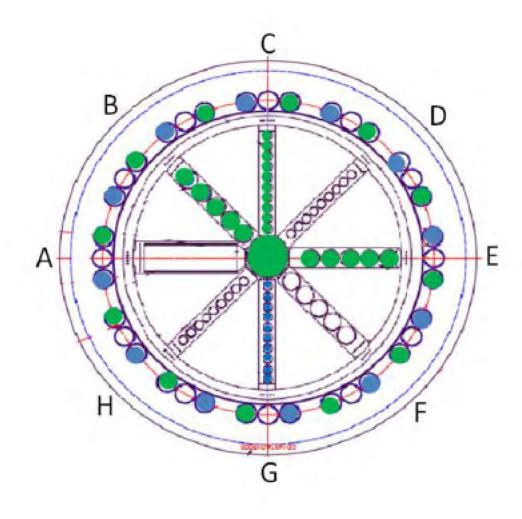
Aux Tel Status plus "Bonus Material"

dkg 25 MAY 2018

- 33 parts coated both sides
- 22 with side 1
- 22 with side 2



AT Observing Strategy

Scheduler will be developed for the AT

- Simplified clone of the LSST schedule
- Based on past and predicted future LSST positions

Spatial aerosol and water variation and their evolution timescale are not well bounded

Effects are band dependent

Full sky characterization versus focusing observations towards LSST positions to be explored

Planning to advance AT science observations to late 2018/early2019

Also need to identify "standard" targets

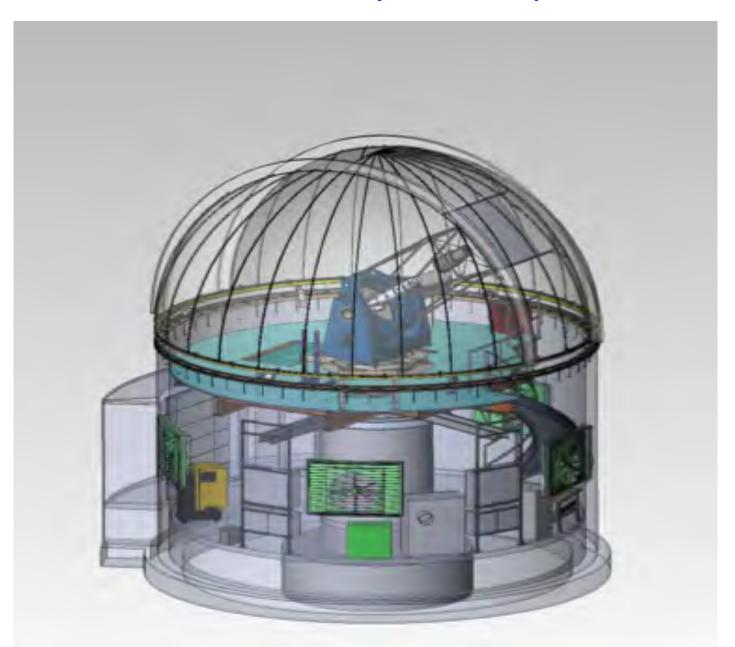
AT will be available for use throughout commissioning

Robotic Auxiliary Telescope System

- Edgar J. Smith Telescope (aka Calypso) donated to LSST
 - 1.2m diameter
 - f/18
 - Two instrument ports
- Currently undergoing integration by ACE, Inc
 - Maintaining compatibility with LSST systems
- Will retain LSST camera systems for ATS (RO/DAQ/CCS)
- Goal is to measure the atmospheric transmission during LSST observations



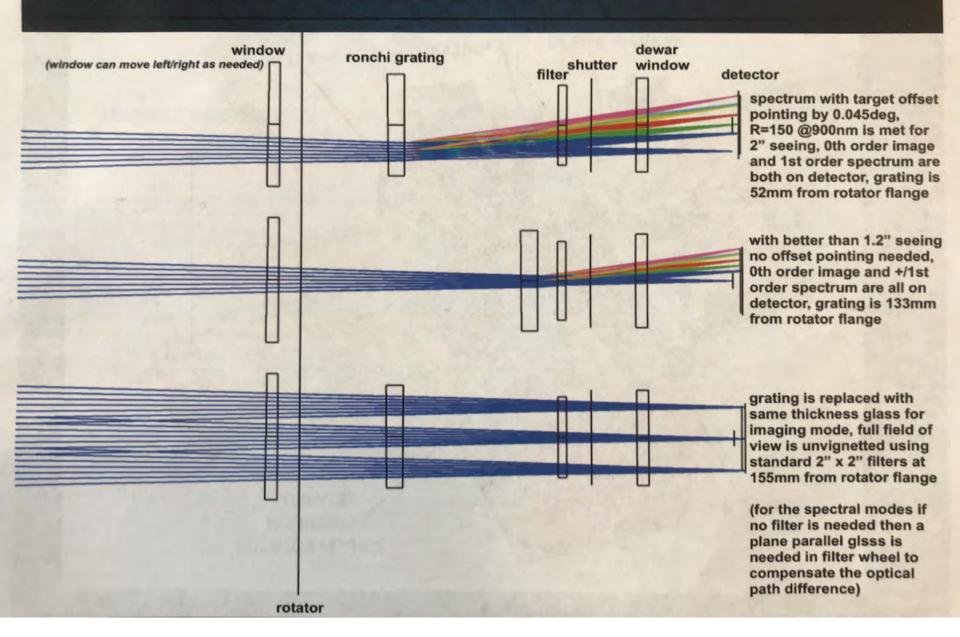
LSST Auxiliary Telescope



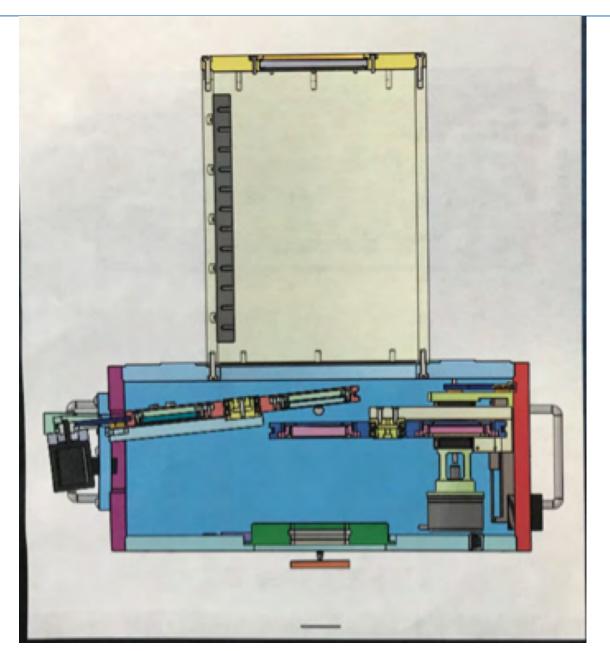
Auxiliary Telescope Optical Model

Meeting the spectral resolution requirement (R~150 @ 900nm) and wavelength range requirement (350-1000nm in a single exposure for all seeing and elevation conditions requires two gratings. Poor seeing or high airmass Median seeing Imaging mode

