

Improving the BAO peak measurement by reconstruction: the Fast Action Minimization Method

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BAO as a standar ruler

Sound Horizon, r_s == radius of perturbation shell

Key idea:

Cosmological model \rightarrow prediction for r_s

\rightarrow Measure r_s to constrain the Cosm. Model

How to do it:

The BAO signature is **imprinted** in the density field

\rightarrow Looking at the density field to “detect the shell”.

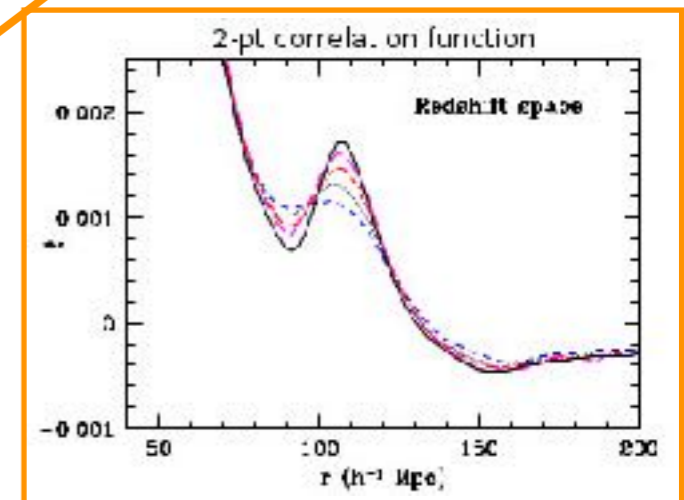
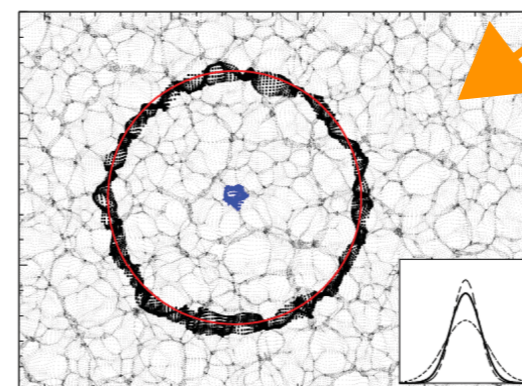
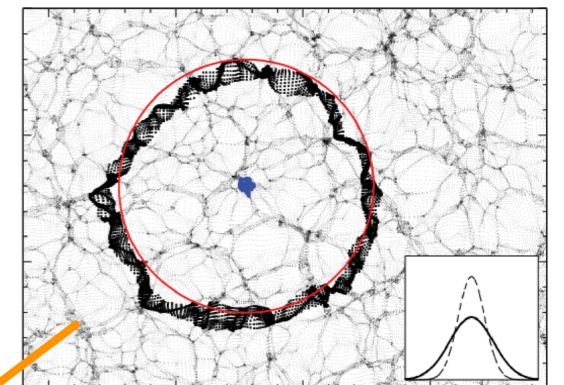
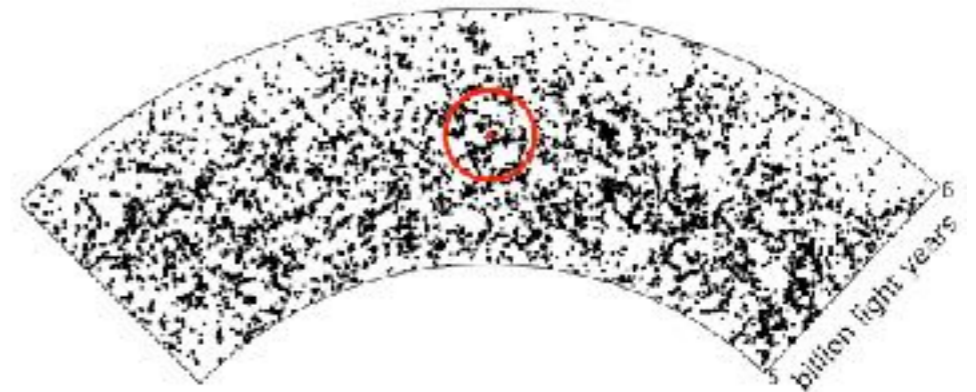
Ex: 2pt-correlation function, **BAO == peak**

