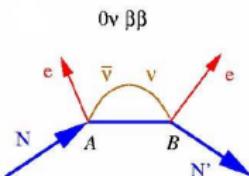


Towards reliable nuclear matrix elements for neutrinoless $\beta\beta$ decay

Frédéric Nowacki



Nuclear physics and neutrinoless $\beta\beta$ decay

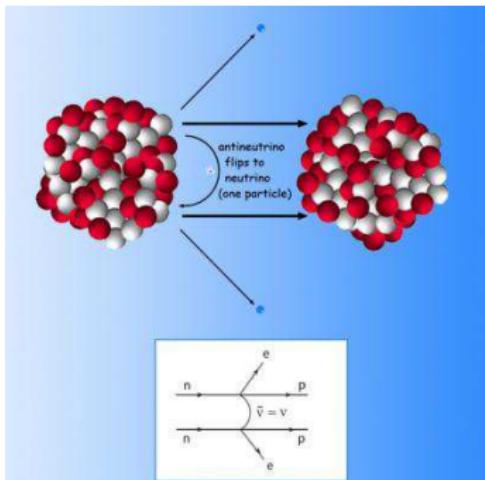
Neutrinos, dark matter studied in experiments using nuclei

Nuclear matrix elements depend on nuclear structure crucial to anticipate reach and fully exploit experiments

$$0\nu\beta\beta \text{ decay: } [T_{1/2}^{0\nu}]^{-1} \propto |M^{0\nu}|^2 \langle m_\nu \rangle^2$$

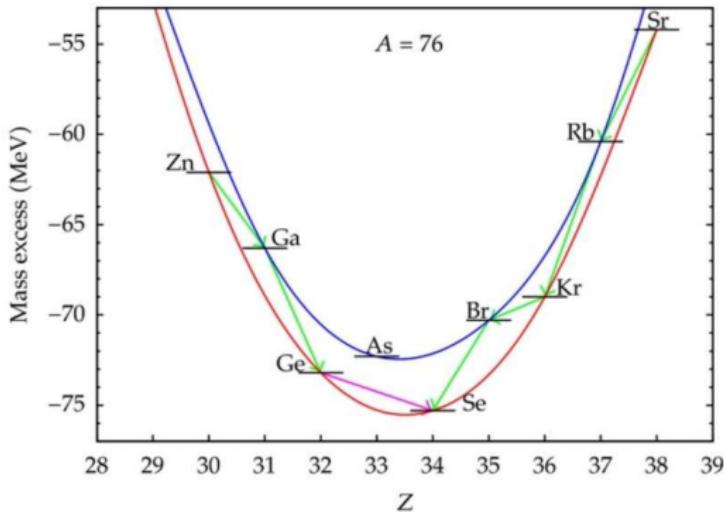
$$\text{Dark matter: } \frac{d\sigma_{\chi N}}{dq^2} \propto |\sum_i c_i \zeta_i \mathcal{F}_i|^2$$

$M^{0\nu}$: Nuclear matrix element
 \mathcal{F}_i : Nuclear structure factor



Neutrinoless $\beta\beta$ decay

Lepton-number violation, Majorana nature of neutrinos
 Second order process only observable in rare cases with
 β -decay energetically forbidden or hindered by ΔJ



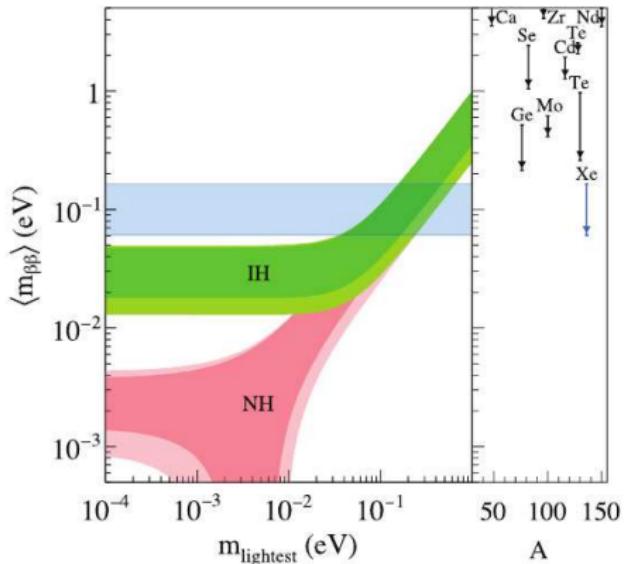
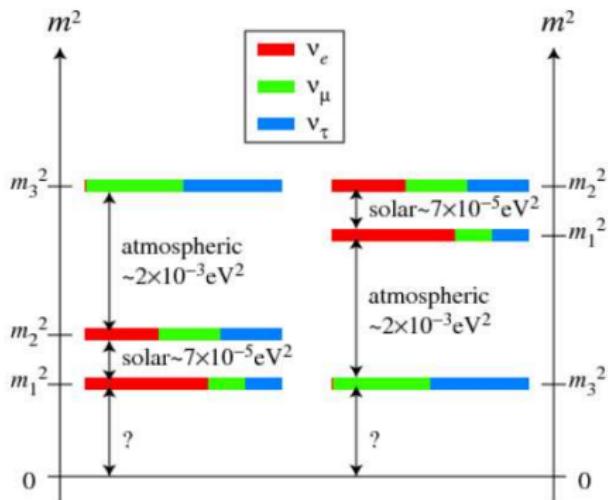
Present best limits $T_{1/2}^{0\nu} \gtrsim 10^{25}$ y.:

^{76}Ge (GERDA, Majorana), ^{130}Te (CUORE), ^{136}Xe (EXO, KamLAND-Zen)

Next generation experiments: inverted hierarchy

The decay lifetime is $[T_{1/2}^{0\nu}(0^+ \rightarrow 0^+)]^{-1} = G_{0\nu}|M^{0\nu}|^2 \langle m_\nu^{\beta\beta} \rangle^2$

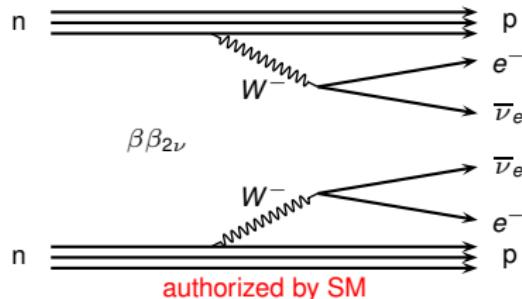
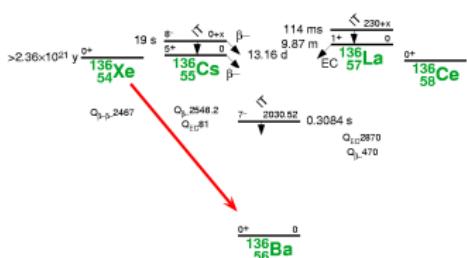
sensitive to absolute neutrino masses, $\langle m_\nu^{\beta\beta} \rangle = \sum_i U_{ei}^2 m_i$



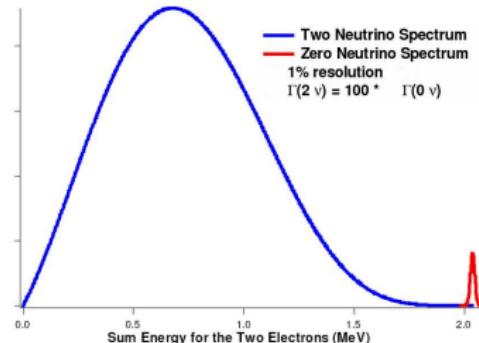
KamLAND-Zen, PRL117 082503 (2016)

Matrix elements needed to make sure next generation ton-scale experiments fully explore “inverted hierarchy”

