

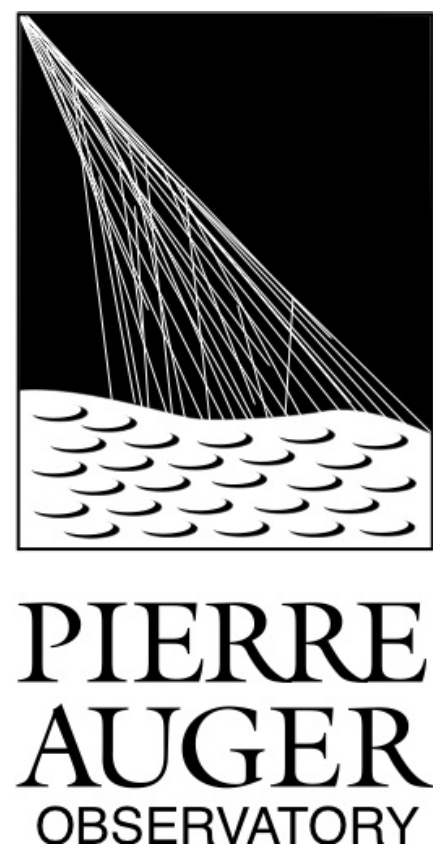
Preliminary results of the AMIGA engineering array at the Pierre Auger Observatory

ALVARO TABOADA^{1,2} for the Pierre Auger Collaboration³

¹Karlsruhe Institute of Technology, Karlsruhe, Germany

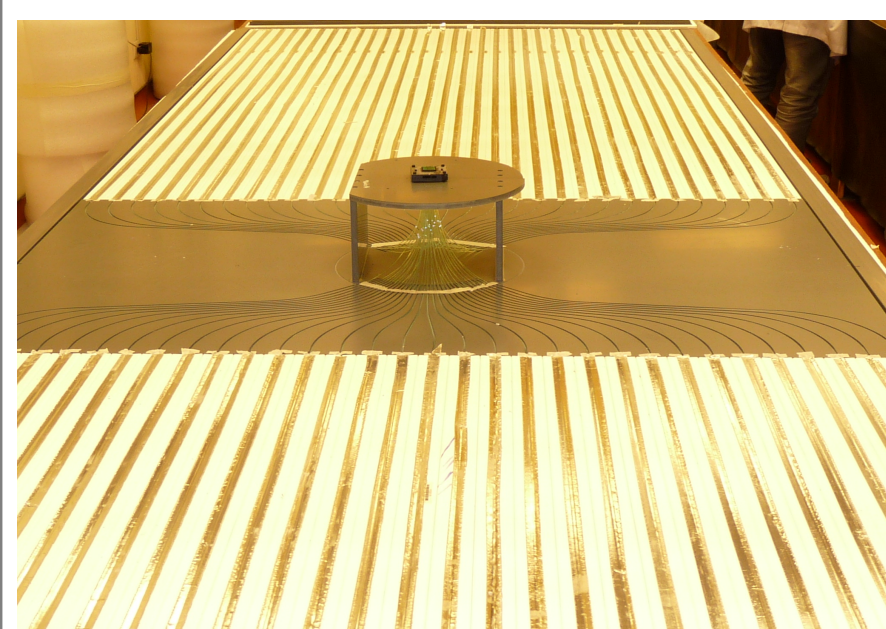
²Instituto de Tecnologías en Detección y Astropartículas, Buenos Aires, Argentina

