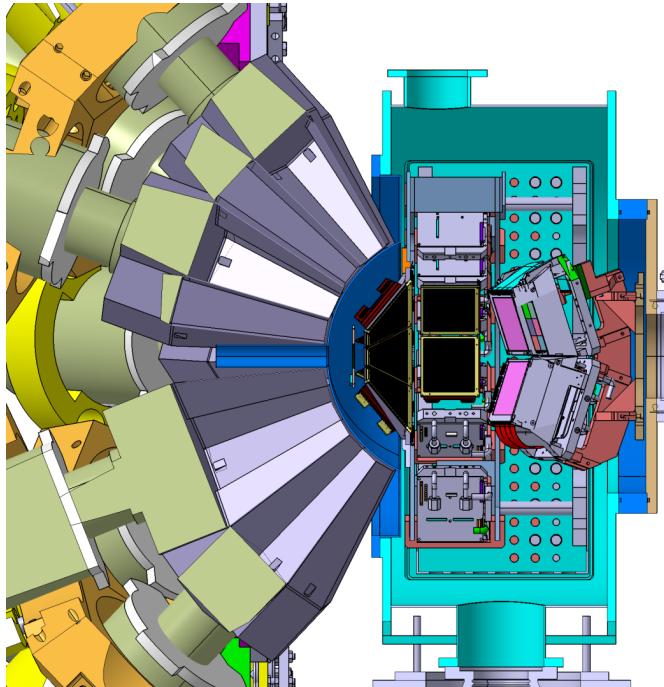




# MUGAST

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- A step towards GRIT
- Physics case & LoI
- Integration : mechanics & detectors /specific targets
- Electronics
- Planning & campaign

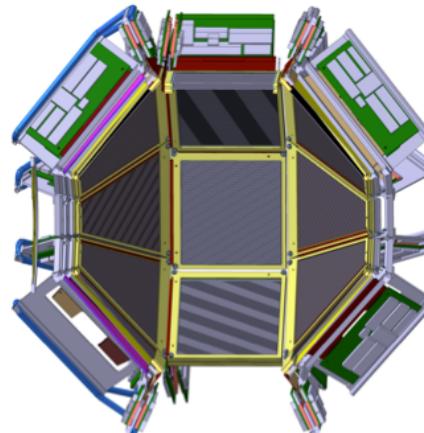
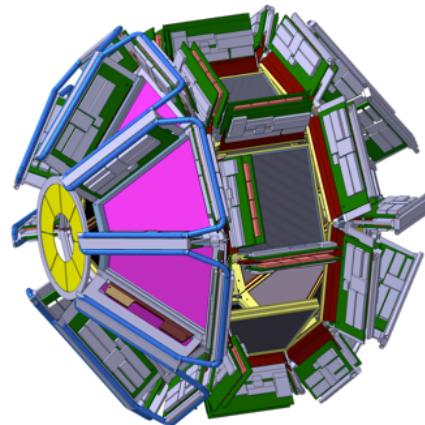
[grit.in2p3.fr/mugast](http://grit.in2p3.fr/mugast)

**M. Assié, IPN Orsay, assie@ipno.in2p3.fr**

# GRIT : Granularity Resolution Identification Transparency

**4 $\pi$  Si array fully integrable in PARIS/AGATA/EXOGAM2  
IPN Orsay / INFN-LNL / LPC Caen / BARC Mumbai / University of Surrey**

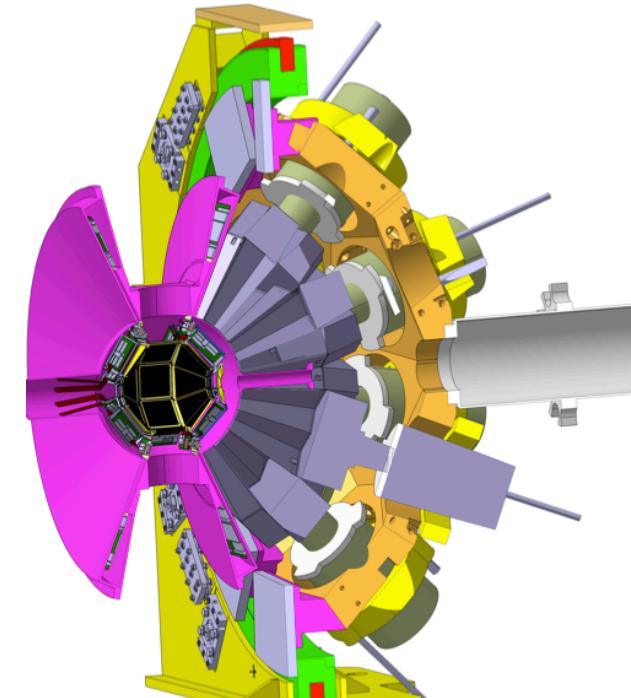
Direct reactions measurement and particle-gamma coincidence



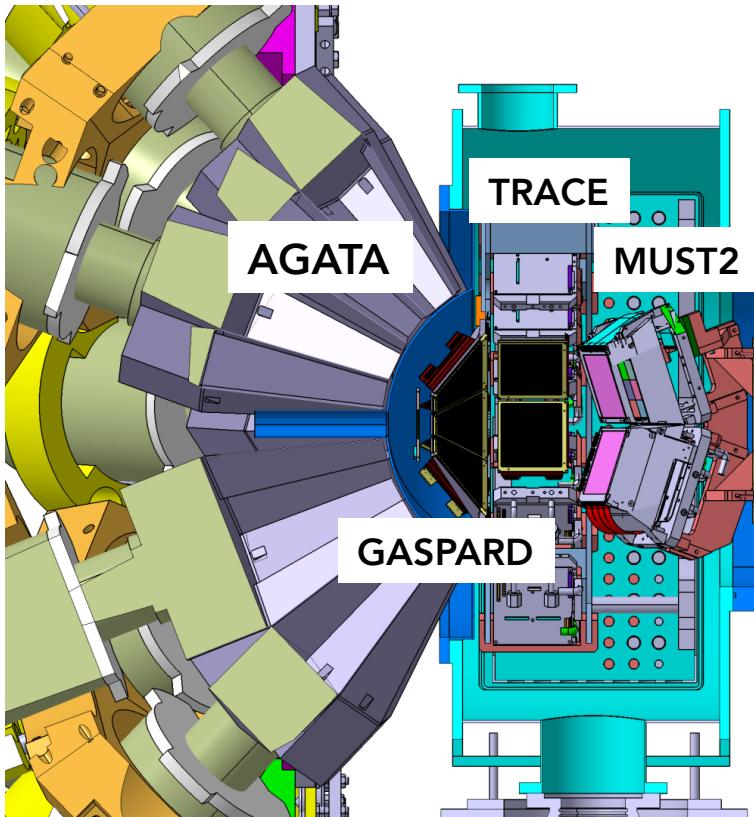
- . Combination of **energy resolution (gamma)** and good efficiency
- . Good **granularity** (pitch <1mm)
- . **PSD** to identify low energy particles (1<sup>st</sup> layer)
- . Integration of **special targets** (cryogenic ,Chymène...)
- . **Integrated electronics** (iPACI, PLAS) designed by IPN & INFN Legnaro