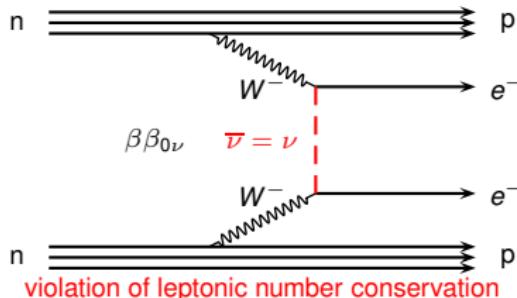
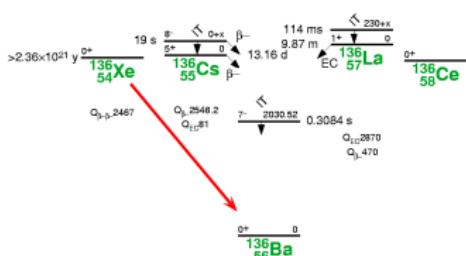
$\beta\beta$ decay



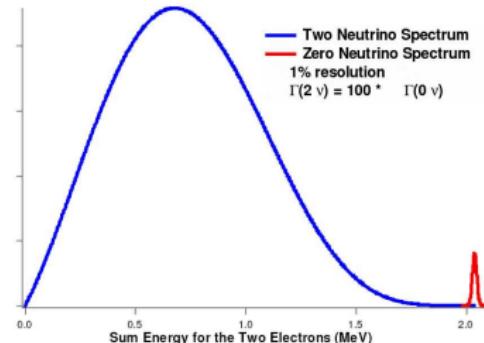
Transition	$Q_{\beta\beta}$ (keV)	Abundance ($^{232}\text{Th} = 100$)
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2013	12
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2040	8
$^{124}\text{Sn} \rightarrow ^{124}\text{Te}$	2288	6
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2479	9
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2533	34
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2802	7
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2995	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3034	10
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	3350	3
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3667	6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4271	0.2



$\beta\beta$ decay



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$(\beta\beta)_{0\nu}$ decay

Specificity of $(\beta\beta)_{0\nu}$:

NO EXPERIMENTAL DATA !!!

prediction for m_ν very **difficult**
easier for $m_\nu(A)/m_\nu(A')$

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What is the best isotope to observe $(\beta\beta)_{0\nu}$ decay ?

What is the influence of the structure of the nucleus on $(\beta\beta)_{0\nu}$ matrix elements ?

Calculating nuclear matrix elements

Nuclear matrix elements needed to study fundamental symmetries

$$\langle \text{Final} | \mathcal{L}_{\text{leptons-nucleus}} | \text{Initial} \rangle = \langle \text{Final} | dx j^\mu(x) J_\mu(x) | \text{Initial} \rangle$$

- Nuclear structure calculation of the initial and final states:

Shell model Betamosa Gaurier FN

Energy-density functional Rodriguez, Yao...

QRPA Vogel Faesller Simkovic Subonen

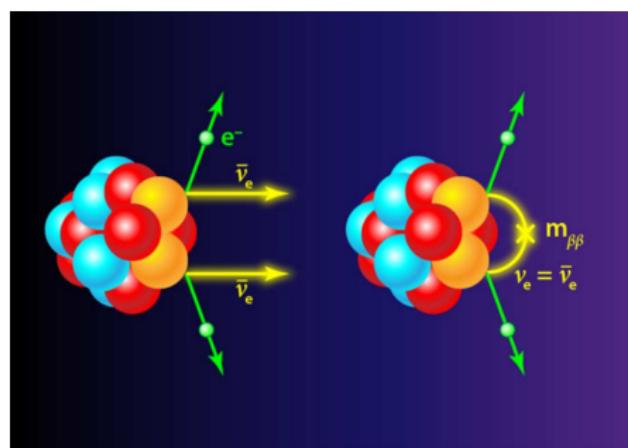
Interacting boson model Iachello, Barea...

Ab Initio many-body methods

Green's Function MC-Coupled-Cluster-JM-SBG

- Lepton-nucleus interaction:

Study hadronic current in nucleus: phenomenological approaches, effec- tive theory of QCD



Two neutrinos mode

The theoretical expression of the half-life of the 2ν mode can be written as:

$$[T_{1/2}^{2\nu}]^{-1} = G_{2\nu} |M_{GT}^{2\nu}|^2,$$

with

$$M_{GT}^{2\nu} = \sum_m \frac{\langle 0_f^+ || \vec{\sigma} t_- || 1_m^+ \rangle \langle 1_m^+ || \vec{\sigma} t_- || 0_i^+ \rangle}{E_m + E_0}$$

- $G_{2\nu}$ contains the phase space factors and the axial coupling constant g_A

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- to quench or not to quench ? ($\sigma\tau_{\text{eff.}}$)

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- $G_{2\nu}$ contains the phase space factors and the axial coupling constant g_A
- summation over intermediate states
- to quench or not to quench ? ($\sigma\tau_{eff.}$)
- does a good 2ν ME guarantee a good 0ν ME ?

$(\beta\beta)_{2\nu}$ structure function

$$M_{GT}^{2\nu} = \sum_m \frac{\langle \ 0_f^+ || \vec{\sigma} t_- || 1_m^+ \rangle \langle 1_m^+ || \vec{\sigma} t_- || \ 0_i^+ \ \rangle}{E_m + E_0}$$

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Calculation in three steps:

- calculate the final and initial states

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Calculation in three steps:

- calculate the **final** and initial states

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- calculate the final and initial states
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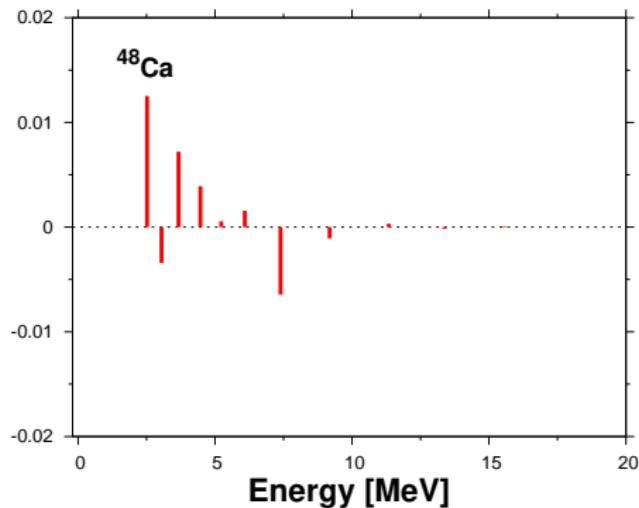
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- take a doorway and use of Lanczos Strength Function method:

- at iteration N, N 1^+ states in the intermediate nucleus, with excitation energies E_m
- overlap with the other doorway, enter energy denominators and add up the N contributions

