



NEWS-G: Direct detection of light WIMPs with a Spherical Proportional Counter

Philippe Gros, Queen's University

Séminaire LLR
2018-09-24

Direct detection of light WIMPs with NEWSG

Neutrino Oscillation Workshop 2018

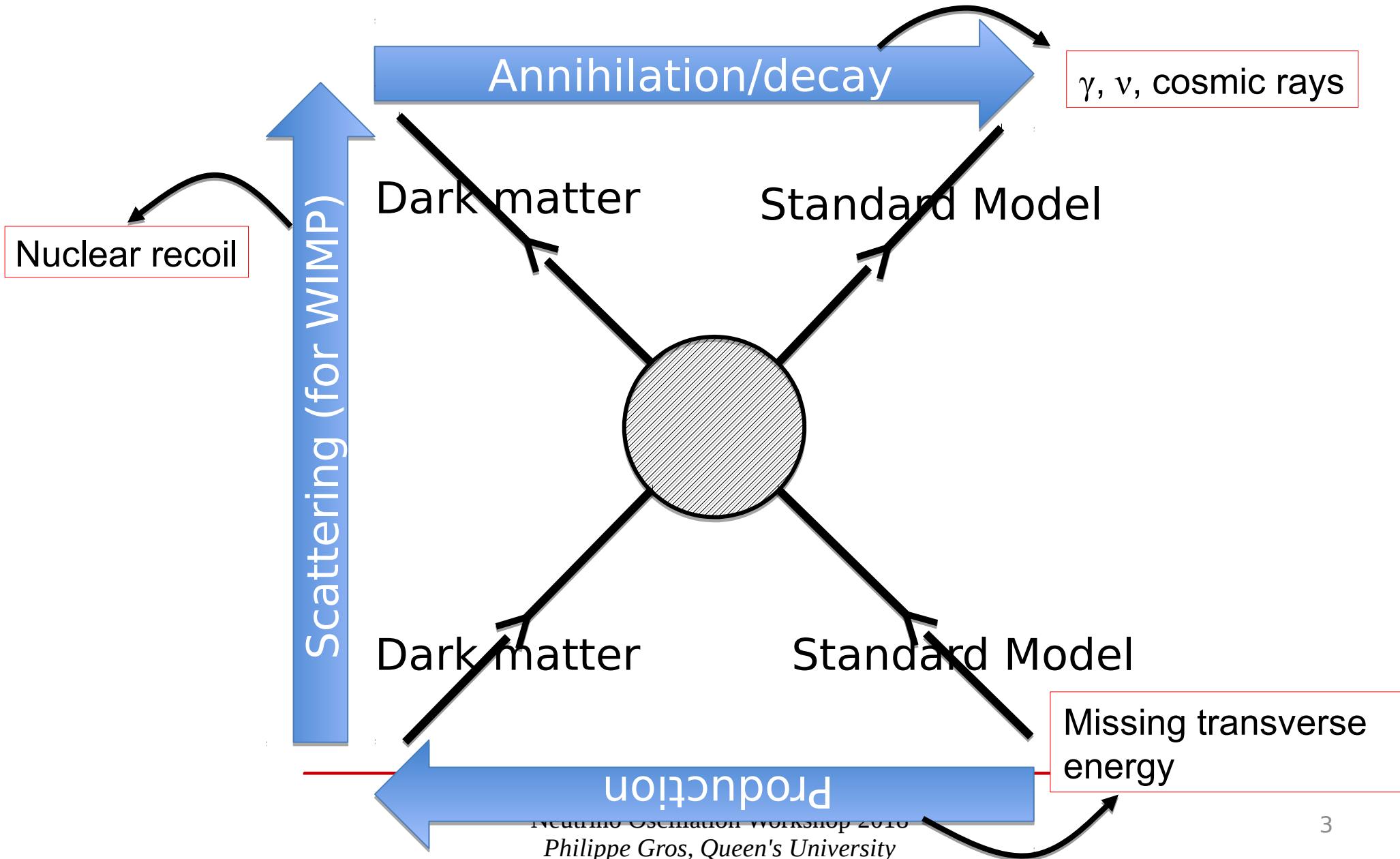
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Dark matter detection

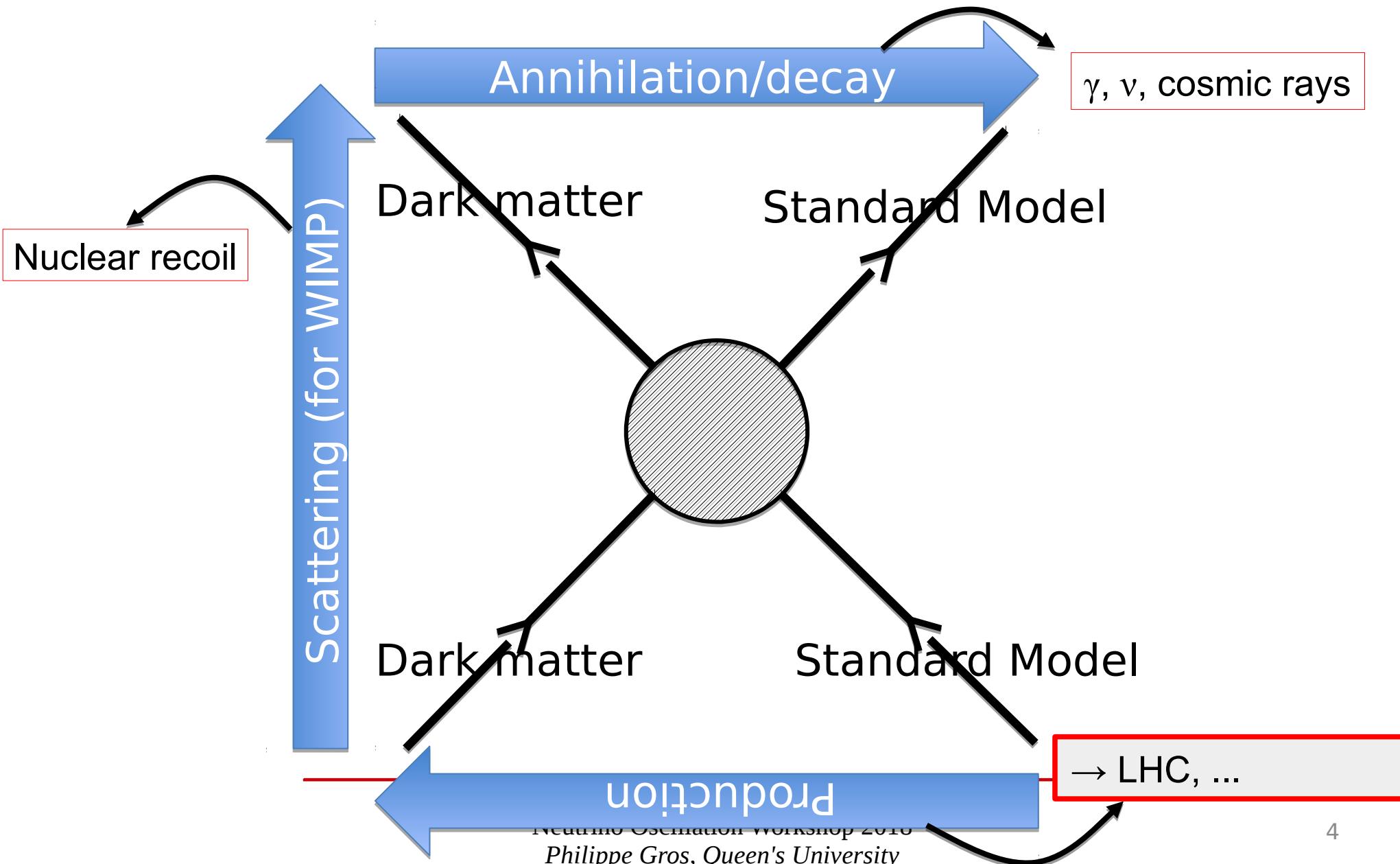


The multiple paths to dark matter



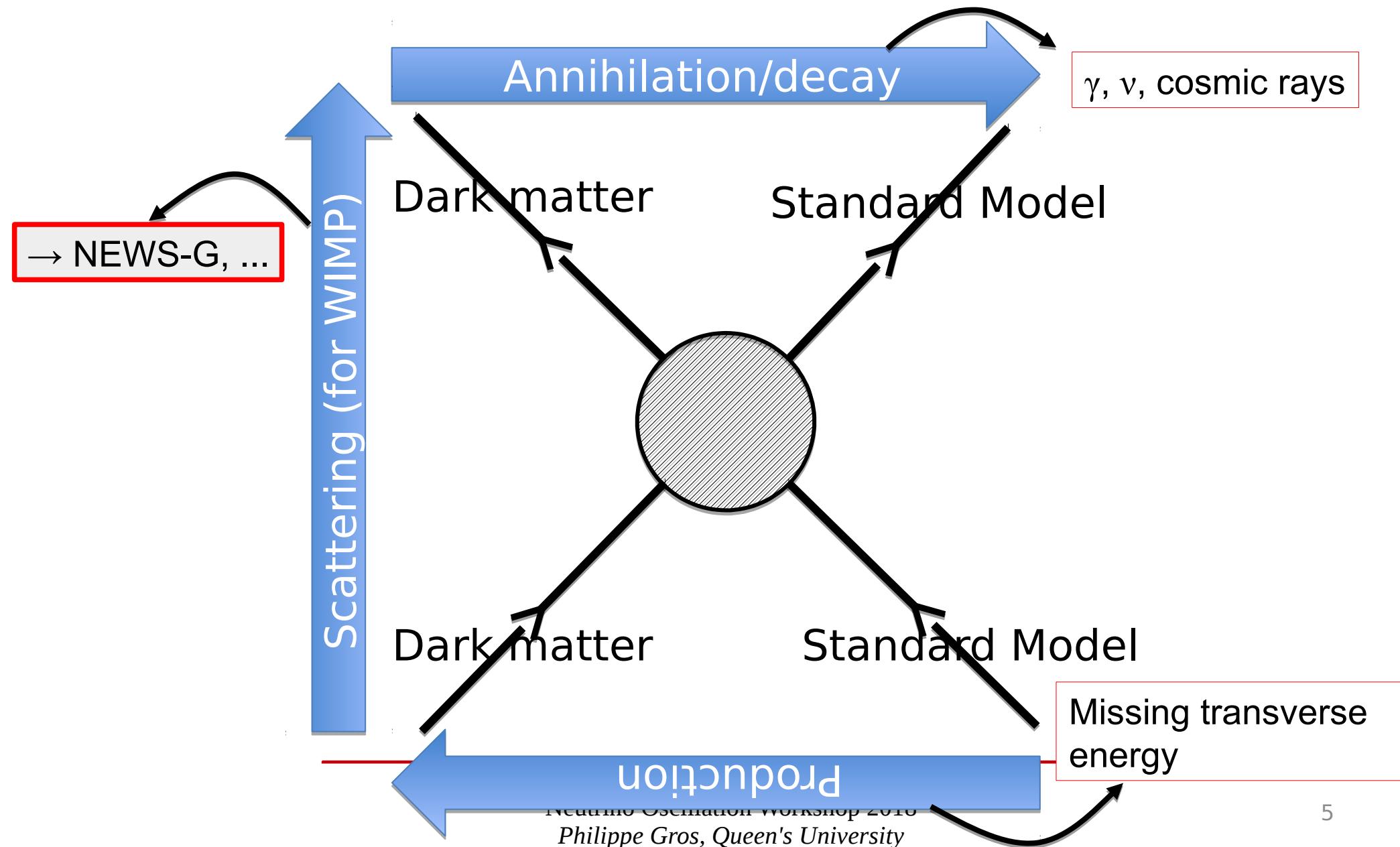


The multiple paths to dark matter





The multiple paths to dark matter

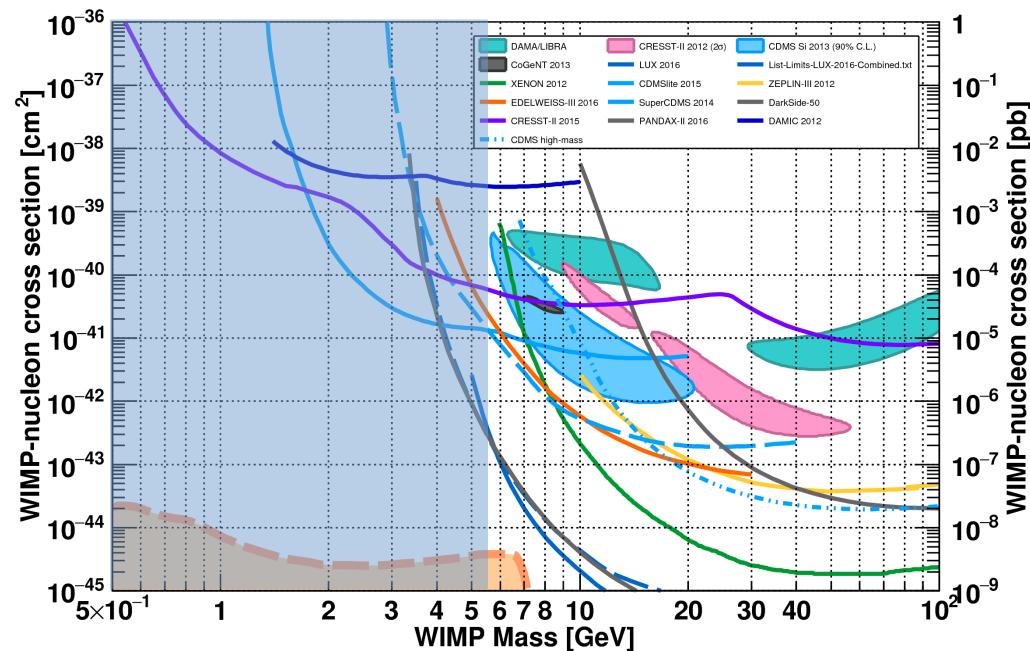




Motivation to search for low mass WIMPs



- Current Upper limits <5GeV orders of magnitude weaker than at higher mass
 - Wide region of the parameter space ($\sigma_{\text{SI}}, m_\chi$) can be probed without ton-scale absorbers
 - current main limitation = energy detection thresholds
- New theoretical approaches
 - e.g. asymmetric dark matter, dark sector...
 - favorite SUSY candidates excluded by LHC



Direct detection of light WIMPs with NEWSG

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How to do it



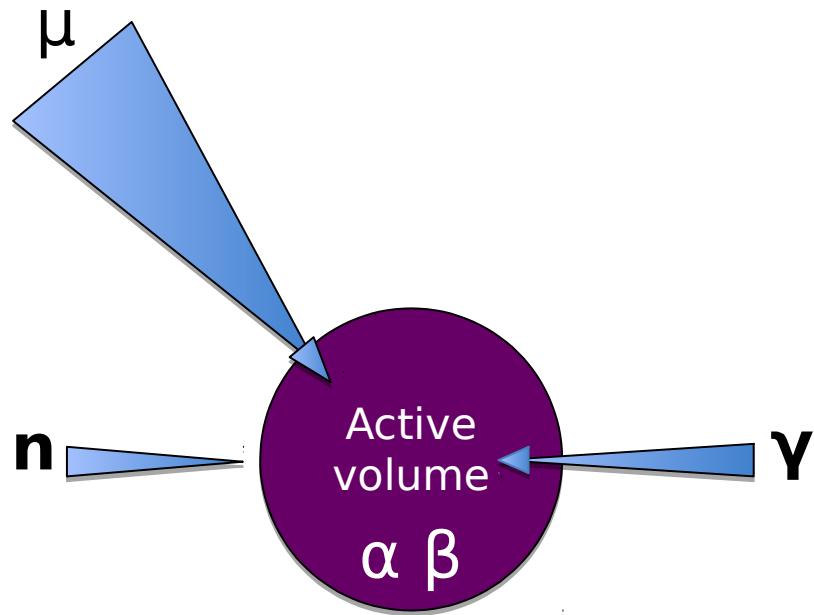
- High sensitivity/Low threshold detector
 - low noise
 - light target material
- Low background
 - cosmic ray → underground
 - local radioactivity → material selection/purification
- (Large exposure)
 - not yet limiting for low mass