**Physics cases :**  
**13 LoI for Spiral2**

- . Shell evolution
- . Pairing
- . Clusters
- . Near Barrier reactions
- . Astrophysics



# MUGAST : a first step towards GASPARD



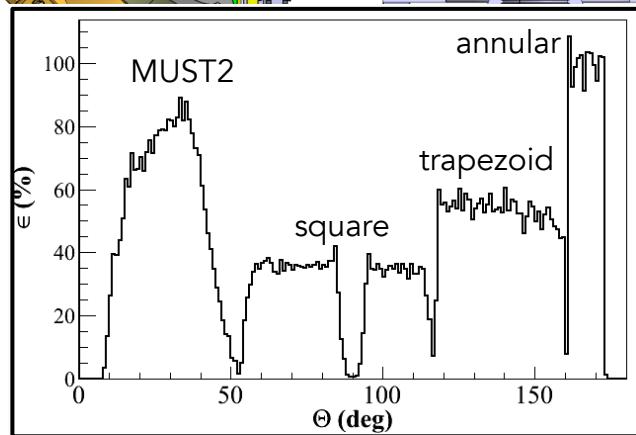
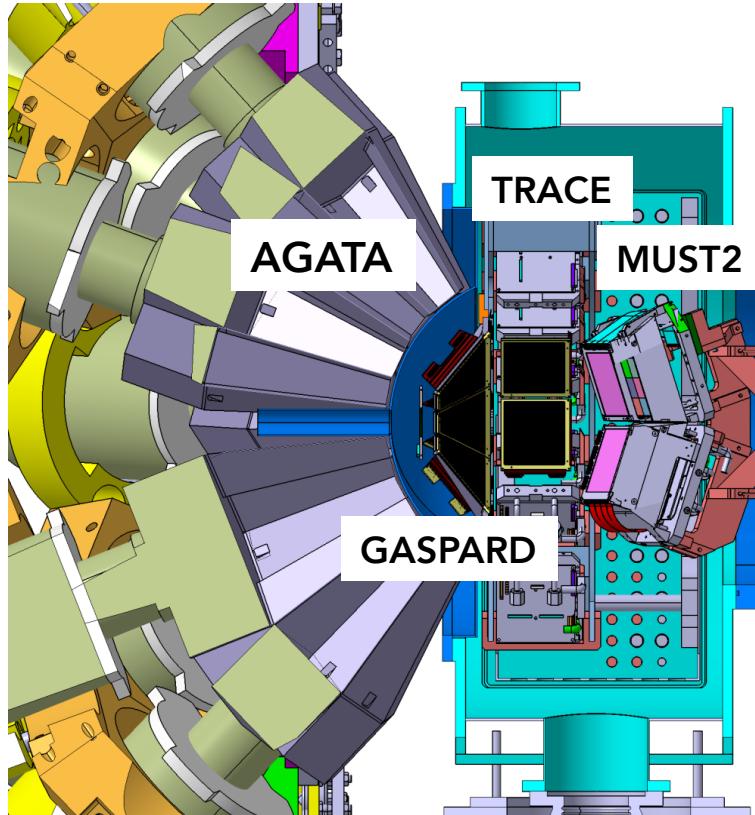
Campaign starting April 2019

- ▶ New Spiral 1 beams (low energy)
- ▶ AGATA
  - . very high energy resolution
  - . good efficiency : 13% at 1.3MeV in 2019 @ 18cm  
(depending on number of clusters)
- ▶ MUGAST
  - . one-layer of Silicon backward & 90 deg.  
--> GRIT detectors (5 trap. + 2 square)
  - + MUST2 (forward direction)  
→ well-suited for stripping measurements
  - . existing electronics (MUFE+MUVI)
  - . specific target (cryogenic)
- ▶ VAMOS : large acceptance spectrometer at 0 degree

**AGATA efficiency ( $E_\gamma = 1.3$  MeV)**

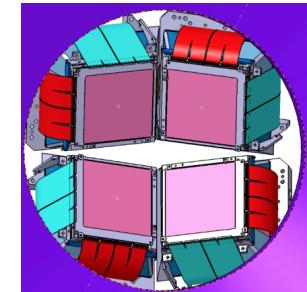
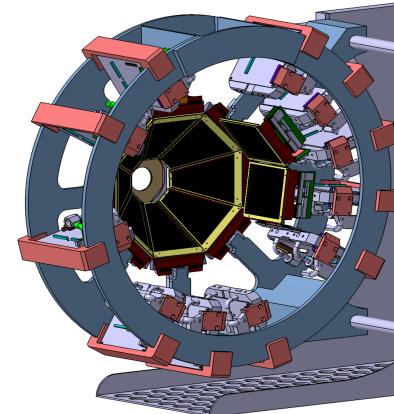
	2014/2015 (24 caps)	Now (32 caps)	2019 (45 caps)
D = 23.5cm	4%	5.3%	7.5%
<b>D = 18cm</b>		<b>8.5%</b>	<b>13%</b>

