

Comprendre l'infiniment grand

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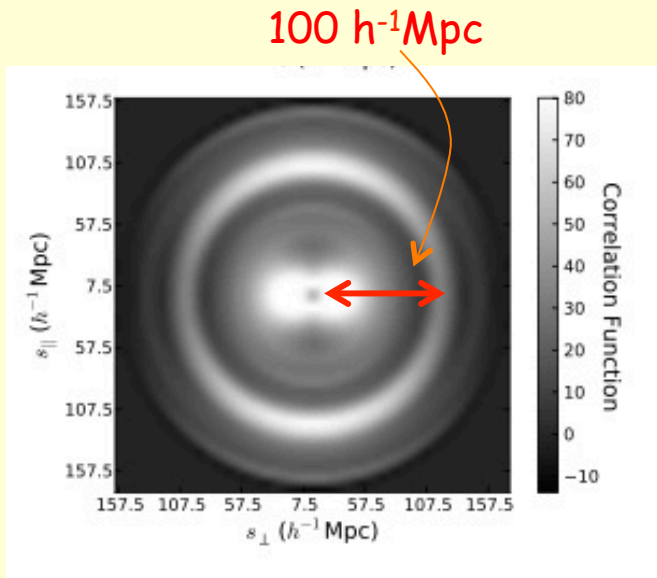
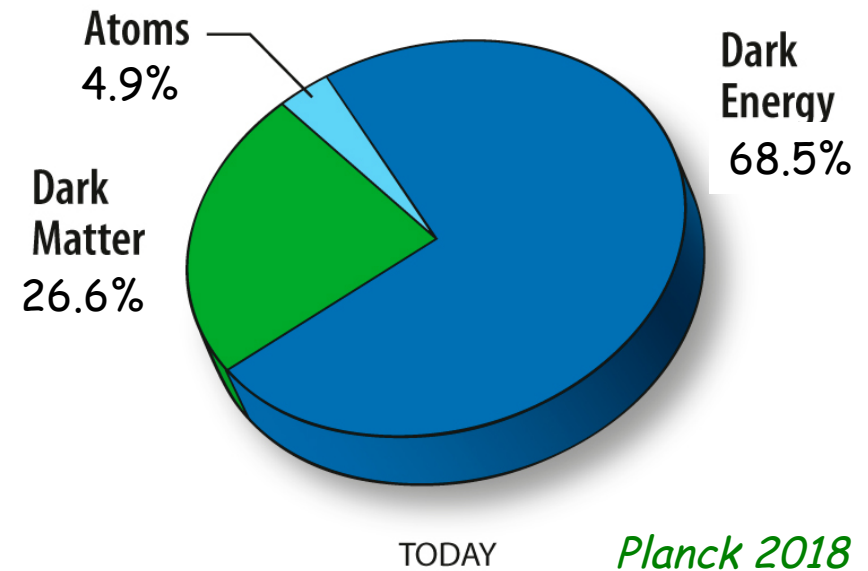
- 1) Du modèle du Big Bang au contenu de l'Univers
- 2) Mesures cosmologiques
- 3) **Les grandes structures de l'Univers**

Large scale structures

- Summary :

General Relativity + Λ and CMB,
SNeIa, BAO data

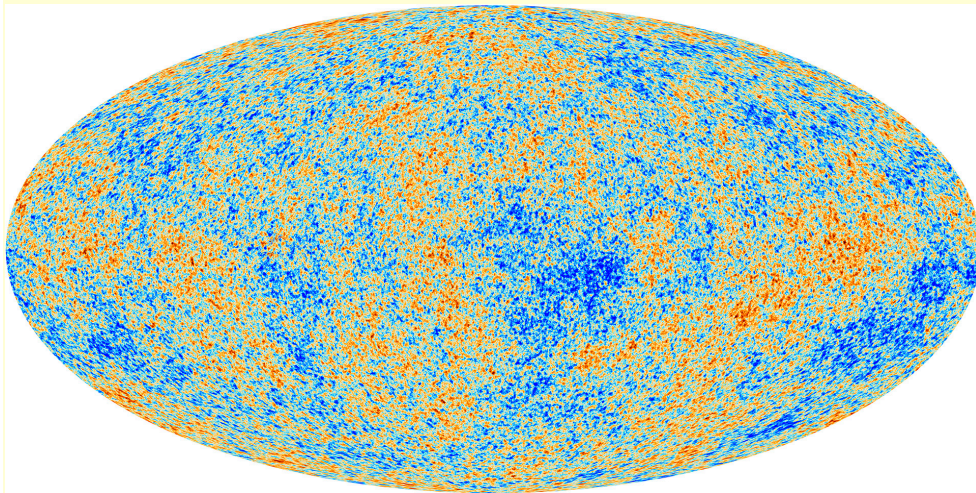
=> Concordance model = Λ_{CDM}



- Beyond BAO: LSS

- Structure formation
- Testing gravity with LSS
- Future LSS surveys : e.g. DESI

1. Structure formation



Planck (2013), all foregrounds subtracted

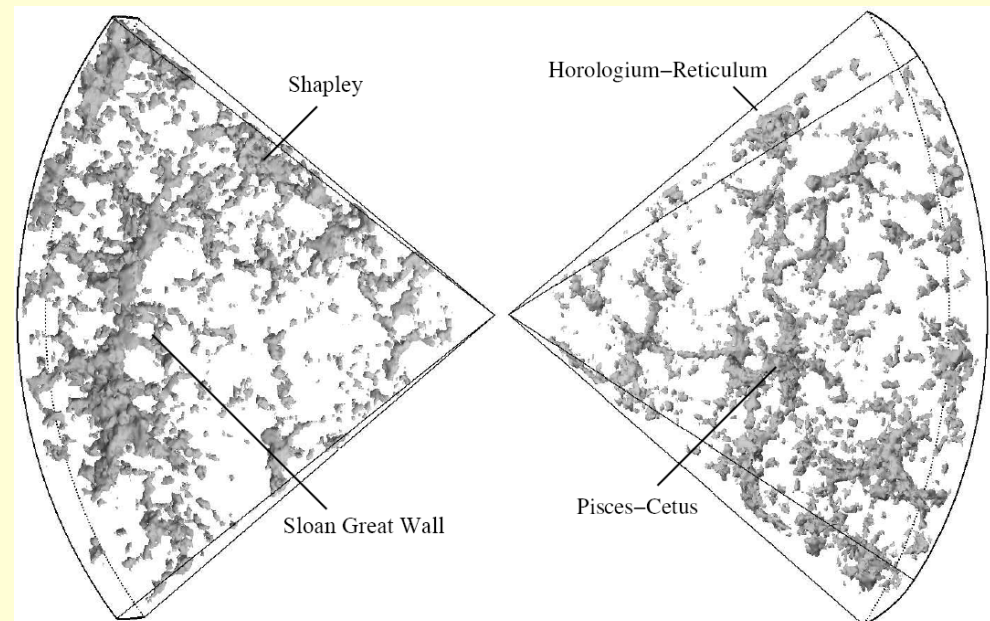
Observational facts

- CMB: very small anisotropies

$$\Delta T/T = 10^{-5}$$

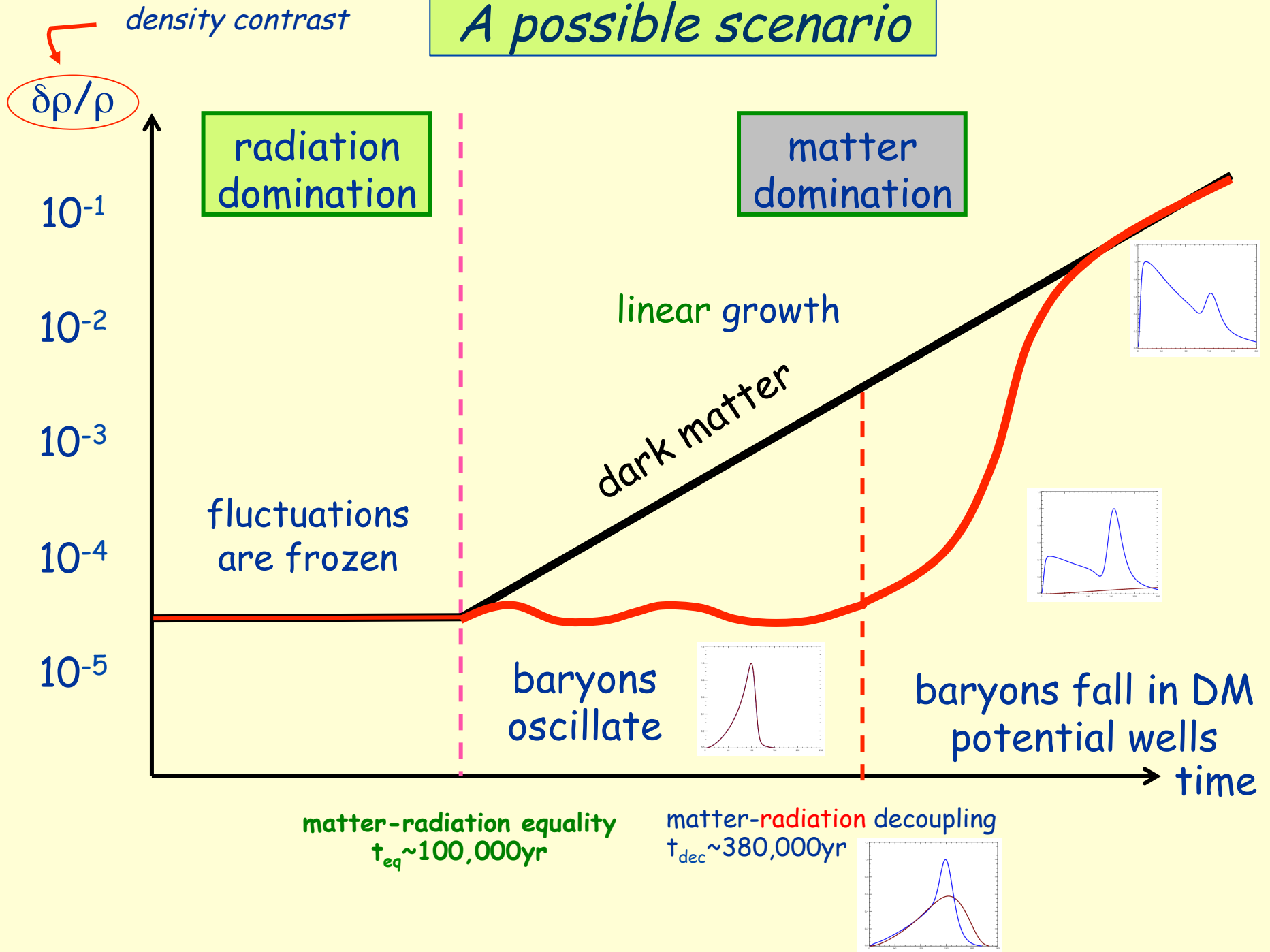
- matter density inhomogeneities,
amplified by gravitation after
decoupling → structures

