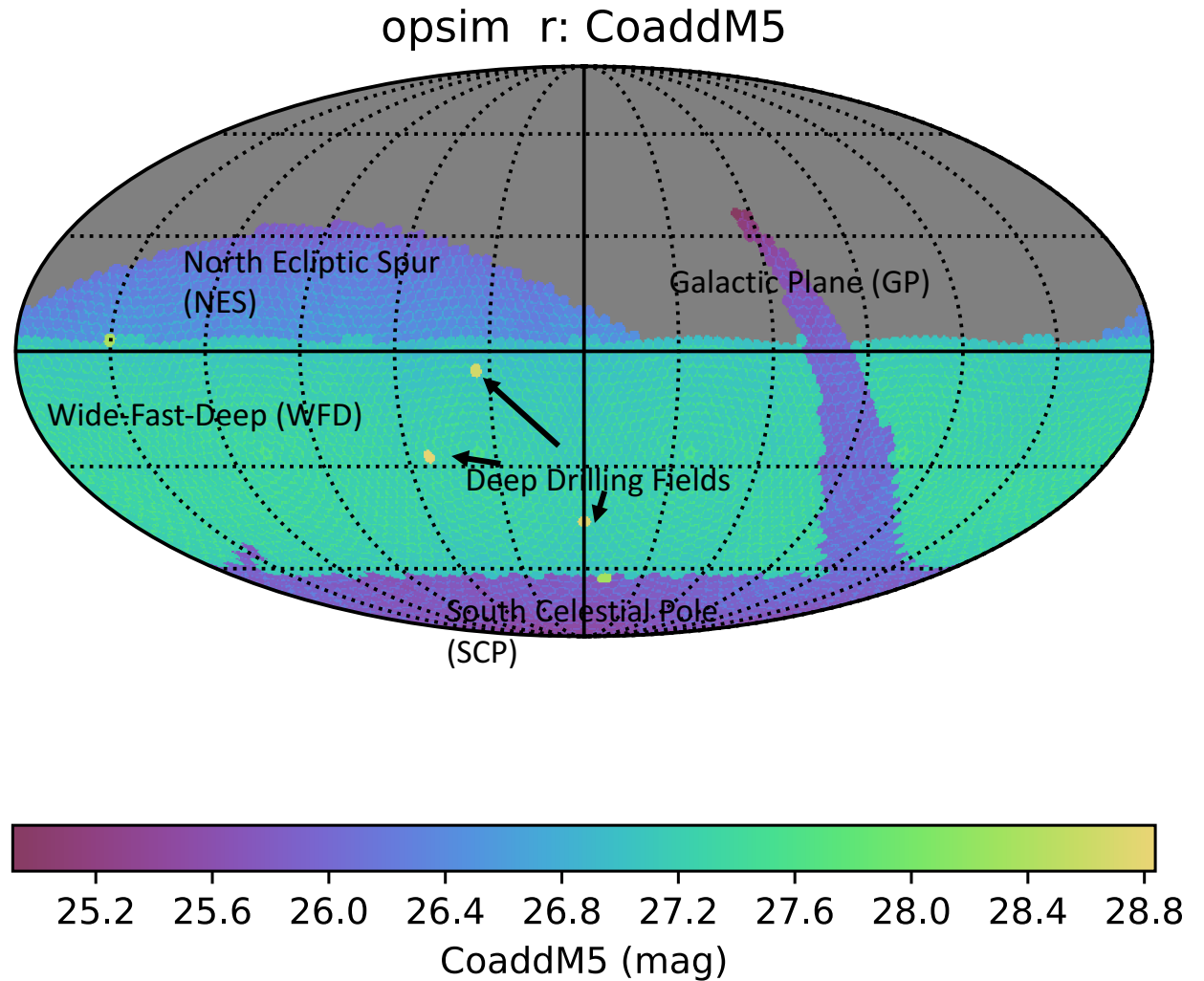


LSST Scheduler Status and Future Plans

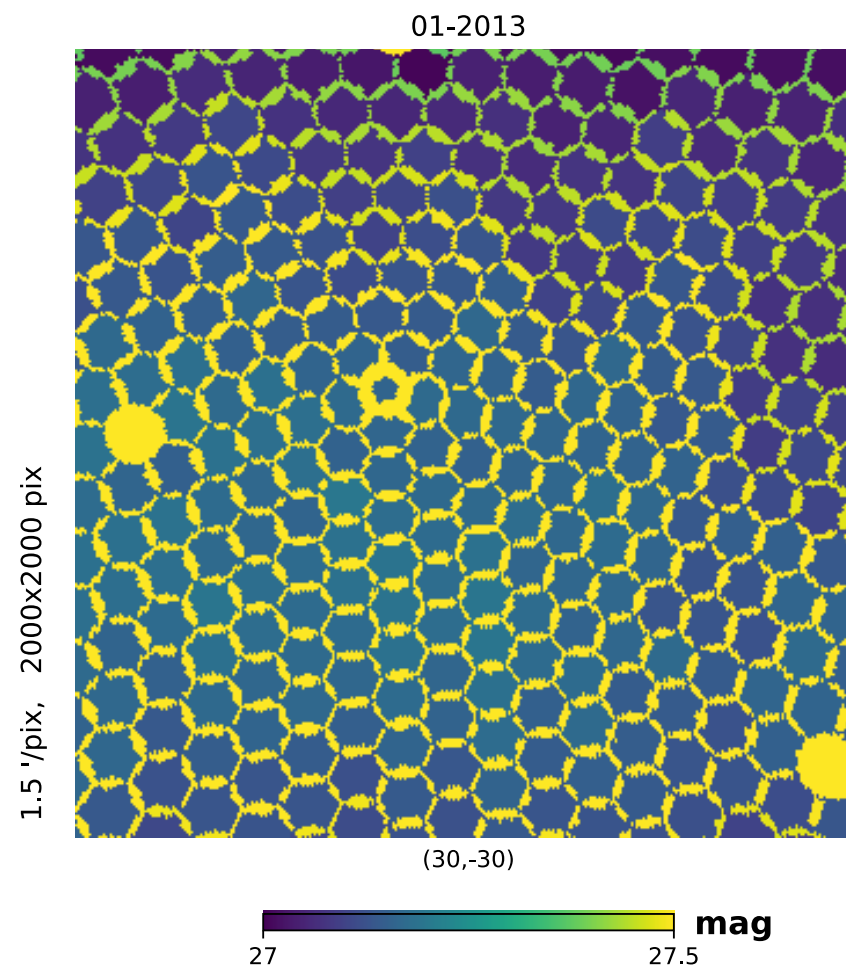
Peter Yoachim
University of Washington

Old proposal based
scheduler

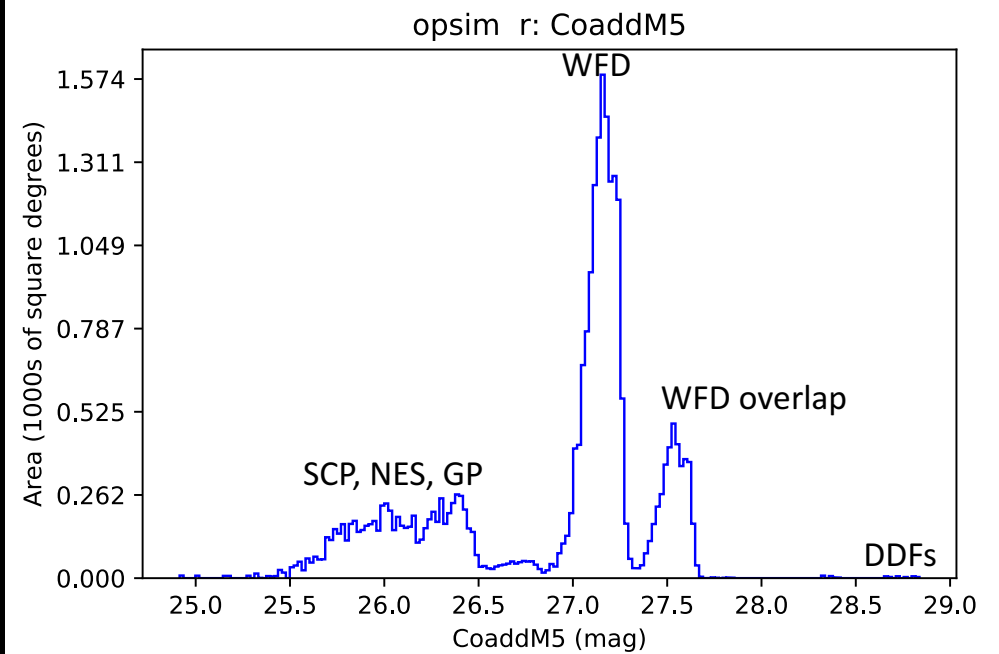
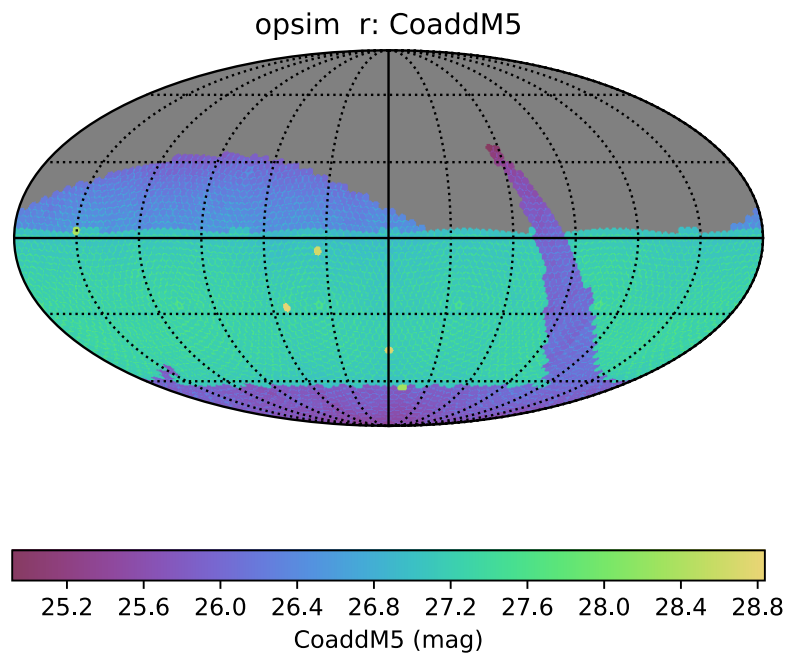
Simulating the motion
of the dome and
telescope, using realistic
weather logs, scheduled
and unscheduled
downtime, seeing
model, sky brightness
model, etc.



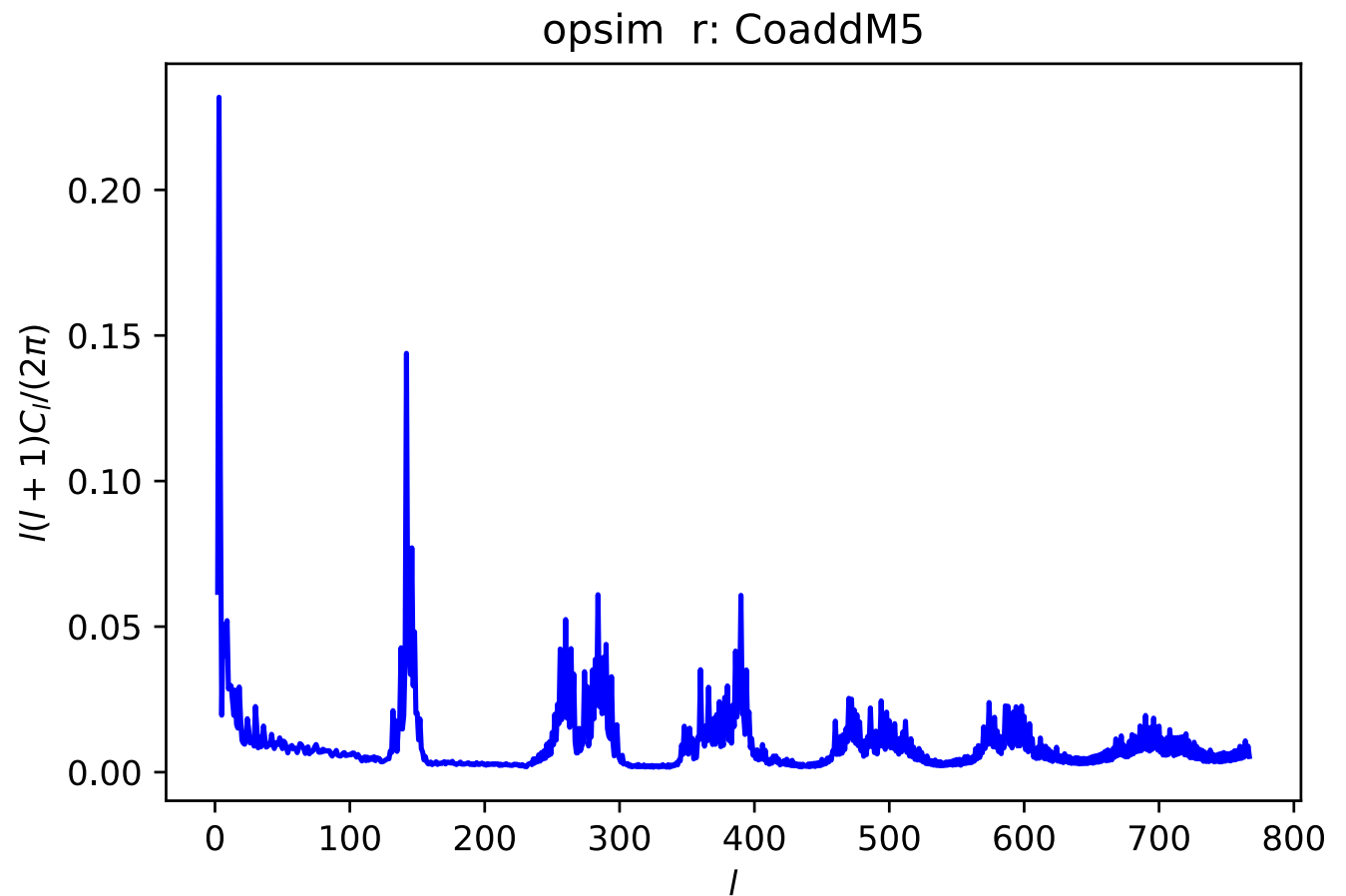
Fixed fields mean overlap areas get
over-observed



Old proposal based scheduler



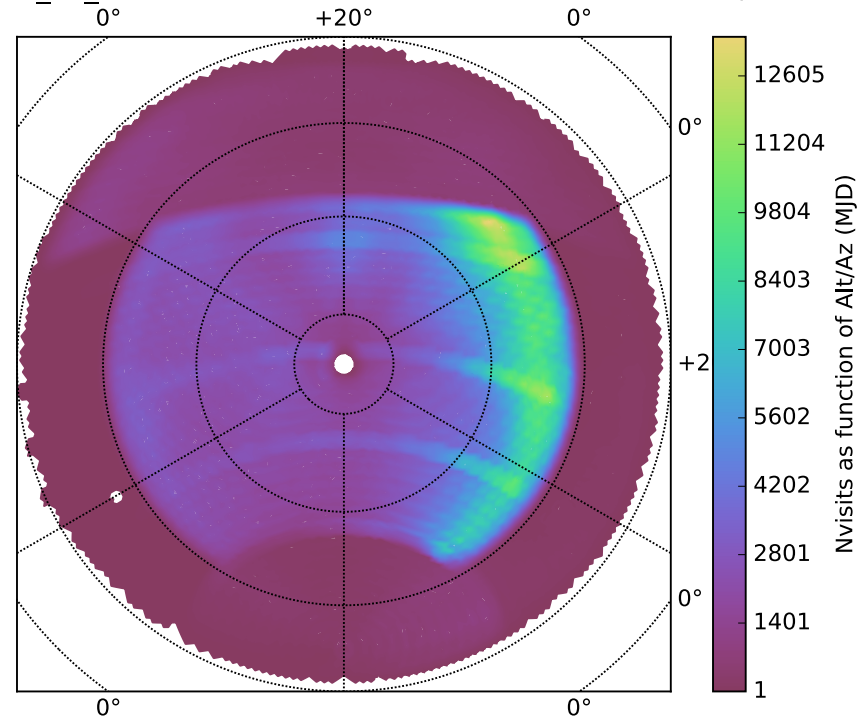
Overlap region
creates a spike in the
power spectrum at ~ 1
degree



Before, we had a “Proposal Based” scheduler.

- Write down all the observations we want (125 visits of Field 23 in r, 90 visits of field 36 in y, ...)
- Pick the “best” observation from remaining list at any given time
- Mostly minimizing slew time

astro_lsst_01_1000 all observations: Nvisits as function of Alt/Az

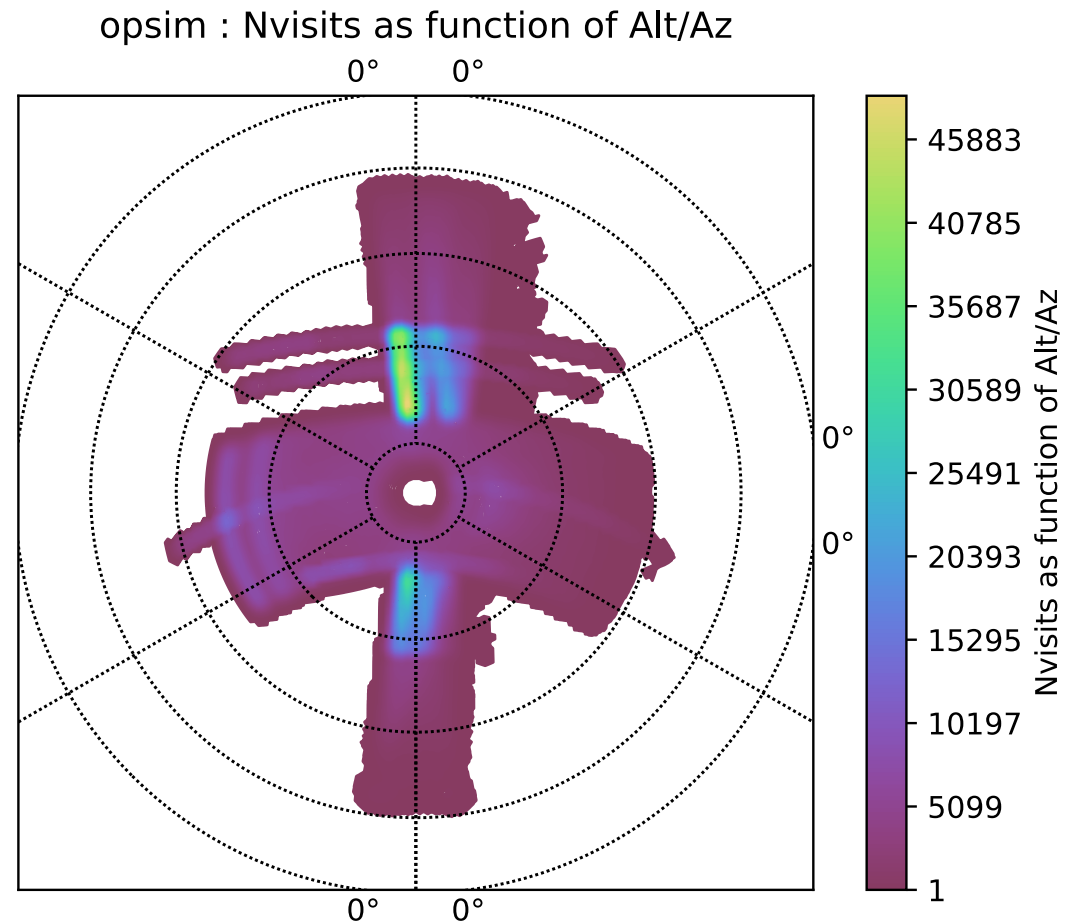


Alt,az distribution for the old scheduler. To minimize slew time, it tends to go to hit the airmass limit and raster in azimuth.

