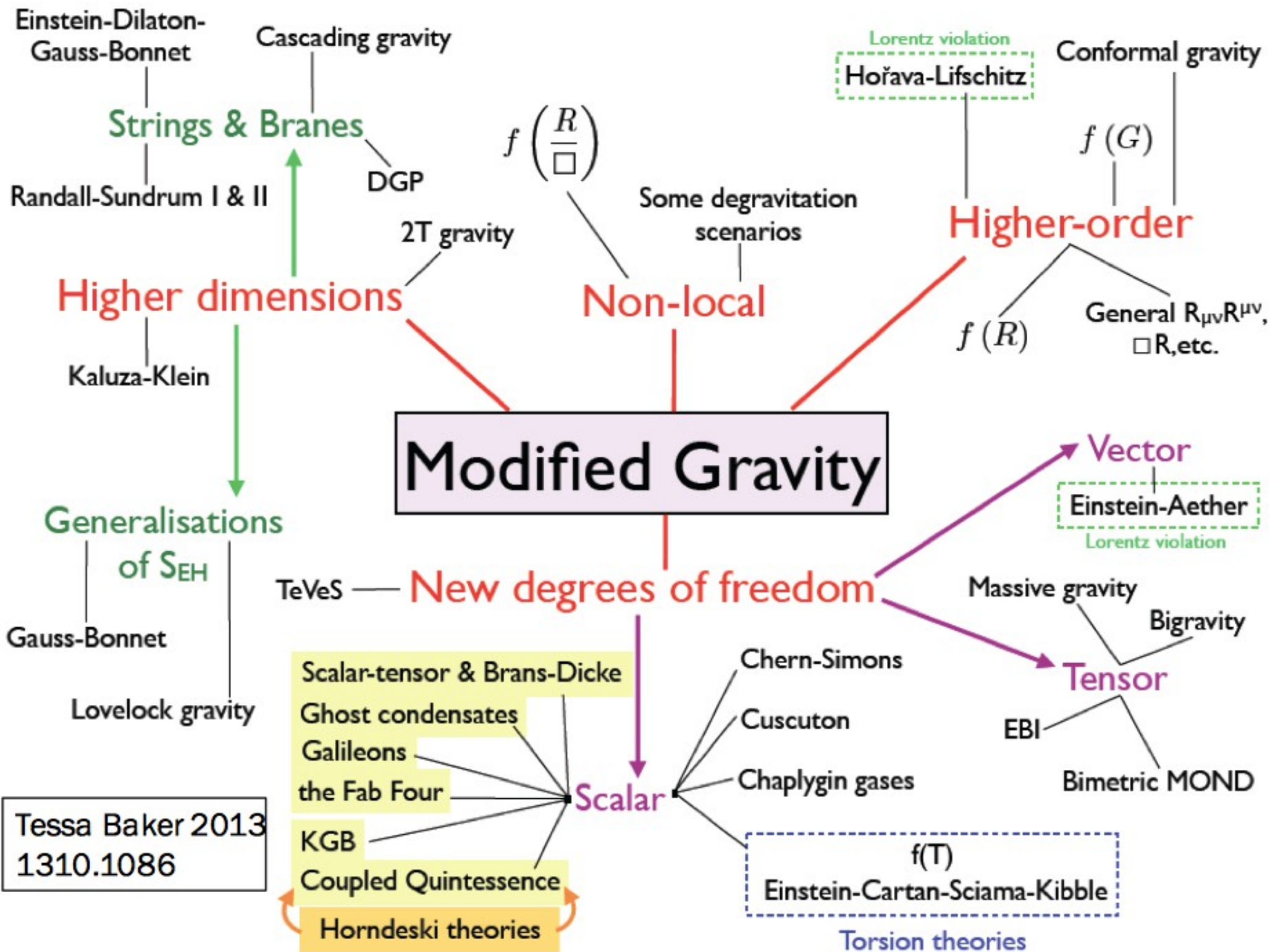


Cosmologie et Gravitation

Possible explanations for $\ddot{a} > 0$

1. It is a cosmological constant, and there is no problem (‘anthropic principle’, ‘string landscape’) – **unsatisfactory but agrees with data**
2. The (supernova) **data is wrong** – **unlikely**
3. We are making a mistake with GR (aka ‘backreaction’) or the Copernican principle is violated (‘LTB’) – **unlikely**
4. It is something evolving, e.g. a scalar field (‘dark energy’)
5. GR is wrong and needs to be modified (‘modified gravity’)

how to
characterize?