- Large scale structures today: galaxies, clusters, superclusters, voids and filaments
- Formation of structures: cold dark matter mandatory



2dF Galaxy Redshift Survey, (2007) 3

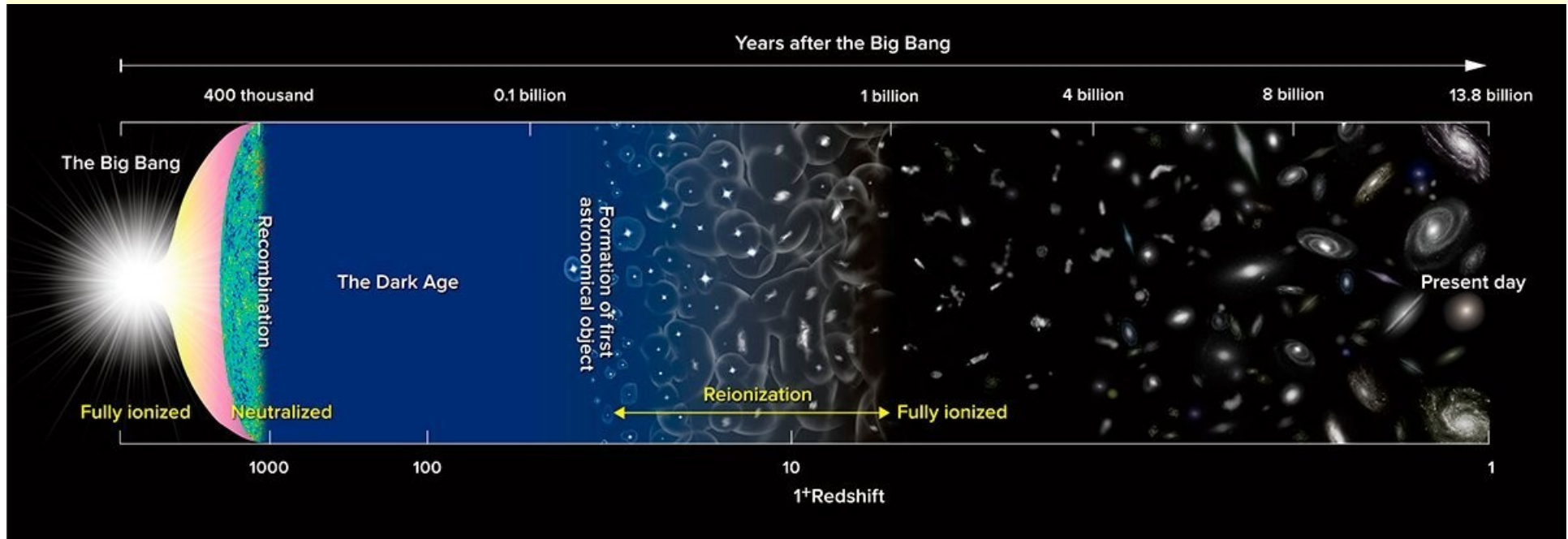
A possible scenario



Cosmic timeline

Dark ages

Galaxy evolution



first stars &
galaxies form

~ 300-500million years

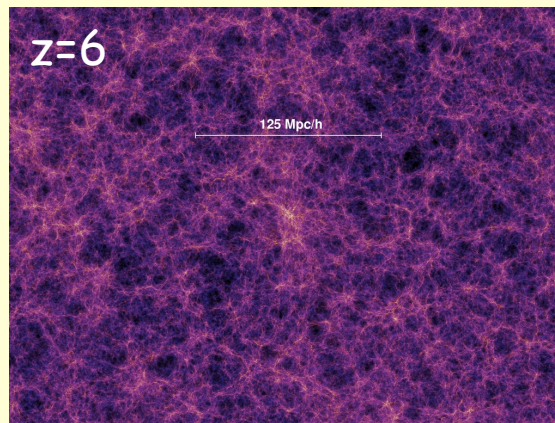
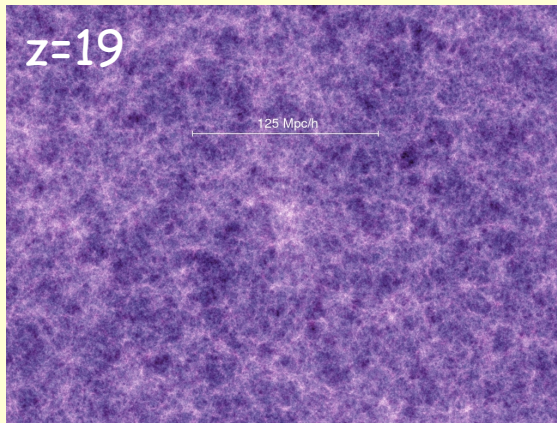
clusters &
superclusters

- Hierarchical formation structure, smaller structures (stars, dwarf galaxies...) collapse earlier than larger ones (clusters)

Numerical simulations, examples

Needed to describe the **non-linear** ($\delta\rho/\rho > 1$) regime of structure growth (galaxies, clusters....).

Scale: 125 Mpc/h

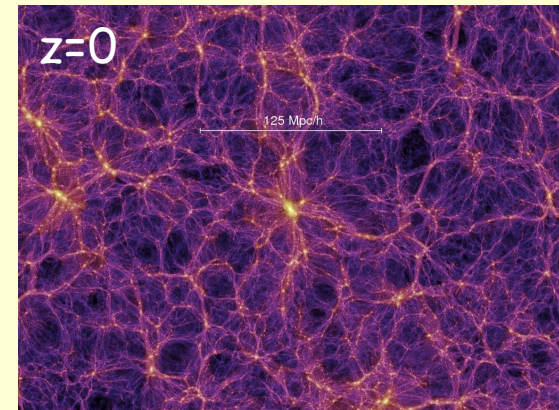
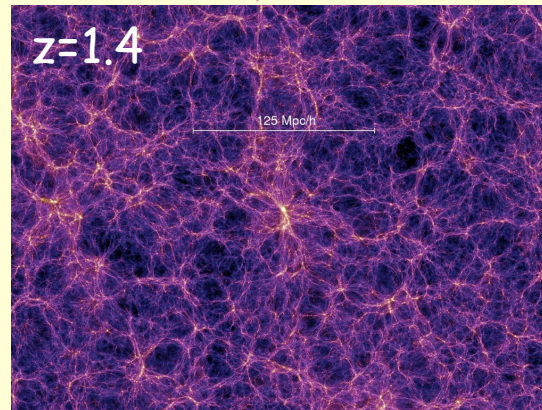


- DM particles
- gravitation
- cosmology
- hydrodynamics

0.2 Gyr

1 Gyr

*Millennium simulation,
Springel et al., 2005.*



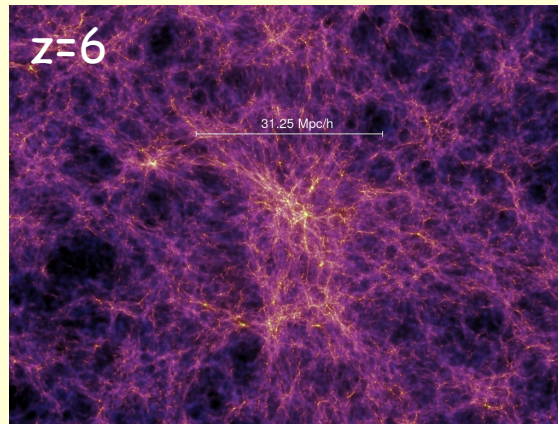
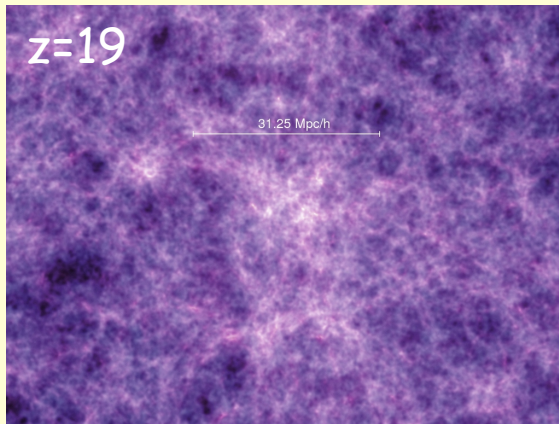
4.7 Gyr

13.6 Gyr

Numerical simulations, examples

Needed to describe the **non-linear** ($\delta\rho/\rho > 1$) regime of structure growth (galaxies, clusters....).