Now, with Markovian Feature Based framework

- Track progress at high resolution (HEALpixels, nside=32, 1.8 degree resolution, 12,288 pixels)
- Compute a reward function for which HEALpixel is most desirable at any given time

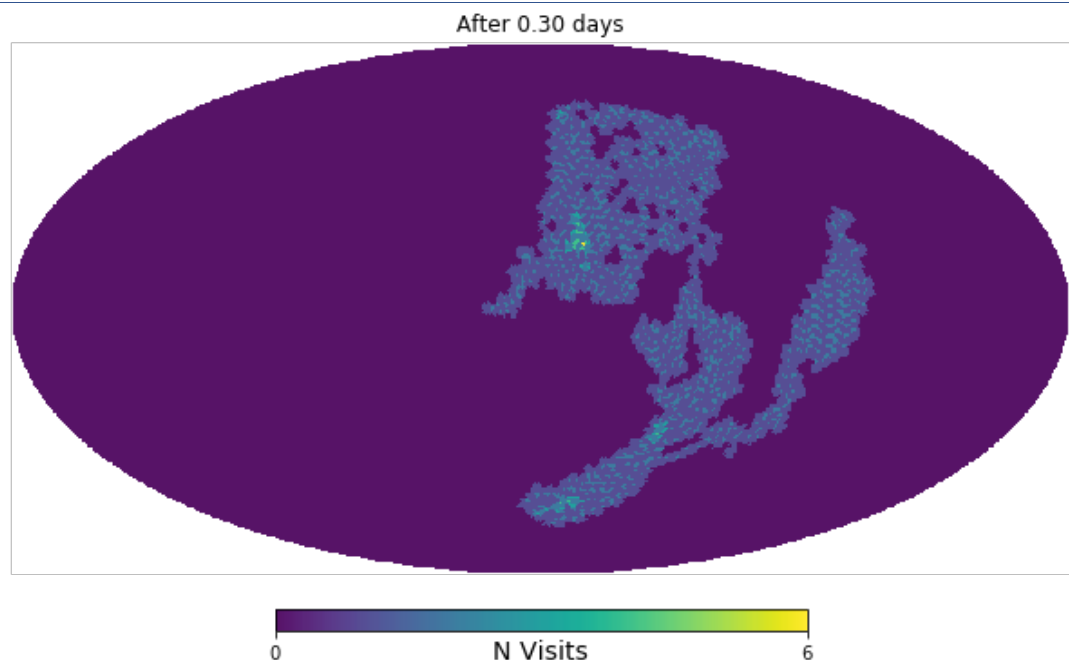
Lots of the work done by Elahe Naghib (Operations Research, Princeton) and Daniel Rothchild (CS, Berkeley)



How do we pick what to observe?

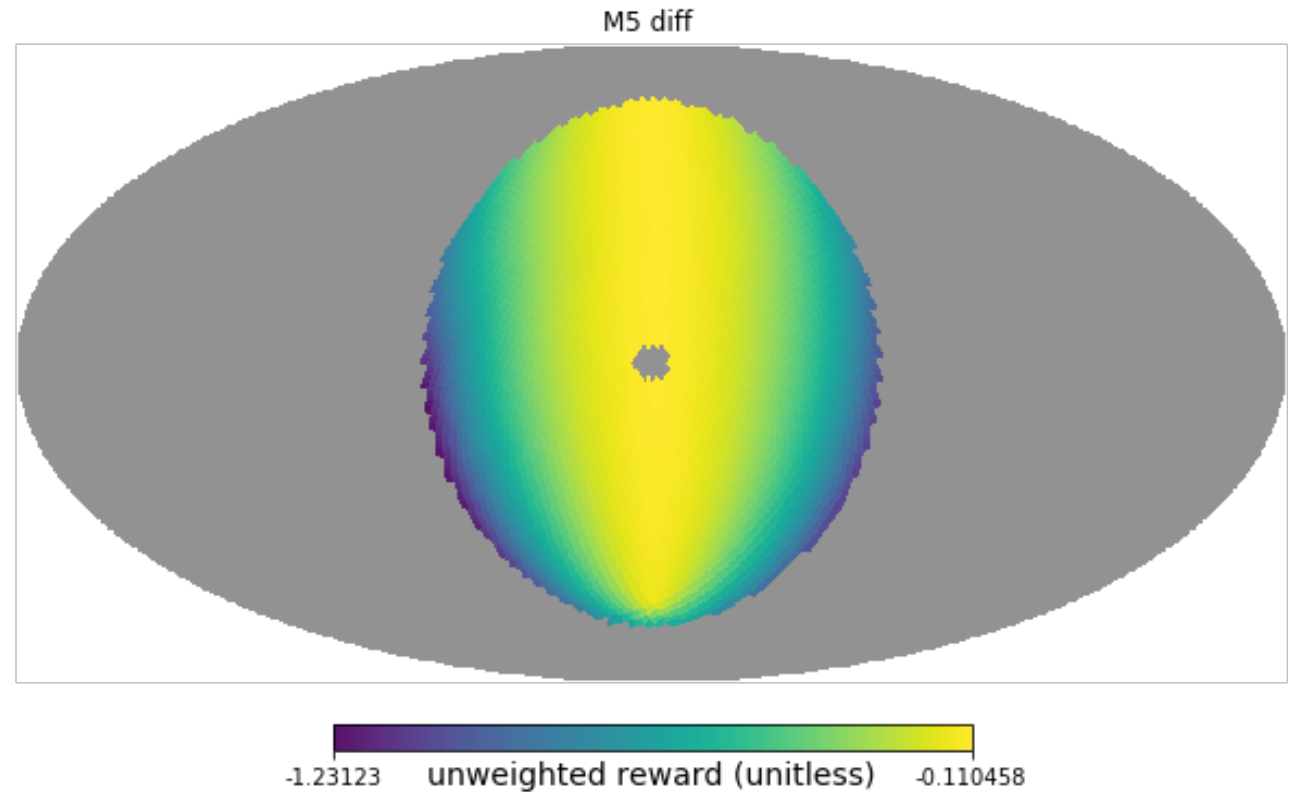
1. What has the least depth degradation (compared to dark time on the meridian)?
2. What area has fallen behind?
3. What's the shortest slewtime?
4. What filter are we in?
5. Mask out bad regions in alt,az space

This let's us pick which healpixel we would most like to observe. To decide where to actually point the telescope, we randomize the tessellation every night (so there is a changing nightly mapping of HEALpix to RA,dec telescope pointing).



Example r-band only survey

5-sigma limiting
depth Compared to
meridian dark time,
using model sky
brightness, airmass,
and seeing



Target goal map, where regions
are defined in terms of the **ratios**.

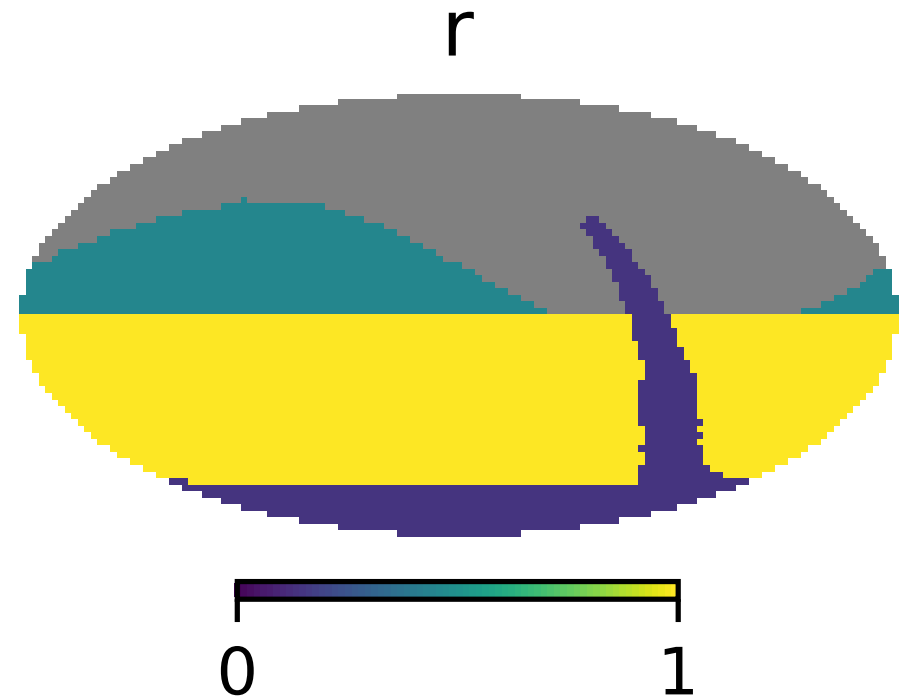
WFD = 1

NES = 0.46

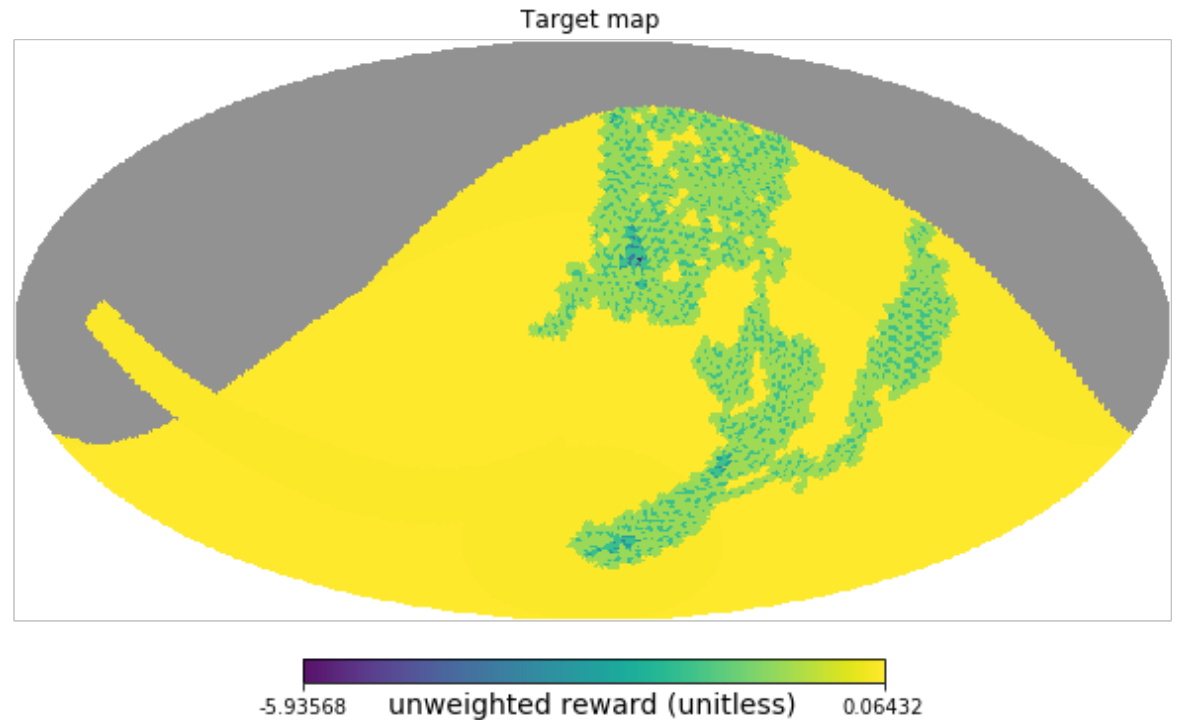
SCP = 0.15

GP = 0.15

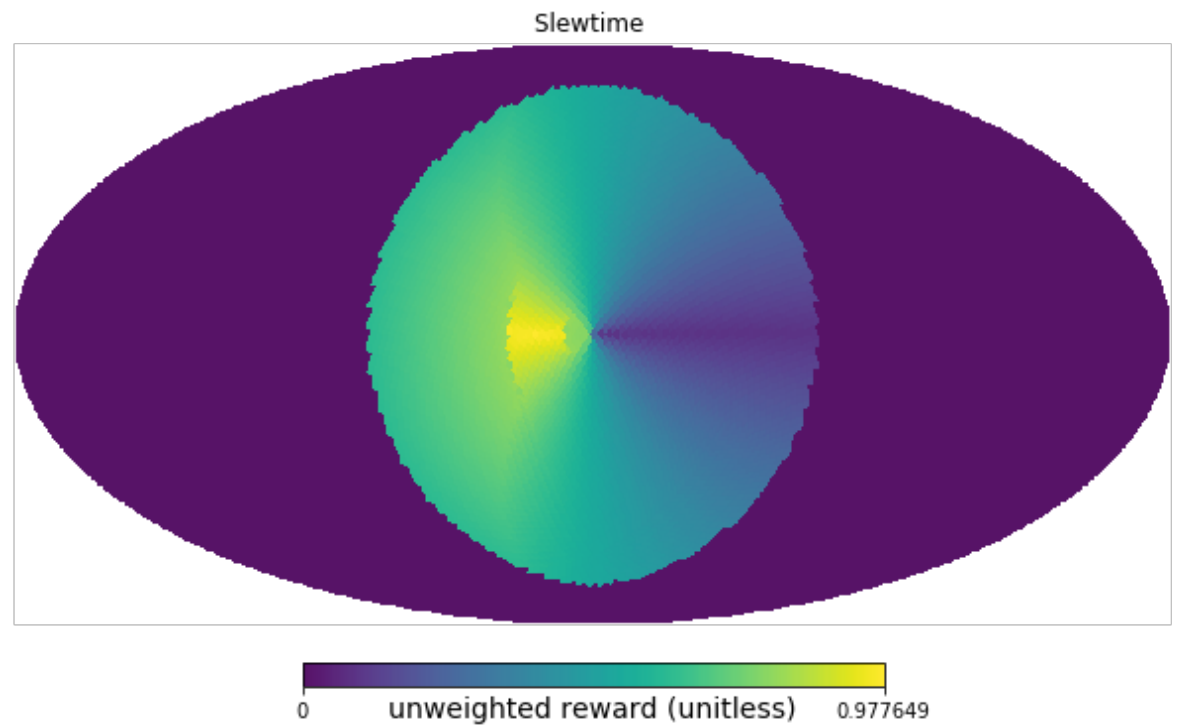
(No running out of observations,
can make a map per filter)



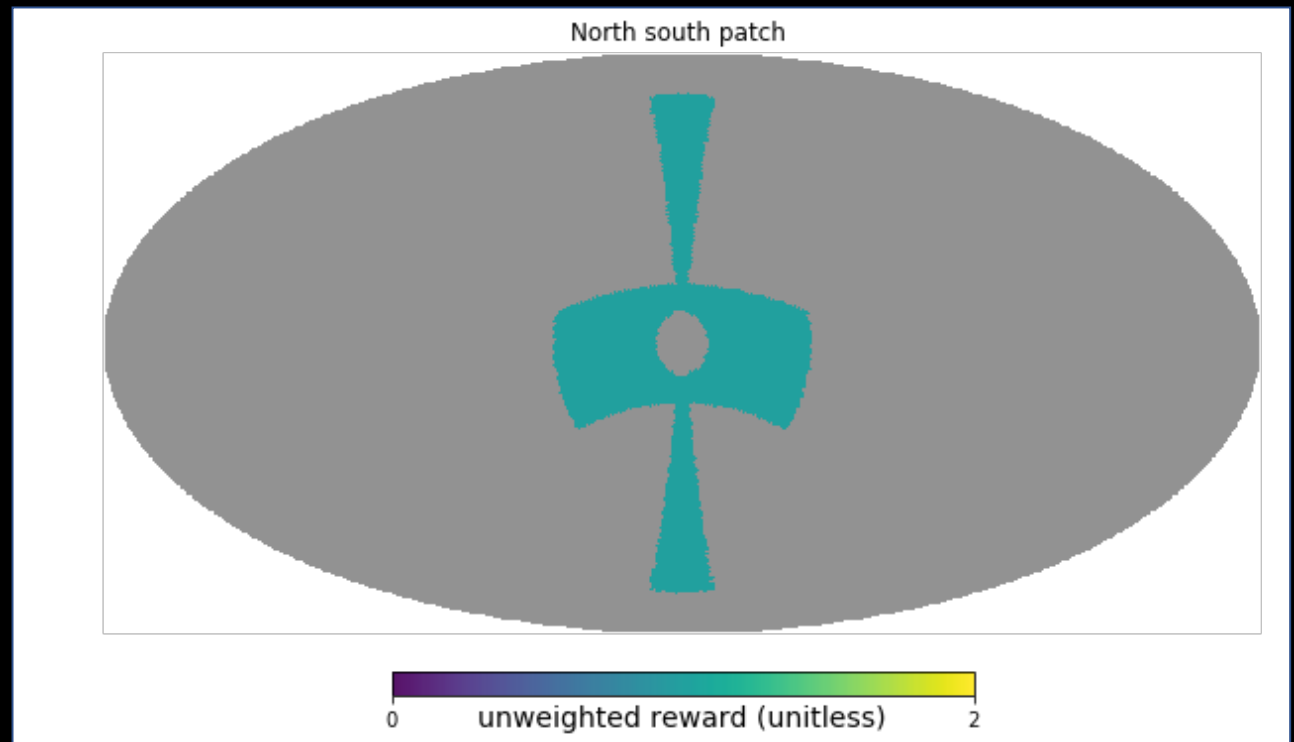
Combine the target goal map with the map of completed observations to generate a map of how far behind HEALpixels are



Time to slew to any
HEALpixels

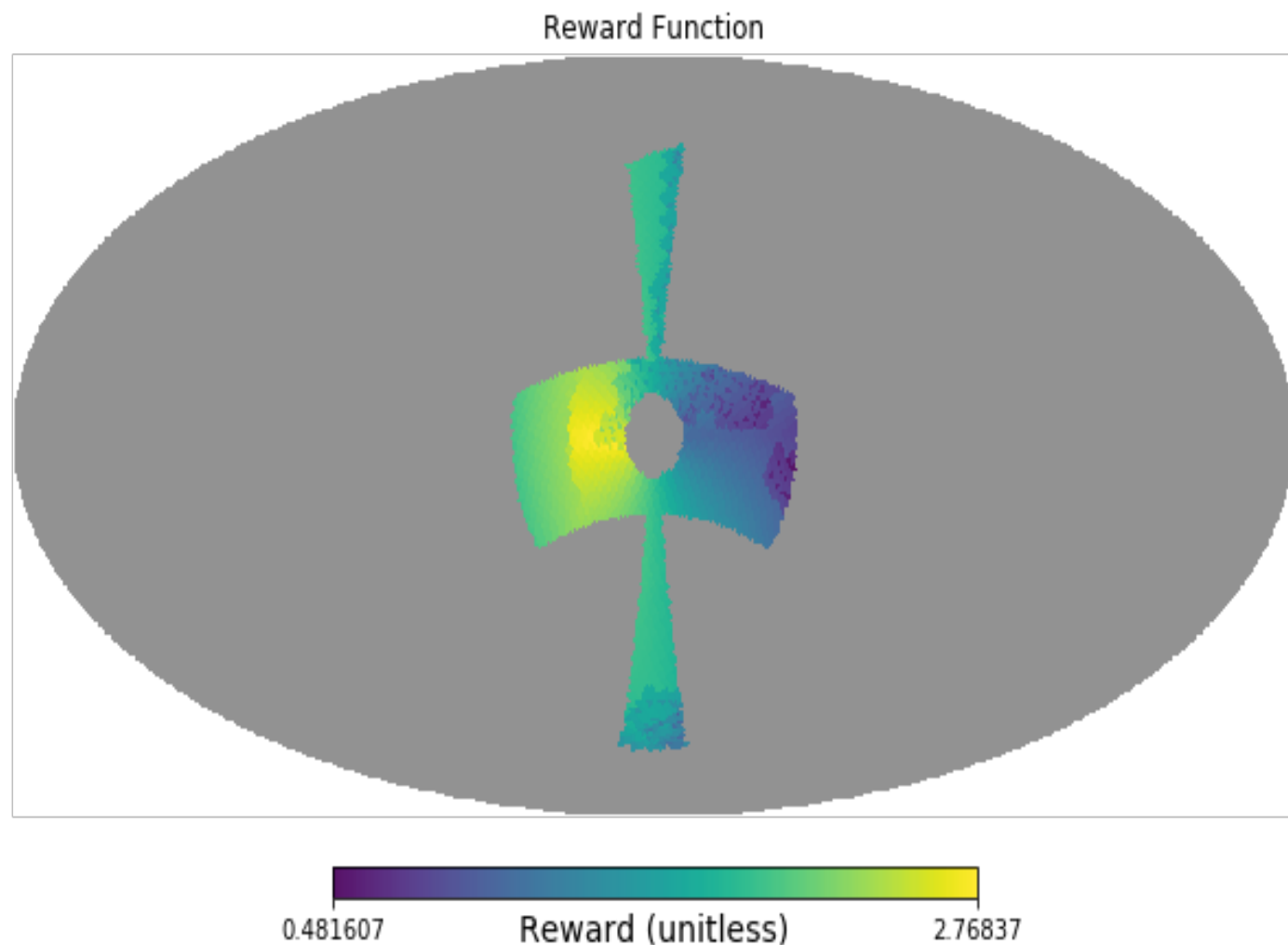


Mask in alt,az space to
prevent getting stuck near
zenith or airmass limit



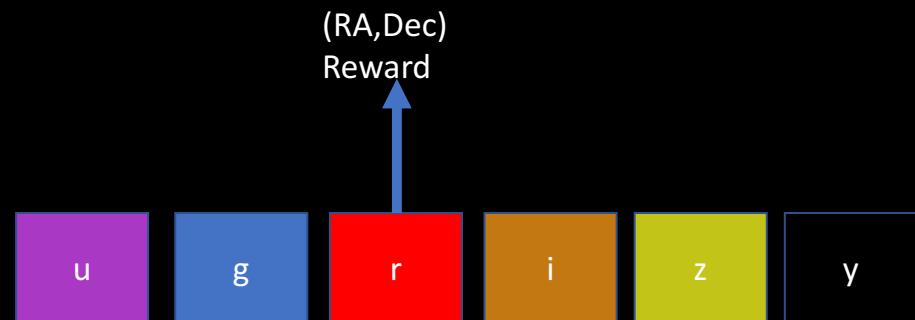
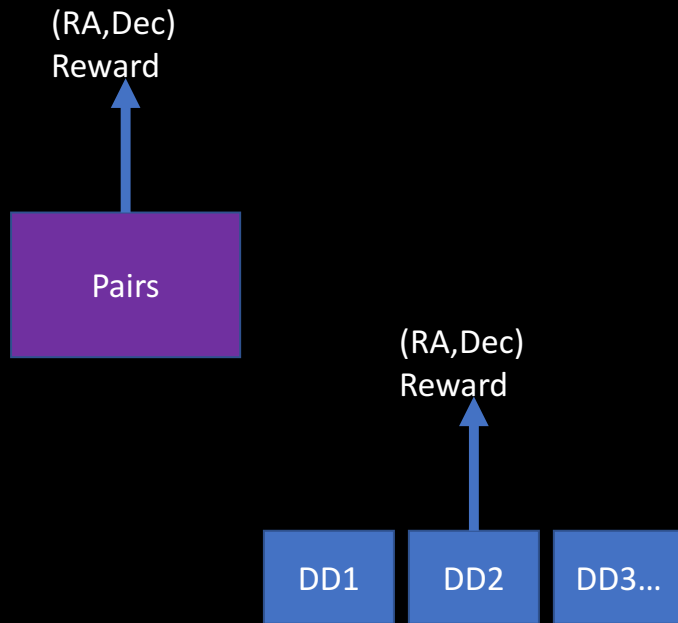
Final reward function is a linear combination of the basis functions.

