

LATTES: simulated performance

Ruben Conceição

on behalf of the LATTES team

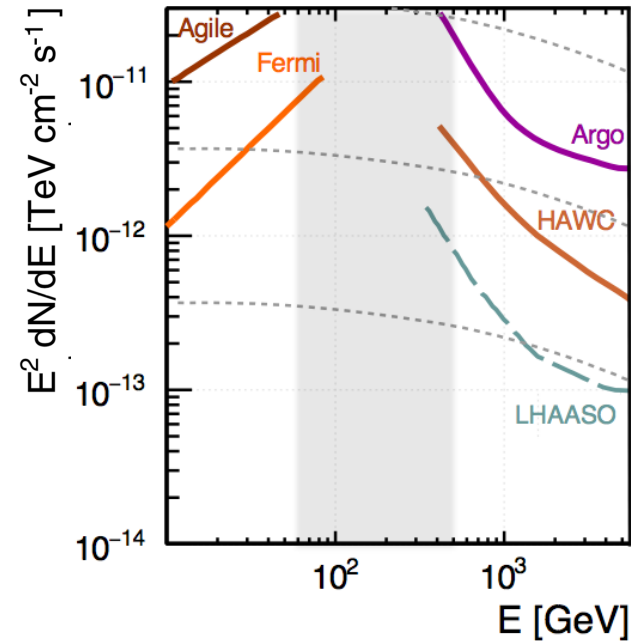
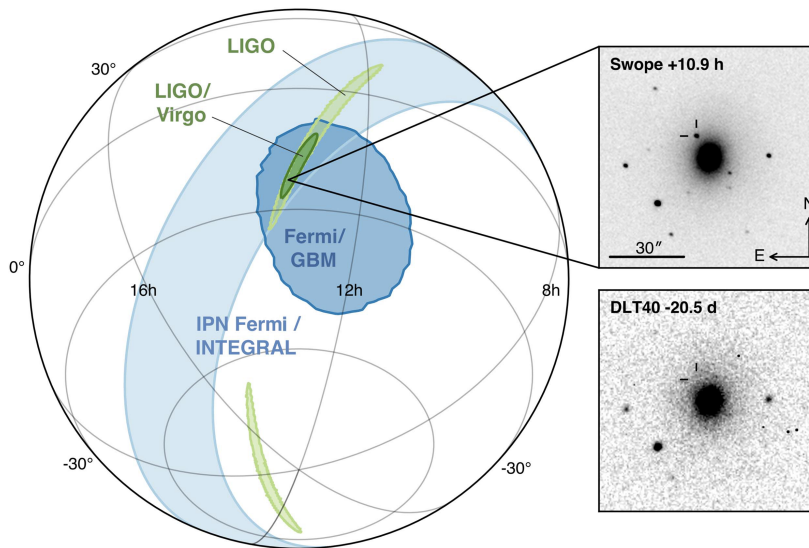


The era of **multi-messenger** observations



Joint publication of LIGO, VIRGO, INTEGRAL, Fermi, IceCube, Pierre Auger ...

Current wide FoV gamma-ray observatories



- ❖ Simultaneous observation of a Gravitational Wave + electromagnetic counterparts
- ❖ Study of transient phenomena in all energy windows is one of the main ingredients

Simulation Framework

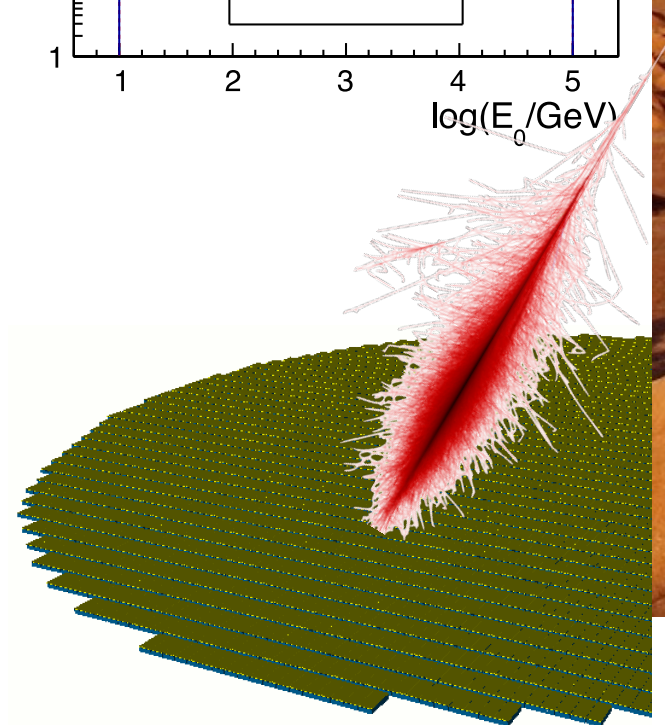
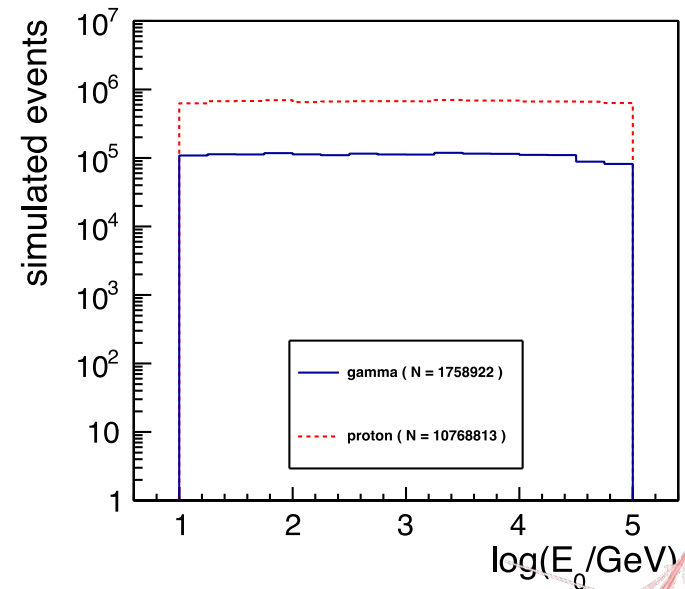
✧ End-to-end realistic simulation

✧ Extensive Air Showers: **CORSIKA**

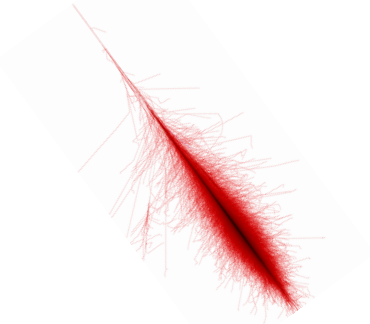
- ✧ v7.6400 with Fluka2011.2c
- ✧ More than 100 000 gamma/proton shower simulated randomly between 10 GeV - 300 TeV
- ✧ Gammas have a fixed zenith angle of 10 degrees
- ✧ Observation level at 5200 m of altitude

✧ Detector simulation: **Geant4**

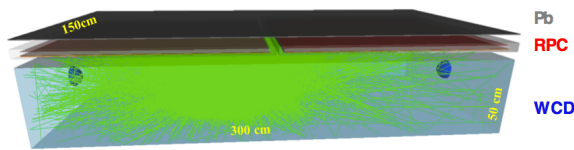
- ✧ v10.1.3
- ✧ Core array 20 000 m²
- ✧ Each shower is resampled 100 times over a big area containing all the array



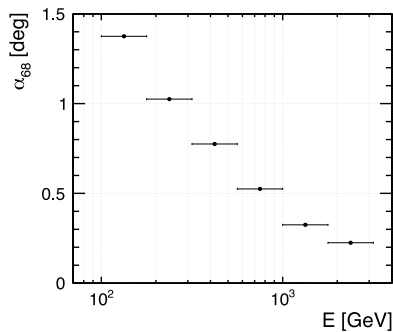
Towards LATTES sensitivity...



Shower simulation
(CORSIKA)



Detector simulation
(Geant4)

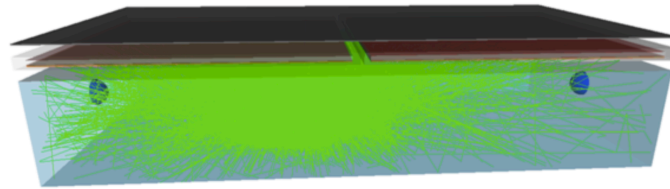


Shower reconstruction
(LATTESrec)

Simulation Framework

✧ Reconstruction

- ✧ **First order analyses** with little optimization only to demonstrate principle



✧ Performance and sensitivity

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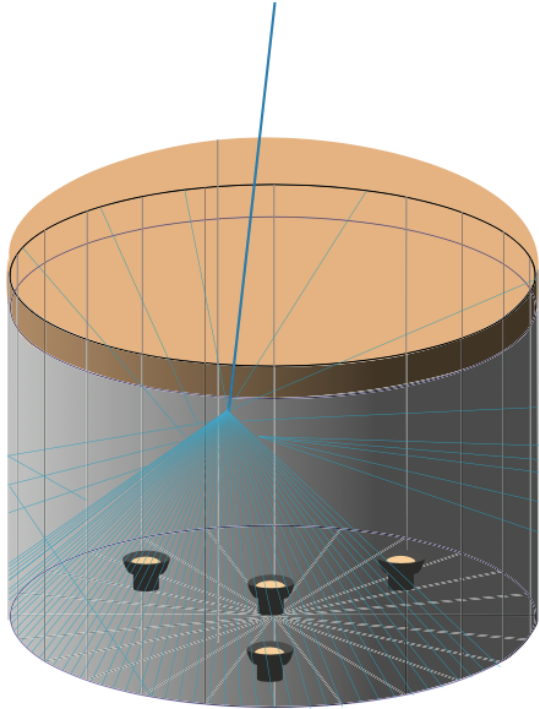


Design and expected performance of a novel hybrid detector for very-high-energy gamma-ray astrophysics

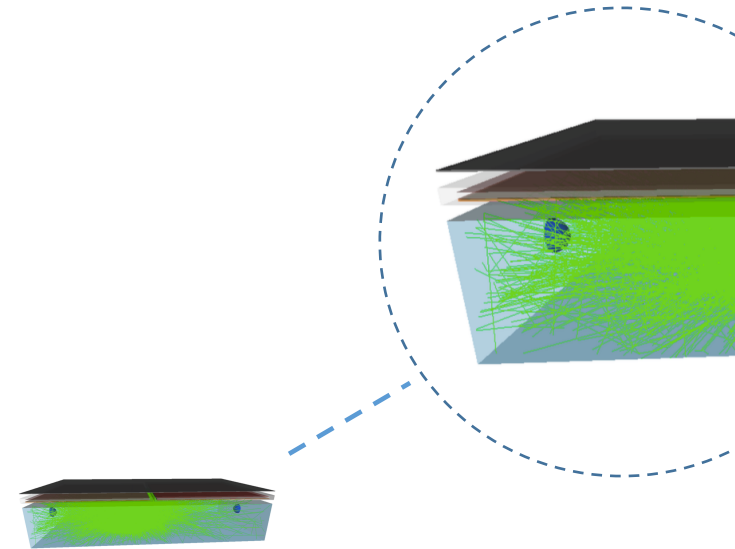
P. Assis^{a,b}, U. Barres de Almeida^c, A. Blanco^d, R. Conceição^{a,b,*}, B. D'Etorre Piazzoli^e,
A. De Angelis^{f,g,b,a}, M. Doro^{h,f}, P. Fonte^d, L. Lopes^d, G. Matthiaeⁱ, M. Pimenta^{b,a}, R. Shellard^c,
B. Tomé^{a,b}



Detector Station



HAWC



LATTES

LATTES expected performance

- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

LATTES expected performance

❖ **Trigger and effective area**

❖ Core reconstruction

❖ Energy reconstruction

❖ Geometry reconstruction

❖ Gamma/hadron discrimination

❖ Sensitivity to steady sources



Effective Area
depends with
quality cuts

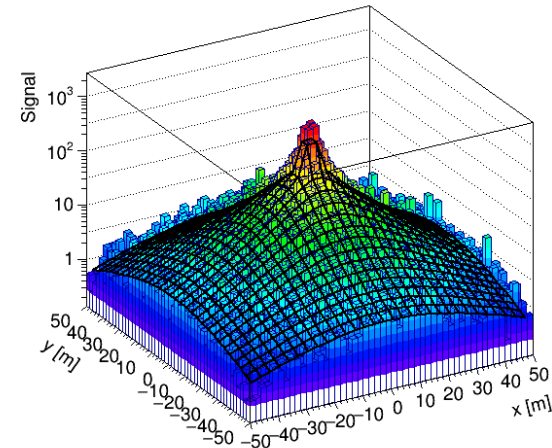
LATTES expected performance

- ✧ Trigger and effective area
- ✧ **Core reconstruction**
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

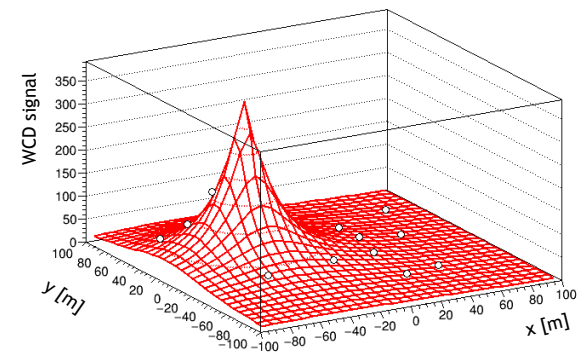
Shower core reconstruction

- ✧ Use the WCD signal
- ✧ Barycenter
 - ✧ Initial guess
 - ✧ Works but the core is always reconstructed inside the array
- ✧ Fit the WCD LDF
 - ✧ Fit photon average LDF to fix the shape
 - ✧ Function inspired in HAWC
 - ✧ Nearly no evolution with energy
 - ✧ Use this form to find the maximum, i.e. the shower core

Average LDF



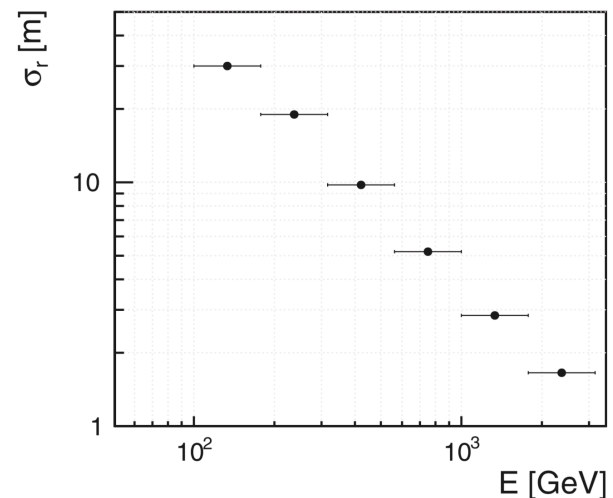
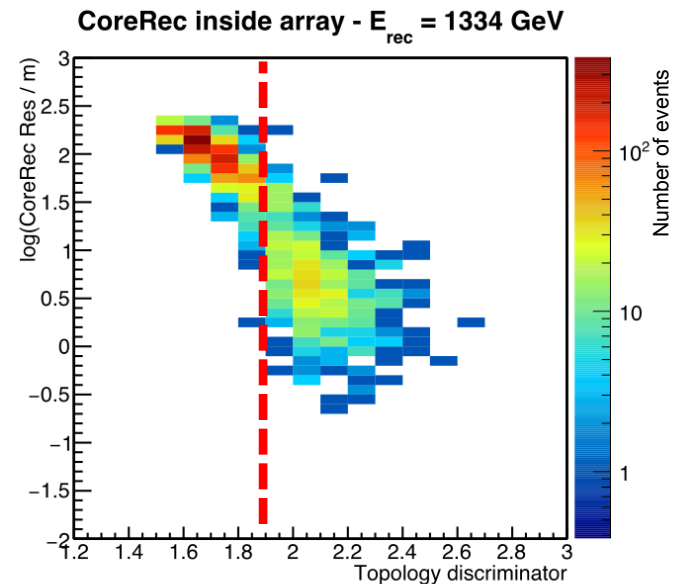
Single event



$$S_i = S(A, \vec{x}, \vec{x}_i) = A \left(\frac{1}{2\pi\sigma^2} e^{-|\vec{x}_i - \vec{x}|^2 / 2\sigma^2} + \frac{N}{(0.5 + |\vec{x}_i - \vec{x}| / R_m)^3} \right)$$