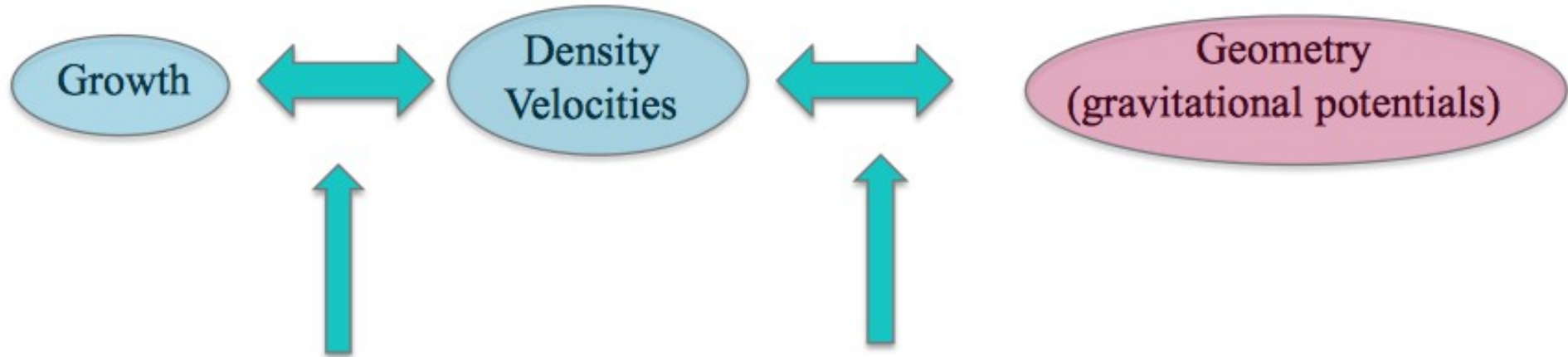




General Relativity



Modified Gravity



(conservation/Euler equation)

$$f \delta = -\frac{\nabla \cdot v}{aH}$$

(Poisson equation)

$$-k^2 \Psi \equiv 4\pi G a^2 \rho \Delta$$

Slide V. Pettorino 2017

Sondes de l'Univers

Sonde	Background Cosmology : w_0, w_a	Perturbations : $f\sigma_8, \gamma$, many others !
Supernova	SN proches et lointaines (diagramme de Hubble)	SN proches (champ de vitesses)
BAO	Pic BAO (150 Mpc / H_0)	Reshift Space Distorsions (vitesses autour des surensités)
Amas	Comptage (peu précis)	Fonction de masse (calibration difficile : X, SZ, lentille)
Clustering	3x2 pt	P(k) densité visible
Lensing	3x2 pt	P(k) potentiels