Problem:

Non linear clustering \rightarrow broadening of the peak

Solution:

RECONSTRUCTION !



eFAM* Results (Preliminary)

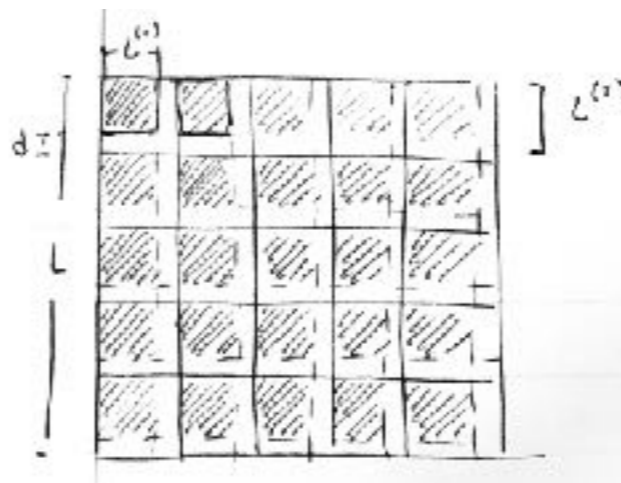
Simulation: DEUS

$z = 0$	$M_{\text{halo}} > 10^{12} M_{\text{sun}}$	$L_{\text{box}} = 21 \text{ Gpc } h^{-1}$	Cosmology: LCDM, WMAP7
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Mocks: sub-cubes

Cutting the Parent simulation into **512** Sub-cubes of $L_{\text{sub}} = 2 \text{ Gpc } h^{-1}$ ($N_{\text{halos}} \sim 23\text{k}$)
Separated by a $0.5 \text{ Gpc } h^{-1}$

* *Non-linear* numerical action method instead of Lagrangian *perturbative* à la Padmanabhan