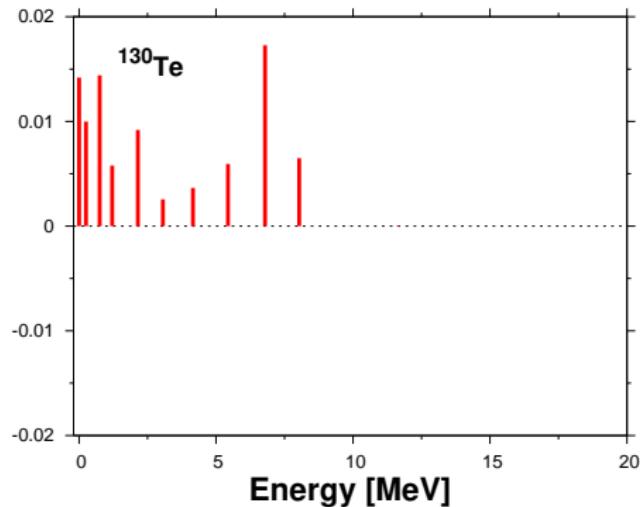
2ν half-lives



2ν strength function in ^{48}Ca , ^{130}Te and ^{136}Xe

Parent nuclei	^{48}Ca	^{76}Ge	^{82}Se	^{130}Te	^{136}Xe
$T_{1/2}^{2\nu}(g.s.)$ th.	$3.7E19$	$1.15E21$	$3.4E19$	$4E20$	$6E20$
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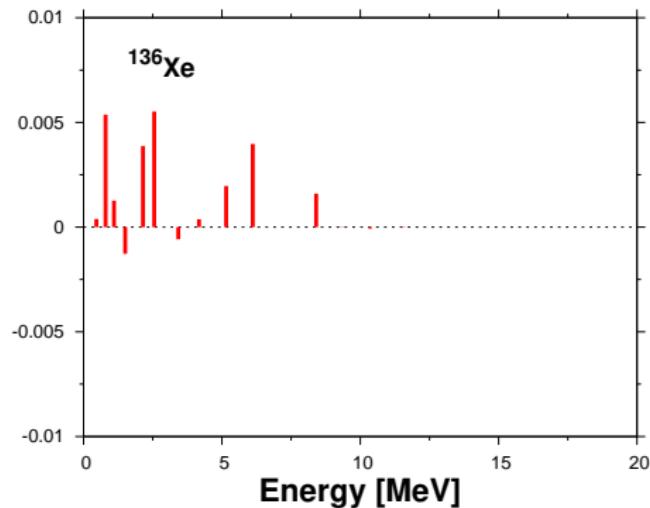
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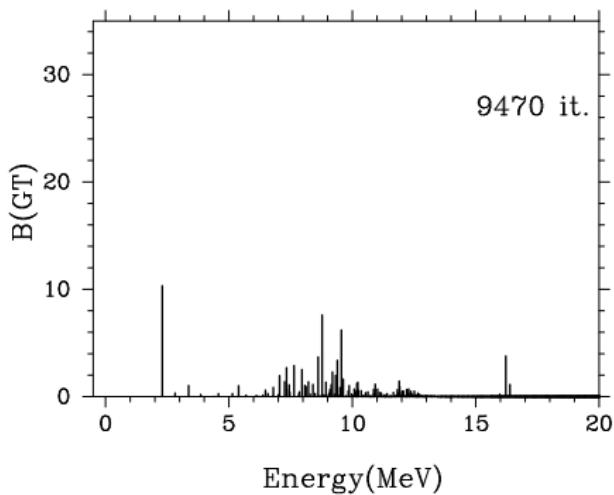
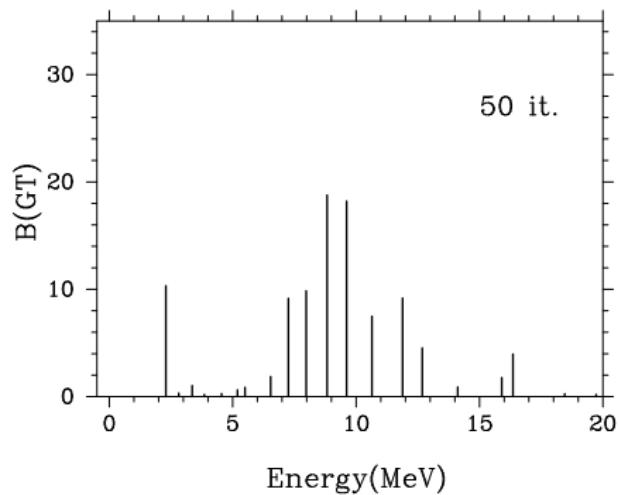


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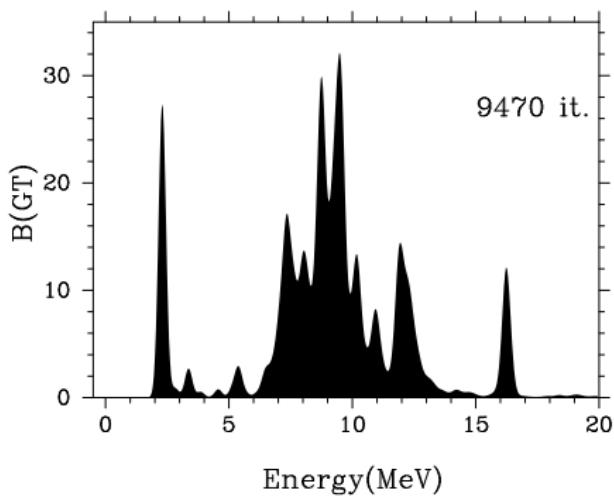
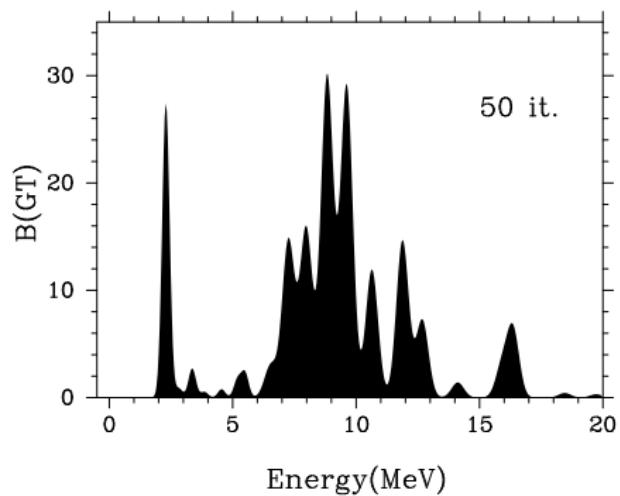
Evolution of Strength Distribution

