

09:30 → 10:00	Accueil	30m	
10:00 → 10:30	Introduction: la stratégie d'upgrade d'Advanced Virgo	30m	
	15 + 15 minutes		
	Orateur: Benoit Mours (LAPP Annecy)		
10:30 → 11:15	AdV+ science	45m	
	La science que nous pouvons faire avec des détecteurs de type AdV+ , A+, Kagra et l'astronomie multimessager; 25+20		
	Orateur: Patrice Hello (LAL-Orsay)		
11:15 → 11:45	Les challenges de l'optique	30m	
	Augmenter la taille des miroirs, la lumière diffusée,... 20+10		
	Orateur: Jérôme Degallaix (LMA)		
11:45 → 12:15	L'augmentation de la puissance du laser et les enjeux associés	30m	
	20+10		
	Orateur: Nelson Christensen (ARTEMIS)		
12:15 → 13:30	Déjeuner sur place (pour les personnes inscrites)	1h 15m	
13:30 → 14:00	La réduction du bruit quantique	30m	
	20+10		
	Orateur: Pierre-François Cohadon (LKB)		
14:00 → 14:30	Noise budget et La réduction des bruits techniques	30m	
	20+10		
	Orateur: Michal Was (LAPP)		
14:30 → 15:15	- L'organisation de l'analyse des données, les contacts avec les partenaires multimessager, le calcul	45m	
	25+20		
	Orateur: Marie Anne Bizouard (LAL)		
15:15 → 15:45	L'organisation de Virgo, le programme de base	30m	
	15+15		
	Orateur: Frédérique MARION (LAPP)		
15:45 → 16:00	Opportunités à EGO	15m	
	10+5		
	Orateur: Stavros KATSANEVAS (CNRS/IN2P3)		
16:00 → 16:30	Discussion	30m	

► Organisée avec le soutien du GdR Ondes Gravitationnelles.

▪ <http://gdr gw.in2p3.fr/>

► Agenda: <https://indico.in2p3.fr/event/17271/>



Groupement de recherche  
Ondes gravitationnelles

# **La stratégie d'upgrade d'Advanced Virgo AdV+**

B. Mours (LAPP-Annecy)

May 31, 2018

# GW science is driven by detector progress

## ► Best BNS range for initial detectors

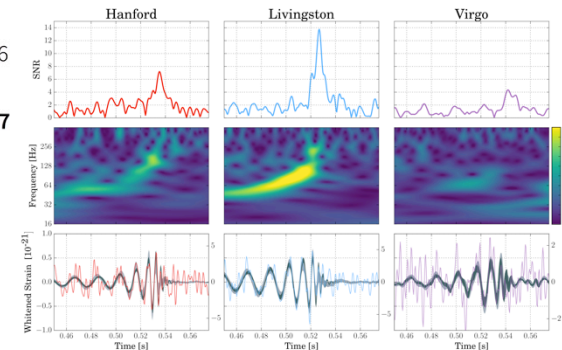
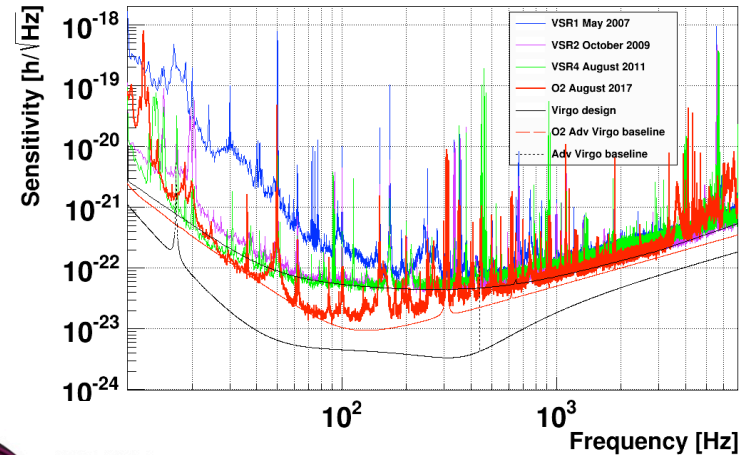
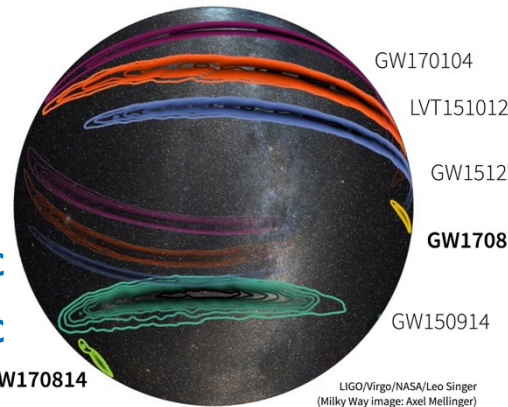
- LIGO:  $\sim 20$  Mpc
- Virgo:  $\sim 12$  Mpc