Single filter survey only has 3* free parameters!

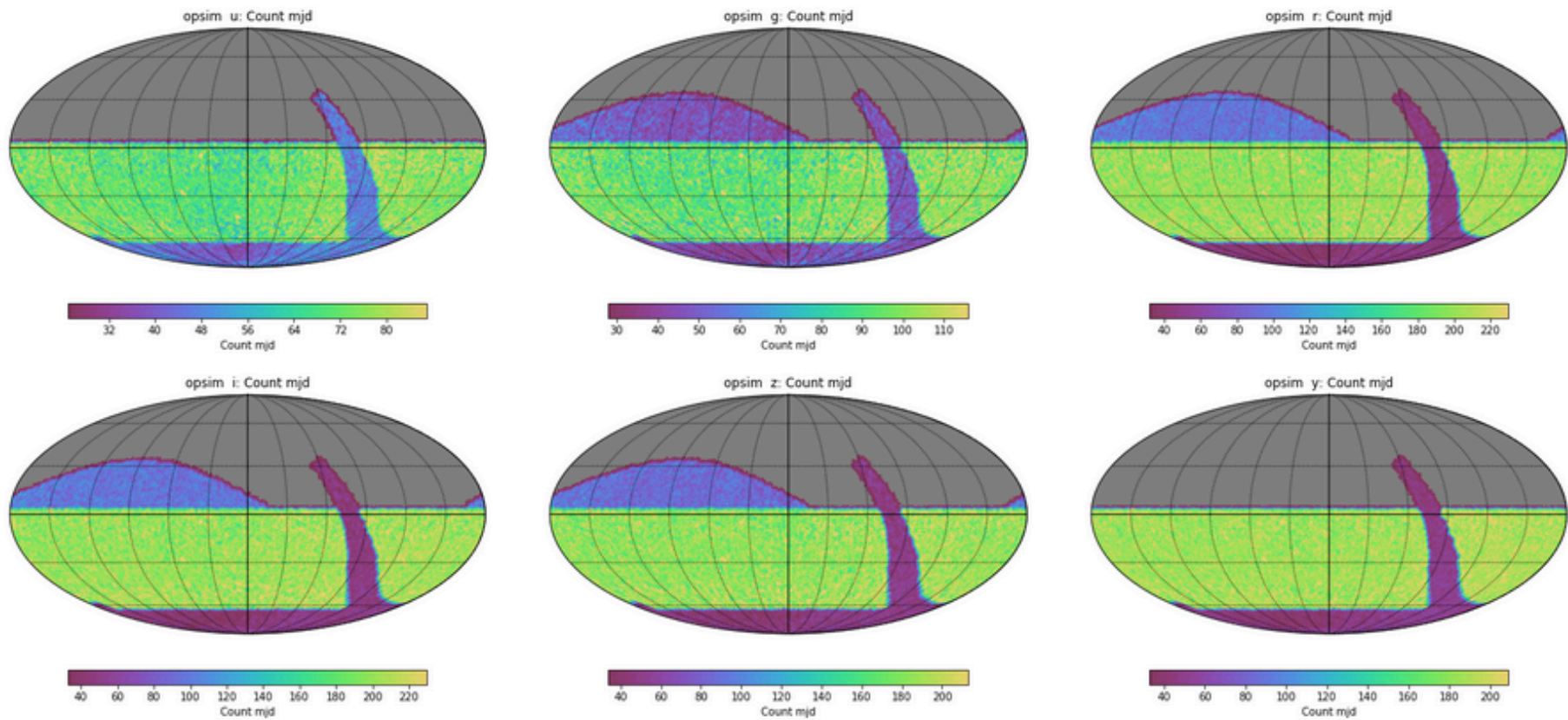


Full Survey

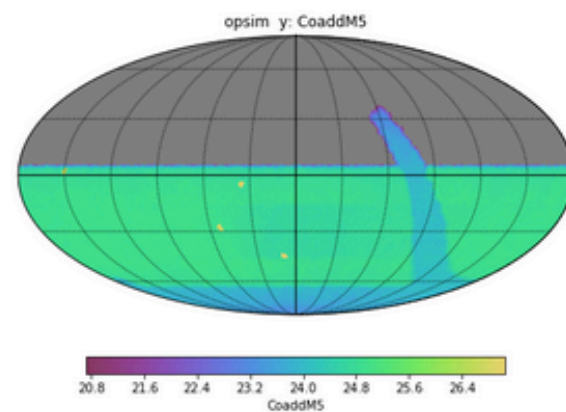
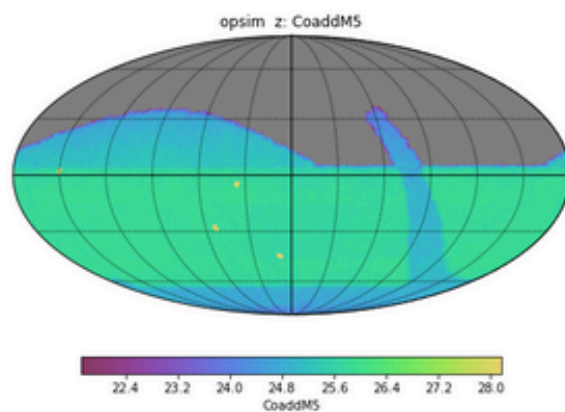
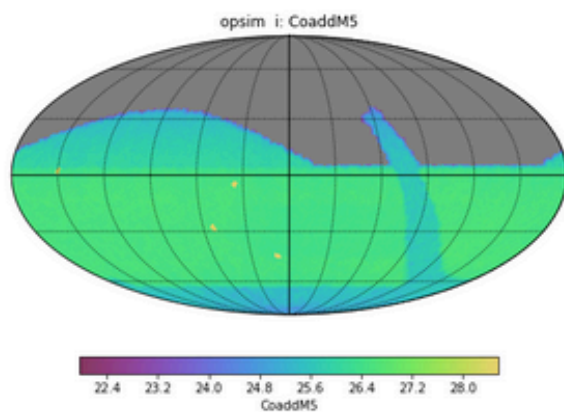
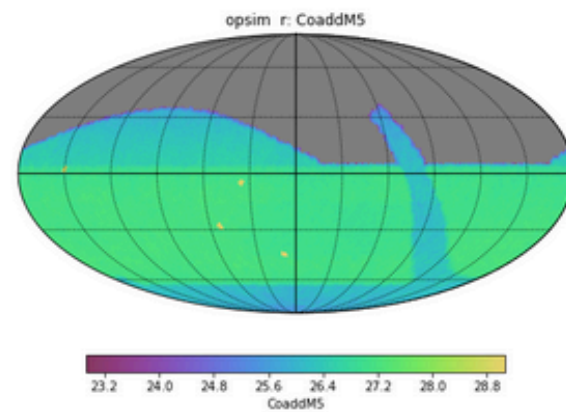
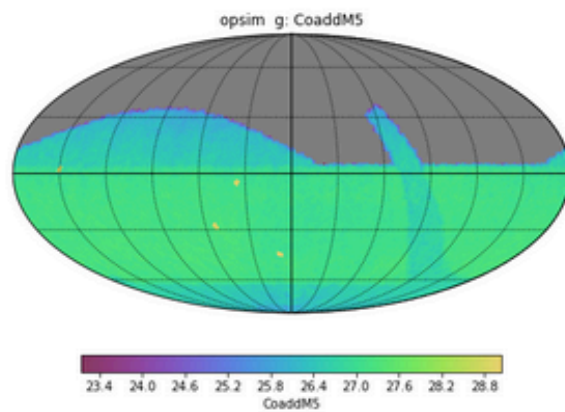
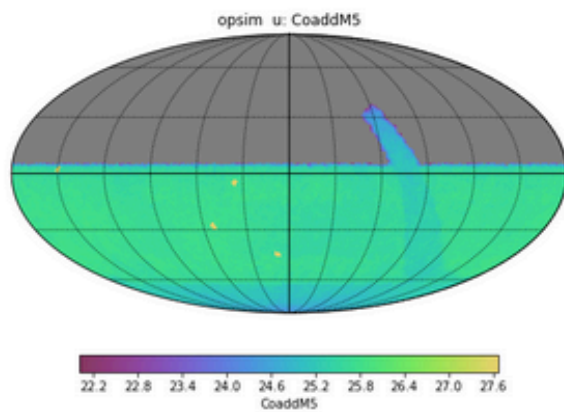
Requesting an observation, each “survey” returns target(s) and a reward value. Highest reward value gets selected.



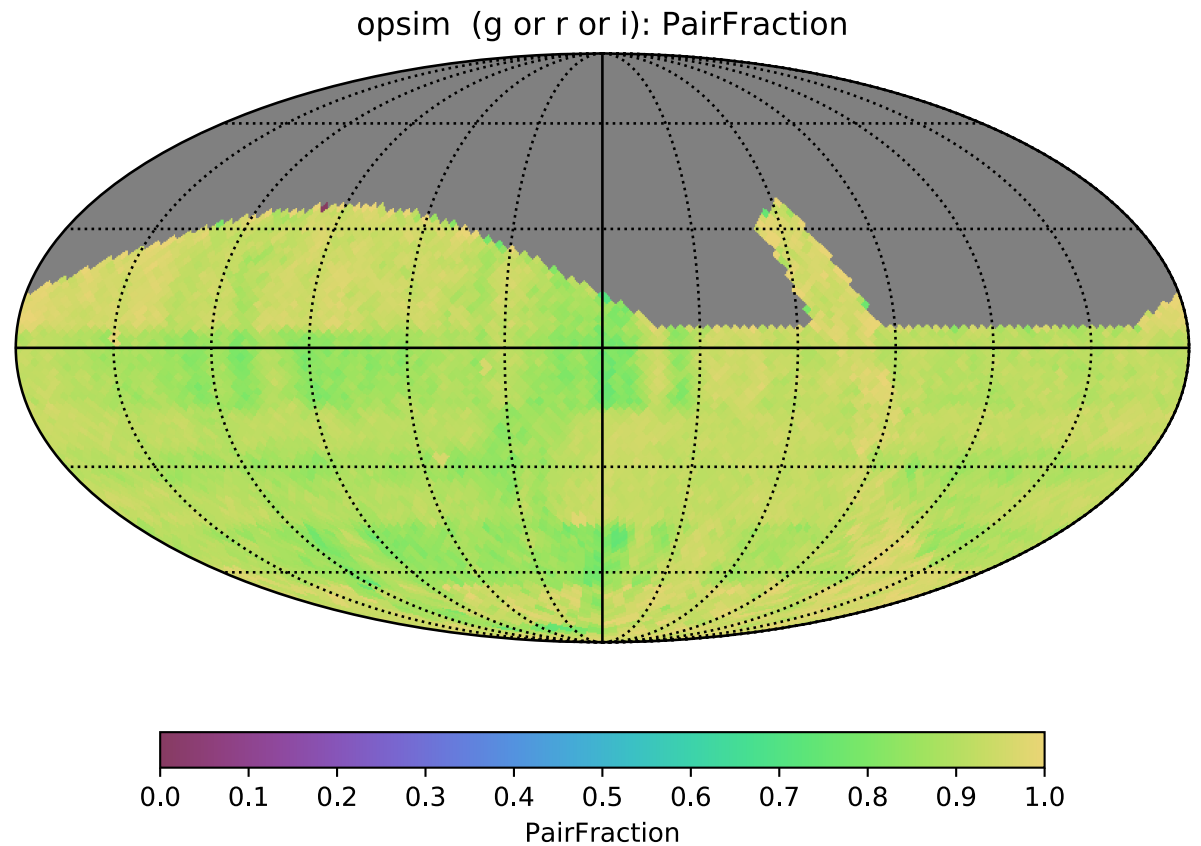
Number of observations in each filter



Coadded Depth

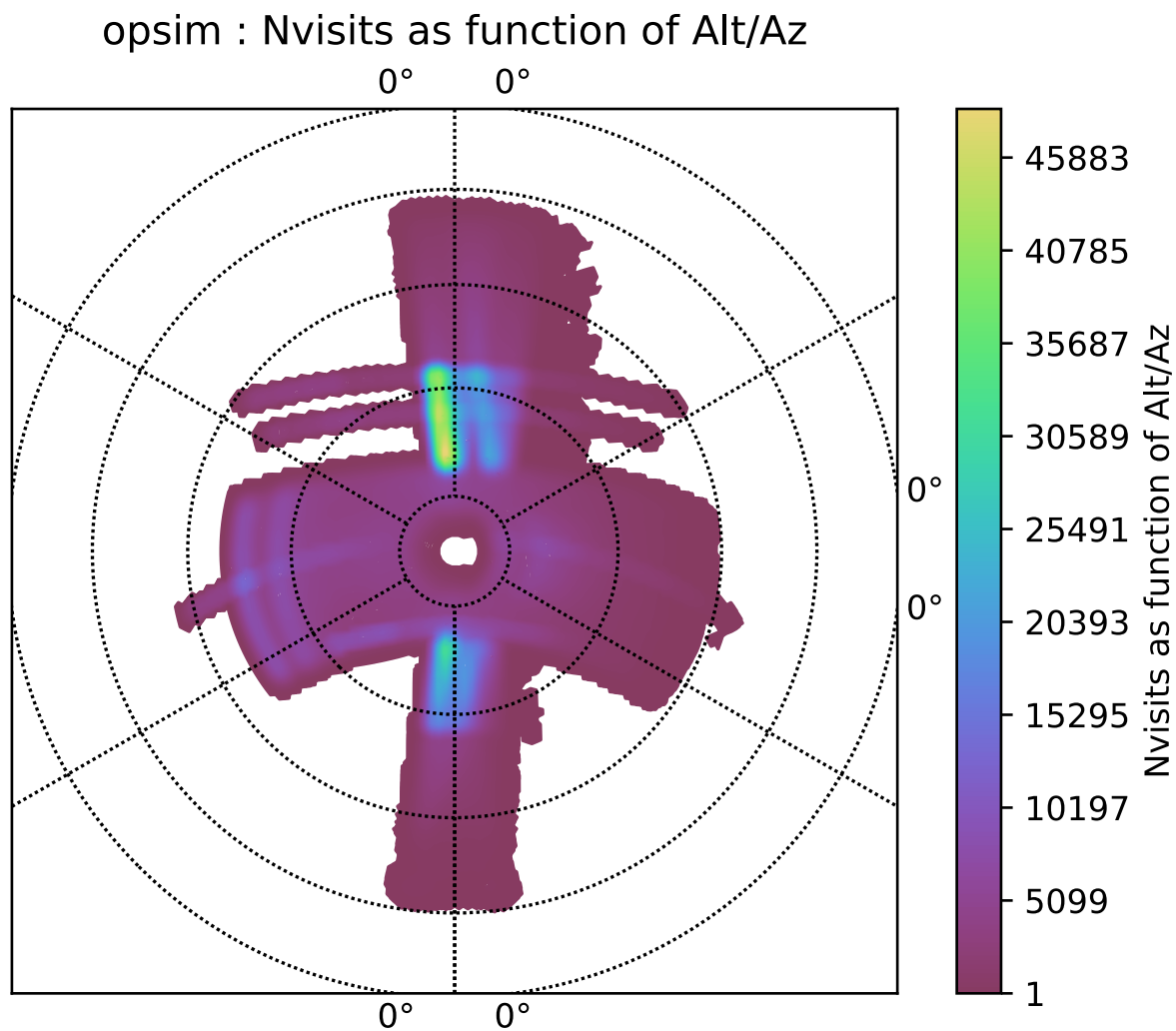


Fraction of observations
that are in a pair that can
be used for solar system
object detection (peak
around 90%)



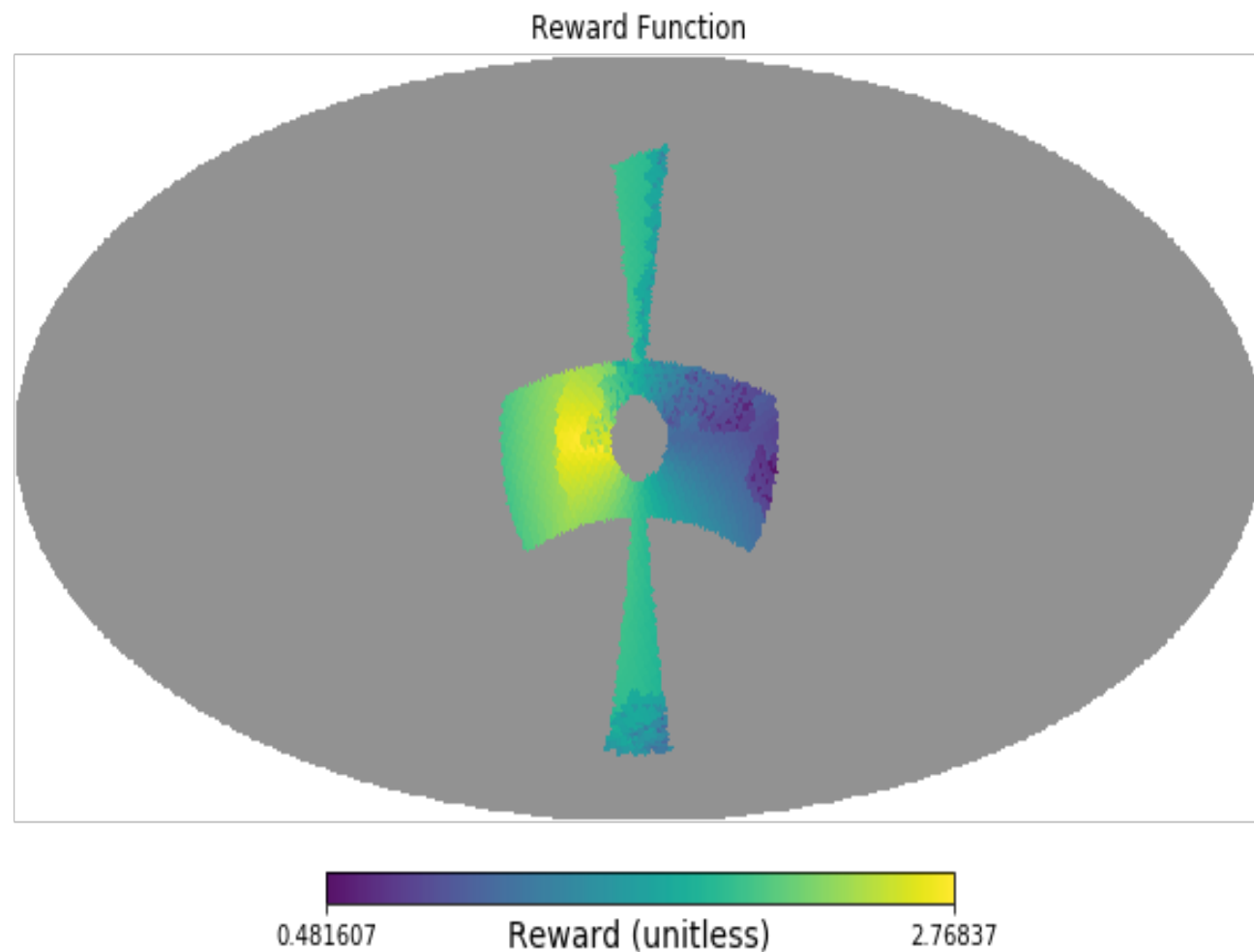
Shutter open 76%
of possible time.
(82% would be the
maximum with no
filter changes and
minimum slew
times)

Observing mostly
near meridian.