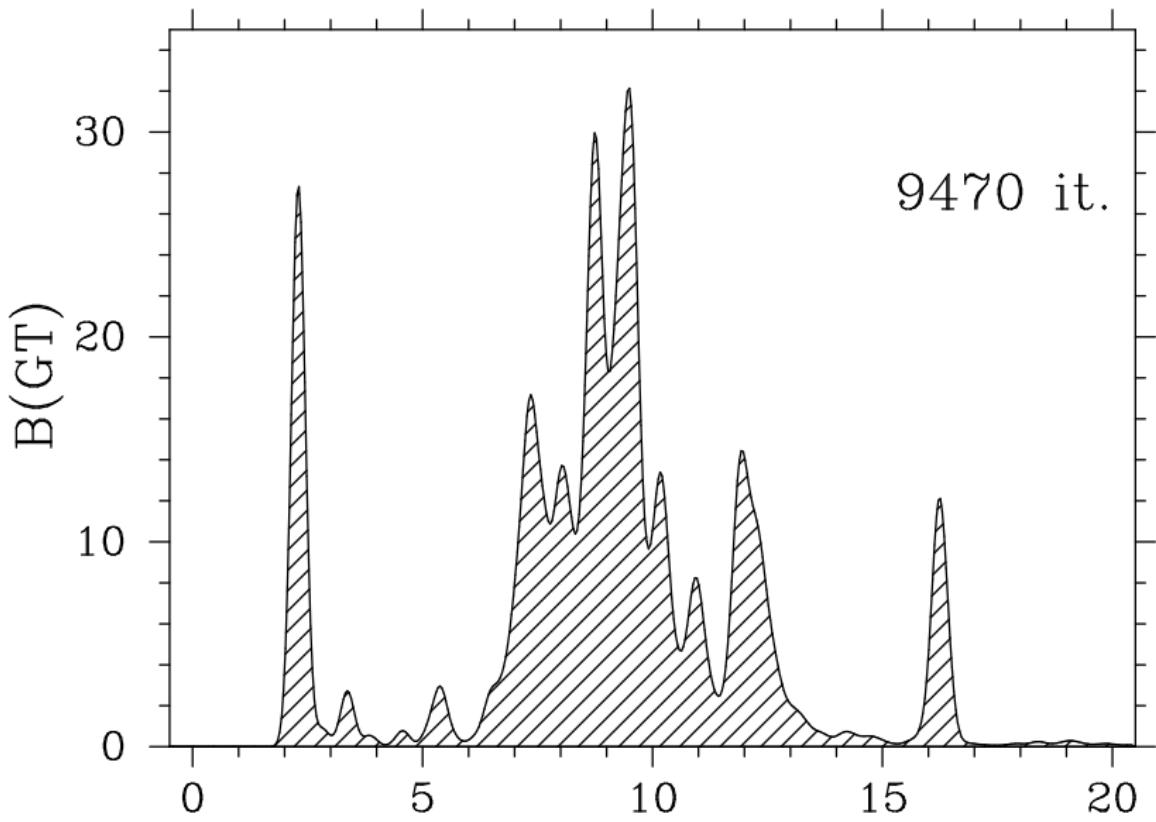
GT Strength in ^{48}Sc

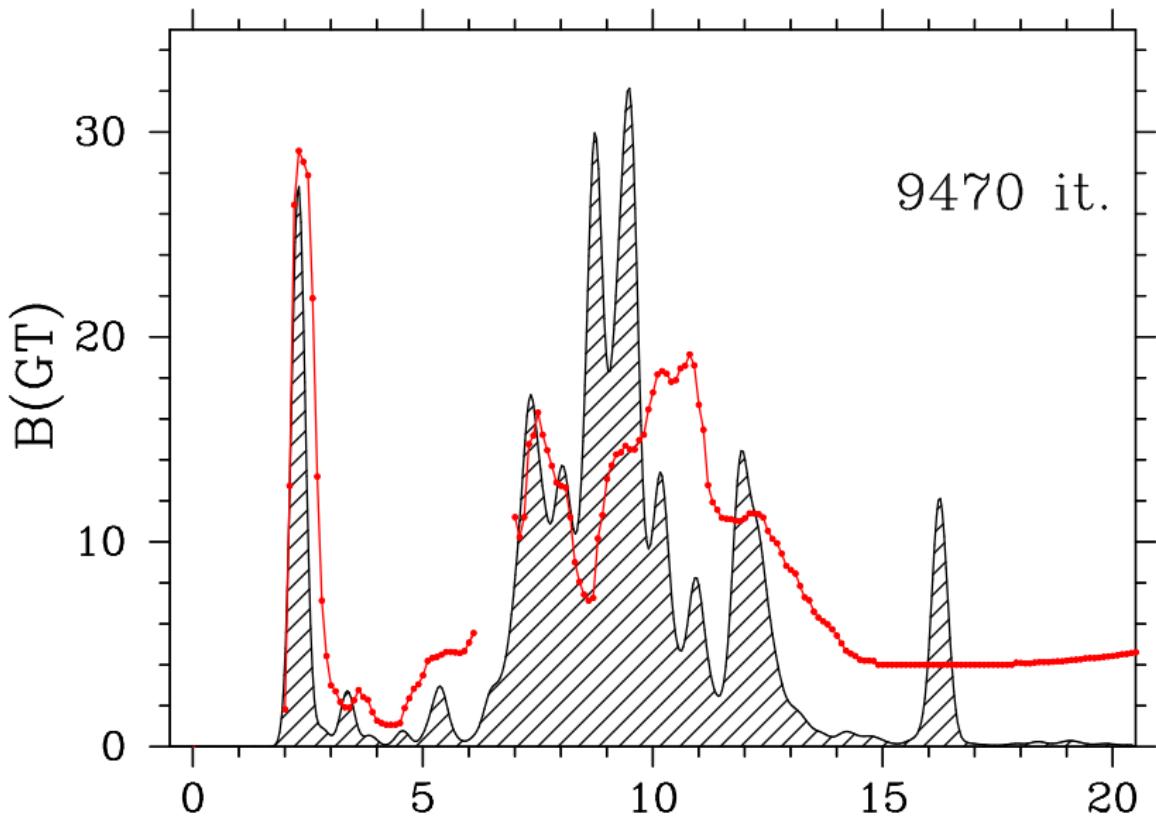


Evolution of Strength Distribution

GT Strength in ^{48}Sc



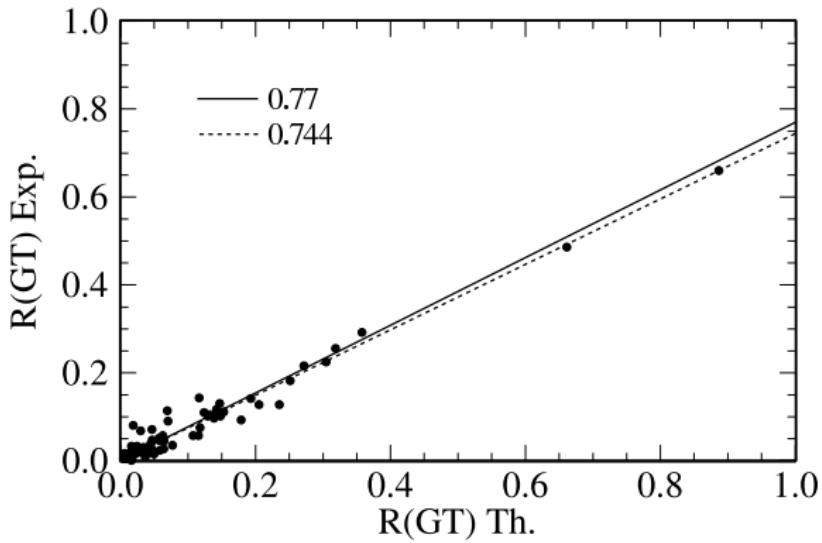
$^{48}\text{Ca}(\text{p},\text{n})^{48}\text{Sc}$ Strength Function

$^{48}\text{Ca}(\text{p},\text{n})^{48}\text{Sc}$ Strength Function

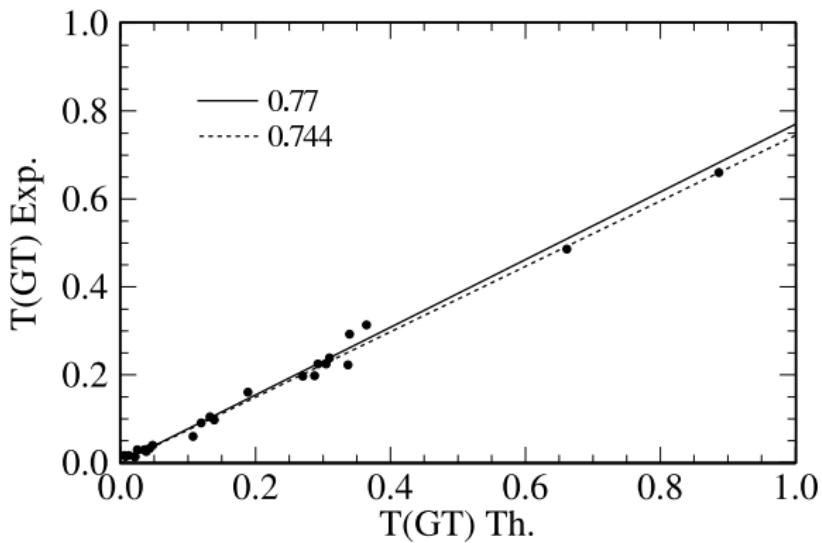
Quenching of GT operator in the *pf*-shell

Nucleus	Uncorrelated	Correlated		Expt.
		Unquenched	$Q = 0.74$	
⁵¹ V	5.15	2.42	1.33	1.2 ± 0.1
⁵⁴ Fe	10.19	5.98	3.27	3.3 ± 0.5
⁵⁵ Mn	7.96	3.64	1.99	1.7 ± 0.2
⁵⁶ Fe	9.44	4.38	2.40	2.8 ± 0.3
⁵⁸ Ni	11.9	7.24	3.97	3.8 ± 0.4
⁵⁹ Co	8.52	3.98	2.18	1.9 ± 0.1
⁶² Ni	7.83	3.65	2.00	2.5 ± 0.1