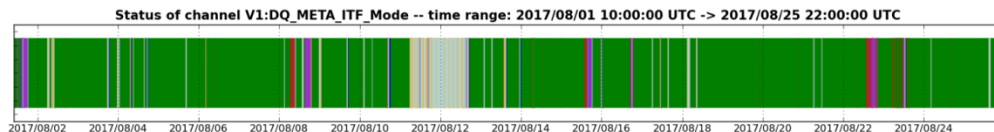
## ► Advanced detectors

- LIGO O1:  $\sim 60$ -80 Mpc
- LIGO O2:  $\sim 70$ -105 Mpc
- AdV O2: 25-28 Mpc

## ► Sources

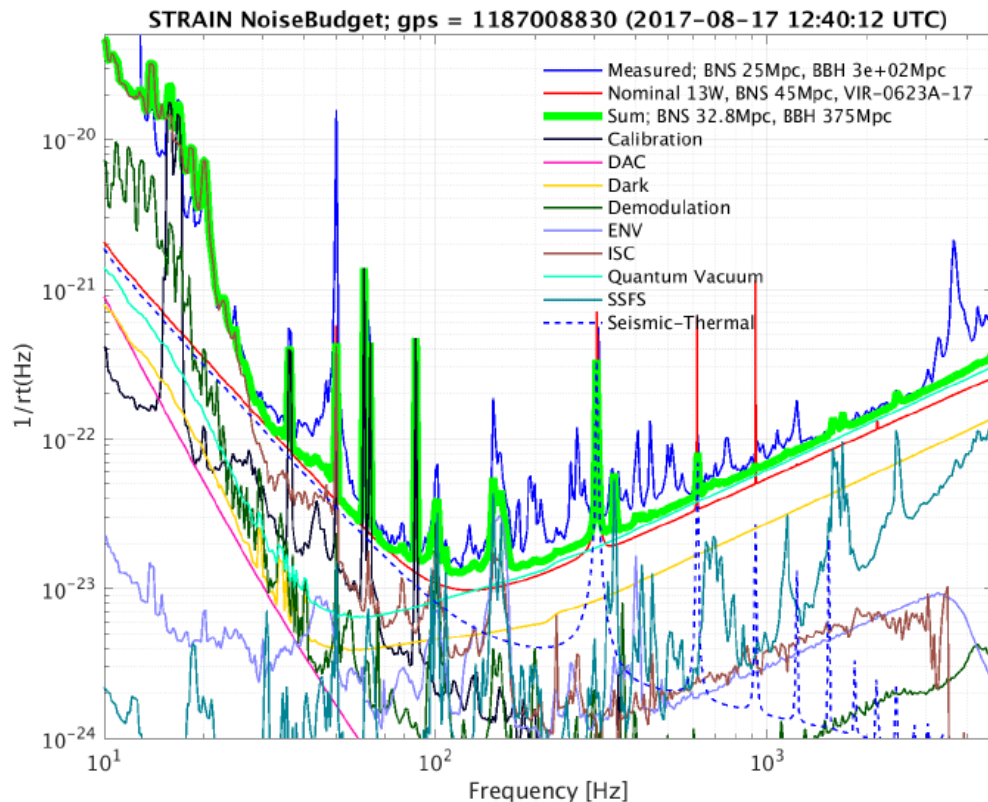
- GW150914: BBH @ 410 Mpc
- GW170814: BBH @ 540 Mpc
- GW170817: BNS @ 40 Mpc