Scale: 31.5 Mpc/h

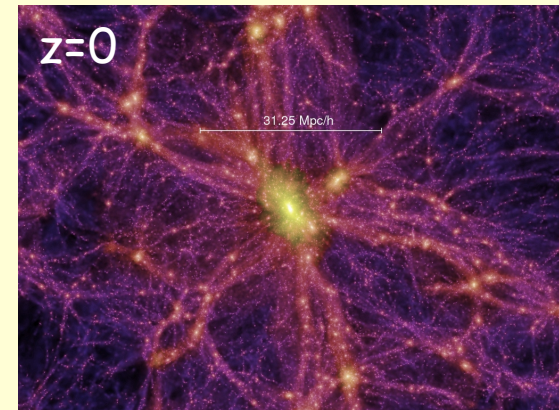
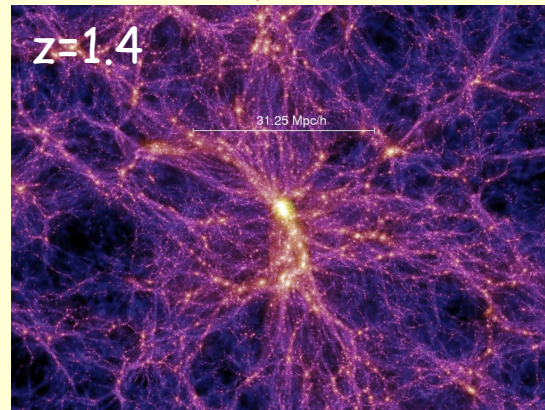


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Springel et al., 2005.*

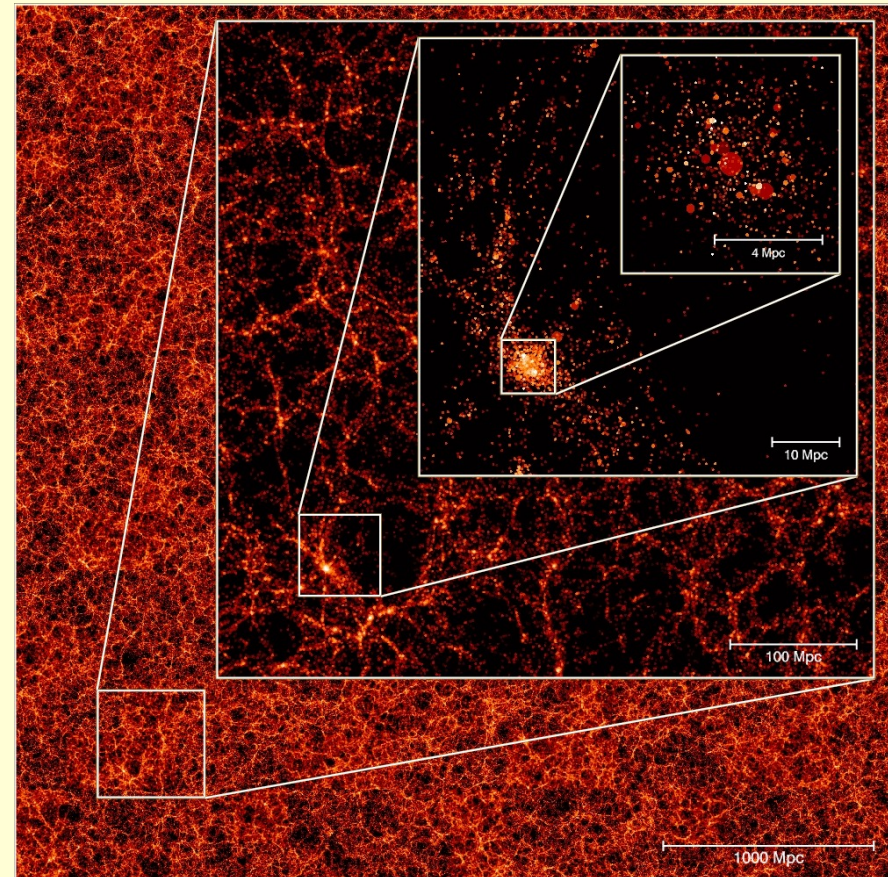
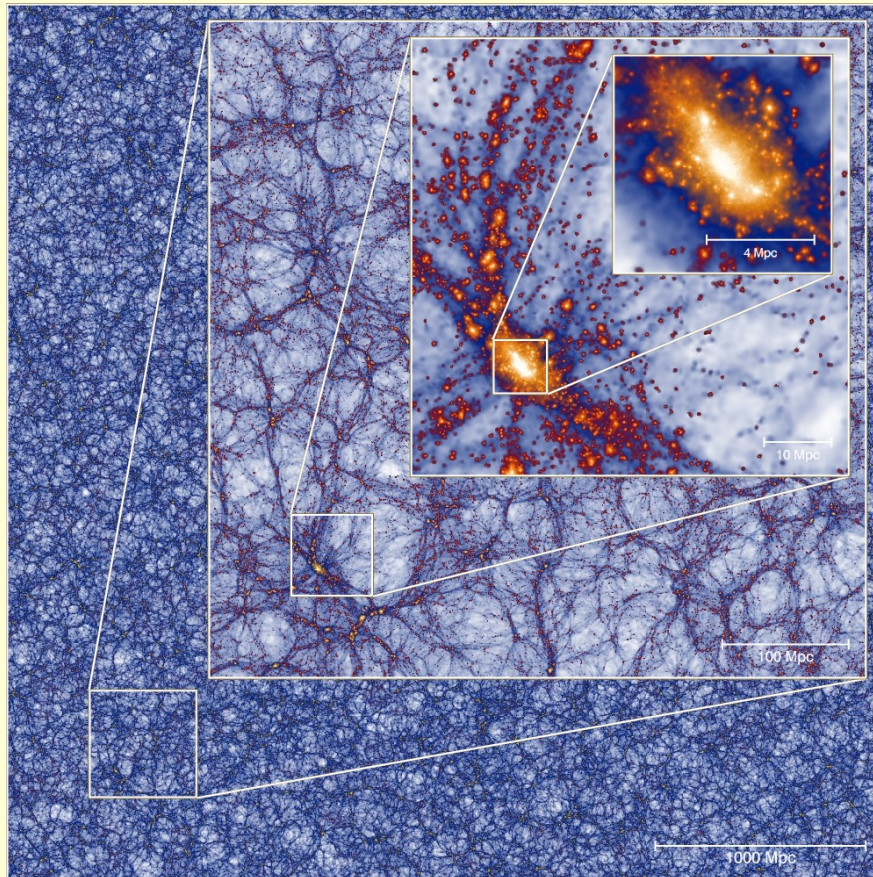


4.7 Gyr

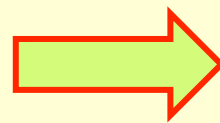
13.6 Gyr

Millenium-XXL simulation (2011)

6720³ DM particles, 4.1 Gpc box, gravitation followed over 13 Gyr



DM density field, $z=0$



galaxy distribution, $z=0$

Structure formation, summary

- Numerical simulations of large scale structure formation (down to galaxy dense halos) from cold dark matter :

voids and filaments (the cosmic web) well reproduced



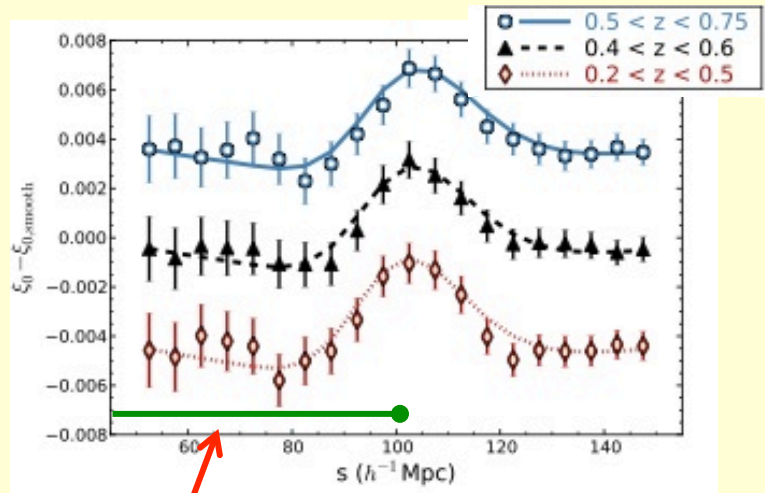
Evolution of matter distribution from numerical simulations

- Current status :

- Large scale success: two-point correlation function and high z luminosity function correctly reproduced
- Small scale issues: too many DM satellites expected, expected DM halo profile too cuspy in galaxy cores, galaxy rotation curves do not agree with data
- By-product : link with particle physics through constraints on neutrinos from large scale structure data ($N_\nu, \Sigma m_\nu$)

2. Testing gravity with LSS

From BAO ...