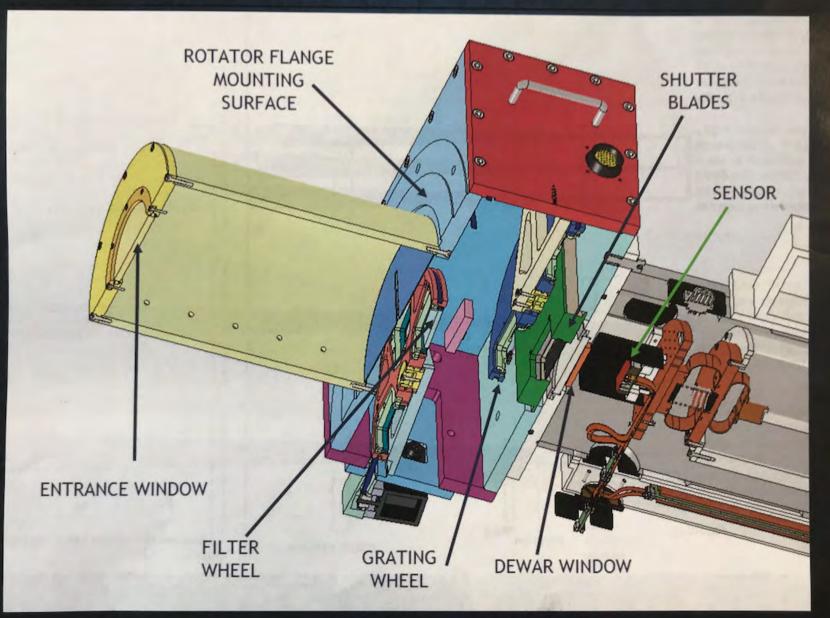
OPTICAL DESIGN - RONCHI GRATINGS



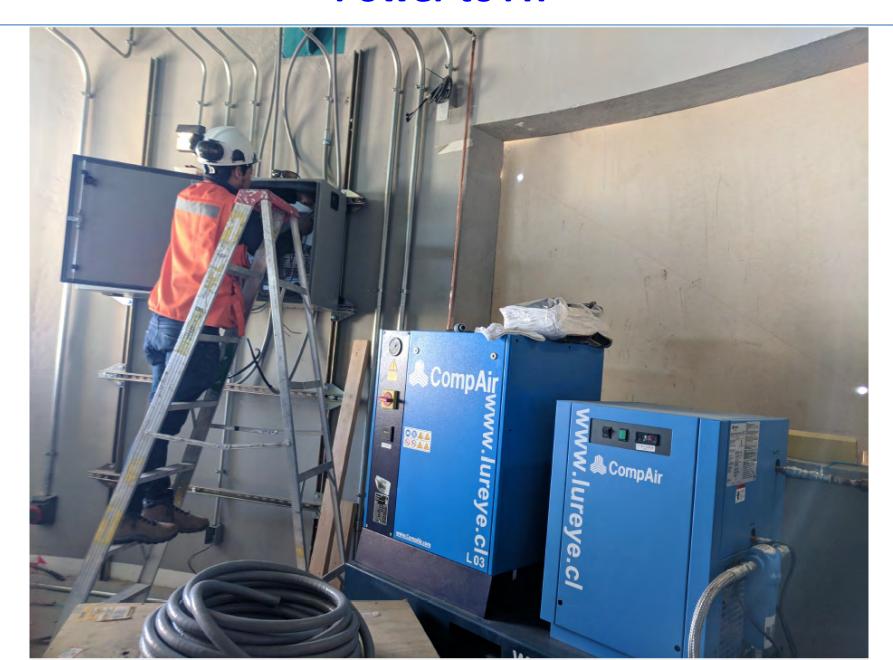
Mechanical Design – Spectrograph Assembly