# MUGAST : a first step towards GASPARD



Campaign starting in April 2019

- ▶ New Spiral 1 beams (low energy)
- ▶ AGATA
  - . very high energy resolution
  - . good efficiency : 13% at 1.3MeV in 2019 @ 18cm  
(depending on number of clusters)
- ▶ MUGAST
  - . one-layer of Silicon backward & 90 deg.  
--> GASPARD-TRACE detectors (5 trap. + 2 square)
  - + MUST2 (forward direction)
    - well-suited for stripping measurements
  - . existing electronics (MUFEET+MUVI)
  - . specific target (cryogenic)
- ▶ VAMOS : large acceptance spectrometer at 0 degree



# Lol for MUGAST

## PAC 2015 & 2016

17 Lol submitted in the "umbrella" Lol

- > mainly stripping reactions (backward)
- > Spiral1 & Spiral1 new beams mostly

7 Lol submitted independently (Spiral1 beams)

### Nuclear astrophysics

- $^{25}\text{Al}(\text{He},\text{d})$  (*N.de Séréville, F. Hammache, IPNO*)
- $^{30}\text{P}(\text{He},\text{d})$  or  $(\text{d},\text{p})$  (*N.de Séréville, F. Hammache, IPNO*)
- $^{79}\text{Se}(\text{d},\text{p})^{80}\text{Se}$  (*G. de Angelis, INFN-LNL*)

### Shell evolution

- $^{56}\text{Ni}(\text{d},\text{p})(\text{d},\text{t})$  (*F.Flavigny, IPNO, O.Sorlin, GANIL*)
- $^{28}\text{Mg}(\text{d},\text{p})$  (*A.Matta, W.Carford, University of Surrey*)
- $^{46,48}\text{Ar}(\text{t},\text{p})$  (*D.Mengoni, University of Padova*)
- $^{67}\text{As}(\text{He},\text{d})$  (*D. Mengoni, A. Goasduff*)

### Clusters, pairing, correlations & others

- $^{56}\text{Ni}(\text{He},\text{p})(\text{Li},\alpha)$  (*M.Assie, IPNO*)

## Letters of Intent Presented to the GANIL PAC, June 2016

### Summary

The science programme described by the Lols was strong. In particular the PAC recognises the opportunity that the combination of MUGAST, VAMOS and AGATA presents and it suggests that this programme be made a priority for future calls for proposals.

### Scientific council of GANIL

In 2019, the priority has to be given either to the VAMOS stand-alone mode or to the MUGAST+VAMOS collaboration, provided that the construction of the MUGAST detector is completed in time.

In view of the large amount of time required to analyze AGATA data, the SC considers that much work still needs to be done to extract the science gathered in the first VAMOS + AGATA campaign and that, hence, a fruitful MUGAST + VAMOS + AGATA campaign would be most productive in 2019.

# MUGAST Proposals & Lol

Reaction	Spokesperson	Comment
Accepted experiments		
$^{15}\text{O}(\text{d},\text{p})^{19}\text{Ne}$	C. Diget, N. de Séréville	PAC 2017
$^{14}\text{O}(\text{p},\text{p}')$	I. Stefan, F. de Oliveira	to be redone
Proposed experiment PAC 2018		
$^{45}\text{Ar}(\text{d},\text{p})^{46}\text{Ar}$	A. Lemasson	
$^{19}\text{O}(\text{d},\text{p})^{20}\text{O}$	E. Clément	DSAM
$^{47}\text{K}(\text{t}, \text{He}^4)\text{Ar}^4$ or $^{46}\text{Ar}(\text{He}^3,\text{d})^{47}\text{K}$	A. Gottardo	cryogenic target
$^{14}\text{O}$ coulomb BU	D. Gupta	
$^{85}\text{Kr}(\text{d},\text{p})^{86}\text{Kr}$	F. Recchia	isomeric state (5h)
$^{46}\text{Ar}(\text{d},\text{p})^{47}\text{Ar}$	D. Mengoni	
$^{28}\text{Mg}(\text{d},\text{p})^{29}\text{Mg}$	A. Matta	new Spiral 1 beam
Lol		
$^{56}\text{Ni}(\text{d},\text{p})^{57}\text{Ni}$	F. Flavigny, O. Sorlin	
$^{56}\text{Ni}(\text{He}^3,\text{p})^{58}\text{Cu}$	M. Assié	
$^{69}\text{Cu}, \text{Ar}^{46}, \text{Sc}^{49-50} (\text{t}, \text{He}^4)$	S. Bottoni	

# Detectors for MUGAST

## Trapezoidal DSSSD

Received from Micron : nTD, 4 deg. cut, 6 inches

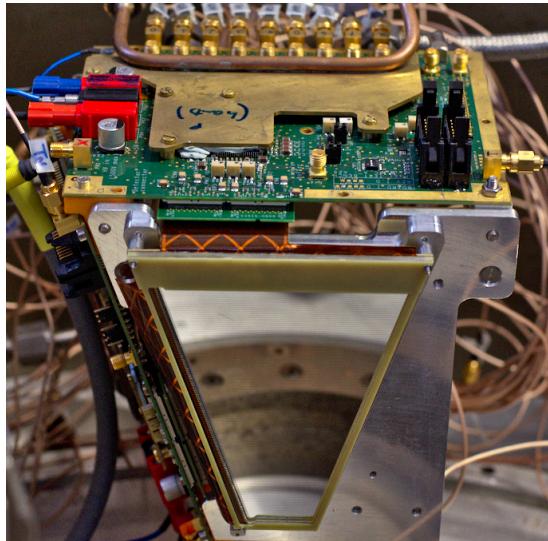
- 2 prototypes 500 um IPNO (june 2015)
- 3 pre-series U. Surrey, Santiago, IPNO (feb. 2017)

Test bench @ IPNO :

- numerical (iPACI): 9X+9Y
- analog (MUFEE) : 128X+128Y

In-beam test @ ALTO

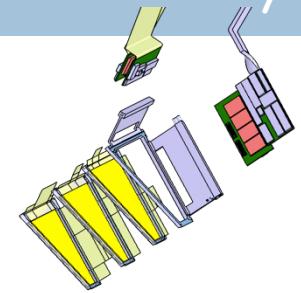
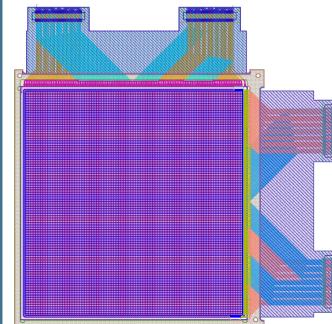
end of 2017.