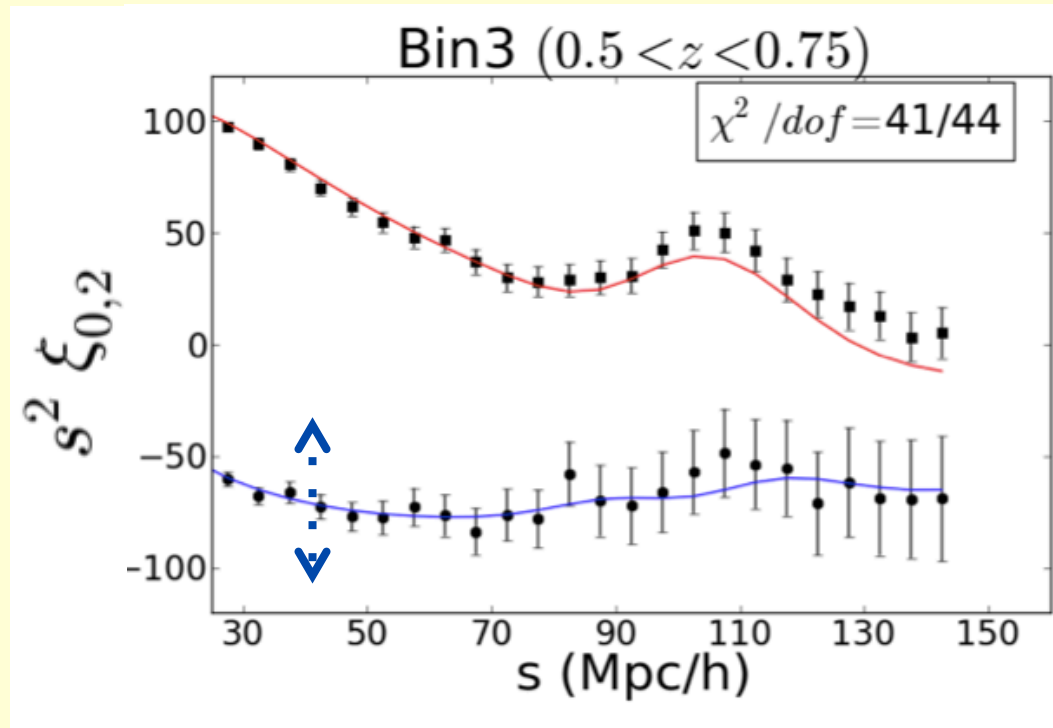


BAO scale

quadrupole amplitude =
gravity strength

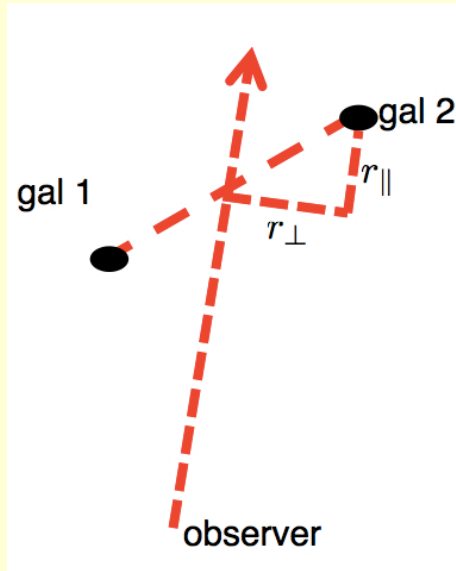
... to full shape analysis

S. Satpathy et al., 2017, MNRAS, 469, 13695



➡ requires understanding of matter clustering on
small scales (i.e. below BAO scale)

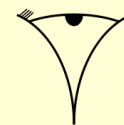
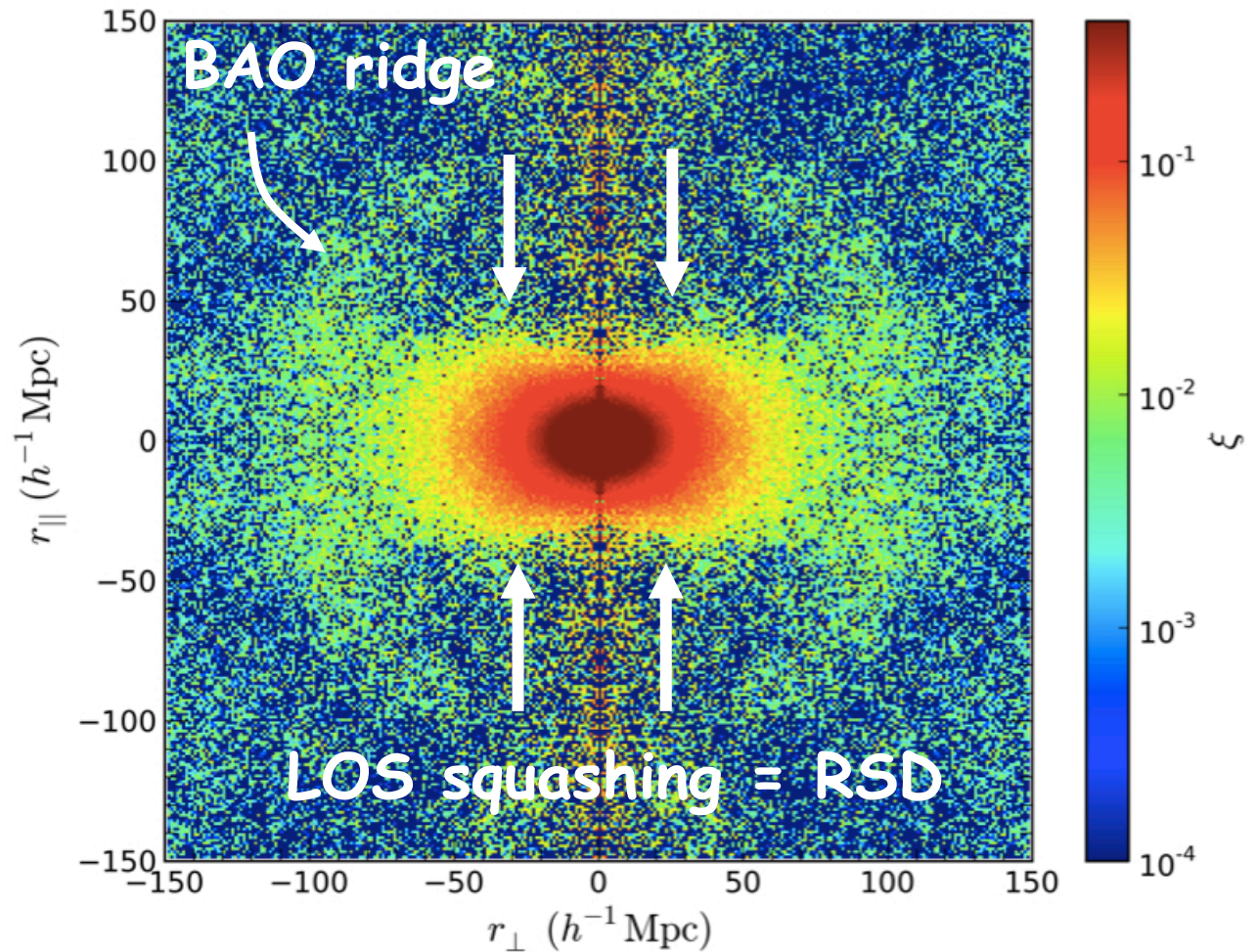
S. Alam et al., 2017, MNRAS, 470, 2617A