Shower core reconstruction

- ✧ Test whether the shower is inside/outside the array
 - ✧ Explore LDF topology
 - ✧ Is maximum observed inside of array?
 - ✧ Currently exploring the quality of the fit
 - ✧ Fixed cut for all energies
- ✧ Resolution better than 10 meters for showers above 300 GeV



LATTES expected performance

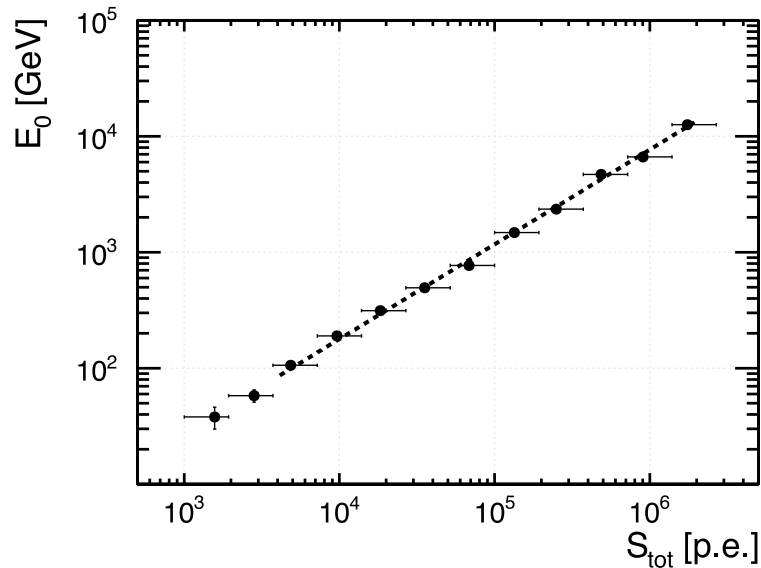
- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ **Energy reconstruction**
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

Energy reconstruction

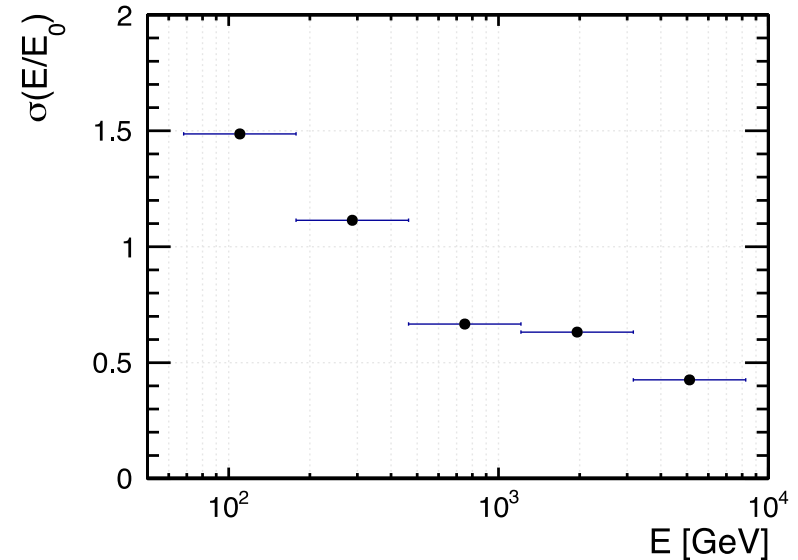
E_0 → Simulated energy

E → Reconstructed energy

Energy Calibration

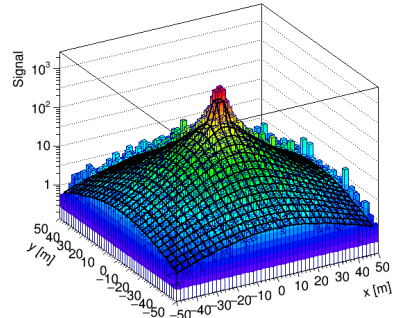
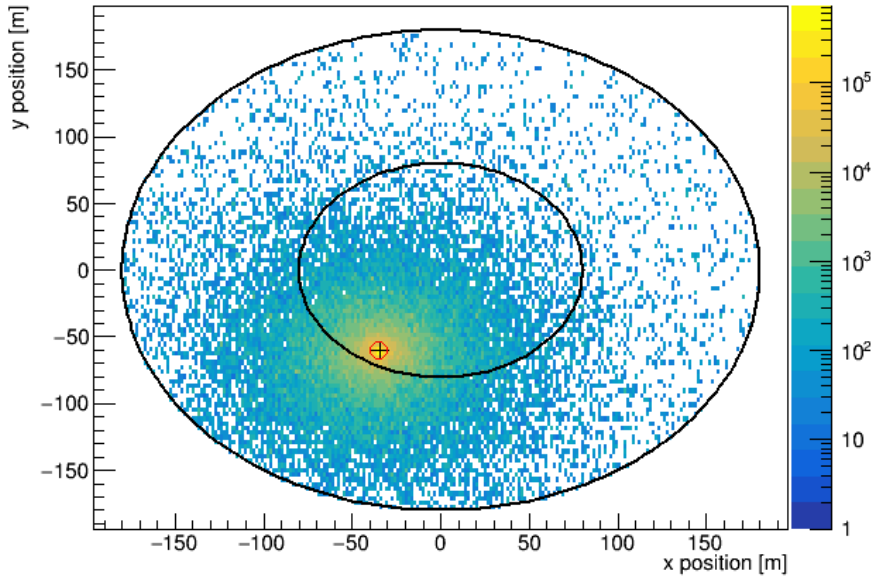


Energy Resolution

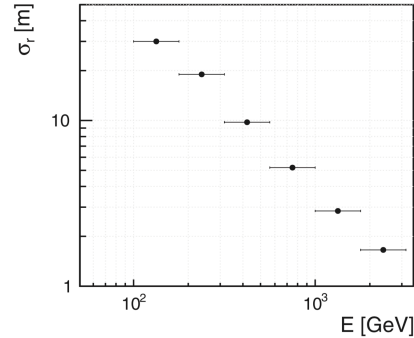


- ✧ Use as **energy estimator** the **total signal** recorded by **WCDs**
 - ✧ Use only shower cores reconstructed inside array
- ✧ Energy resolution at low energy dominated by shower fluctuations

Towards a more sophisticated energy reconstruction



Average LDF



Core resolution

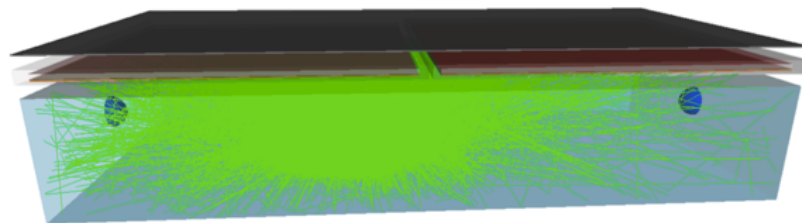
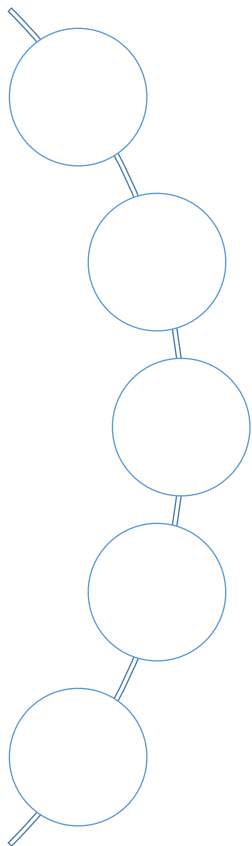
❖ Combine the core position with an average LDF to estimate the amount of energy outside of the array

LATTES expected performance

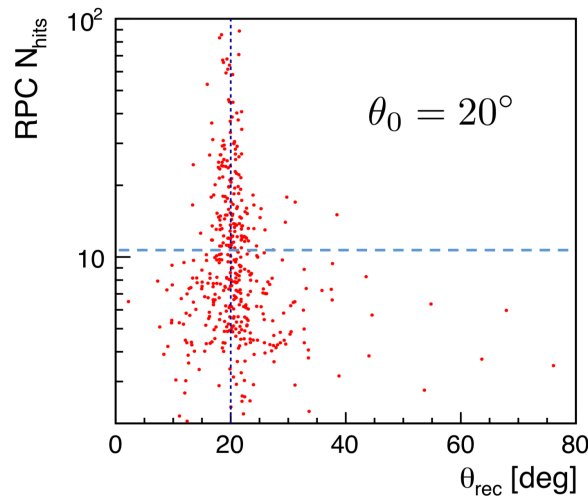
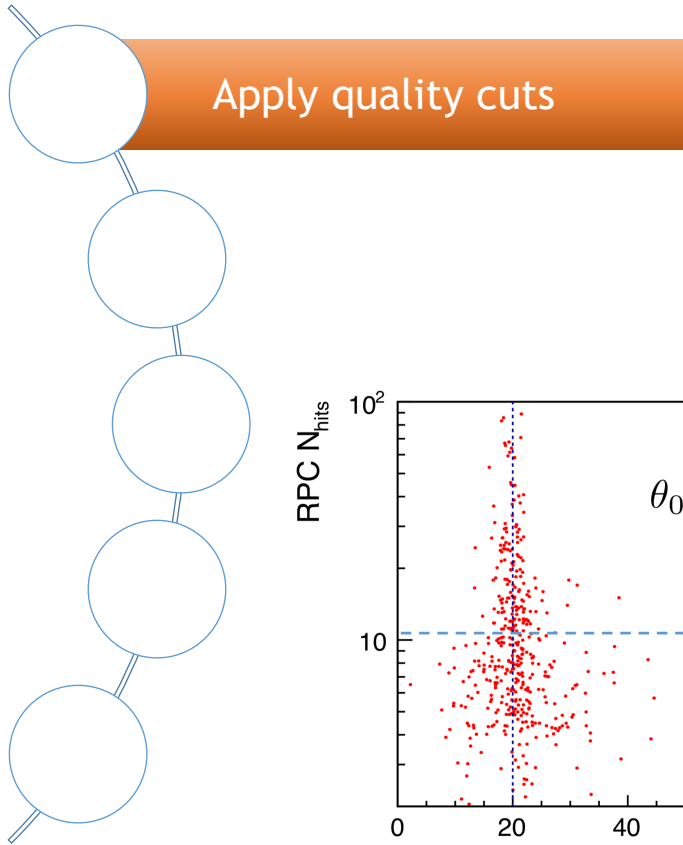
- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ **Geometry reconstruction**
- ✧ Gamma/hadron discrimination
- ✧ Sensitivity to steady sources

Reconstruction of shower geometry

- ✧ Use **RPC hit time** information
 - ✧ Take advantage of high spatial and time resolution
 - ✧ Used time resolution of 1 ns

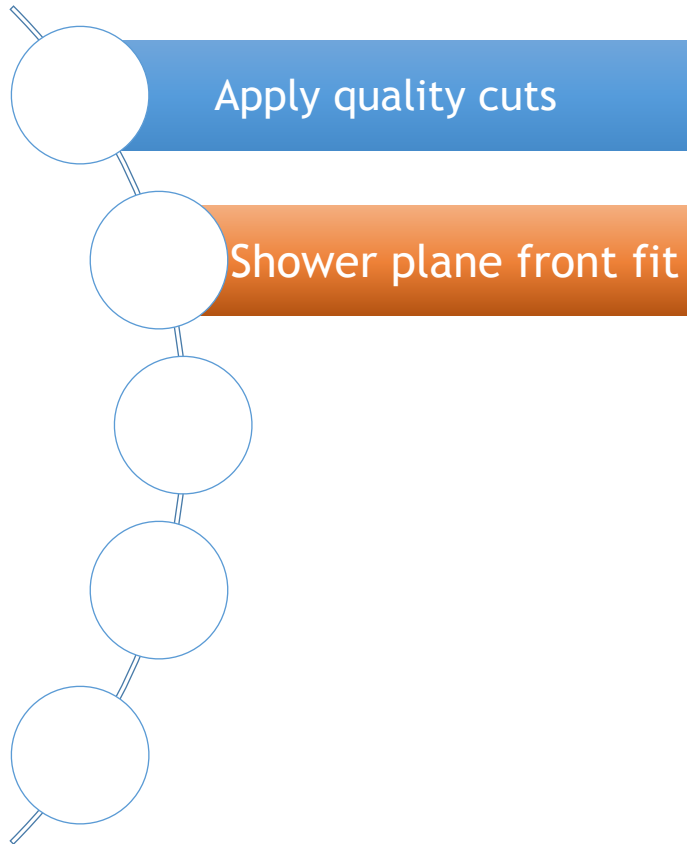


Reconstruction of shower geometry

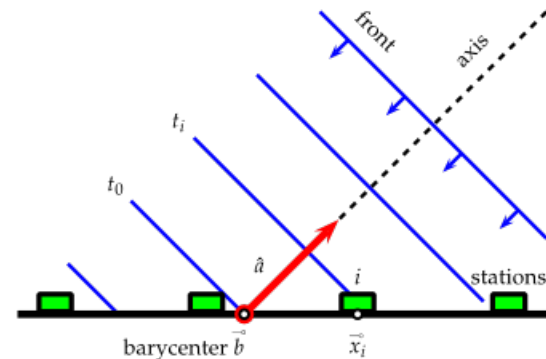


- ✧ Use **RPC hit time** information
 - ✧ Apply previous shower rec quality cuts
 - ✧ Apply cuts on the number of registered hits on the RPCs
 - ✧ Consider only RPCs in triggered WCD stations