## Squared DSSSD

Received from Micron Semiconductors :

- 2 prototypes 500um INFN (beg. 2017)
- 1 prototype 1.5mm INFN (beg. 2017)



# Detectors for MUGAST

## Trapezoidal DSSSD

Received from Micron : nTD, 4 deg. cut, 6 inches

- 2 prototypes 500 um IPNO (june 2015)
- 3 pre-series U. Surrey, Santiago, IPNO (feb. 2017)

Test bench @ IPNO :

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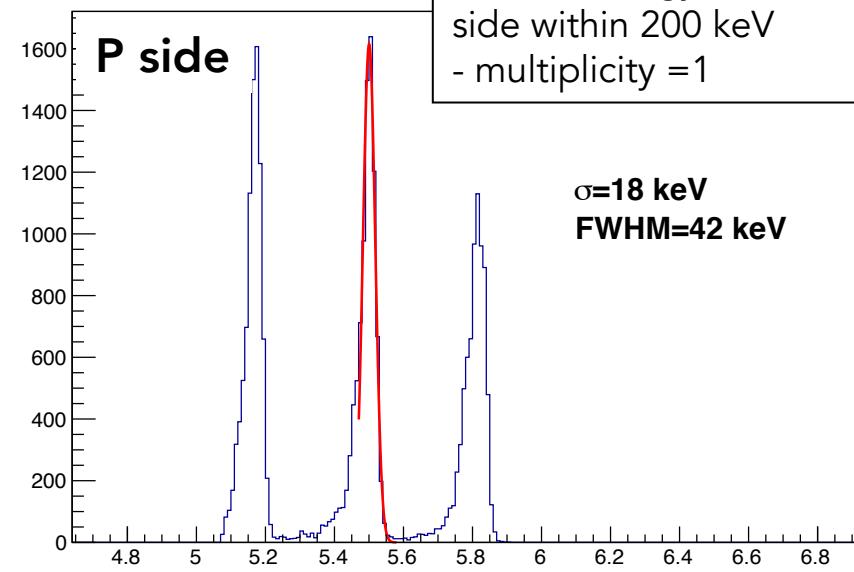
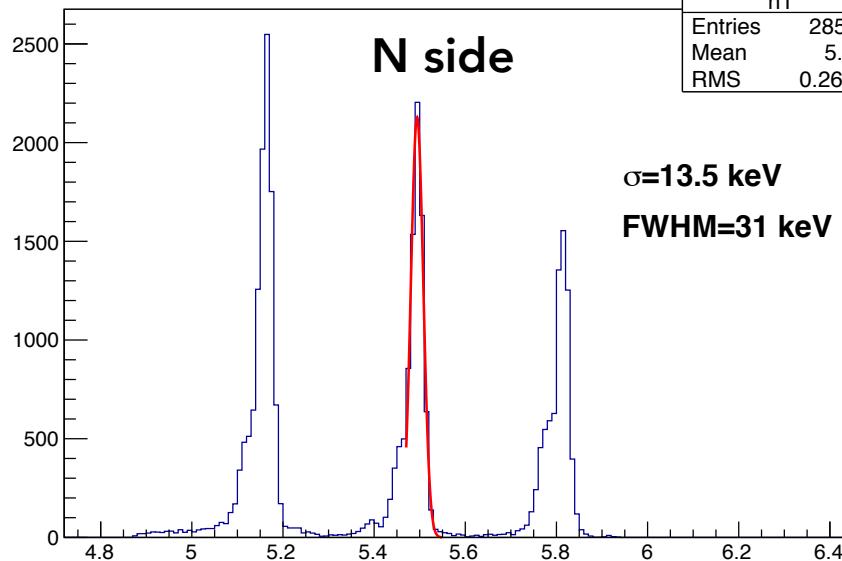
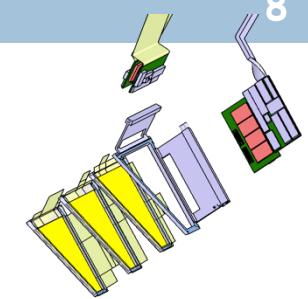
In-beam test @ ALTO

end of 2017.

## Squared DSSSD

Received from Micron Semiconductors :

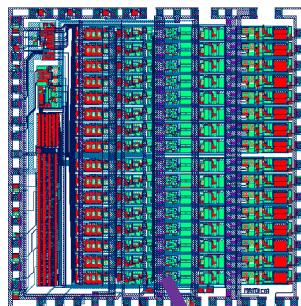
- 2 prototypes 500um INFN (beg. 2017)
- 1 prototype 1.5mm INFN (beg. 2017)



Conditions applied :

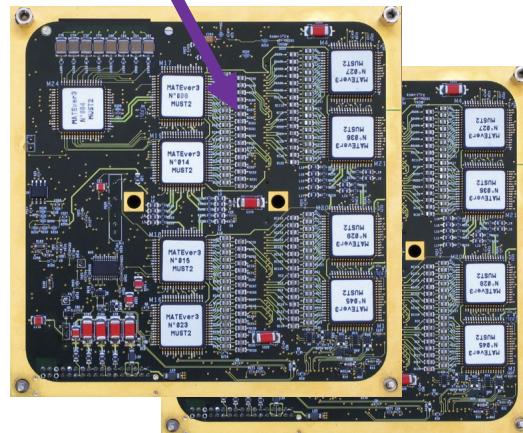
- same energy on both side within 200 keV
- multiplicity = 1

# Electronics for MUGAST



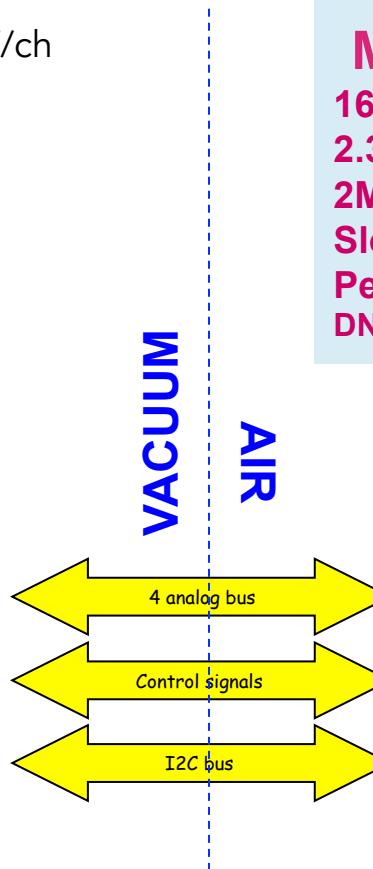
- 16 channels 28 mW/ch
  - Energy & Time
  - Si, Si(Li) and CsI
  - Multiplexer
  - I2C interface
  - High linear. pulser
  - T sensor

