



The Cosmic Ray Energy Spectrum between 2 PeV and 2EeV Observed with the TALE Detector in Monocular Mode.

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(University of Utah)

Currently visiting Jerusalem

for the TA Collab.

Presented by

Charlie Jui

Oct 9, UHECR 2018

Paris, France



Tareq's son Ziad from Jerusalem, Oct 8



Telescope Array collaboration

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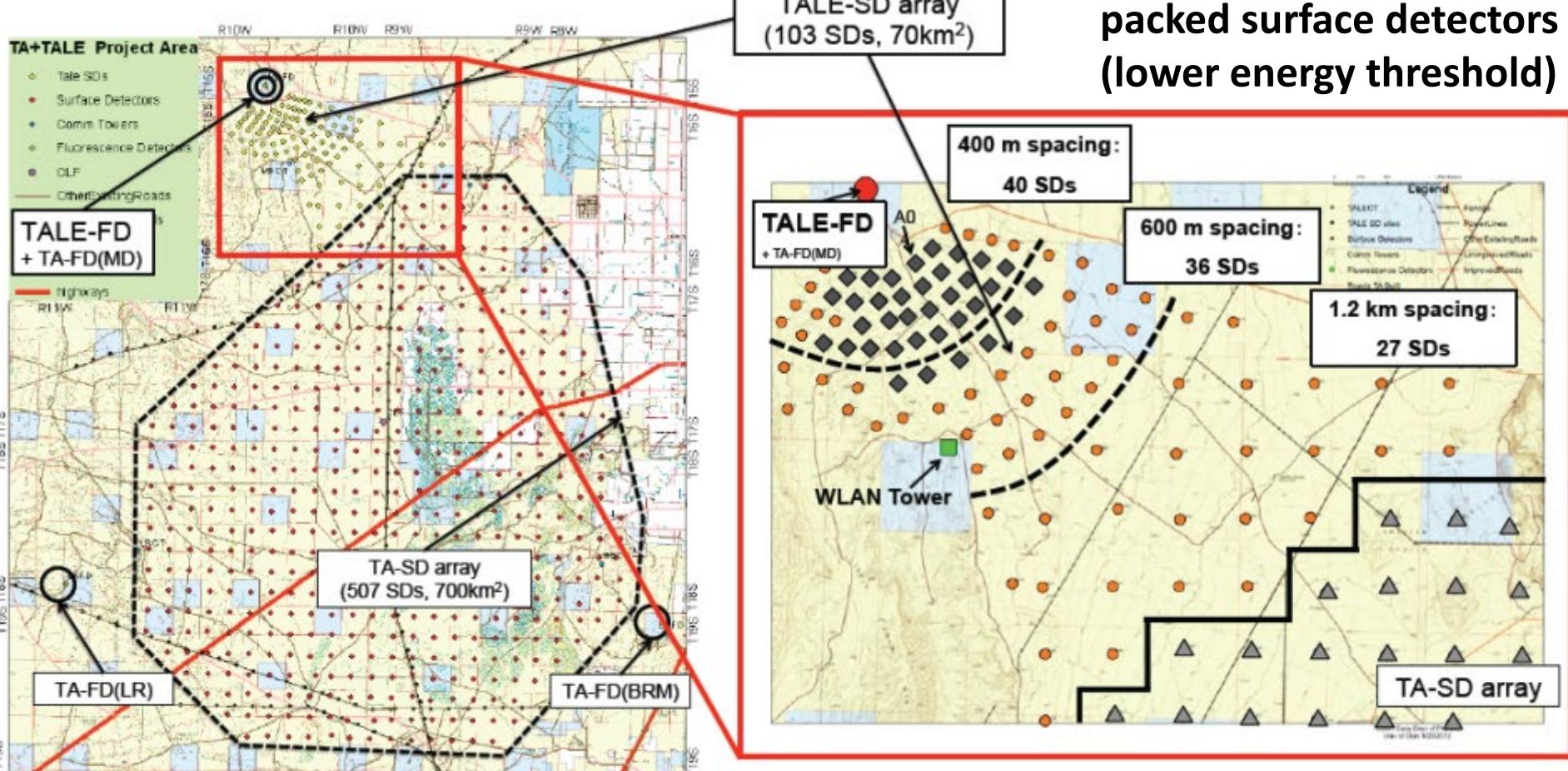
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TA Low Energy Extension (TALE)



10 new telescopes to look higher in the sky ($31\text{--}59^\circ$) to see shower development to much lower energies