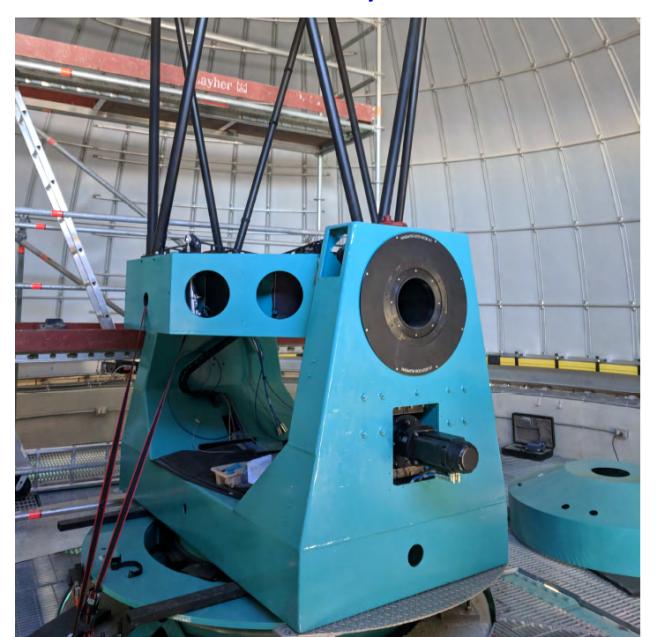
OPTICAL DESIGN - RONCHI GRATINGS



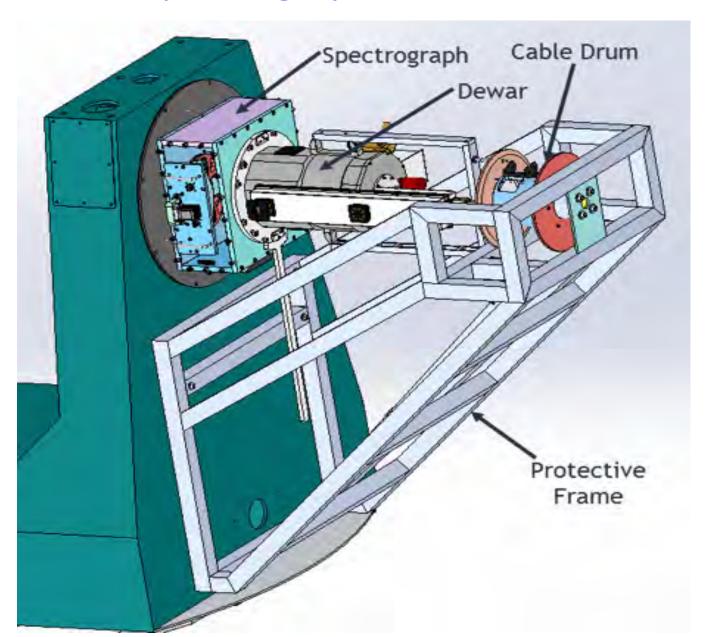
Power to AT



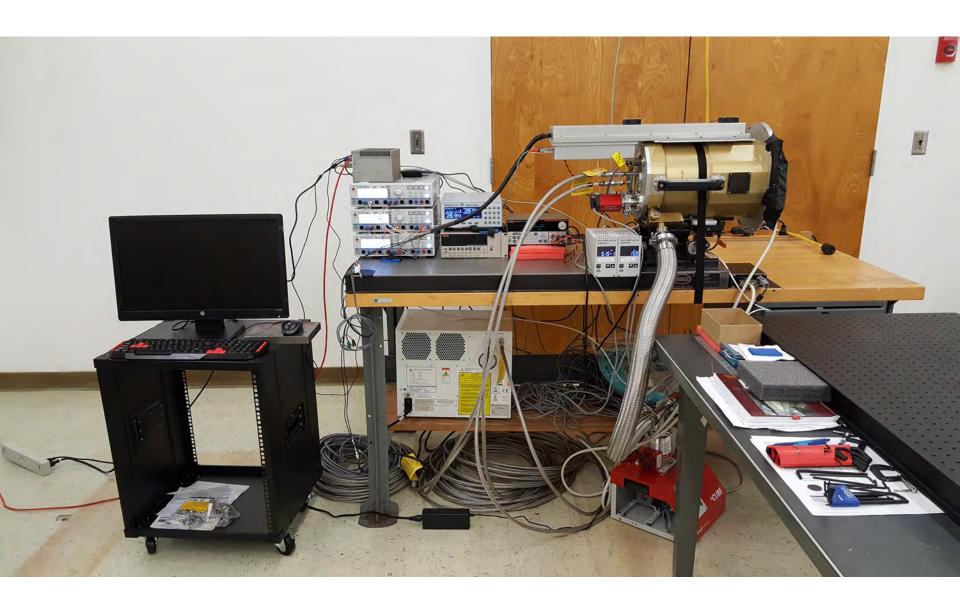
Aux Tel May 2018



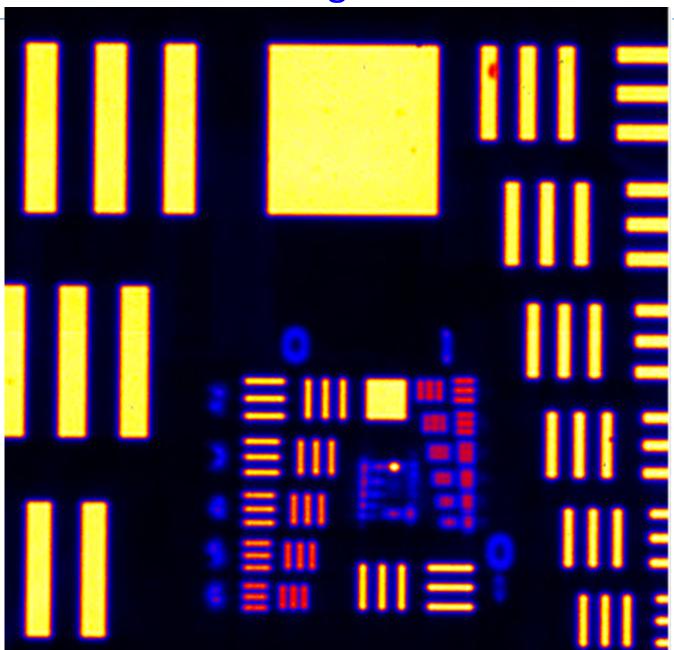
AT Spectrograph and Camera



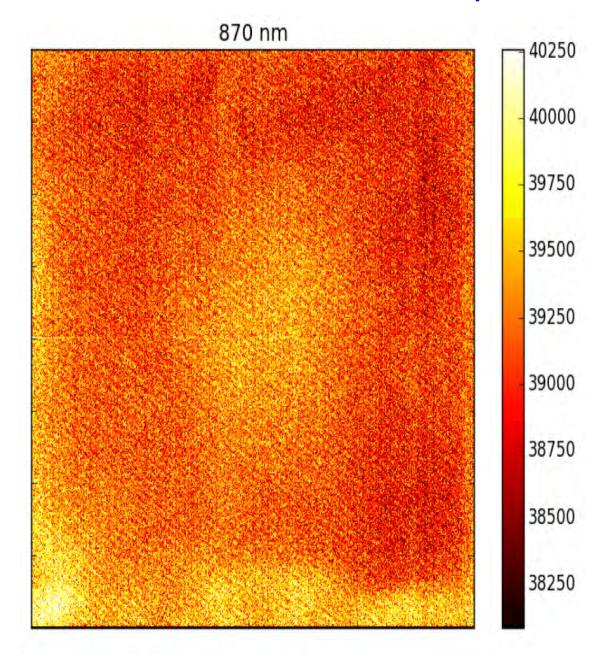
ATS SRS in Tucson

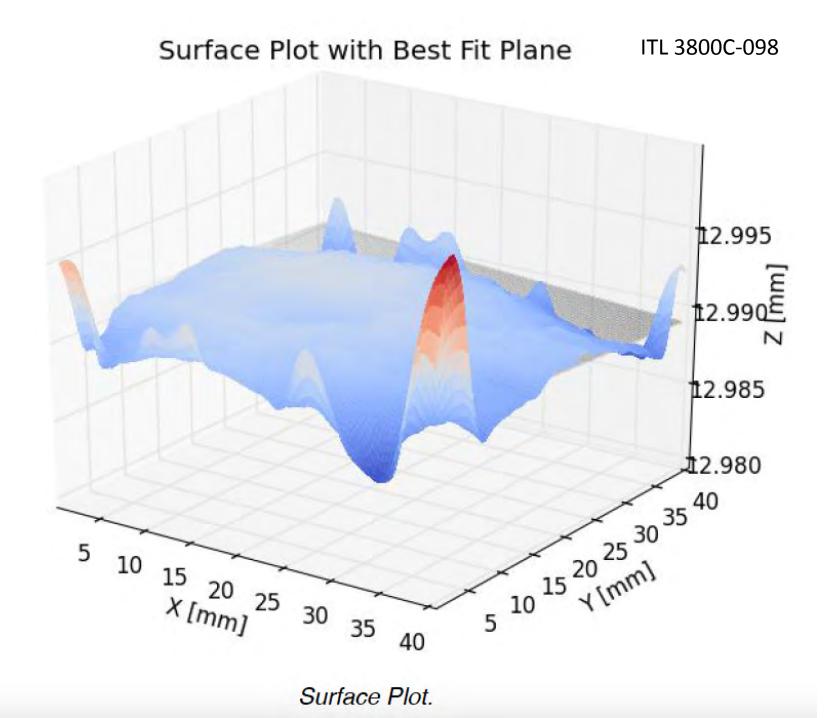


Calibration Target in Tucson



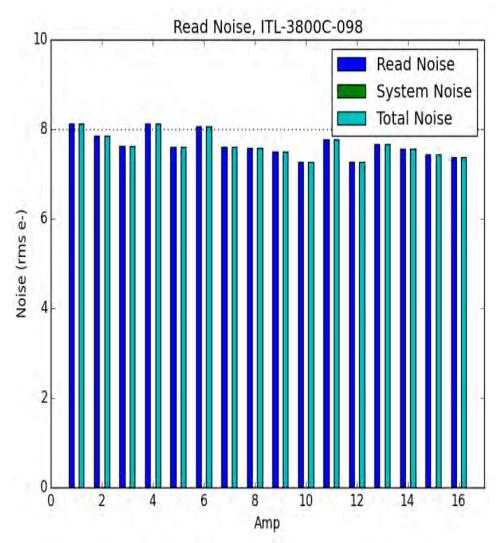
CCD – ITL-098: 870nm Response