observed
redshift: Hubble
expansion +
peculiar velocity
due to **gravity**



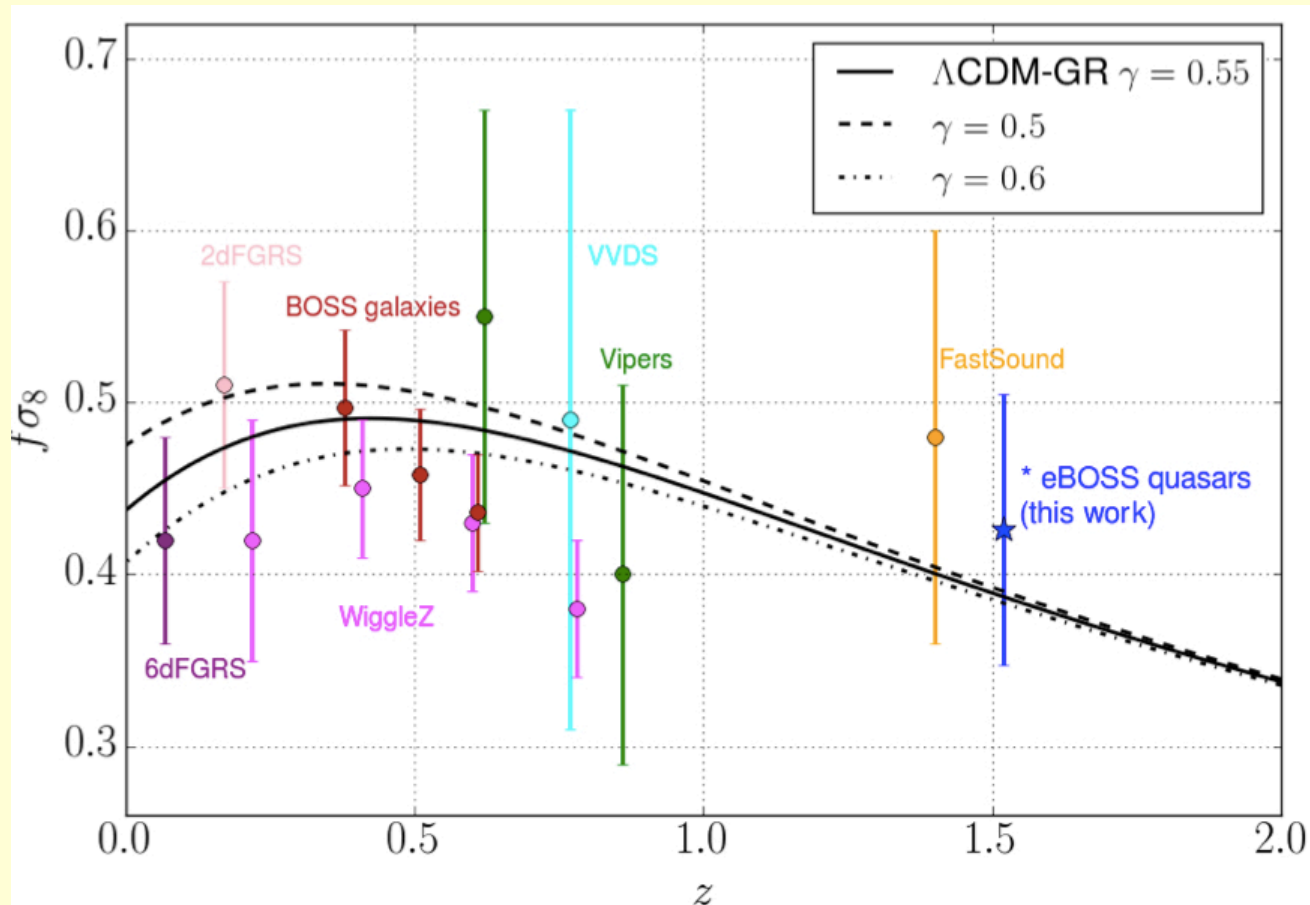
Redshift Space Distortions : a way to test gravity -> **full shape analysis** and **accurate modeling** of correlation function required (numerical simulations)



*L. Samushia et al, 2014,
MNRAS, 439, 3504.*

Results : growth rate of structure measurements

$$f(z) \equiv \frac{d \ln D}{d \ln a}, \quad D(a) \equiv \frac{\delta \rho(a)}{\delta \rho_0}, \quad \sigma_8 \equiv \text{power spectrum normalization today}$$



P. Zarrouk et al., 2018, MNRAS, 477, 1639

At present : **few** measurements, best accuracy \approx **6-10%**

3. Future LSS surveys

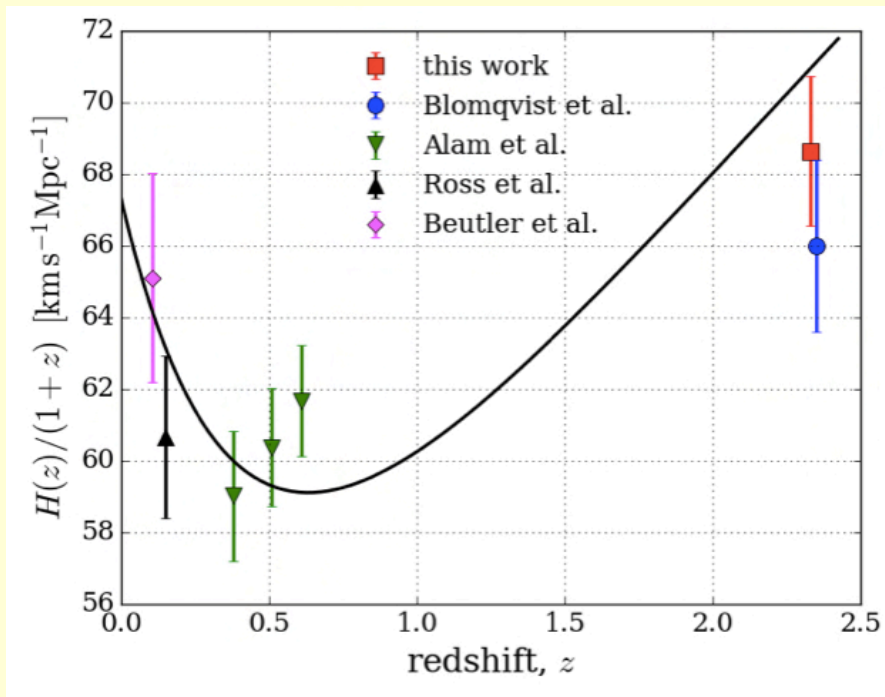
The Dark Energy Spectroscopic Instrument (2020)

- From SDSS to DESI: data $\times 15$
35 million galaxies, quasars / 14,000 sq.degrees / $0 < z < 4$

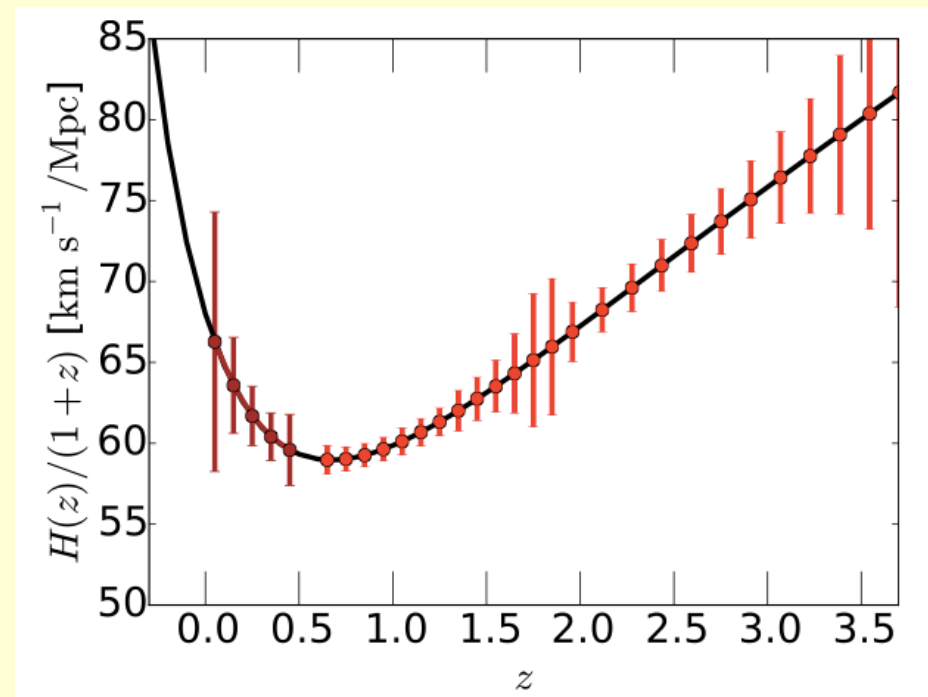
BAO along the line of sight $\Rightarrow H(z)$