## MOTHER BOARDS (IPNO)



## MUFEE X & Y for 1 telescope

## 10 MUFEE pairs available for MUST2 (10 tel)



MUVI : VXI board (GANIL)

16 ADC14 bits

## 2.3K parameters

2MHz

## Slow Control I2C

## Pedestal subtraction

## DNL correction

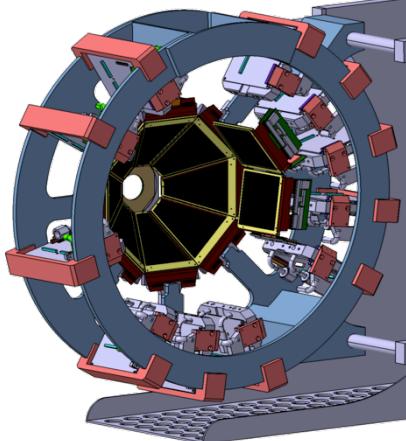
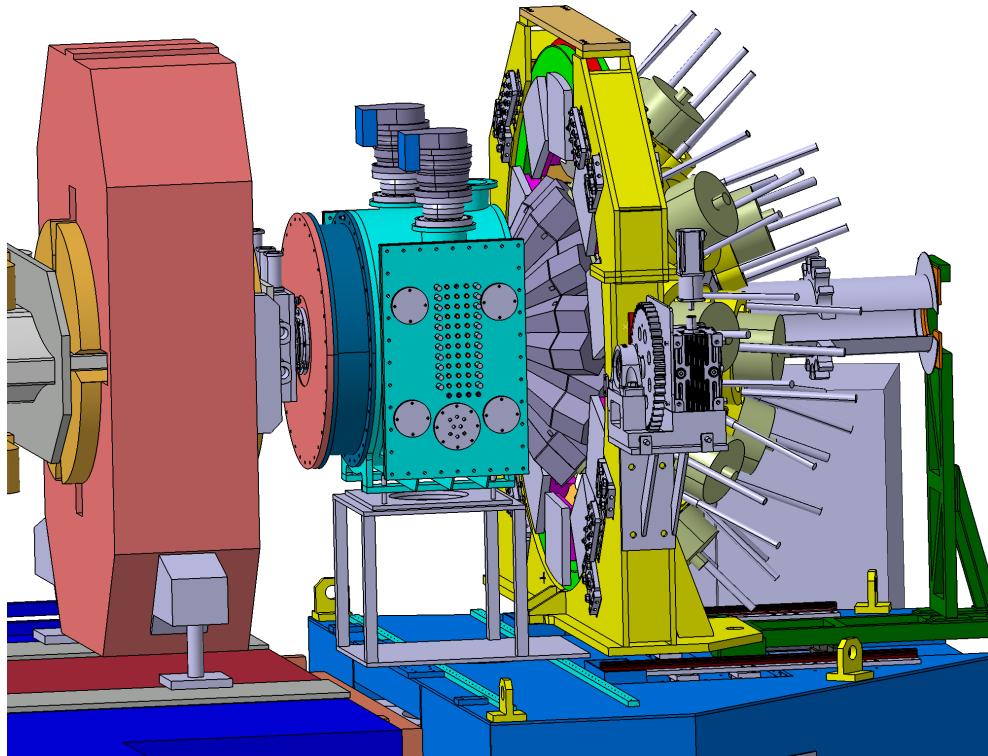


# 4 telescopes

## 4 MUVI available for MUST2 (16 telescopes)

**7 MUFEE pairs (X&Y) redone in 2016 (components & cabling)  
tested with pulser at IPN Orsay in 2018  
Coupling to be done at GANIL now !**