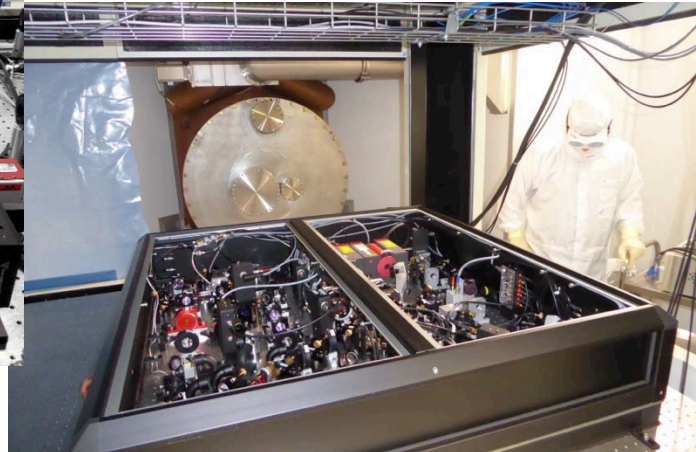
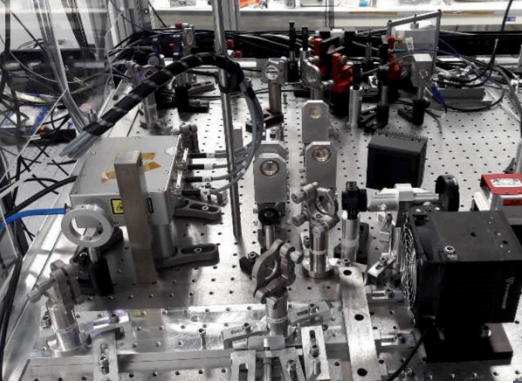
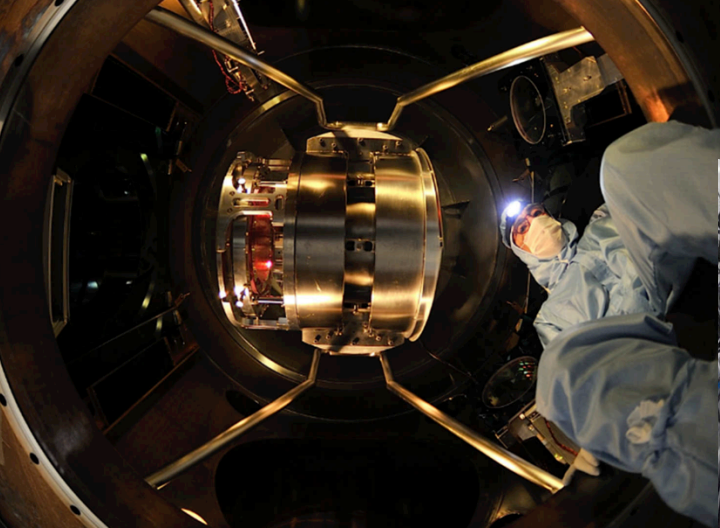