2019

DESI



V.de Sainte Agathe et al., arXiv:1904.03400



DESI prospects, arXiv:1611.00036

3. Future LSS surveys

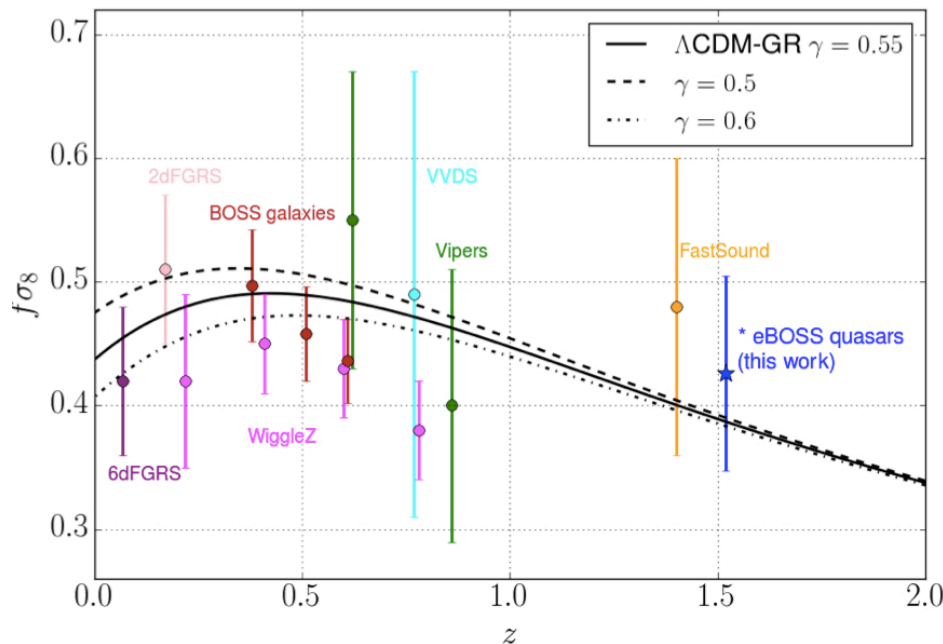
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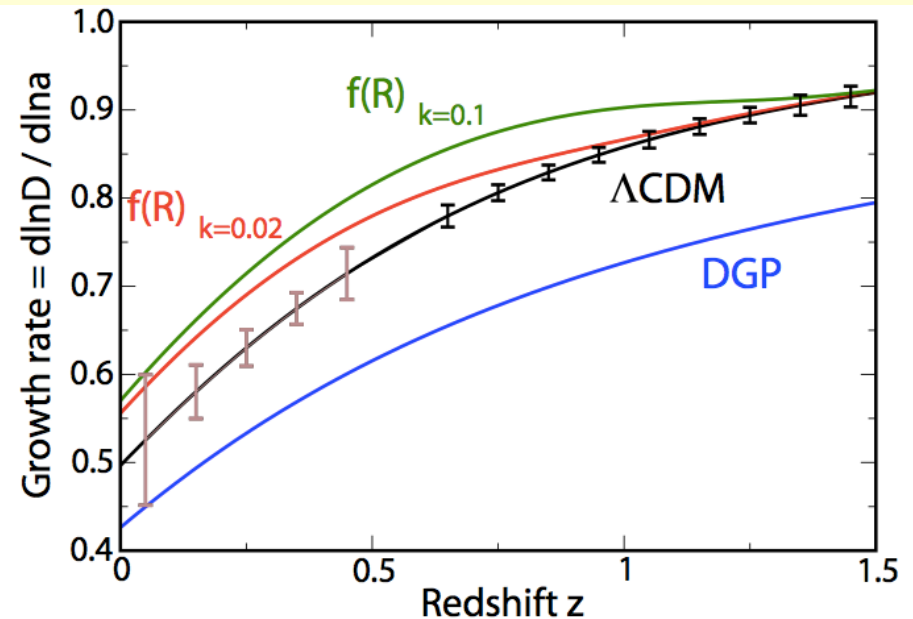
Linear growth of structure rate