³Observatorio Pierre Auger, Av. San Martín Norte 304, 5613 Malargüe, Argentina



1. The AMIGA Engineering Array

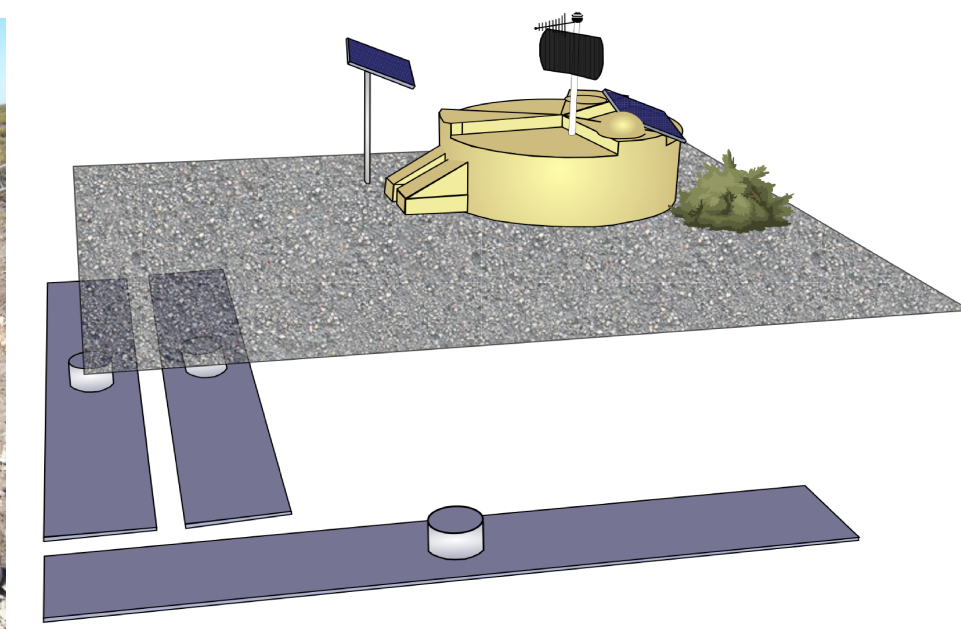
The AMIGA muon detector array (MD) will provide Muon Counters for the whole SD-750 m spacing array. Between March 2015 - November 2017, an engineering array (EA) of the MD was operating as a proof of concept of the detector and to optimize the design, where the main upgrade was the change of PMTs by SiPMs and the introduction of an specific channel to measure high muon densities close to the shower core. Each station on the array is composed by a pair of Water-Cherenkov Detector and a 30 m² Muon Counter buried at 2.3 m depth. The Muon Counters at the EA were a combination of 5 m² and 10 m² modules, each filled with 64 scintillator bars and wavelength-shifting fibers which guide the scintillation light to a SiPM array. The AMIGA The deployment of the AMIGA MD array with SiPMs started in January 2018. So far, 7 stations were installed. The production rate is 10 m²/week.