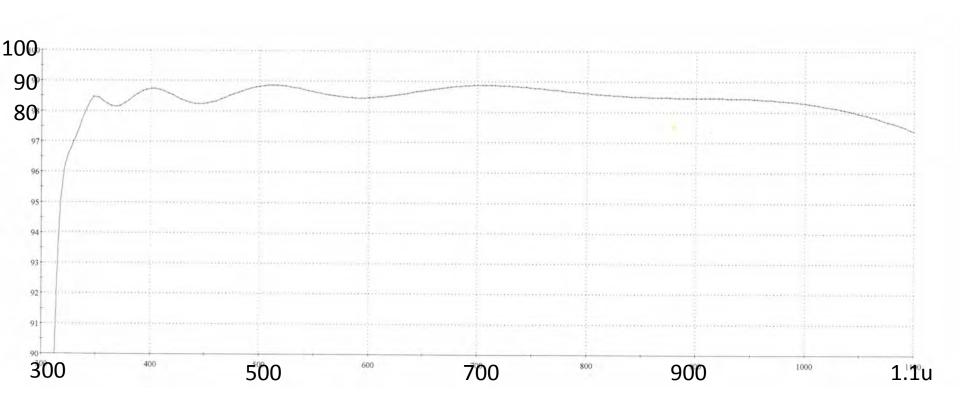


CCD – ITL-098: Read Noise

Amplifiers 1,4, and 6 fail read noise



Camera Input Window BBAR Coating



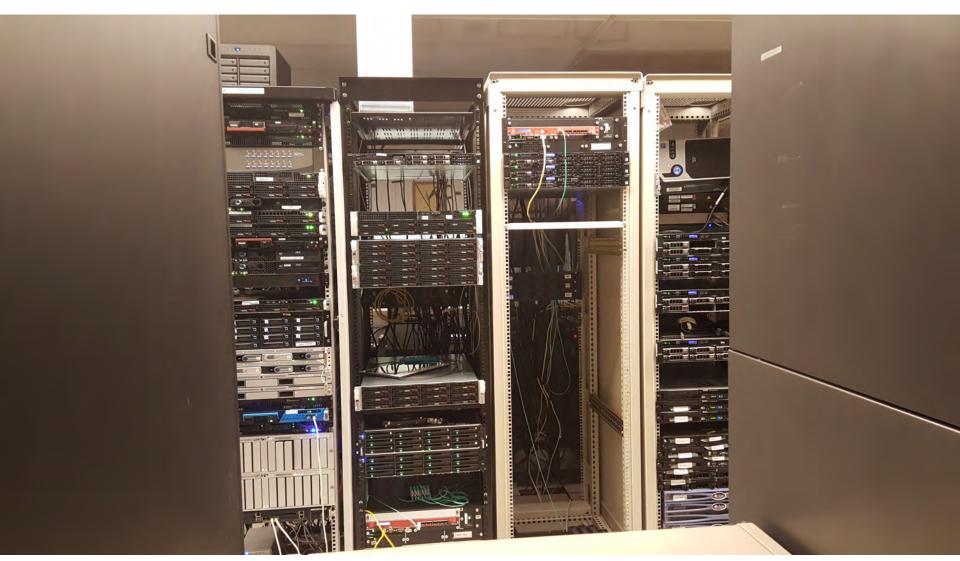
AT Spectrograph



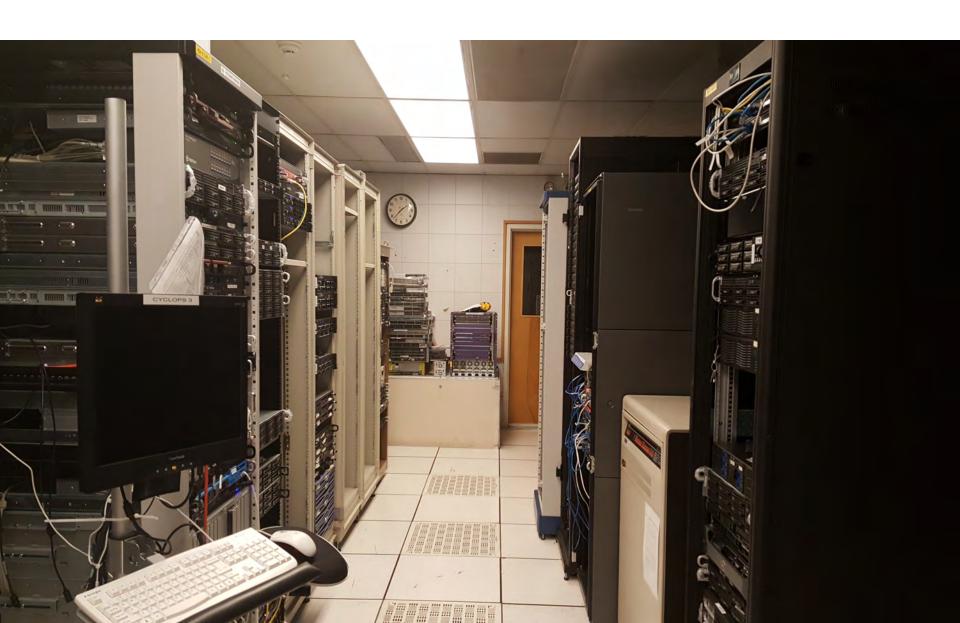
Spectrograph control

Control hardware will be a compact Rio. ACE will provide LabVIEW-based software to drive and communicate with the devices. The interface between ACE and LSST will be via TCP/IP/