Low Background



Multiple background suppression strategies

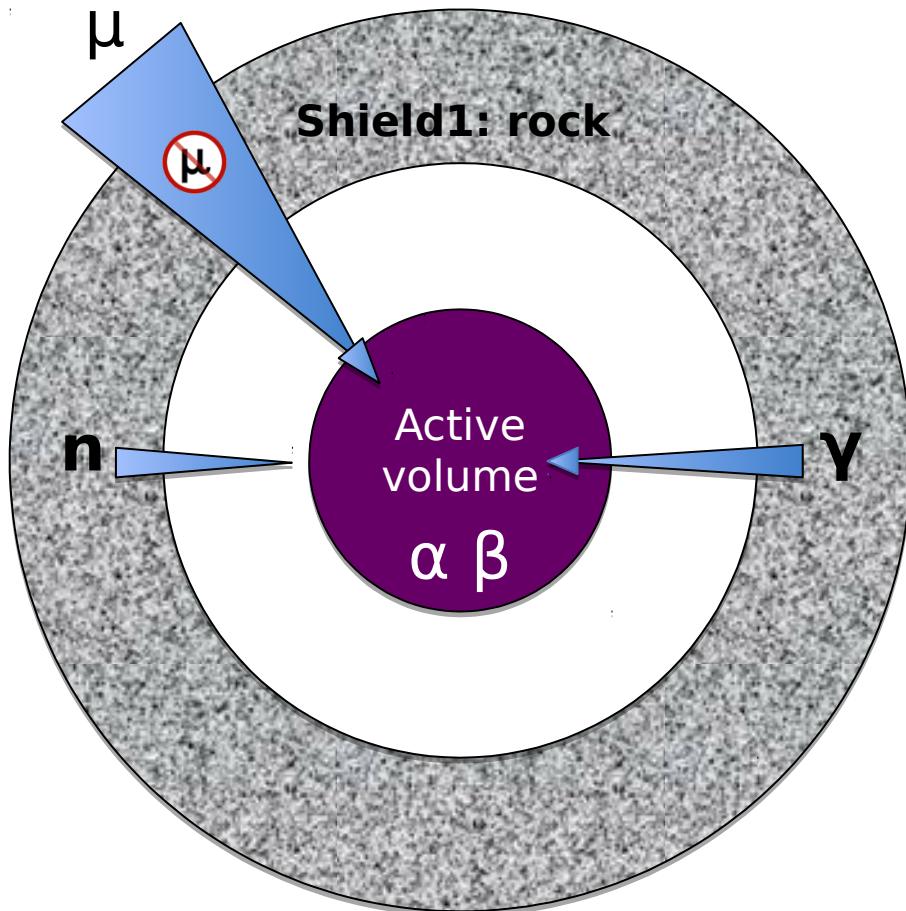




Multiple background suppression strategies



Passive shielding

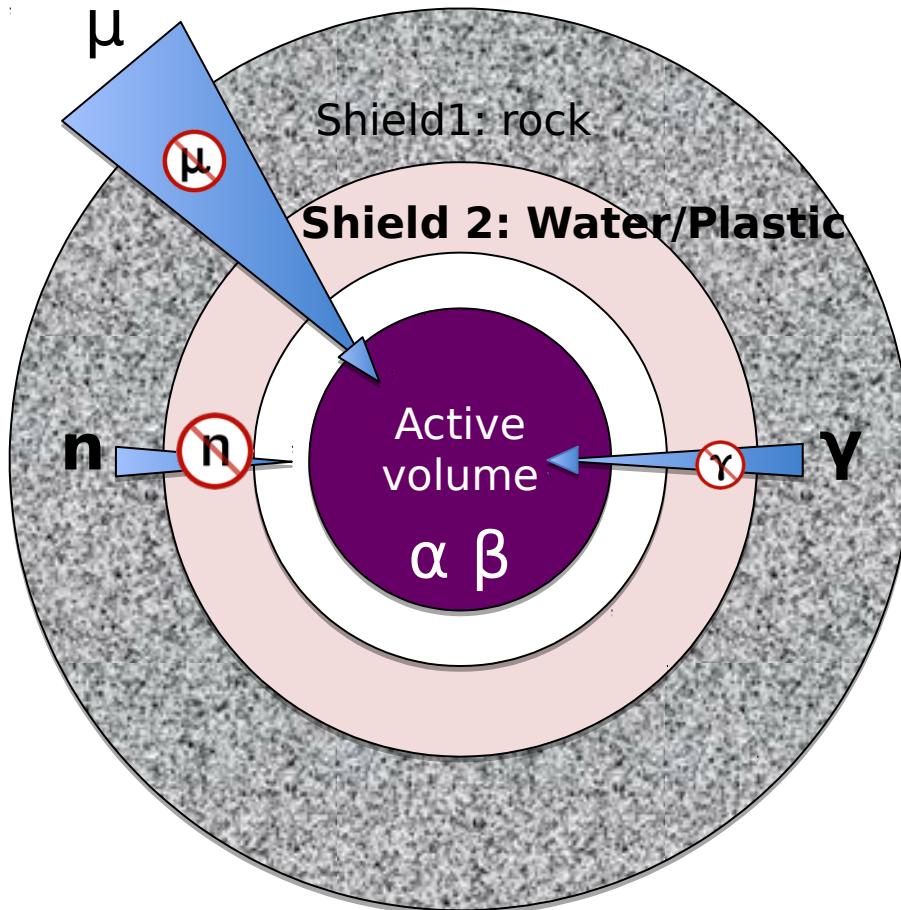




Multiple background suppression strategies



Passive shielding

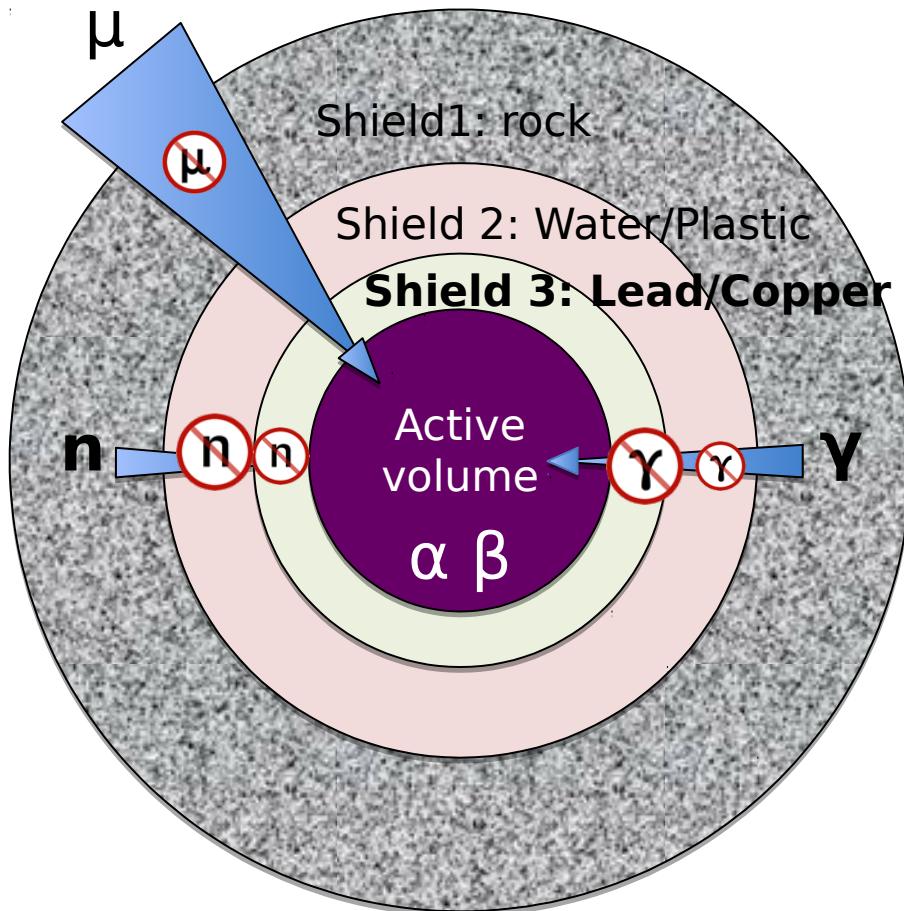




Multiple background suppression strategies



Passive shielding

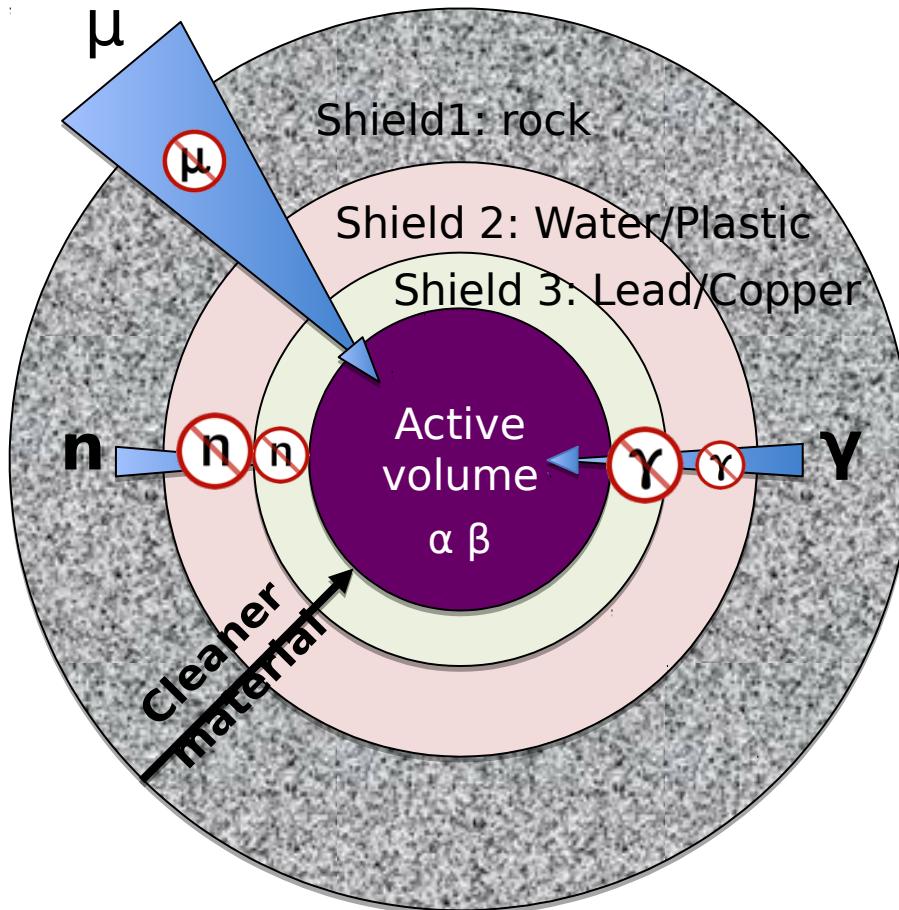




Multiple background suppression strategies



Passive shielding

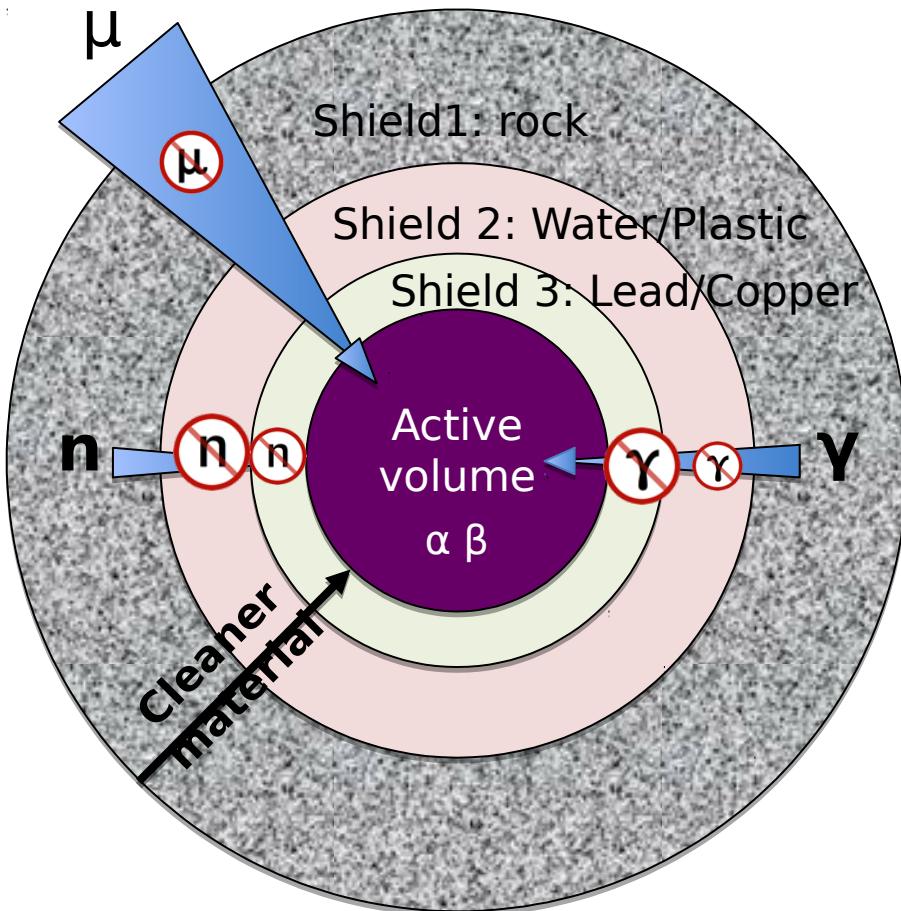




Multiple background suppression strategies

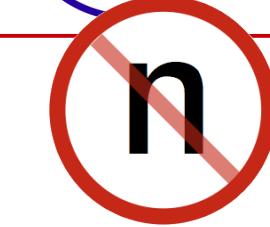


Passive shielding



Discrimination

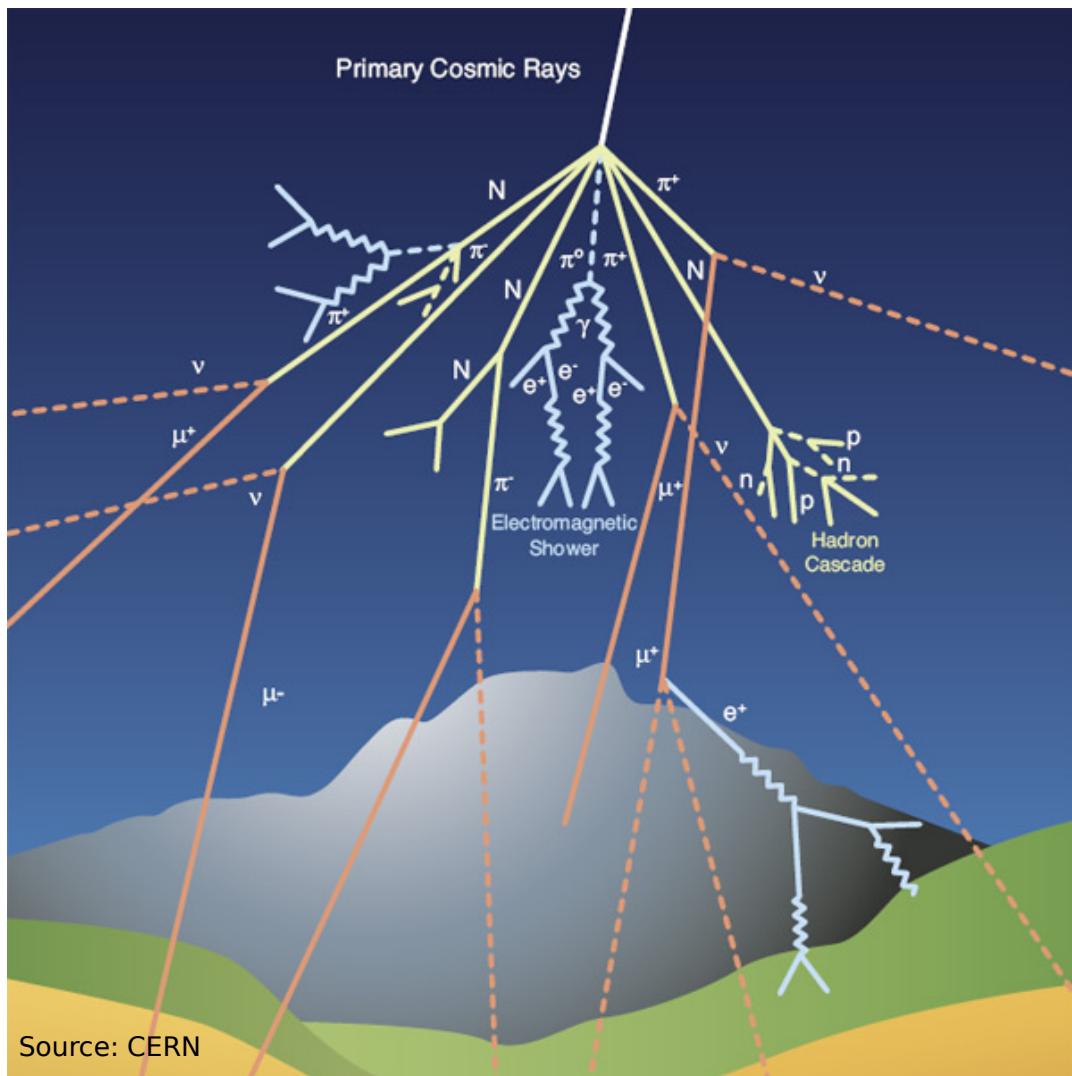
- Active veto
- Energy
- Position
- Pulse shape
- Multi signal
- ...



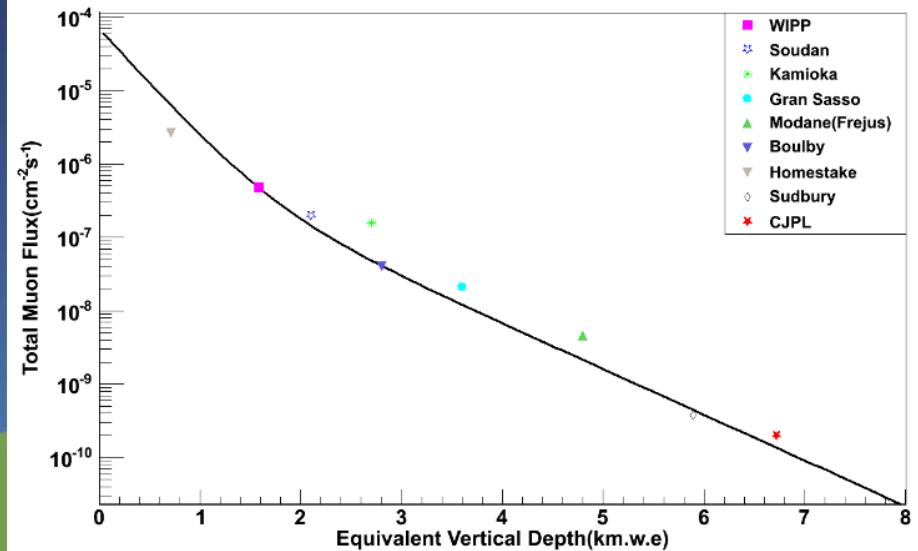
The cosmic rays problem



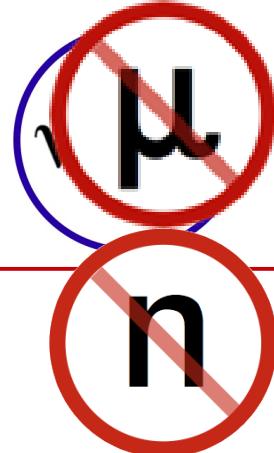
Queen's
UNIVERSITY



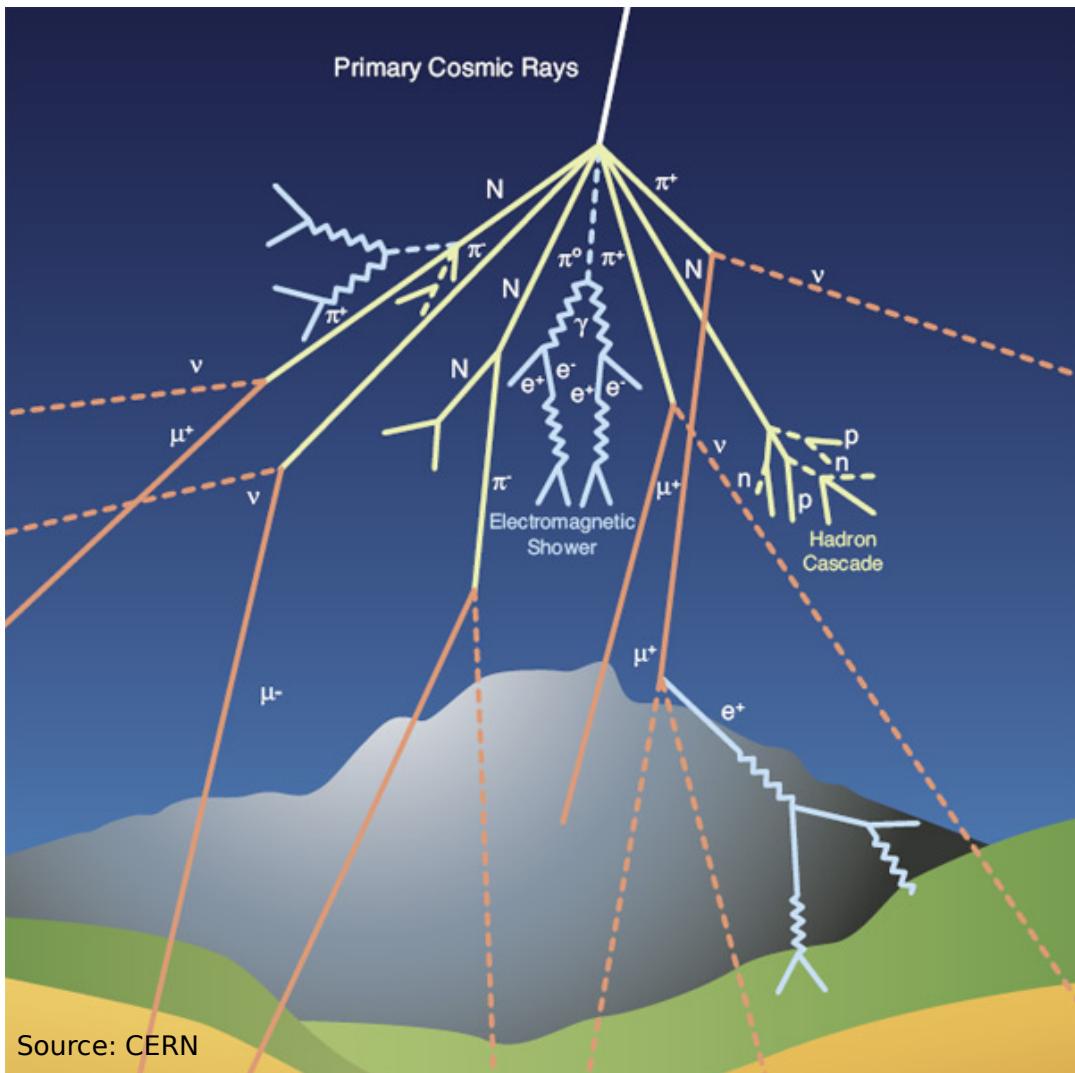
- High energy muons travel deep
- Production of unstable isotopes
- Muon-induced spallation=> neutrons



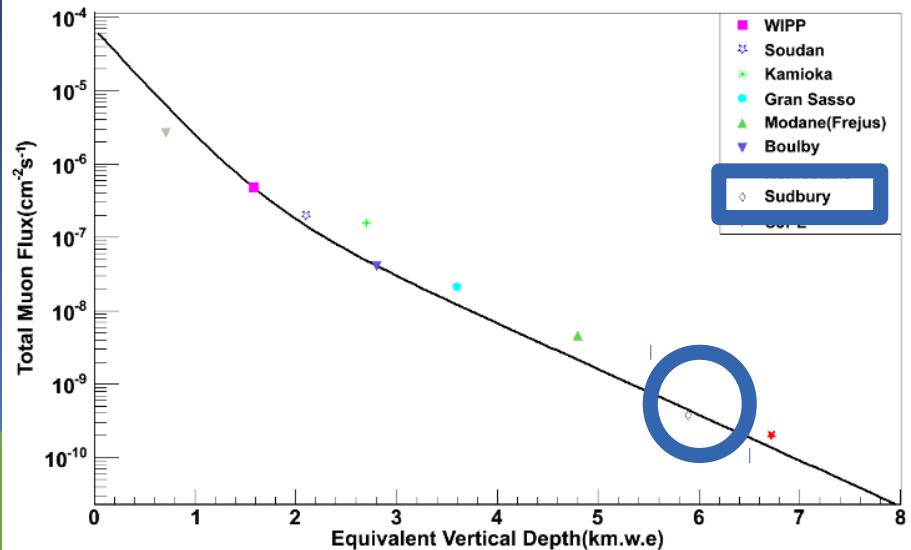
Measurement of Cosmic Ray Flux in China JinPing underground Laboratory
Arxiv:1305.0899



The cosmic rays problem



- High energy muons travel deep
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Measurement of Cosmic Ray Flux in China JinPing underground Laboratory
Arxiv:1305.0899

SNOLAB Surface facility

- Offices
- Clean laboratory
- Warehouse
- Machine shop



Shaft headframe



Greater Sudbury (Ontario)