Sondes de l'Univers

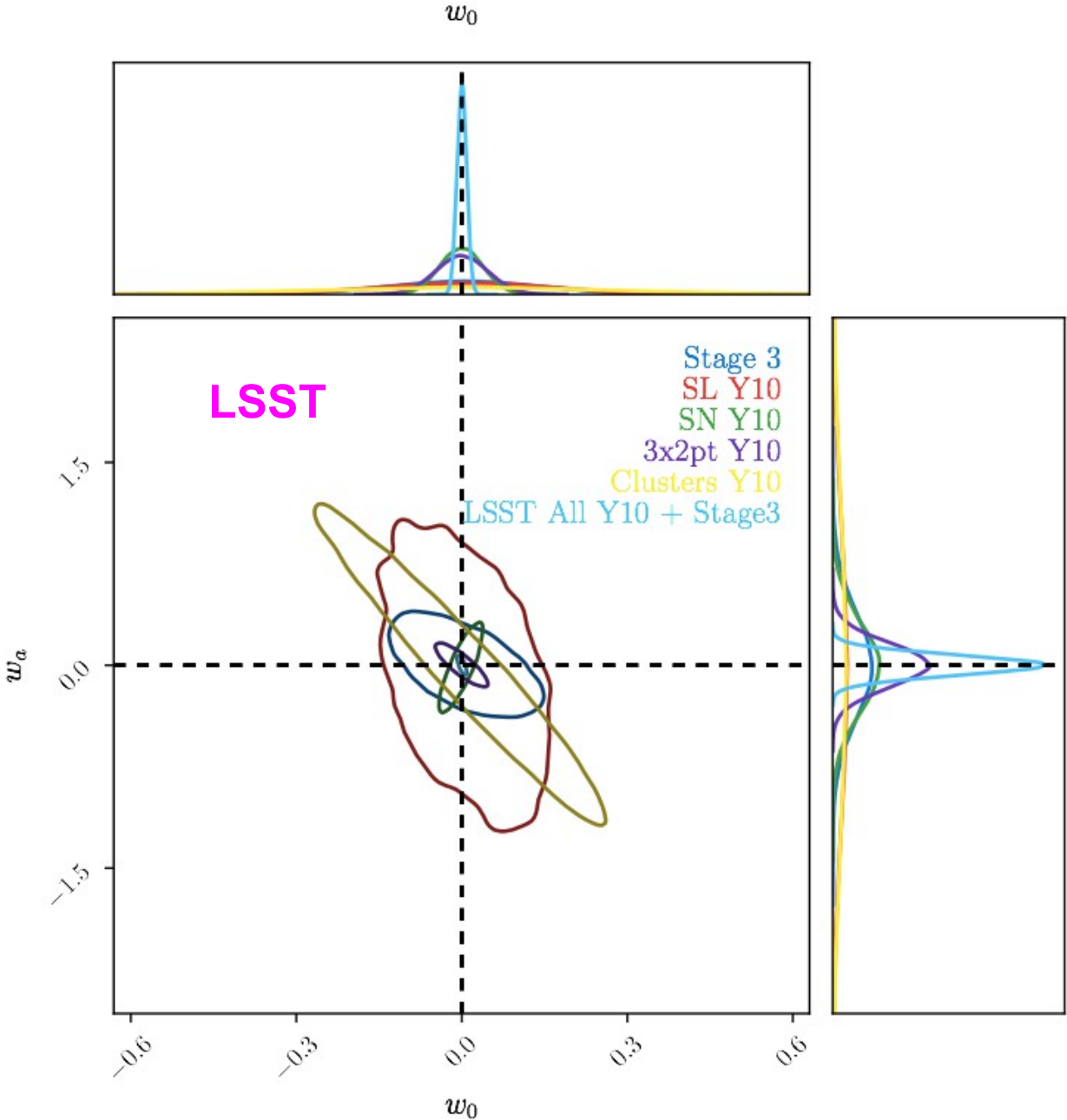
Sonde		Background Cosmology : w_0, w_a	Perturbations : $f\sigma_8, \gamma$, many others !
Supernova		SN proches et lointaines (diagramme de Hubble)	SN proches (champ de vitesses)
BAO	DESI	Pic BAO (150 Mpc / H_0)	Redshift Space Distorsions (vitesses autour des surdensités)
Amas		Comptage (peu précis) LSST	Fonction de masse (calibration difficile : X, SZ, lentille)
Clustering		3x2 pt	P(k) densité visible
Lensing		3x2 pt	P(k) potentiels EUCLID

WFIRST

SKA ?

Sondes de l'Univers

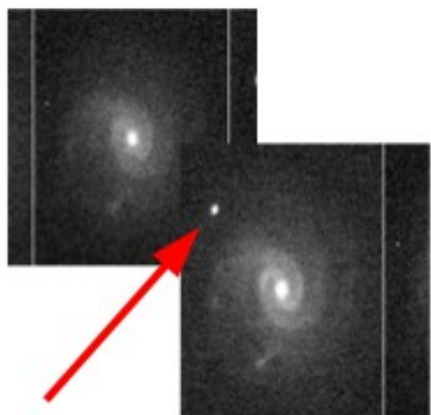
Sonde	
Supernova	
BAO	DES
Amas	
Clustering	
Lensing	