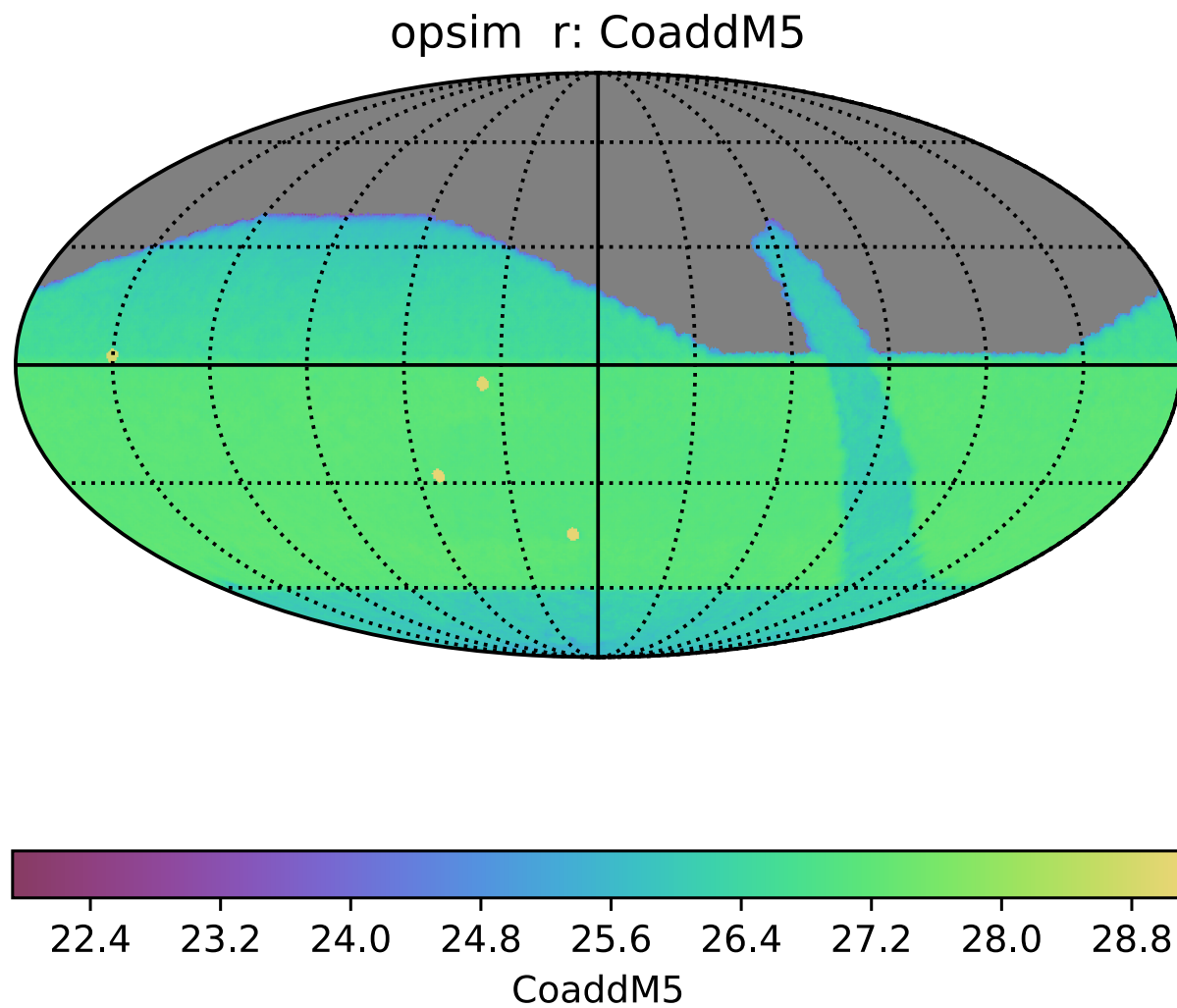


The reward function shows the best HEALpixel to observe.

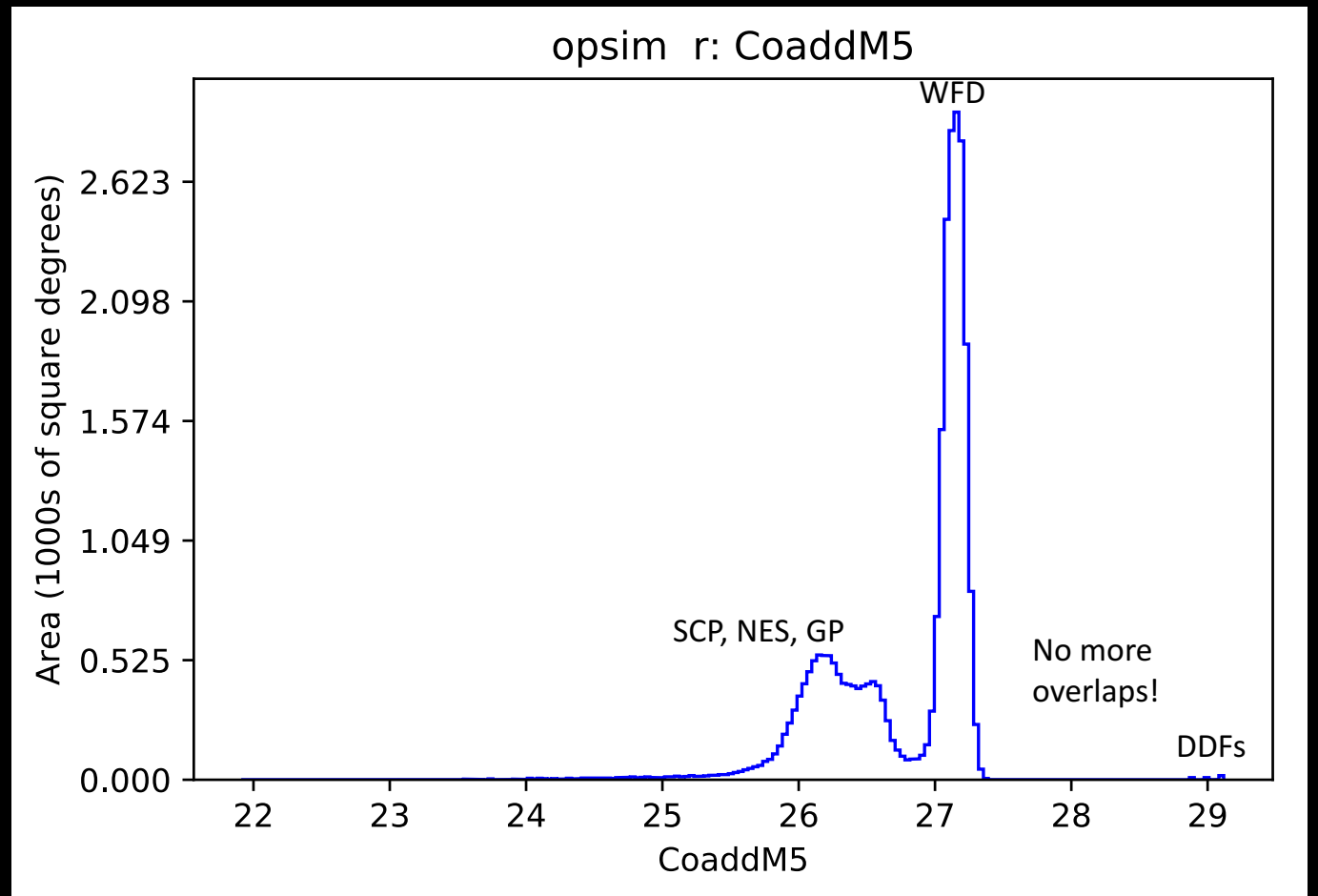
Each night, we randomly rotate the tessellation and create a new mapping of HEALpix to pointings



No more over-exposed
overlap regions

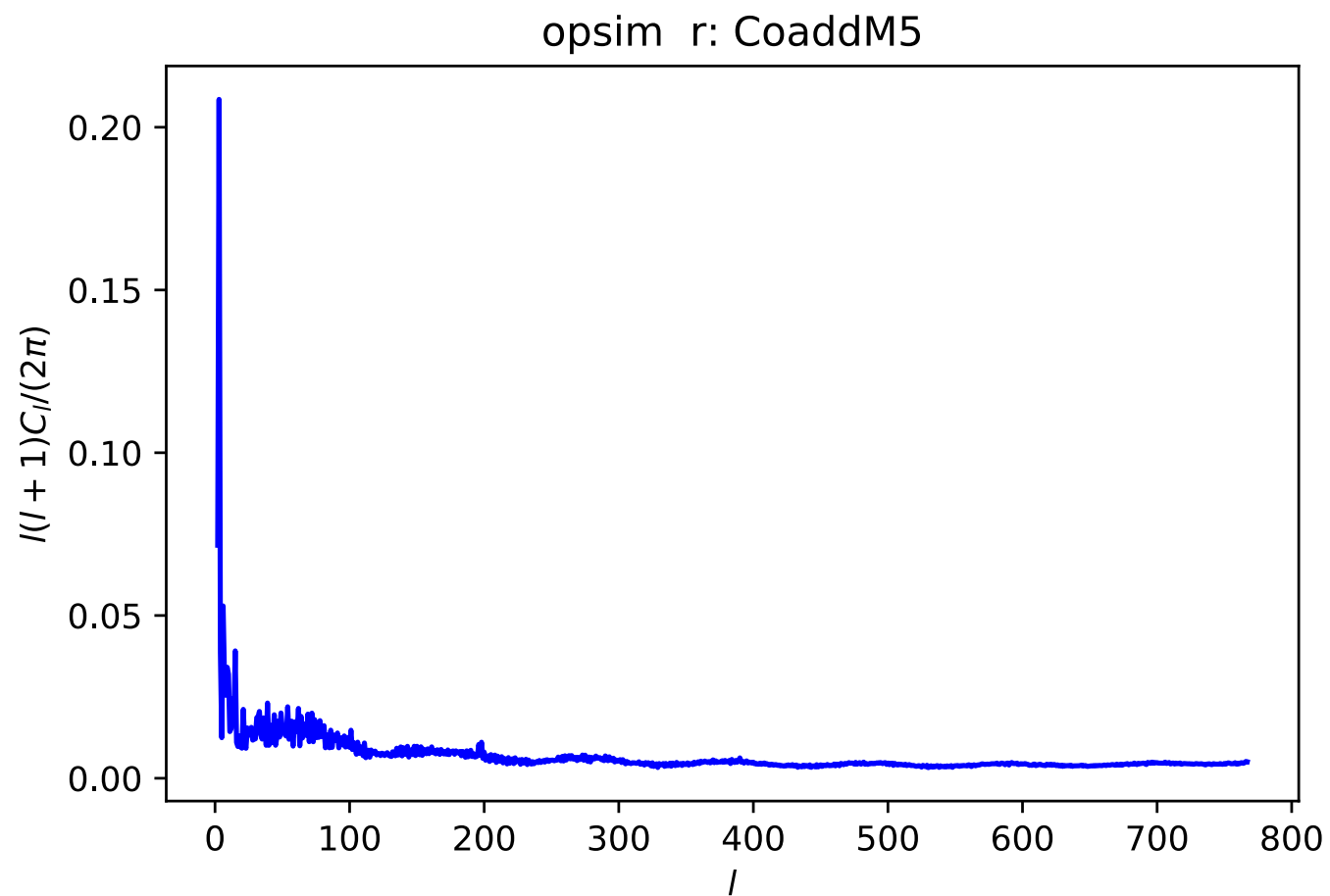


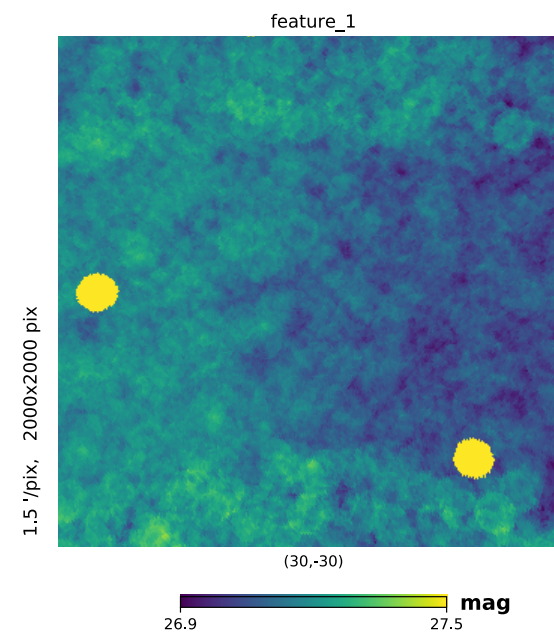
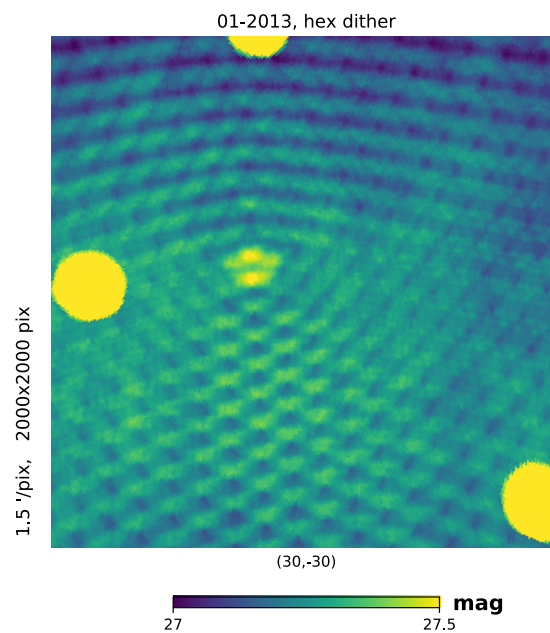
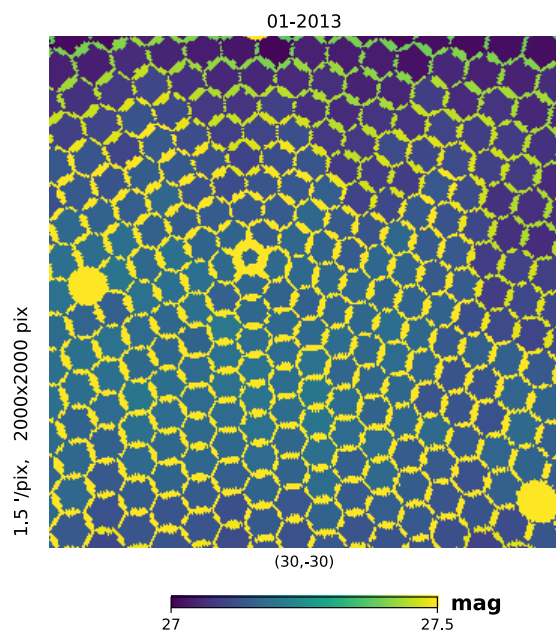
Feature based
scheduler



Feature based
scheduler.

No more giant power
spike at 1 degree
($l=180$)



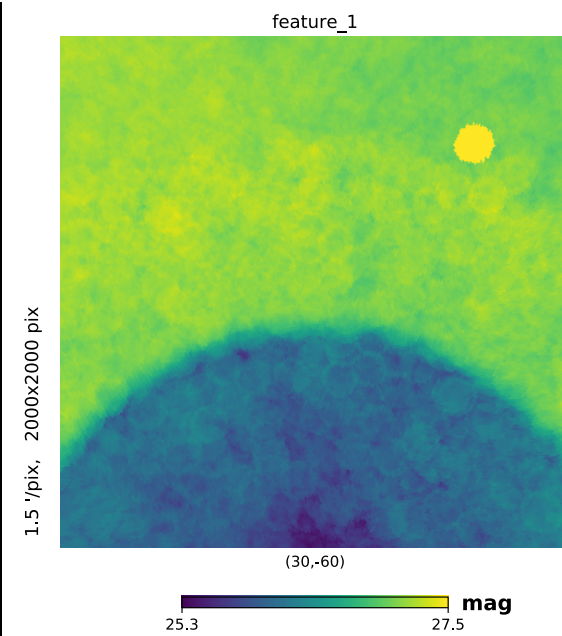
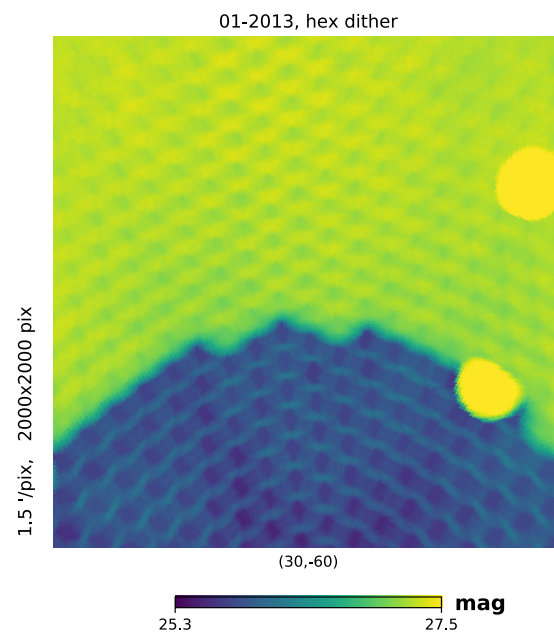
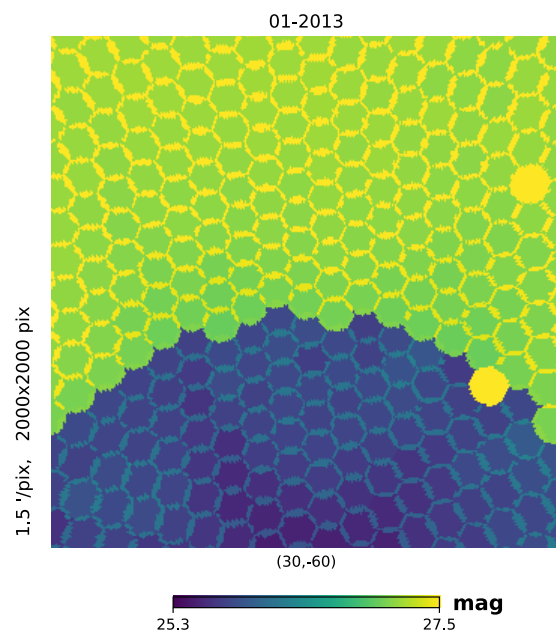


Undithered

Sub field dithers

Feature Based scheduler

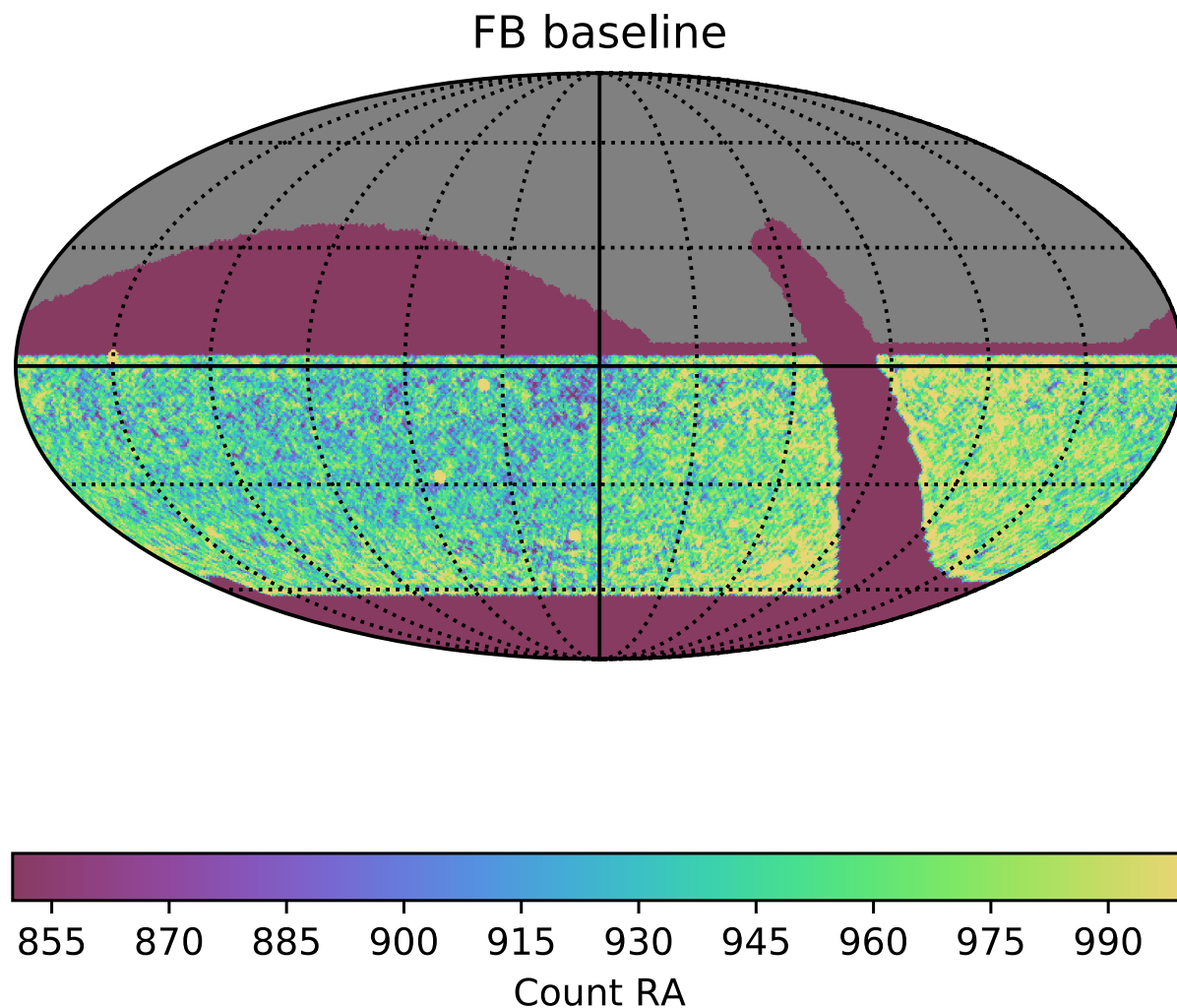
So far, we have only tried to keep the number of visits even. We could try to keep coadded depth even.



