



# Prospects for CCSN alert triggering with KM3NeT

Workshop on core-collapse supernova neutrino detection

Massimiliano Lincetto

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Aix-Marseille Université, CNRS/IN2P3, CPPM, Marseille (France)

## KM3NeT design key points





Figure 1: KM3NeT DOM and building block

- 31 PMTs per DOM, 18 DOM per line (detection unit, DU), 115 lines per block;
- Two blocks for the ARCA (IT) and one block for the ORCA (FR) detector;
- local **coincidence**  $\equiv$  cluster of hits on the same DOM within 10 ns;
- multiplicity ≡ number of hit PMTs in a coincidence.

## KM3NeT DAQ system

#### Overview

- all data to shore: every detected hit over 0.3 PE (L0) is recorded with ns time accuracy and sent to the on-shore datafilter;
- the **optical datafilter** builds local coincidences (L1) and produces corresponding data streams (data downsampling is foreseen);
- the **trigger algorithms** process the hit/cluster datas treams and produce triggered events for GeV-TeV scale physics.

#### Supernova monitoring

- Dedicated supernova coincidence datastream including high-multiplicity (≥ 4) coincidences;
- Sampling of PMT single rates on sub-ms scale is foreseen to compensate the downsampling of low-level data.

## KM3NeT background for the CCSN signal

- PMT single rates: 250 kHz per DOM from <sup>40</sup>K decays in sea water (+ bioluminescence contribution);
- Coincidences: dominated by <sup>40</sup>K decays for low multiplicities and atmospheric muons for high-multiplicities.



Figure 2: Multiplicity rates for background and signal (Garching model) in KM3NeT.

- Effective volume for single hit detection is larger but correlation of bioluminescence noise is difficult to constrain without real data from a multi-line detector → study of single rates is undergoing;
- S/N does not benefit from the inclusion of low-level multiplicities → selection of high-level coincidences (6-10 multiplicity).

#### KM3NeT preliminary

### Bioluminescence

- Localised bursts affecting neighbors PMTs → channels recording a hit rate over 20 kHz are suppressed (high-rate-veto);
- Diffuse glows are well-suppressed by the coincidence selection.

## <sup>40</sup>K in seawater

- Constant rate uncorrelated background that affects to 6-fold coincidences;
- Application of an **angular cut** to the hits allowed in coincidence has a slight reduction effect.

#### Atmospheric muons

• Cherenkov tracks from muons produce high-multiplicity correlated concidences across DOMs on a  $\mu {\rm s}$  time scale.

## **ORCA** muon rejection filter

Same approach used both offline and online.

- Reject events with at least two M ≥ 4 coincidences within 1 μs on different DOMs.
- Introduces a 2% dead time dominated by  $\rm ^{40}K$  random coincidences.



Figure 3: Multiplicity rates before and after filter

Multiplicity	6	7	8	9	10	$\sum_{6-10}$
ORCA rej. eff.	40%	80%	90%	90%	92%	62%
ARCA rej. eff.	12%	26%	35%	35%	35%	17%

**Foreseen development** (ARCA): use of physics events identified by the trigger to suppress the DOMs in correspondence with muons.

## KM3NeT CCSN sensitivity



Figure 4: Sensitivity of KM3NeT to a CCSN supernova neutrino burst

- This sensitivity estimation assumes a **known arrival time** of the CCSN neutrino burst.
- The signal is evaluated on a defined time window τ after the exact time of the event.

(p-value) 
$$\equiv P(X \ge X_T) = \sum_{X=X_T}^{+\infty} \mathcal{P}(\rho_B \cdot \tau, X)$$

What if we don't know the onset time?

#### Configuration & false rate

- sliding time window  $\equiv \tau$
- sampling frequency  $\equiv f$

$$R_B(X \ge X_T) = f \sum_{X=X_T}^{+\infty} \mathcal{P}(\rho_B \cdot \tau, X)$$

#### Robustness

Switch to **DOM counting** instead of coincidence counting (maximum of one coincidence in the SN cut per DOM is accepted in the time window).

#### Significance?

No longer well-defined as long as the observation time T is indefinite.

Fixing T a post-trial significance can be recalculated multiplying the p-value by the resulting **trial factor** fT.

- Integral over time is profitable as long as the signal expectation is above the background expectation;
- In general this is distance-dependent;
- In practice the peak of interest (accretion) is **4-500 ms wide**.



Figure 5: Expected time profile of a CCSN neutrino signal in 1 KM3NeT line, 1 timeslice = 100 ms

#### First approach

- If too low: risk of losing signal events due to the time discretization;
- If too high: unnecessary increase of the number of trials;
- Natural choice for KM3NeT is the DAQ time slicing (100 ms  $\leftrightarrow$  10 Hz).

#### **Getting smarter**

- *f* can be increased arbitrarily once the coincidence stream has been reconstructed;
- Instead of fixing a sampling frequency, update the time window with each detected coincidence → statistical model is not trivial, toy MC investigation is needed;
- Use of **bayesian blocks** approach (cfr. IceCube).

- DOM counting regime is shown to be very stable and perfectly predictable with Poisson statistics;
- Bioluminescence does not impact the trigger in terms or false alerts but produces loss of efficiency due to vetoed channels.



Figure 6: False trigger rate for ORCA 1 line (100 ms time window) and for ARCA 2 lines for (400 ms time window).



**Figure 7:** False alert rate expressed in 1/week as a function of the SN distance for a 27  $M_{\odot}$  progenitor, resp. for ORCA 10 lines (left) and ORCA 115 lines (right).

Under the **SNEWS** requirement of 1/week false alert rate requirement a coverage up to **20 kpc** is reached with 1 ORCA-like KM3NeT building block.

The offline analysis for the CCSN detection in KM3NeT can be translated with minor adaptations to a real-time online approach for the SN triggering.

The optimization of the time search is still under study and a more sophisticated **time-domain analysis** could bring further improvements.

According to the current estimation, a single KM3NeT block is able to trigger a 27  $M_{\odot}$  CCSN at **20 kpc within the false alert requirements for SNEWS**.