# O2 Detector Summary

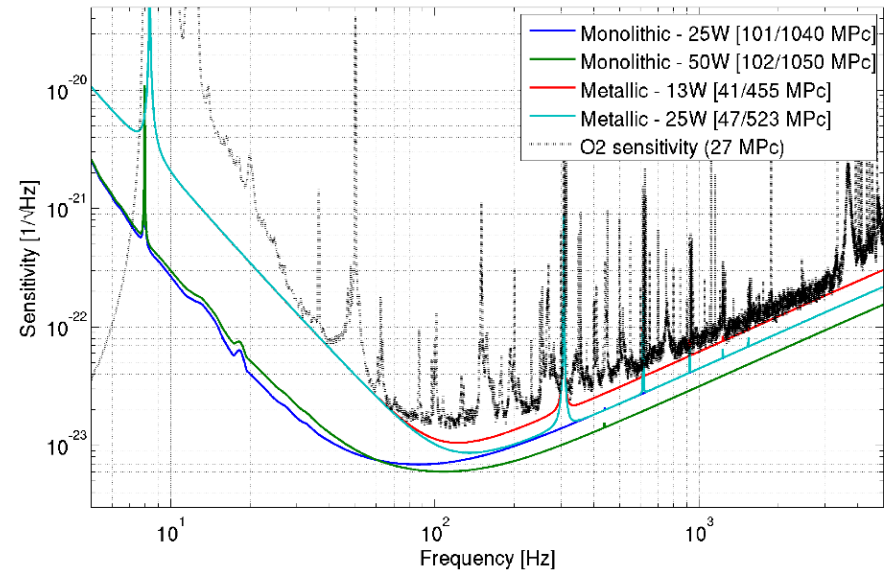
- ▶ 85 % duty cycle
  - Longest lock segment: 69 hours
- ▶ Mean BNS range: 25-28 Mpc
  - Limit was 45 Mpc
- ▶ Noise budget
  - Many bumps and lines and some extra broadband noise