Quenching of GT strength in the *pf*-shell

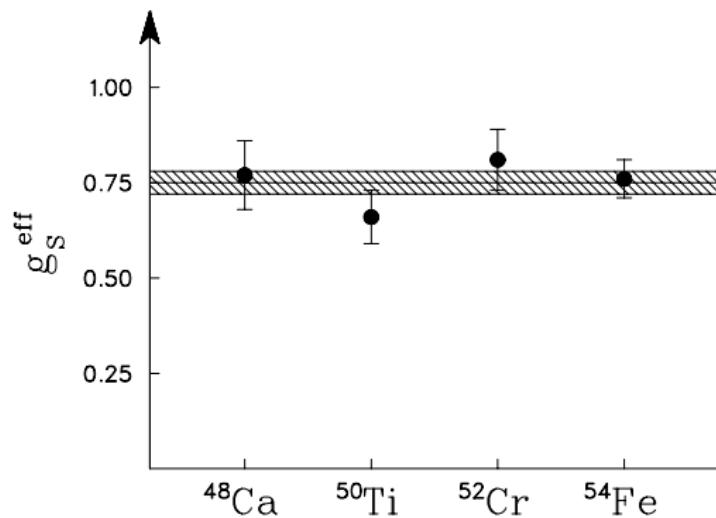


Quenching of GT strength in the *pf*-shell



Quenching of M1 operator in the *pf*-shell

KB3 interaction

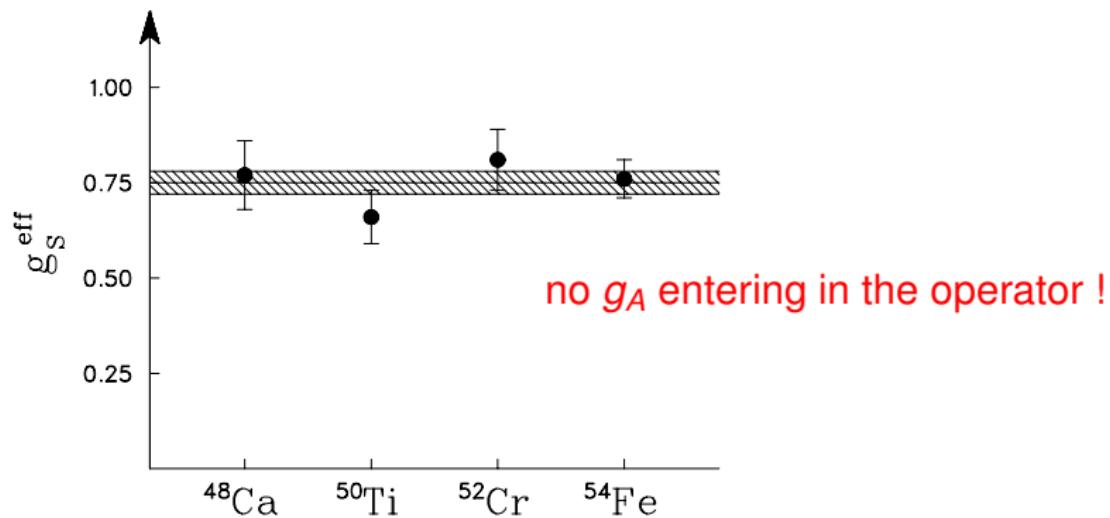


Neumann-Cosel et al.

Phys. Lett. **B433** 1 (1998)

Quenching of M1 operator in the *pf*-shell

KB3 interaction



Neumann-Cosel et al.

Phys. Lett. **B433** 1 (1998)

Quenching of M1 operator in the pf -shell

KB3 interaction

Definitions

The talk is based on "Value of the axial-vector coupling strength in β and $\beta\beta$ decays: A review" to appear in **Frontiers in Physics**.

Quenching:

$$q = g_A/g_A^{\text{free}}$$

Free value of g_A (Particle Data Group 2016):

$$g_A^{\text{free}} = 1.2723(23)$$

Effective value of g_A :

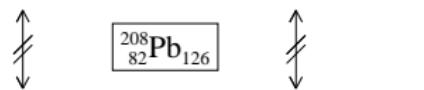
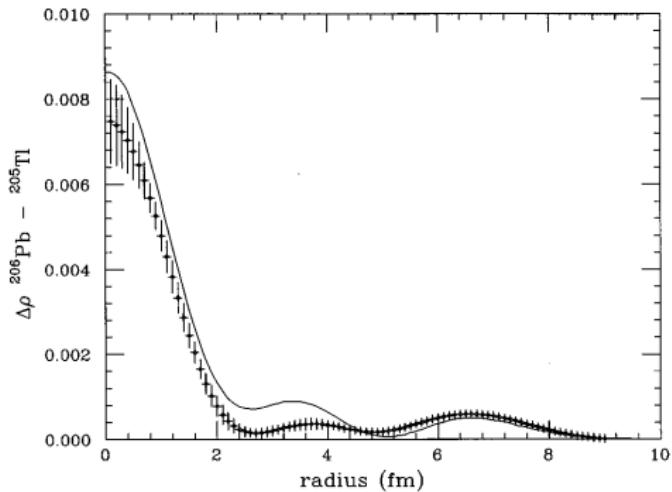
$$g_A^{\text{eff}} = q g_A^{\text{free}}$$

From J. Suhonen (VTT, Finland)

CNNP2017 5/38

Notes From J. Suhonen, CNNP 2017, Catania
 Phys. Lett. **B455** 1 (1999)

Quenching of GT operator in the *pf*-shell



$3s_{1/2}$		—	-8013
$2d_{3/2}$		—	-8364

$1h_{11/2}$		—	-9361
$2d_{5/2}$		—	-9696

$1h_{9/2}$		—	-10781
------------	--	---	--------

$1g_{7/2}$		—	-11487
------------	--	---	--------

Protons

$^{132}_{50}\text{Sn}_{82}$

Neutrons

FIG. 3. Density difference between ^{206}Pb and ^{205}Tl . The experimental result of Cavendon *et al.* (1982) is given by the error bars; the prediction obtained using Hartree-Fock orbitals with adjusted occupation numbers is given by the curve. The systematic shift of 0.0008 fm^{-3} at $r \leq 4 \text{ fm}$ is due to deficiencies of the calculation in predicting the core polarization effect.

V. R. Pandharipande, I. Sick and P. K. A. deWitt
Huberts, Rev. mod. Phys. **69** (1997) 981

Quenching of GT operator in the pf -shell

If we write

$$|\hat{i}\rangle = \alpha|0\hbar\omega\rangle + \sum_{n \neq 0} \beta_n|n\hbar\omega\rangle,$$

$$|\hat{f}\rangle = \alpha'|0\hbar\omega\rangle + \sum_{n \neq 0} \beta'_n|n\hbar\omega\rangle$$

then

$$\langle \hat{f} \parallel \mathcal{T} \parallel \hat{i} \rangle^2 = \left(\alpha \alpha' T_0 + \sum_{n \neq 0} \beta_n \beta'_n T_n \right)^2,$$

- $n \neq 0$ contributions negligible

- $\alpha \approx \alpha'$

 projection of the physical wavefunction in the $0\hbar\omega$ space is $Q \approx \alpha^2$

 transition quenched by Q^2

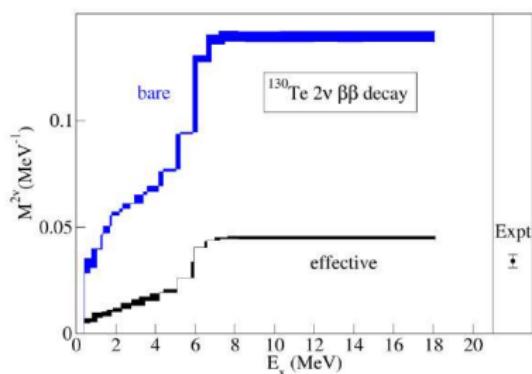
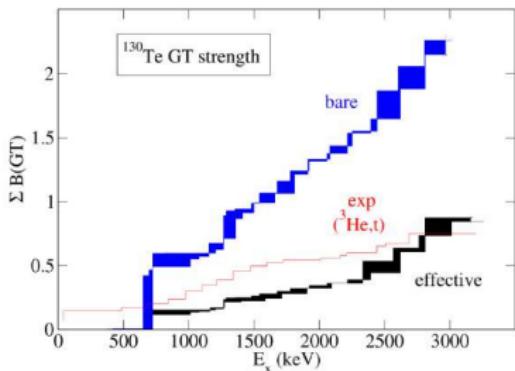
Renormalisation of the GT operator by MBPT

PHYSICAL REVIEW C 95, 064324 (2017)



Calculation of Gamow-Teller and two-neutrino double- β decay properties for ^{130}Te and ^{136}Xe with a realistic nucleon-nucleon potential

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Renormalisation of the GT by Many-Body Perturbation Theory