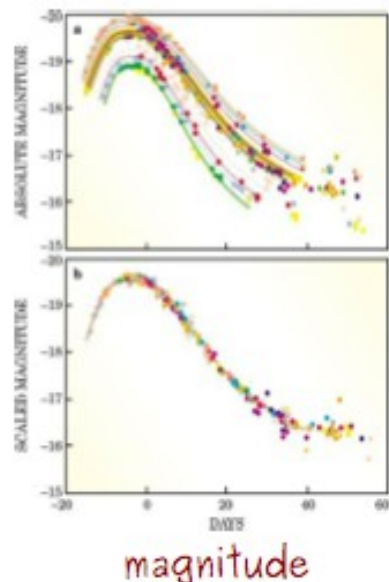
!
orsions
s
EUCLID
RST

Supernova Cosmology

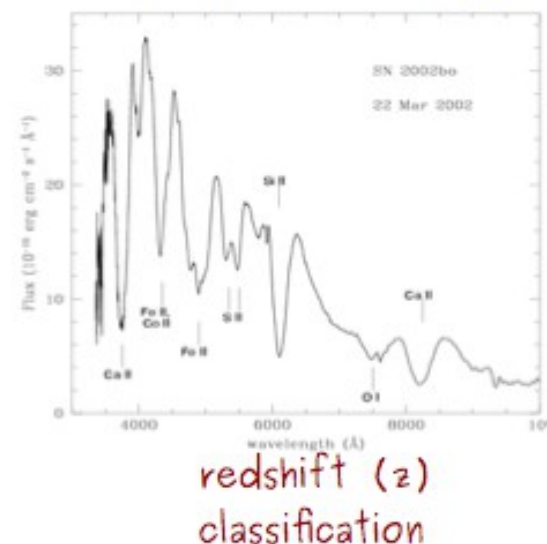
1. detection



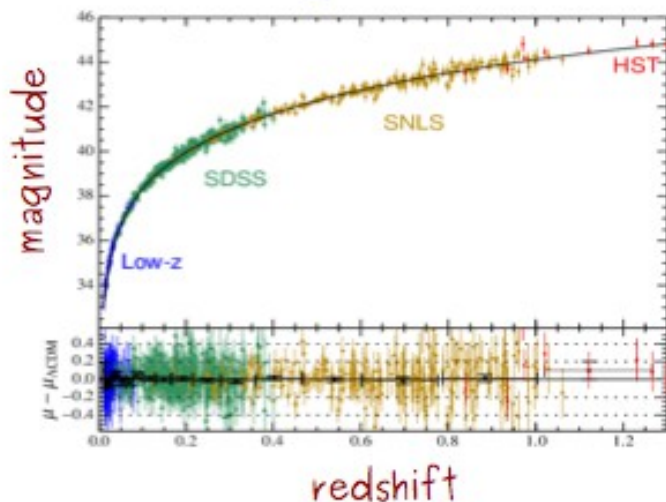
2. photometry



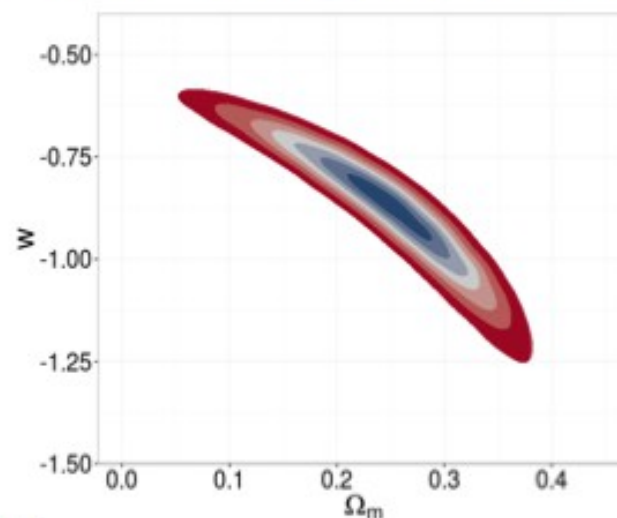
3. spectroscopy



4. standardization + cosmological fit



5. Dark Energy constraints



Supernova Cosmology

1. detection

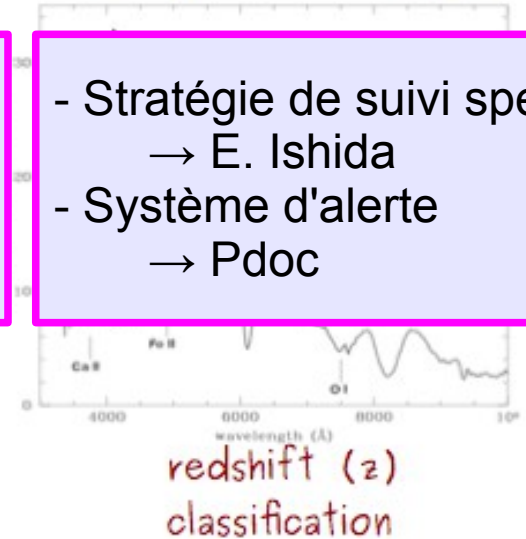
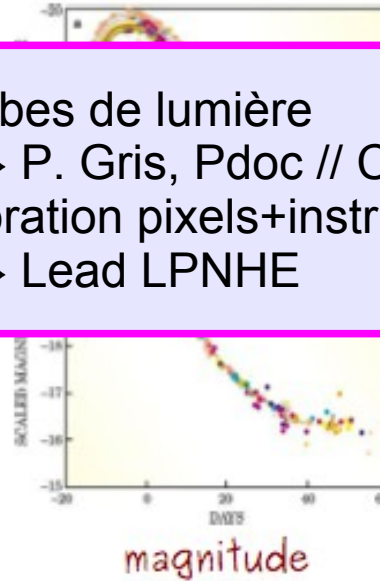
- Stratégie d'observation
→ P. Gris