## From O2 to O3

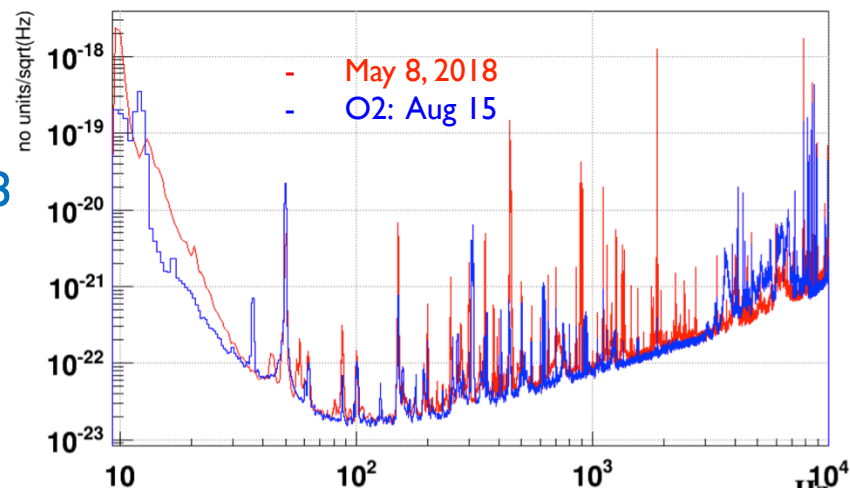


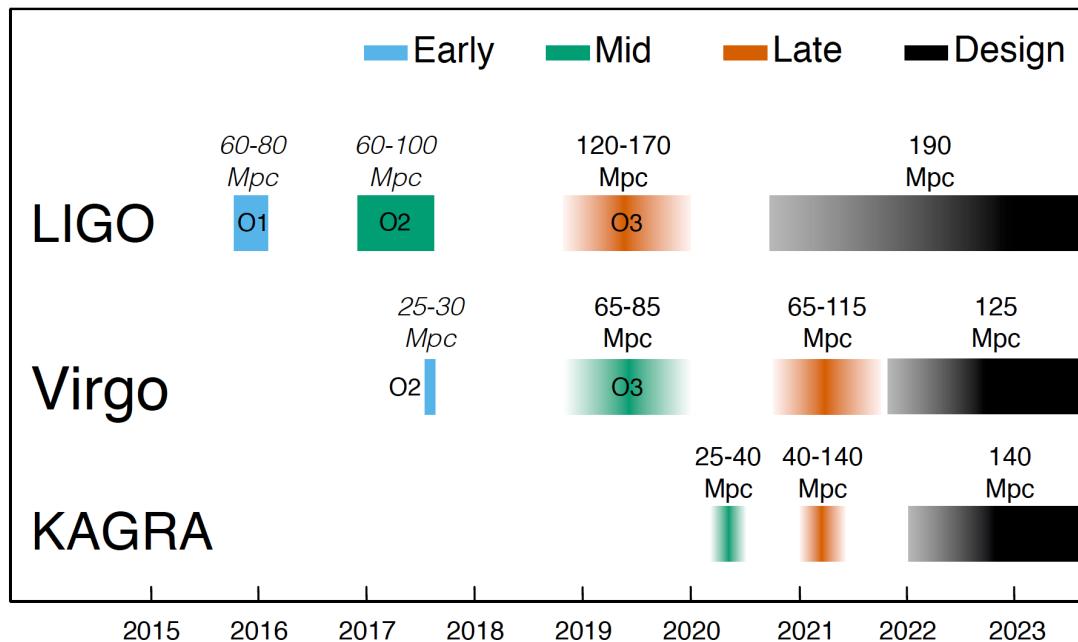
- ▶ Main changes
  - Monolithic suspensions
  - Laser power: 70 W  $\rightarrow$  100 W
  - Frequency independent squeezer
- ▶ O3 Goal: 60 Mpc
- ▶ Installation completed on April 19
  - Back to commissioning



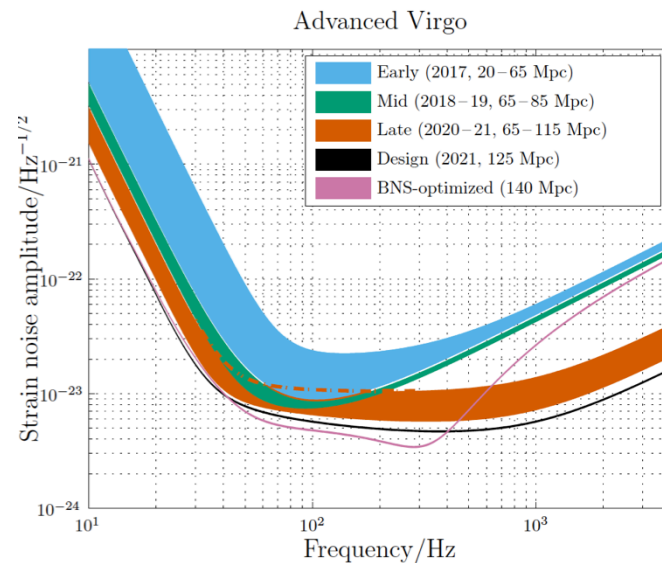
# Toward O3: back to commissioning

- ▶ Back in “low noise 3” on May 2<sup>nd</sup>,
  - Reach up to 25 Mpc (BNS range)
- ▶ Long commissioning period
  - No major shutdown scheduled before O3
  - O3 start: Feb. 2019, aligned with LIGO
  - Commissioning activities
    - ▶ Usual noises/glitches reduction
    - ▶ Robustness
    - ▶ Commissioning of the squeezer
    - ▶ Power increase; max for O3: 50W
- ▶ Engineering Runs
  - Once per month over a weekend
  - Long engineering run prior to O3