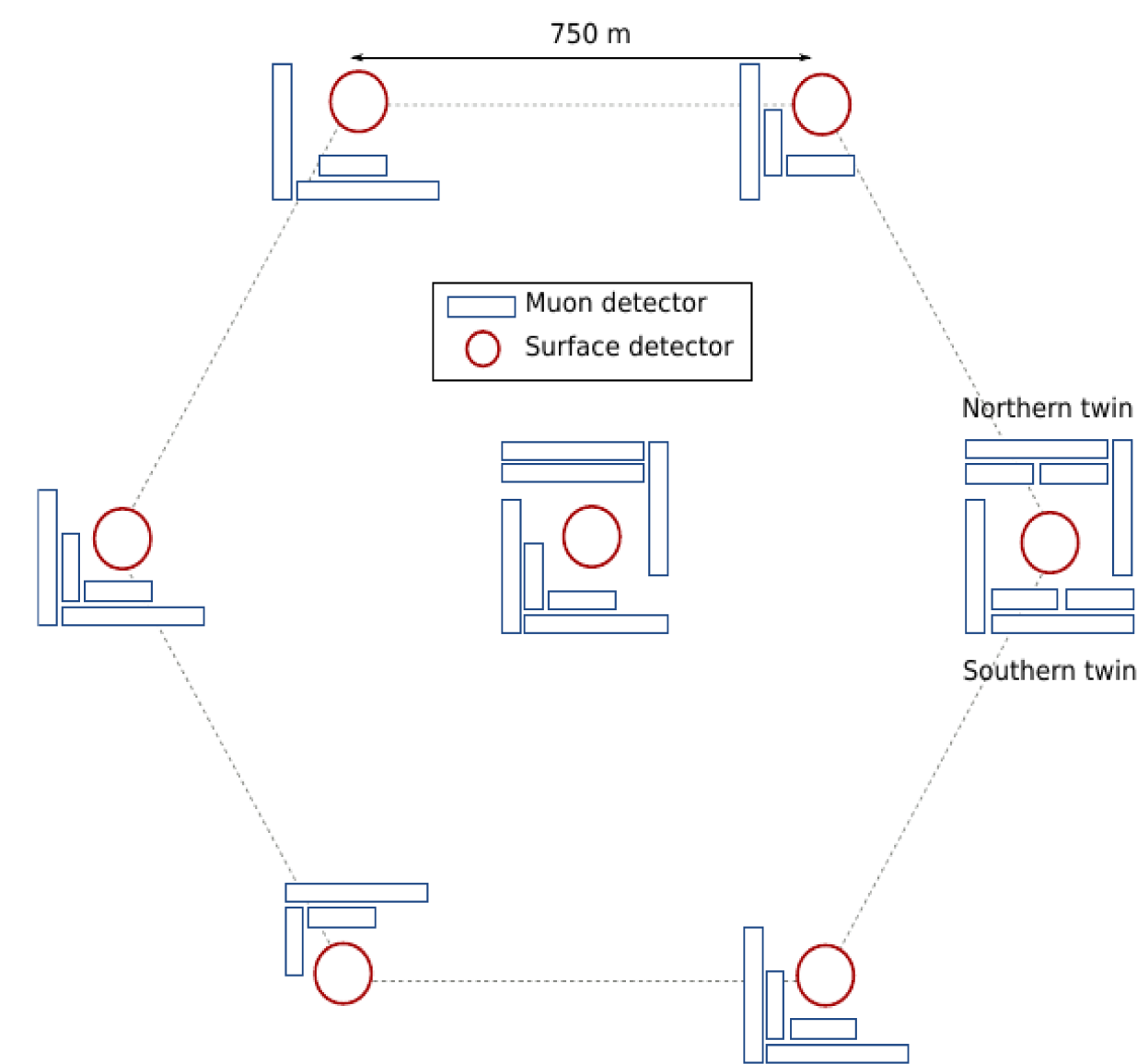
64 scintillator bars and fibers per module



Deployment of AMIGA module



Sketch of the 30 m² AMIGA station



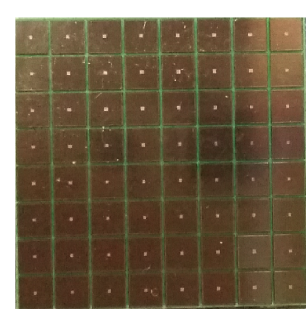
Map of the AMIGA Engineering Array

2. SiPMs

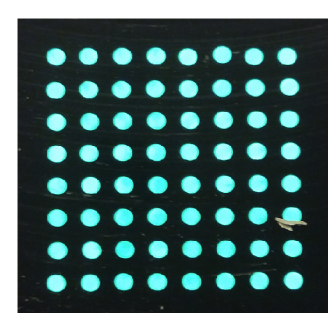
64 SiPMs array coupled to WLS fibers.