**Infill surface detector
array of more densely
packed surface detectors
(lower energy threshold)**



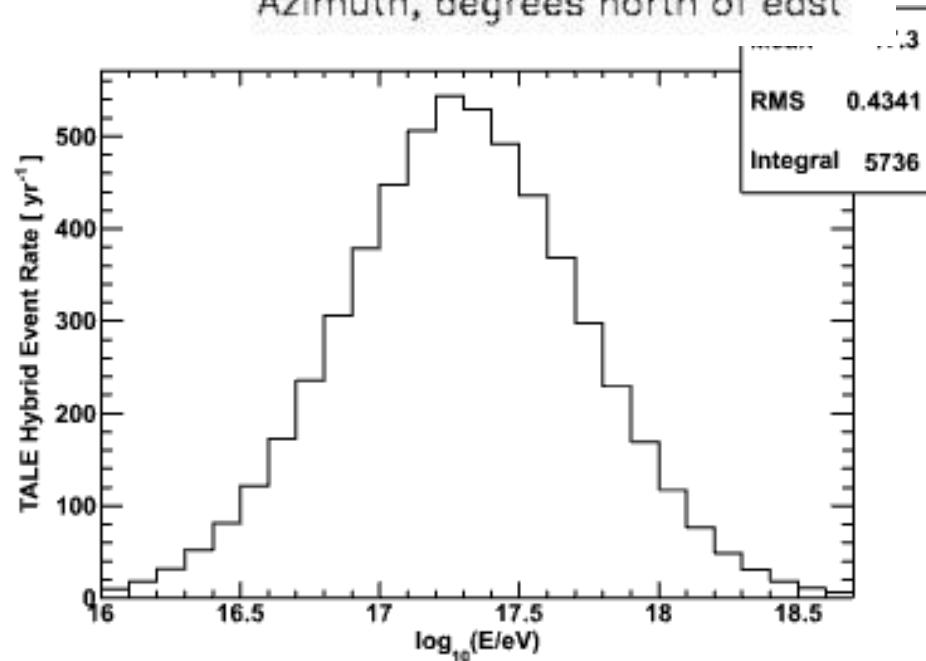
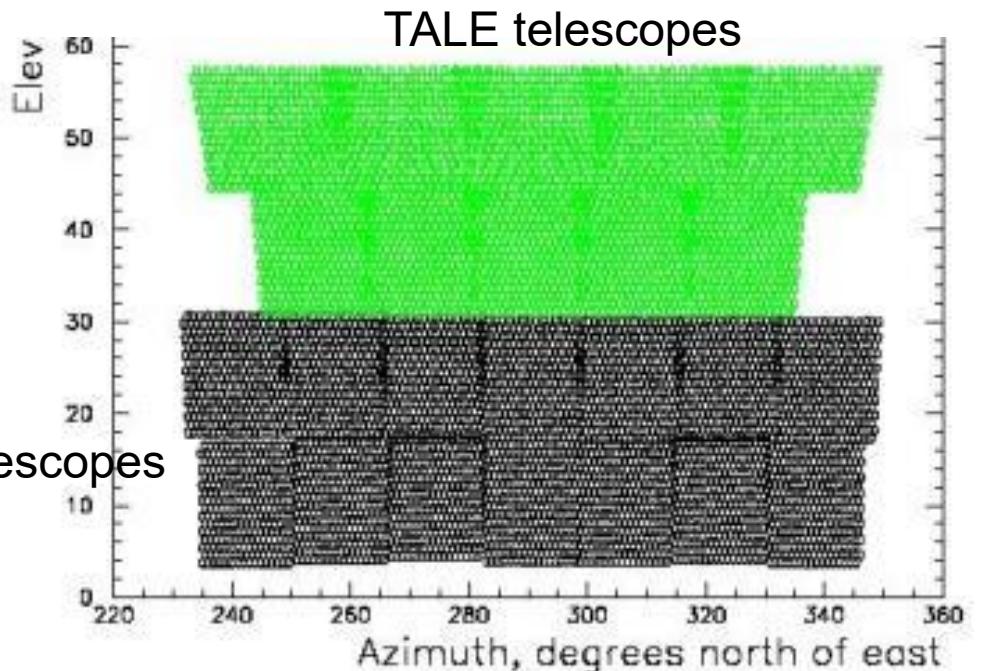


- All 10 TALE FD telescopes installed in 2013.
 - Shake-down 2013-2014
 - **Stable operation since fall 2014**
- **103 TALE SD counters deployed as of fall 2017**



TALE Fluorescence Detector

- 10 high-elevation telescopes at the Middle Drum site, looking from 31° - 59° in elevation.
- Originally designed for monocular and hybrid observations down to $\sim 10^{16.5}$ eV.