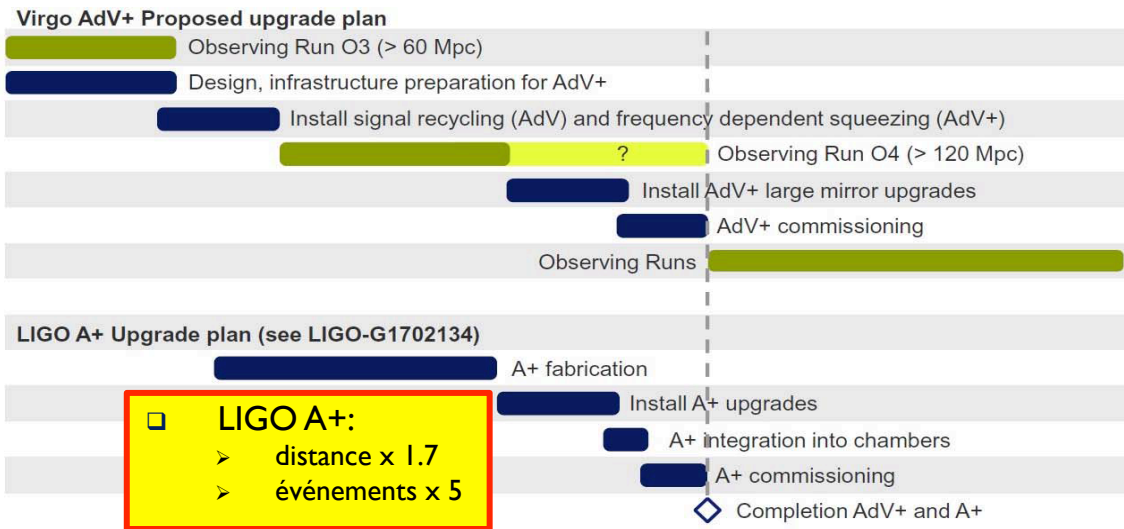


## Beyond O3 ?

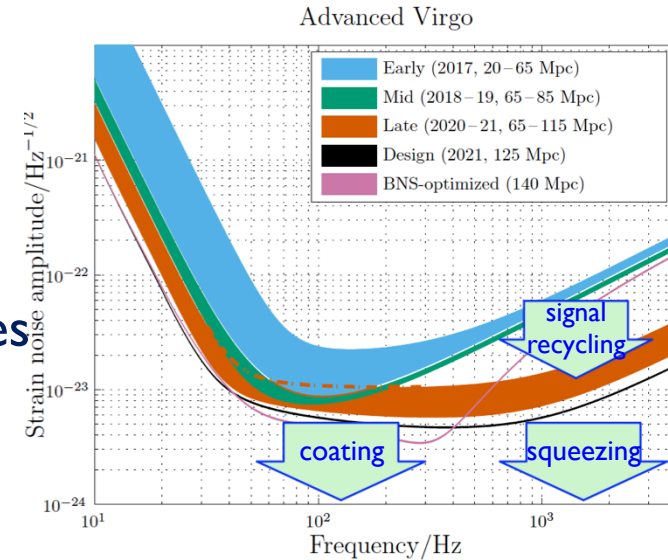


- AdV design range: 125 Mpc
- ▶ Could we do better than AdV?
- ▶ Could we stay competitive?
- ▶ Could we prepare the way to ET?