Sources of uncertainty

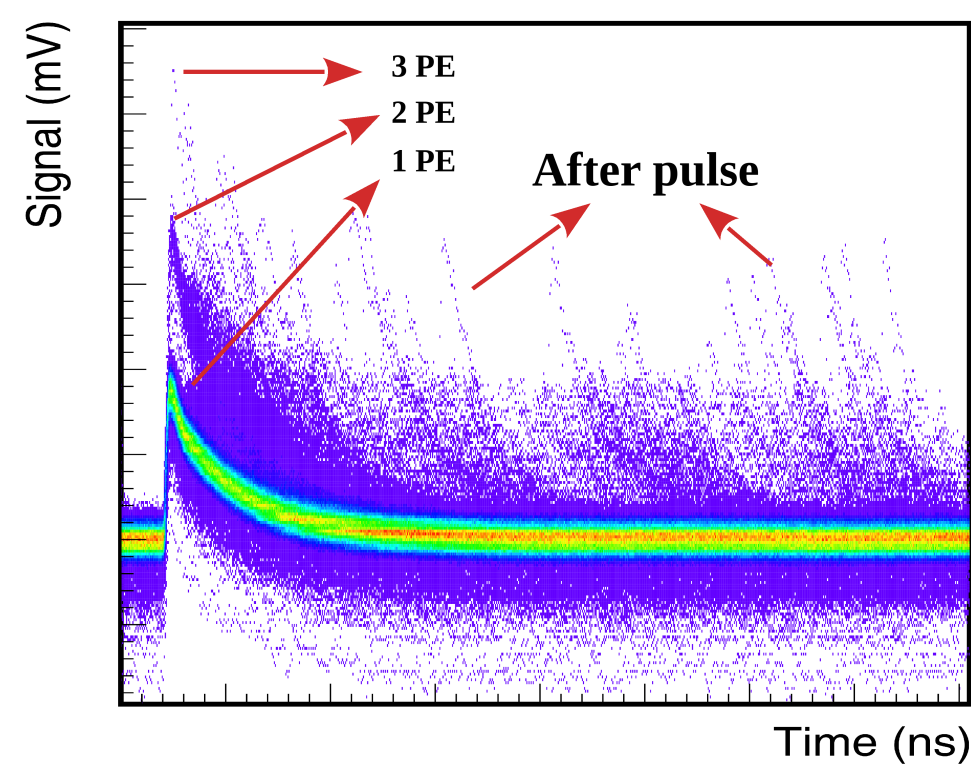
- Dark Rate
- Cross-Talk
- Temperature Dependency