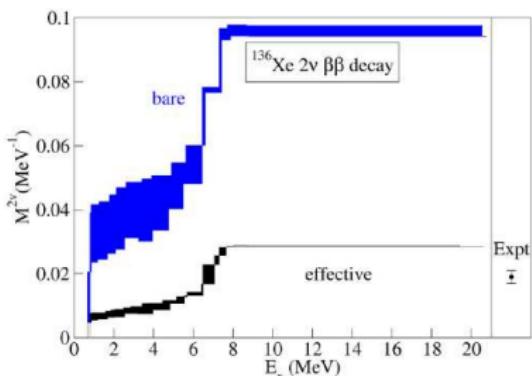
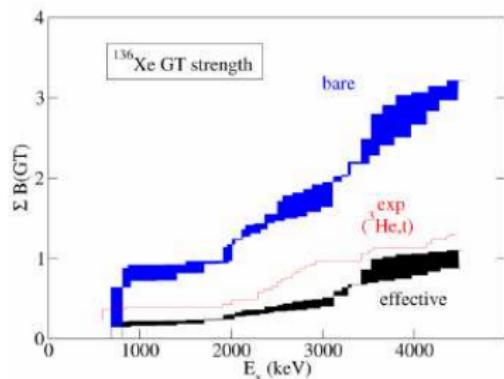
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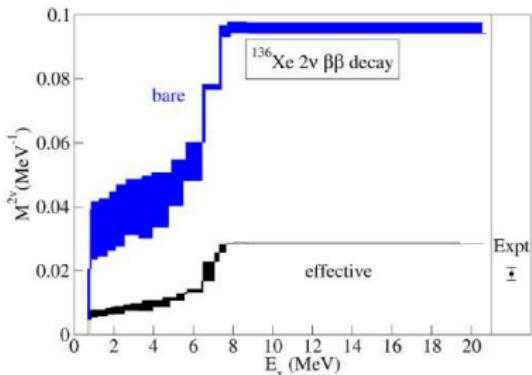
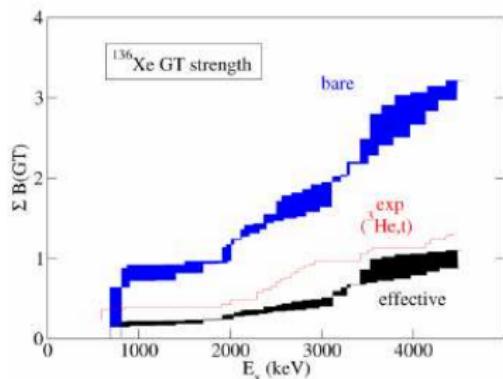
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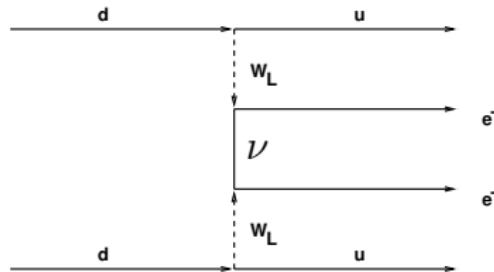
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Renormalisation of the GT by Many-Body Perturbation Theory
Further step, apply MBPT to Neutrinoless operator

Neutrinoless mode:

Exchange of a light neutrino, only left-handed currents



The theoretical expression of the half-life of the 0ν mode can be written as:

$$[T_{1/2}^{0\nu}(0^+ \rightarrow 0^+)]^{-1} = G_{0\nu} |M^{0\nu}|^2 \langle m_\nu \rangle^2$$

Neutrinoless mode:

CLOSURE APPROXIMATION then

$$\langle \Psi_f | |\mathcal{O}^{(K)}| |\Psi_i \rangle \quad \text{with} \quad \mathcal{O}^{(K)} = \sum_{ijkl} W_{ijkl}^{\lambda,K} \left[(a_i^\dagger a_j^\dagger)^\lambda (\tilde{a}_k \tilde{a}_l)^\lambda \right]^K$$

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two-body operator

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two-body operator

We are left with a “standard” nuclear structure problem

$$M^{(0\nu)} = M_{GT}^{(0\nu)} - \left(\frac{g_V}{g_A}\right)^2 M_F^{(0\nu)} - M_T^{(0\nu)}$$

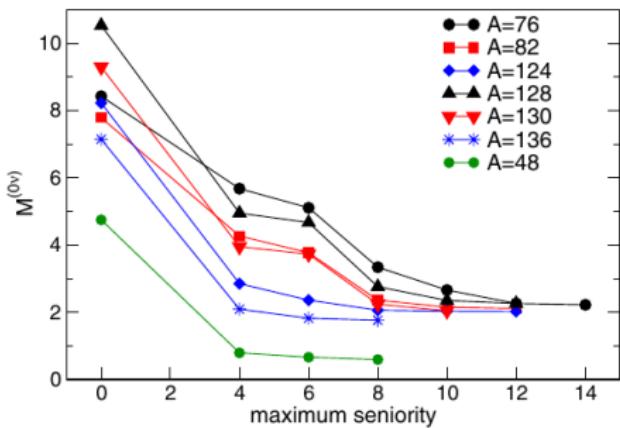
SM results for $(\beta\beta)_{0\nu}$

emitter	$\langle m_\nu \rangle$ ($T_{\frac{1}{2}} = 10^{25}$ y.)	$M_{0\nu}^{tot}$ (UCOM)
⁴⁸ Ca	0.63	0.85
⁷⁶ Ge	0.72	2.81
⁸² Se	0.37	2.64
⁹⁶ Zr		
¹⁰⁰ Mo		
¹¹⁰ Pd		
¹¹⁶ Cd	0.46	1.60
¹²⁴ Sn	0.37	2.62
¹²⁸ Te	1.32	2.88
¹³⁰ Te	0.28	2.65
¹³⁶ Xe	0.38	2.19
¹⁵⁰ Nd	heavy and deformed !	

Pairing correlations and $0\nu\beta\beta$ decay

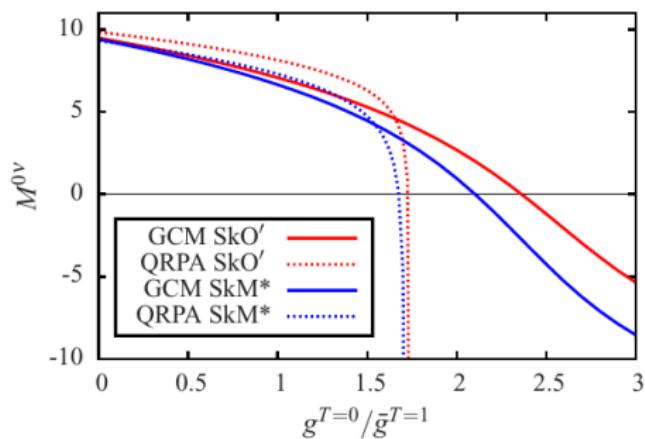
$0\nu\beta\beta$ decay favoured by proton-proton, neutron-neutron pairing,
but it is disfavored by proton-neutron pairing

Ideal case: superfluid nuclei
reduced with high-seniorities



E. Caurier et al., PRL100 052503 (2008)

Addition of isoscalar pairing
reduces matrix element value



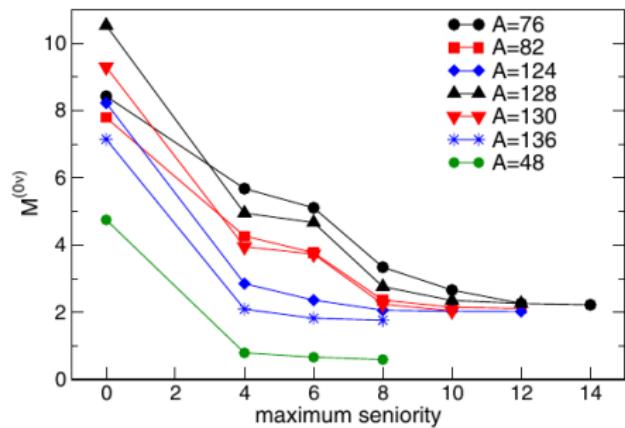
Hinohara, Engel, PRC 90 031301 (2014)

Related to approximate SU(4) symmetry of the $\sum H(r)\sigma_i\sigma_j\tau_i\tau_j$ operator