# Mechanics for MUGAST



*Internal mechanics to hold electronics and detectors*

**Reaction chamber design :**

IPN by E. Rindel & L Vatrinet

- 2 parts : main for GRIT

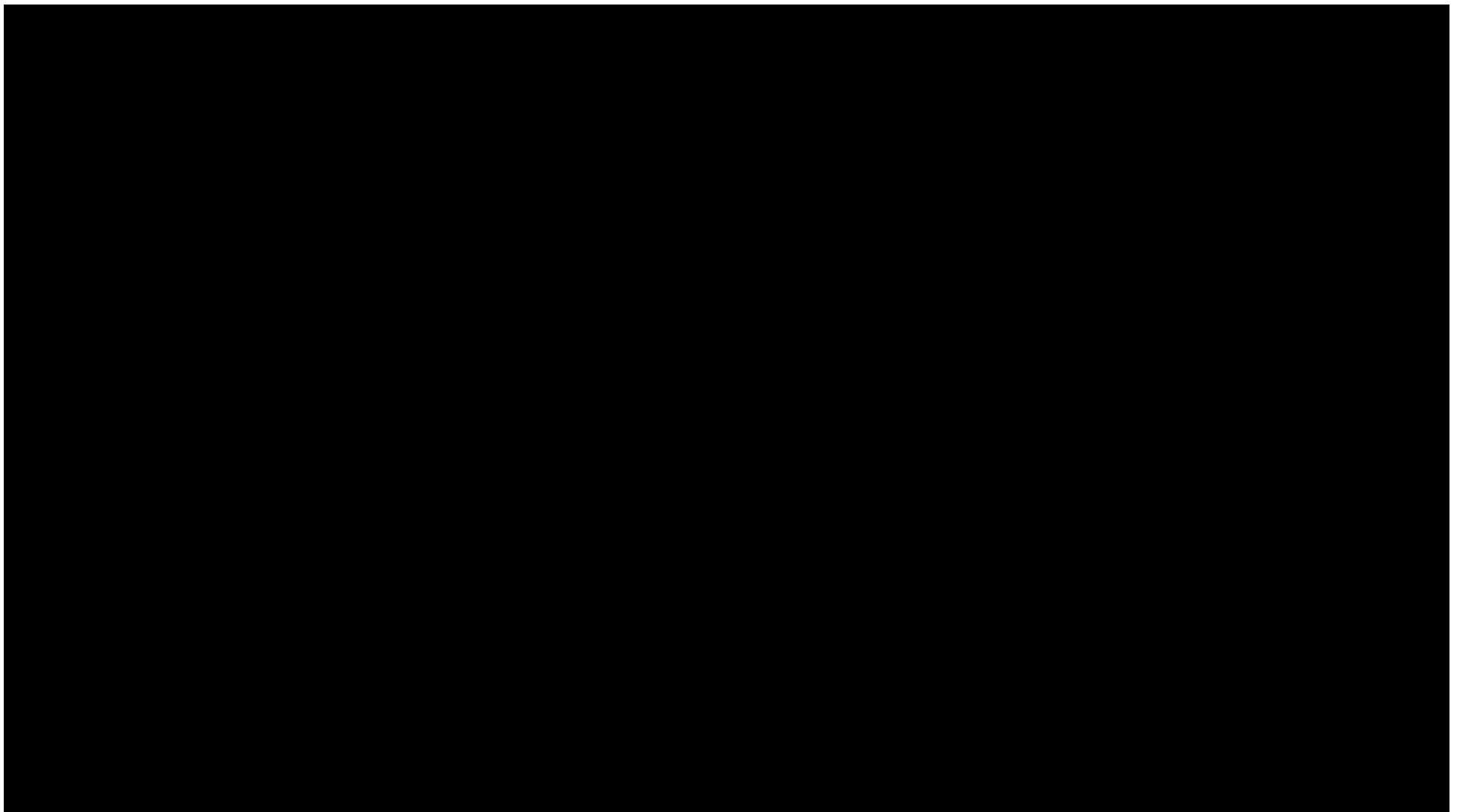
extension for MUST2 (dark blue)  
with Vamos at 0 degree

- Mechanical support for detectors in order to  
maximize transparency

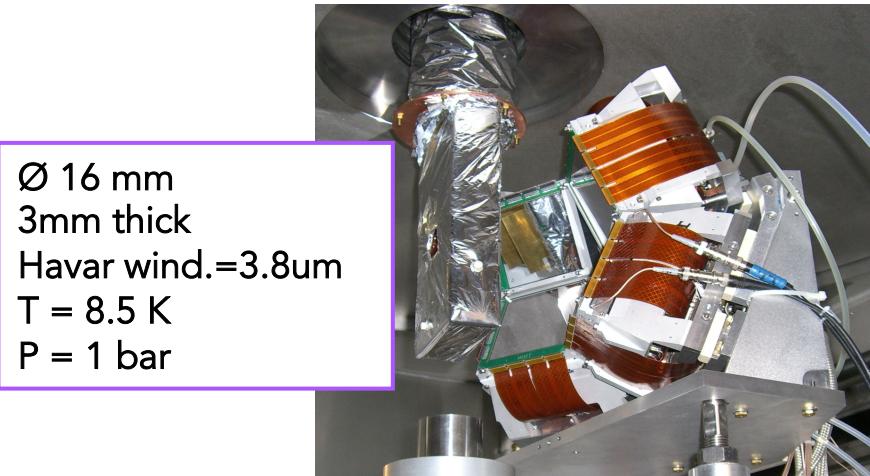
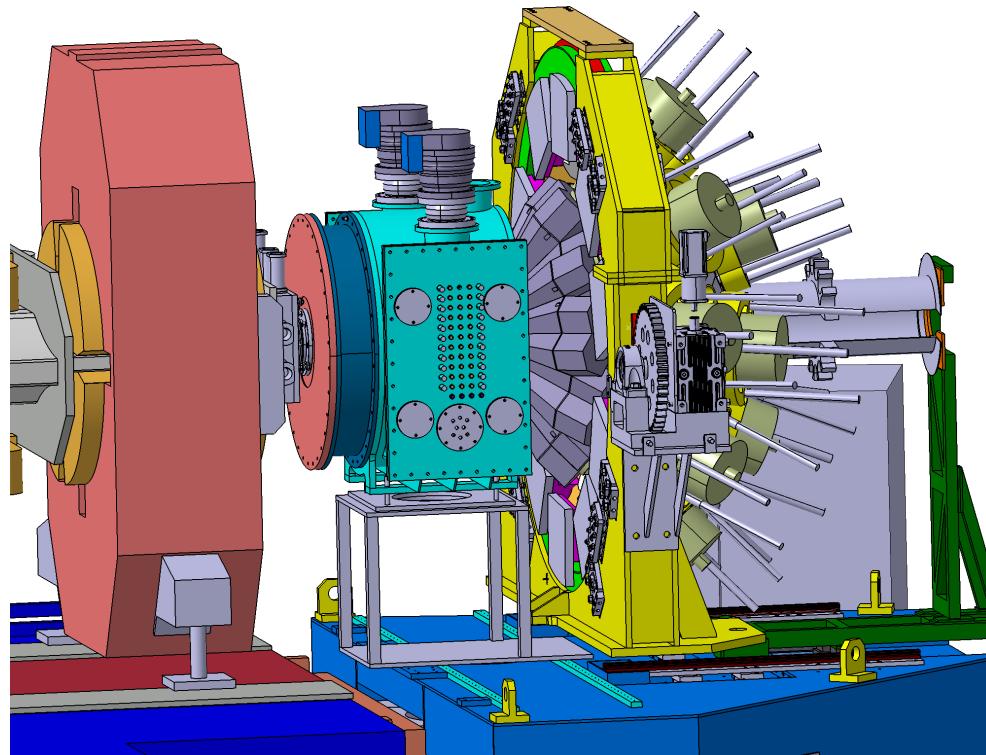