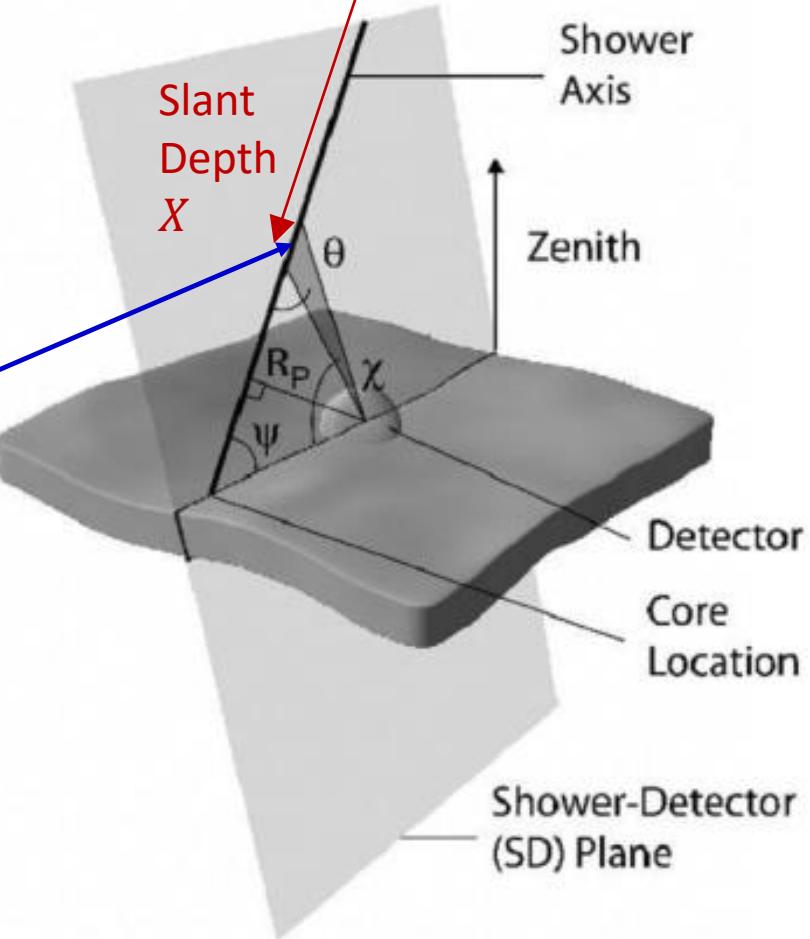
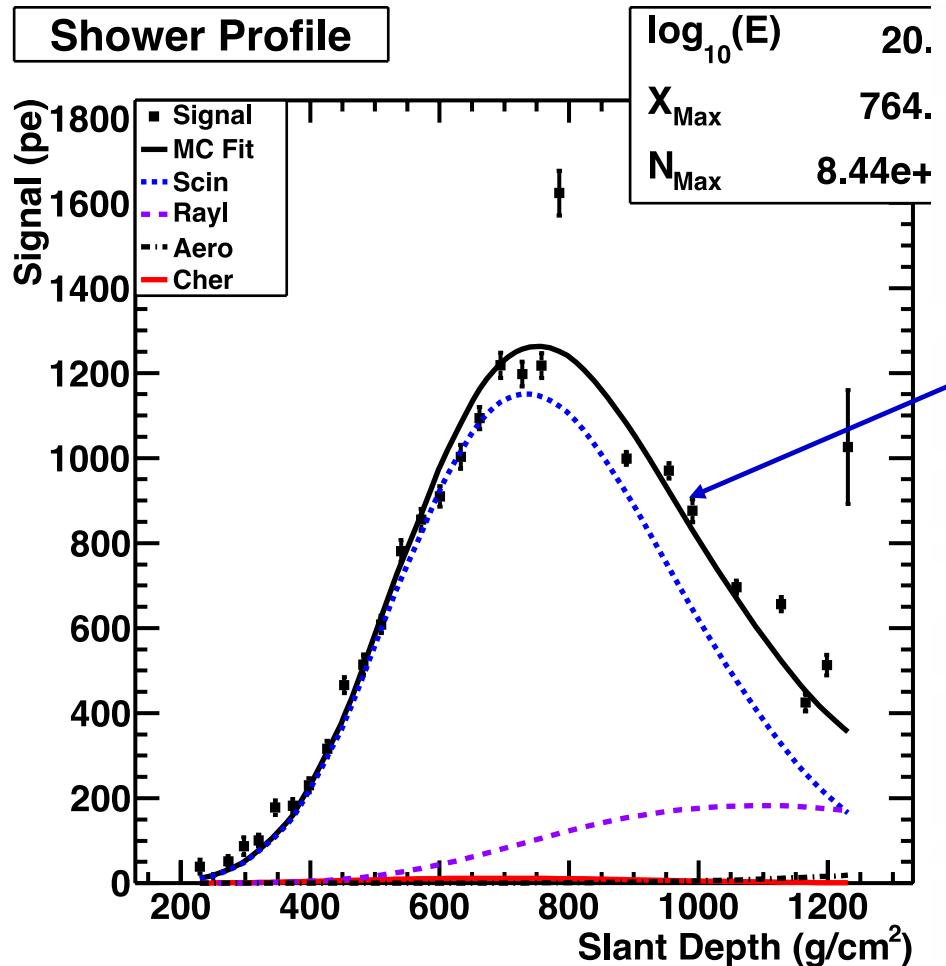
Expected TALE hybrid events per year