- Imagerie
→ CPPM

2. photometry

- Courbes de lumière
→ P. Gris, Pdoc // CPPM
- Calibration pixels+instrument+ciel
→ Lead LPNHE

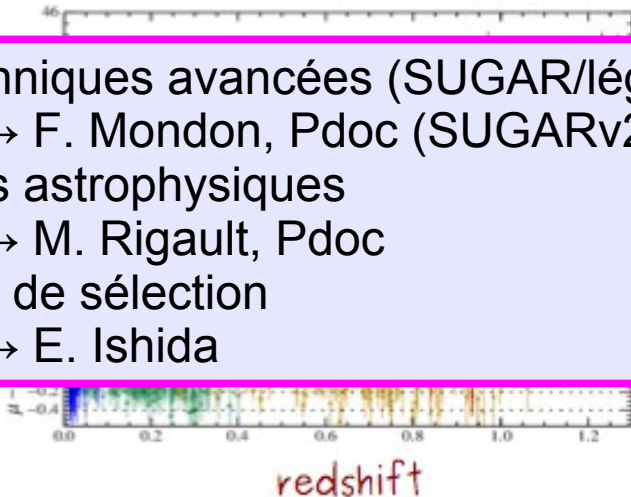
3. spectroscopy

- Stratégie de suivi spectro
→ E. Ishida
- Système d'alerte
→ Pdoc



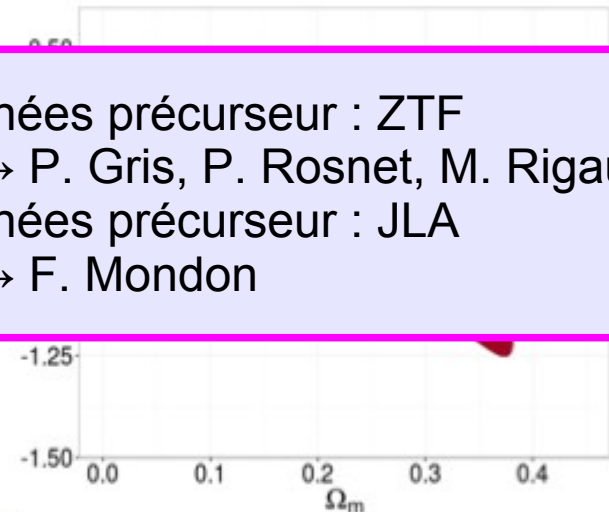
4. standardization + cosmological fit

- Techniques avancées (SUGAR/léget):
→ F. Mondon, Pdoc (SUGARv2)
- Biais astrophysiques
→ M. Rigault, Pdoc
- Bais de sélection
→ E. Ishida