2018

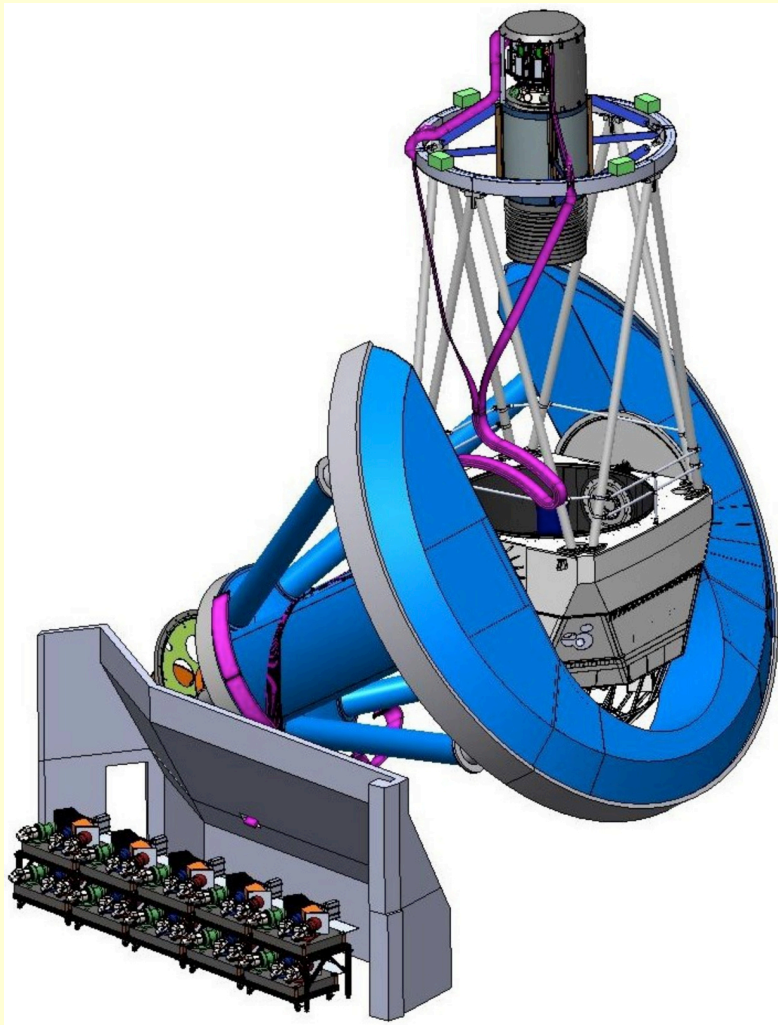
DESI



P. Zarrouk et al., 2018, MNRAS, 477, 1639

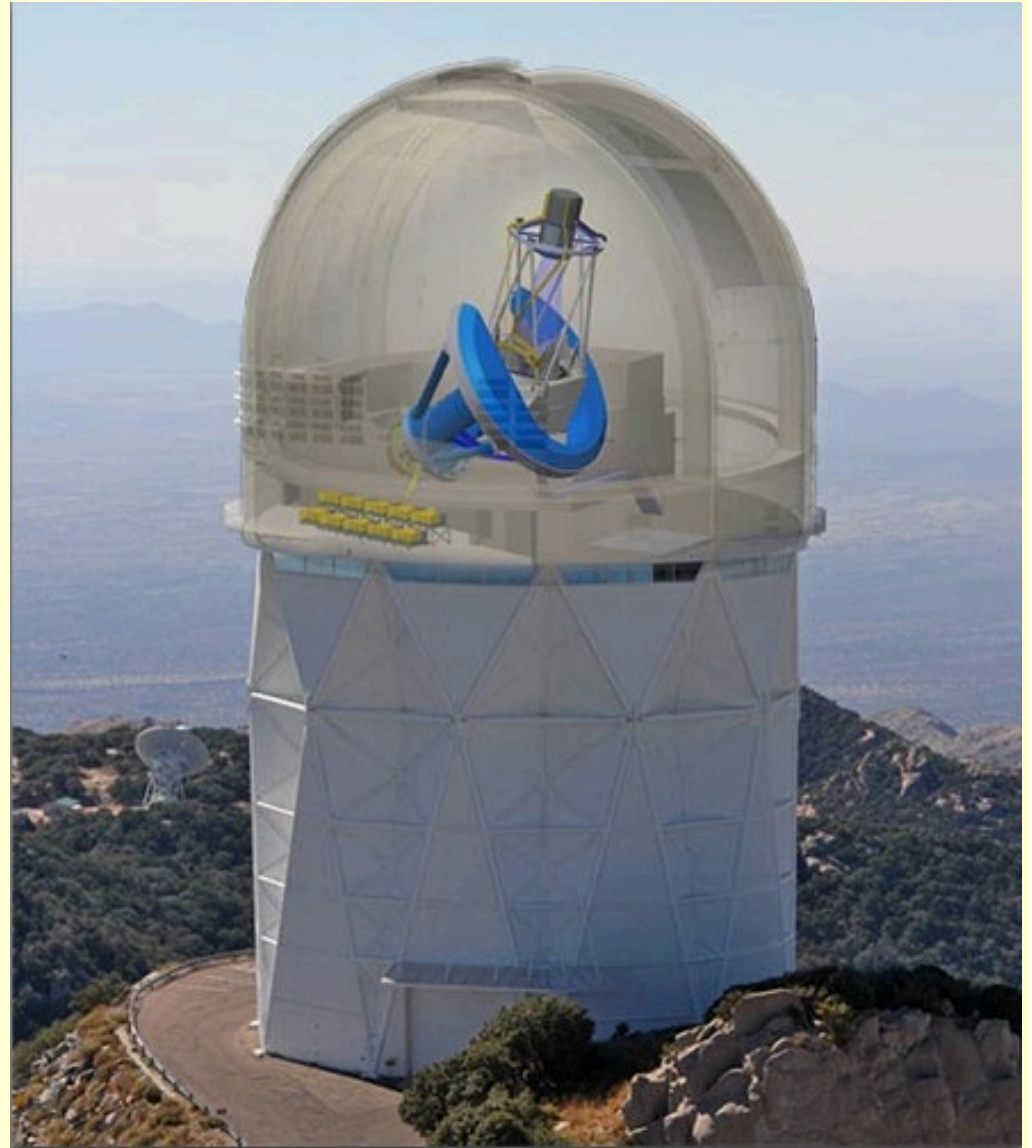


DESI prospects, arXiv:1611.00036



*Dark Energy
Spectroscopic
Instrument*

DESI



The Mayall telescope at Kitt Peak observatory





DESI corrector

summer 2018

*+ commissioning
instrument*

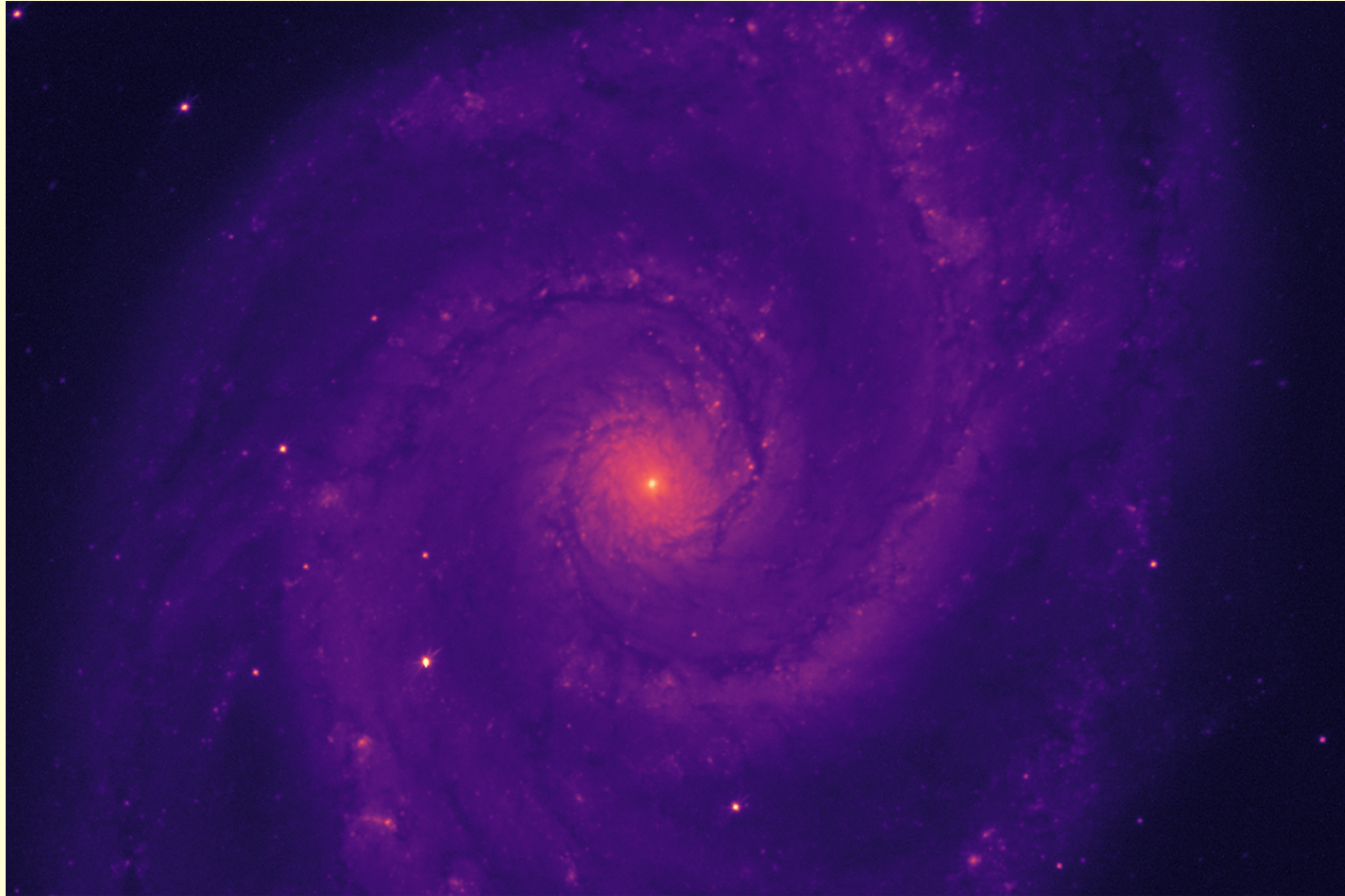
spring 2019





DESI corrector commissioning

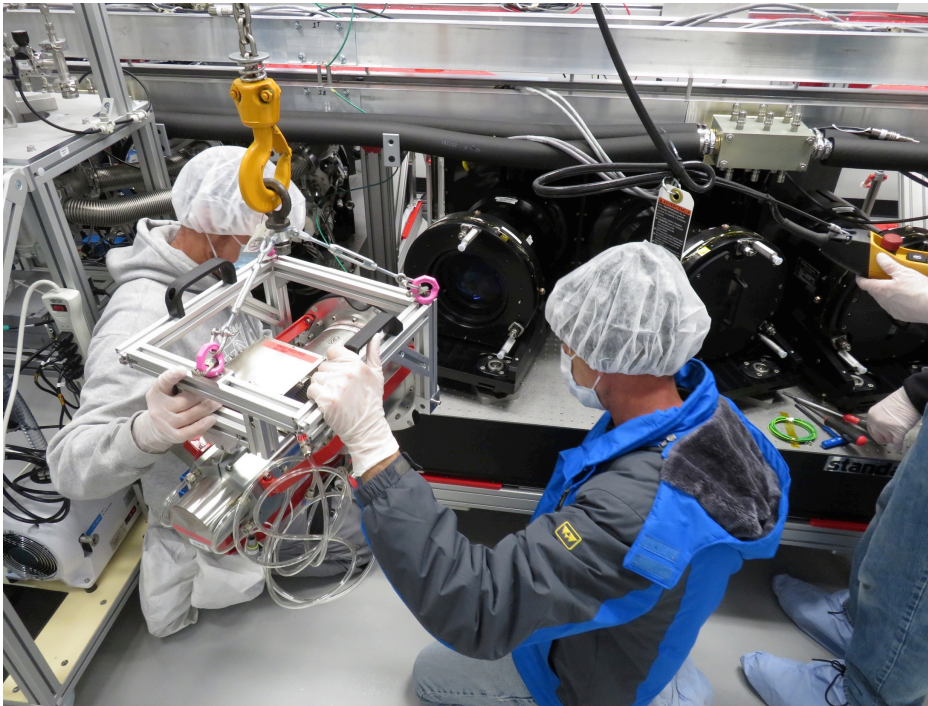
April 2019



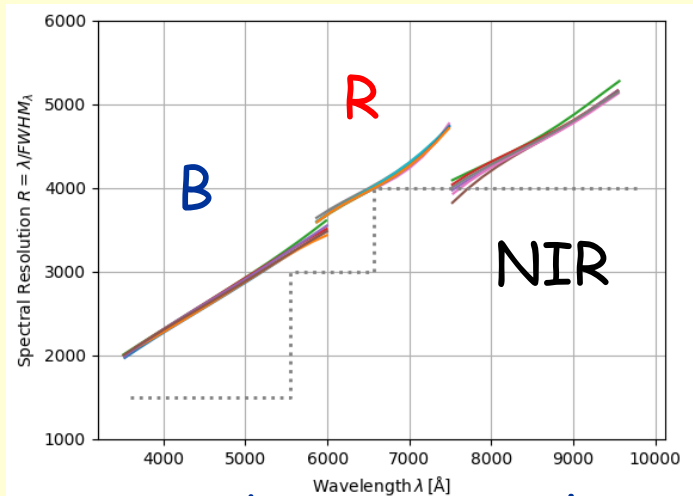
the Whirlpool galaxy seen through the DESI corrector lenses

February 2019

DESI spectrographs



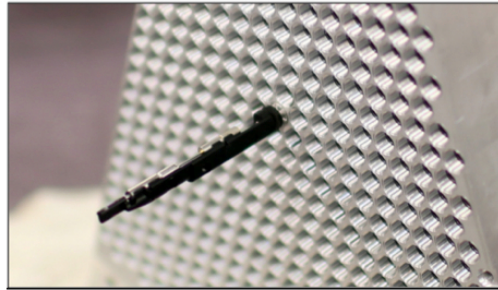
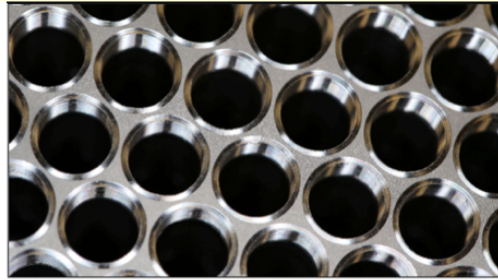
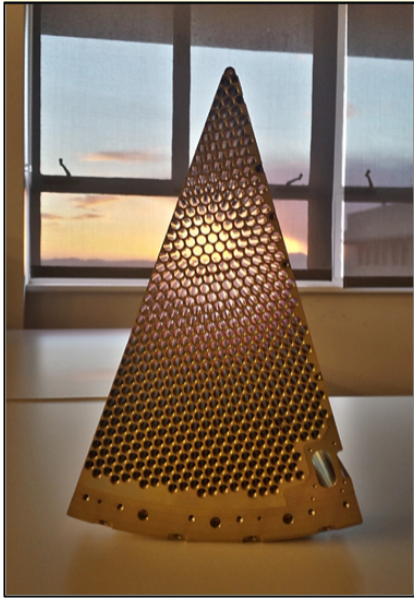
6 spectrographs installed