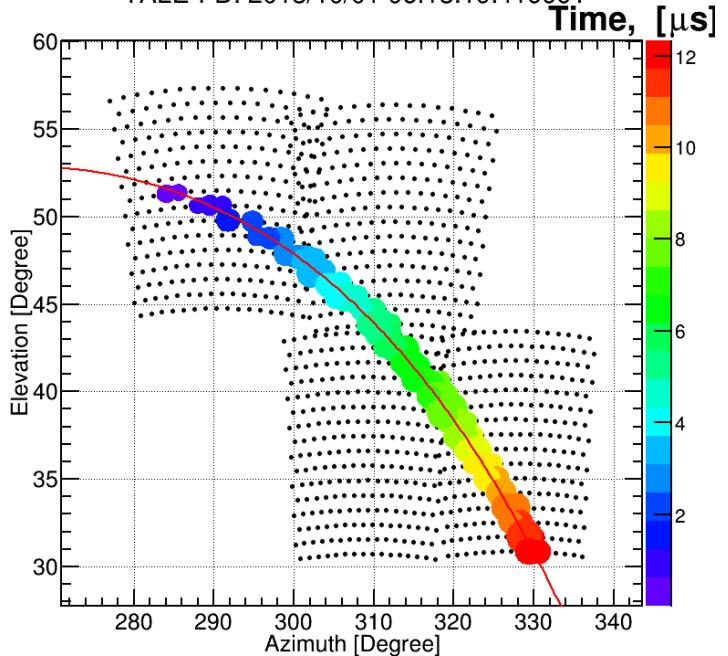




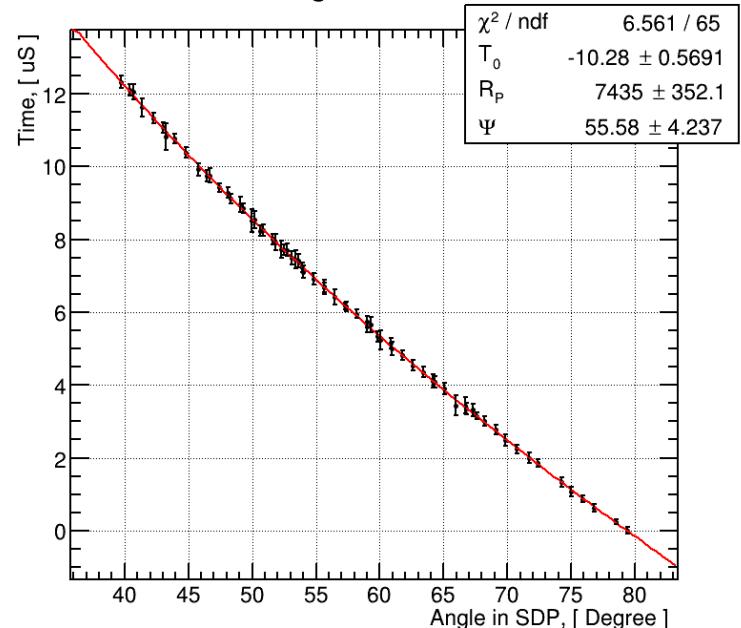
Combined Tim-Profile Fit

From Top of Atmosphere