# Mechanics for MUGAST

E. Rindel, J. Bettane IPN



# Mechanics for MUGAST

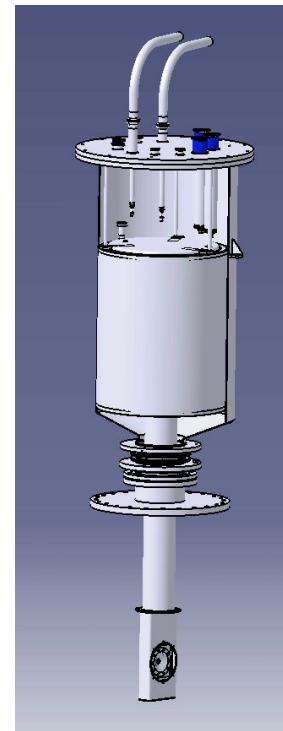


**Reaction chamber design : done ! (2016)**

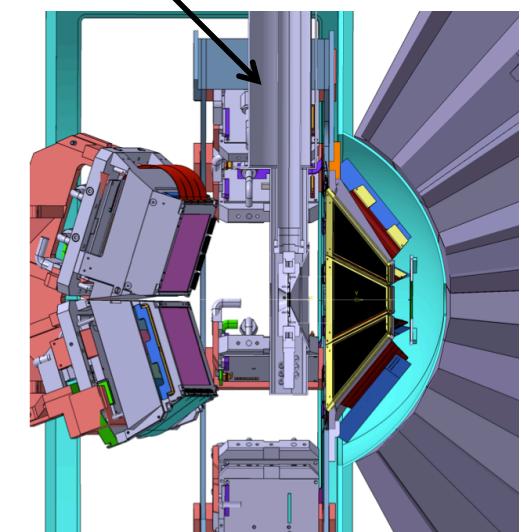
- 2 parts : main for GRIT/MUGAST  
extension for MUST2 (dark blue)

- Mechanical support for detectors in order to maximize transparency

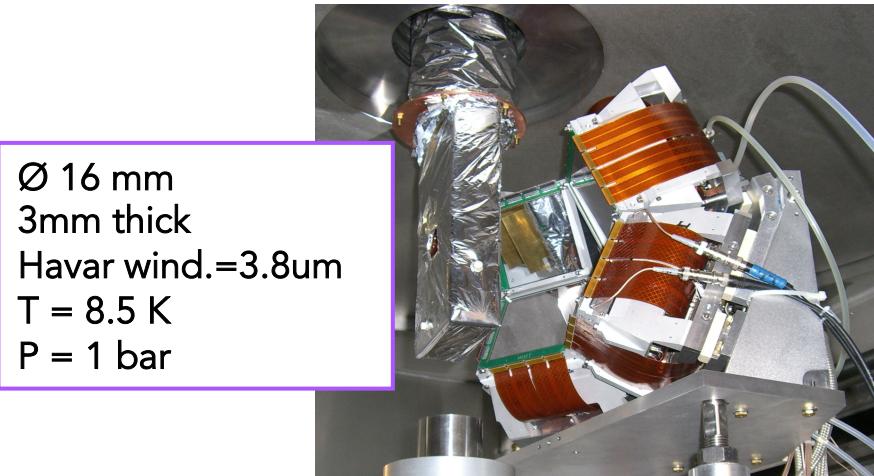
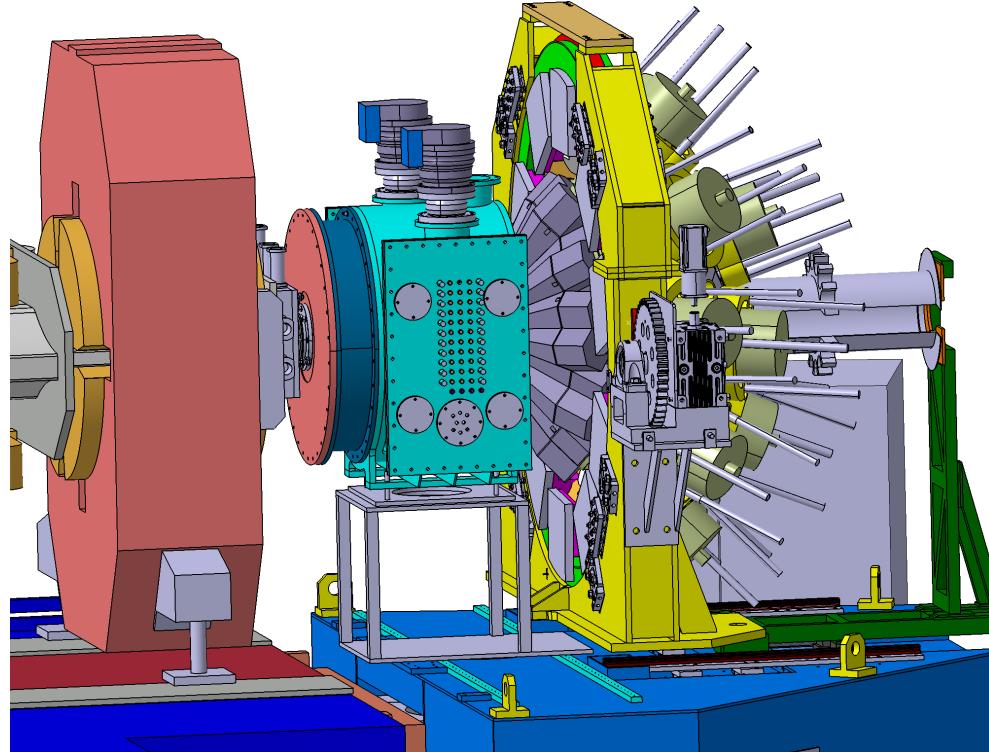
- integration of special target (cryogenic)
  - . redesign of existing target to fit
  - . conversion for  $^3\text{He}$  (IPN – 2018)