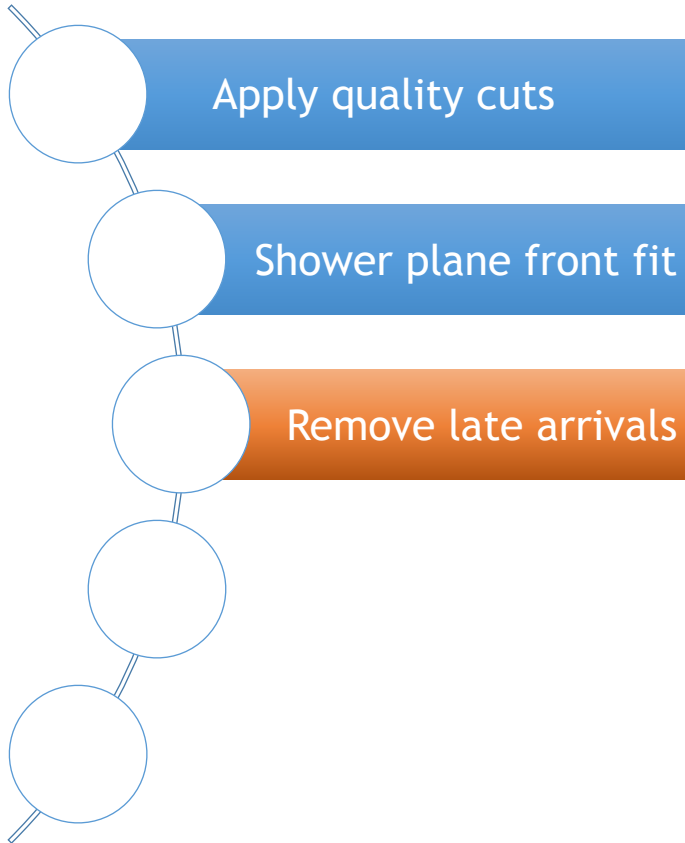
Reconstruction of shower geometry



- ✧ Use **RPC hit time** information
- ✧ Perform shower reconstruction
- ✧ Use shower front plane approximation
- ✧ Analytical procedure



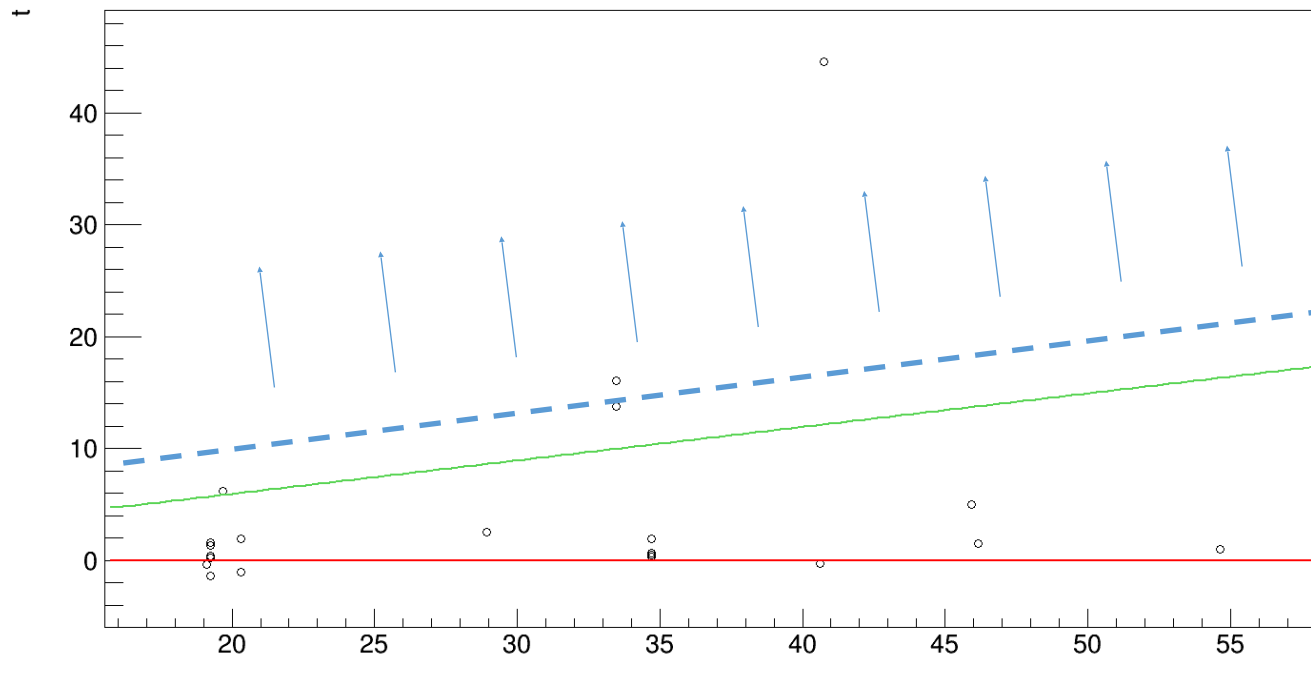
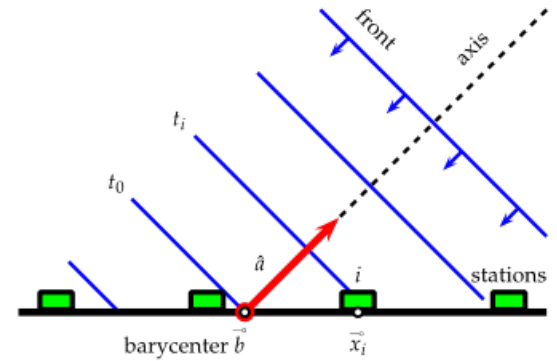
Reconstruction of shower geometry



- ✧ Use **RPC hit time** information
 - ✧ Identify late arrivals with respect to Rec Shower Front
 - ✧ Mainly low energy electrons that lost correlation with shower front

Removal of late arrivals

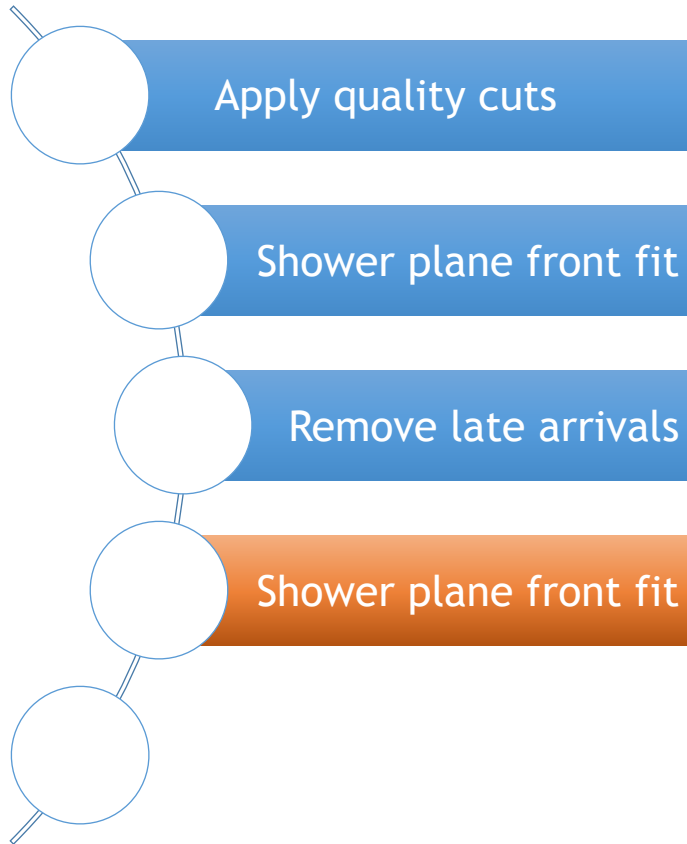
- ✧ Example of a vertical gamma shower
- ✧ Plot depicts arrival time (ns) distance to simulated shower core (m)



$$\theta_{rec} = 5.1^\circ$$

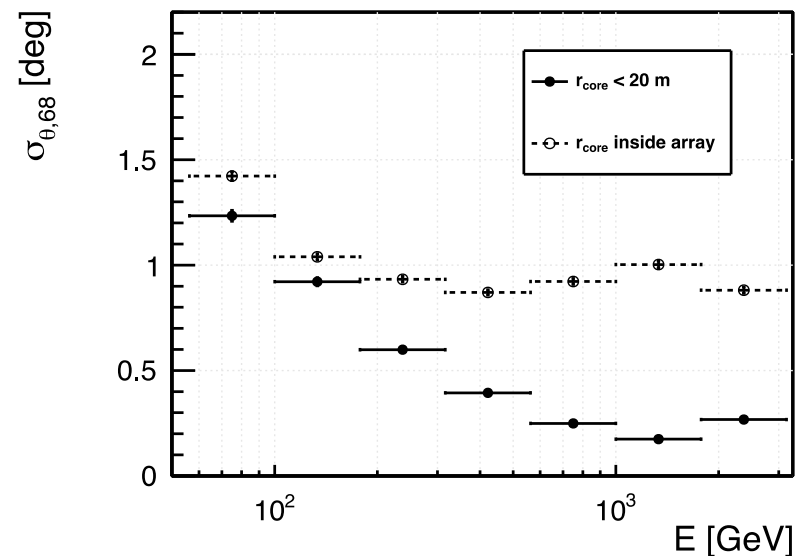
$$\theta_0 = 0^\circ$$

Reconstruction of shower geometry

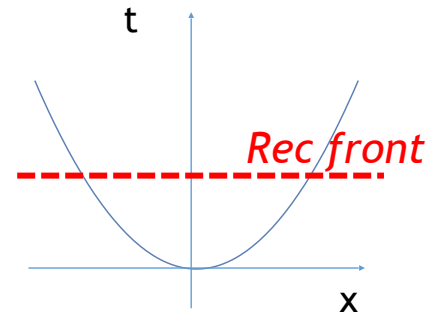
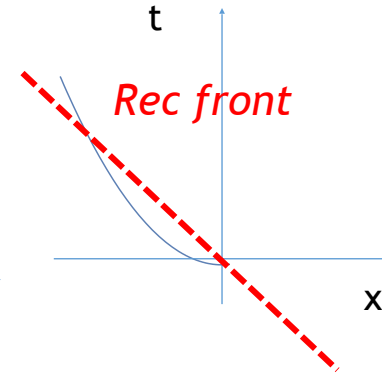
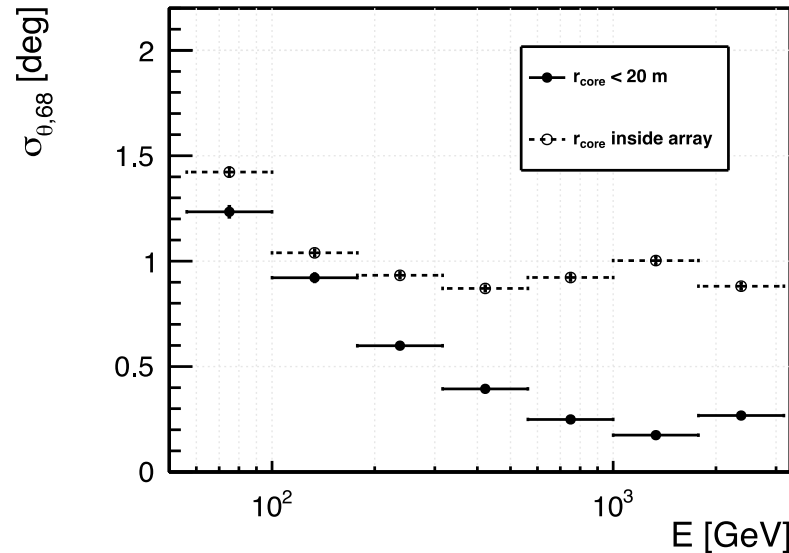


- ✧ Use **RPC hit time** information
 - ✧ Repeat fit without arrivals
 - ✧ Initial guess for next step

γ – showers; $\theta = 10^\circ$



Impact of shower curvature

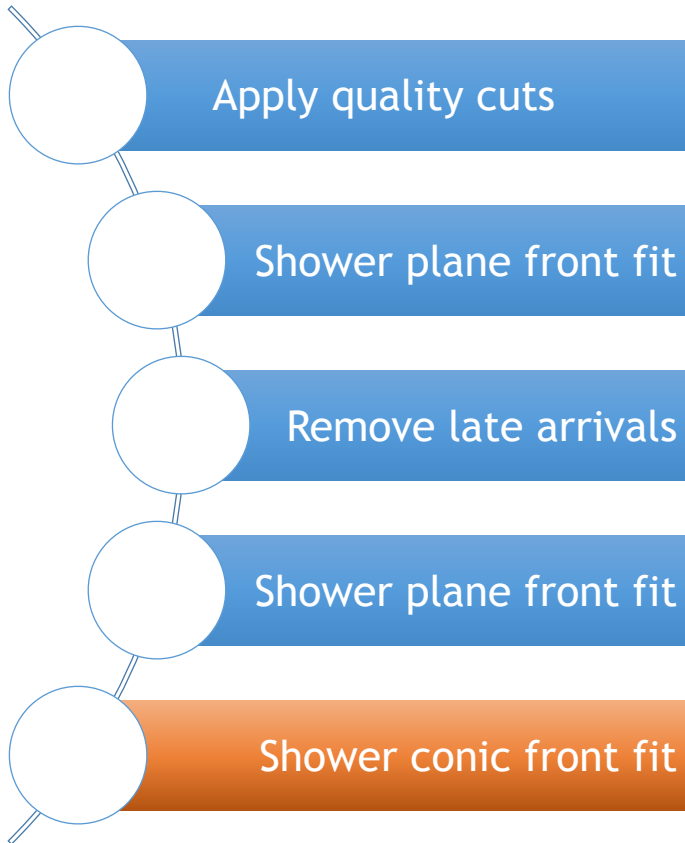


Center of the array Border of the array

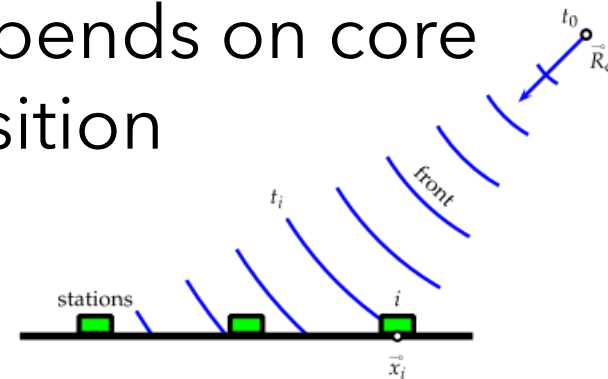
Solution: implement a conic fit instead of fitting a plane

$$\chi^2 = \sum (c \cdot (T_n - T_0) - X_n \cdot -Y_n \cdot m - R_n \cdot \alpha)^2$$

Reconstruction of shower geometry

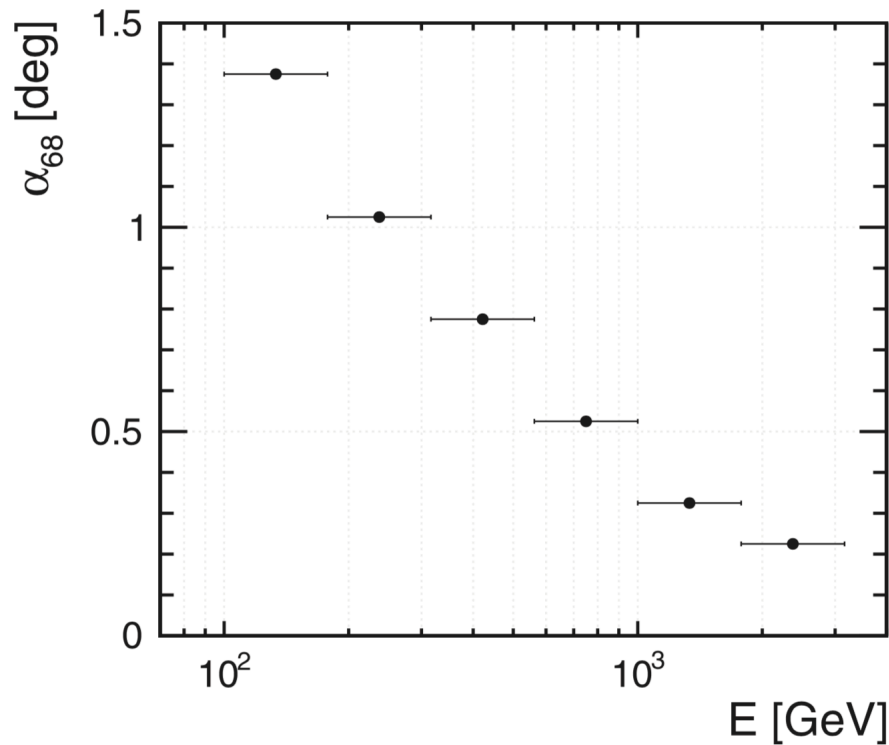


- ✧ Use **RPC hit time** information
- ✧ Fit the shower geometry using a shower conic front model
- ✧ Depends on core position