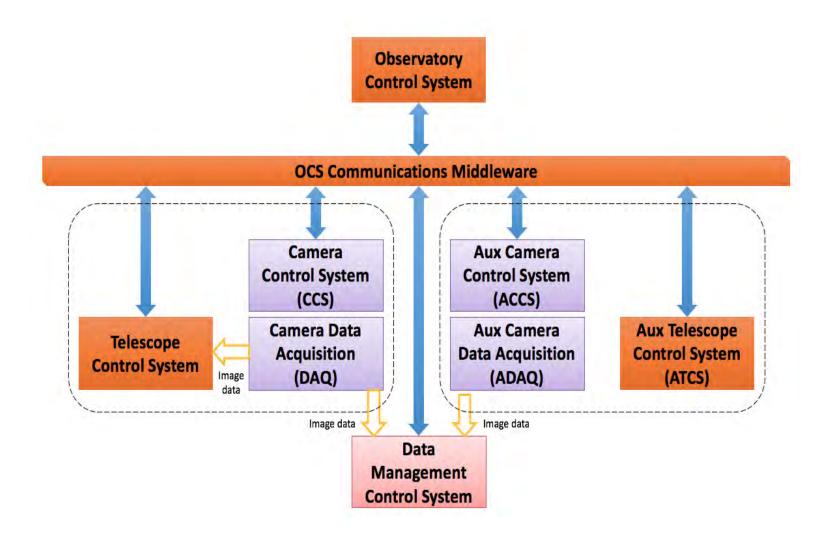
ATS SRS CCS/DAQ rack in LSST Server room



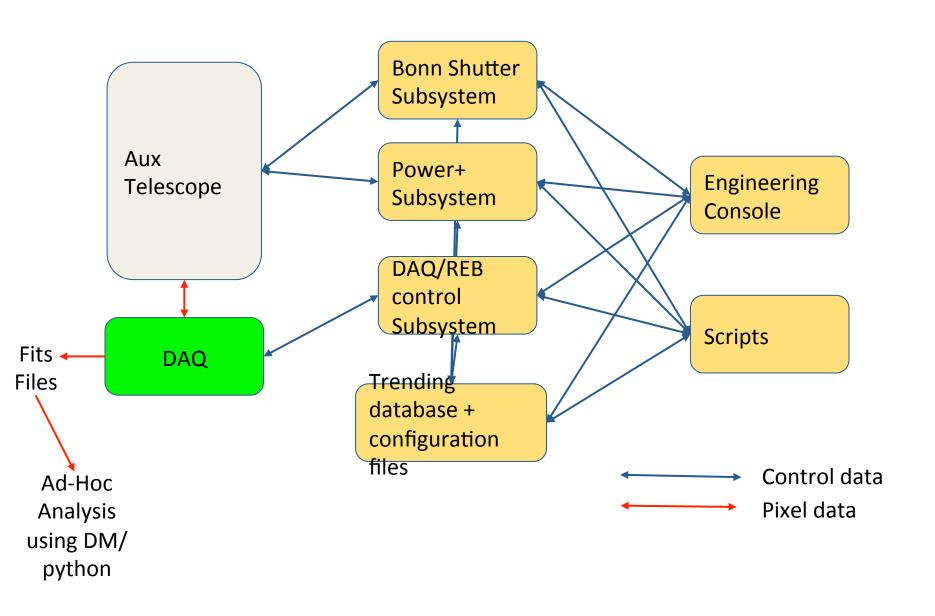
LSST (Tucson) Server Room



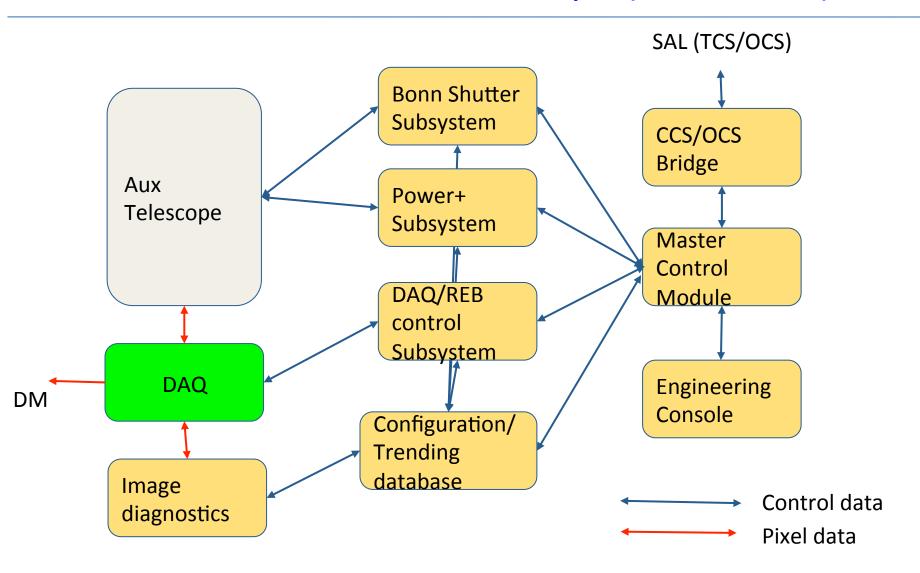
AT Control Software Systems



Software for Aux Telescope (for Harvard)



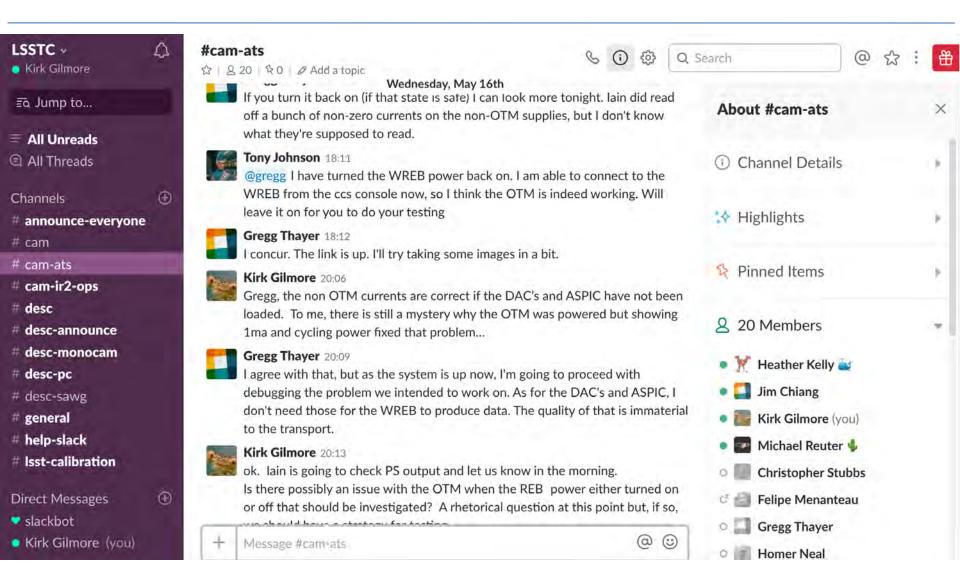
Software for Aux Telescope (for Tucson)