Pairing correlations and $0\nu\beta\beta$ decay

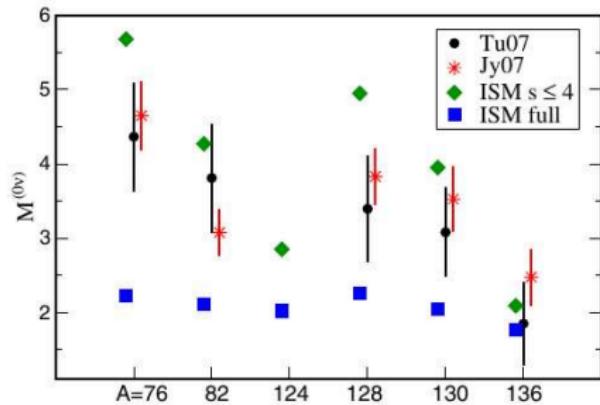
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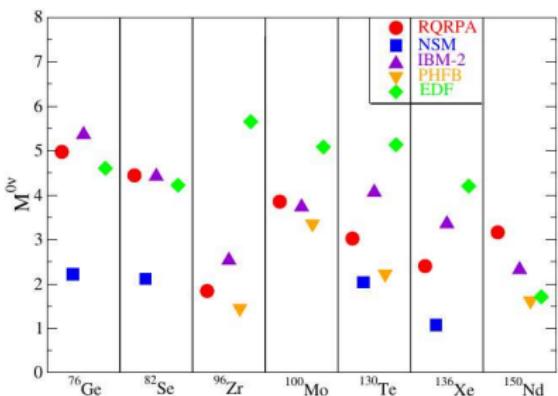


E. Caurier et al., PRL100 052503 (2008)

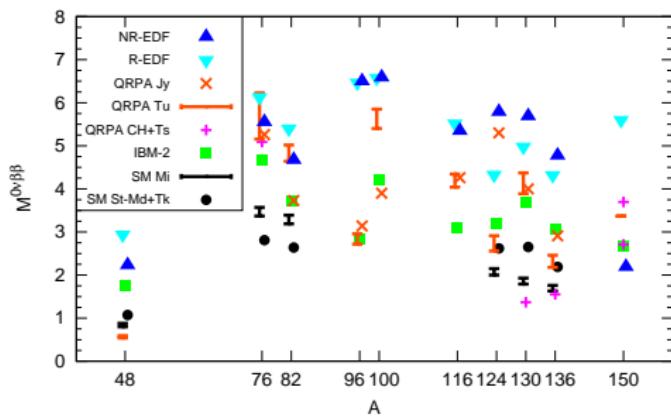
Related to approximate SU(4) symmetry of the $\sum H(r)\sigma_i\sigma_j\tau_i\tau_j$ operator

$0\nu\beta\beta$ matrix elements: last 5 years

Comparison of nuclear matrix elements calculations: 2012 vs 2017



P. Vogel, J. Phys. G39 124002 (2012)



J. Engel, Rep. Prog. Phys. 80 046301 (2017)

What have we learned in the last 5 years ?

Shell model configuration space: two shells

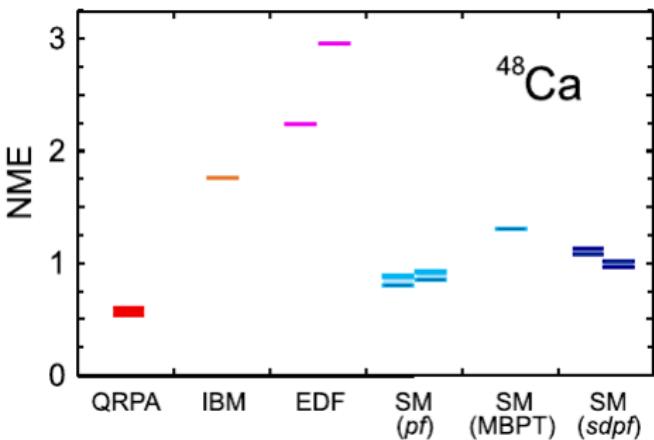
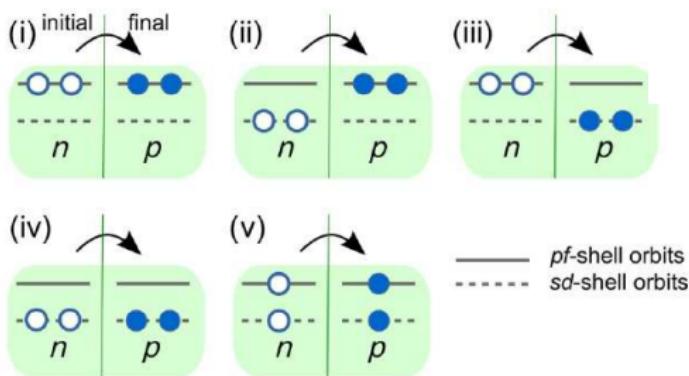
For ^{48}Ca enlarge configuration space from pf to $sdpf$

4 to 7 orbitals, dimension 10^5 to 10^9

increases matrix elements

but only moderately 30%

Iwata et al. PRL116 112502 (2016)



Contributions dominated by pairing

2 particle - 2 hole excitations

enhance the $\beta\beta$ matrix element,

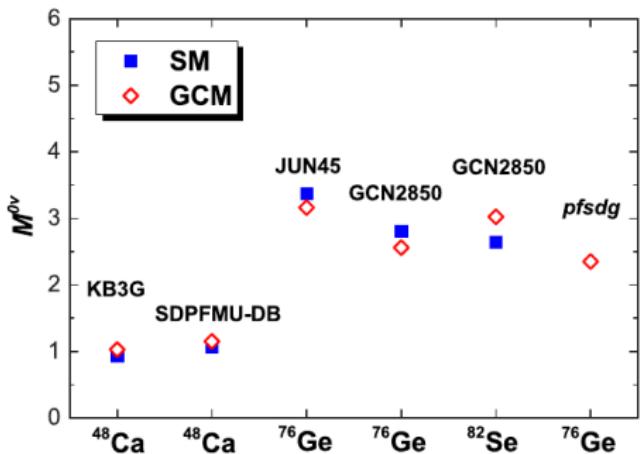
Contributions dominated by

1 particle - 1 hole excitations

suppress the $\beta\beta$ matrix element

^{76}Ge matrix element in two shells: approximate

Large configuration space calculations in 2 major oscillator shells include all relevant correlations: isovector/isoscalar pairing, deformation
 Many-body approach: Generating Coordinate Method (GCM)



GCM approximates shell model calculation

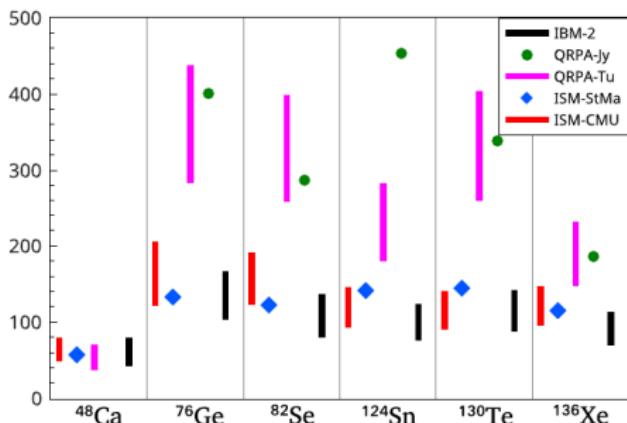
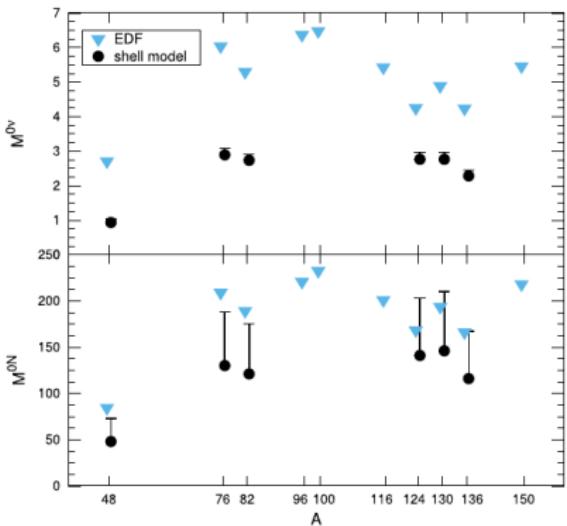
Degrees of freedom,
 or generating coordinates
 validated against
 exact shell model
 in restricted configuration space

Jiao et al., PRC96 054310 (2017)

^{76}Ge nuclear matrix elements in 2 major shells
 very similar to shell model nuclear matrix element in 1 major shell

Heavy-neutrino exchange nuclear matrix elements

Contrary to light-neutrino exchange, for heavy-neutrino exchange decay shell model, IBM and EDF matrix elements agree reasonably!