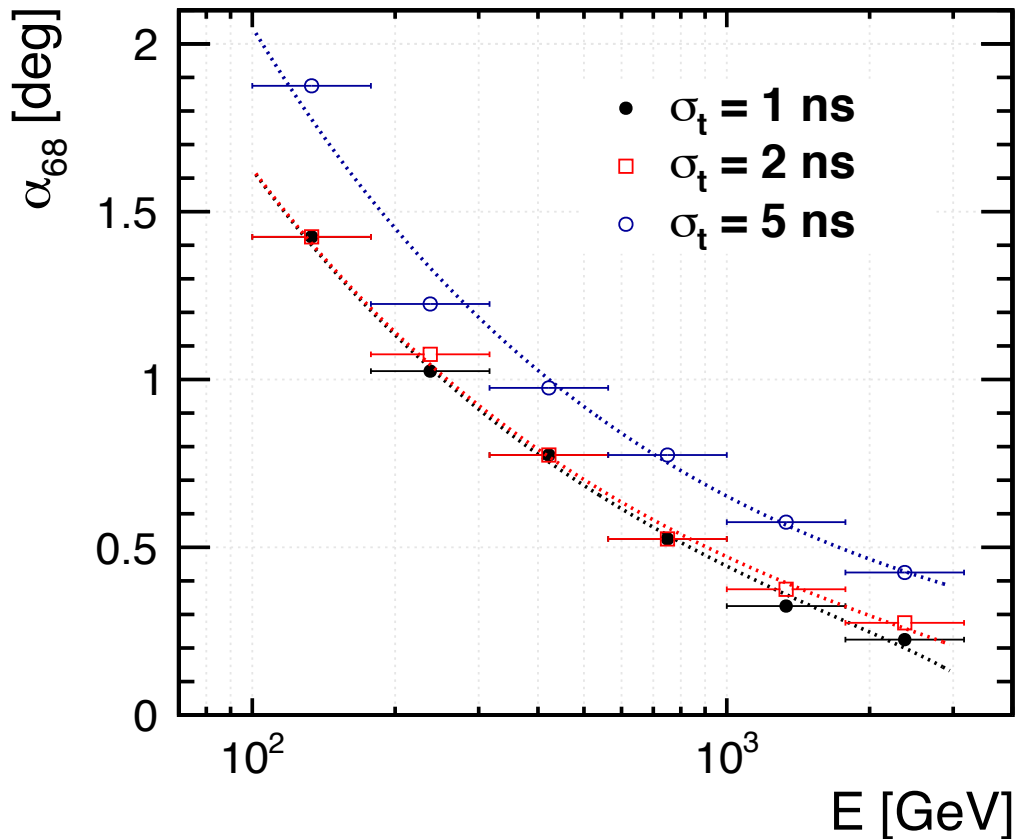
Shower geometry reconstruction

α_{68} = angular distance that contains 68% of the events



A good angular resolution can be achieved for all events reconstructed inside the array

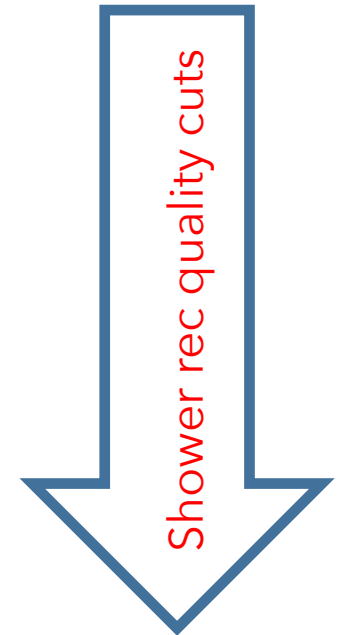
Geom Rec: RPC time resolution



Need of a time resolution of 1-2 ns to obtain a good geometry reconstruction

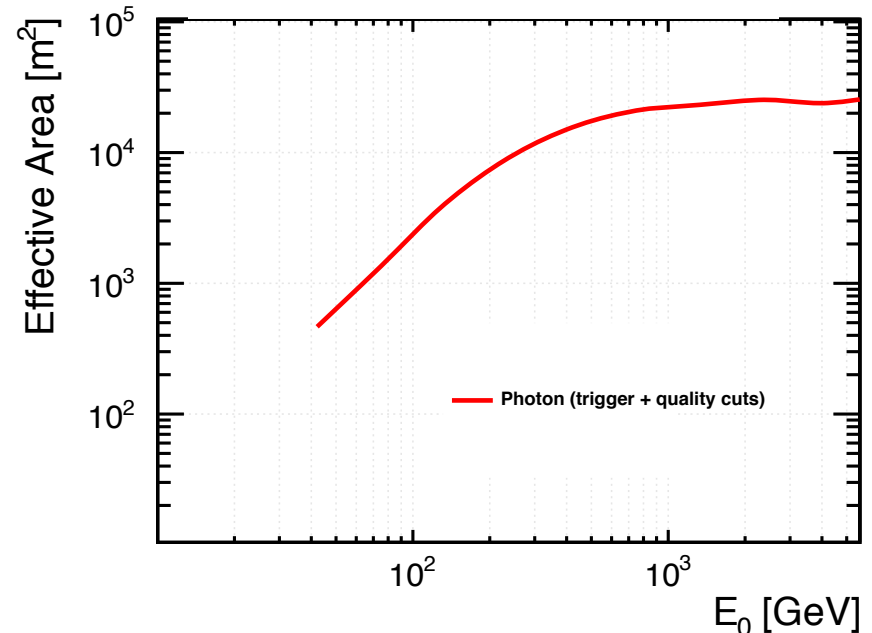
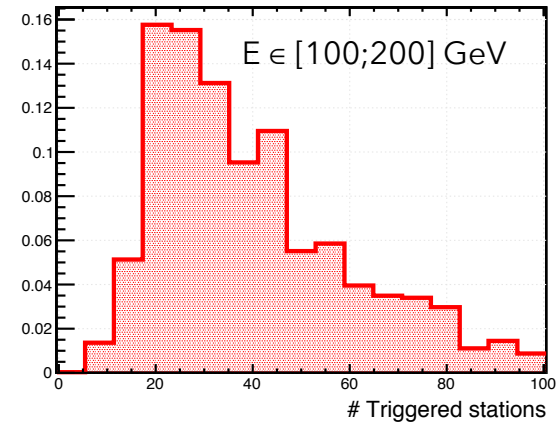
LATTES expected performance

- ❖ Trigger and effective area
- ❖ Core reconstruction
- ❖ Energy reconstruction
- ❖ Geometry reconstruction
- ❖ Gamma/hadron discrimination
- ❖ Sensitivity to steady sources



Effective Area

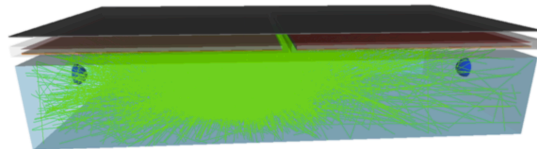
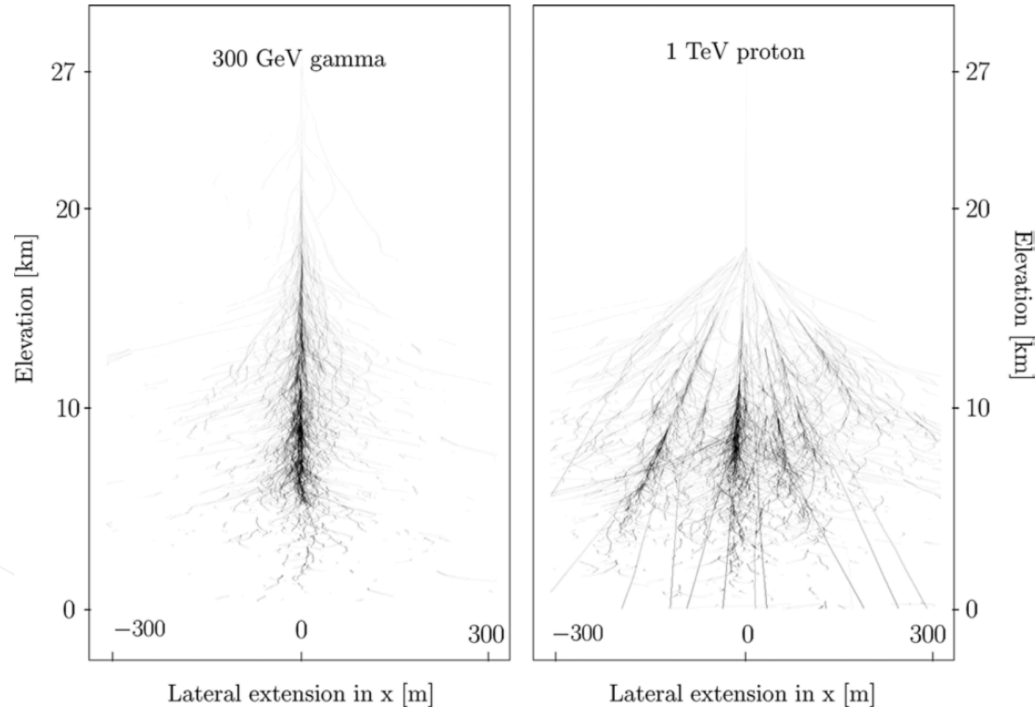
- ✧ Station Trigger
 - ✧ 5 p.e. in each WCD PMT
- ✧ Event Trigger
 - ✧ 3 stations
- ✧ Quality cuts
 - ✧ Good core rec
 - ✧ Core in array
 - ✧ 10 hits in RPCs pads (belonging to active WCDs)
 - ✧ Good geom rec
- ✧ After applying all quality cuts
LATTES gets an effective area of **$\sim 1000 \text{ m}^2$ for $E = 100 \text{ GeV}$**



LATTES expected performance

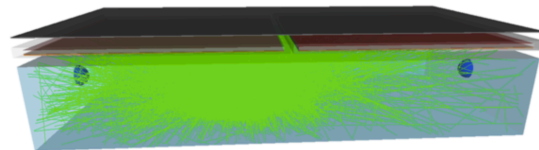
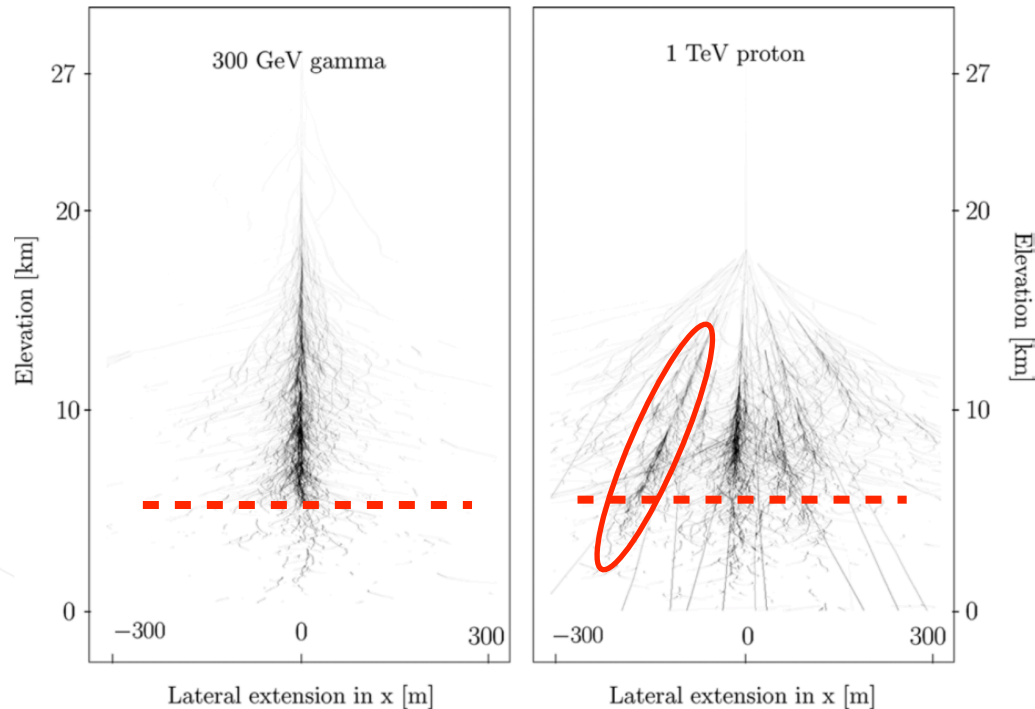
- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ **Gamma/hadron discrimination**
- ✧ Sensitivity to steady sources

Shower characteristics



A pure electromagnetic shower (gamma) has distinct features from a shower with an hadronic component (hadron)

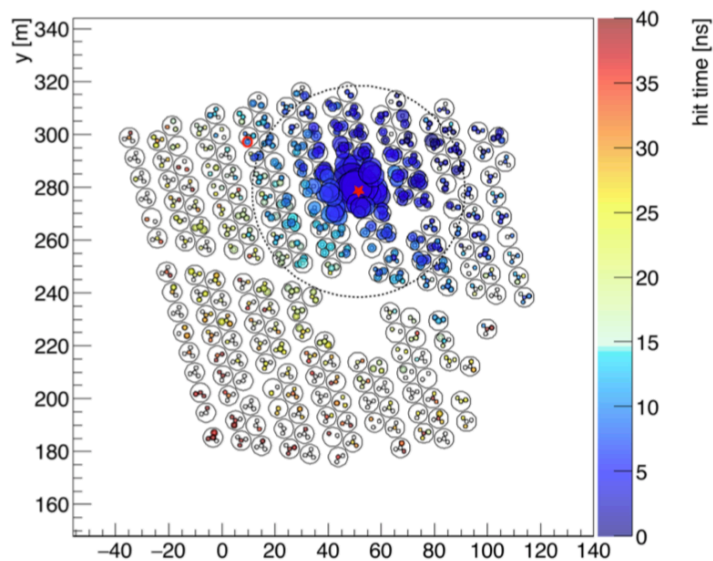
Shower characteristics



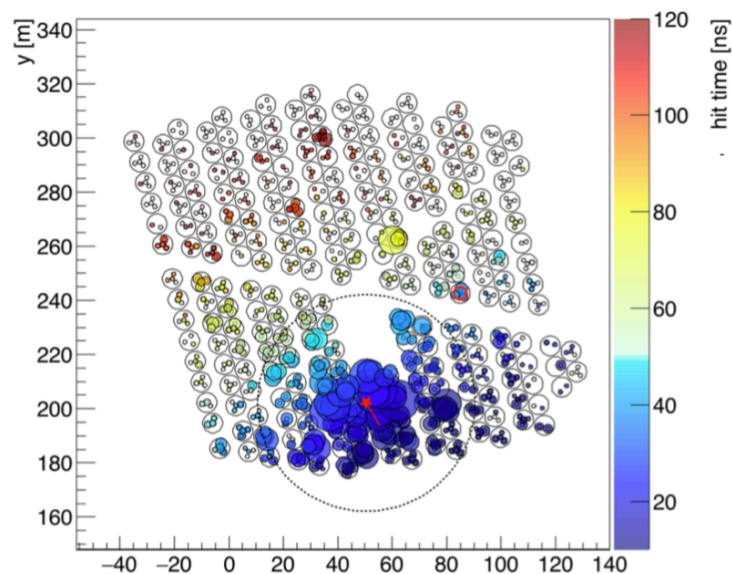
A pure electromagnetic shower (gamma) has distinct features from a shower with an hadronic component (hadron)