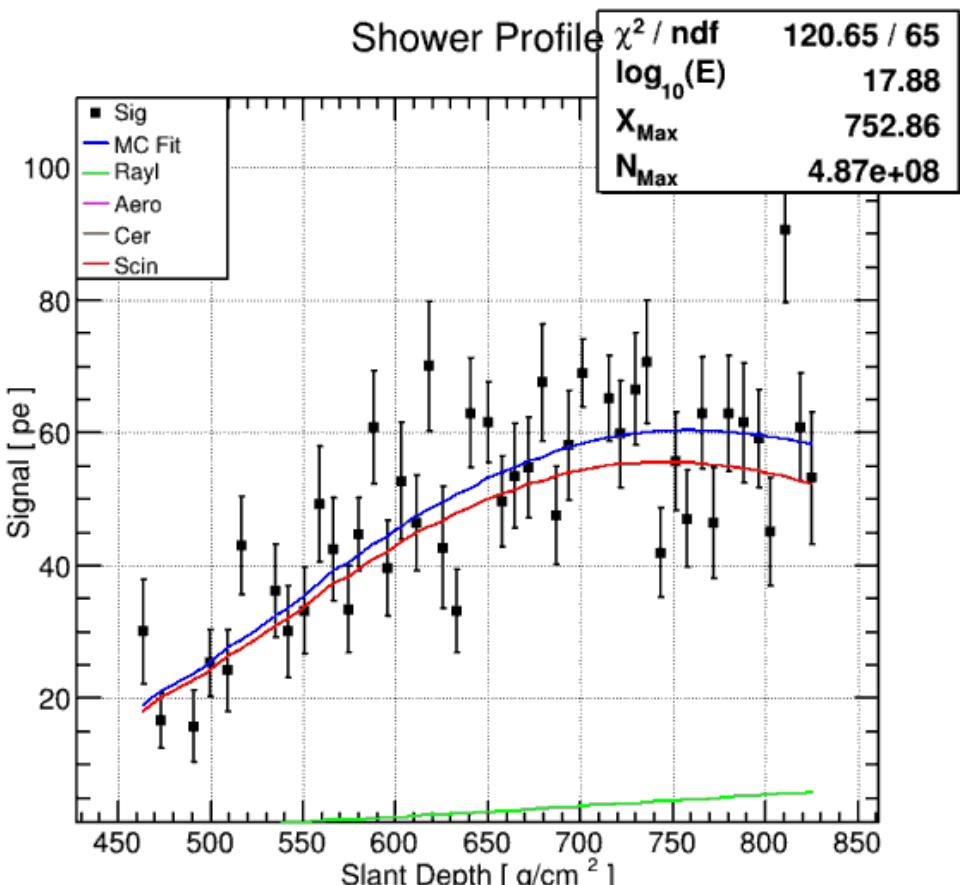
Time vs Angle, TALE-FD-Mono

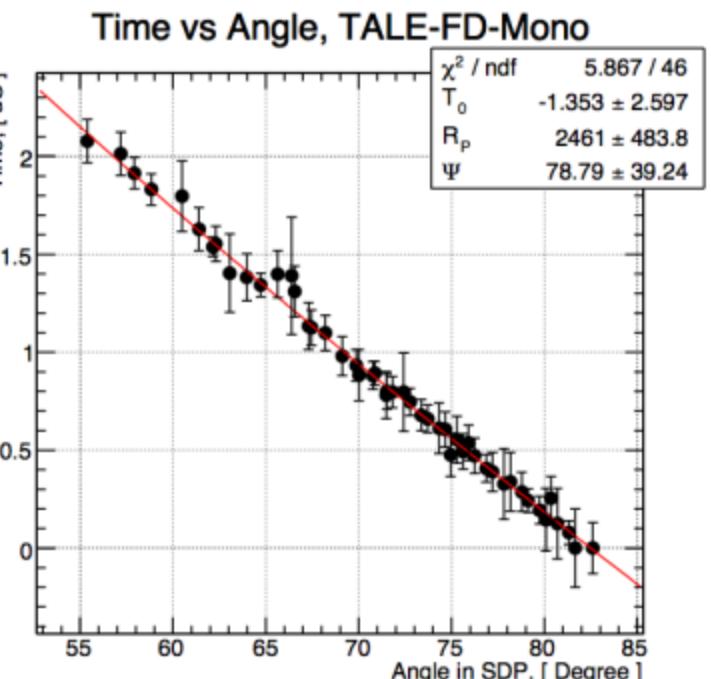
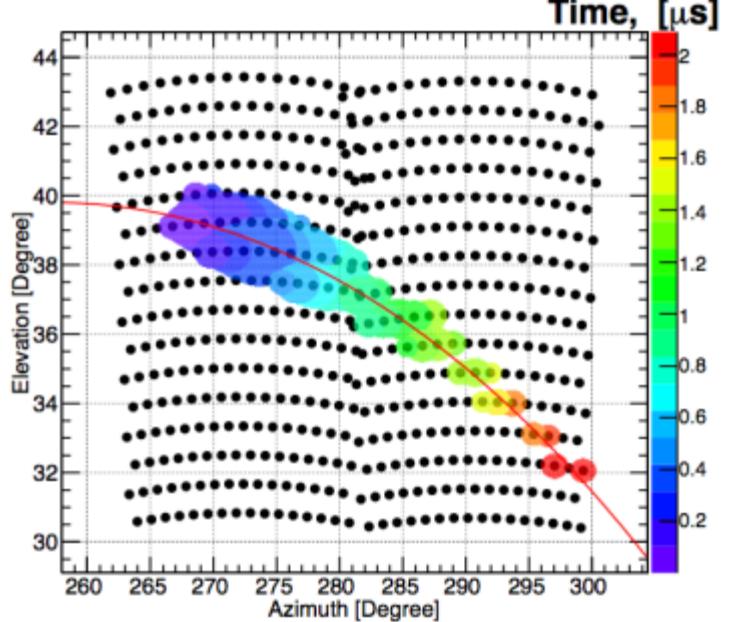