Pathfinder Info

- Pathfinder exercises are being used to test interoperability of
 - Camera Control System (CCS)
 - Camera Data Acquisition System (DAQ)
 - Observatory Control System (OCS)
 - Telescope Control System (TCS)
 - Data Management (DM)
- Series of exercises
 - Starting with simulations
 - Moving to improved fidelity of simulations
 - Real data
 - Auxiliary Telescope (Spectrograph)
 - ComCam

SLACK channel for ATS

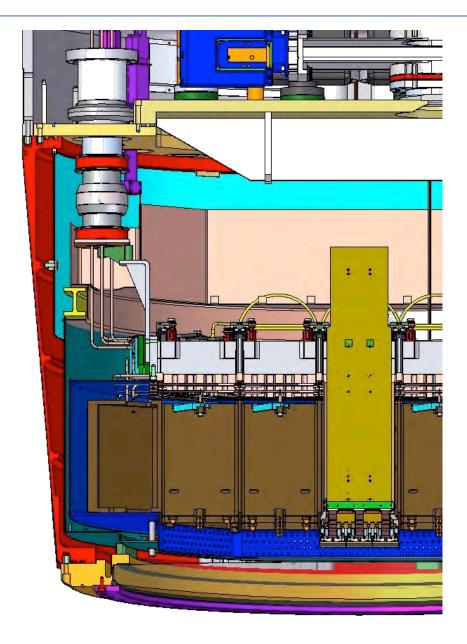