Looking for high- p_T sub-showers

gamma shower

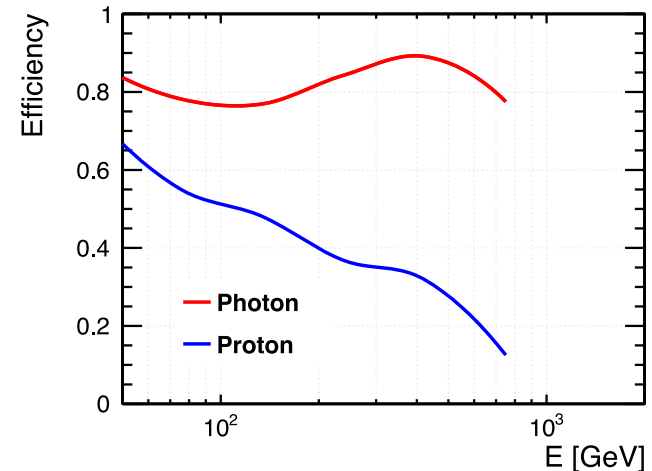
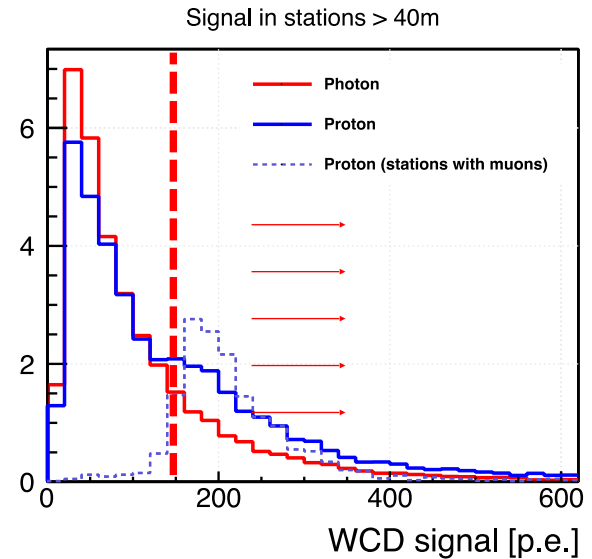


hadron shower



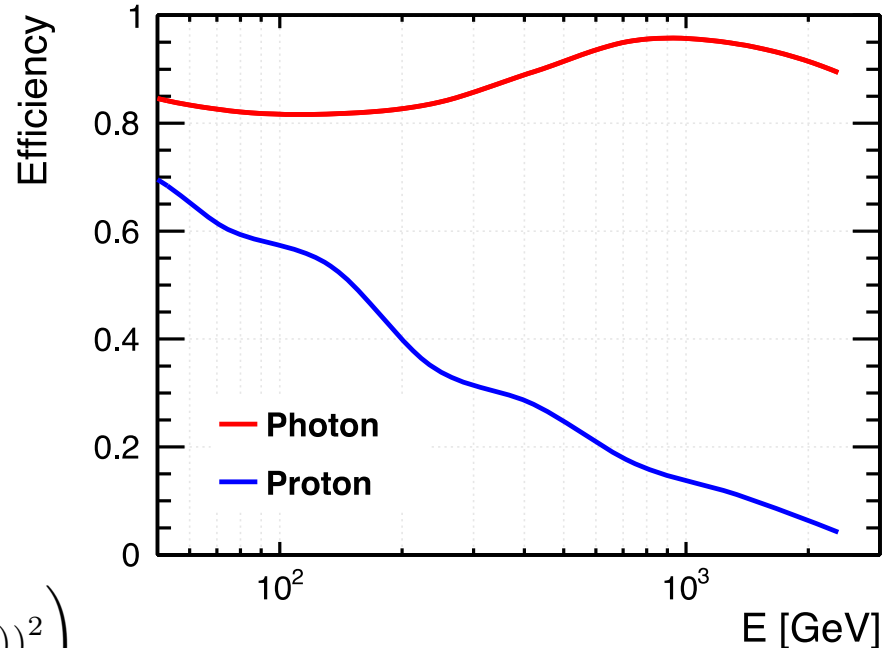
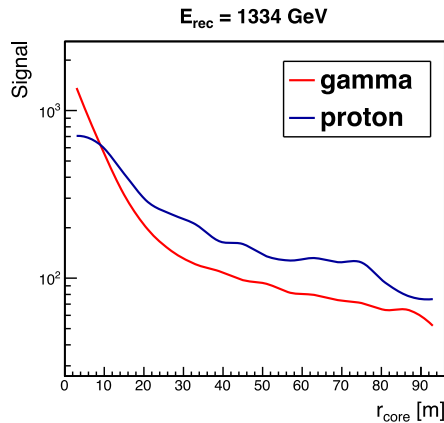
Looking for high p_t sub-showers

- ✧ LATTES g/h discrimination
 - ✧ Use only stations with a distance above 40 m
 - ✧ **S₄₀**: sum all WCD stations signal
 - ✧ **S_{40_high}**: sum all WCD stations that have a signal above the muon energy threshold
 - ✧ Compute **S_{40_high} / S₄₀**
 - ✧ Not optimized...



High-energy discrimination strategy

Average LDF



✧ Compute event-by-event:

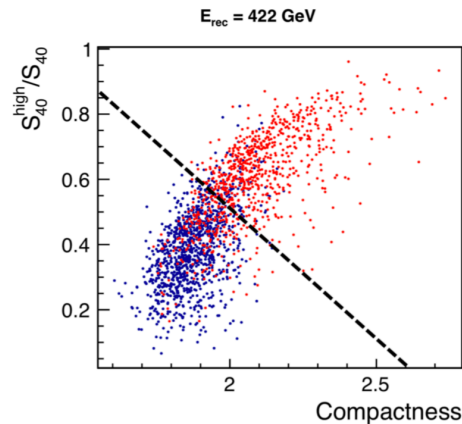
$$\text{Compactness} = \log_{10} \left(\sum_i^n (\langle LDF \rangle(r_i) - y(r_i))^2 \right)$$

✧ Lateral distribution function (LDF)

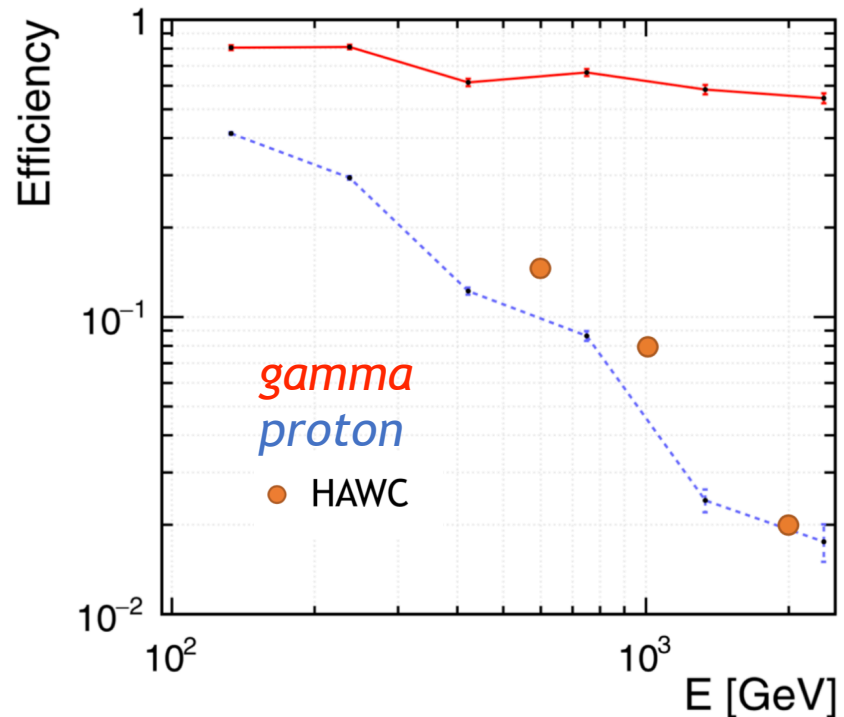
✧ LDF of gamma showers is more steep and less bumpy than the LDF of hadron showers

LATTES g/h discrimination

Using only the WCD



Compactness = LDF steepness
S40 = Signal outside 40 m

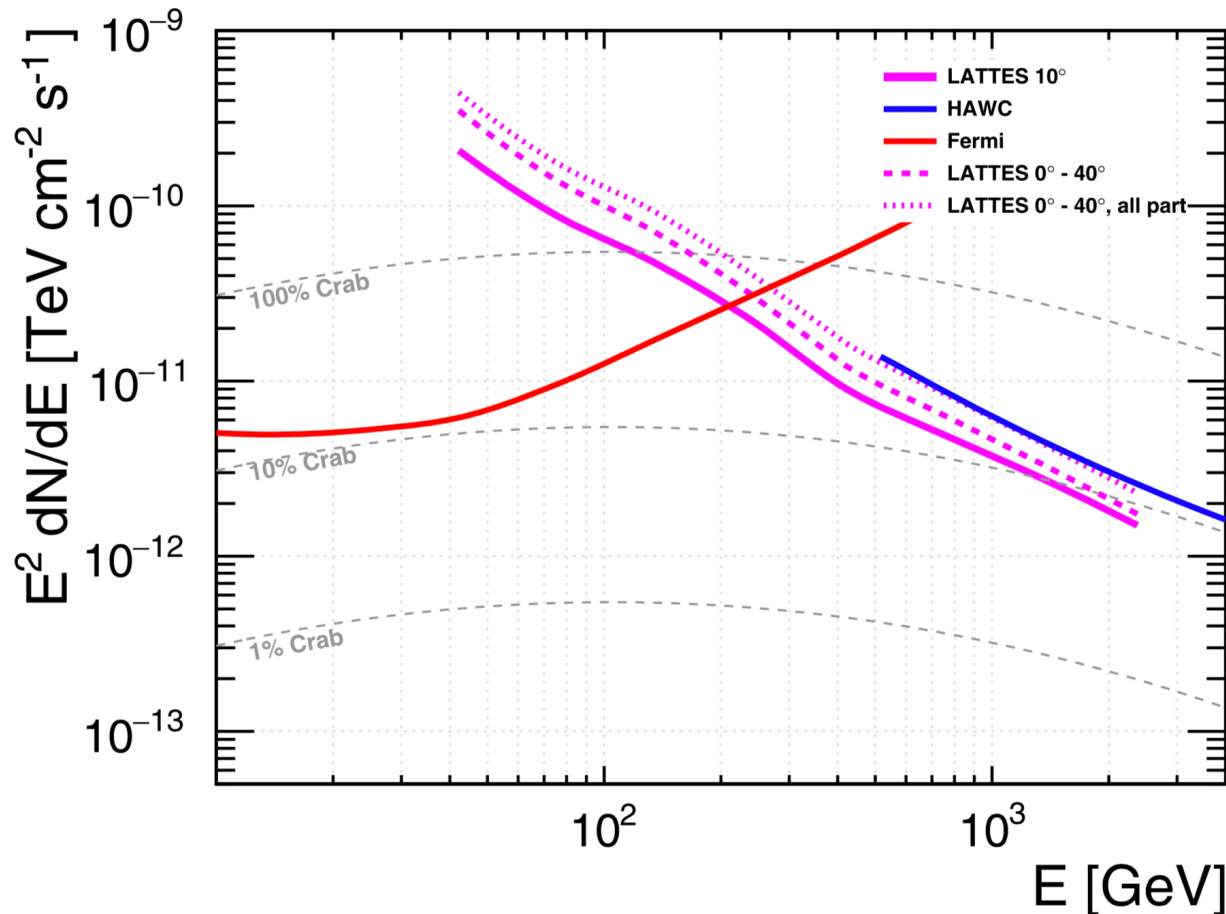


- ✧ Although not optimized the gamma/hadron discrimination results are already very encouraging
- ✧ Starting to investigate more sophisticated tools (ANN: pattern at ground ; cuts ; ...)

LATTES expected performance

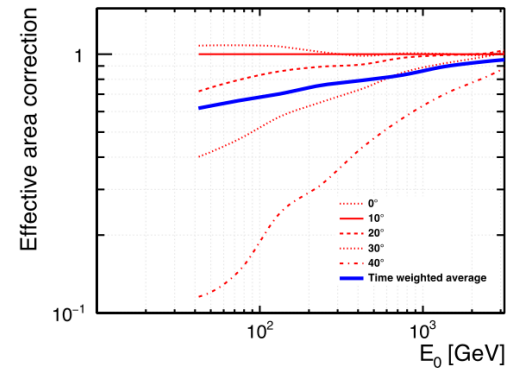
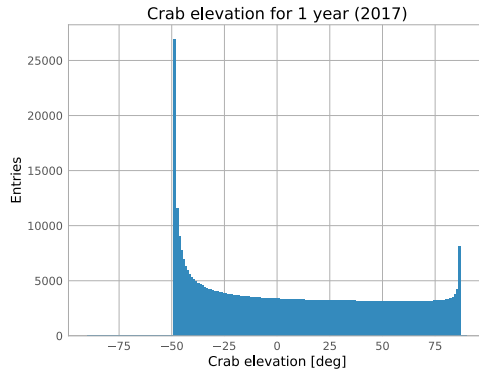
- ✧ Trigger and effective area
- ✧ Core reconstruction
- ✧ Energy reconstruction
- ✧ Geometry reconstruction
- ✧ Gamma/hadron discrimination
- ✧ **Sensitivity to steady sources**