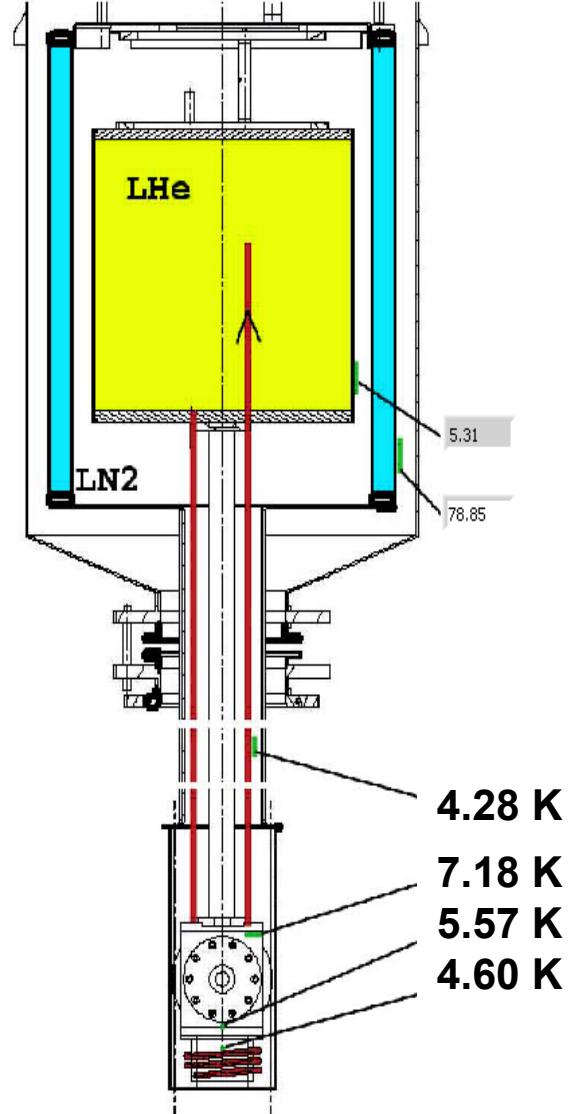
**cryogenic target**



# Mechanics for MUGAST

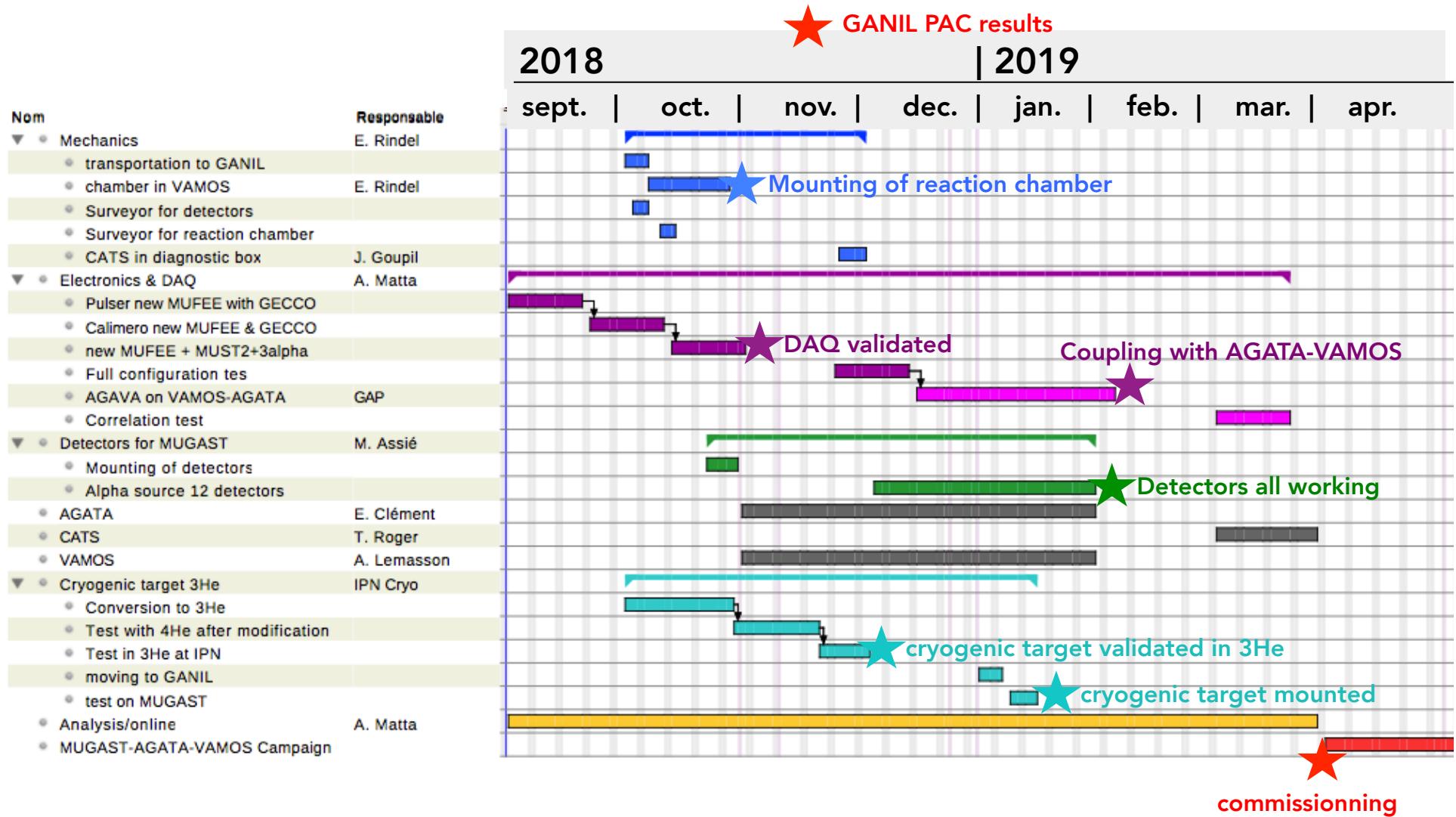


Ø 16 mm  
3mm thick  
Havar wind.=3.8um  
 $T = 8.5 \text{ K}$   
 $P = 1 \text{ bar}$



F. Galet, M. Pierens, H. Saugnac, IPN

# Planning



# Conclusion

- AGATA-VAMOS-MUGAST campaign starting spring 2019  
experiments submitted to the forthcoming PAC (Nov. 2018)  
campaign of 3 to 5 experiments foreseen.

Mounting starts Oct. 2018

Tests of detectors Dec. 2018

Coupling with VAMOS-AGATA jan-fev 2018

Commissioning : April 2019 ?

- Cryogenic 4He/3He target will be available
- Tritium target will be ordered if proposals accepted
- Possibility to have a second campaign ?  
more Spiral1 new beams available  
LoI for these new beams submitted