6800 ft
2000 m
6060 m.w.e

SNOLAB Underground facility



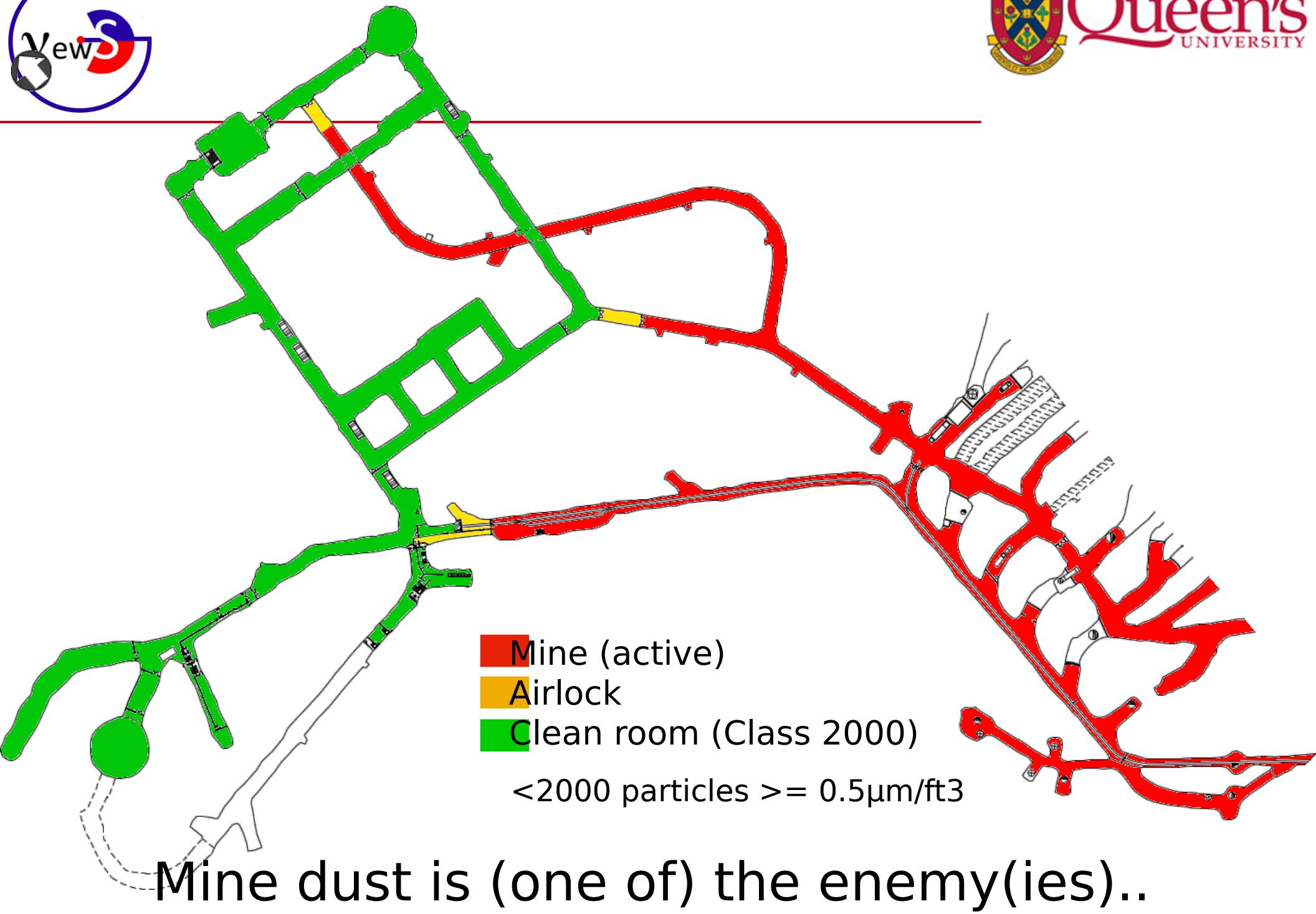
0.27 muon/m²/day



Lab entrance: boot wash

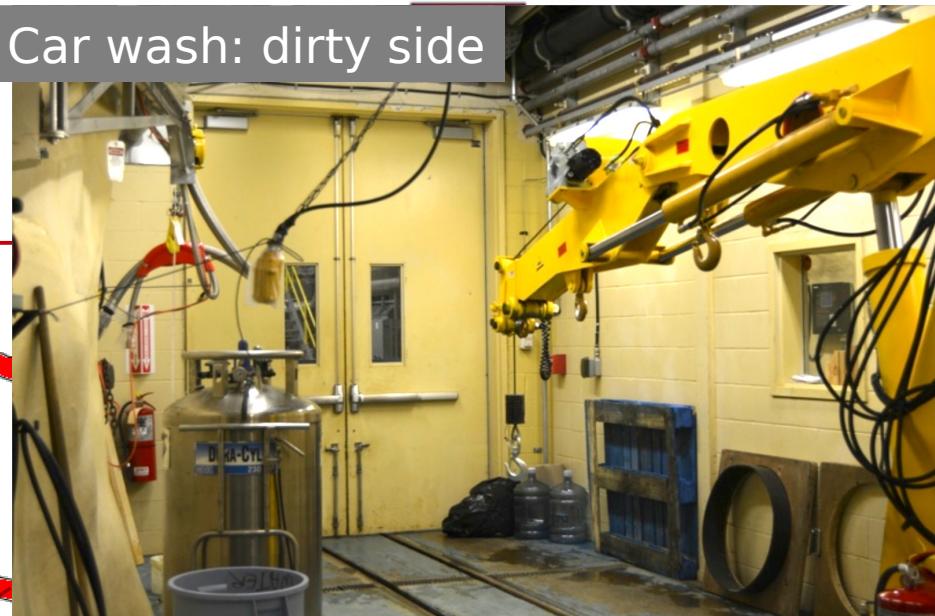


Credit: Inconsiderate Hat blog





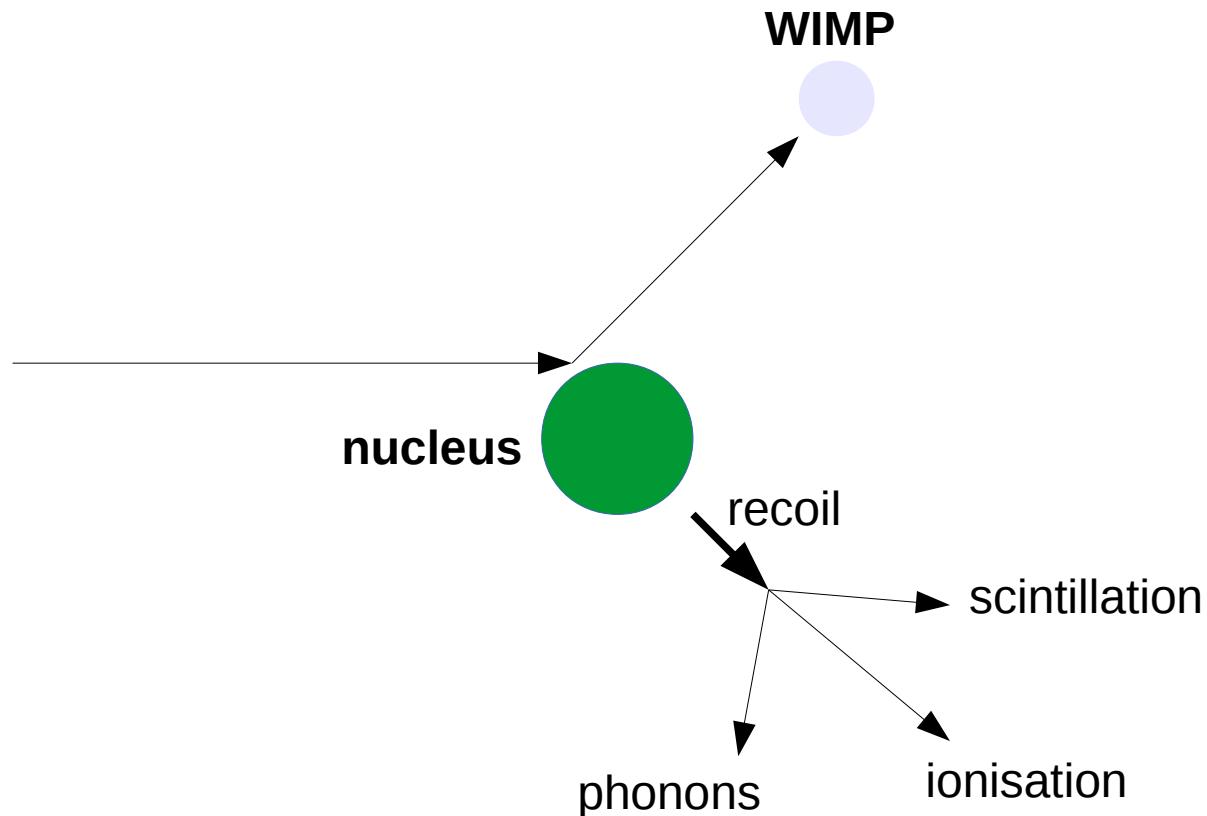
Car wash: dirty side





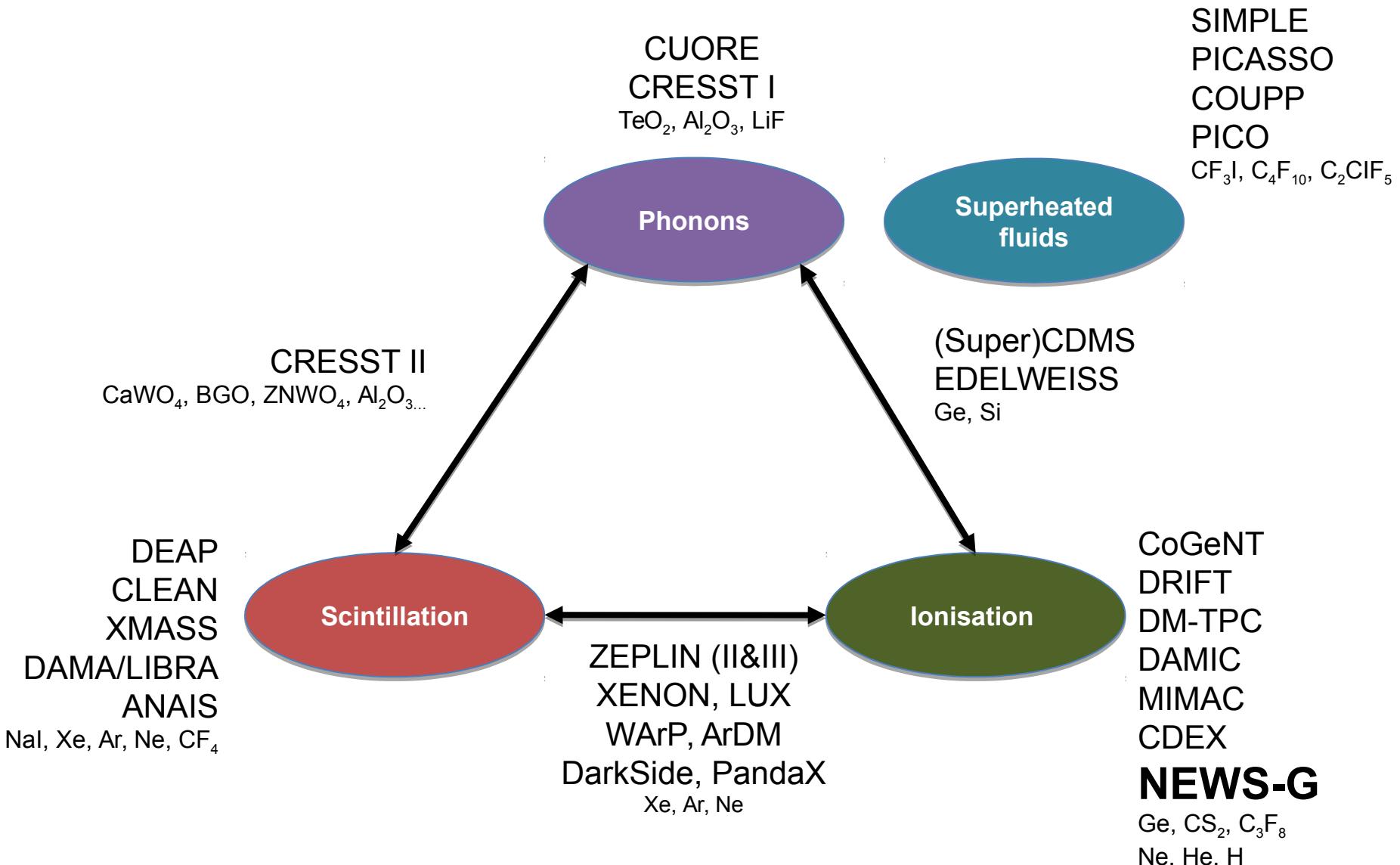
A low threshold detector: Spherical Proportional Counter (SPC)

Detection principle



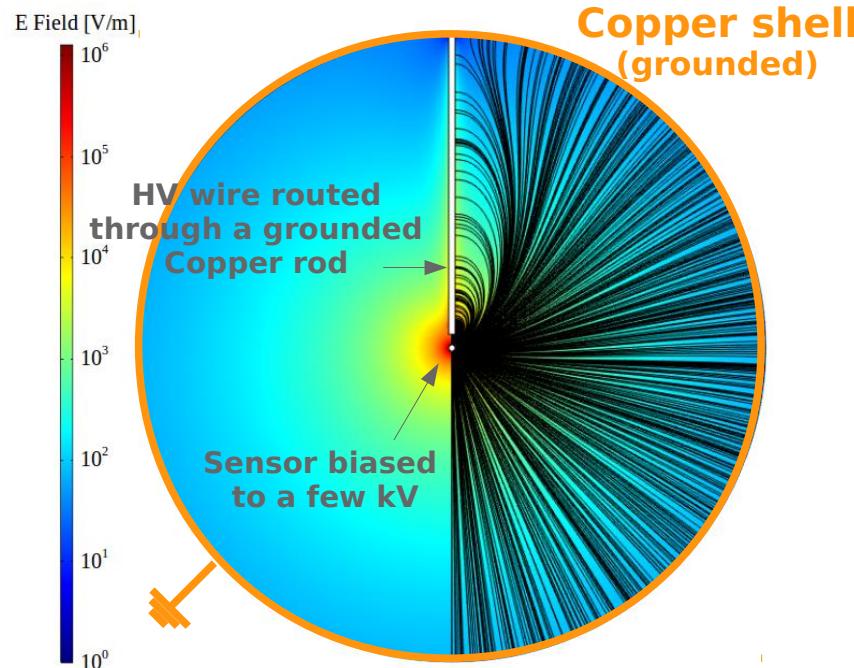


Direct detection techniques

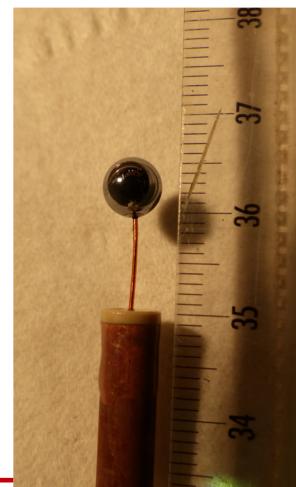




SPC



Vessel
60 cm Ø NOSV Copper

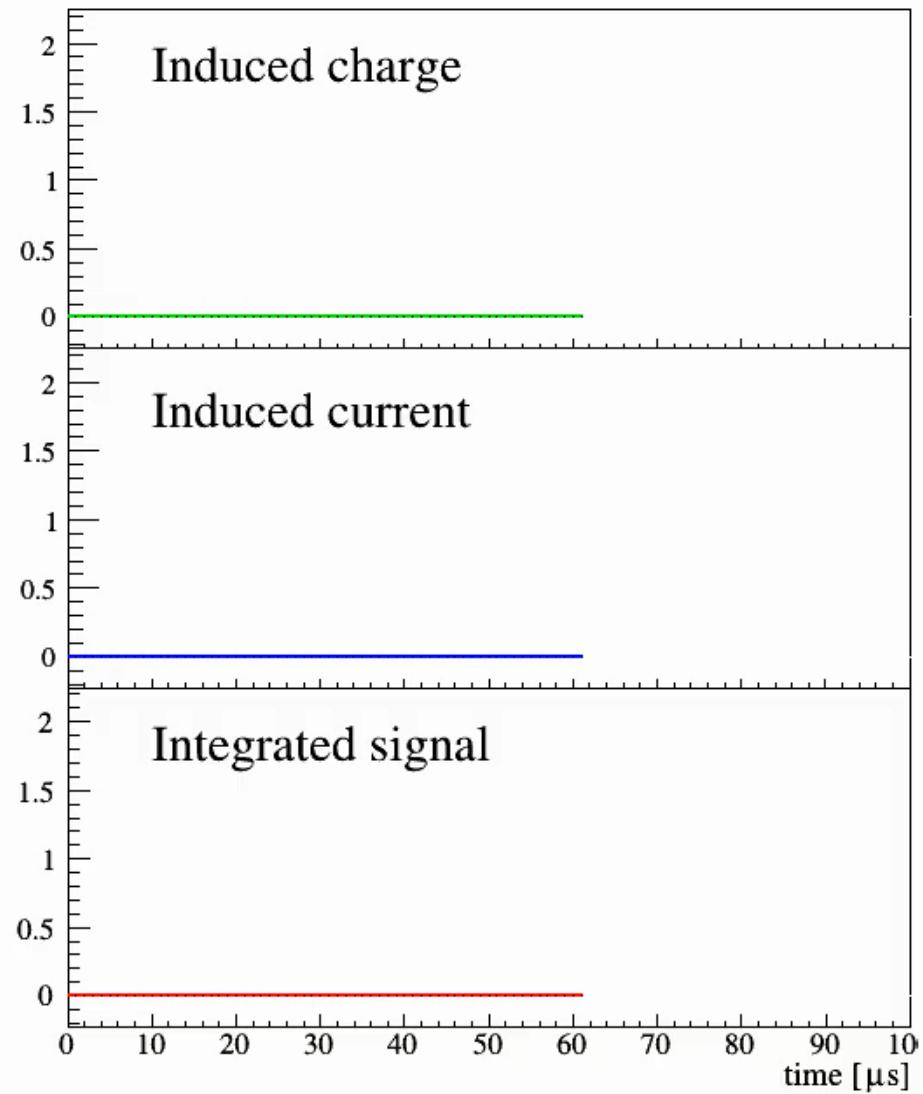
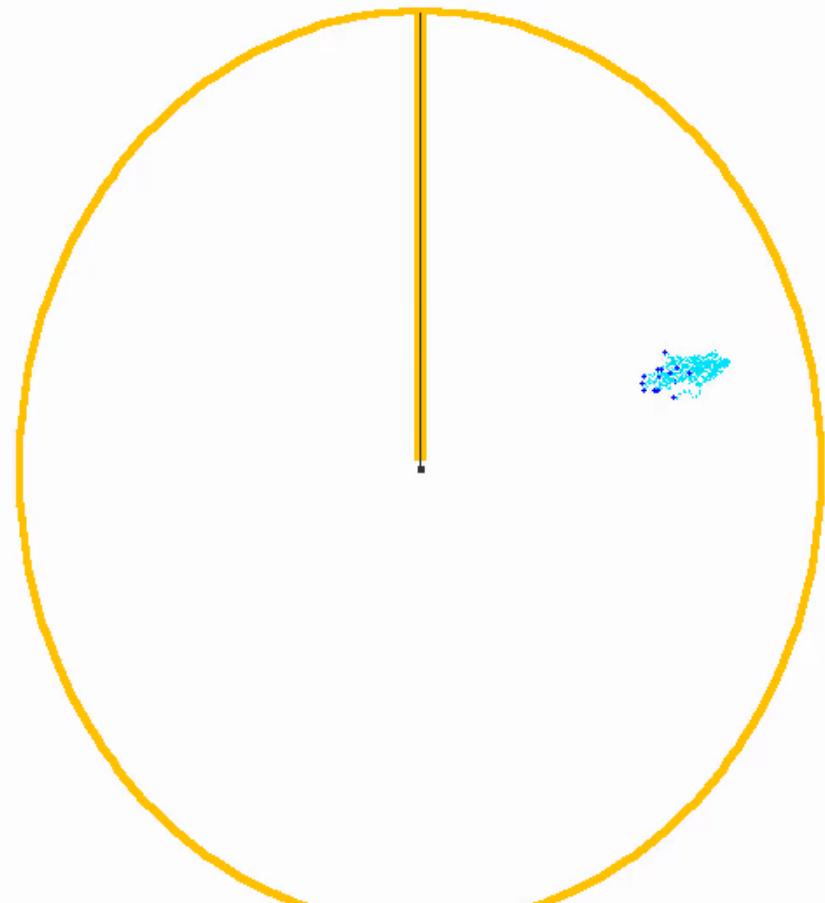


Sensor
6.3 mm Ø

- Gaseous proportional counter
 - Copper sphere
 - Central ball with HV
- Very low capacitance
 - $<1\text{ pF}$
- Drifting electrons from ionisation
- Light gas better for low mass WIMPs
 - elastic scattering energy transfer



SPC signal

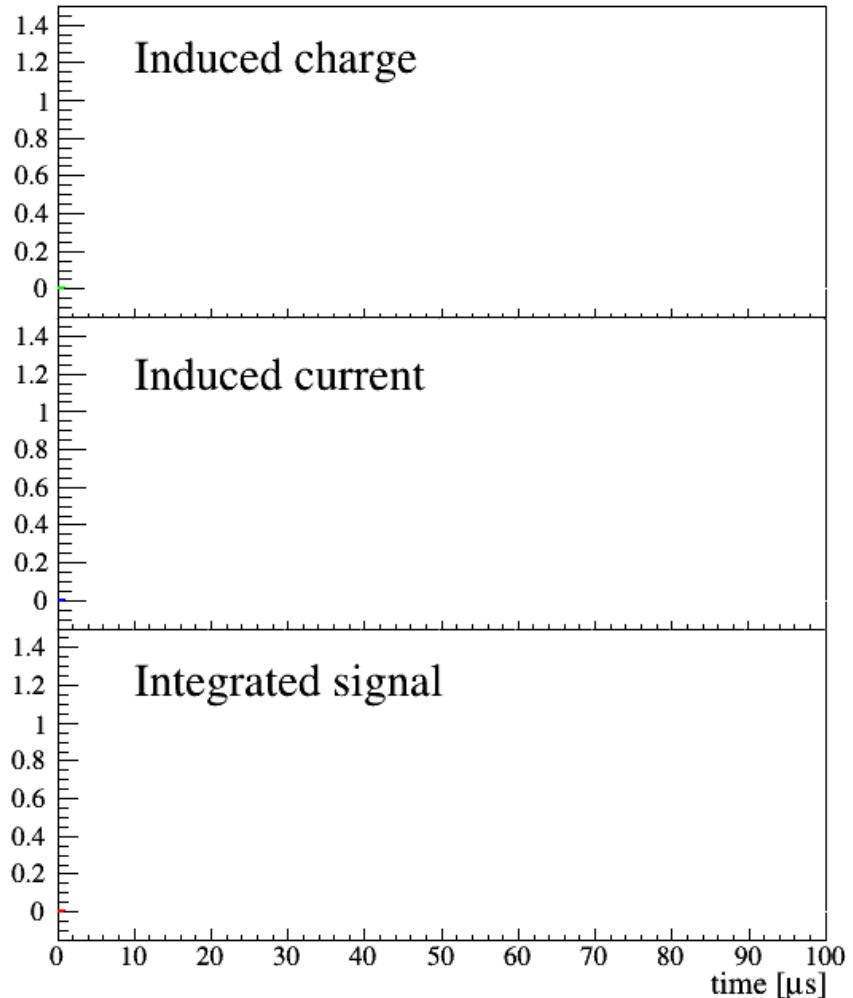
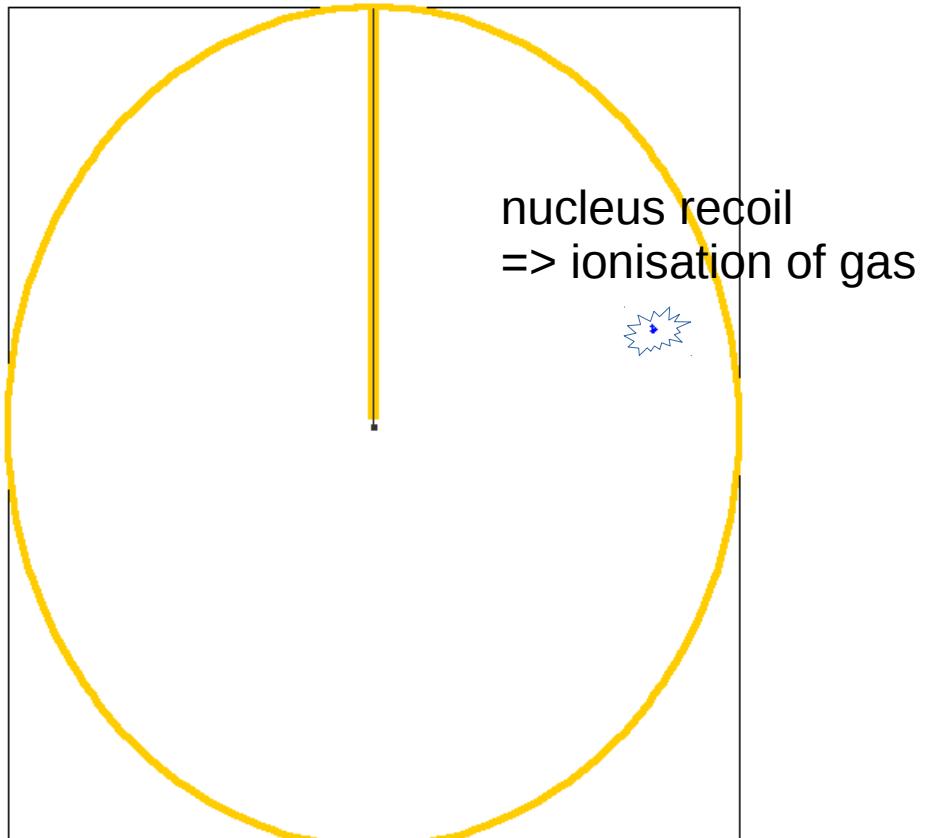


Direct detection of light WIMPs with NEWSG

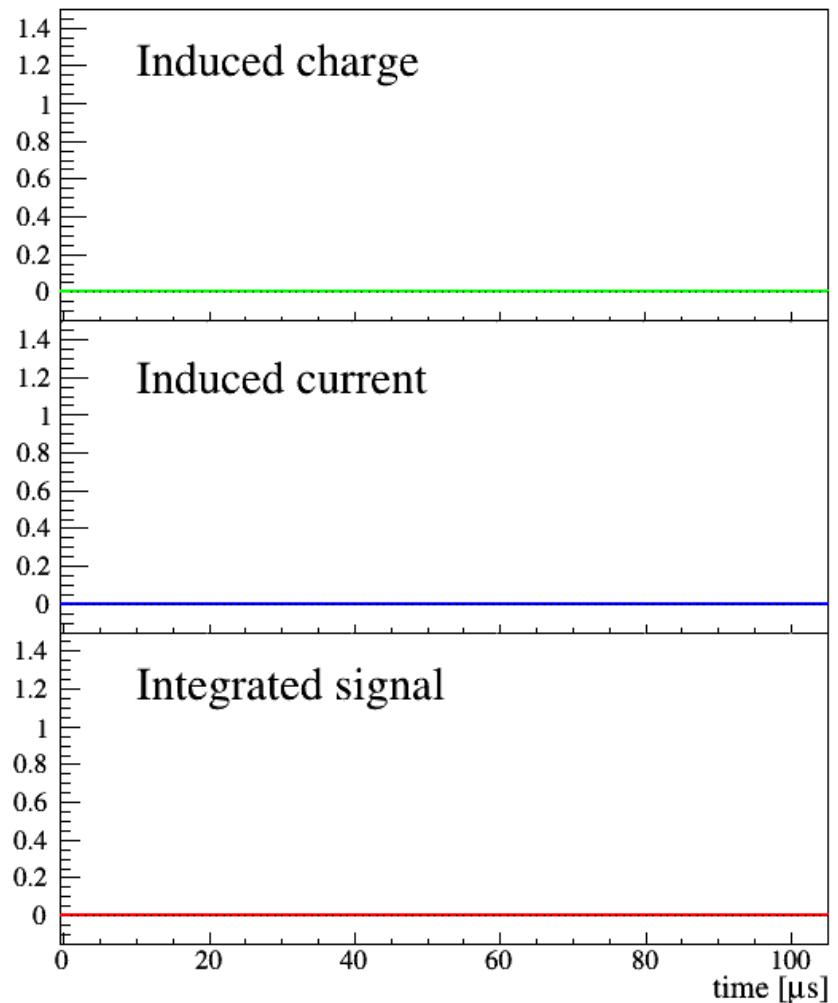
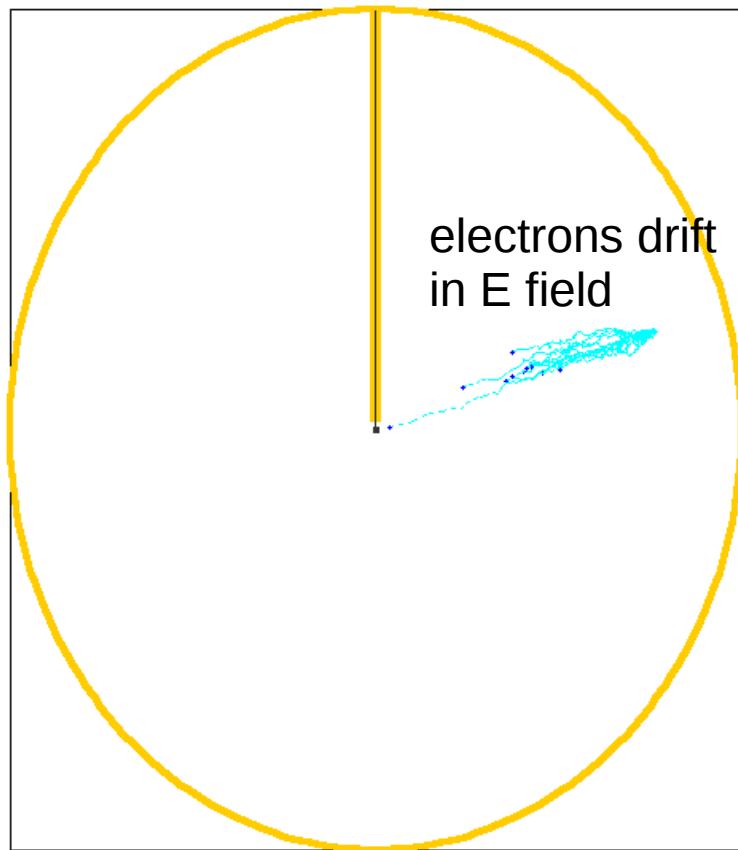
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SPC signal