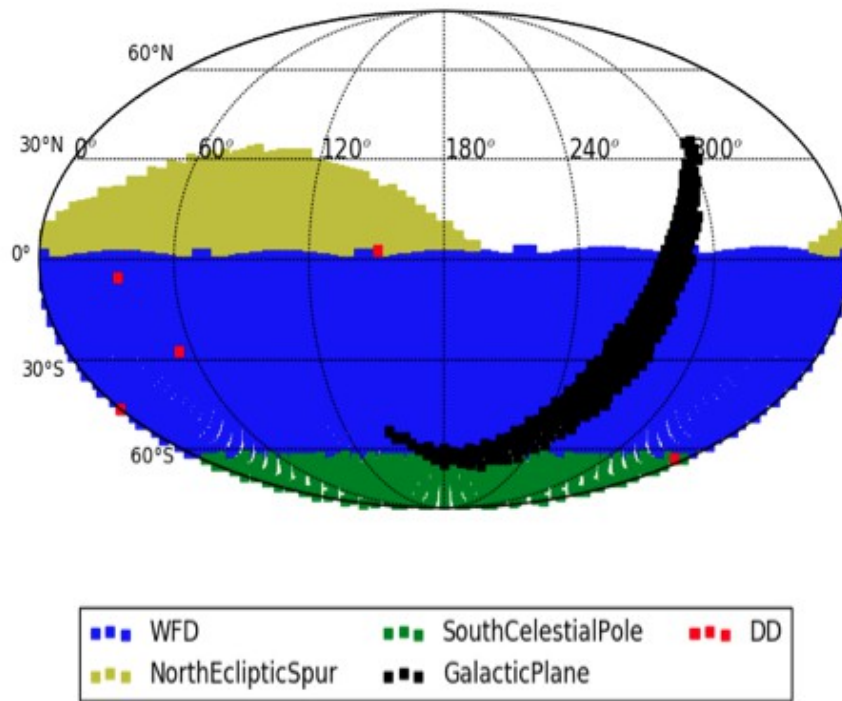


5. Dark Energy constraints

- Données précurseur : ZTF
→ P. Gris, P. Rosnet, M. Rigault, Pdoc
- Données précurseur : JLA
→ F. Mondon



LSST Observing Strategy



Current baseline (minion_1016)

Total number of visits : 2,448,282
(1 visit = 2 exposures of 15s)

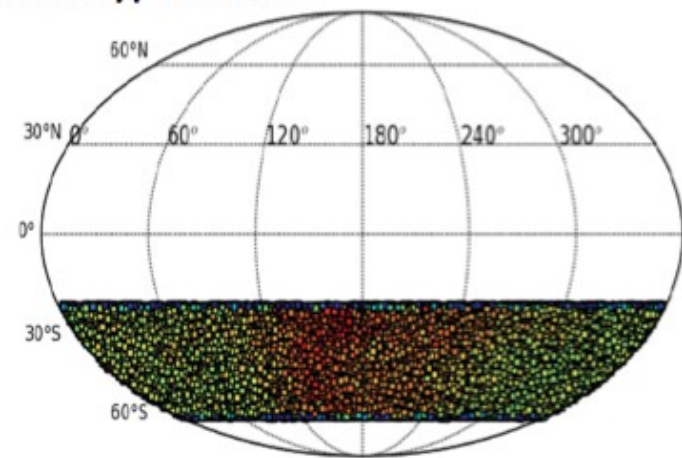
3 types of surveys:

- universal (WFD) ~ 90% -> Supernovae
- mini-surveys : ~5%
- Deep (DDF) : ~5% -> Supernovae

WFD

Current strategy : not suitable for Sne Ia science
Work in progress:

- Rolling cadence



- Number of Sne Ia ?
- Completeness ?

DDF

4 DDF -> ~ 10000 Sne Ia (10 years)
-> completeness : $z \sim 0.6$

Work in progress:

- number of DDFs (probably 8 to 10)
- synergy with other surveys (Euclid, WFIRST)
- spectroscopy measurements (4MOST, ESO-NTT, ...)

PLAsTiCC

Photometric LSST Astronomical Time-series Classification Challenge

A data challenge aimed to prepare
a larger community for the LSST data paradigm

- PI: Renee Hlozek, simulations: Rick Kessler
- SNANA simulations → Light curves in observer-frame (no images!)
- 3 years worth of LSST data, ~ 100 MB
- ~ 10^7 objects
- Around 20 transient models
(galactic and extra-galactic, periodic and non-periodic)

Expected release date:

Summer/Fall 2018

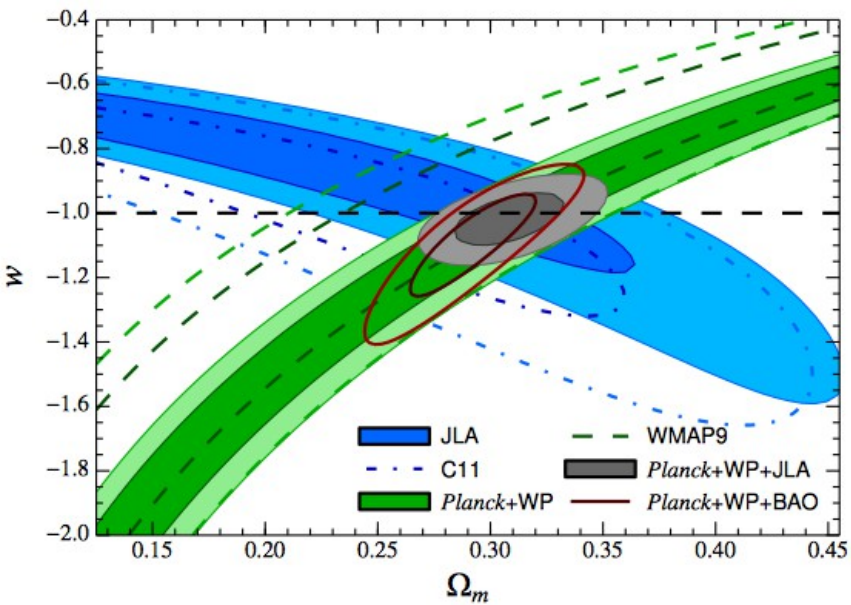
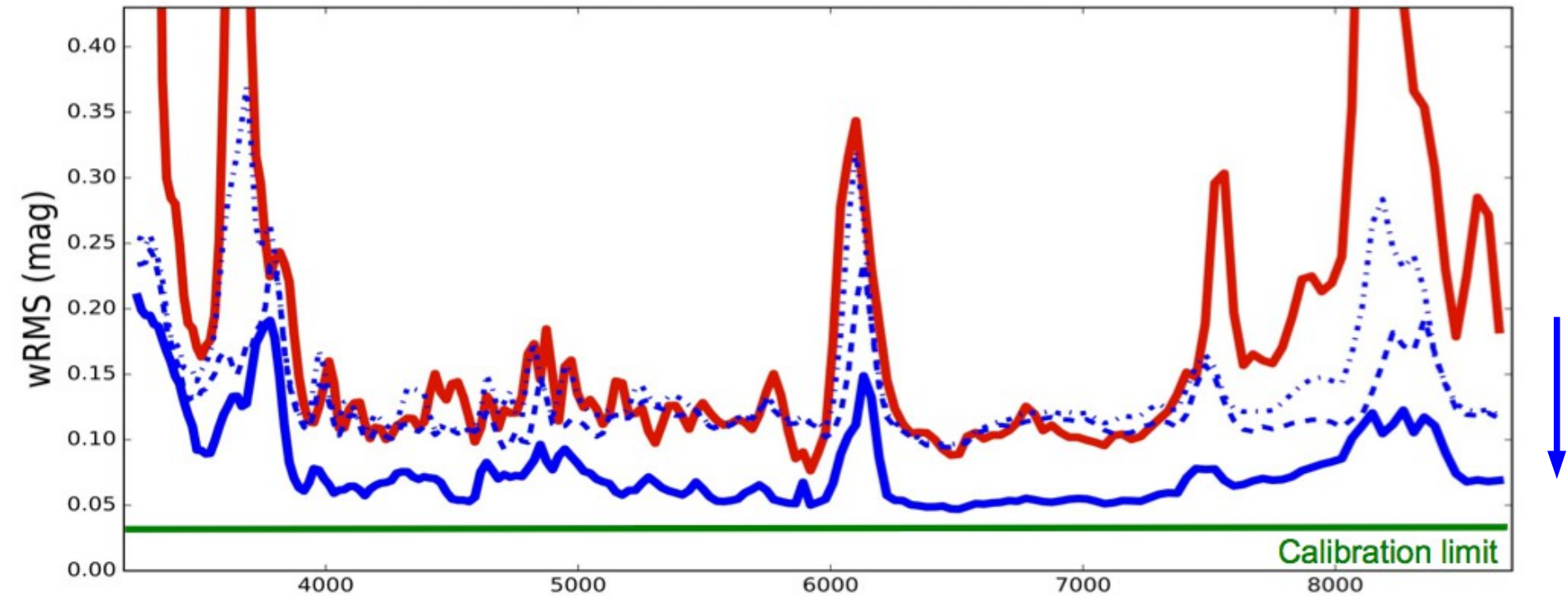
- Please respect model-information policy:
“don't ask, don't tell”



