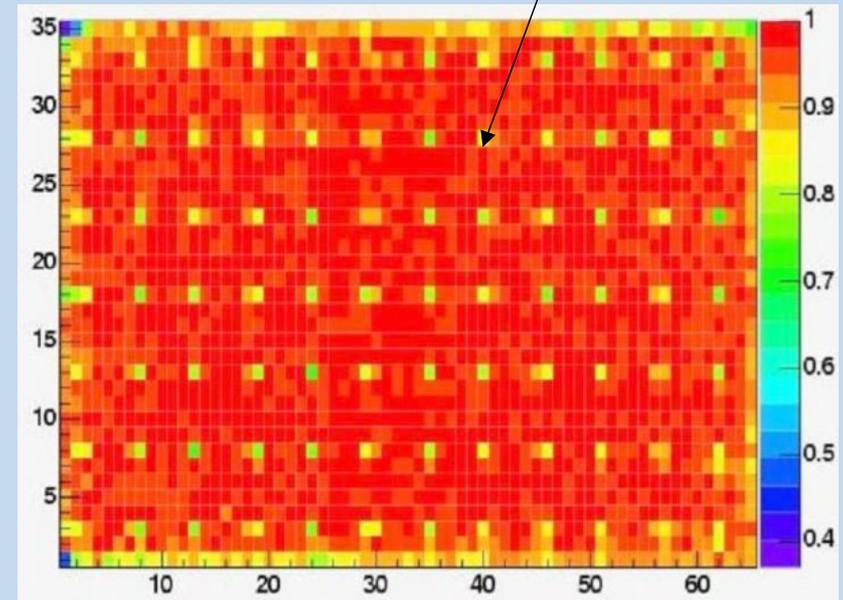
In *Streamer mode* we chose **8200 V** and **8100 V**.

In *Avalanche mode* we chose only **10800 V**.

The cells for efficiency maps are about $2 \times 2 \text{ cm}^2$ large. With a 1000000 events run ($\sim 10 \text{ h}$ acquisition time), the statistics is of about 500 events in central cells, 100 in peripheral cells , 50 in the very side cells. The resolution is of the order of the centimeter, so that, in the efficiency map, even the spacers that keep the distance between the electrodes constant (whose diameter is 1 cm) can be resolved.



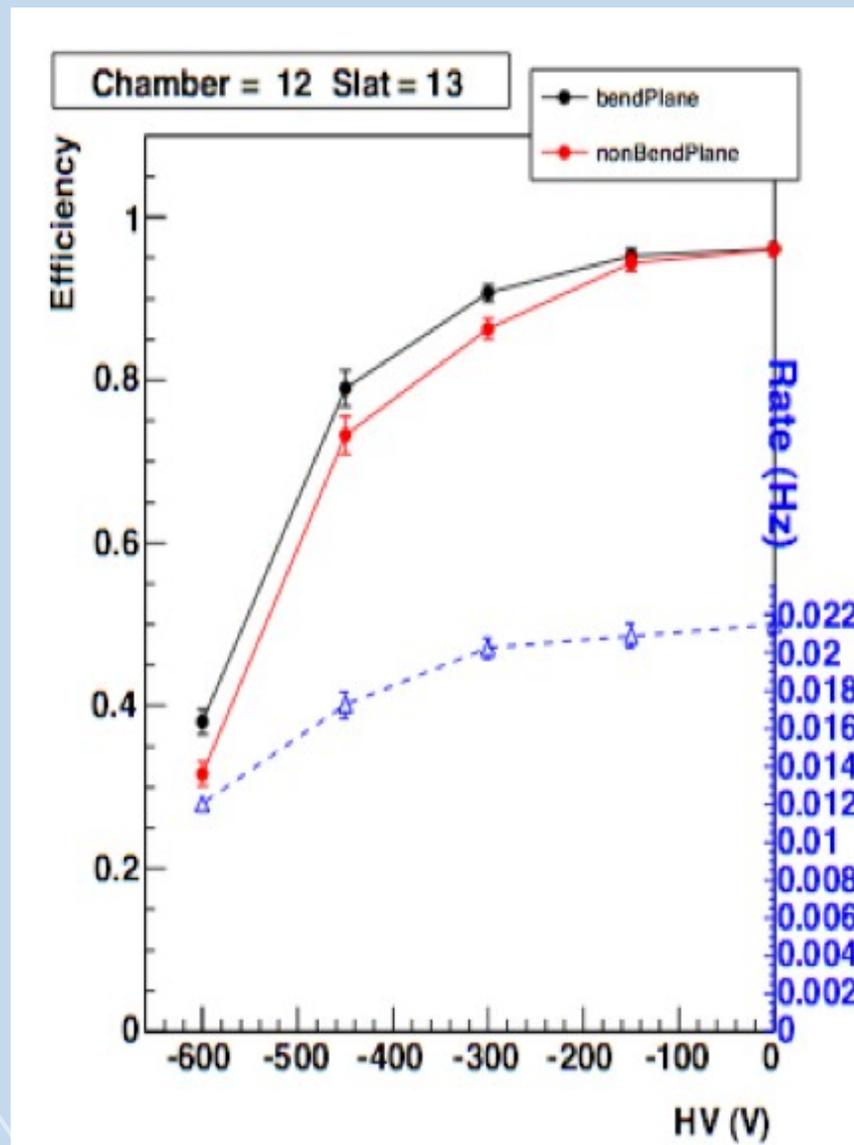
Distribution of triggered events for efficiency measurement over the surface of a half chamber. Units are given in cells. The area of the cells is $2 \times 2 \text{ cm}^2$



Efficiency map of a half chamber operated at 8200 V . Units are given in cells. The area of the cells is $2 \times 2 \text{ cm}^2$

Temperature - Pressure HV correction

$$V_{eff} = V \frac{T}{T_0} \frac{p_0}{p}$$



$$N_{3/3}^{12} =$$