SPC signal

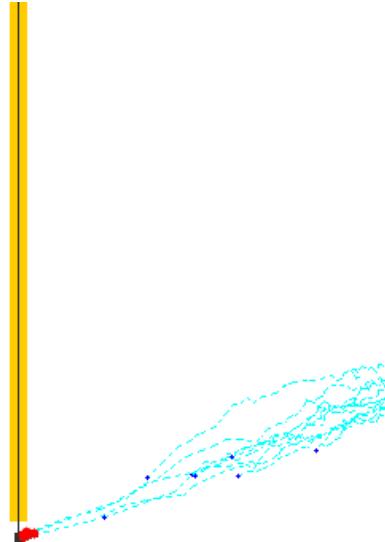


Direct detection of light WIMPs with NEWSG

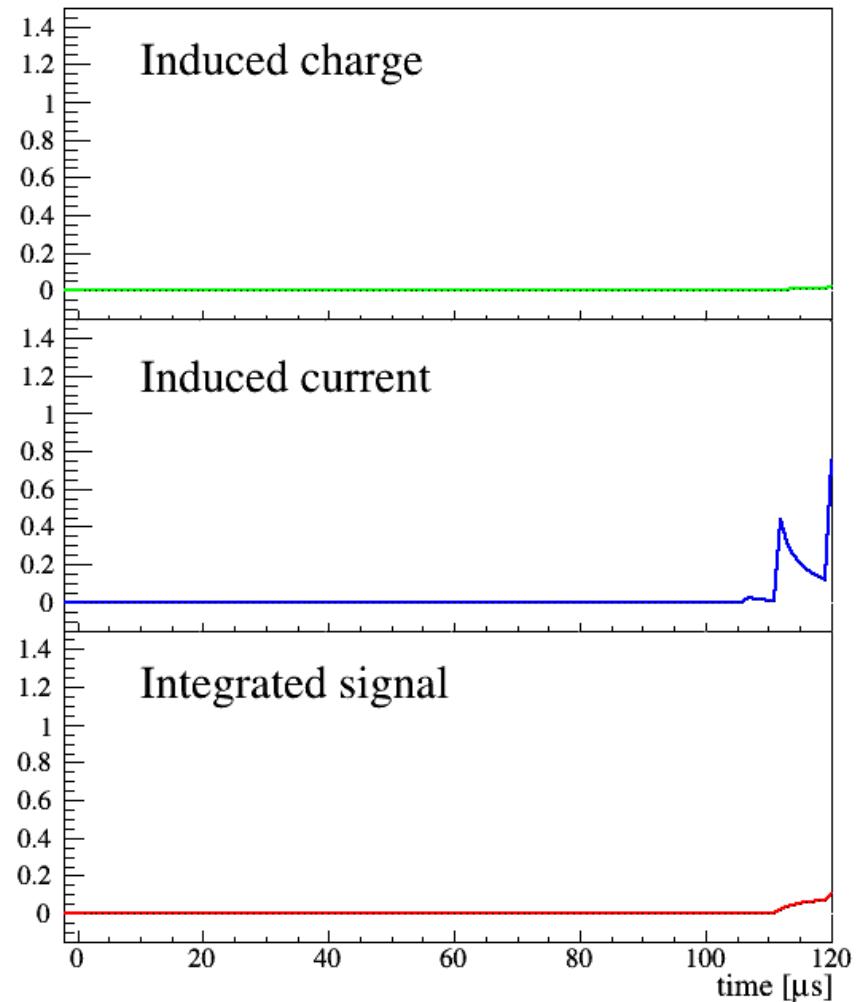
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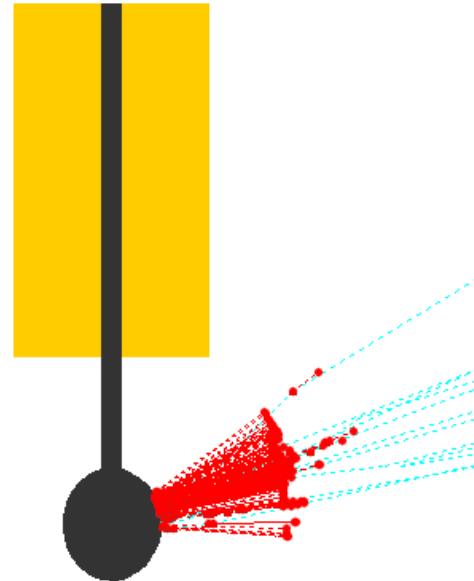
SPC signal



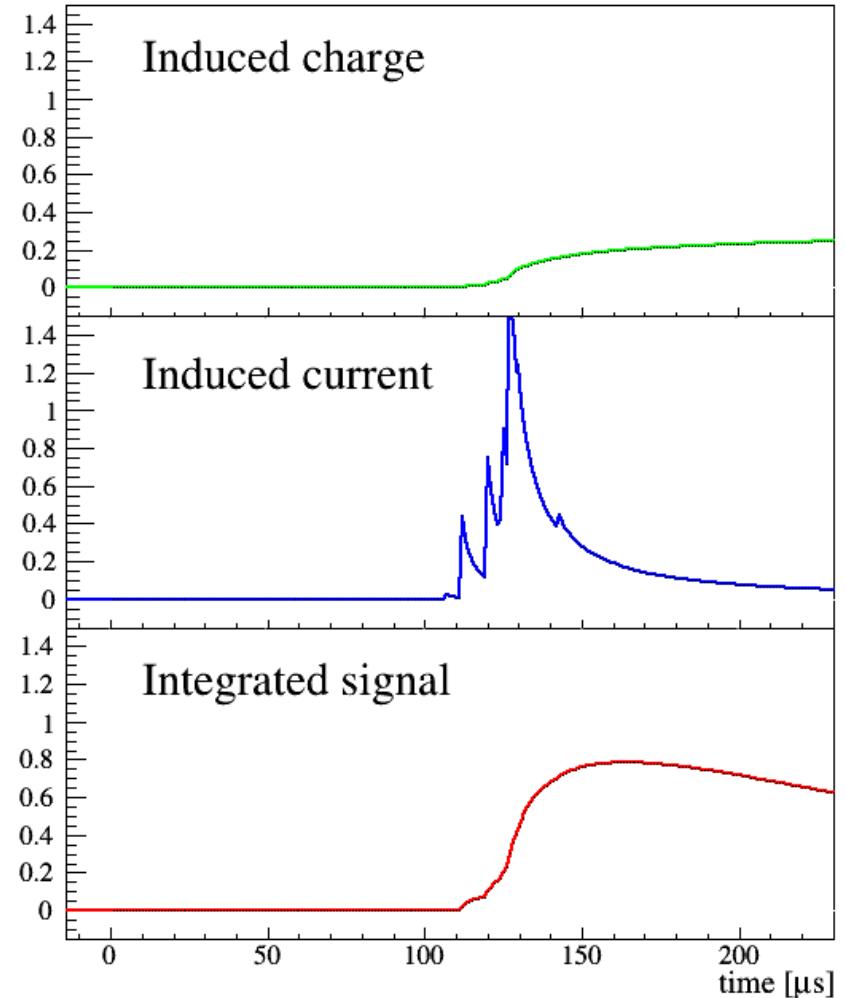
avalanche in high field near sensor



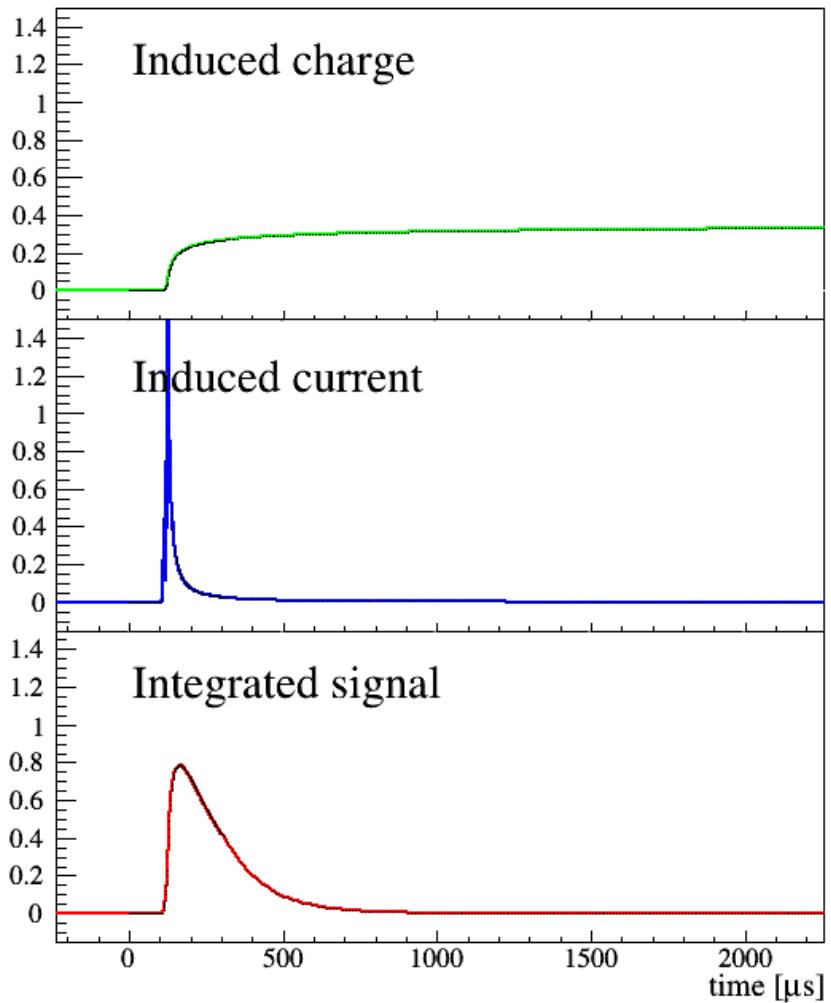
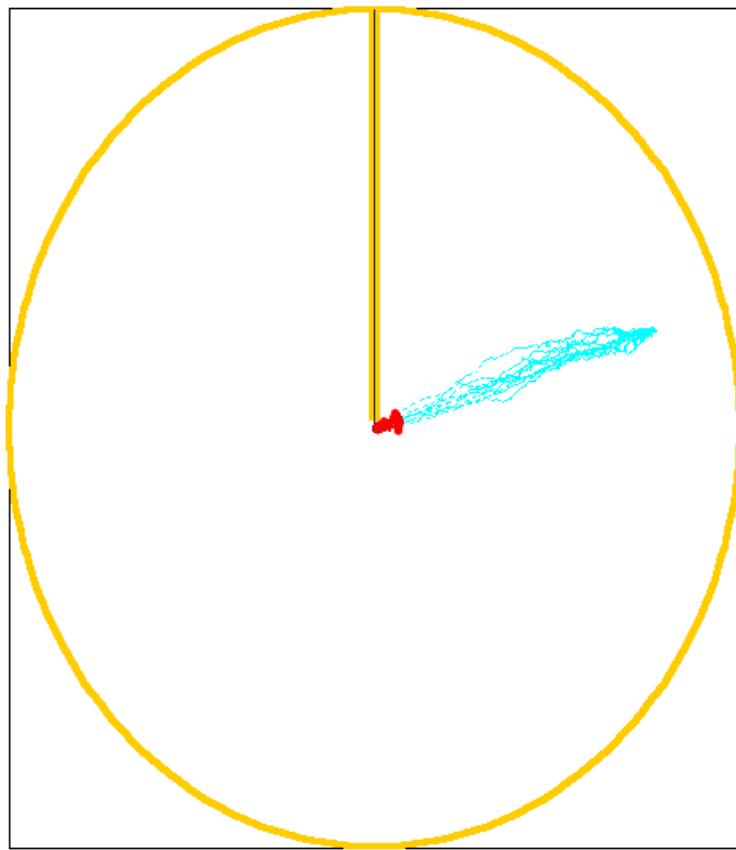
SPC signal



Signal induced by ions drifting back



SPC signal



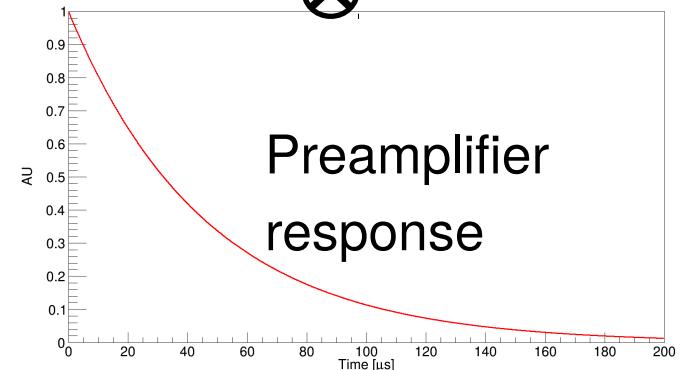
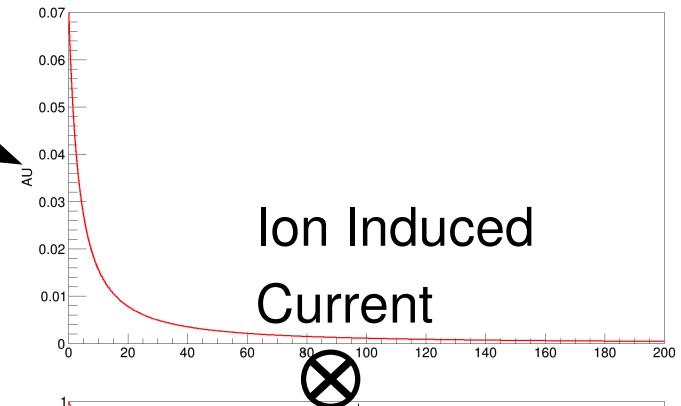
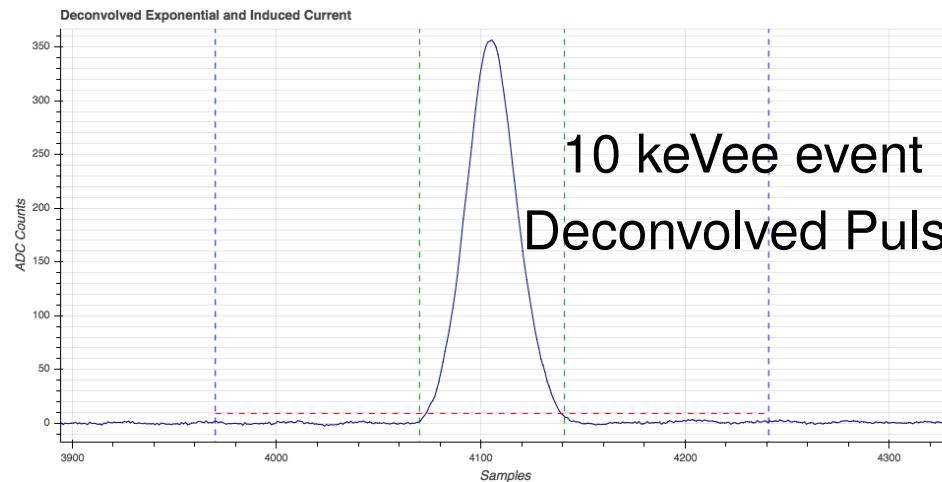
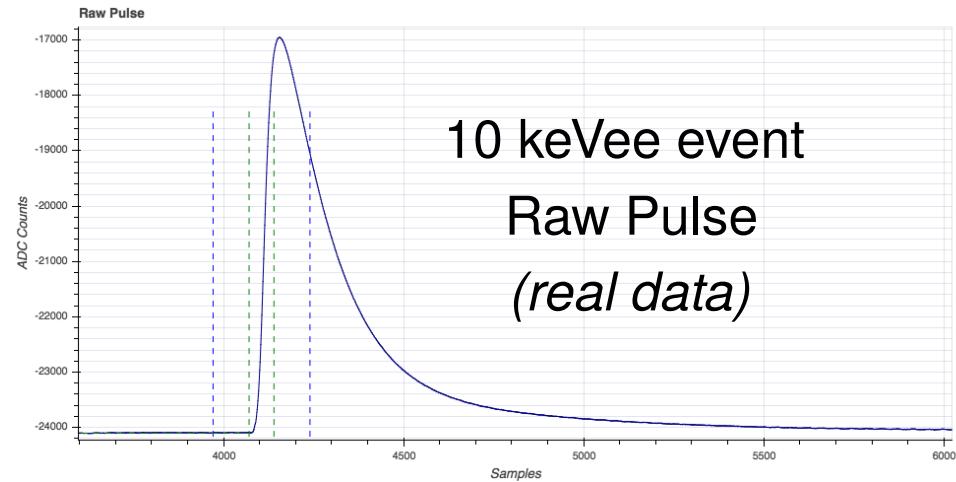
Direct detection of light WIMPs with NEWSG

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Pulse treatment



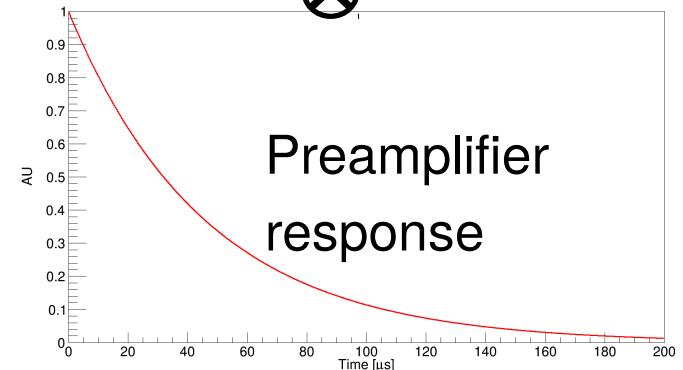
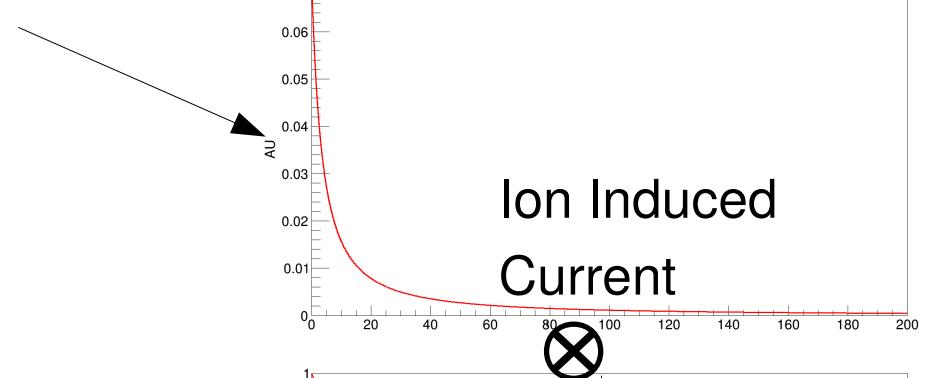
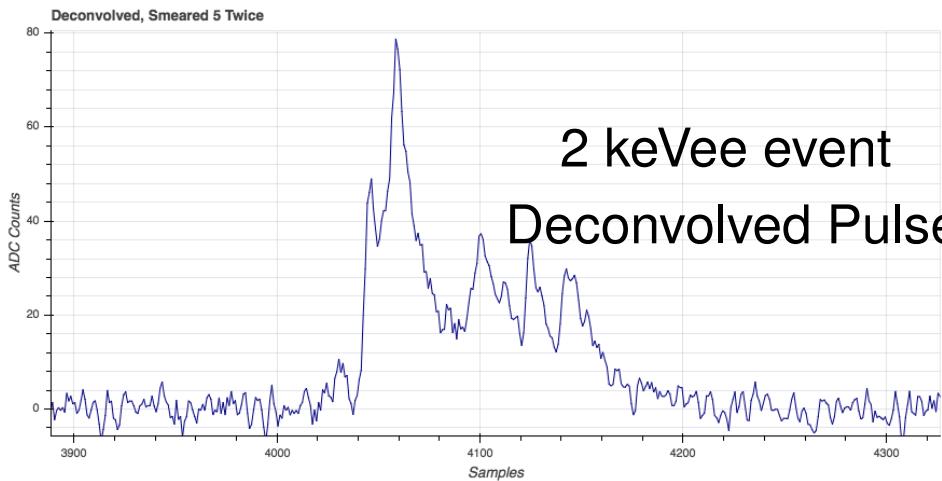
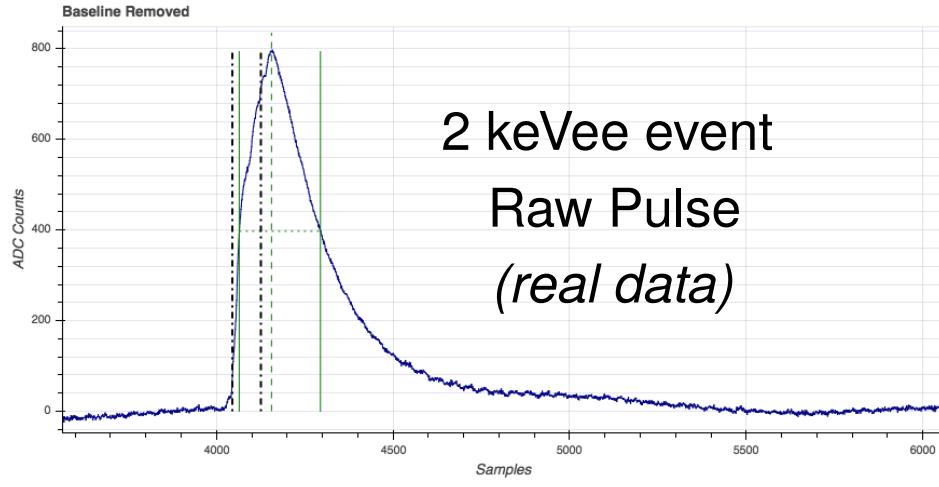
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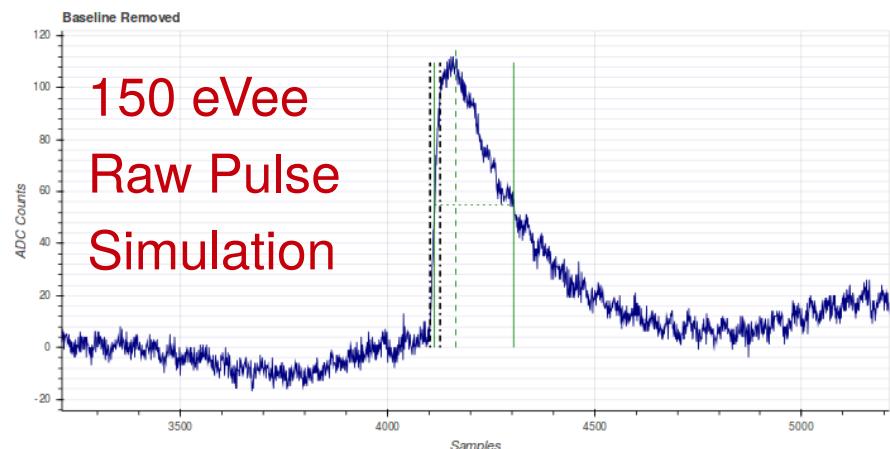
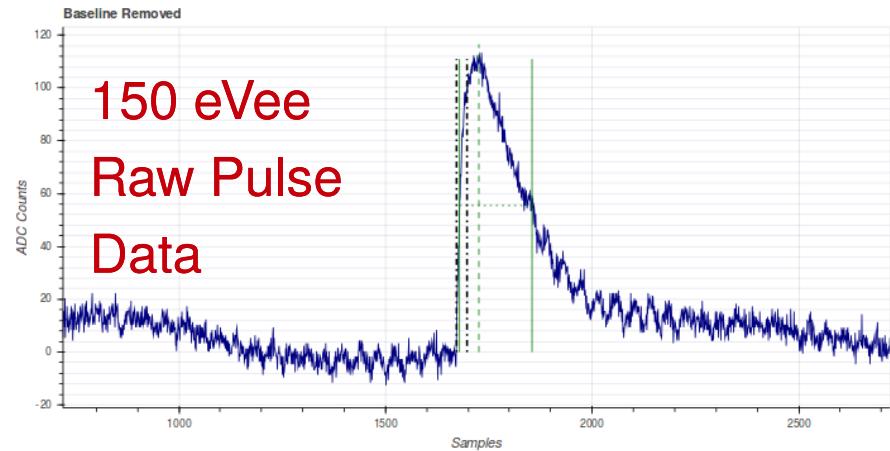
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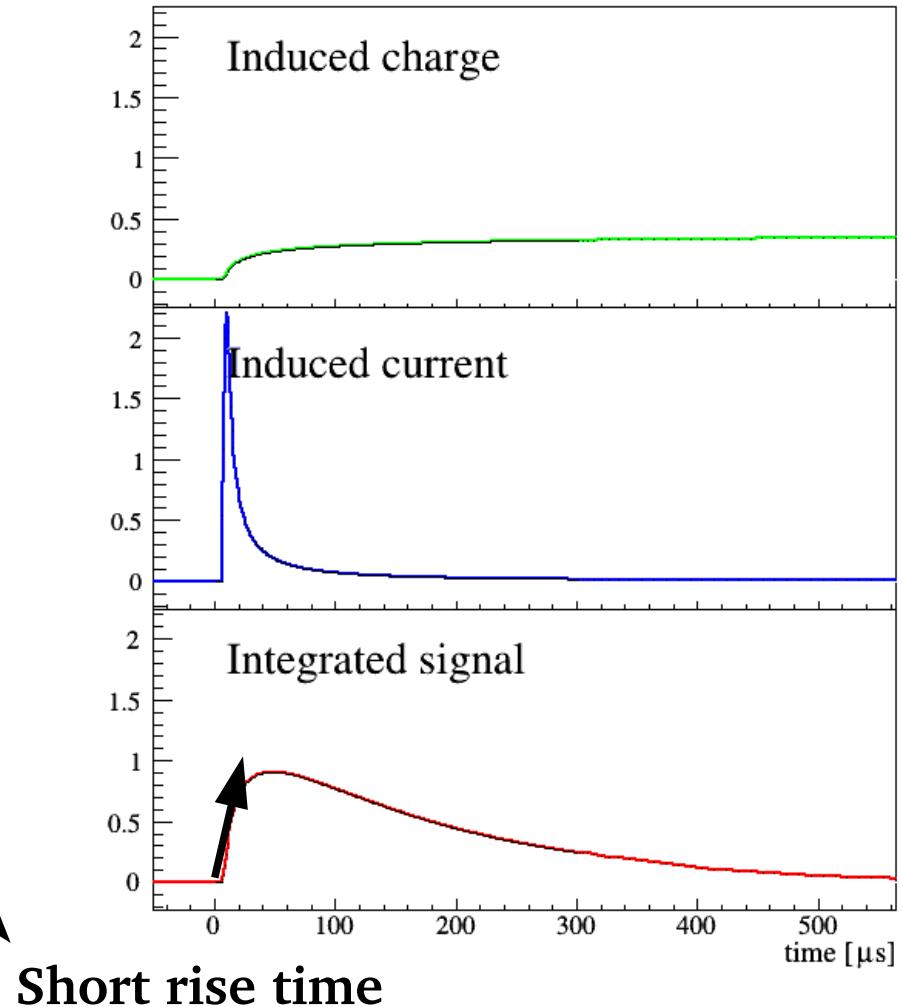
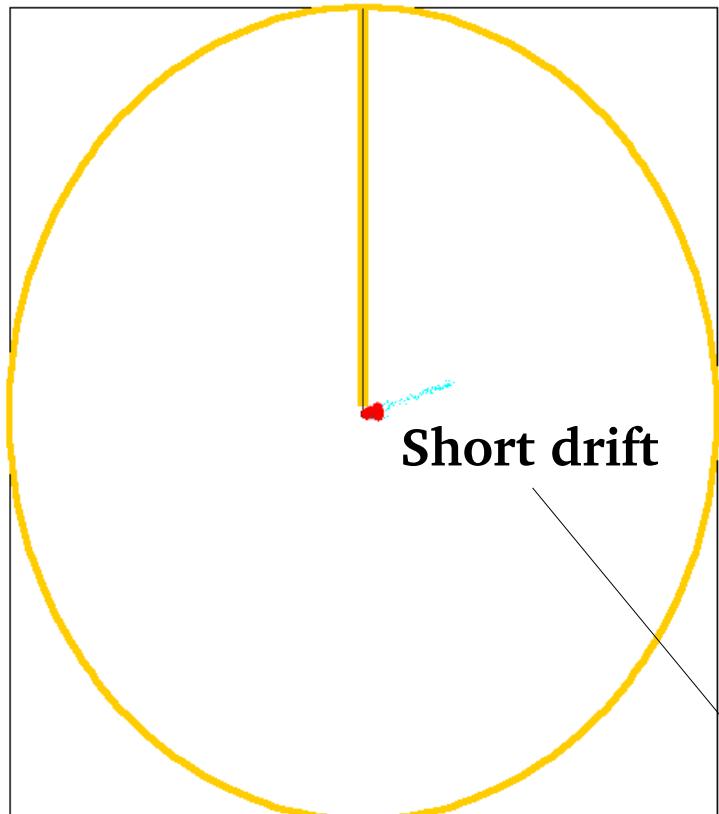
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Simulation