and operational



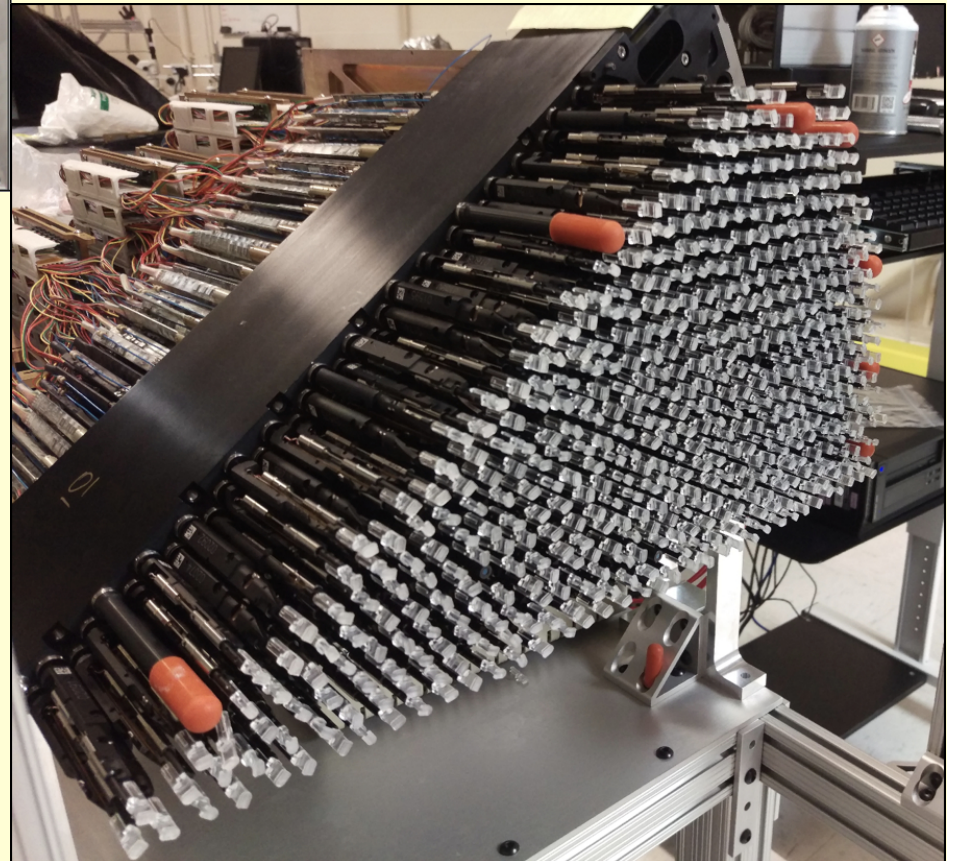
June 2019



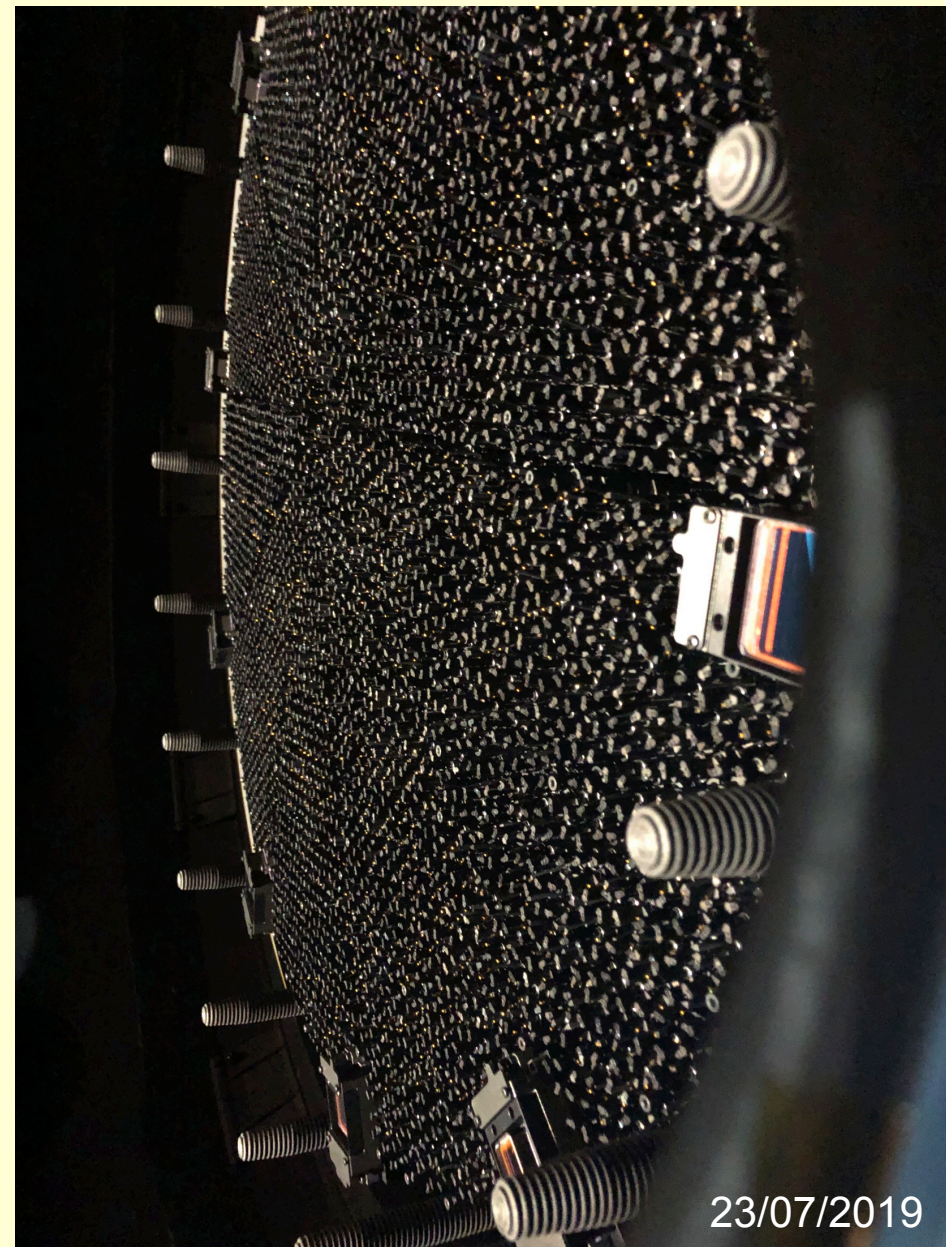
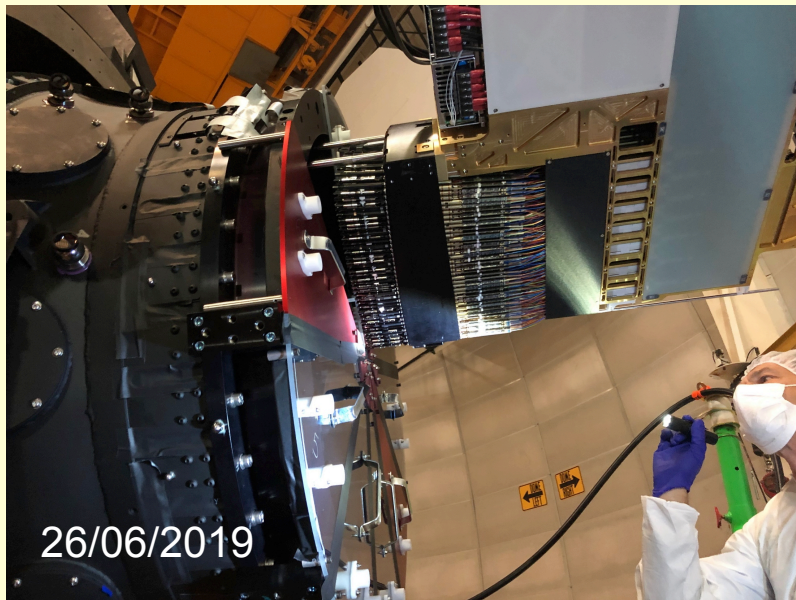
DESI focal plane

Focal Plane = 10 petals, each with 500 fiber-optic cables moved by robotic positioners

each fiber collects light from a separate sky object

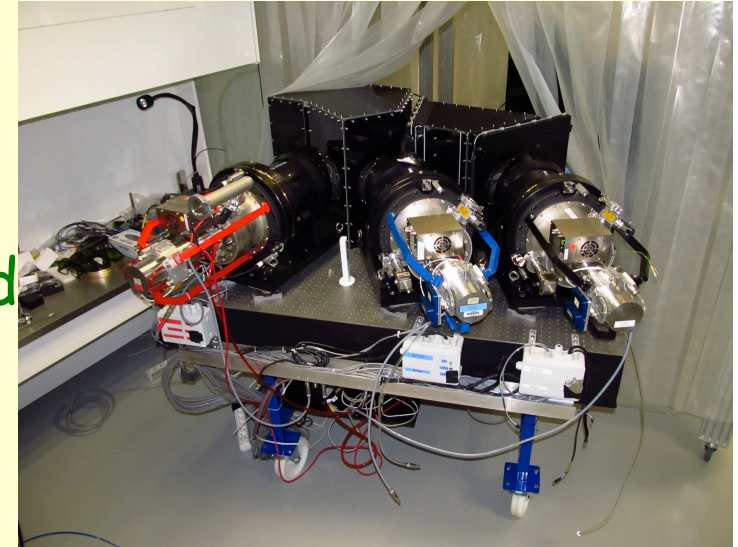


DESI focal plane installation



DESI timeline

- 2010: proposal for a large science program on the Mayall telescope at Kitt Peak observatory (Arizona)
- 2010-2016: R&D and prototype period
- 2016-2017: construction



- 2018-2019: installation at the telescope
- Oct 2019-Feb 2020: commissioning
- Feb 2020-summer 2020: validation survey
- 2020-2025: science survey

CONCLUSIONS (4)

- Cosmological measurements **today**: CMB, SNeIa, BAO
- Standard cosmological model: Λ_{CDM}
- Very near future: go beyond BAO with **large scale structure** surveys (DESI, Euclid, LSST, WFIRST...)



The Mayall telescope

DESI prospects