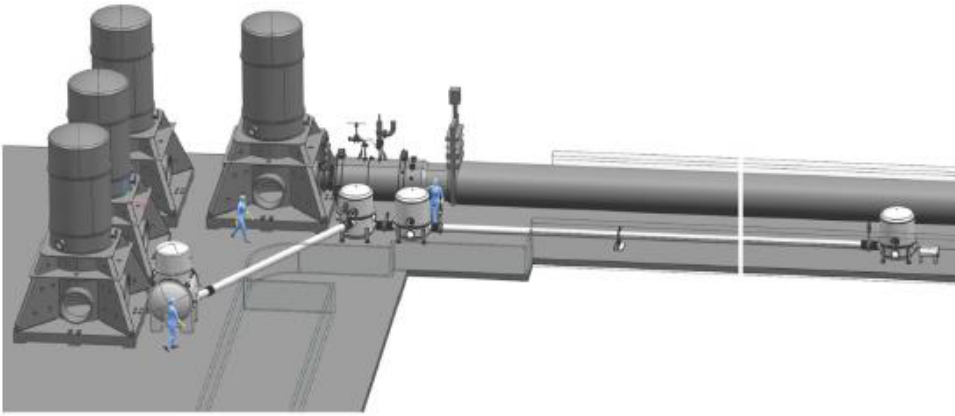
## Beyond O3: AdV+



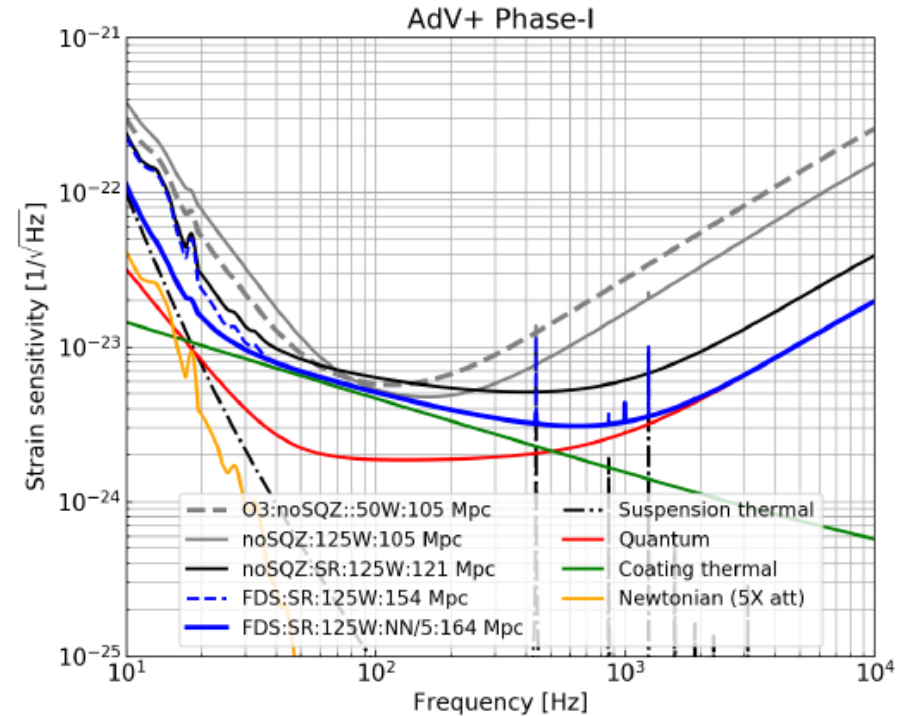
- ▶ Strategy: long observing runs and AdV+ upgrades
- ▶ A two step approach:
  - For O4: Frequency dependent squeezing
    - ▶ + Complete AdV: signal recycling
  - For O5: Reduce the coating thermal noise
    - ▶ Larger beam: Four test masses or just the end test masses?
    - ▶ Improved coating
- ▶ AdV+ proposed to the EGO Council last December