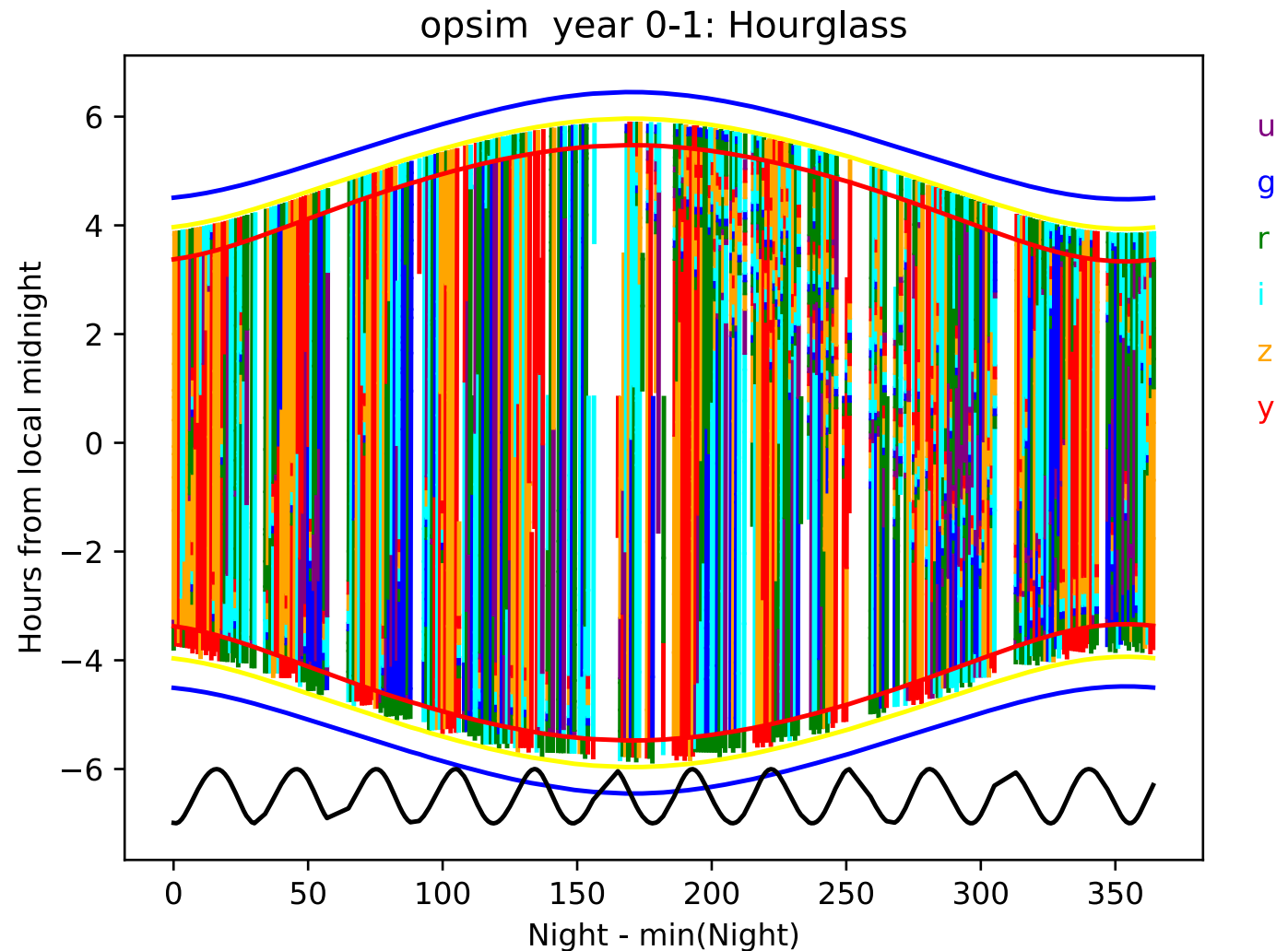
Another view at the WFD-SCP border

We are a little under-subscribed in winter, little over-subscribed in summer.

May want to adjust WFD area a bit to compensate. Make sure we don't pile-up DD fields too much in RA.



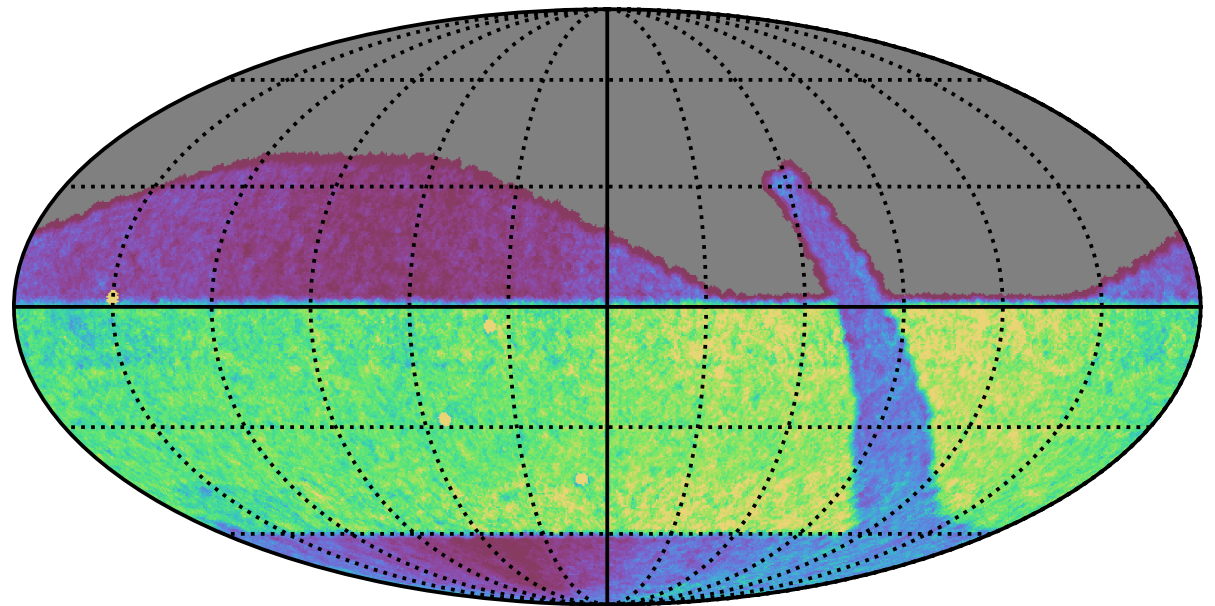
Filter choice currently gets driven largely by lunar phase. So we get several days of red followed by days of blue.



Some initial experiments in rolling cadence
(proof of concept, I did these in a day)

WFD~90-100
observations/yr

baseline, year 1



30

45

60

75

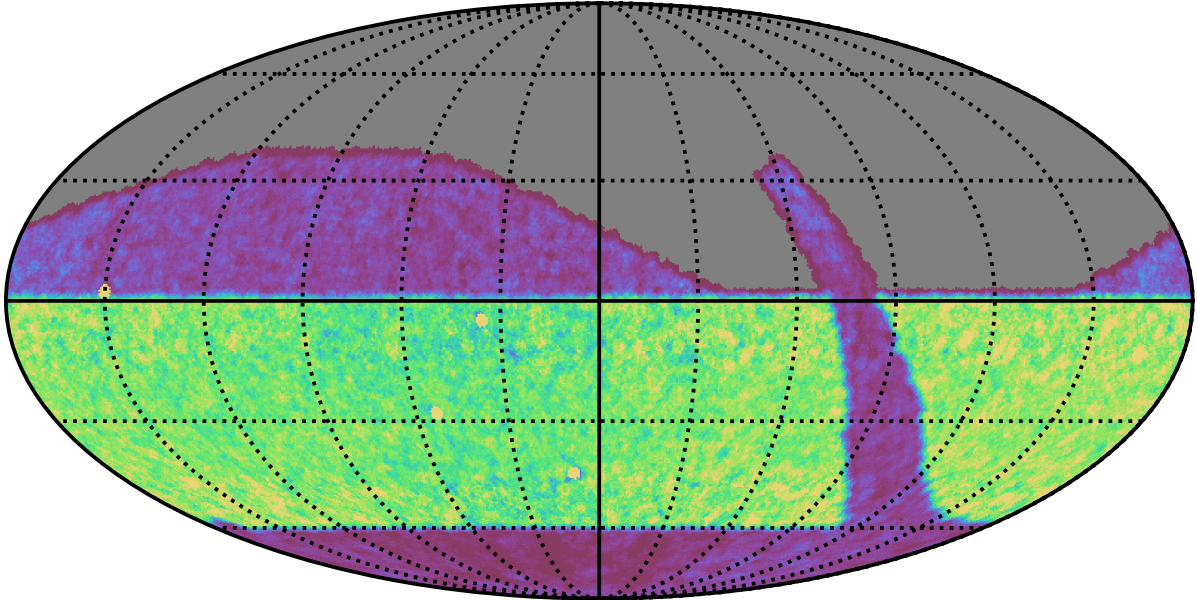
90

105

120

Count RA

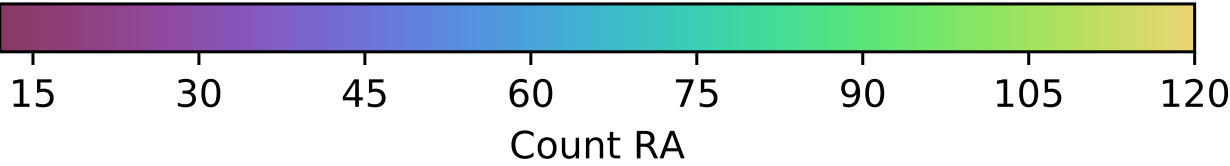
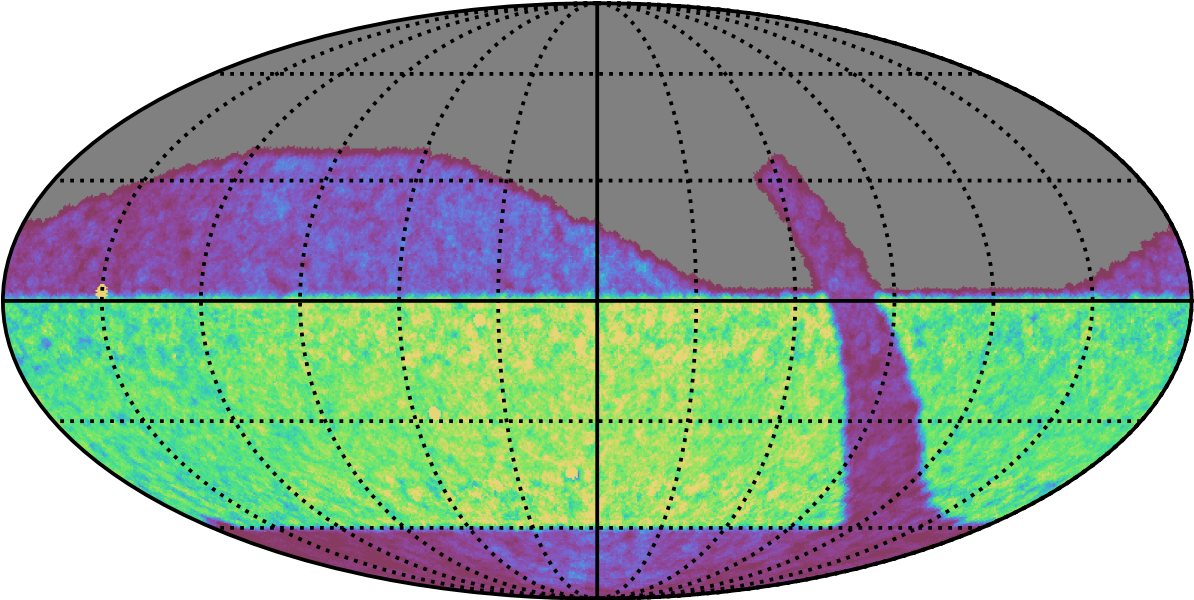
baseline, year 2



15 30 45 60 75 90 105 120

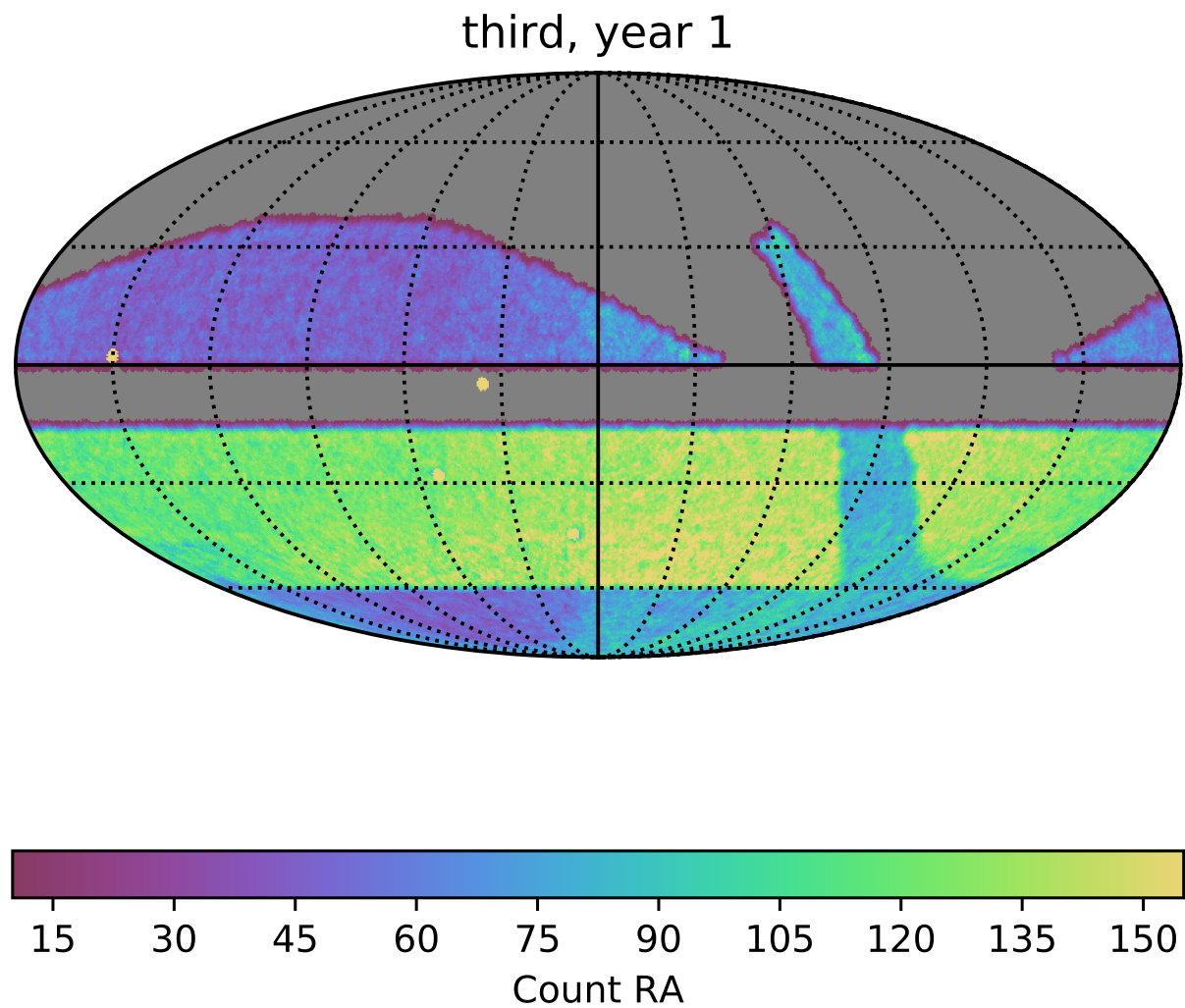
Count RA

baseline, year 3

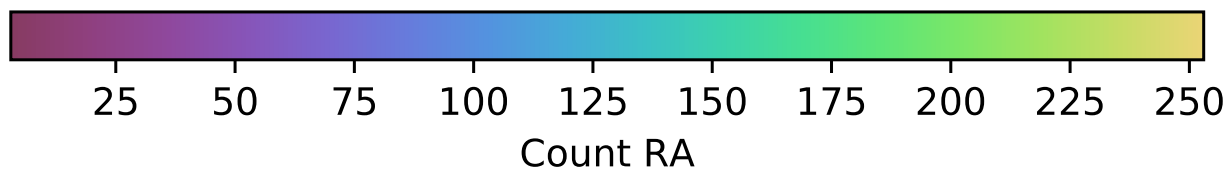
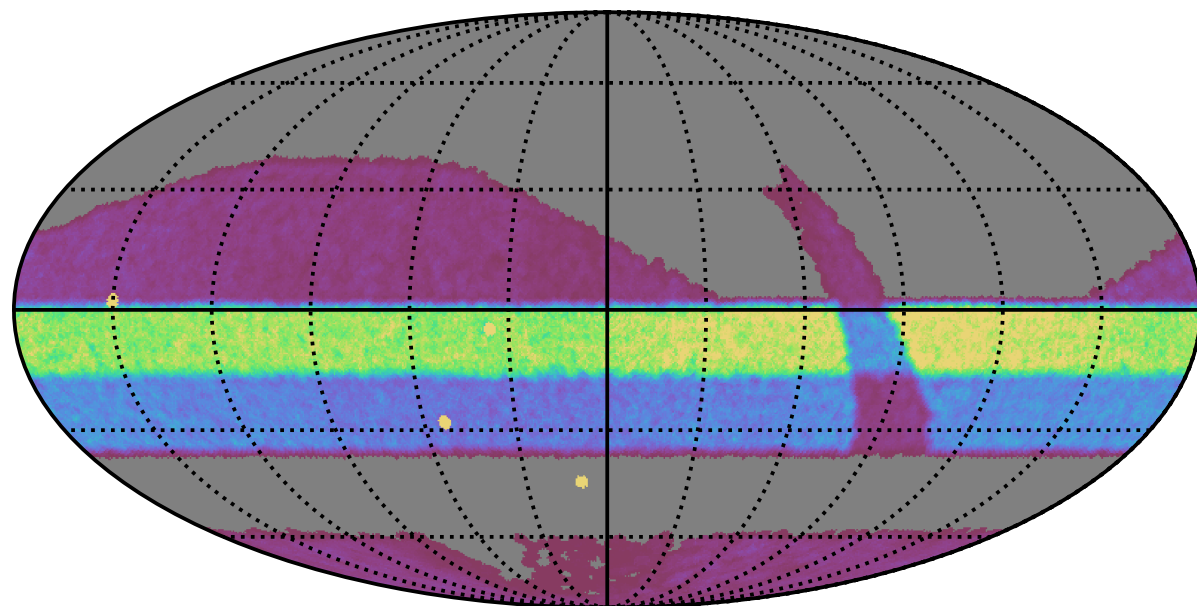