CPU time eFAM < 3h

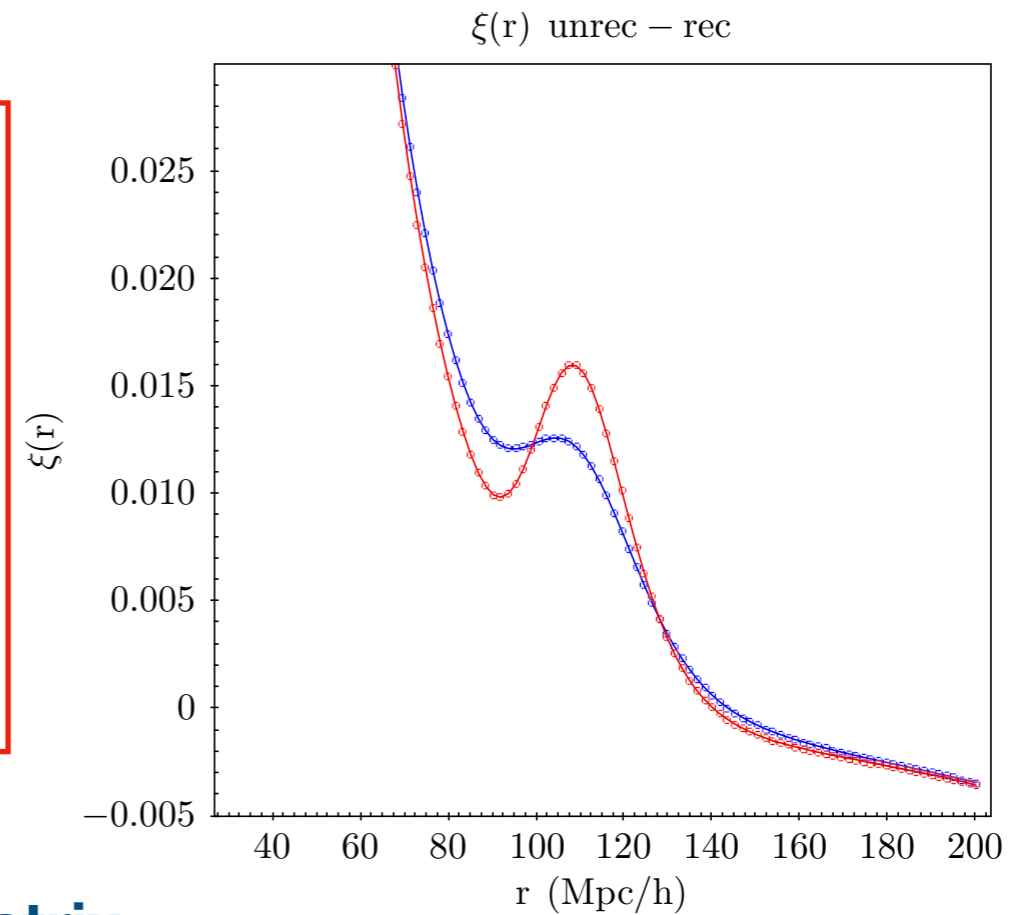
2pt correlation function

- Routine: CosmoBolognaLib (F. Marulli)
- Estimator: Landy and Szalay $\xi(r) = \frac{DD(r)+RR(r)-2DR(r)}{RR(r)}$