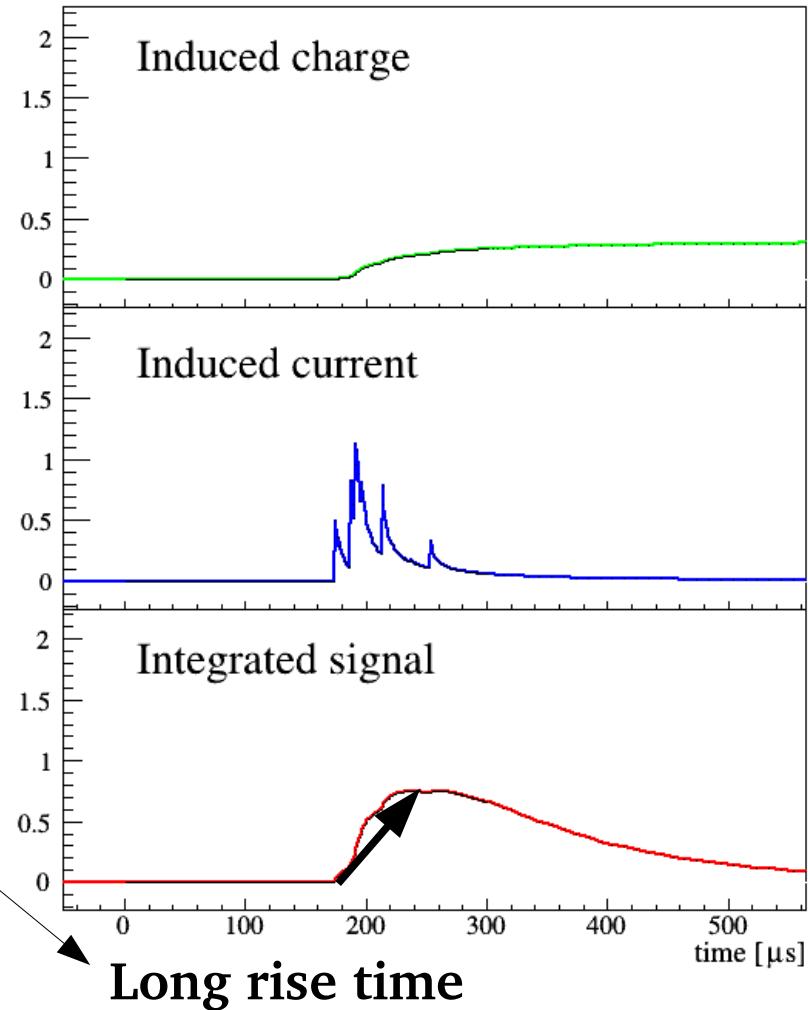
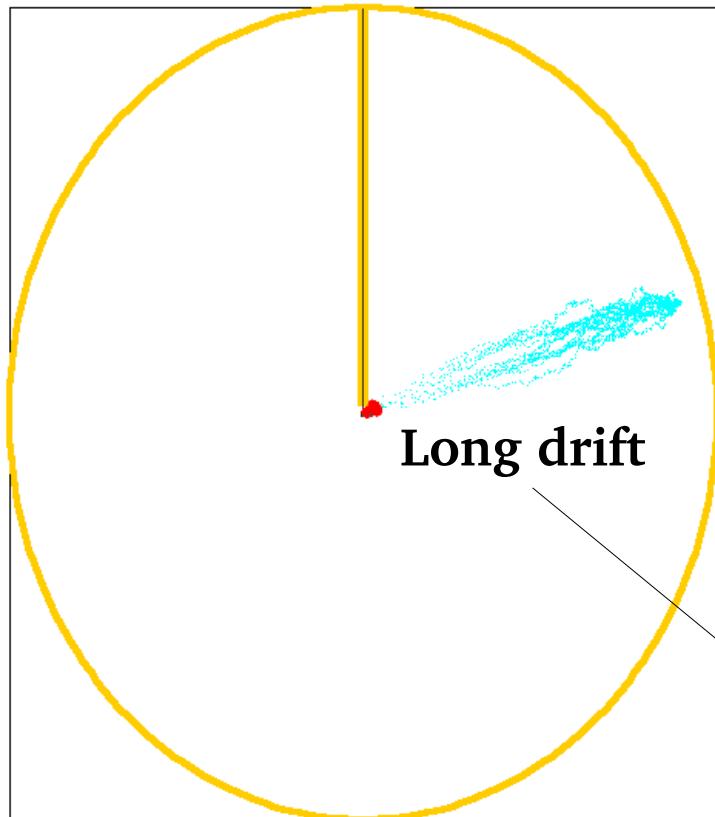
- Drift of individual electrons
 - COMSOL, Magboltz
- Quenching:
 - SRIM, parametrisation
- Avalanche
 - Polya distribution (Garfield++)
- Simulated amplifier pulses
- Noise from data templates
- Same processing as real pulses



Event discrimination: Rise time

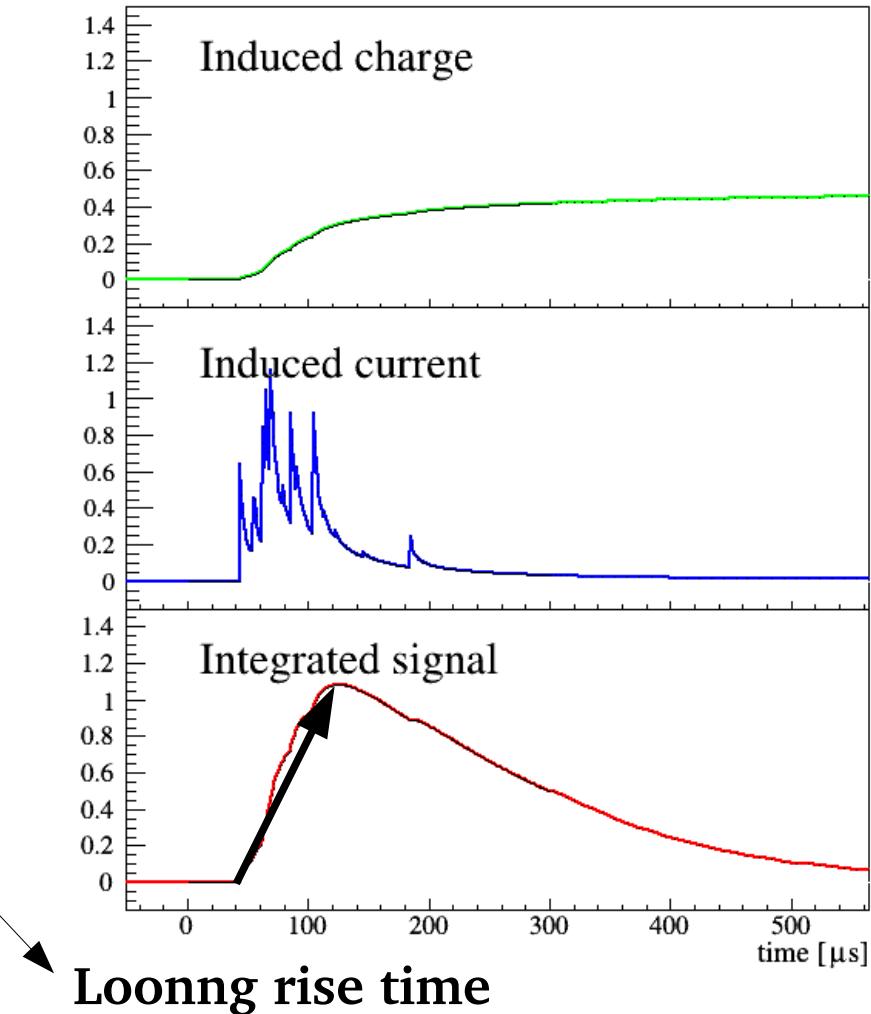
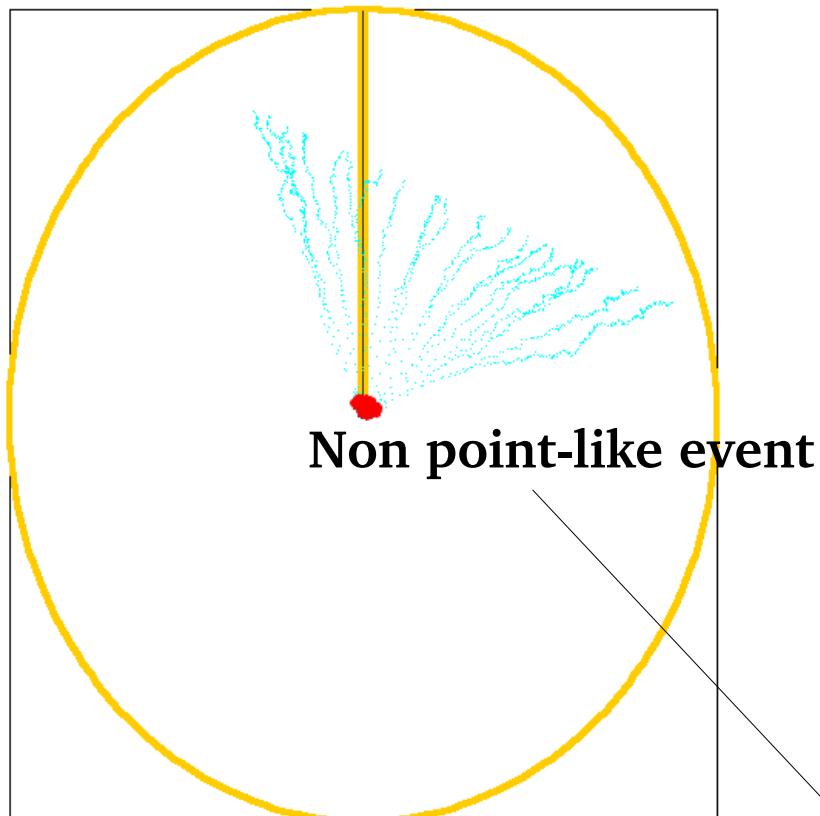


Event discrimination: Rise time





Event discrimination: Rise time



Direct detection of light WIMPs with NEWSG

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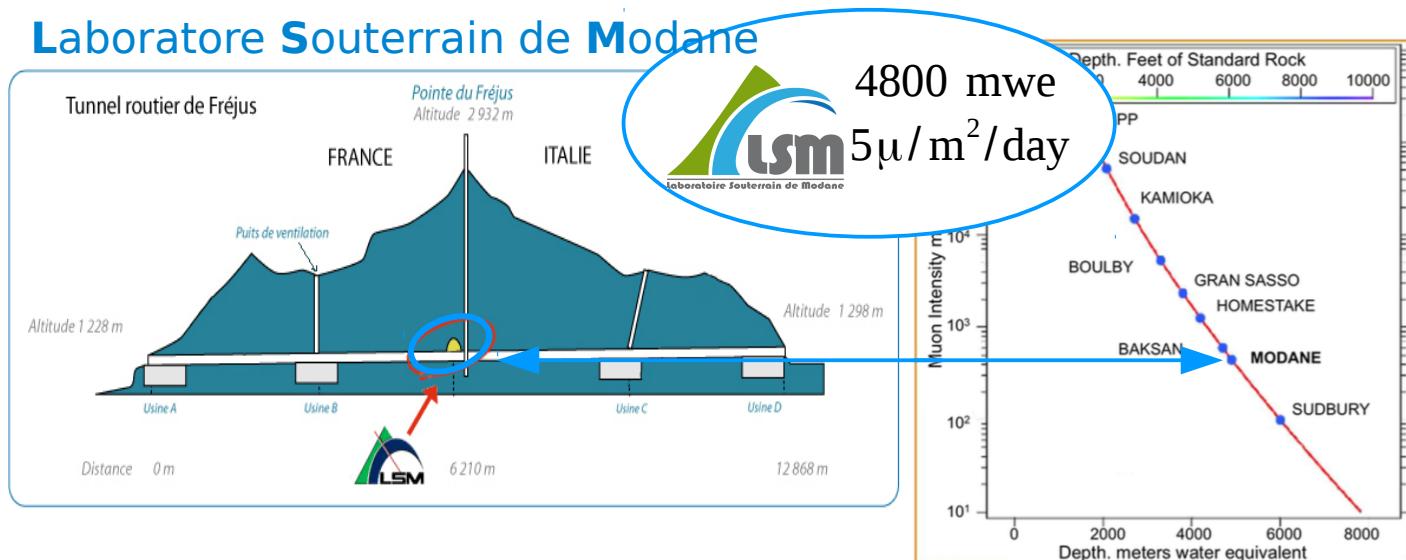
Results with SEDINE (NEWS-G at LSM)



SEDINE at LSM



Laboratoire Souterrain de Modane



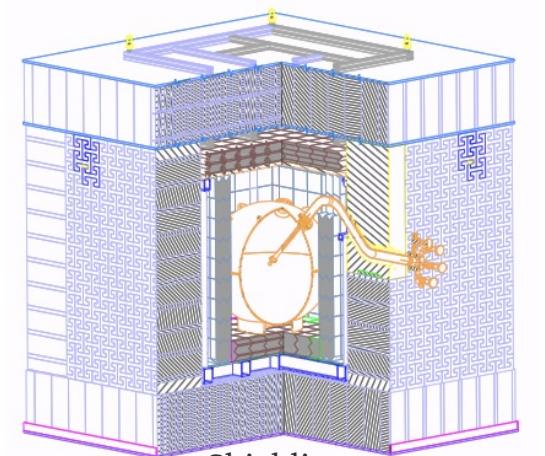
Vessel

60 cm Ø NOSV Copper



Sensor

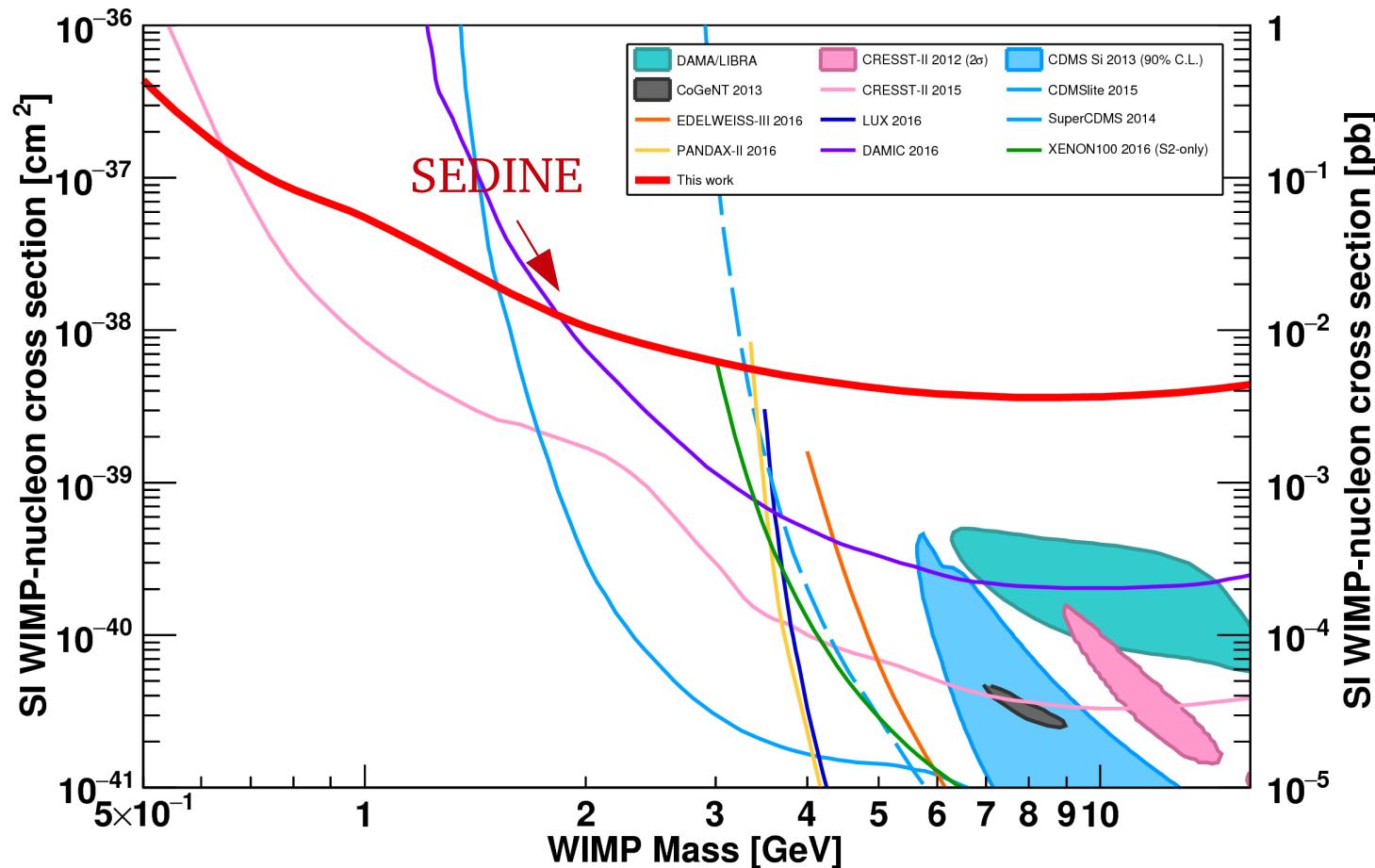
6.3 mm Ø



Shielding
30cm PE, 10-15cm Pb, [3-8]cm Cu



First results from NEWS-G at LSM



Q. Arnaud et al. (NEWS-G), Astropart. Phys. 97, 54 (2018)

doi: 10.1016/j.astropartphys.2017.10.009

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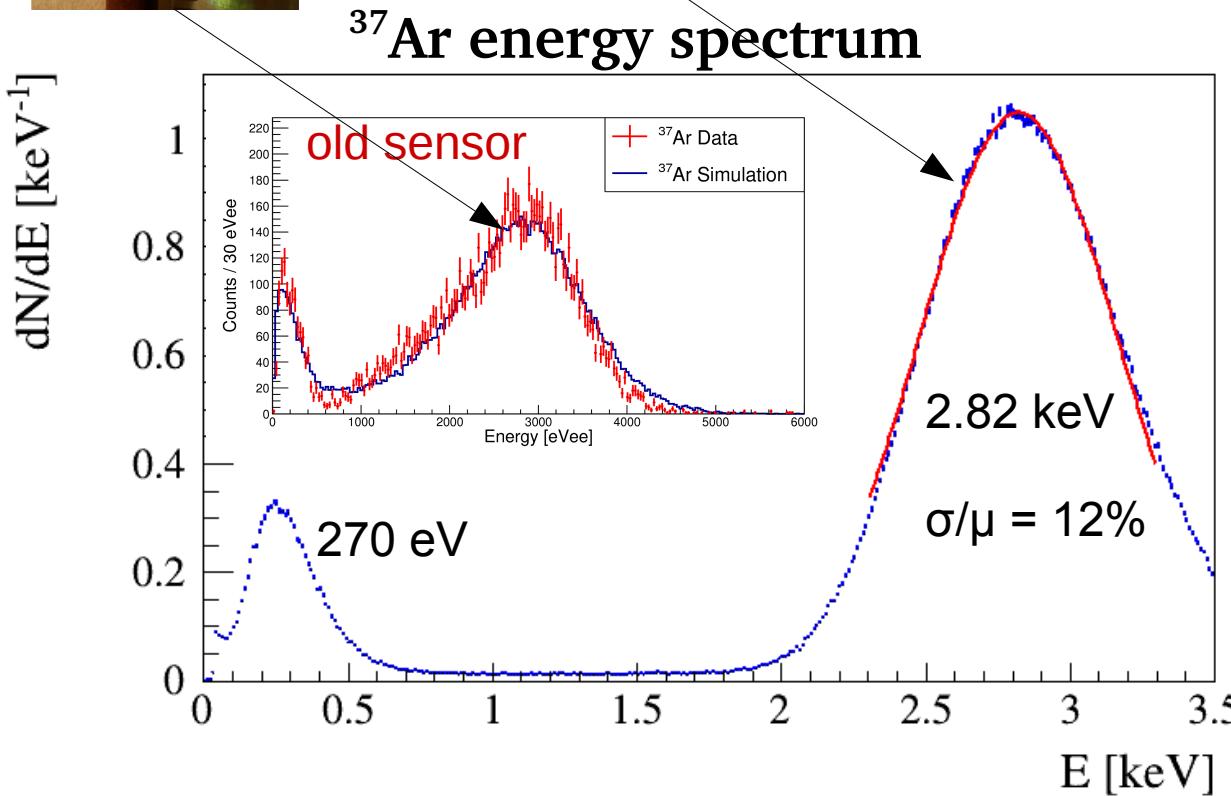
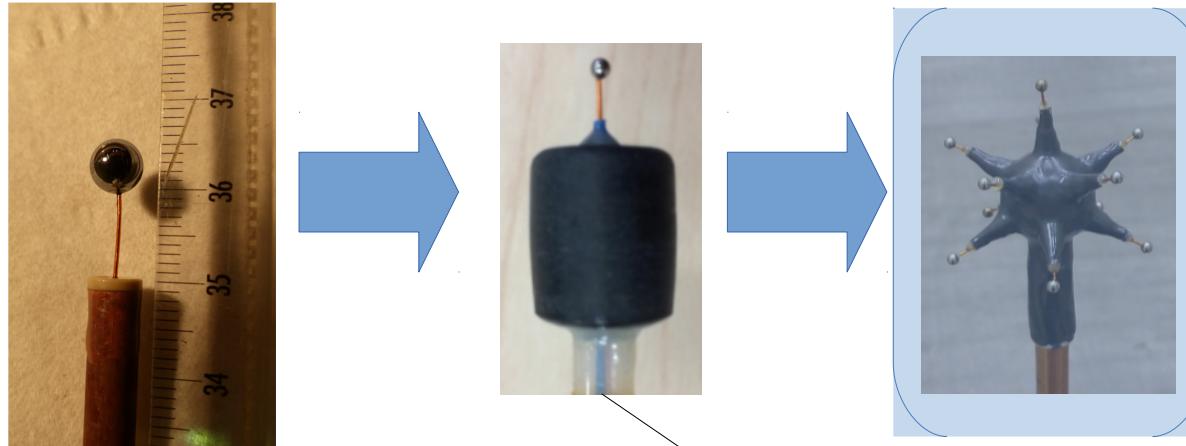
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Towards a 140cm Low background Sphere at SNOLAB