TALE FD Event



For TALE FD reconstruction: we combined the time and profile fit: simultaneous **Profile Constrained Geometry Fit (PCFG)** originally developed for HiRes monocular analysis

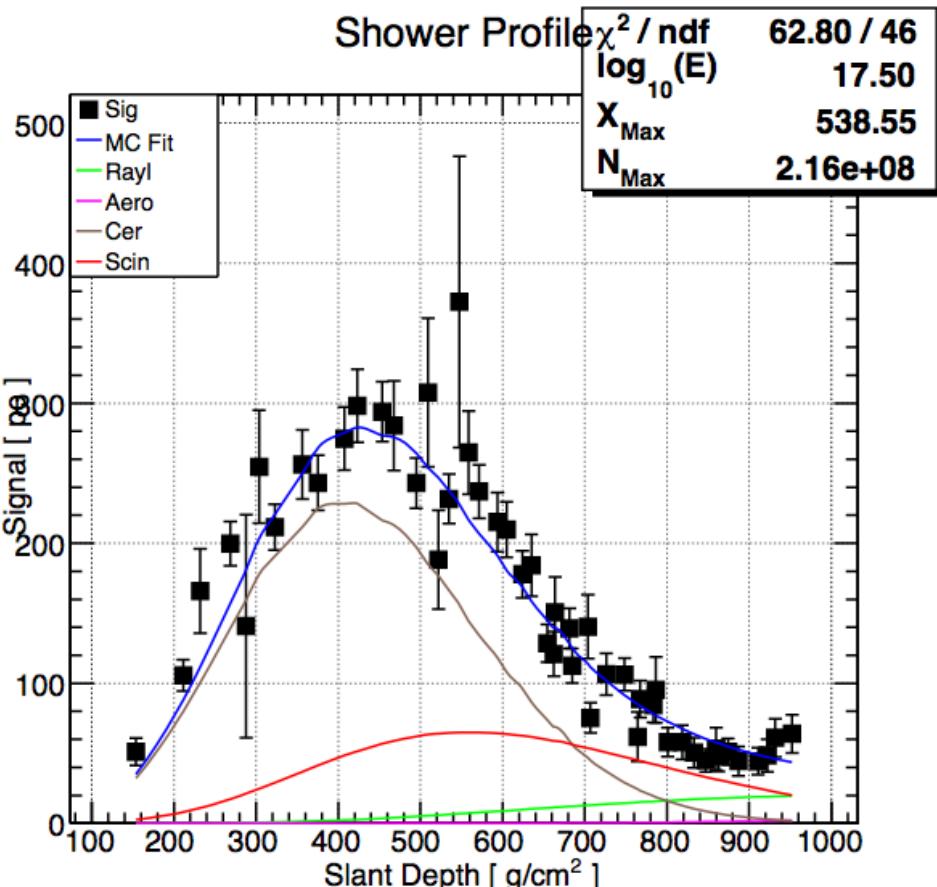




TALE Cherenkov Event



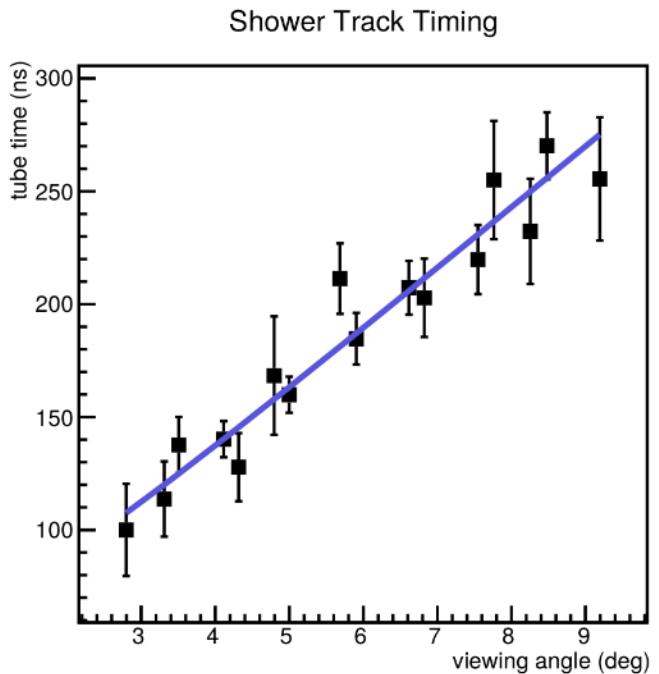