- Power Spectrum calculator: CAMB
- Minimization method: MCMC

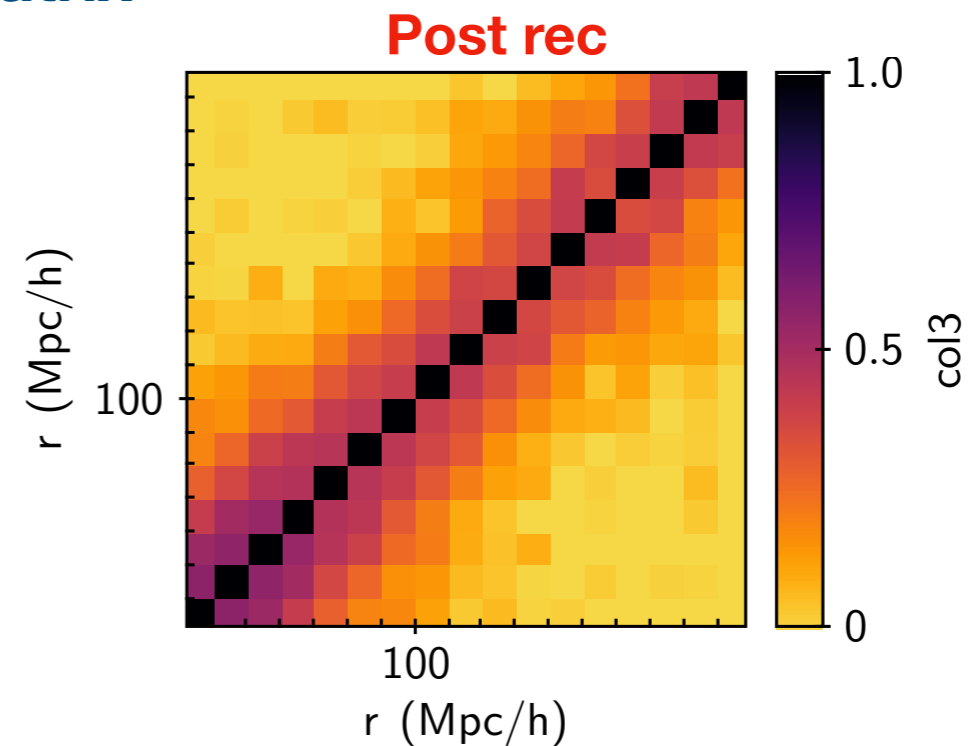
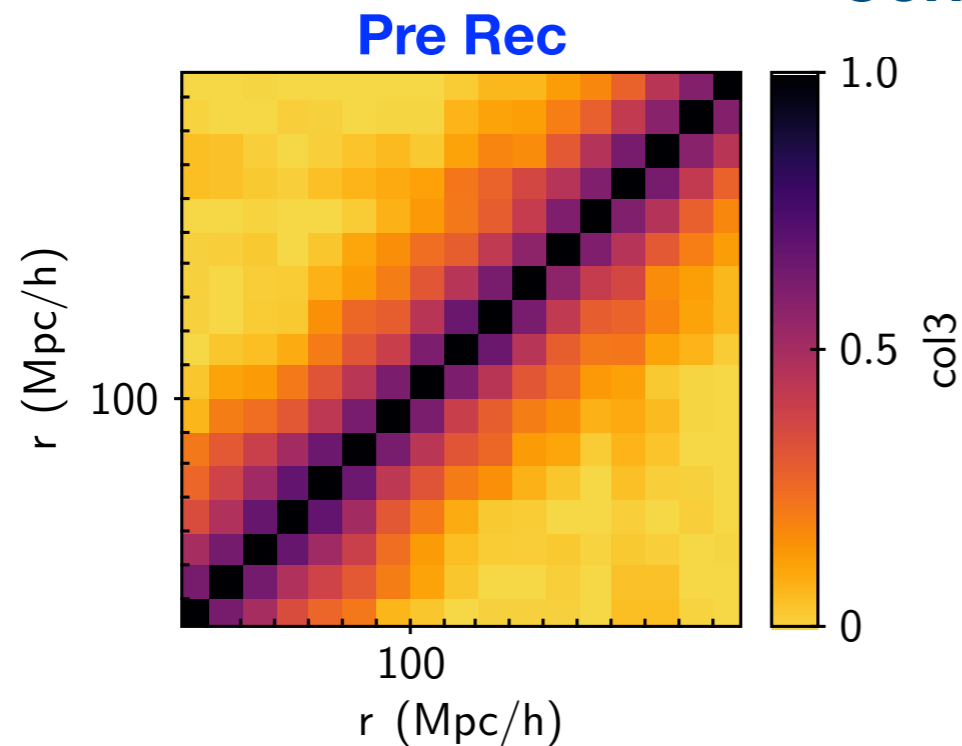
eFAM Results: fitting the averaged 2p-cf