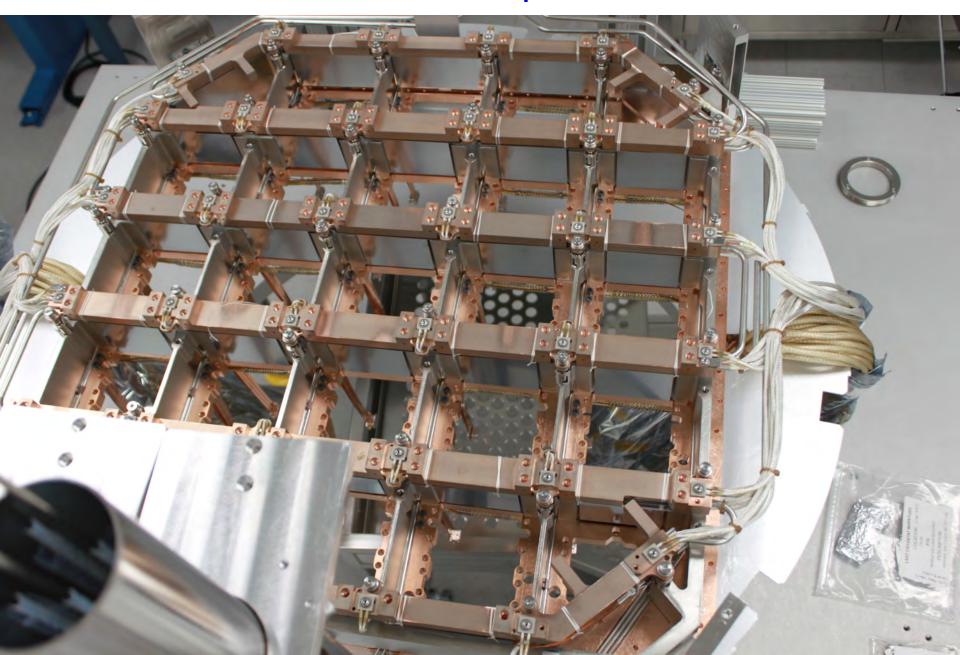
LSST – May 17, 2018



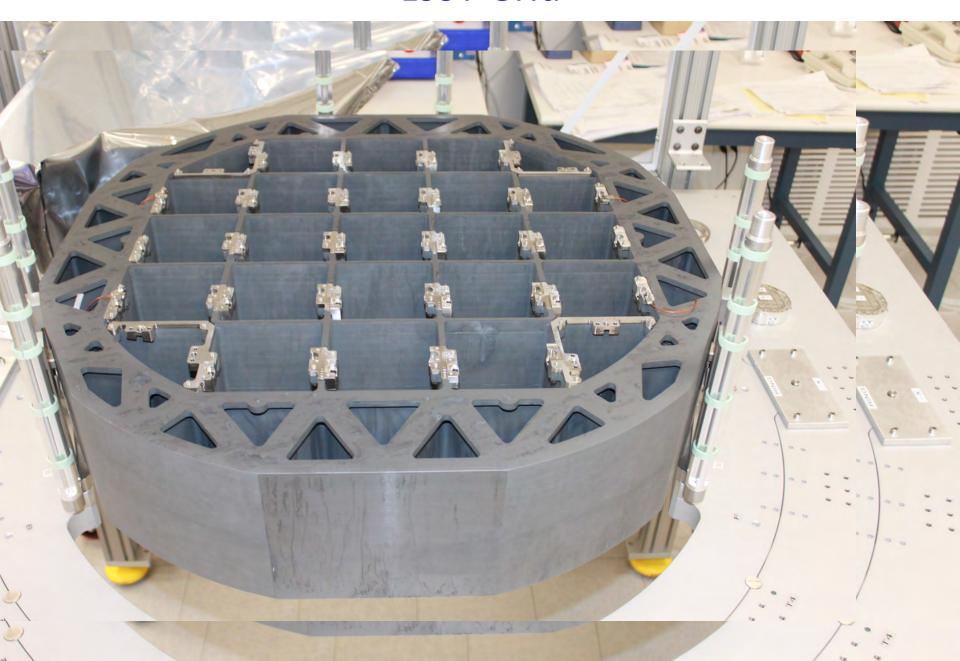
Camera Assembly



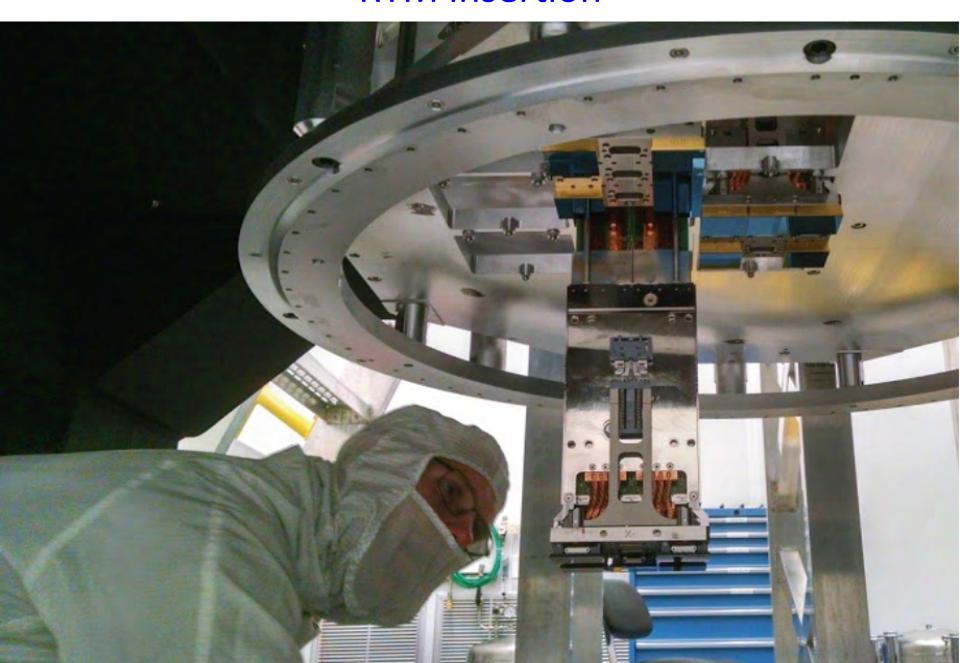
LSST cold plate



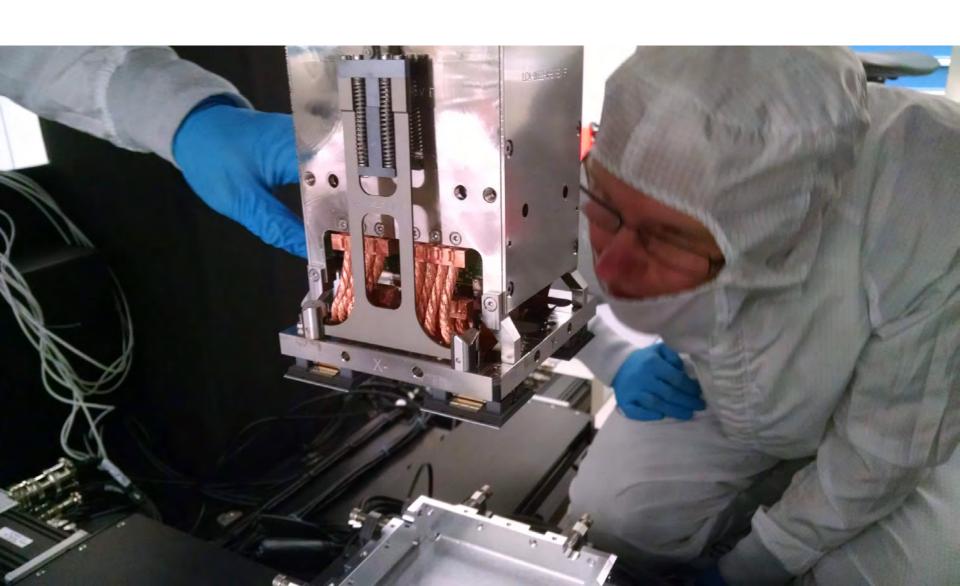
LSST Grid



RTM insertion



RTM Insertion



LSST Camera – back flange



The I&T BOT – LSST cleanroom



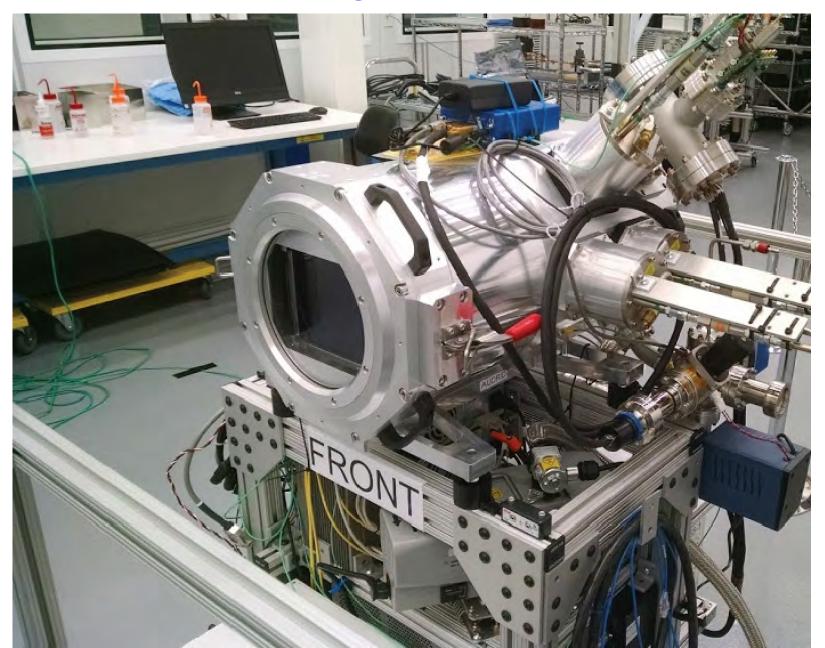
Shutter Blades



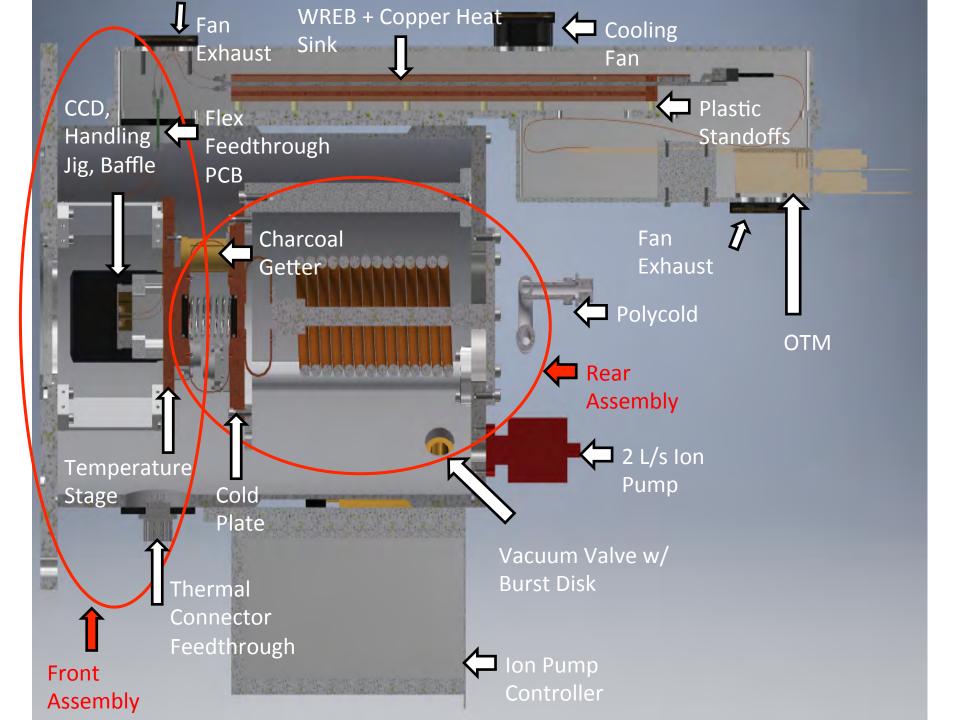
Shutter Rails



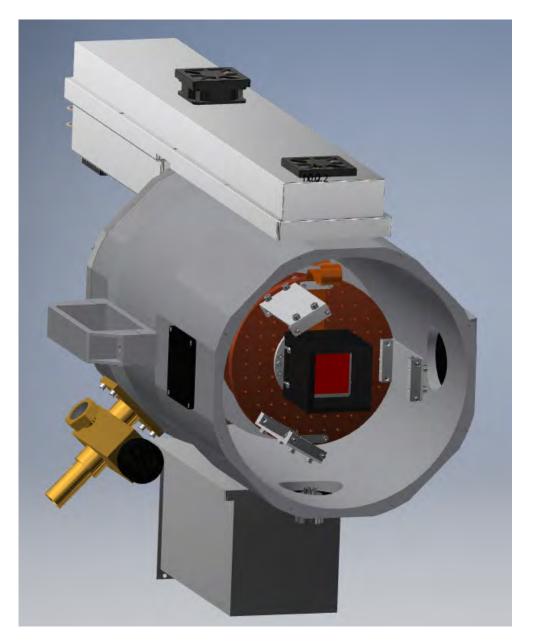