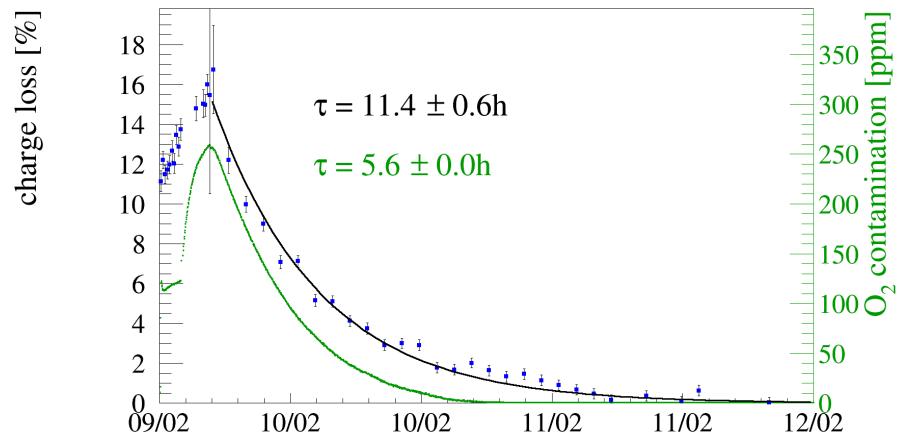
Improved electric field: new sensor rod



- Second electrode improves field uniformity
- Resistive material for spark protection
 - bakelyte
 - glass
- Significant resolution improvement
- Development of star shaped sensor
 - better for large detector
 - difficult to build

Gas purification

- O_2 contamination induces charge loss
- Simple circulator
 - Design J. Prudent
 - No noticeable contamination
- Purification with SAES getter
 - radon contamination
- Radon trap to be added



Direct detection of light WIMPs with NEWSG

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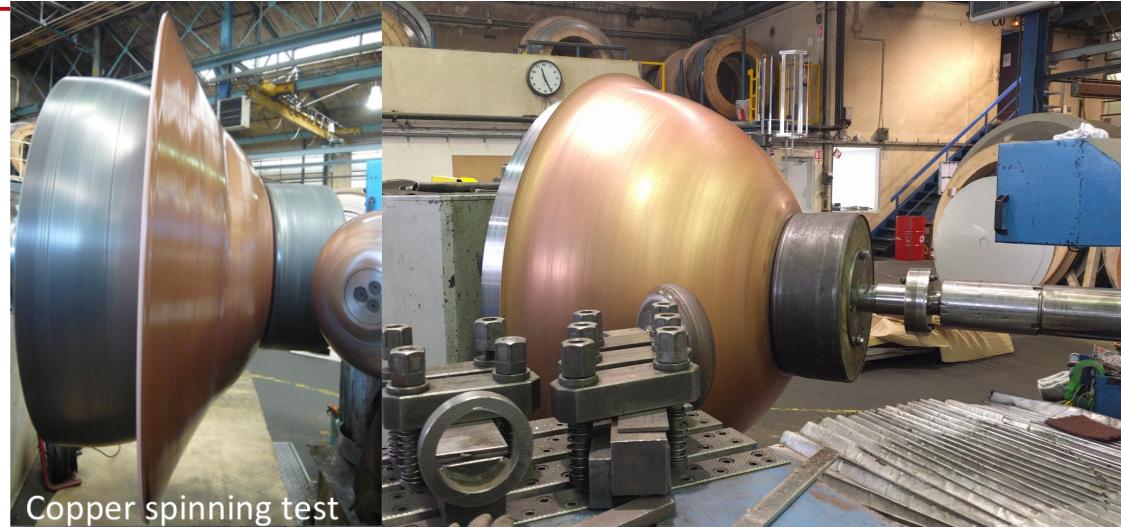
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High purity copper sphere



- Low activity copper C10100
 - 7 to 25 $\mu\text{Bq}/\text{kg}$ Th
 - 1 to 5 $\mu\text{Bq}/\text{kg}$ U



Copper spinning test



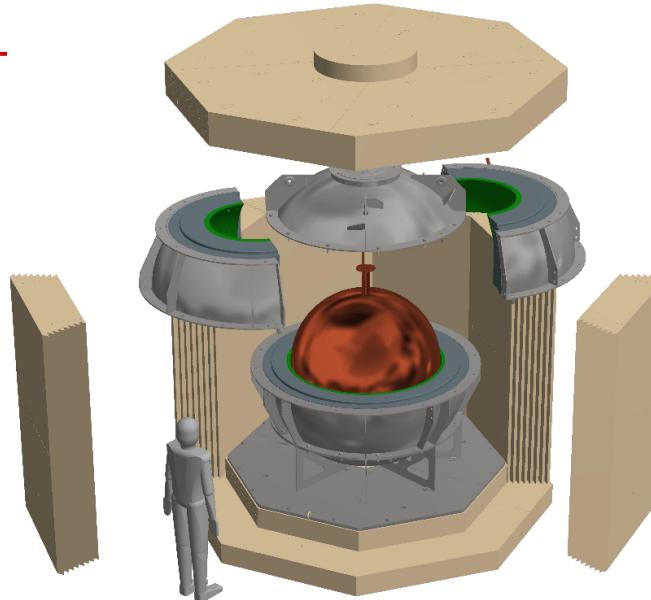
- Electropolished and Electroplated at LSM
 - 500 μm pure Copper

Direct detection of light WIMPs with NEWSG

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Shielding and others

- Compact shield
 - 3cm roman lead
 - 22cm VLA lead ($1\text{Bq/kg } ^{210}\text{Pb}$)
 - Air tight SS envelope with N_2 flushing
 - 40cm HDPE
 - *Under construction in France*
- Seismic platform *under construction at SNOLAB*
- Glove Box for rod change *built in Saclay*
- Gas purification system *tested in Kingston*
- Laser calibration *tested in Kingston*



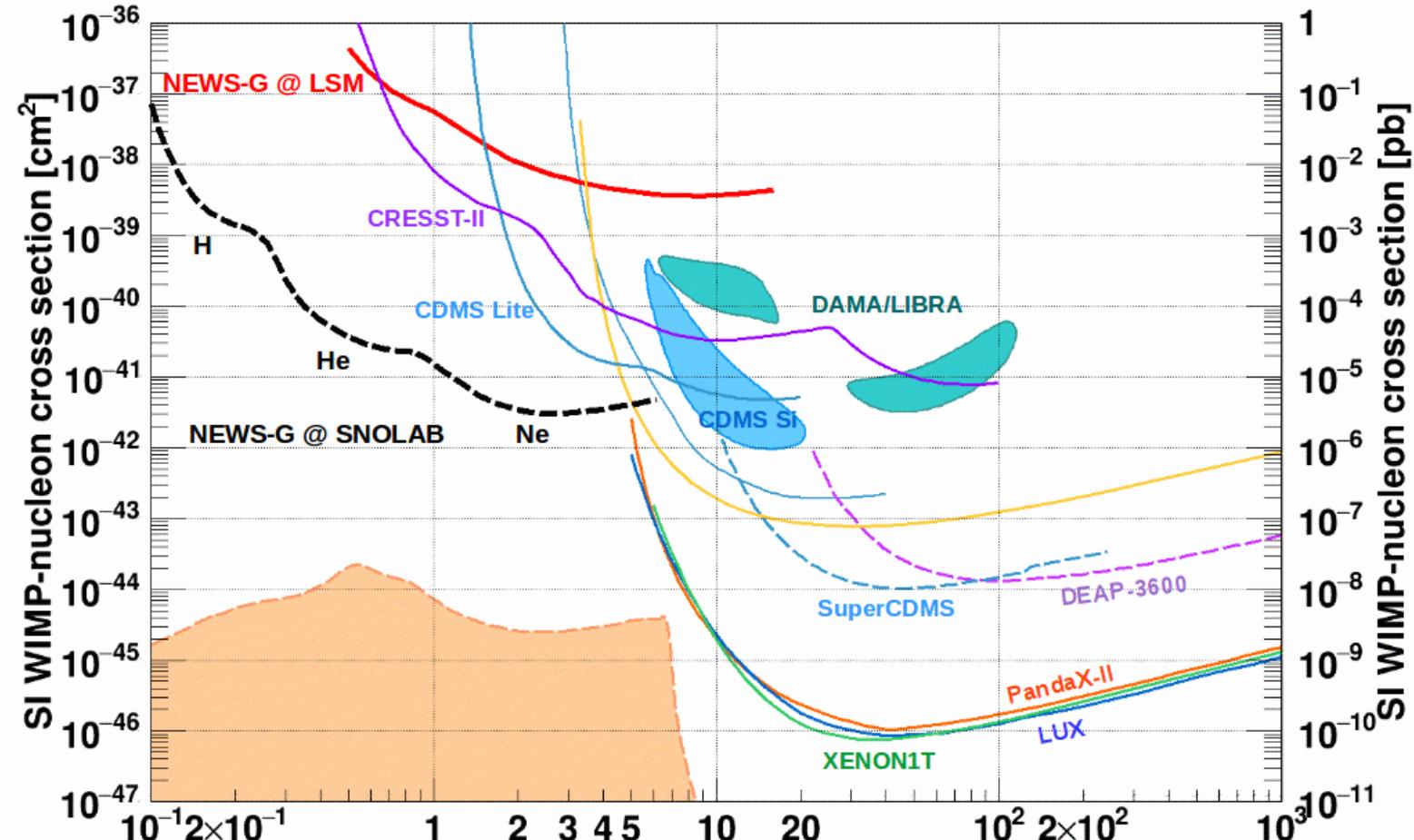
***Installation in Cube Hall at SNOLAB
summer of 2019***

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Sensitivity projection



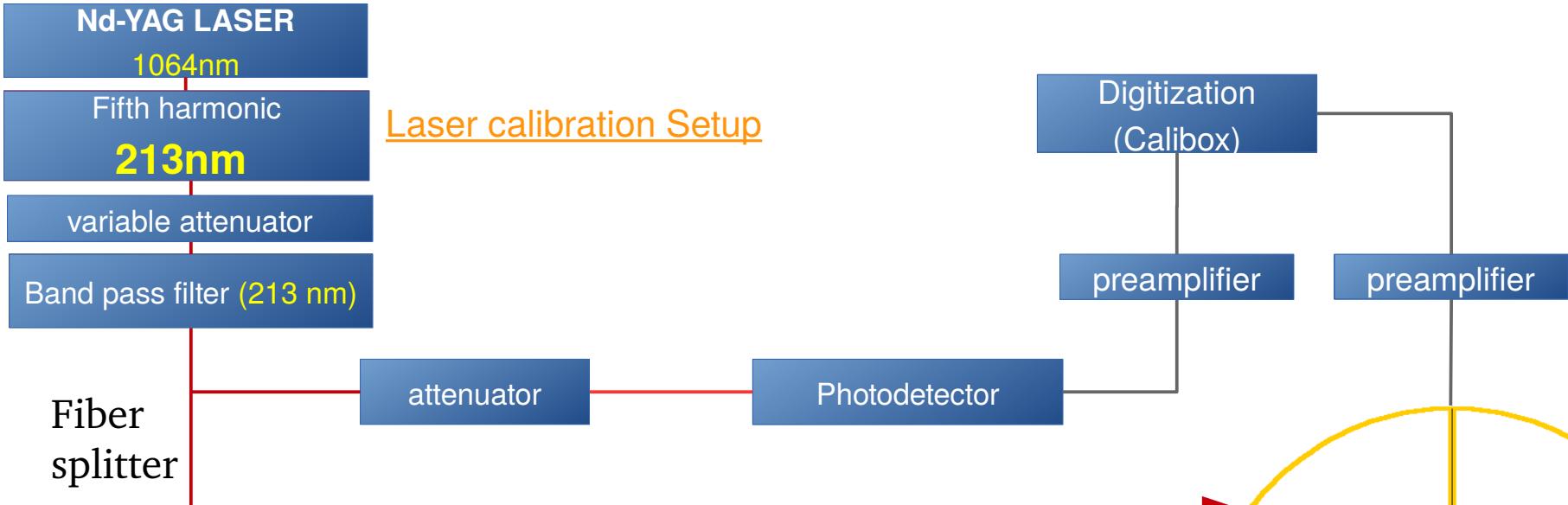
100 kg.days, 200eVee ROI above threshold @ 1 electron.

(Not accounting for sensitivity improvement from resolution effects and RT cuts)



R&D and related activities

Single electron response with UV laser

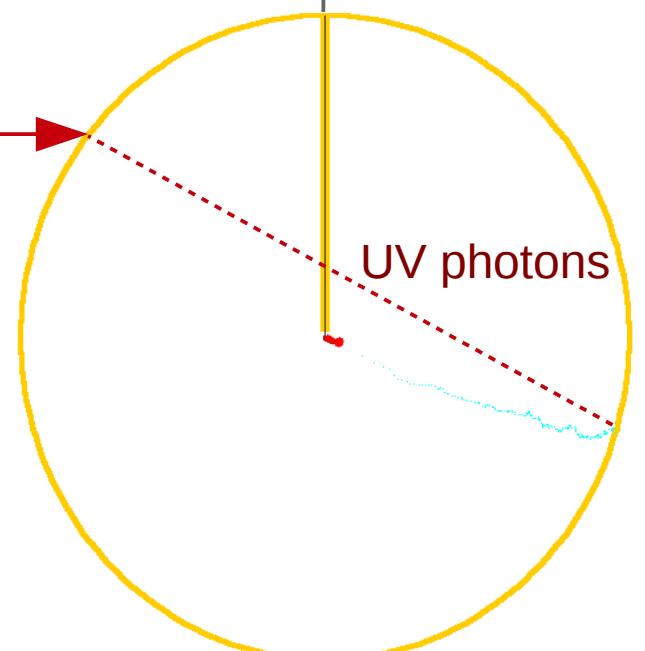


Laser Data

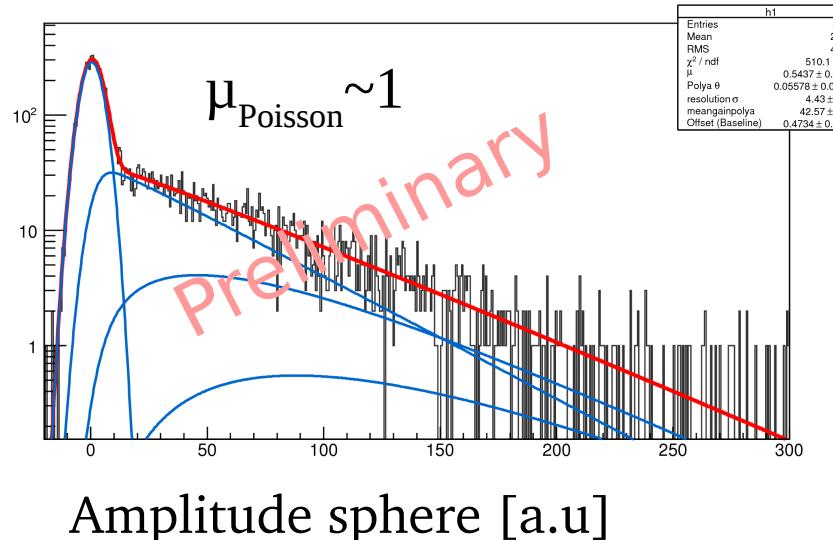
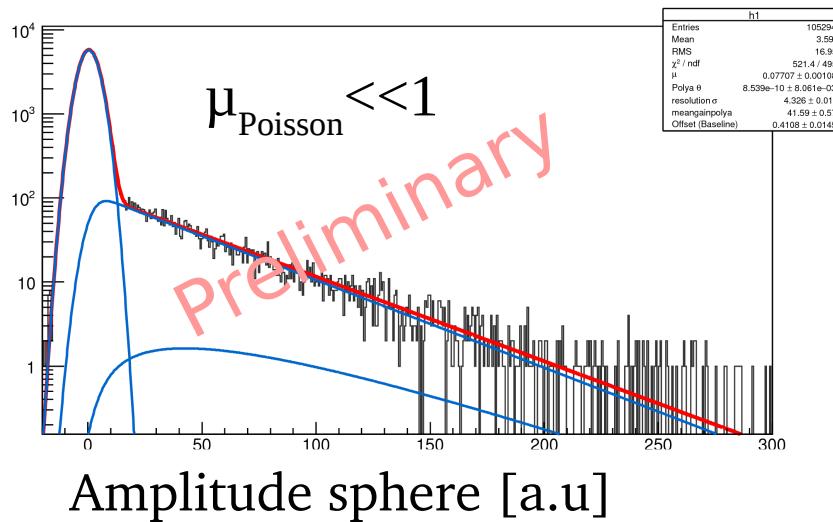
- Single electron response parametrization (θ of Polya)
- Energy calibration / W-value measurements
- Monitoring of the stability of the detector response over time
- Drift and Diffusion time measurements

Photodetector (PD)

- Monitoring of the stability of laser
- Start Time (in drift time measurements)

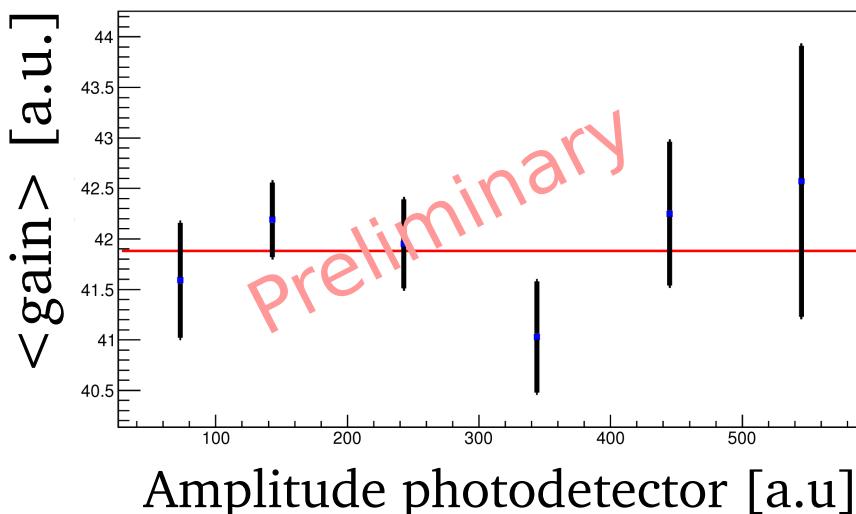
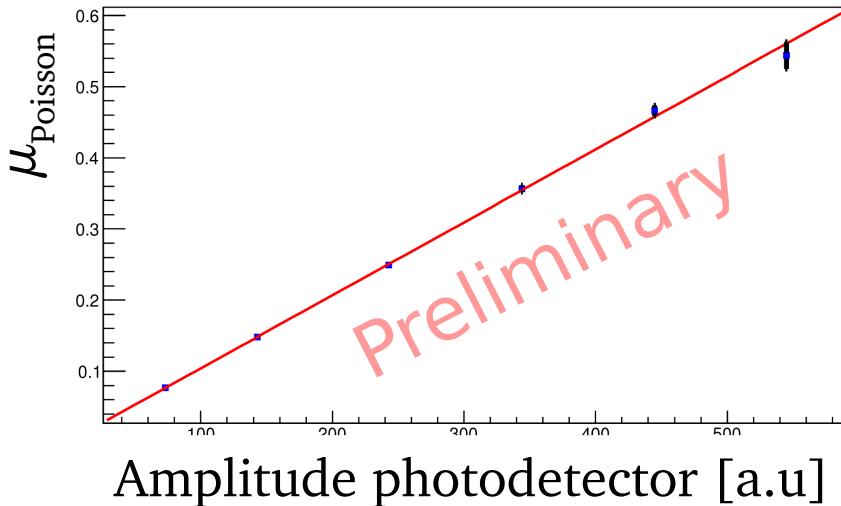


Single electron spectra



- Fitted with full model of the detector (4 parameters)
 - Poisson fluctuations: μ_{Poisson}
 - gain fluctuation: <gain>, θ
 - noise σ_{noise}
- Fit valid for multiple electrons (Poisson)
- only μ_{Poisson} depends on laser intensity

Extraction of detector parameters



- Mean number of electrons measured is proportional to laser intensity
 - single photon photoelectric process
- Mean gain and θ consistent for all intensities
 - robust fit
- Simultaneous measurement of ^{37}Ar
 - measurement of mean number of ionisation electrons for the gas mixture (first results in argon consistent with literature)



Quenching factor measurement



- Definition
 - Energy(ionisation)/Energy(recoil)
- Crucial to measurement
 - Converts energy measured to energy deposited
 - No existing measurement for Neon or Helium
 - No reliable extrapolation
- Measurement with monochromatic neutron beam at TUNL
 - preliminary measurement down to 5keV Ne recoil
 - new beam campaign planned for Ne and He



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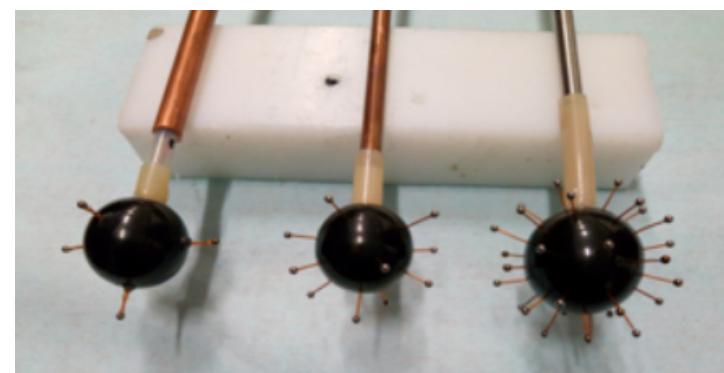
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Future sensor: Achinos

- Amplification driven by the ball size
 - Smaller ball → higher amplification.
 - Drift field far from the sensor proportional to sensor radius
 - Smaller ball → lower drift field
- Achinos structure: small balls, large sensor

$$E(r) \approx \frac{V}{r^2} r_{anode}$$



Direct detection of light WIMPs with NEWSG

Neutrino Oscillation Workshop 2018

Philippe Gros, Queen's University



Conclusions and outlook



- Low mass WIMPs are an interesting DM candidate
 - investigation only starting
- NEWS-G has an ambitious program of detection with SPCs
- SEDINE Prototype gave competitive results
- Intensive work provided great improvement in detector performance and understanding
 - improved electric field and gas purity
 - detailed measurement of amplification and drift parameters
- Construction under way for installation at SNOLAB in **summer 2019**
- And still more work to improve the future



the NEWS-G Collaboration

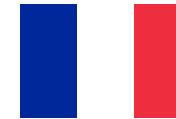


- Queen's University Kingston – G. Gerbier, P. di Stefano, R. Martin, G. Giroux, T. Noble, D. Durnford, S. Crawford, M. Vidal, A. Brossard, P. Vasquez de Sola, Q. Arnaud, K. Dering, J. McDonald, M. Chapellier, A. Ronceray, P. Gros, C. Neron, A. Rolland



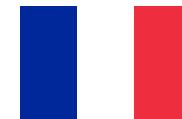
- Copper vessel and gas set-up specifications, calibration, project management
- Gas characterisation, laser calibration, on surface prototypes
- Simulation/Data analysis

- IRFU (Institut de Recherche sur les lois Fondamentales de l'Univers)/CEA Saclay – I. Giomataris, M. Gros, C. Nones, I. Katsioulas, T. Papaevangelou, J.-P. Bard, J.-P. Mols, X.-F. Navick



- Sensor/rod (low activity, optimisation of E field)
- Electronics (low noise preamps, digitization, stream mode)
- DAQ/soft

- LSM (Laboratoire Souterrain de Modane), IN2P3, U. of Chambéry – F. Piquemal, M. Zampaolo, A. Dastgheibi-Fard



- Low activity archeological lead
- Coordination for lead/HDPE shield and copper sphere



- Thessaloniki University – I. Savvidis, A. Leisos, S. Tzamarias

- Simulation, neutron calibration
- Studies on sensors

- LPSC (Laboratoire de Physique Subatomique et Cosmologie) Grenoble – D. Santos, J.-F. Muraz, O. Guillaudin