Sensitivity to steady sources



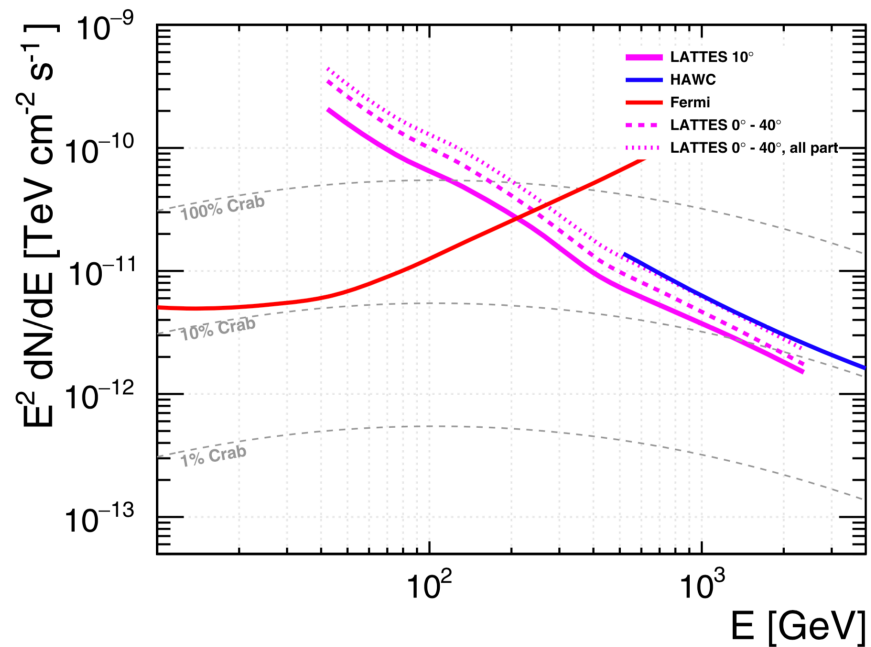
✧ *Full line*: full MC calculation for a source at 10 degrees in zenith

Sensitivity to steady sources



✧ *Dashed line*: Crab transit as seen by HAWC

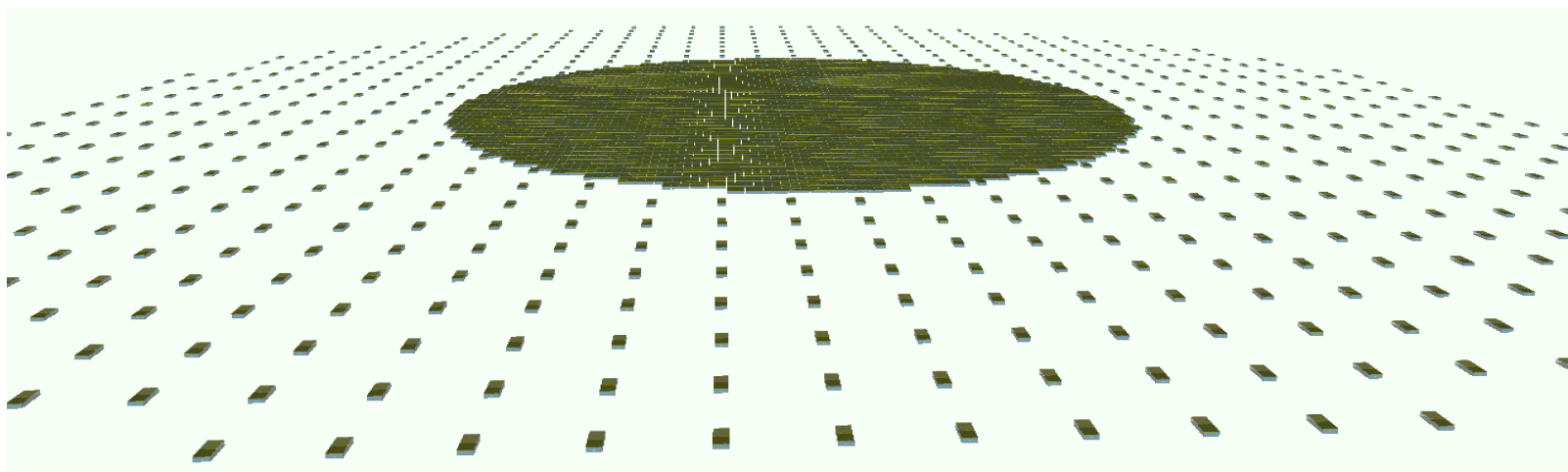
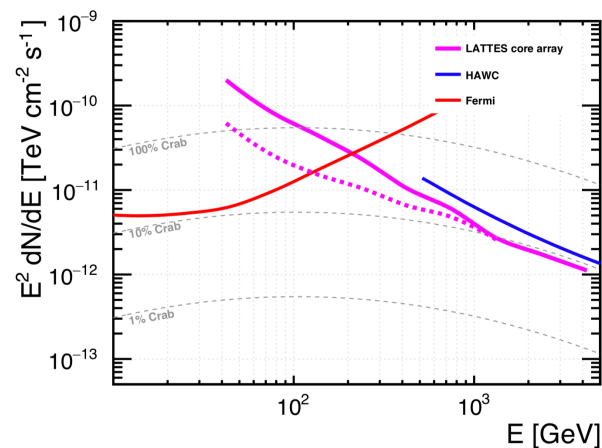
✧ Degradation of effective area with zenith angle estimated from electromagnetic energy at ground



Sparse Array

- ✧ Use a sparser array (100 000 m²)
 - ✧ Collect more events at higher energies
 - ✧ Remove high energy events that fall outside of the core array
- ✧ Built LATTES fastsim:
 - ✧ Use particle tracklength in water to generate number of photons collected by PMTs

Vetoing showers that fall outside the array



Summary

- ✧ LATTES shower reconstruction performance has been evaluated yielding very good results
 - ✧ Shower trigger (effective area)
 - ✧ Shower core reconstruction
 - ✧ Shower energy reconstruction
 - ✧ Shower geometry reconstruction
 - ✧ Gamma/hadron discrimination
- ✧ LATTES capabilities are far from being fully explored
 - ✧ Possible improvements already identified
 - ✧ Sparse array to veto far away high-energy showers (main background source)
 - ✧ Use RPC patterns to discriminate g/h
 - ✧ Better assess LATTES ability to reconstruct
 - ✧ Inclined showers
 - ✧ Heavier primaries induced showers
 - ✧ ...

Acknowledgements



**REPÚBLICA
PORTUGUESA**

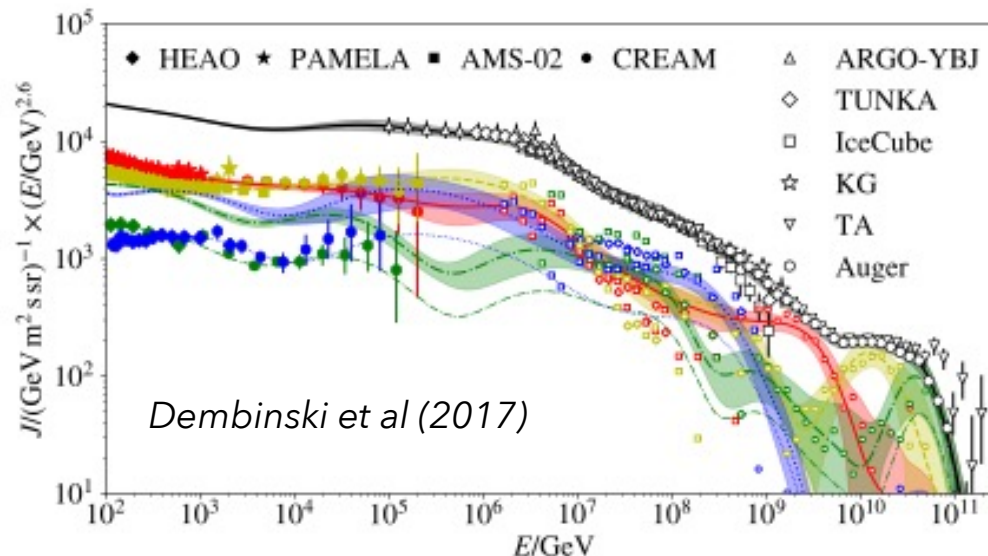
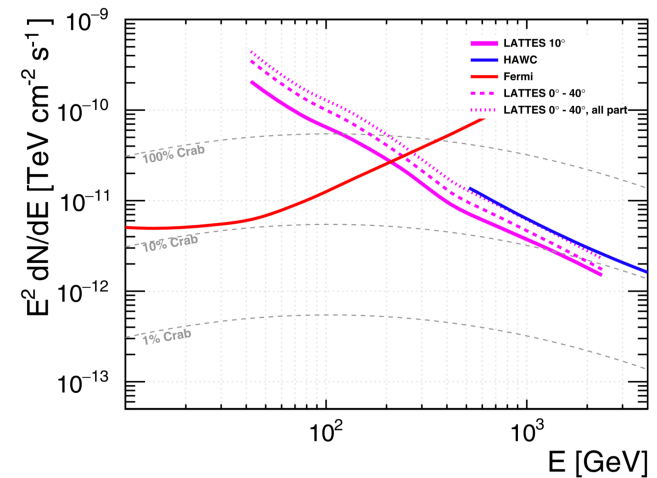


**TÉCNICO
LISBOA**

Backup slides

Sensitivity to steady sources

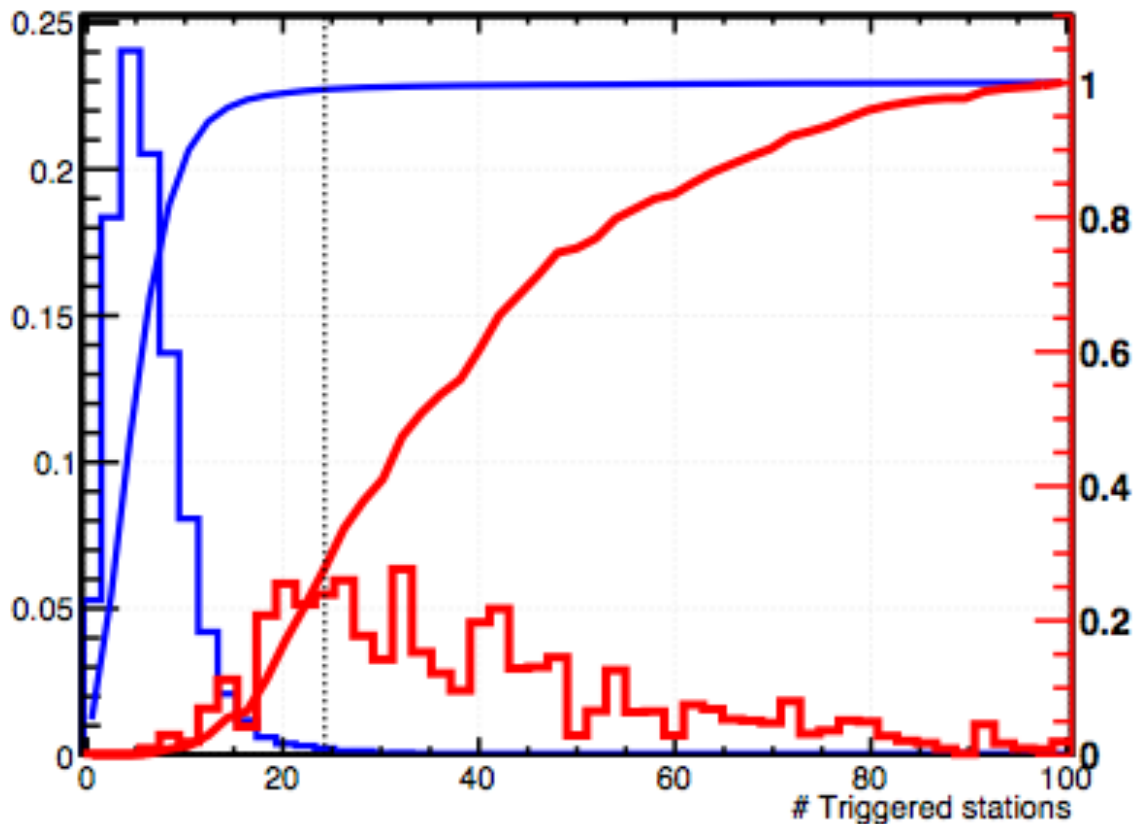
- ✧ *Dotted line*: CR all-spectrum
- ✧ Additional elements (He, N, Fe...)
- ✧ Assume that LATTES cannot distinguish gammas from irons



Cosmic rays trigger rate vs photon triggers

Cosmic rays triggered stations

Crab photons with energy in [100 GeV, 200 GeV]



> 18 stations : 98% background rejection (440 kHz) ; 90% photon efficiency

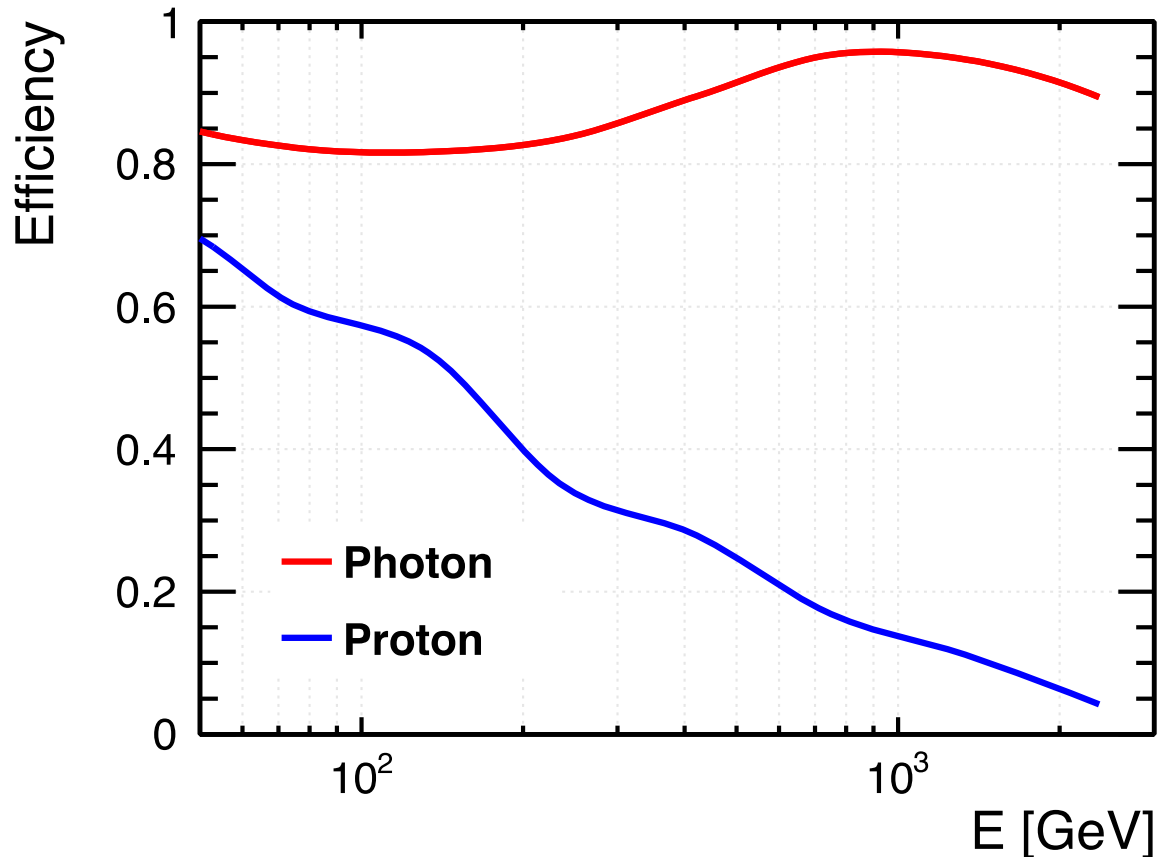
> 25 stations : 99% background rejection (220 kHz) ; 75% photon efficiency