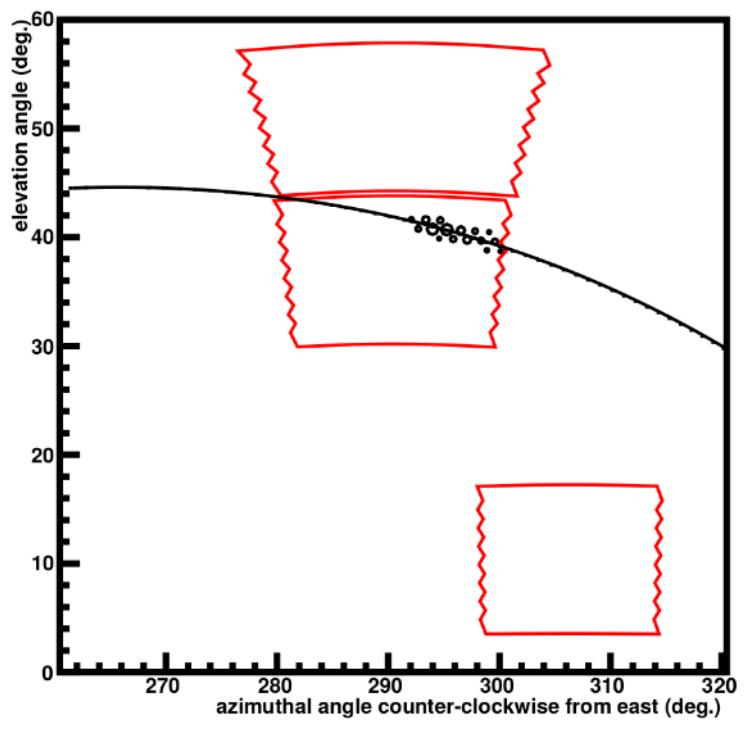
PCGF turns out to work very well on **Cherenkov light dominated events**



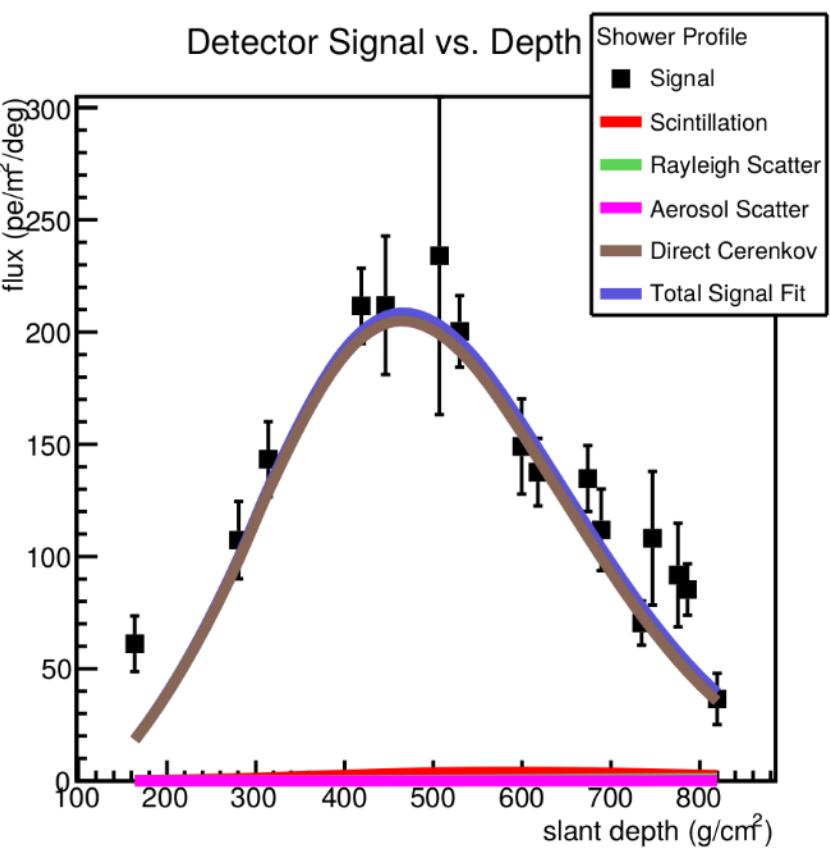


TALE Cherenkov Event

Cherenkov light dominated events
allowed TALE to reach more than another
decade lower in energy than designed:
Down to $10^{15.3}$ eV