	Pre Rec	Post Rec
α	1.125	1.121
σ_α	0.00250413	0.00205138
Σ_{nl}	$(10.4 \pm 0.9) \text{ Mpc}^{-1}$	$(3.28 \pm 1.4) \text{ Mpc}^{-1}$

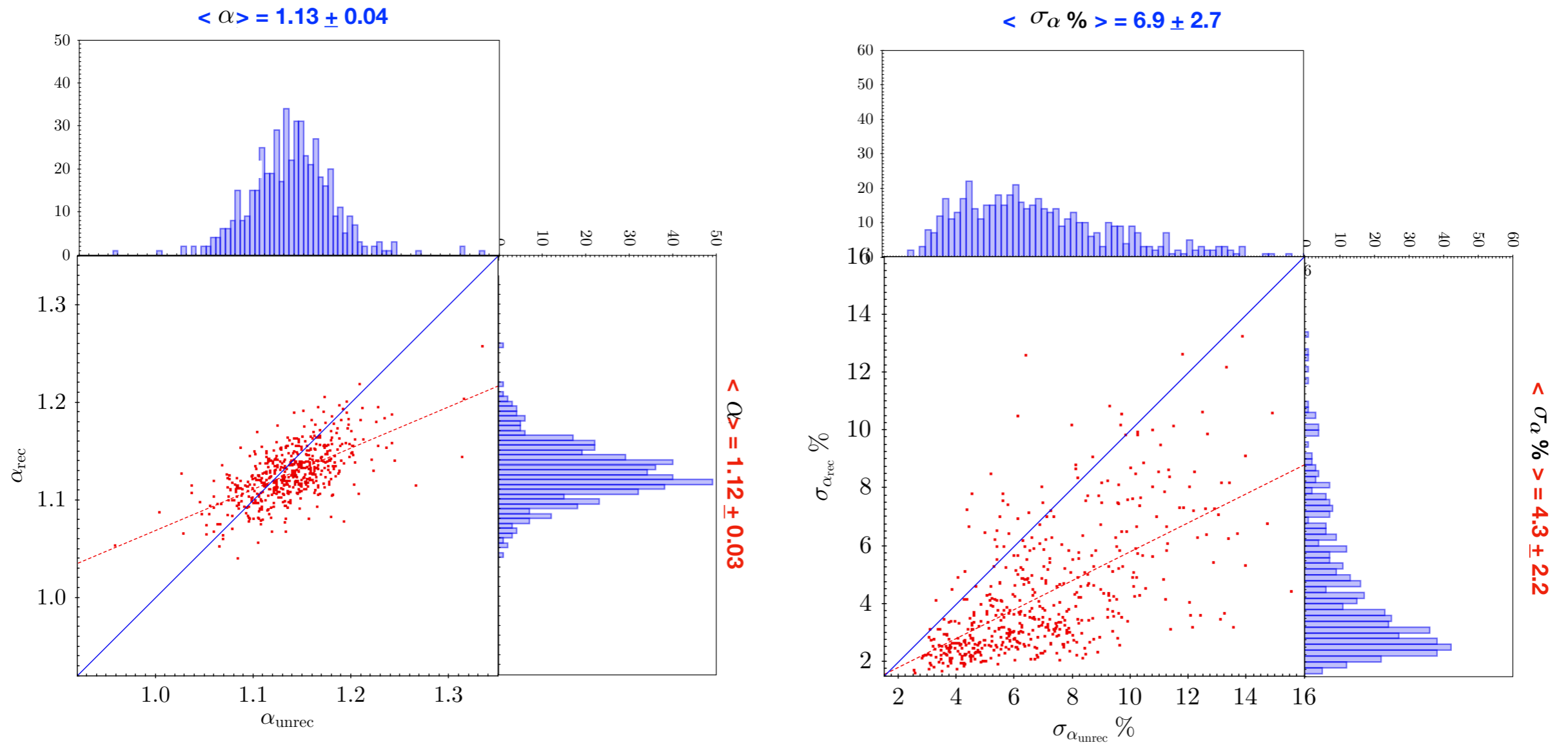
Improving factor: $I = \frac{\sigma_\alpha^{\text{pre}}}{\sigma_\alpha^{\text{Post}}} \sim 1.22$