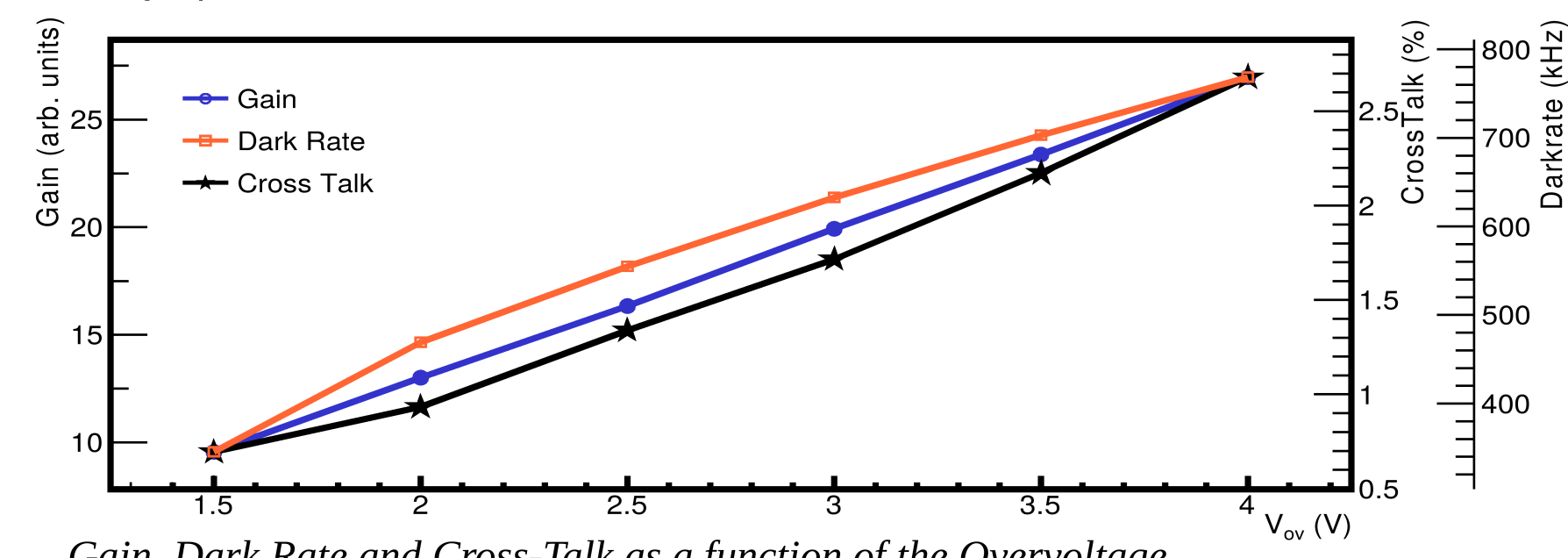
Array of SiPMs



Optical coupling

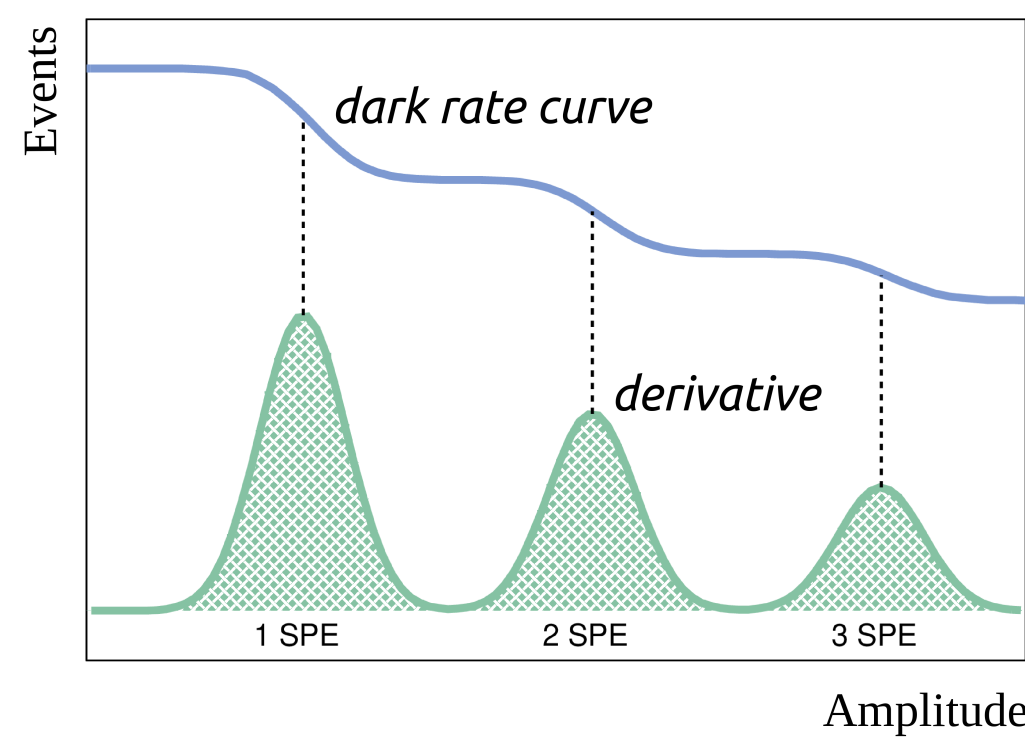


Dark rate pulses



Gain, Dark Rate and Cross-Talk as a function of the Overvoltage

3. Calibration



The curve of dark rate is obtained varying the threshold level and its derivative gives the amplitude of PE. This is used to calibrate the SiPM. As the Gain depends on the Overvoltage (V_{Ov}), in order to achieve a uniform gain over the array of SiPMs, the Breakdown voltage (V_{Br}) is calculated for each channel and the V_{Ov} is set as:

$$V_{Ov} = V_{BIAS} - V_{Br}$$

4. Event Reconstruction

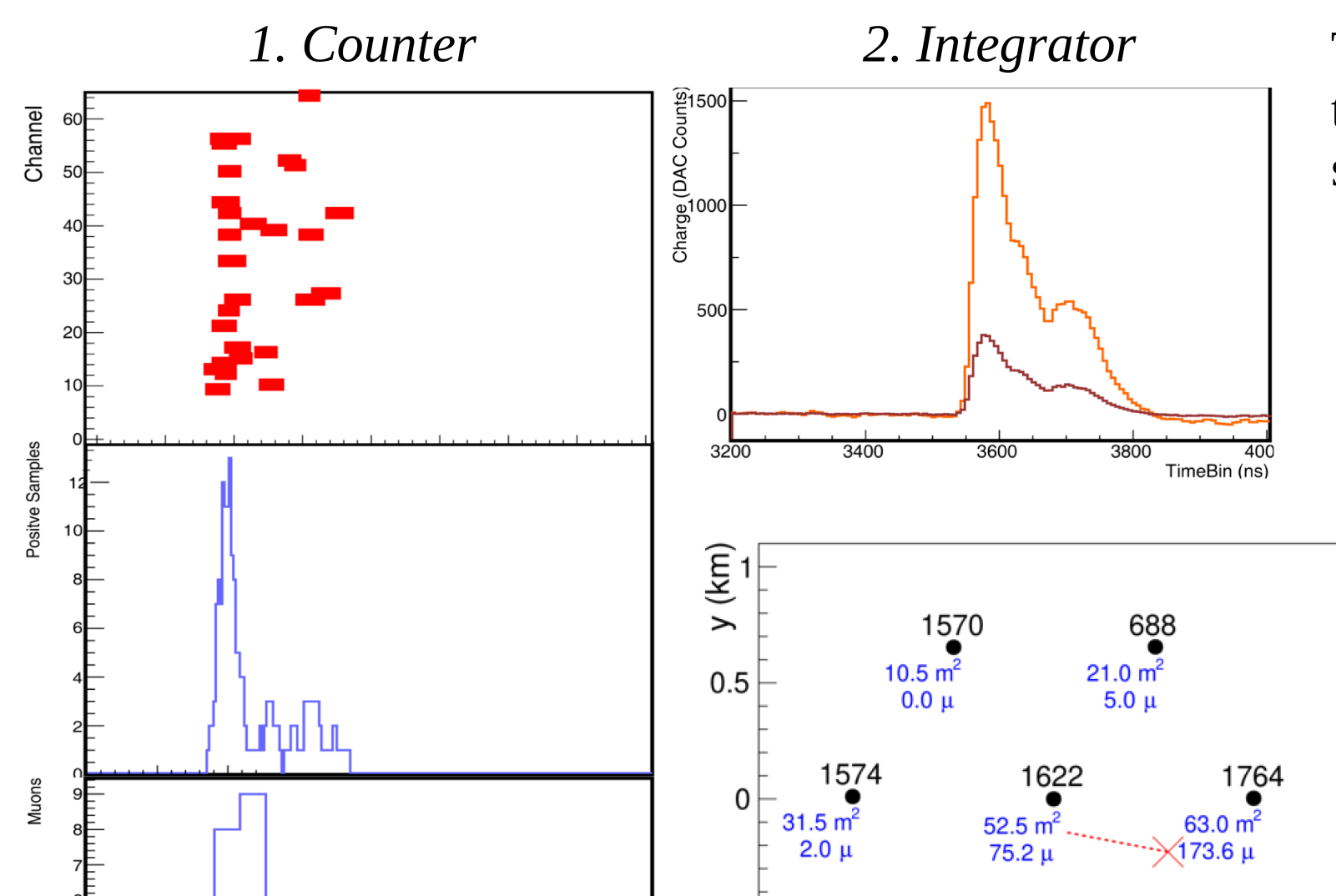
The muon density ρ_μ is estimated via two different ways: **counter & integrator**

1. Counter

SiPMs signal consists in a trace of bits. When signal is above some discrimination threshold muons are counted. The counter is useful to measure **low densities** with minimal fluctuations.

2. Integrator

Signal is obtained by integrating the



The **event reconstruction** procedure consists in **fitting** the **muon density** as a function of the distance to the shower core of the form:

$$\rho_\mu(r, E, \theta) \propto \left(\frac{r}{r_*}\right)^{-\alpha} \left(1 + \frac{r}{r_*}\right)^{-\beta(E, \theta)} \left(1 + \left(\frac{r}{10r_*}\right)^2\right)^{-\gamma}$$