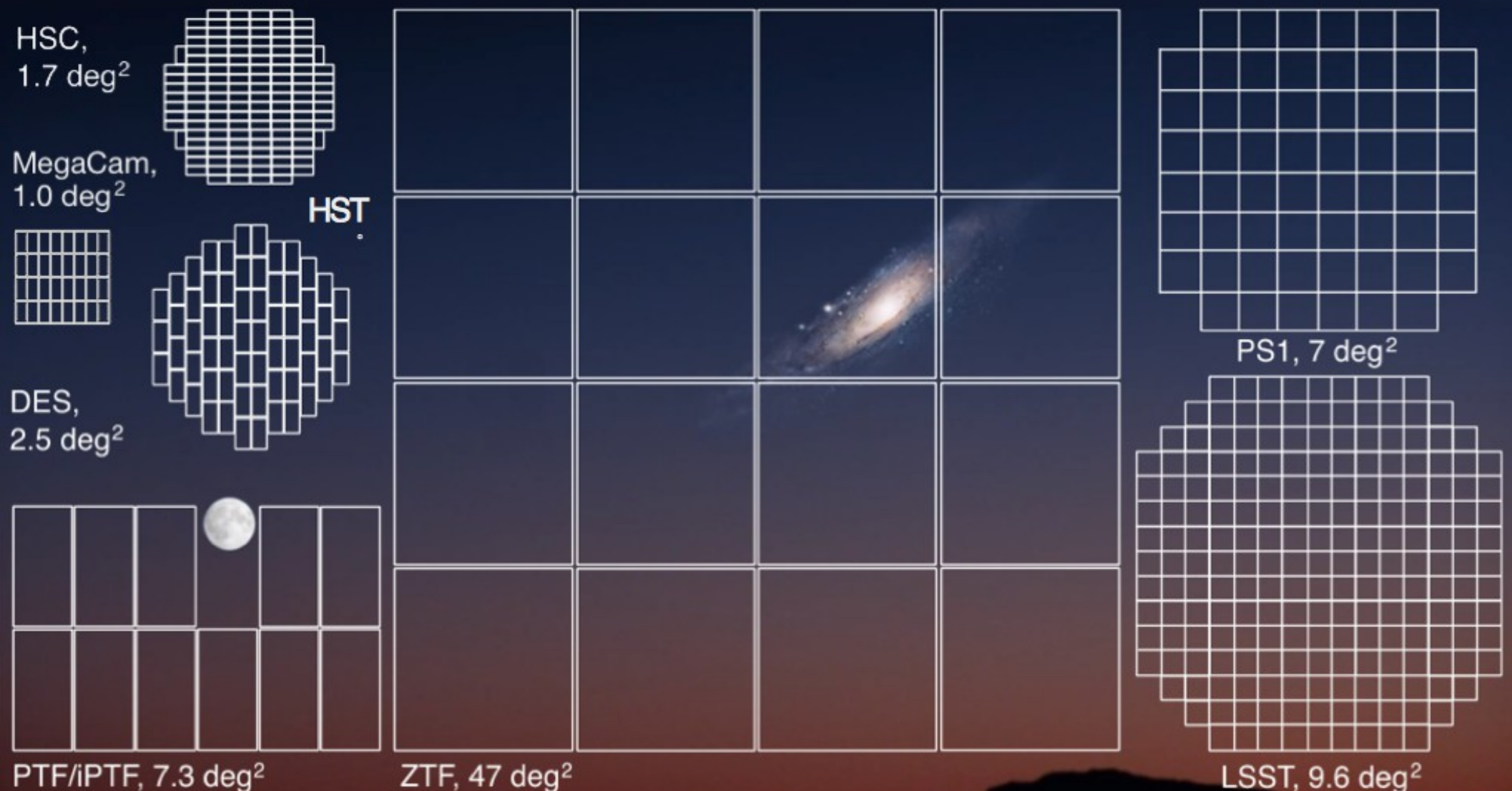
- Not all models will be present in the training sample
- Supervised classification + novelty detection
- Deployment: [kaggle](#) + [SRAMP](#)

Using SUGAR as a fitter :

Spectral reconstruction accuracy



New Surveys and Time Domain Astronomy



1990 → 2000 ~ 100 SNe Ia

2000 → 2015 ~ 10³ SNe Ia

2017 → 2030 ~ 10⁶ SNe Ia

Discovery of Dark Energy

Looks like a Cosmological Constant

Is it really?