Correlation Matrix



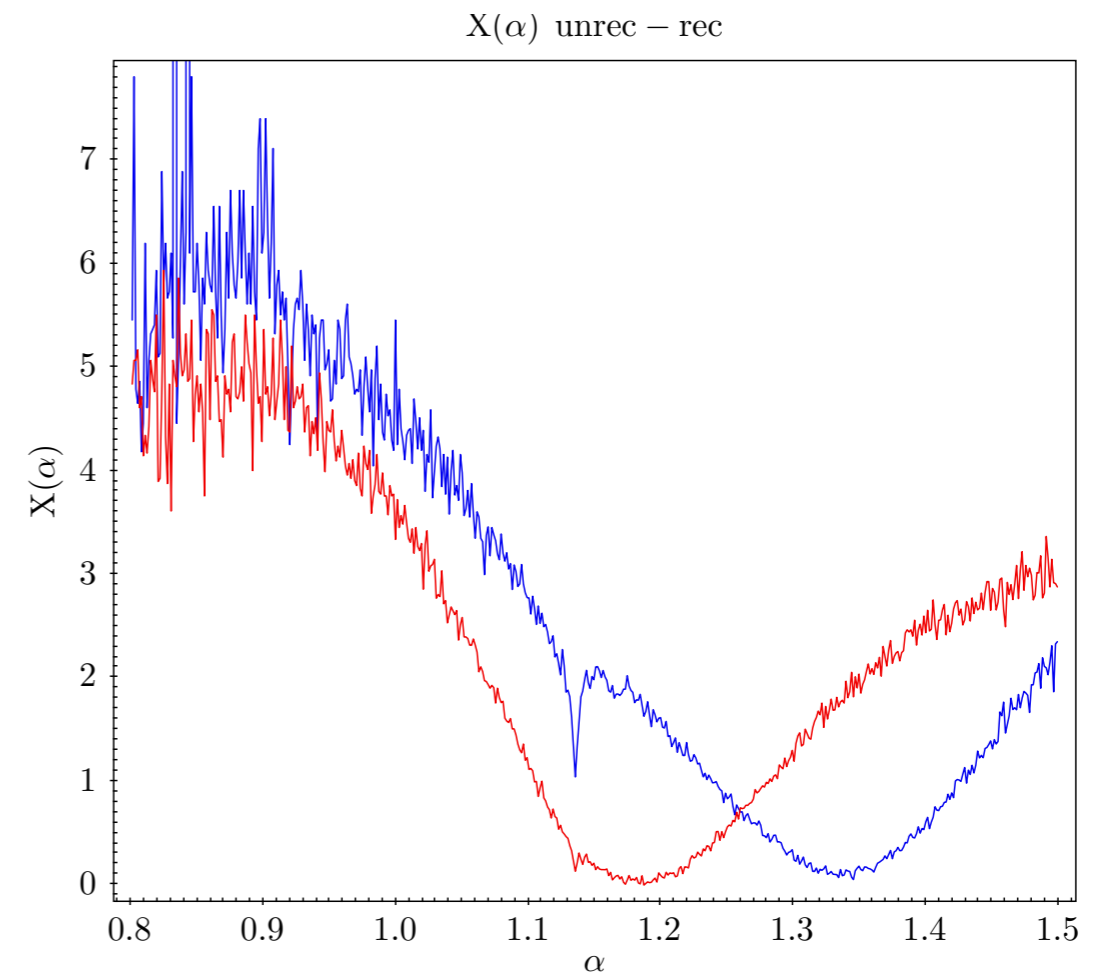
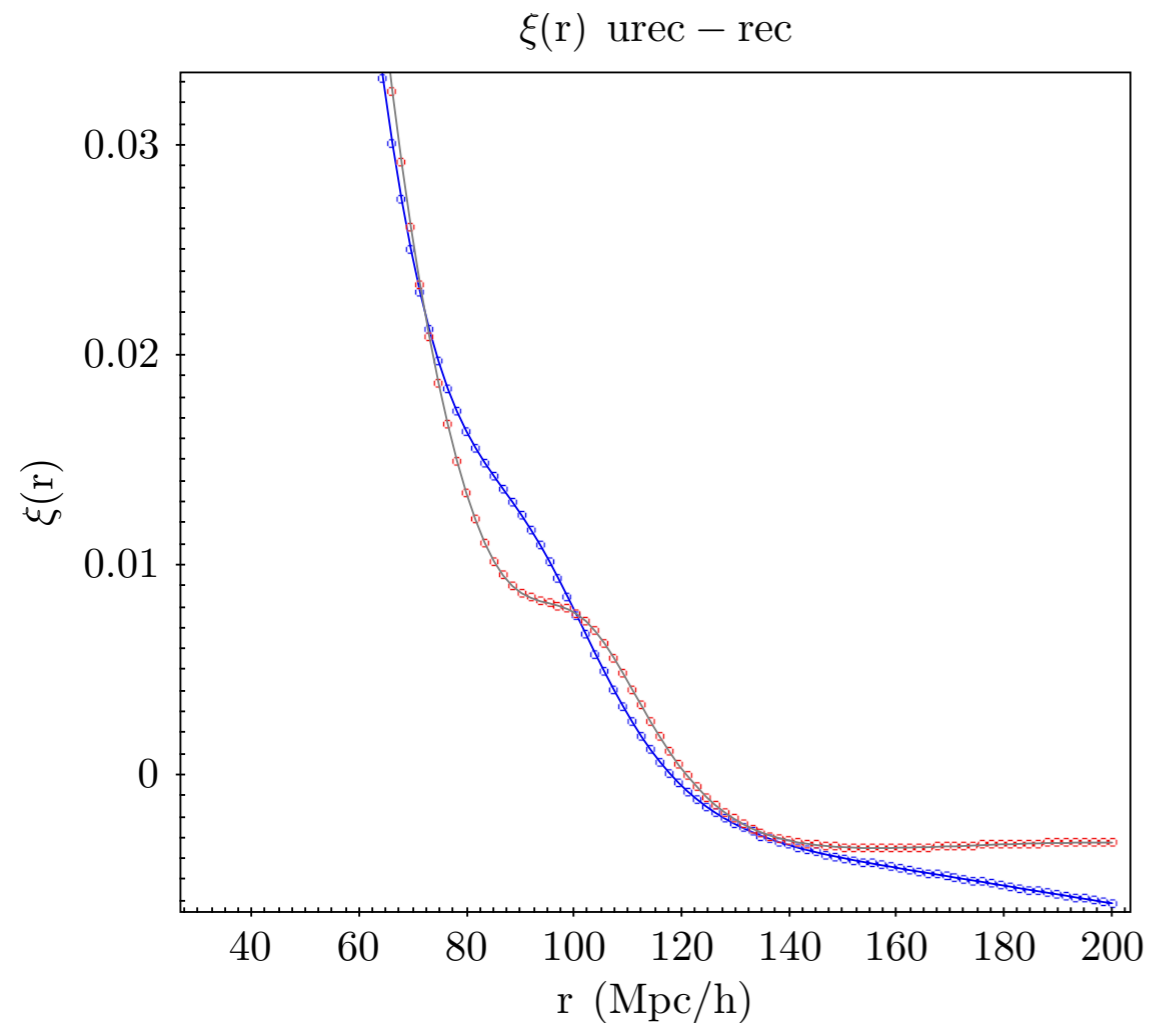
eFAM Results: fitting mocks 2p-cf