A. Neacsu et al., PRC 93 024308 (2016)

Suggests differences in treating
longer-range nuclear correlations
dominant in light-neutrino exchange

Heavy-neutrino matrix element

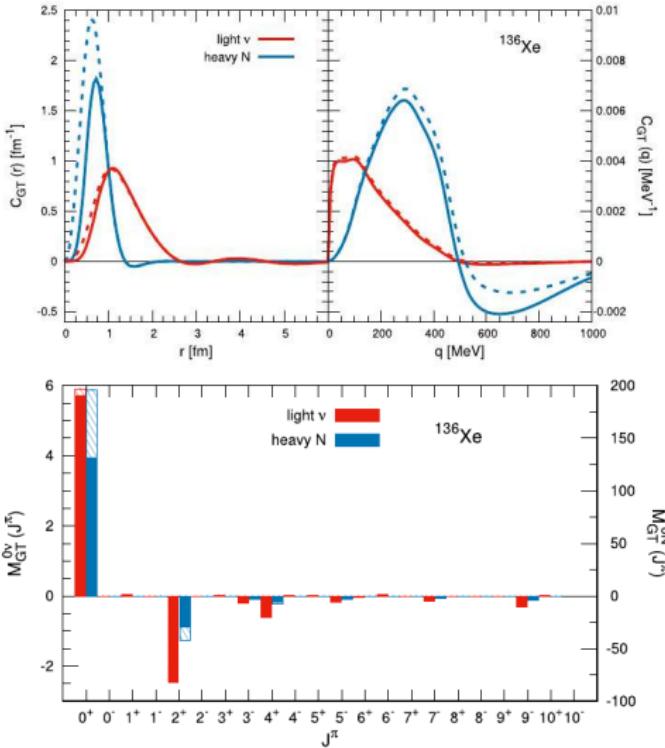
Compared to
light-neutrino exchange

heavy neutrino exchange
dominated by shorter inter-
nucleon range,
larger momentum transfers

heavy neutrino exchange
contribution
from $J > 0$ pairs smaller:
pairing most relevant

Long-range correlations
(except pairing)
not under control

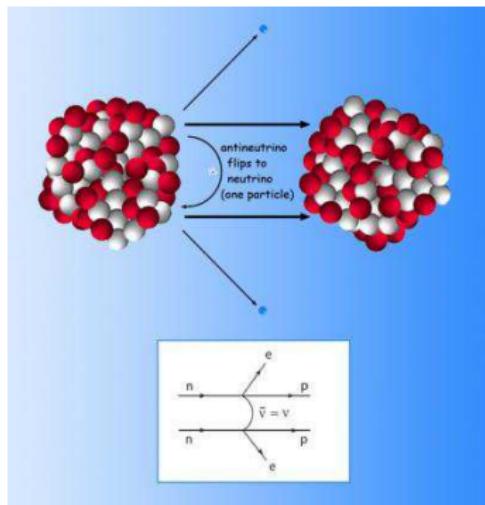
J. Menendez, JPG 45 014003 (2018)



Summary

Reliable nuclear matrix elements needed to plan and fully exploit impressive experiments looking for neutrinoless double-beta decay

- Matrix elements differences between present calculations, factor 2-3 besides additionnal “quenching” ?
- ^{48}Ca and ^{76}Ge matrix elements in larger configuration space increase $\lesssim 30\%$, missing correlations introduced in IBM, EDF
- First Ab-initio calculations of β decays do not need additionnal “quenching”, Ab-initio ^{48}Ca matrix elements in progress
- $2\nu\beta\beta$ decay, μ -capture/ ν -nucleus scattering and double Gamow-Teller transitions can give insight on $0\nu\beta\beta$ matrix elements



$(\beta\beta)_{0\nu}$ matrix elements

$$\begin{aligned}
 M_{GT}^{(0\nu)} &= \langle 0_f^+ | \sum_{n,m} h(\sigma_n \cdot \sigma_m) t_{n-} t_{m-} | 0_i^+ \rangle, & \chi_F &= \langle 0_f^+ | \sum_{n,m} h t_{n-} t_{m-} | 0_i^+ \rangle \left(\frac{g_V}{g_A} \right)^2 / M_{GT}^{(0\nu)}, \\
 \chi'_{GT} &= \langle 0_f^+ | \sum_{n,m} h'(\sigma_n \cdot \sigma_m) t_{n-} t_{m-} | 0_i^+ \rangle / M_{GT}^{(0\nu)}, & \chi'_F &= \langle 0_f^+ | \sum_{n,m} h' t_{n-} t_{m-} | 0_i^+ \rangle \left(\frac{g_V}{g_A} \right)^2 / M_{GT}^{(0\nu)}, \\
 \chi_{GT}^\omega &= \langle 0_f^+ | \sum_{n,m} h_\omega(\sigma_n \cdot \sigma_m) t_{n-} t_{m-} | 0_i^+ \rangle / M_{GT}^{(0\nu)}, & \chi_F^\omega &= \langle 0_f^+ | \sum_{n,m} h_\omega t_{n-} t_{m-} | 0_i^+ \rangle \left(\frac{g_V}{g_A} \right)^2 / M_{GT}^{(0\nu)}, \\
 \chi_T &= \langle 0_f^+ | \sum_{n,m} h' [(\sigma_n \cdot \hat{r}_{n,m})(\sigma_m \cdot \hat{r}_{n,m}) - \frac{1}{3} \sigma_n \cdot \sigma_m] t_{n-} t_{m-} | 0_i^+ \rangle / M_{GT}^{(0\nu)}, \\
 \chi_P &= \langle 0_f^+ | i \sum_{n,m} h' \left(\frac{r_{+n,m}}{2r_{n,m}} \right) [(\sigma_n - \sigma_m) \cdot (\hat{r}_{n,m} \times \hat{r}_{+n,m})] t_{n-} t_{m-} | 0_i^+ \rangle \frac{g_V}{g_A} / M_{GT}^{(0\nu)}, \\
 \chi_R &= \frac{1}{6} (g_{-\frac{1}{2}}^s - g_{\frac{1}{2}}^s) \langle 0_f^+ | \sum_{n,m} h_R(\sigma_n \cdot \sigma_m) t_{n-} t_{m-} | 0_i^+ \rangle \frac{g_V}{g_A} / M_{GT}^{(0\nu)}.
 \end{aligned}$$

back

$(\beta\beta)_{0\nu}$ matrix elements

$$\begin{aligned}
 h(r, \langle \mu \rangle) &= \frac{R_0}{r} \phi(\langle \mu \rangle m_e r), \\
 h'(r, \langle \mu \rangle) &= h + \langle \mu \rangle m_e R_0 h_0(\langle \mu \rangle r), \\
 h_\omega(r, \langle \mu \rangle) &= h - \langle \mu \rangle m_e R_0 h_0(\langle \mu \rangle r), \\
 h_R(r, \langle \mu \rangle) &= -\frac{\langle \mu \rangle m_e}{M_i} \left(\frac{2}{\pi} \left(\frac{R_0}{r} \right)^2 - \langle \mu \rangle m_e R_0 h \right) + \frac{4\pi R_0^2}{M_p} \delta(r), \\
 h_0(x) &= -\frac{d\phi}{dx}(x), \\
 \phi(x) &= \frac{2}{\pi} [\sin(x) C_{int}(x) - \cos(x) S_{int}(x)], \\
 \frac{d\phi}{dx} &= \frac{2}{\pi} [\sin(x) C_{int}(x) + \cos(x) S_{int}(x)].
 \end{aligned}$$

$S_{int}(x)$ and $C_{int}(x)$ being the integral sinus and cosinus functions,

$$S_{int}(x) = - \int_x^\infty \frac{\sin(\zeta)}{\zeta} d\zeta, \quad C_{int}(x) = - \int_x^\infty \frac{\cos(\zeta)}{\zeta} d\zeta$$

back