Commissioning Camera (ComCam)



Extra Slides



Full Dewar – No Input Window Flange

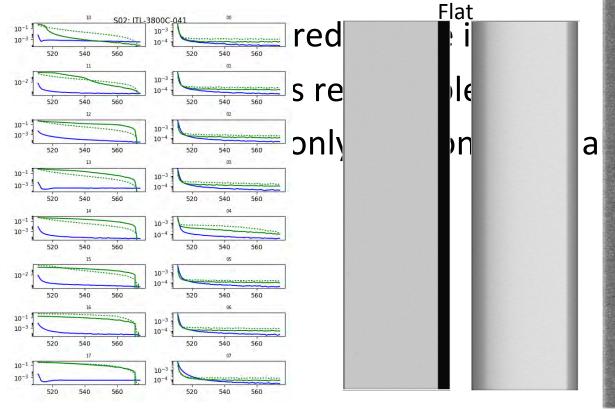


Changes in ETU#2

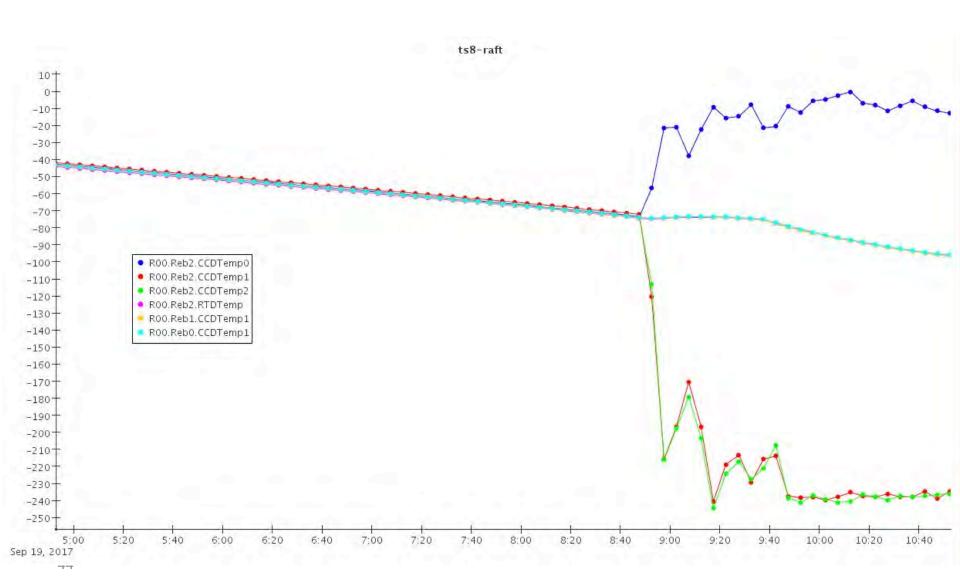
Likely a Serial Clock phase is missing: Flex Cable?

• ½ of S02 ITL-041 is broken, was working fats

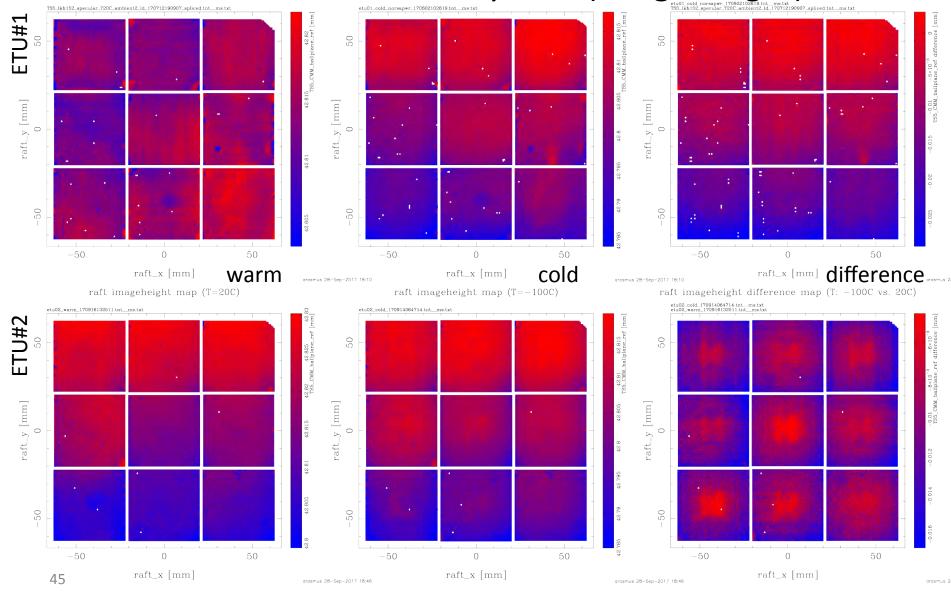
BNL



Changes in ETU#2



raft image EgTU#1 & ETU#2, side-by-side (image-height) map (T: -95C vs. 20C)



MOUNTED ON NASMYTH PORT