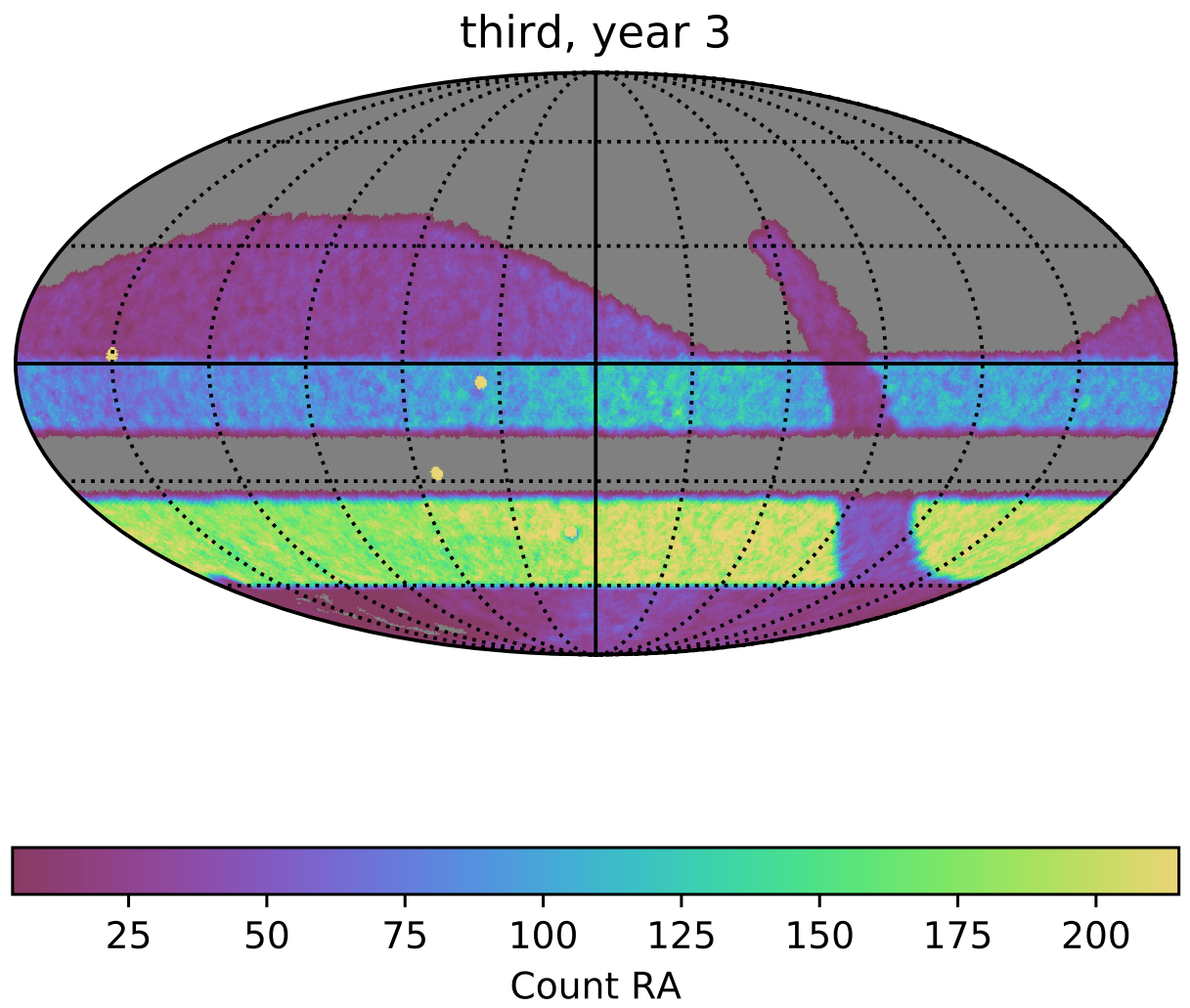
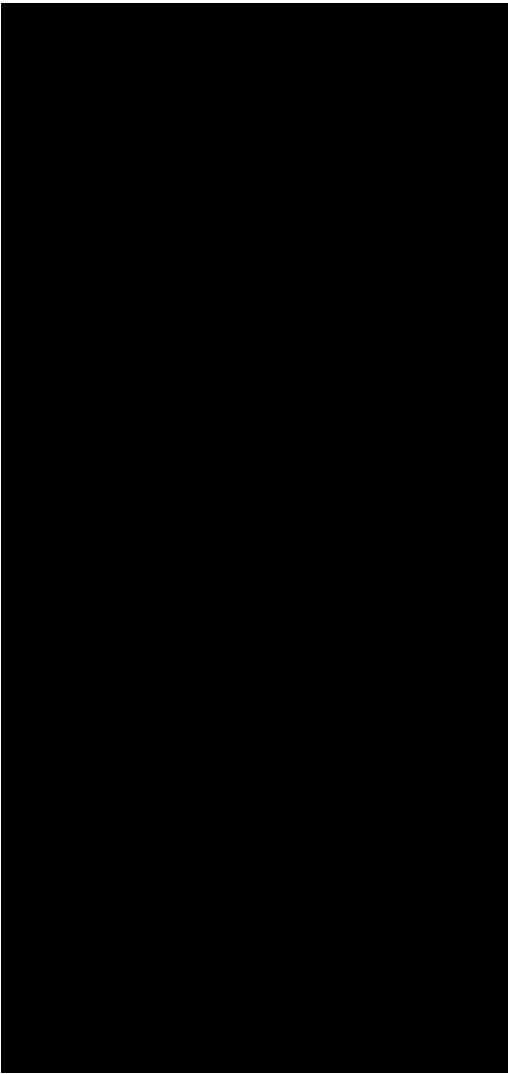
Rolling Cadence:
Mask 1/3 of WFD
region each year.

Now 120-130
observations/yr

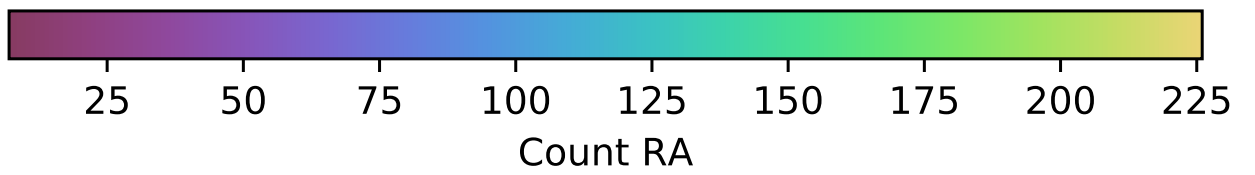
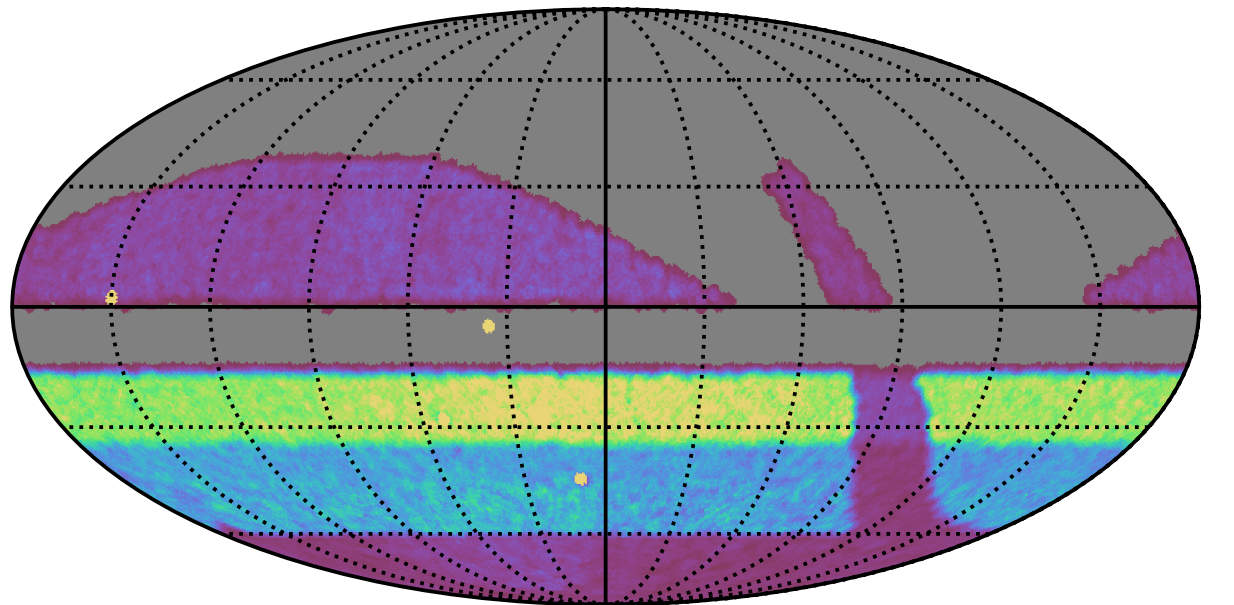


third, year 2



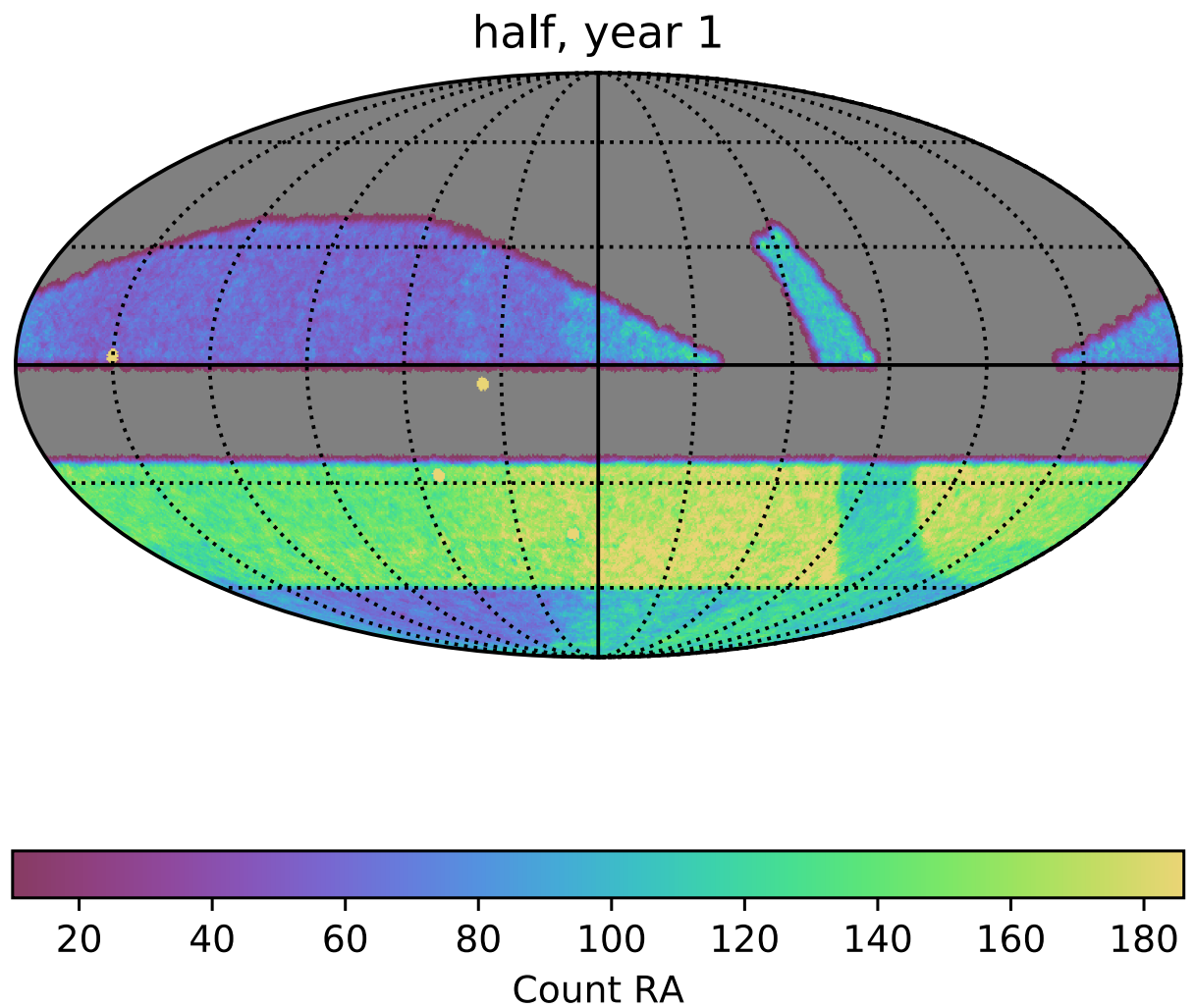


third, year 4

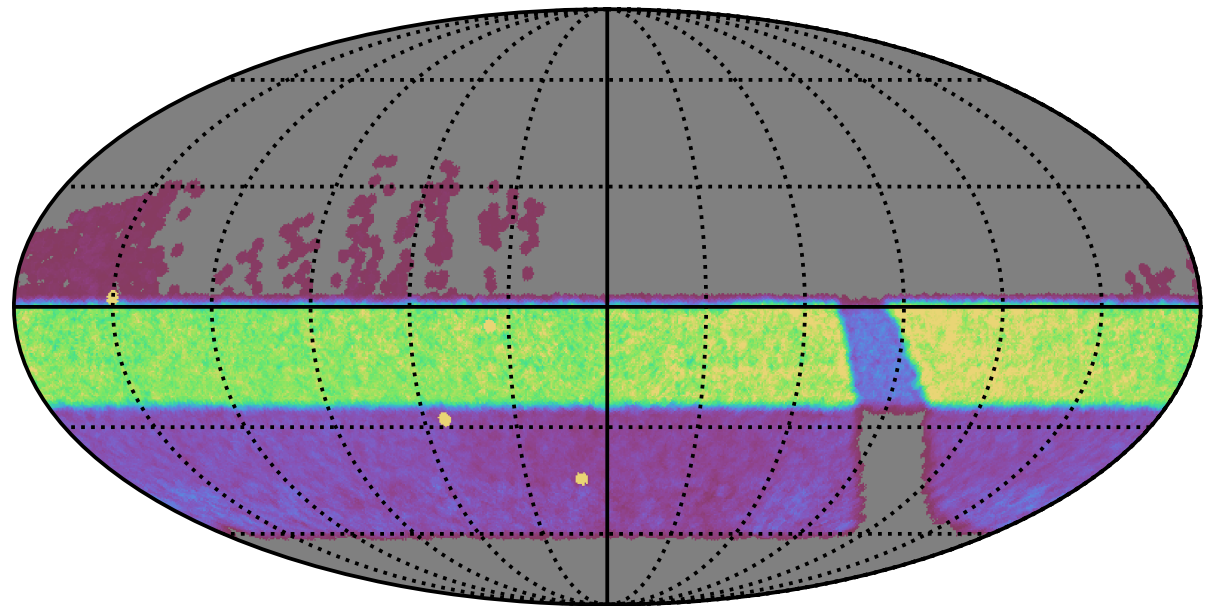


Now Mask half the
WFD region. Then
unmask next year

140-160 obs/yr



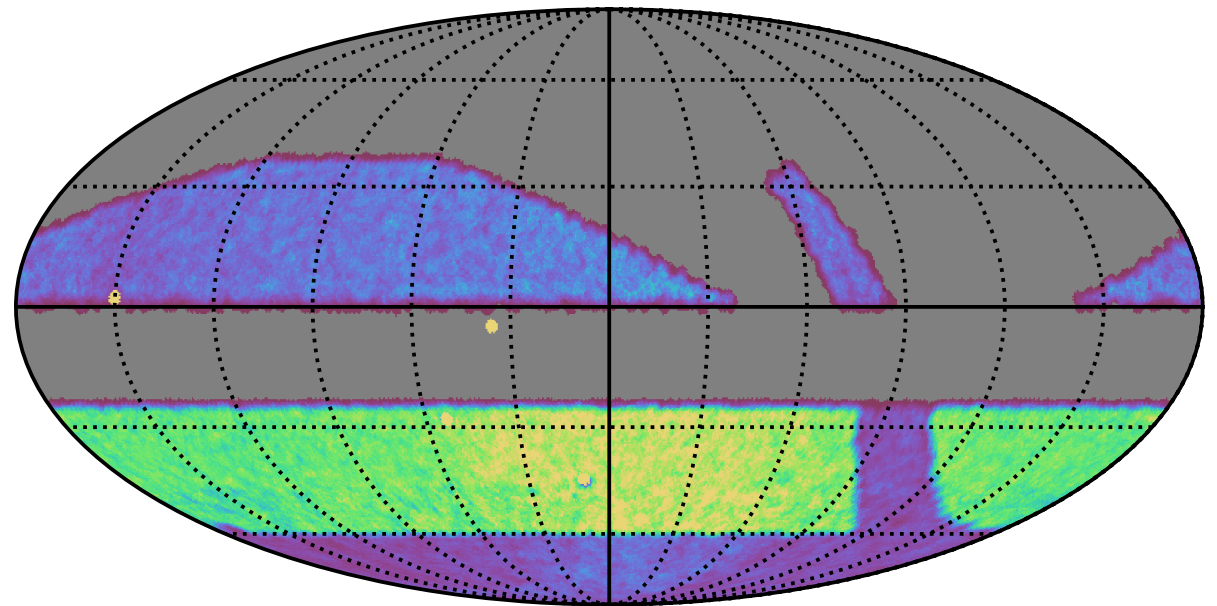
half, year 2



25 50 75 100 125 150 175 200

Count RA

half, year 3



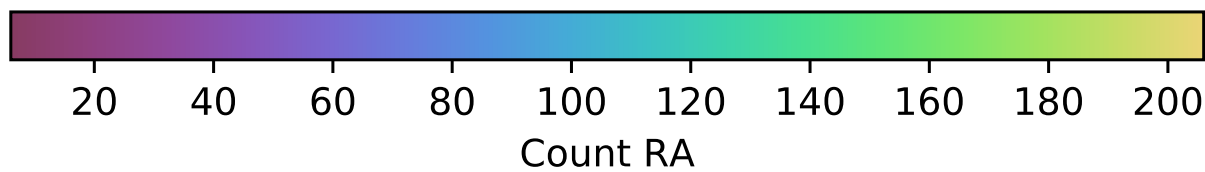
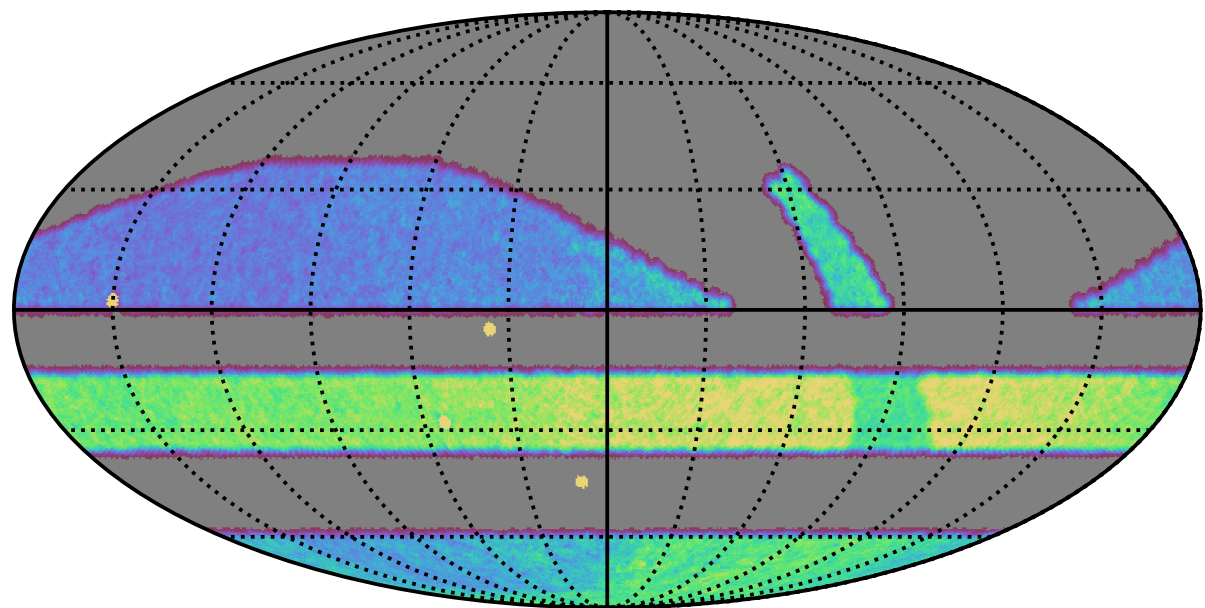
20 40 60 80 100 120 140 160 180