# AdV+ phase I



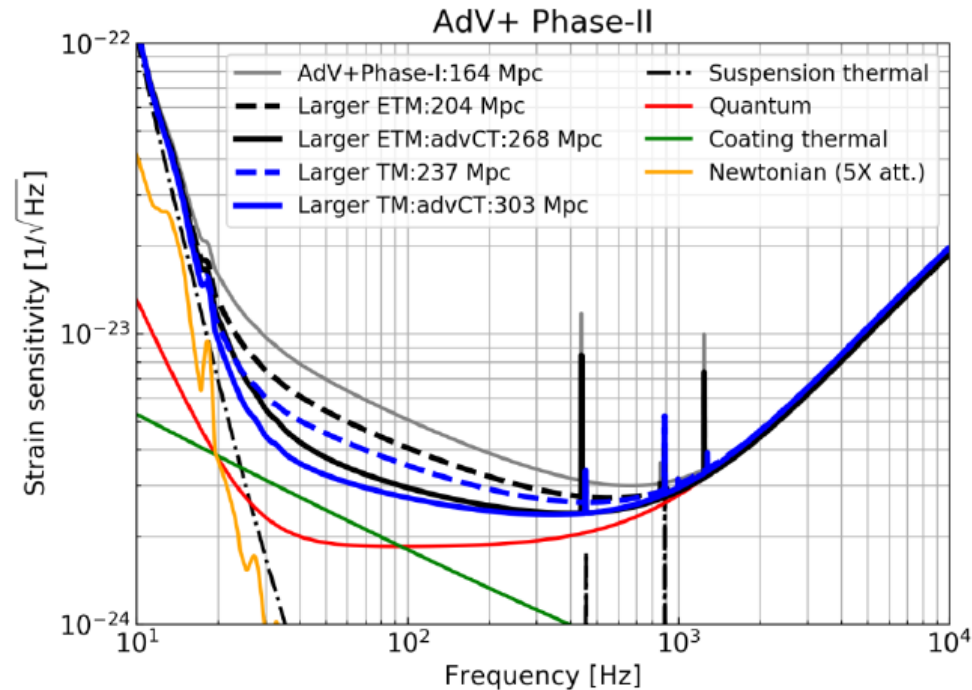
- ▶ Complete the AdV program:
  - 200 W laser; 125W at the ITF input
  - Signal recycling  $\rightarrow$  120 Mpc
- ▶ Frequency dependent squeezing
  - New filtering cavity
  - $\rightarrow$  150 Mpc
- ▶ Newtonian noise cancellation
  - $\rightarrow$  160 Mpc



# AdV+ Phase II



- ▶ Larger mirrors
  - Diameter: 550 mm, thickness: 200 mm, mass: 105 kg (?)
  - Scenario 1: ETM-only → 200 Mpc
  - Scenario 2: full upgrade → 230 Mpc
- ▶ Coating improvements
  - If factor three reduction in CTN:
    - ▶ Scenario 1: ETM-only → 260 Mpc
    - ▶ Scenario 2: full upgrade → 300 Mpc
- ▶ Many challenges and activities
  - Grand Coater upgrade
  - Vacuum, infrastructure
  - Payloads and superattenuators
  - Aberration control



# Budget

Phase	Item	Completed	BNS range [ Mpc ]	Cost [ M€ ]
AdV-O3		2019	65 - 80	
AdV design	Tuned signal recycling + 125 W	2021	120	
Phase I	Frequency dependent squeezing (FDS)	2021	150	3.5
	Newtonian noise cancellation (NNC)	2021	160	1.0
Phase II	Scenario 1: Large mirrors, ETM-only (LM1)			11.7
	Scenario 1 net cost: FDS + NNC + LM1			16.2
	Contingency (20 %)			3.2
	<b>Scenario 1 - ETM-only upgrade cost</b>	2023	200 - 260	<b>19.5</b>
	Scenario 2: Large mirrors, full upgrade (LM2)			20.8
	Scenario 2 net cost: FDS + NNC + LM2			25.3
	Contingency (20 %)			5.1
	<b>Scenario 2 - Full upgrade cost</b>		230 - 300	<b>30.4</b>
	Risk reduction NDRC			8.5

Not yet reviewed

- ▶ Hope to get (some) approval during the June EGO Council meeting...

# AdV+: a step toward 3G

- ▶ 3G detectors
  - Could see some sources all over the Universe
- ▶ Detector and Data Analysis challenges