High-energy discrimination strategy

- ✧ Get the **gamma average LDF** for each reconstructed energy bin
- ✧ Fit the average LDF to each single event
 - ✧ Absorb the **normalization** factor
- ✧ Compute the shower **compactness**
 - ✧ Event LDF "distance" to the gamma average LDF

$$\text{Compactness} = \log_{10} \left(\sum_i^n (\langle LDF \rangle (r_i) - y(r_i))^2 \right)$$

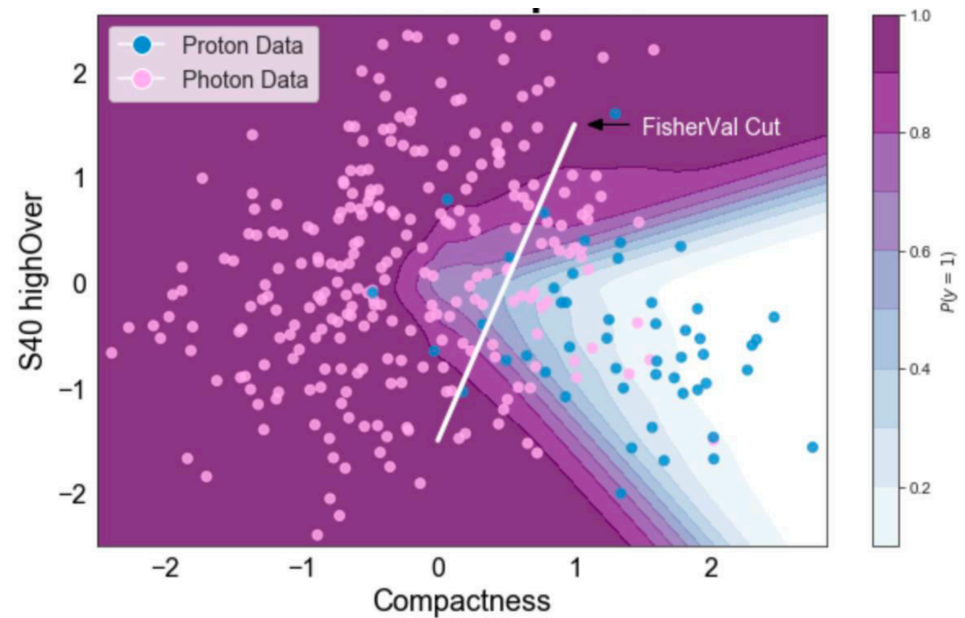
High-energy discrimination strategy



Shower **compactness** discrimination variable allows for a good background rejection which increases with energy

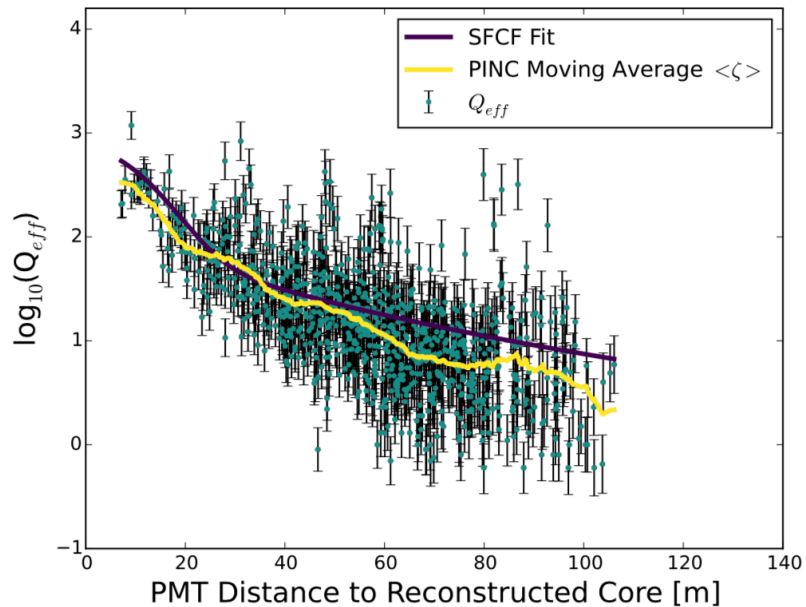
G/H discrimination and ANN

- ❖ Linear Discriminante (Fisher) allows a good separation
- ❖ Simple artificial neural networks can improve g/h discrimination
- ❖ Keras + Scikit-learn + ANN with 5 layers
- ❖ More simulation statistics necessary to apply parametric cuts
- ❖ Test at lower energies...

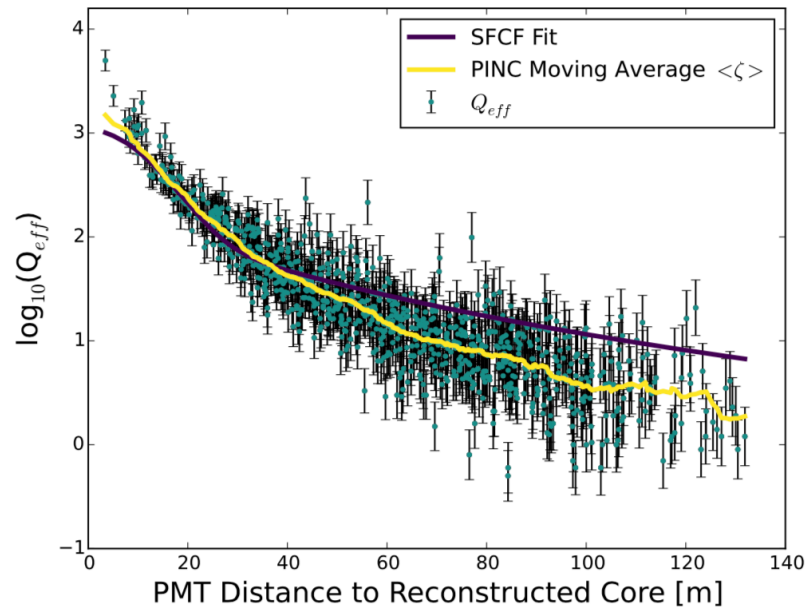


HAWC g/h discrimination

Cosmic ray



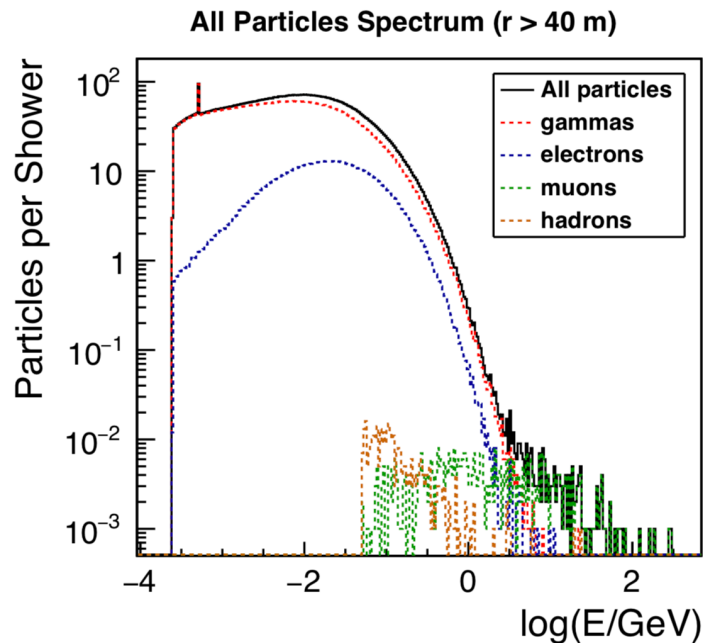
Gamma-ray



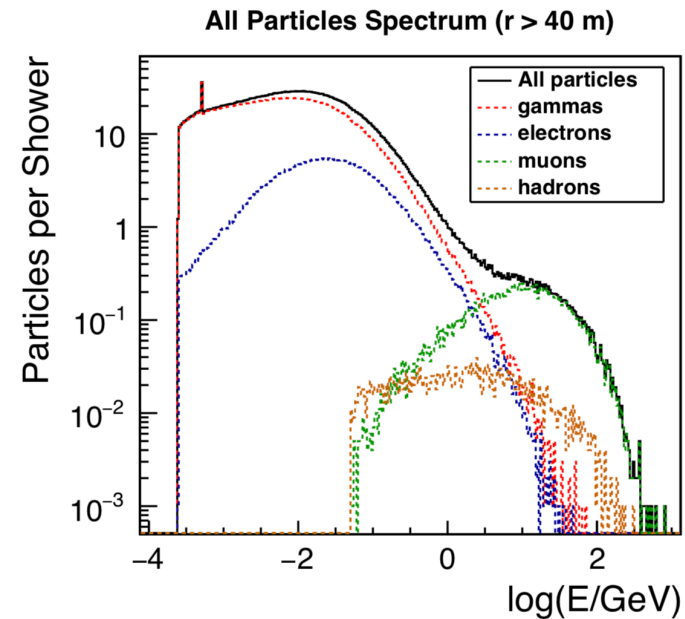
Shower calorimetric information

E=5 TeV

Gamma induced showers

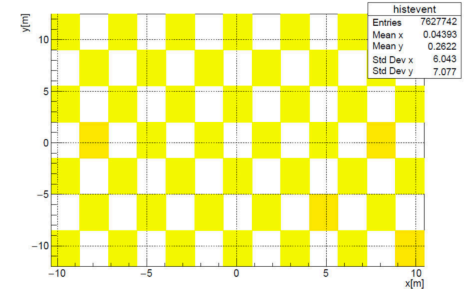
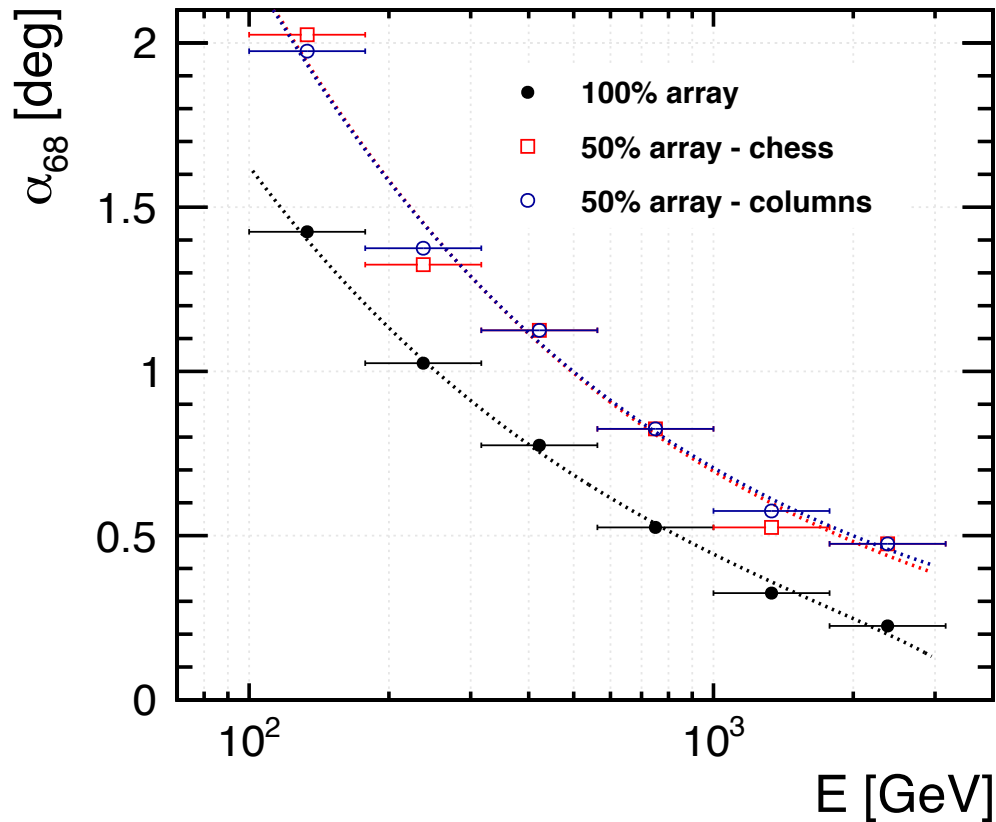


Proton induced showers



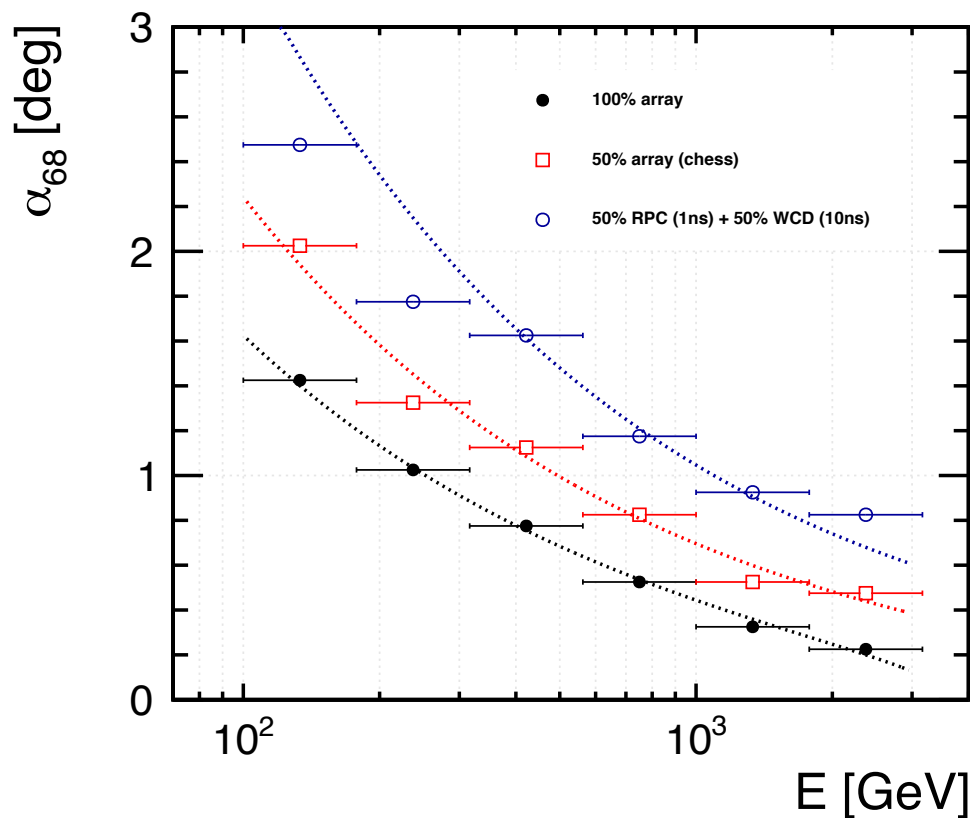
- ✧ High p_T sub-shower carry large amounts of energy
- ✧ Look for energetic clusters far from the shower core (> 40 m)
 - ✧ Muons and high-energy photons/electrons