- Quenching factor measurements, Copper electroforming

- PNNL (Pacific Northwest National Lab) – E. Hoppe, R. Bunker

- Low activity measurements, Copper electroforming

- RMCC (Royal Military College Canada) Kingston – D. Kelly, E. Corcoran

- ^{37}Ar source production, sample analysis

- SNOLAB, Sudbury – P. Gorel

- Calibration system, slow control



- University of Birmingham – K. Nikolopoulos, P. Knights



- Simulation, analysis, R&D

- Associated labs: TRIUMF – F. Retiere





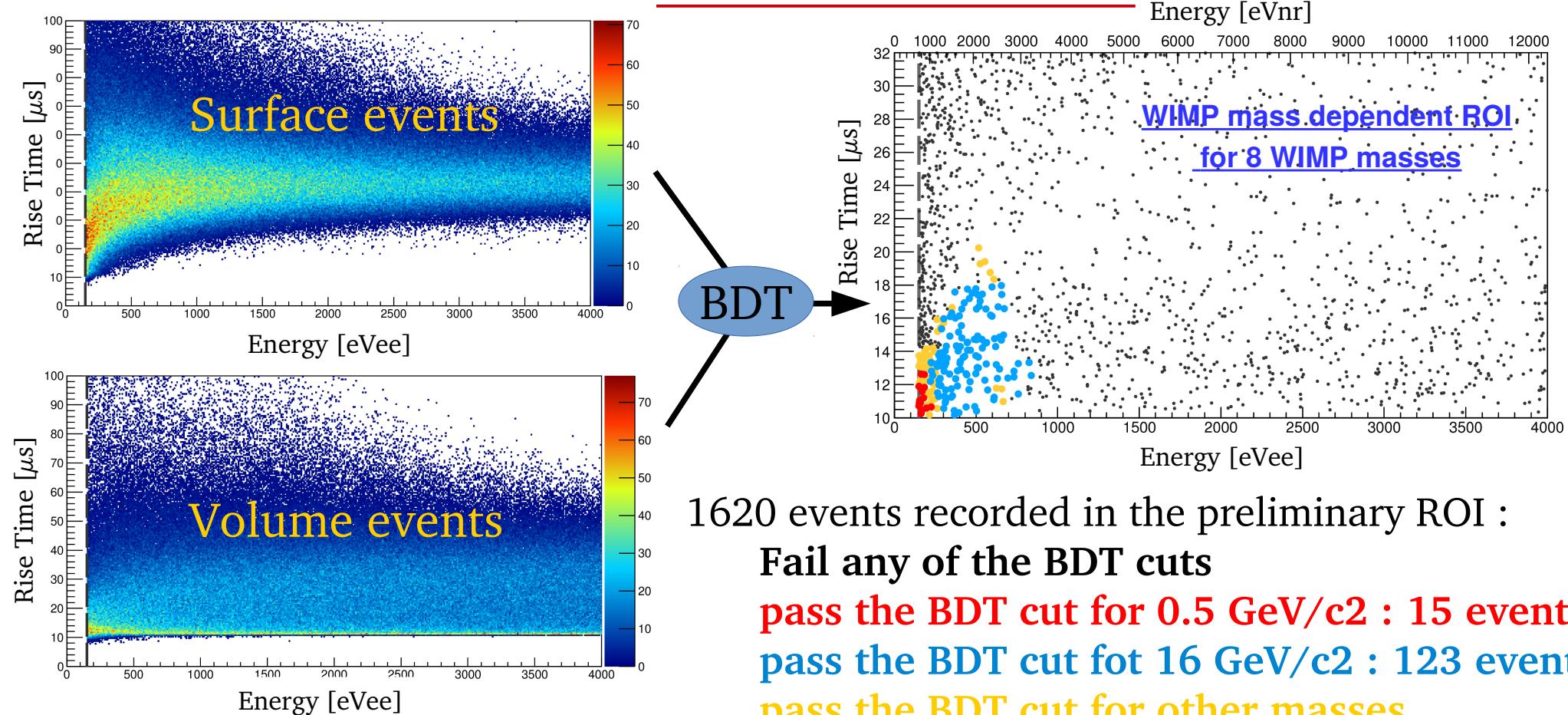
Back up

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Data analysis



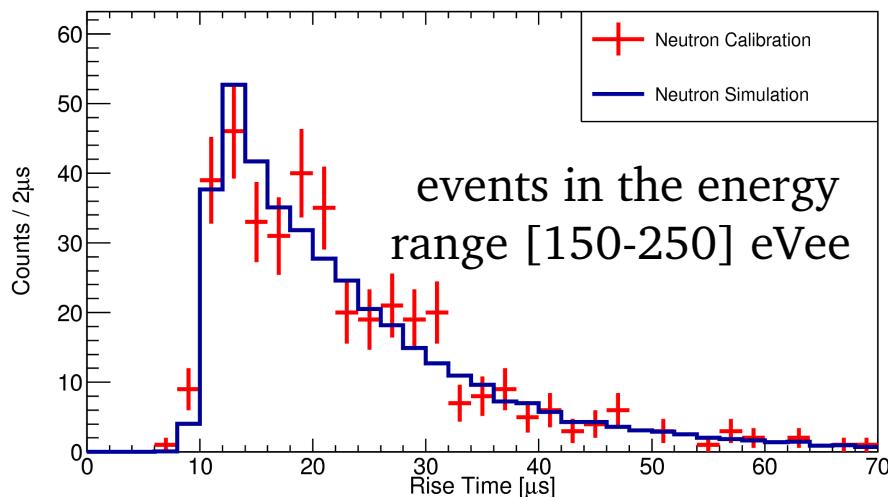
1620 events recorded in the preliminary ROI :
Fail any of the BDT cuts
pass the BDT cut for 0.5 GeV/c² : 15 events
pass the BDT cut for 16 GeV/c² : 123 events
pass the BDT cut for other masses

Analysis methodology robust against background mis-modeling:
If BDT trained with inaccurate bkg models, ROI not optimized

Simulation Validation

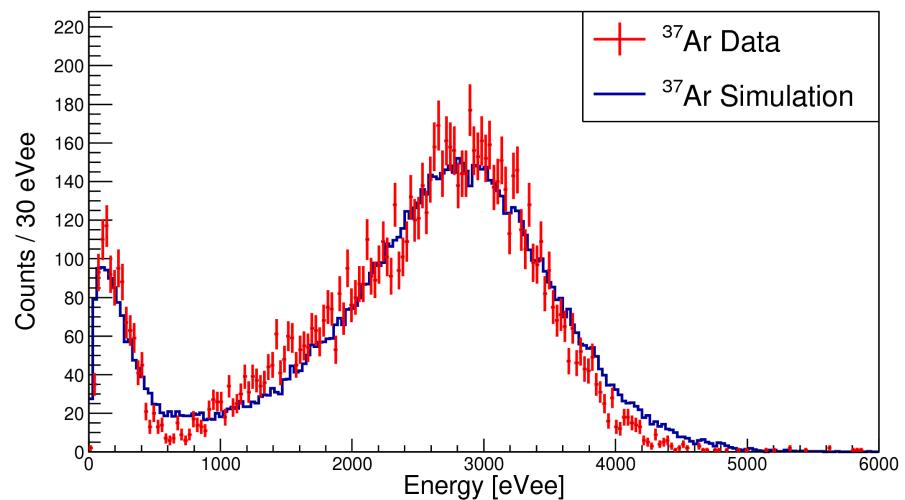
Am-Be neutron source

Nuclear recoils
homogeneously distributed
in the volume



^{37}Ar gas added to the mixture

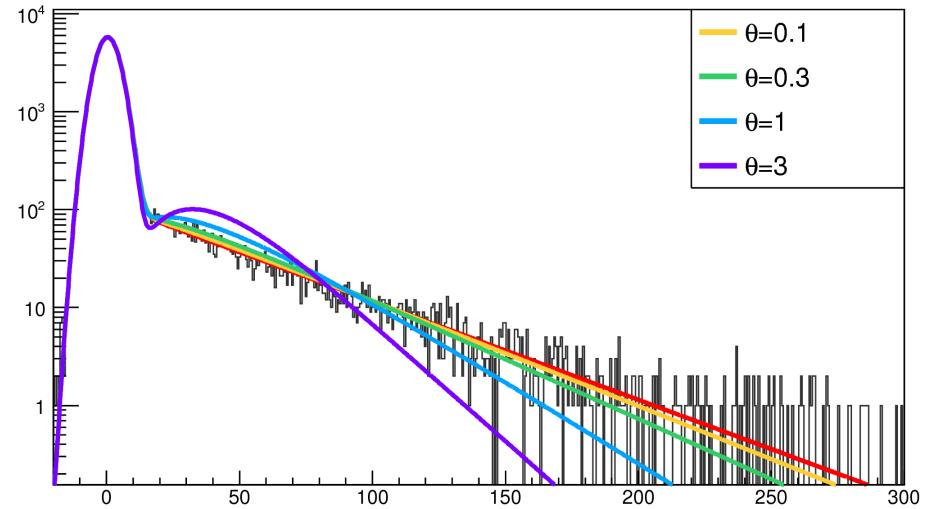
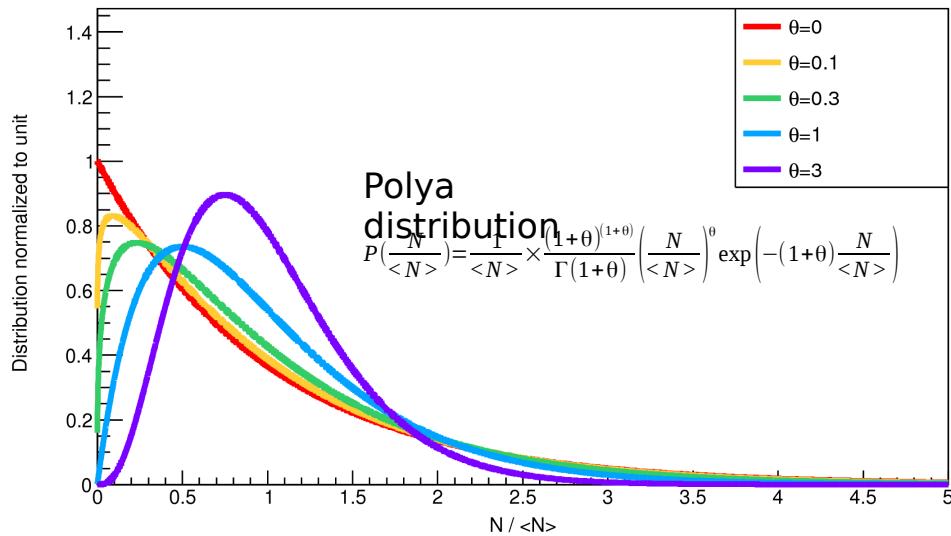
2.82 keV and 270 eV X-rays from
the electron capture
in the K- and L-shells respectively



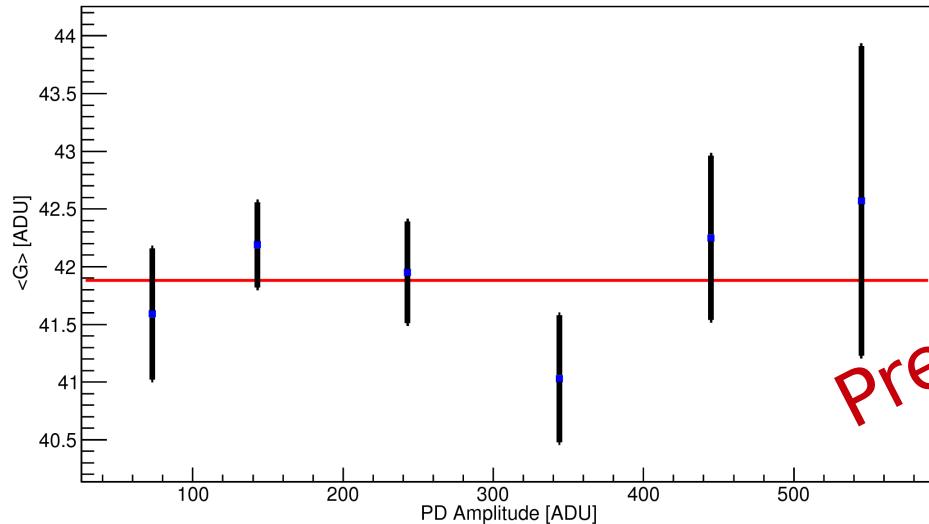
The overall agreement allows us to confidently derive our sensitivity from simulated WIMP events

Laser calibration measurements

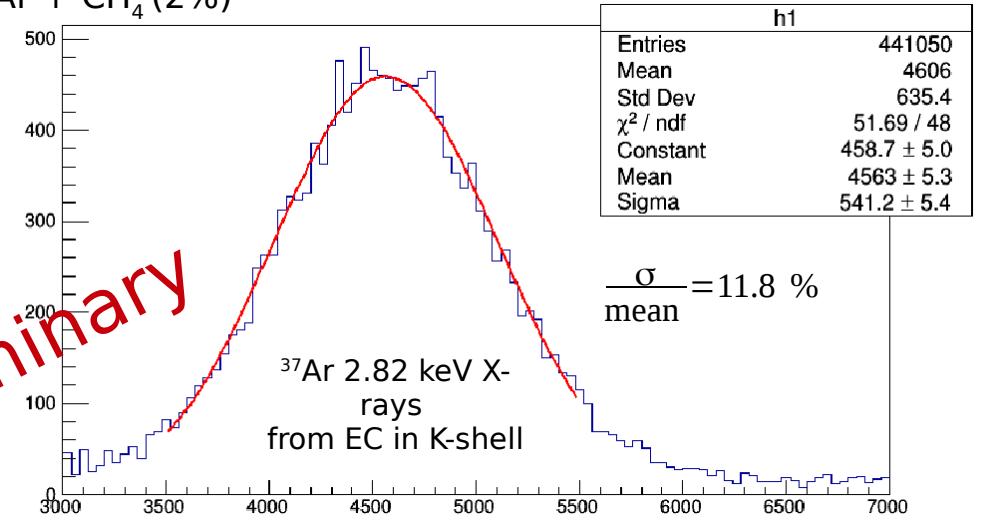
Parametrization of the Single Electron Response (SER)



W-value measurement and upper limit on the Fano factor in 500 mbar of Ar + CH₄ (2%)



Preliminary



$$\left(\frac{\sigma}{\text{mean}}\right)^2 \times \frac{E_R}{W} = \left(F + \frac{1}{1+\theta}\right)$$

$F < 0.53$

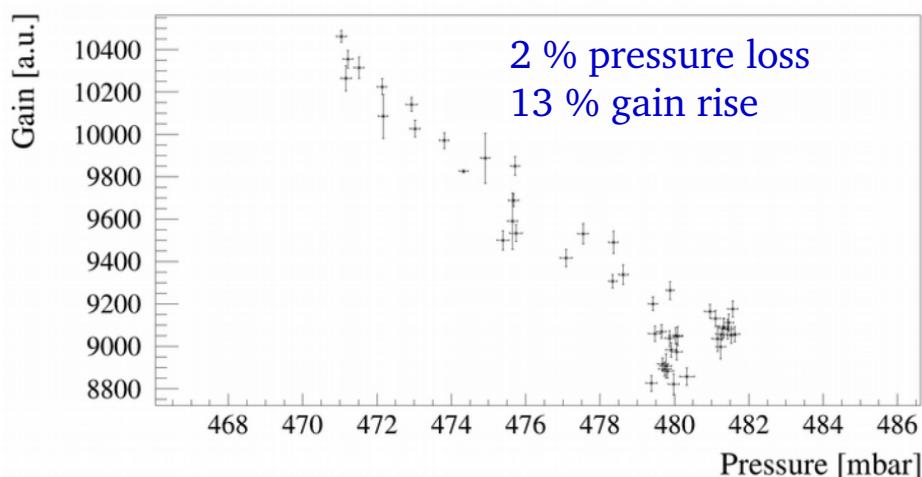
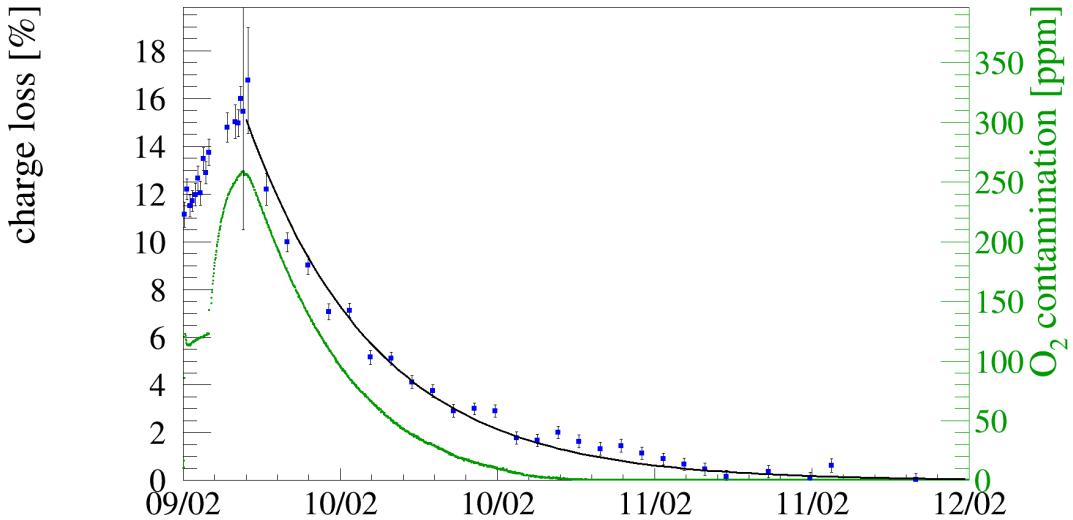
$\langle \text{Gain} \rangle \sim 41.9 \text{ ADU}$

$\&$

2820 keV peak @ 4563 ADU

Gas quality

- Oxygen captures drifting electrons → signal loss
- RGA monitors oxygen contamination
- Gas purifier removes oxygen from gas mixture



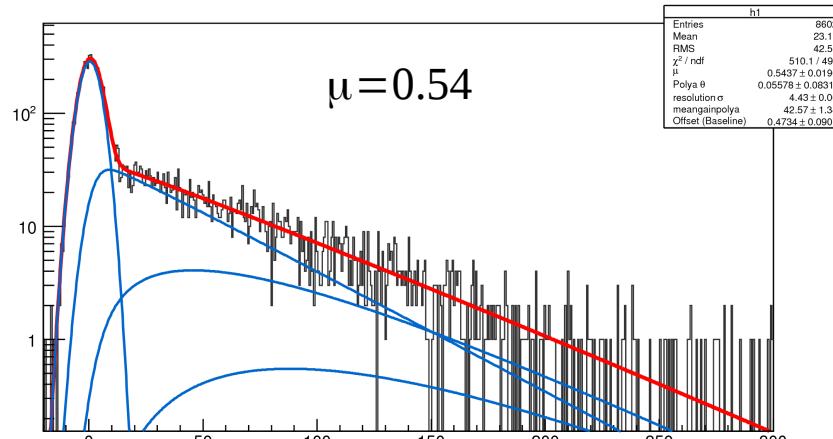
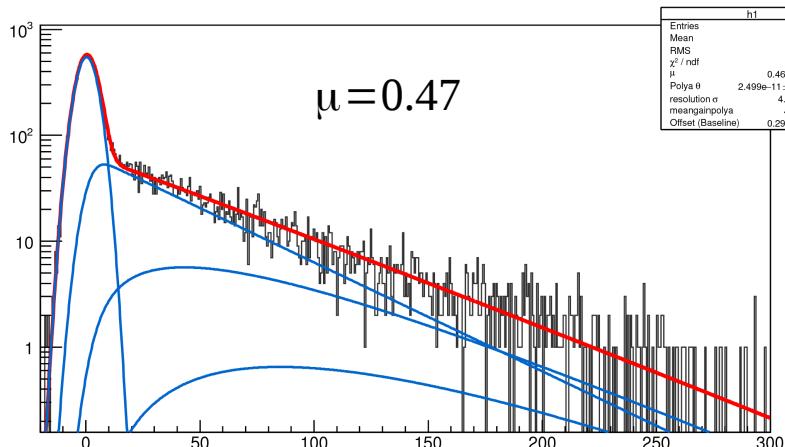
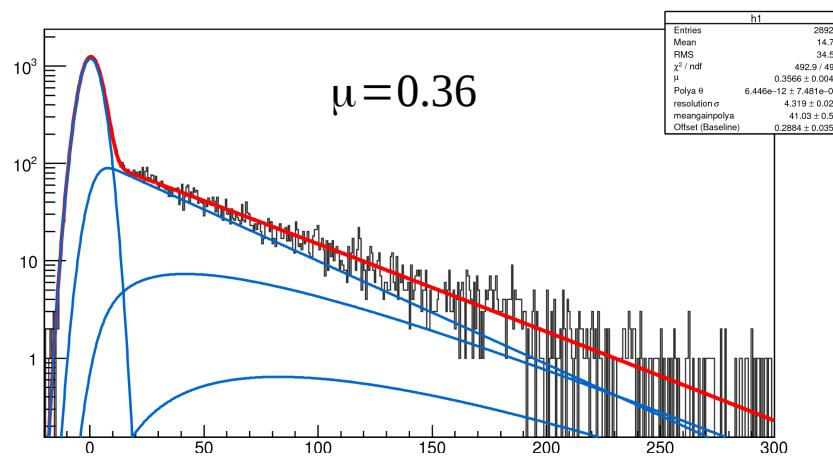
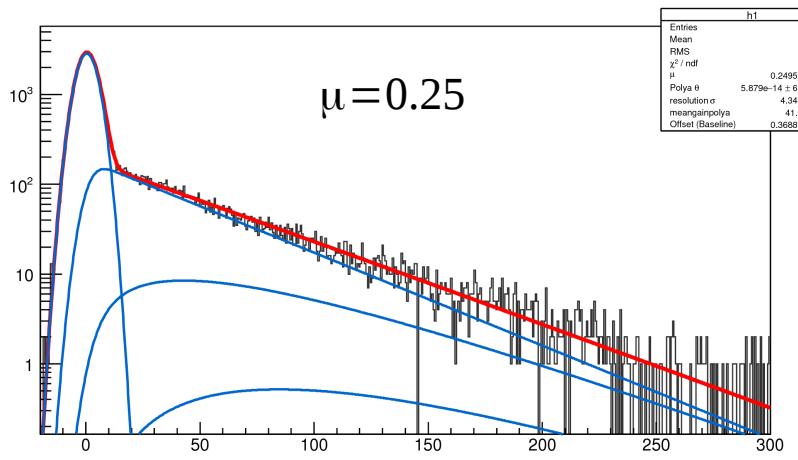
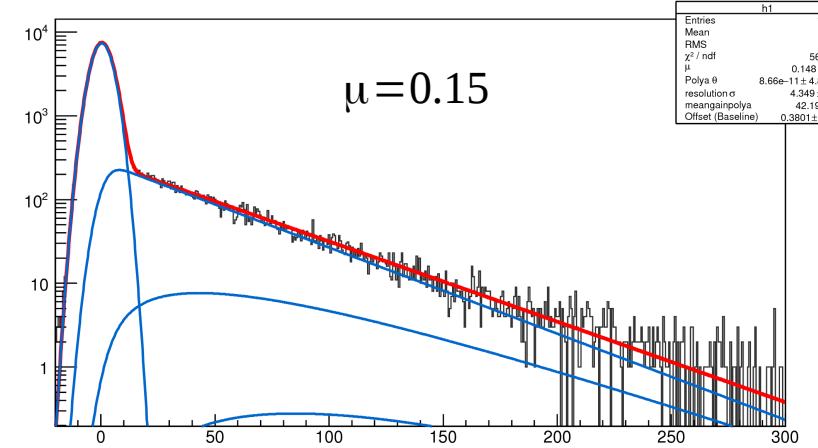
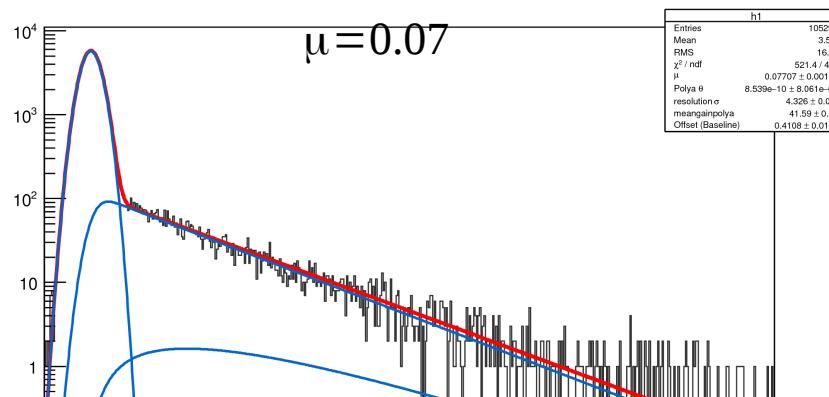
- Gain very sensitive to gas pressure
- Continuous monitoring of pressure
- measurement of dependence