N. of mocks w/ $\sigma_\alpha > 7\%$: **69** \longrightarrow **14**

eFAM Results: fitting mocks 2pcf

Analysis of “bad-constrained” mocks

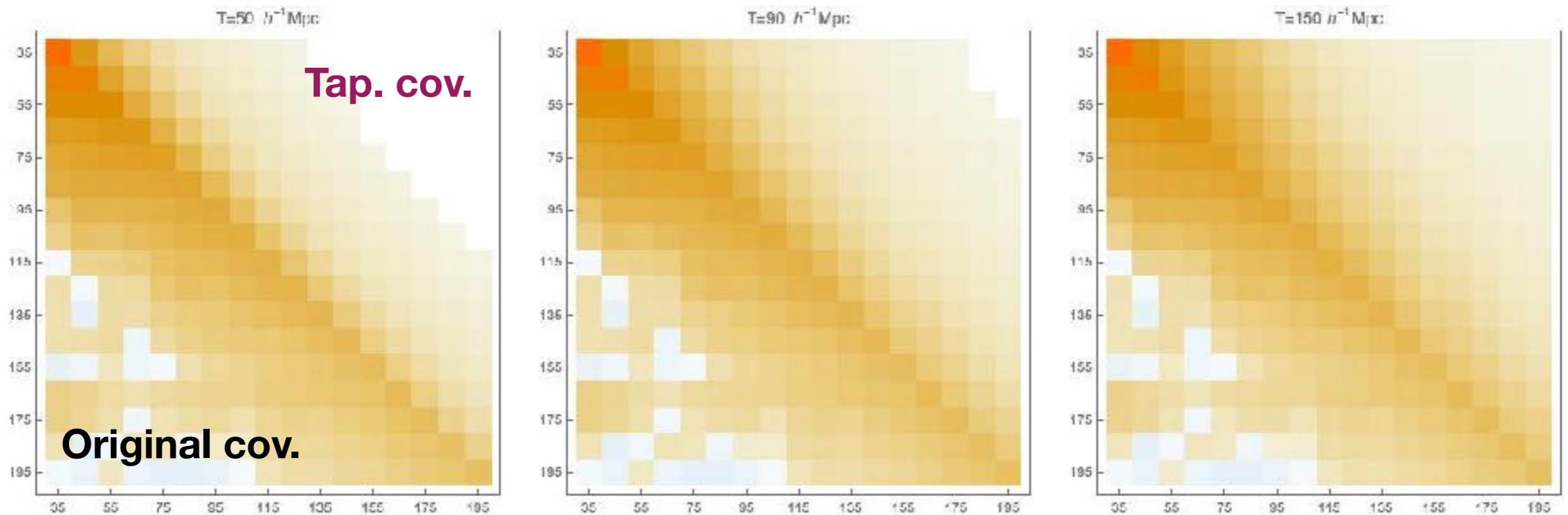


	Pre Rec	Post Rec
α	1.315	1.203
σ_α	0.092	0.090

eFAM Results: coming soon ...

Covariance matrix regularisation: **tapering**

(Paz & Sánchez 2015)



Summary

- ❖ 10^5 particles, CPU time = 8h
- ❖ Polynomial expansion at 10th order (ZA: 1st order)
- ❖ LCDM DEUS halos simulation instead of 2-LPT mocks (!!!)
- ❖ Work in progress:
redshift space computations,
fit with tapered Con matrix,
modelling od RSD

Thank you for the attention

eFAM Results: fitting mocks 2pcf