Geom Rec: array configuration



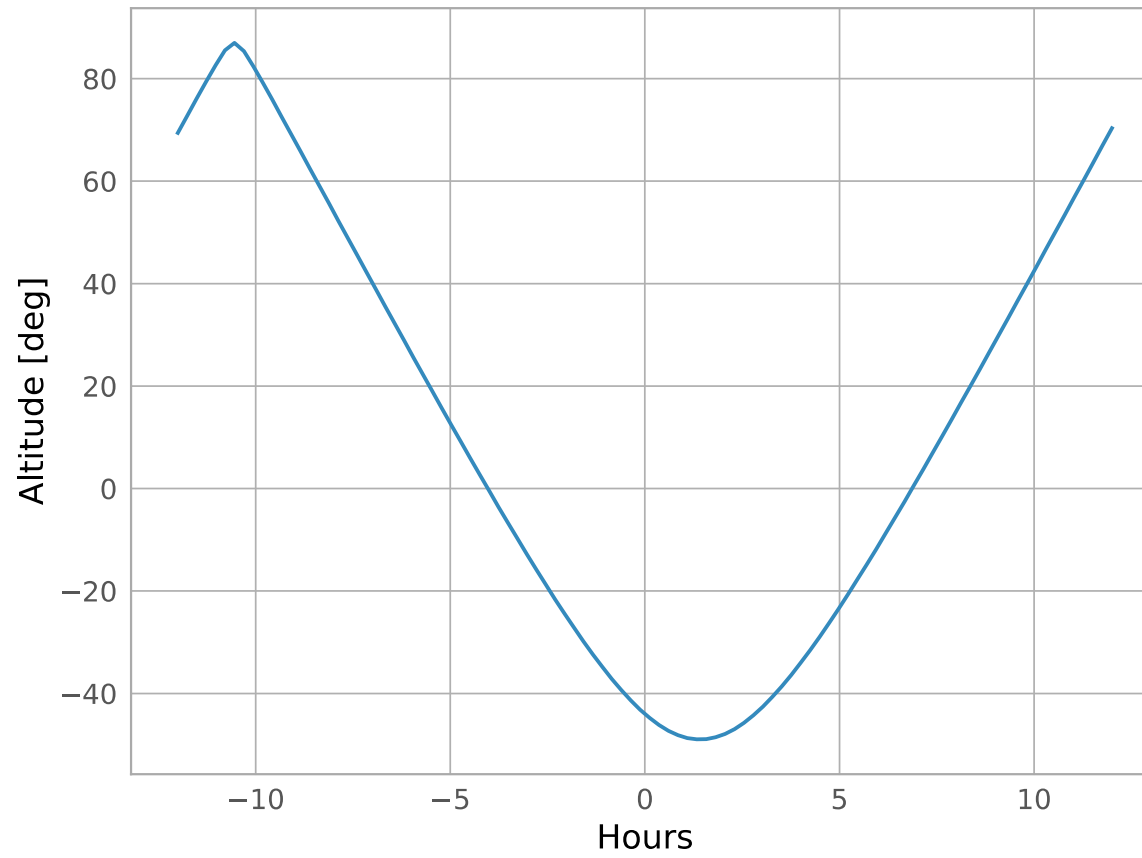
✧ It seems important to have RPCs on all stations

Geom Rec: RPC + WCD

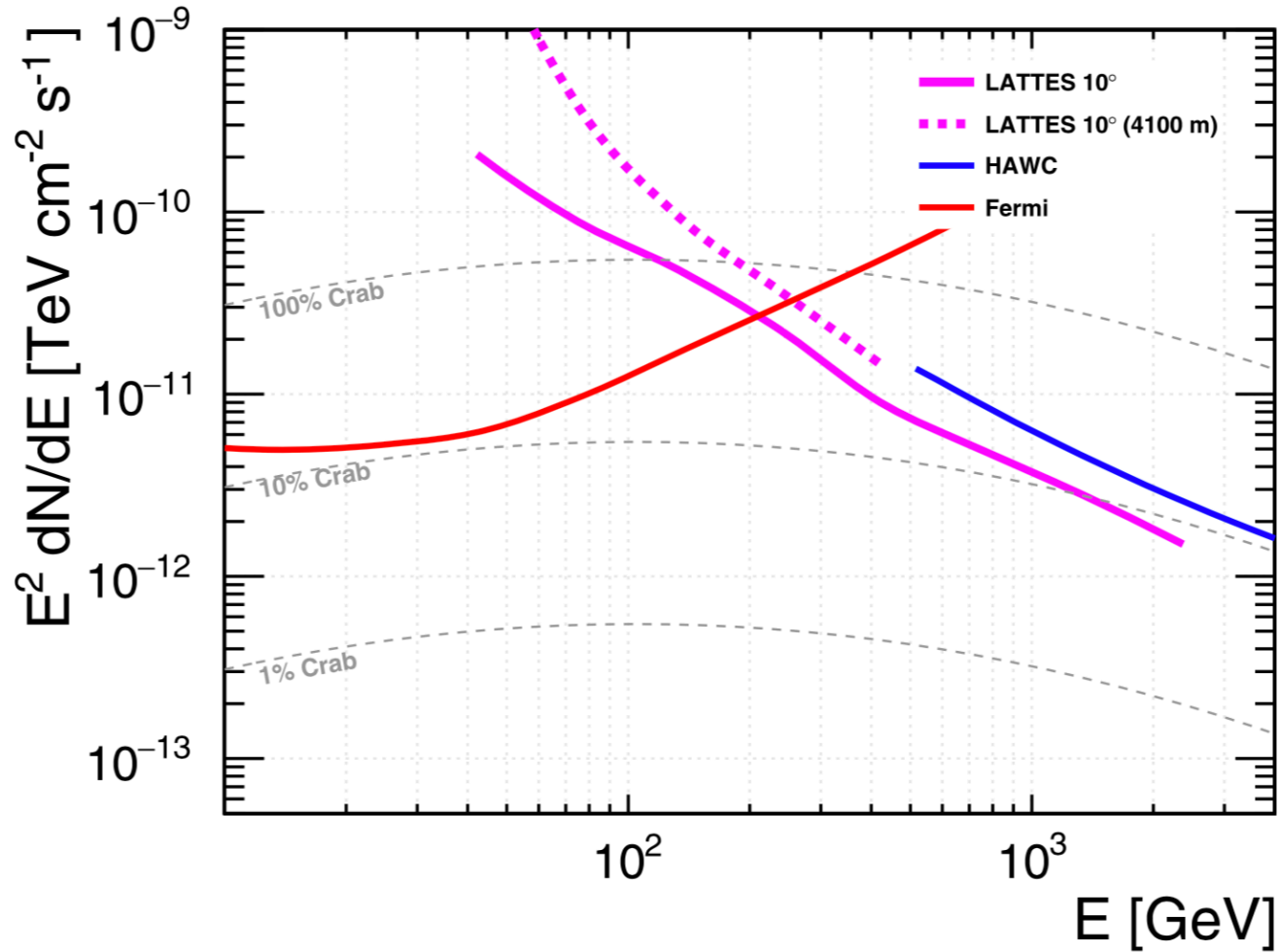


- ◇ Next steps: use only first hit in pad (trade-off between higher correlation with shower front and event statistics)

Crab



Impact of altitude



LATTES: a hybrid detector

❖ Thin lead plate

- ❖ To convert the secondary photons
- ❖ Improve geometric reconstruction

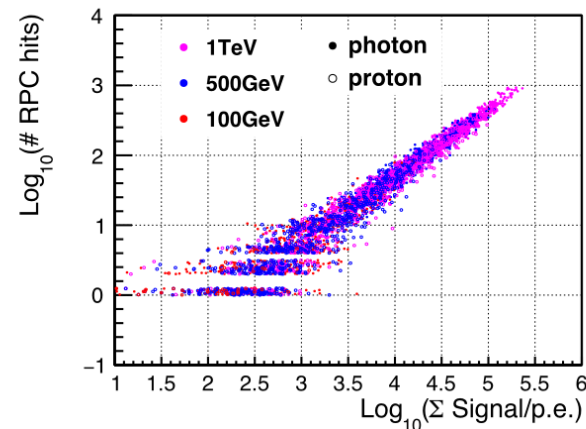
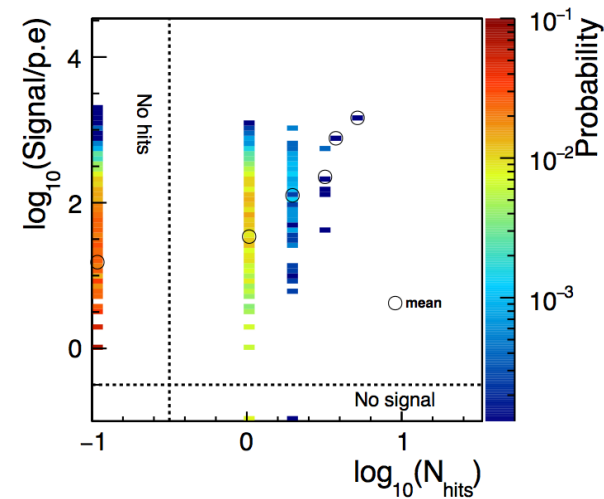
❖ Resistive Plates Chamber

- ❖ Sensitive to charged particles
- ❖ Good time and spatial resolution
- ❖ Improve geometric reconstruction
- ❖ Explore shower particle patterns at ground

❖ Water Cherenkov Detector

- ❖ Sensitive to secondary photons and charged particles
- ❖ Measure energy flow at ground
- ❖ Improve trigger capability
- ❖ Improve gamma/hadron discrimination

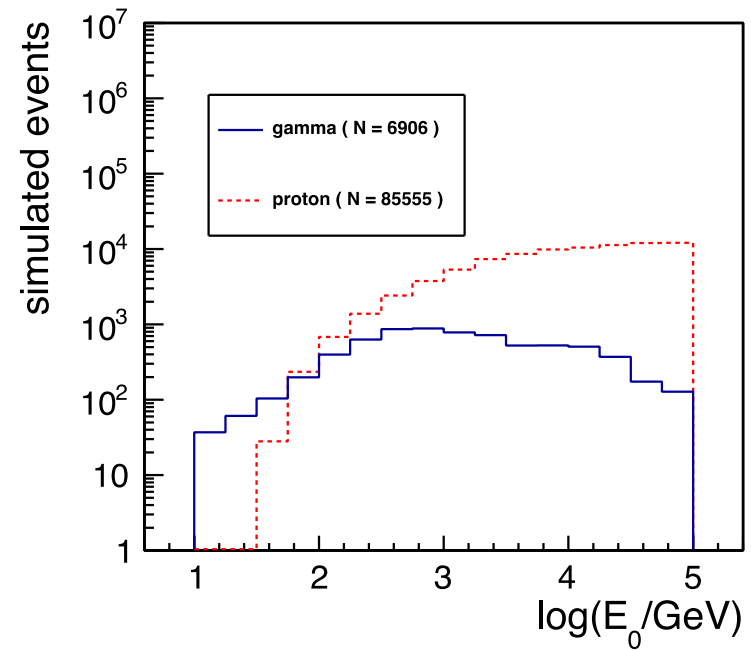
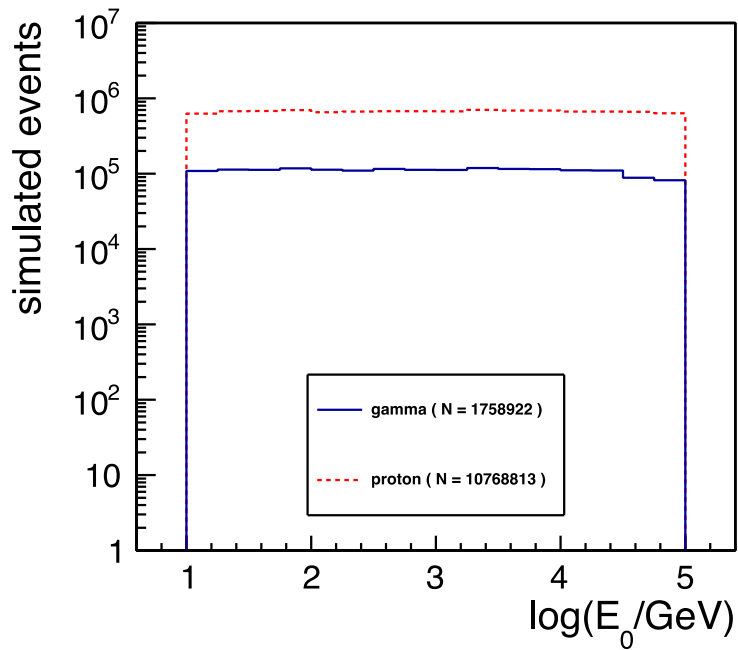
WCD vs RPC (station level)



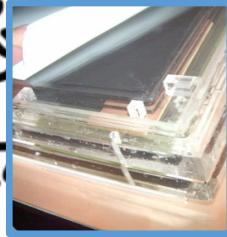
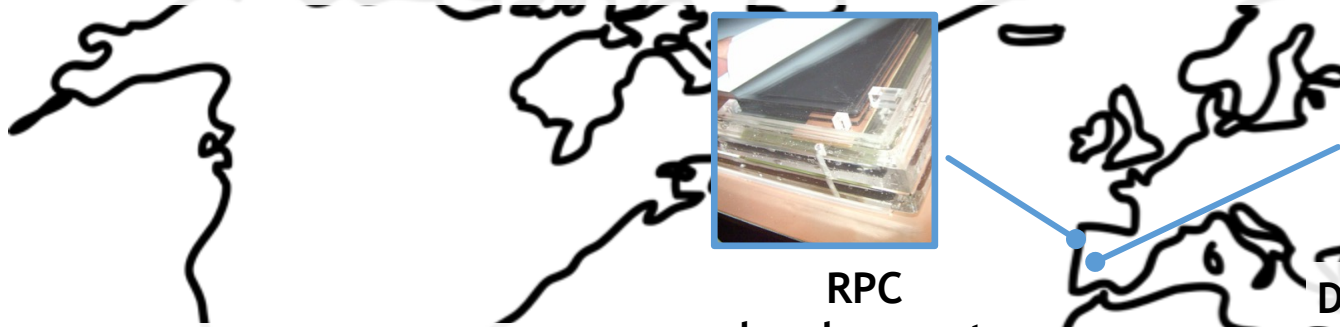
Complementarity

Inter-calibration

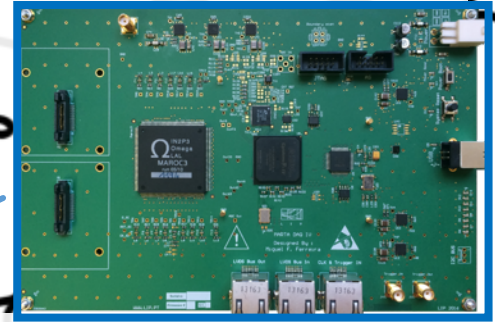
Reconstruction efficiency



Ongoing developments and tests on RPCs, electronics and read-out systems

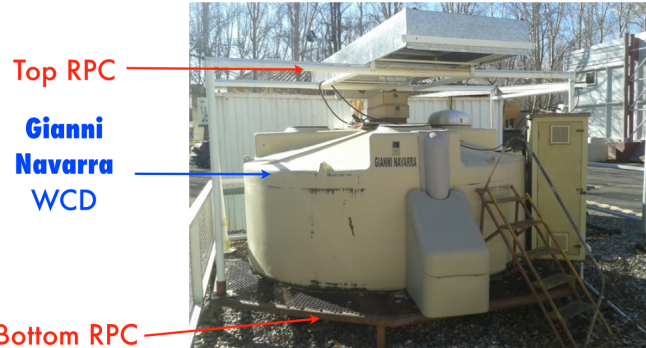


RPC
developments
Construction and Assembling



DAQ Engineering prototype

RPC based muon hodoscope for precise studies of the Auger WCD



RPC hodoscope

RPCs in the field @ Auger



Conceição