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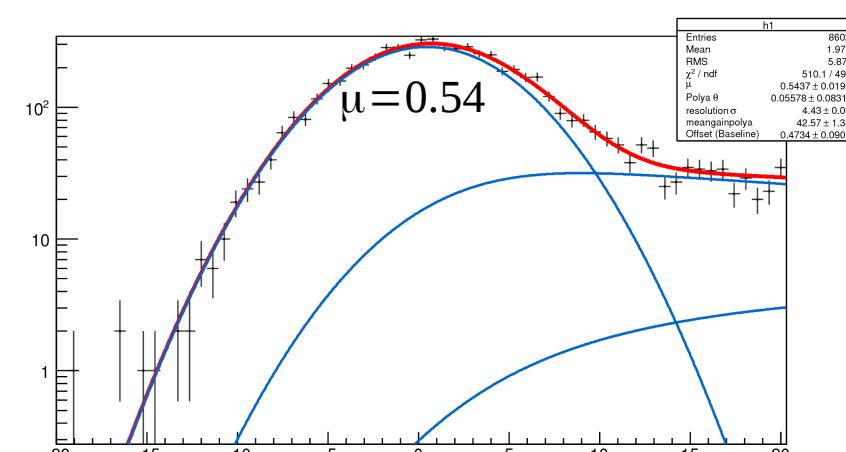
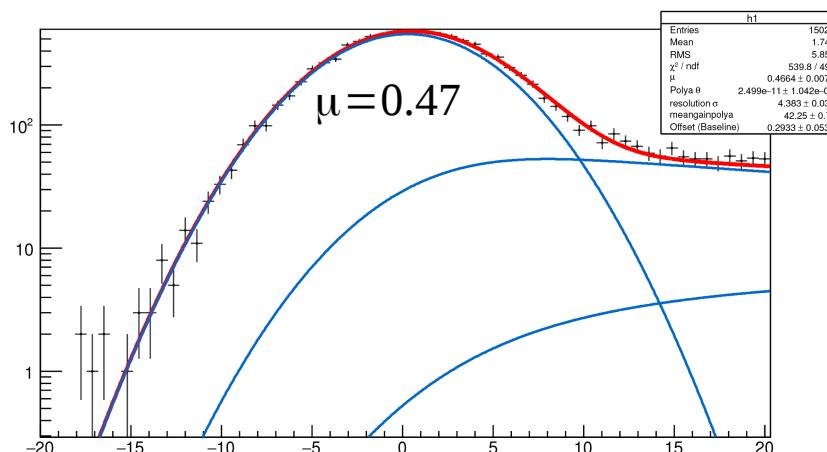
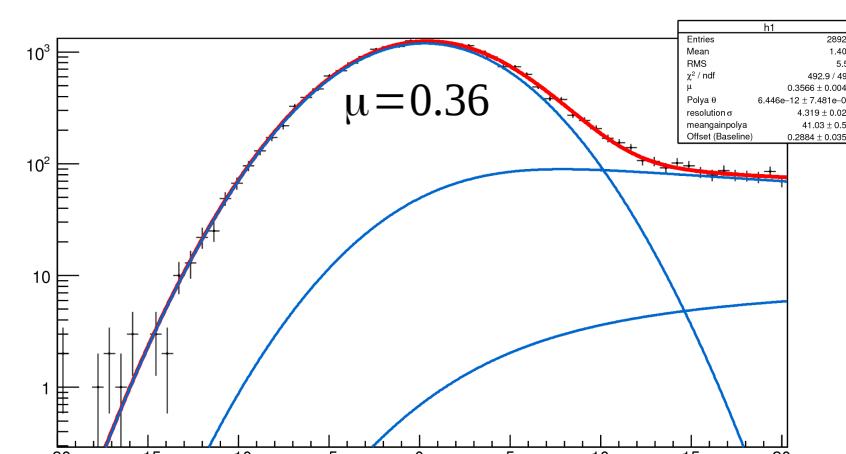
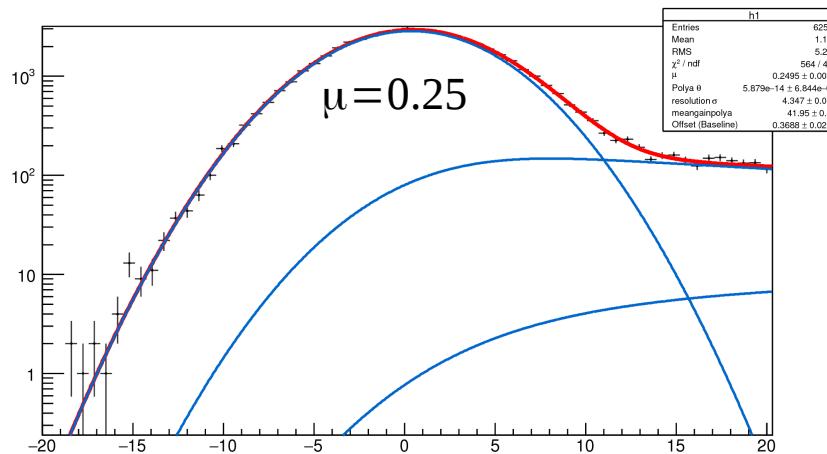
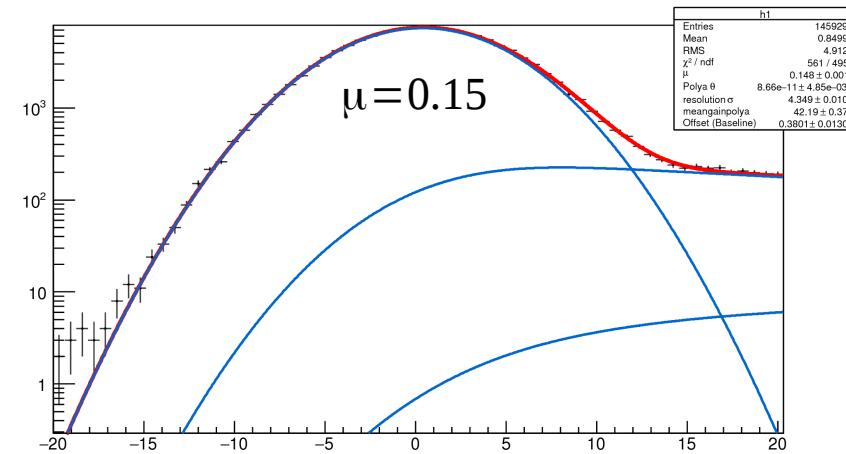
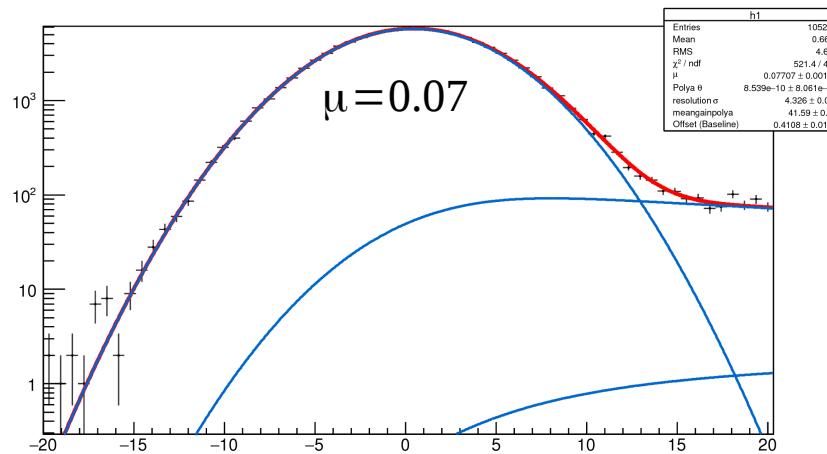
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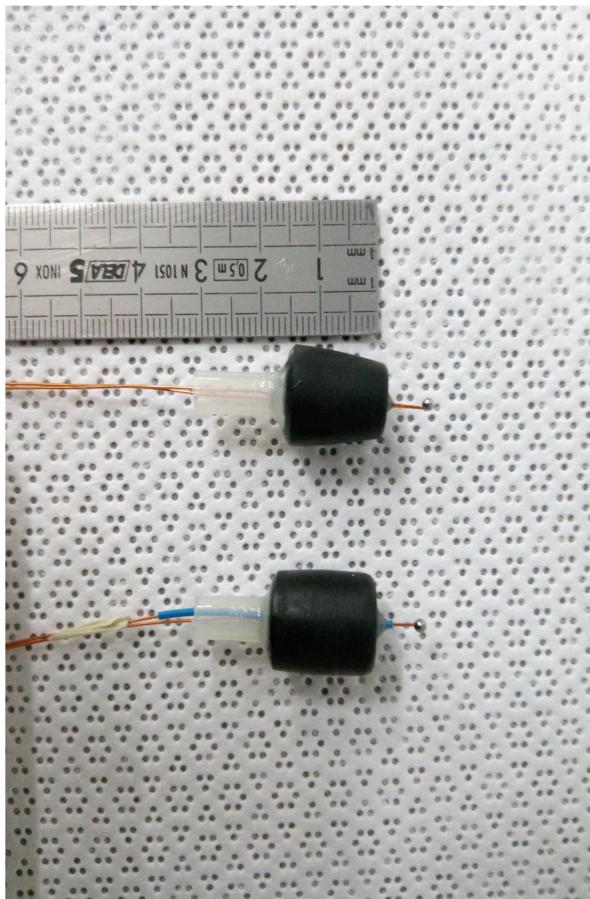
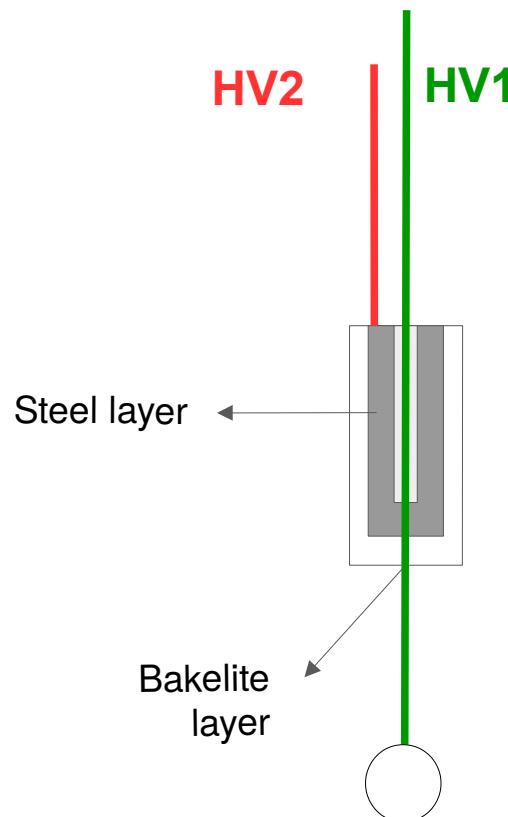
Fit of our model to Real data



Fit of our model to Real data (zoom in the low energy region)



The bakelite resistive umbrella



Bakelite
Chemical Formula:
 $(C_6H_6-O-C-H_2-O)_x$

Thermosetting phenol formaldehyde resin, formed from a condensation reaction of phenol with formaldehyde.

Advantages:

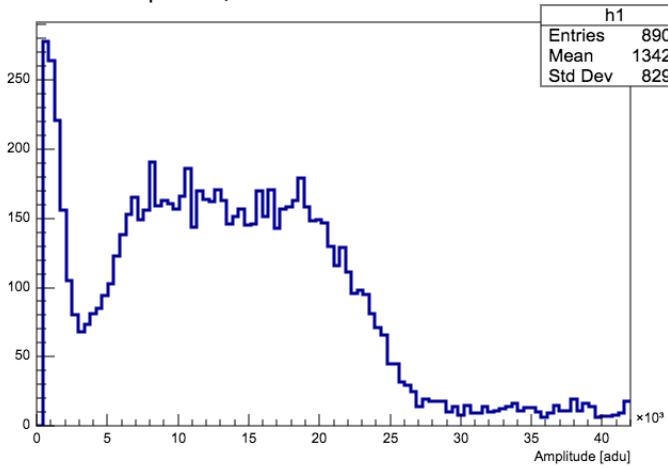
- Bakelite resistivity up to $\sim 10^{12} \Omega\text{cm}$
- Compact and homogenous material



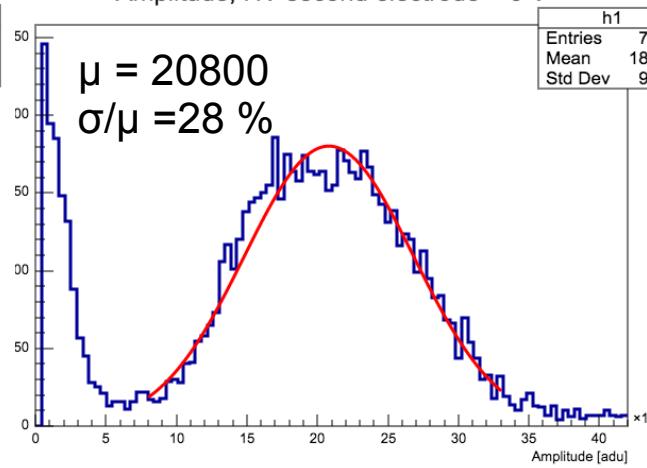
30 cm diameter sphere / Gas mixture: Ar + 2% CH₄ @ 500 mbar
Source: ³⁷Ar Electronic capture released 0.27 or 2.8 keV



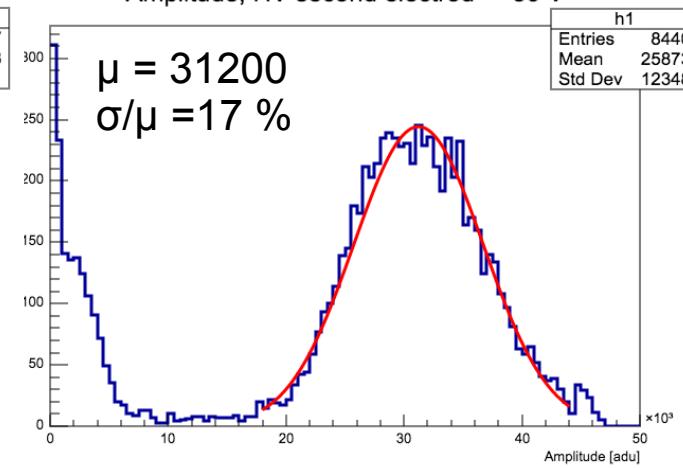
Amplitude, HV second electrode = +50 V



Amplitude, HV second electrode = 0 V

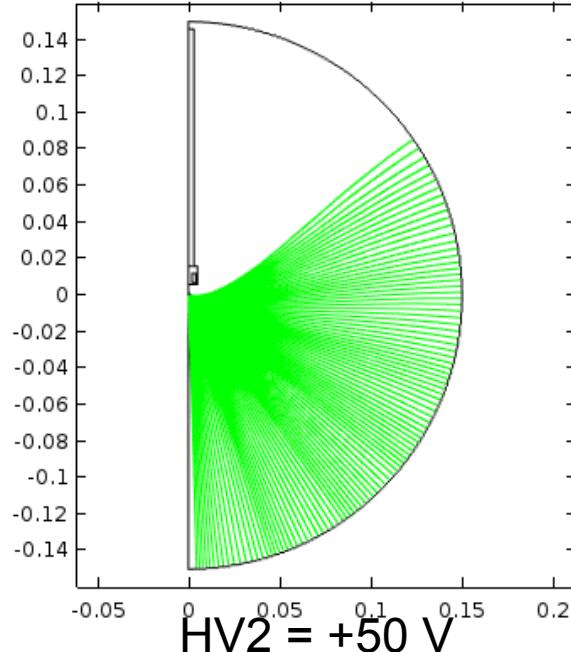


Amplitude, HV second electrode = -50 V

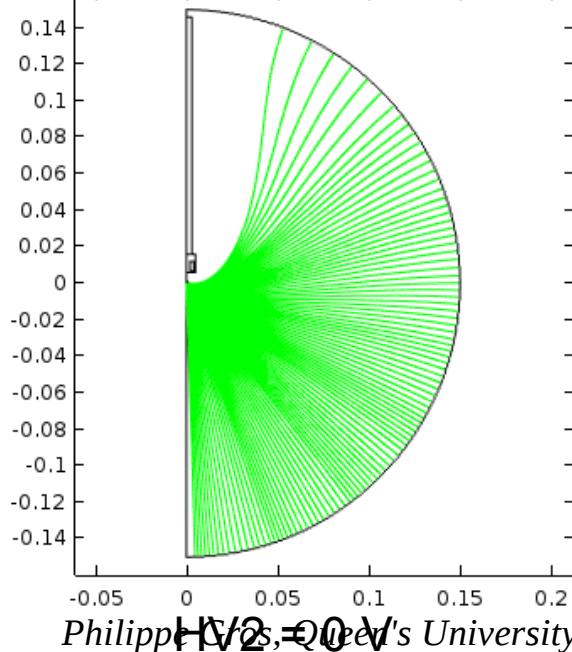


Electric field lines reaching the lower half of the sensor

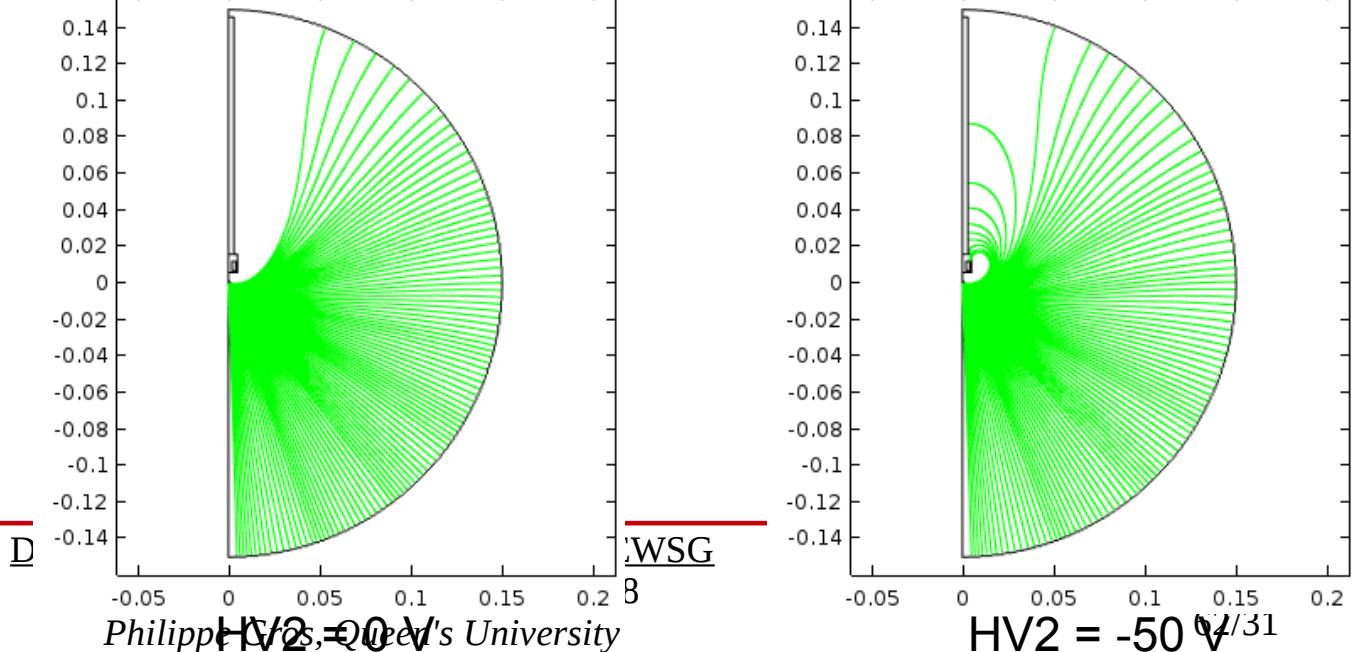
Streamline: Electric field

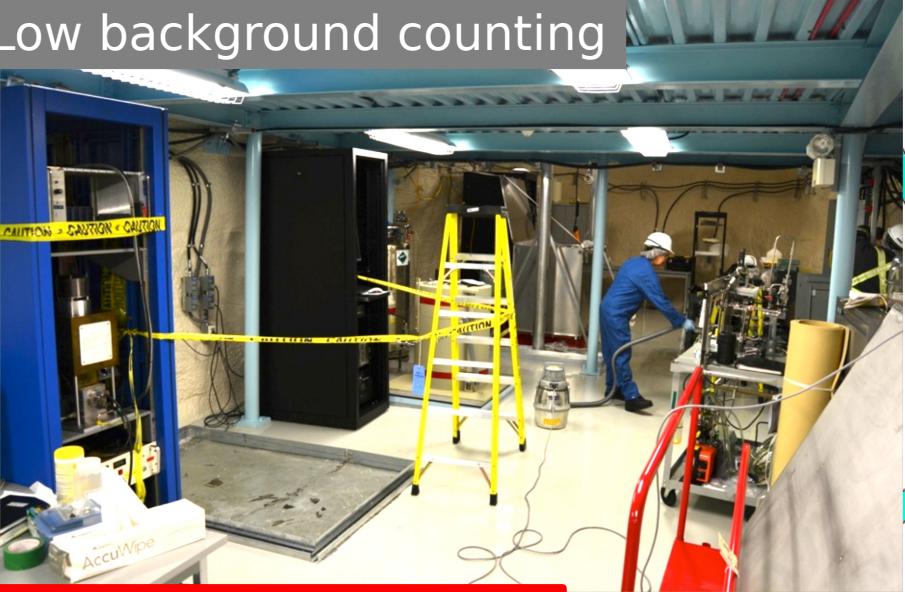
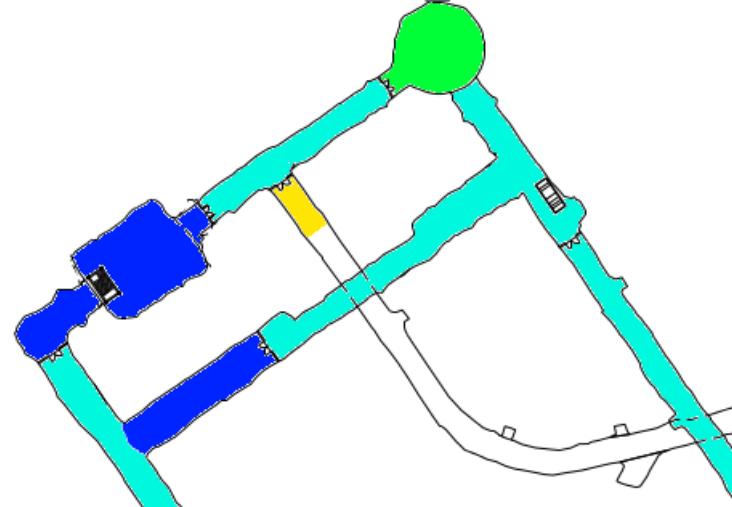
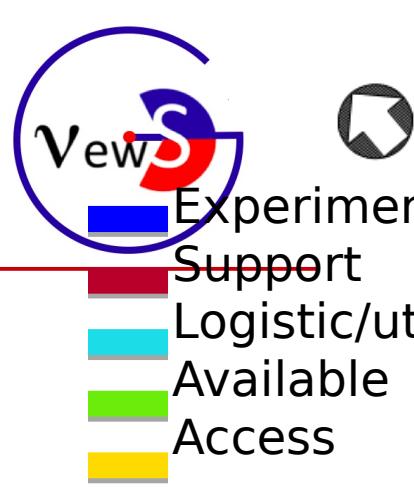


Streamline: Electric field

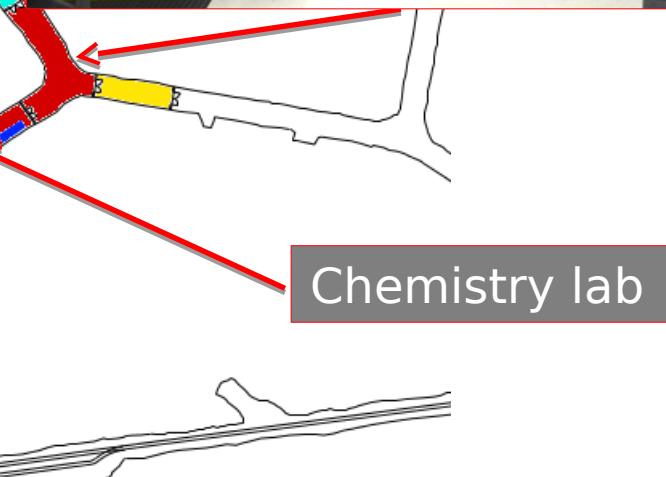


Streamline: Electric field





- 3 Ge counter
- Well detector
- Alpha counter (proj)
- Low Rn lab (proj)
- Emanation chamber
- XR fluorescence



Refuge/lunch room