Count RA

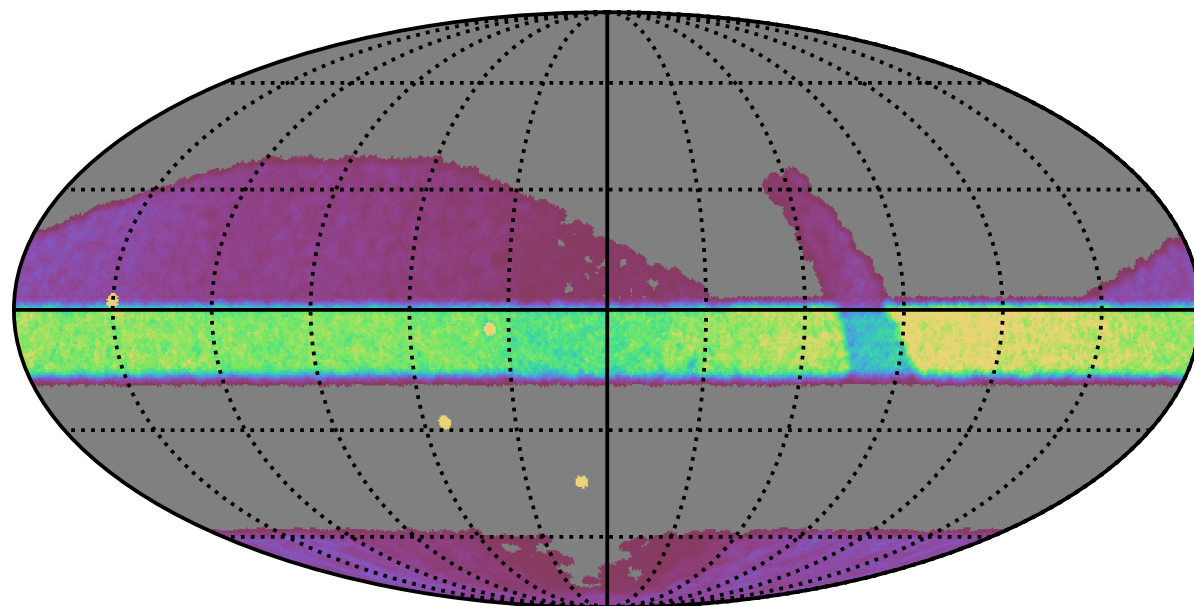
Mask 2/3 of WFD

160-180/yr

two thirds, year 1



two thirds, year 2



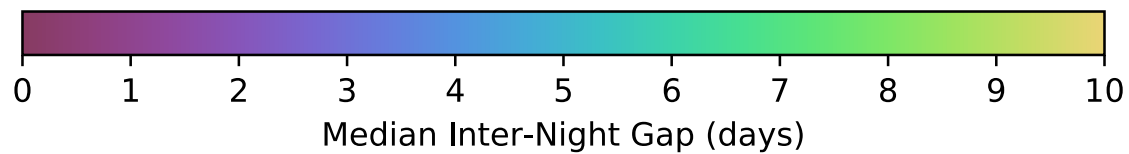
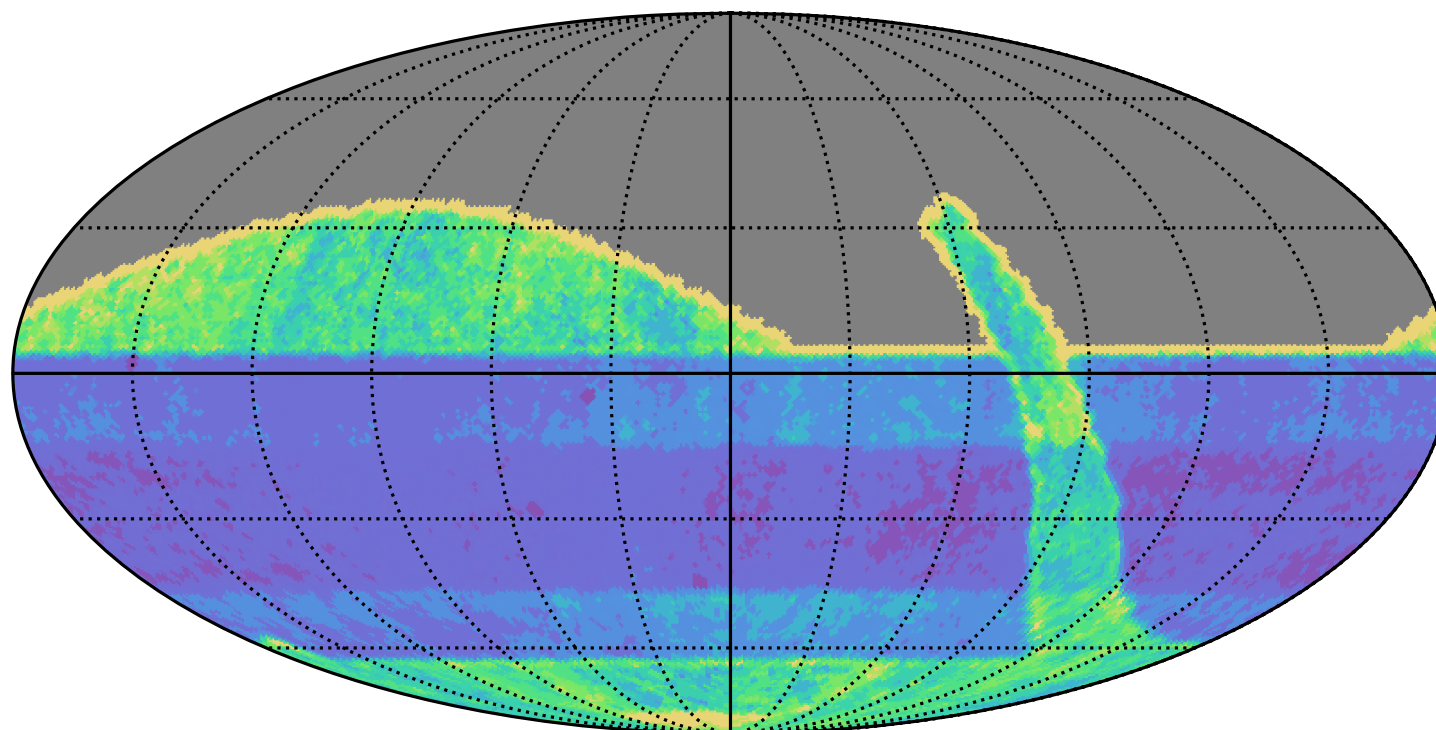
40 80 120 160 200 240 280 320

Count RA

Baseline:

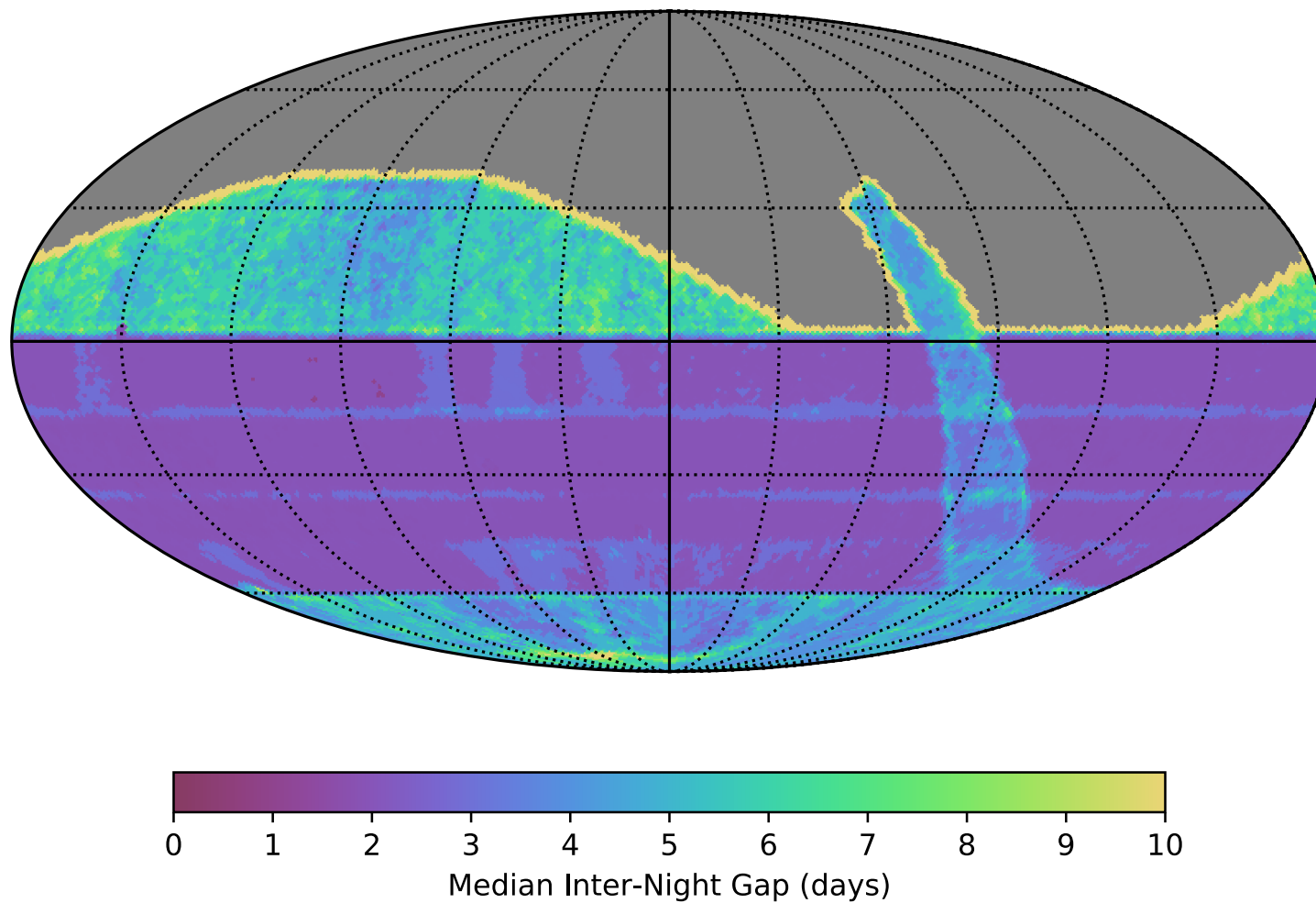
Median
inter-night
gap of 3.95
days

opsim all bands: Median Inter-Night Gap



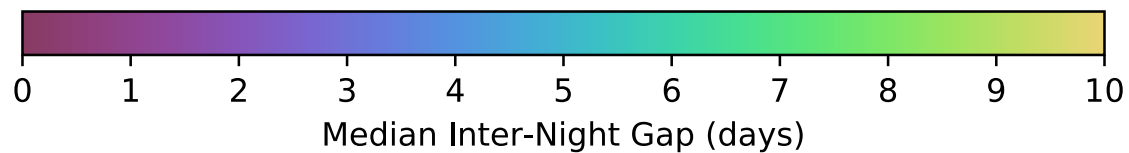
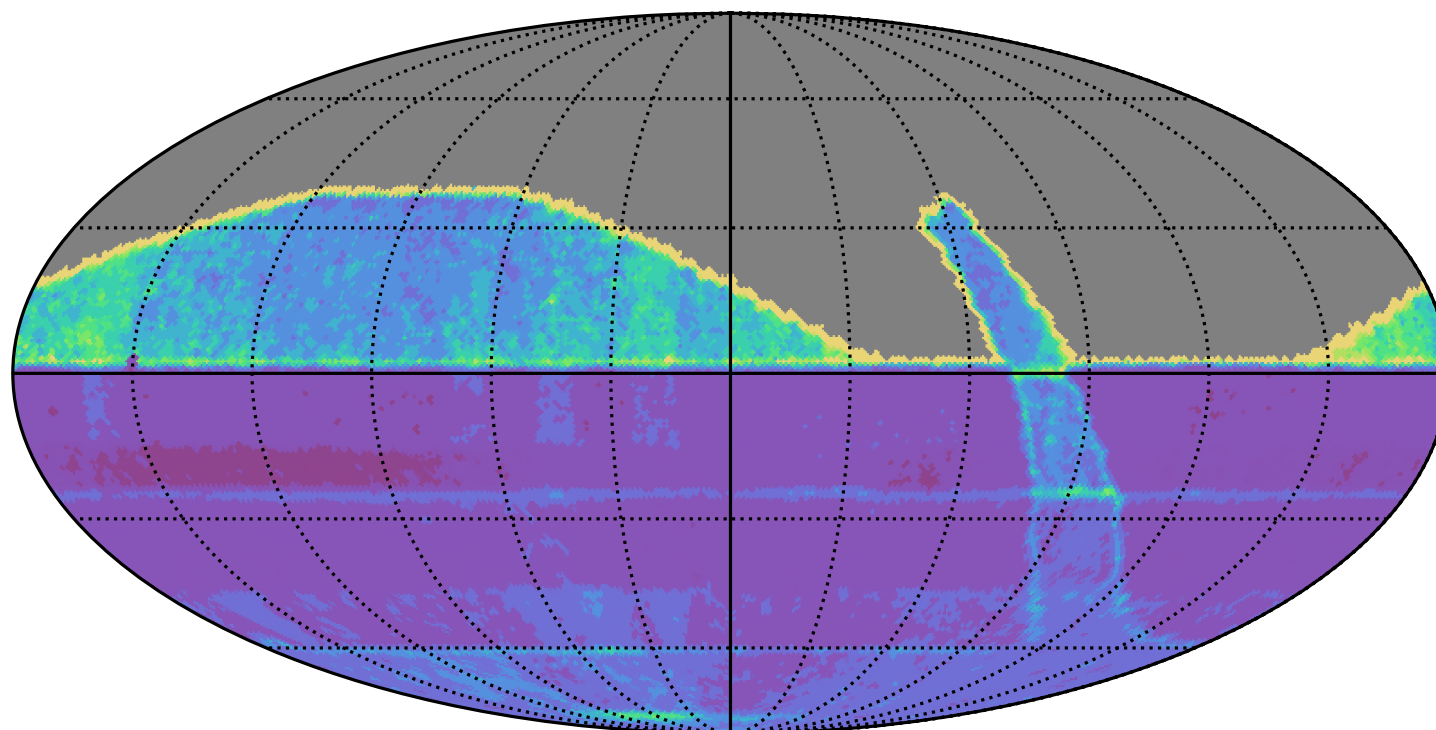
1/3 rolling
cadence
Median gap
of 2.06 days

opsim all bands: Median Inter-Night Gap



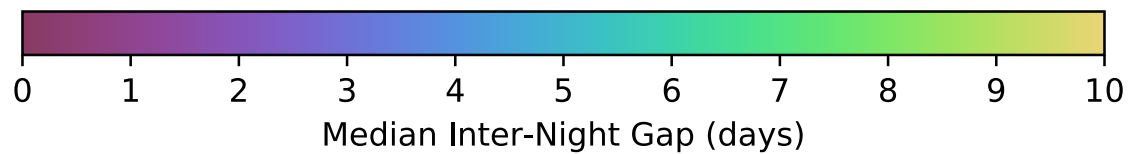
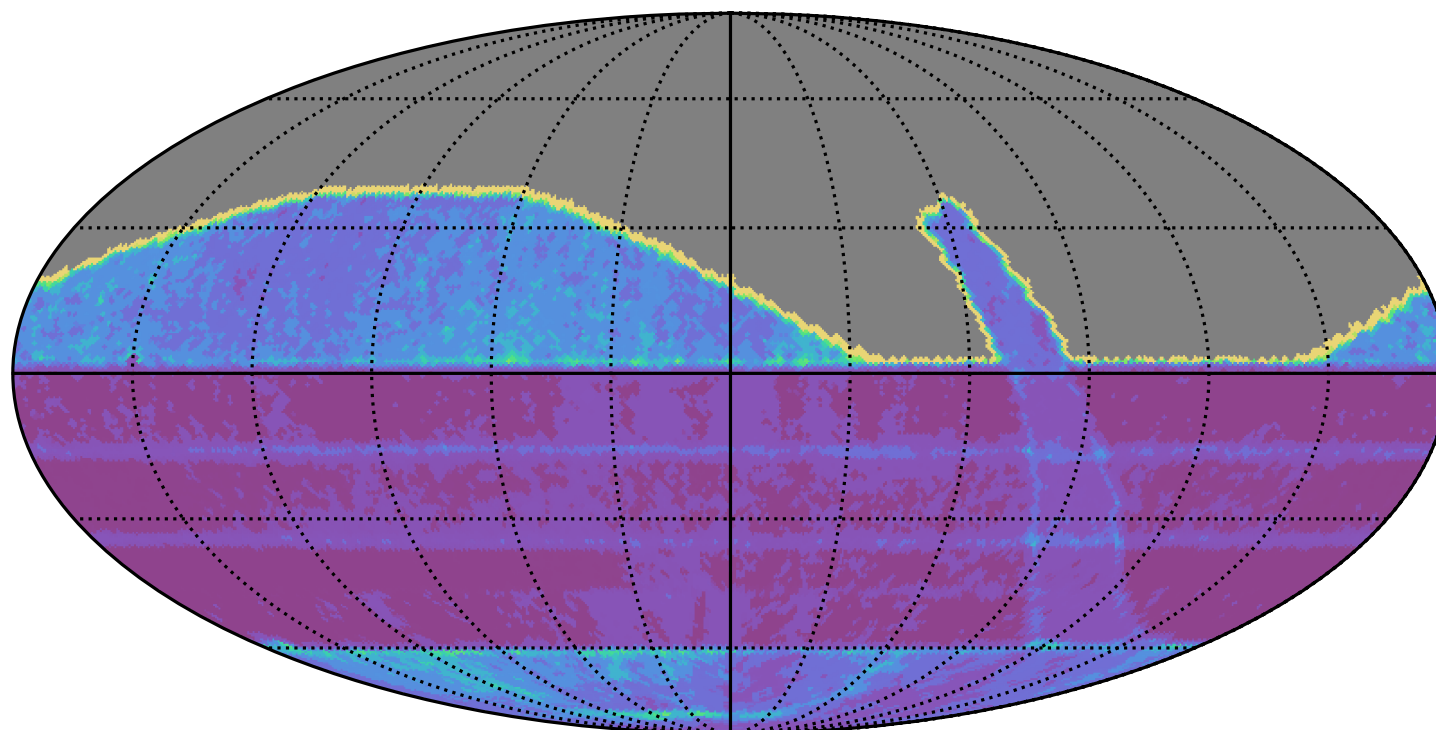
Half-sky
rolling
cadence:
Median gap
2.01 days

opsim all bands: Median Inter-Night Gap



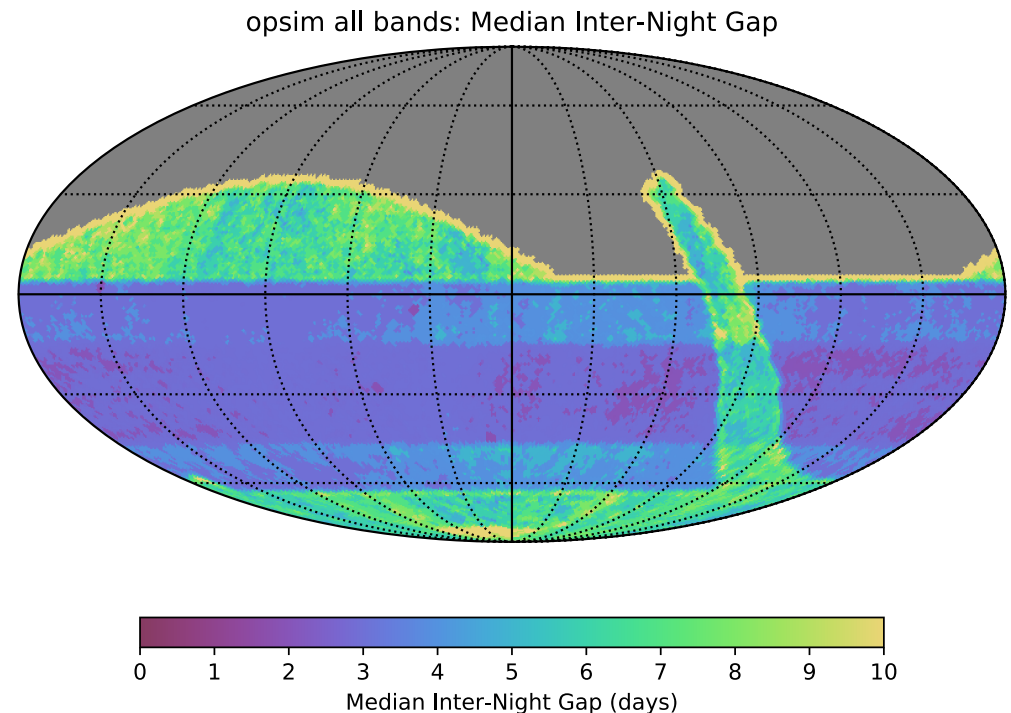
2/3 rolling
cadence:
Median gap
1.97 days

opsim all bands: Median Inter-Night Gap



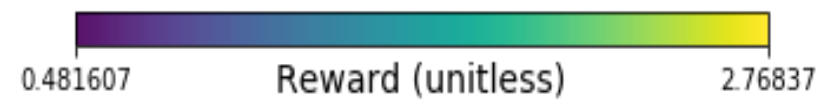
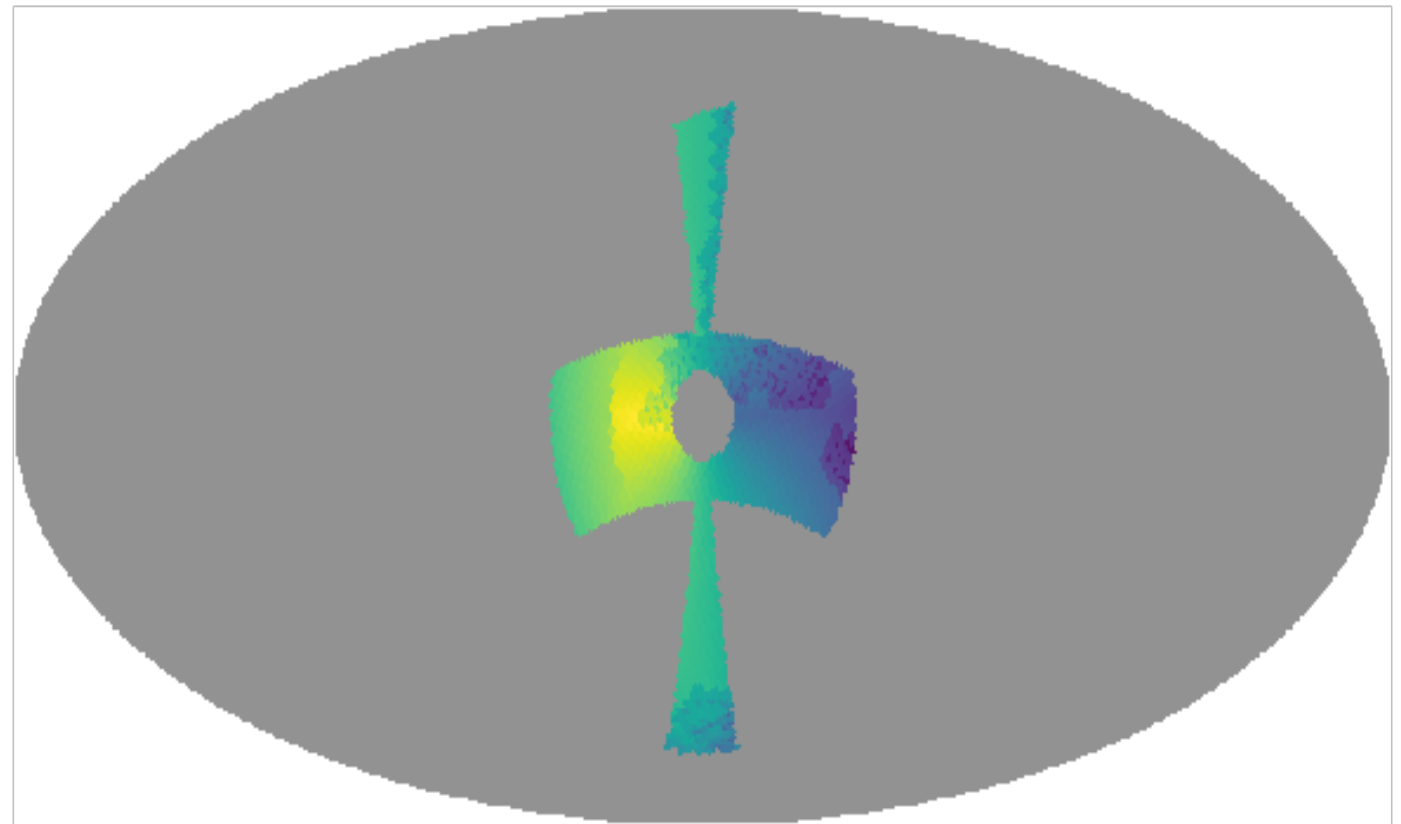
Rolling cadence takeaway:

- The baseline tends to have 4-day gaps (in all filters)
- Overly aggressive rolling cadence can cut that down to 2
- I often head 3 is the magic number for SNe, so we should be able to get close

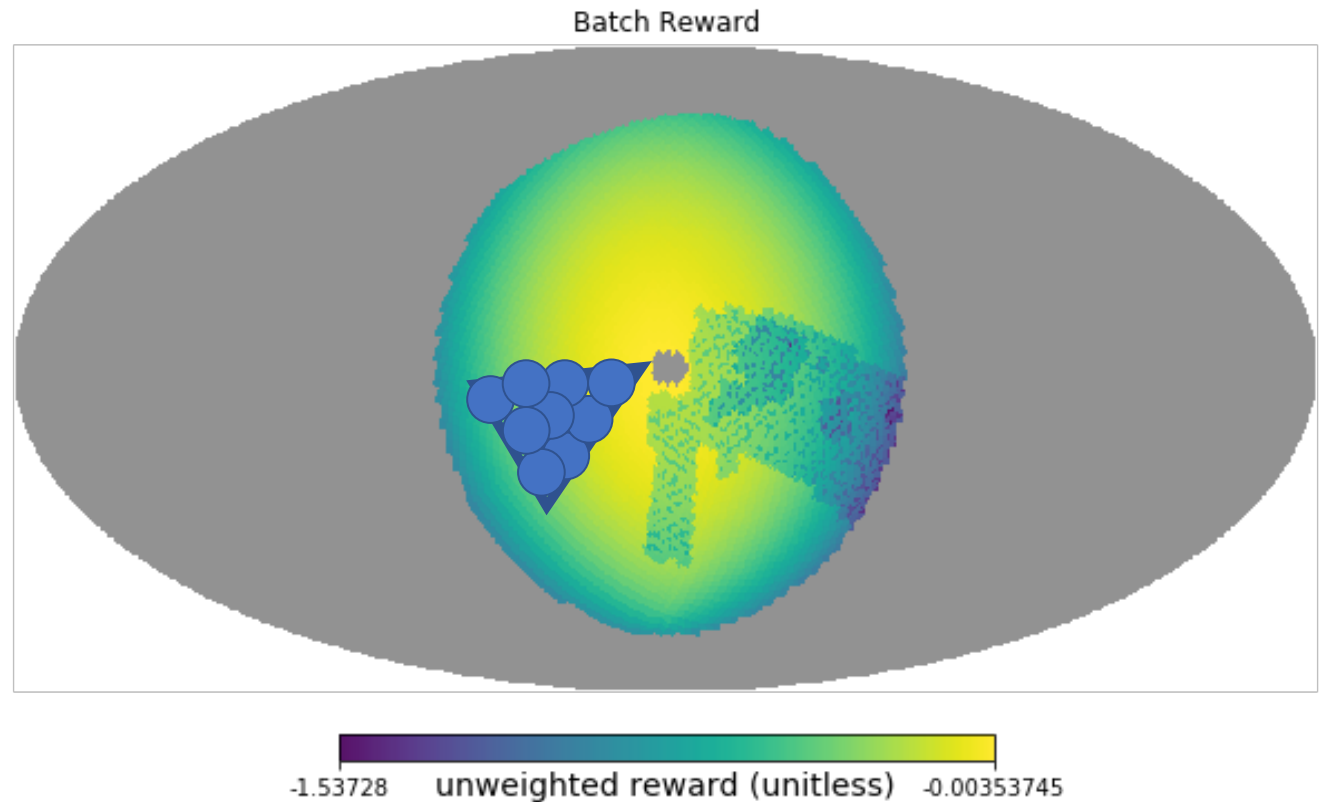


We need to do the slightly more sophisticated experiment where we increase cadence but still observe the full sky each year (rather than masking the target map, just set it to a lower value)

Reward Function

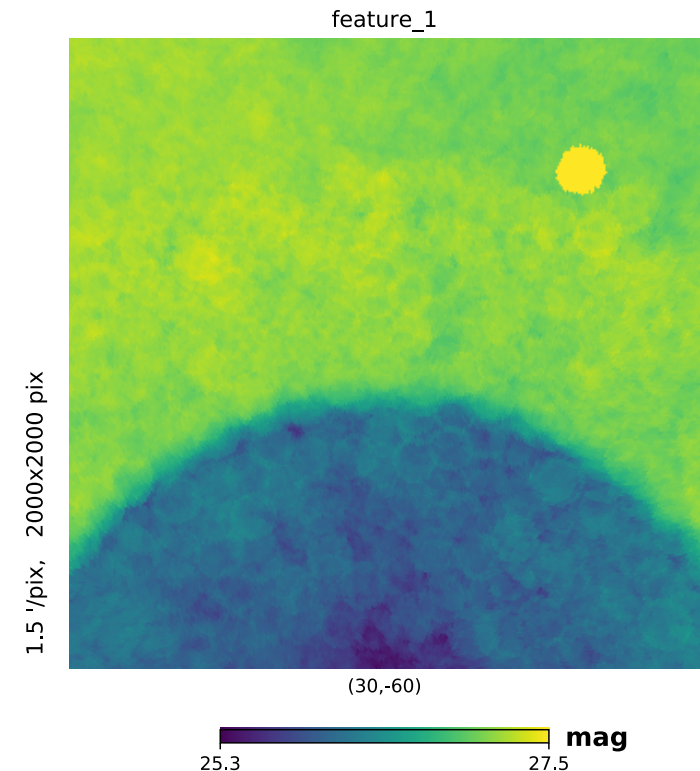


Rather than have a greedy algorithm pick single observations, have a greedy algorithm select a large chunk of sky. Then solve the traveling salesperson to keep slewtime down. Go through the block twice to get pairs.

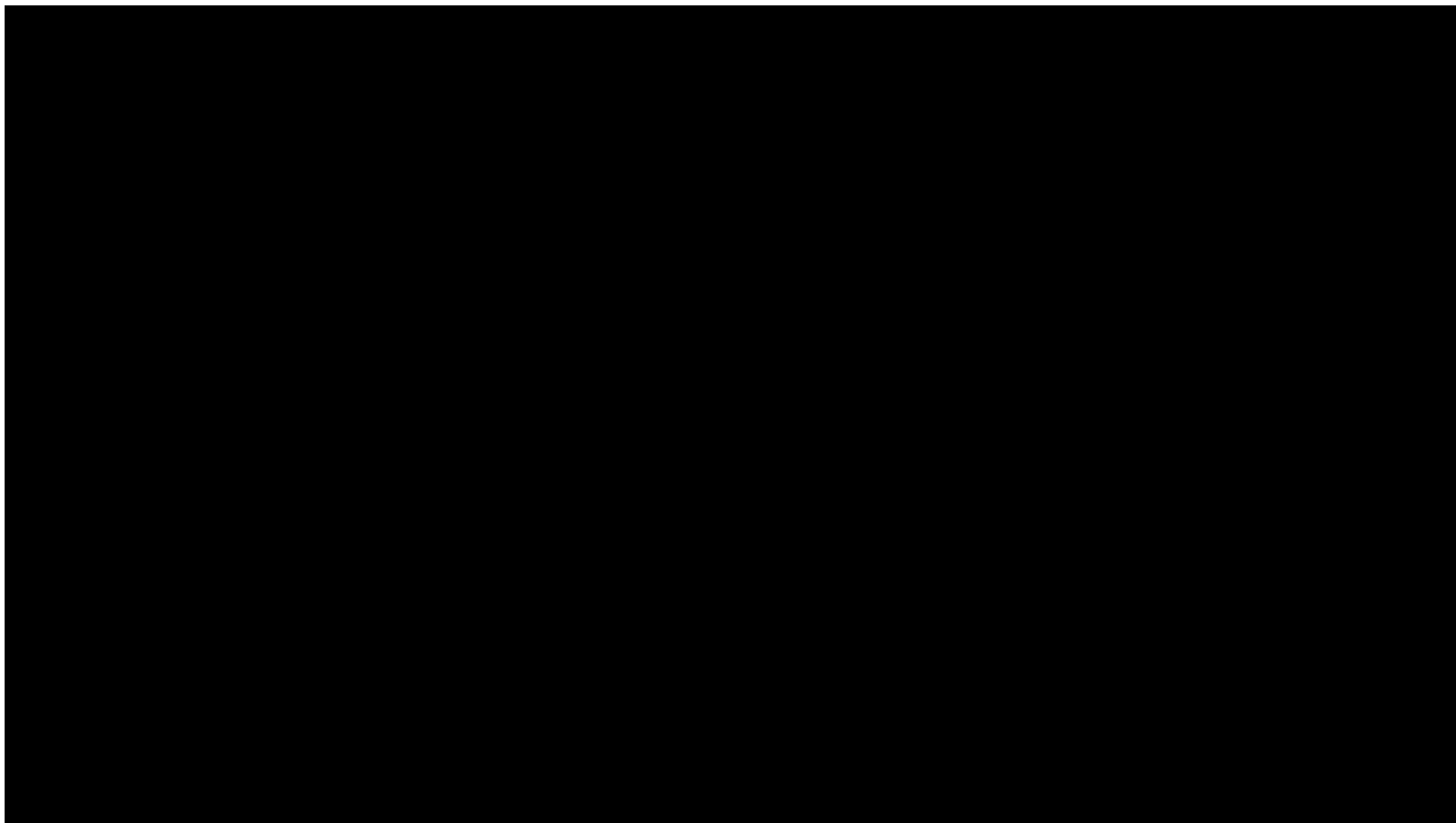


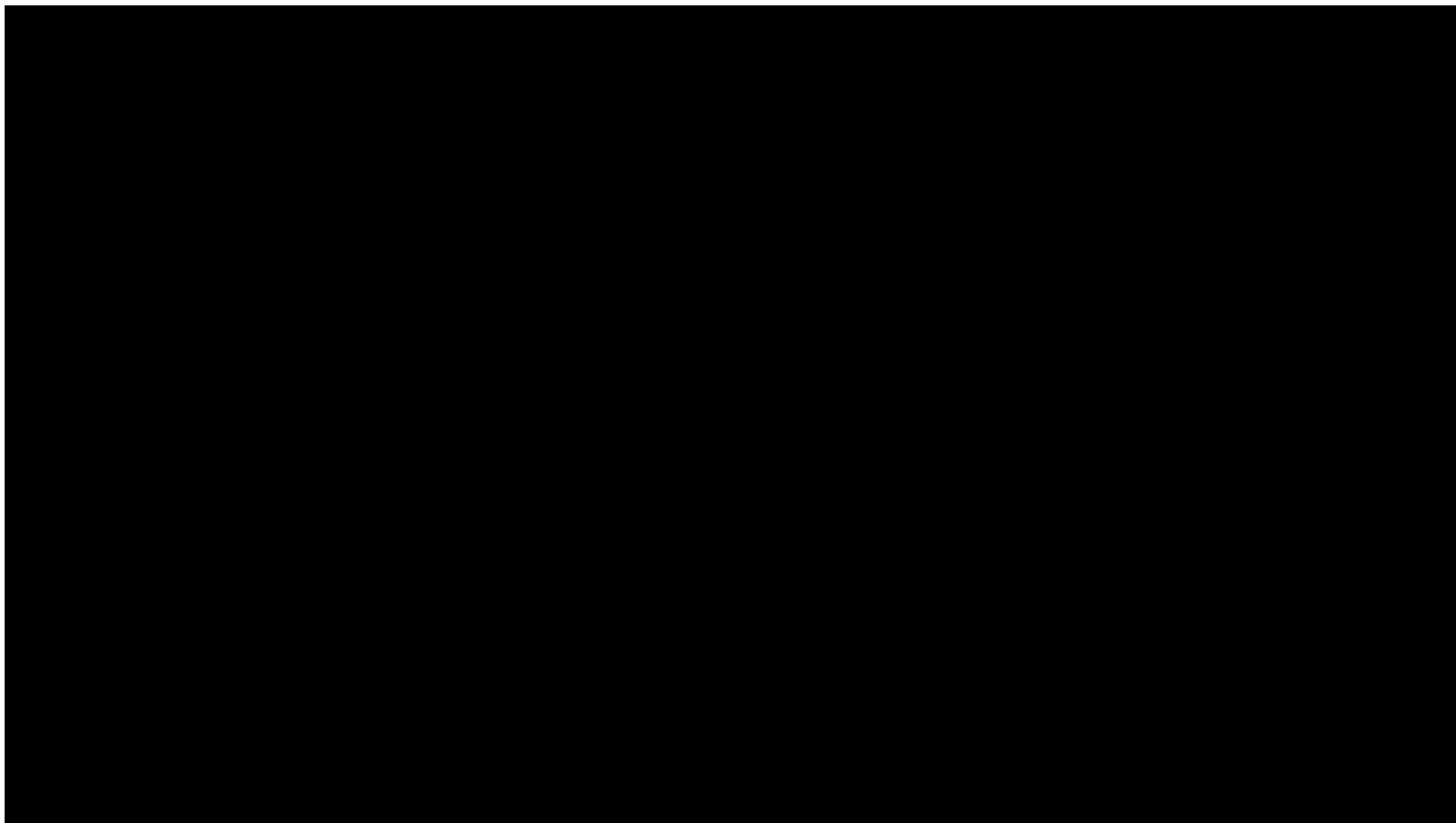
Future Plans:

- Double check that all-sky re-tessellation is good for ubercal (it should be great)
- Schedule large areas rather than single visits
- More sophisticated DD cadence
- We (still) need a dither strategy for the DD fields
- We (still) need a strategy for dithering in camera rotator angle
- More sophisticated rolling cadence
- Adjust WFD area for summer/winter
- Galactic plane as part of WFD
- Try to enforce particular WFD cadence with a new basis function?
- Depth uniformity basis function?
- Special 1st year strategy to get templates
- I'd like to try new sky tessellations



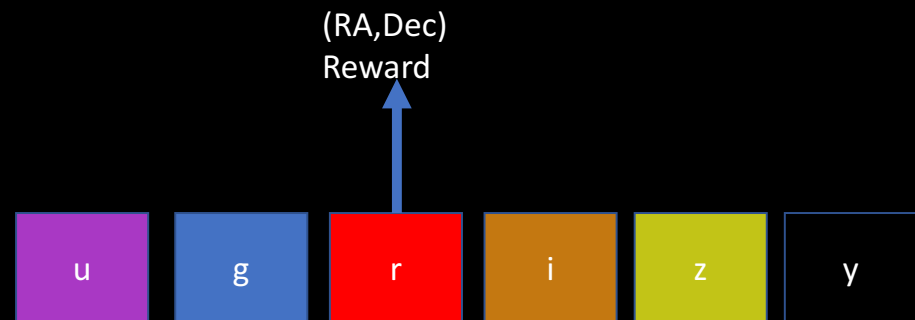
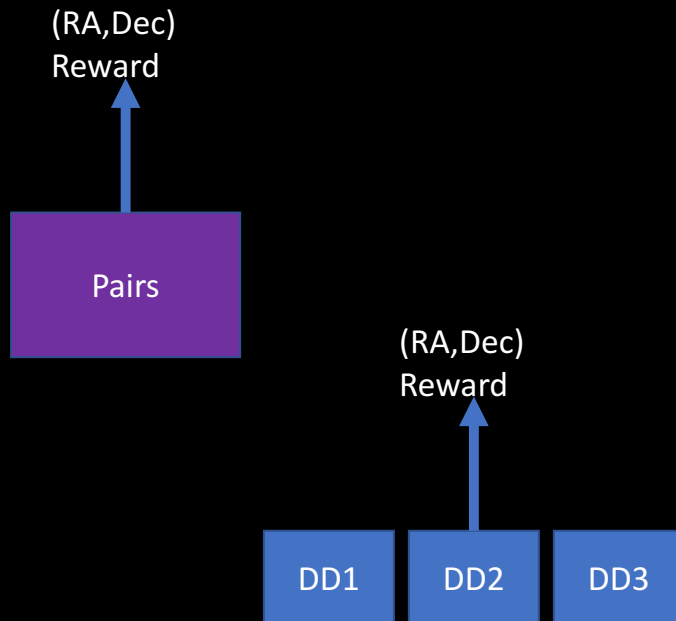
Red = I think DESC could help a lot with that





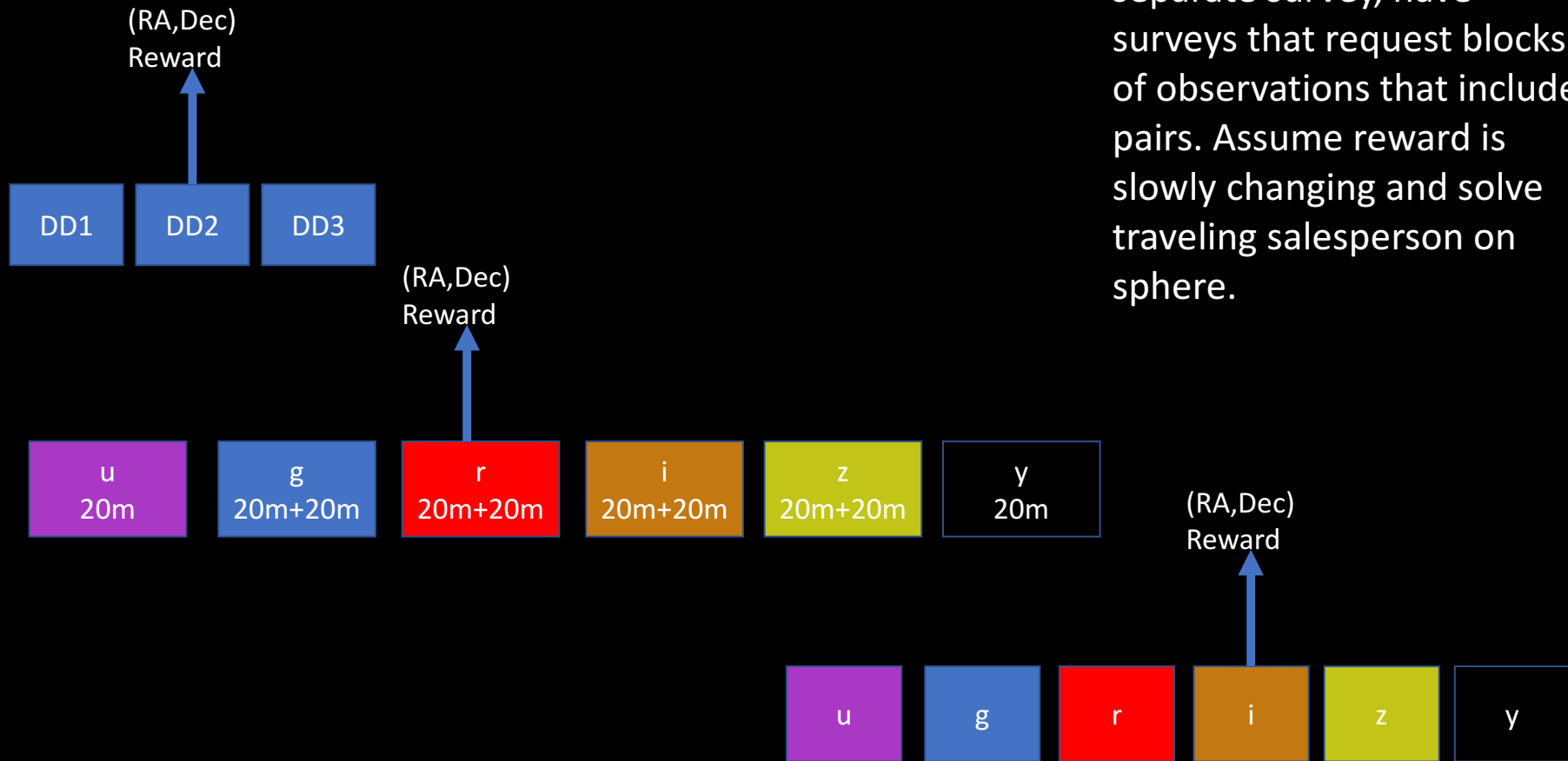
Full Survey

Requesting an observation, each “survey” returns target(s) and a reward value. Highest reward value gets selected.



Full Survey

Rather than pairs as a separate survey, have surveys that request blocks of observations that include pairs. Assume reward is slowly changing and solve traveling salesperson on sphere.



Full Survey

Could include
pairs in different
filters.