$10^{15.6}$ eV
Event





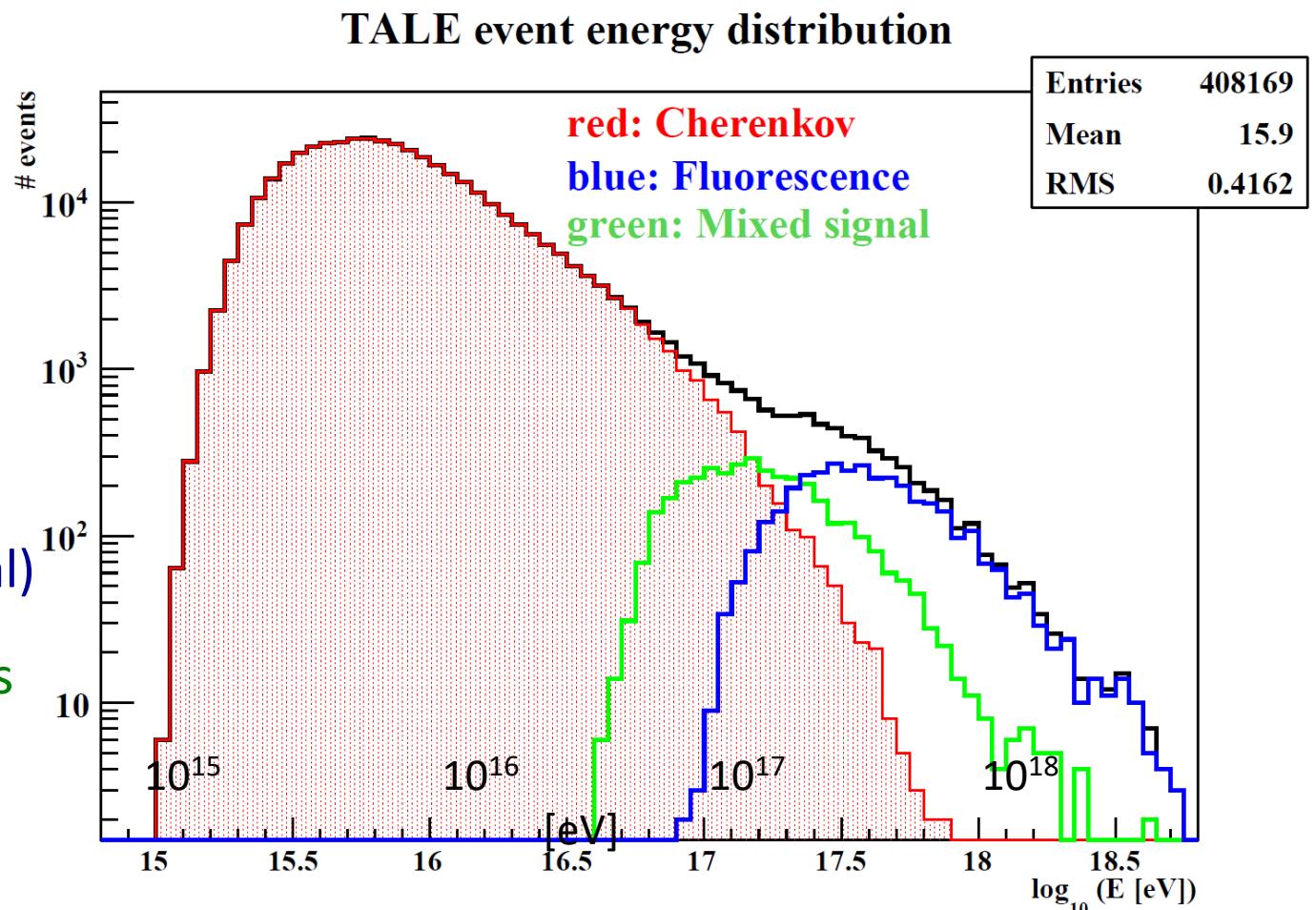
Events

Between June 2014 and March 2016 TALE collected
220 173 events between $10^{15.5}$ and $10^{16.0}$ eV
98 677 events between $10^{16.0}$ and $10^{16.5}$ eV
23 285 events between $10^{16.5}$ and $10^{17.0}$ eV

Cherenkov
dominated events
(C'kov > 75% of
total)

Fluorescence
dominated events
(Scin. > 75% of total)

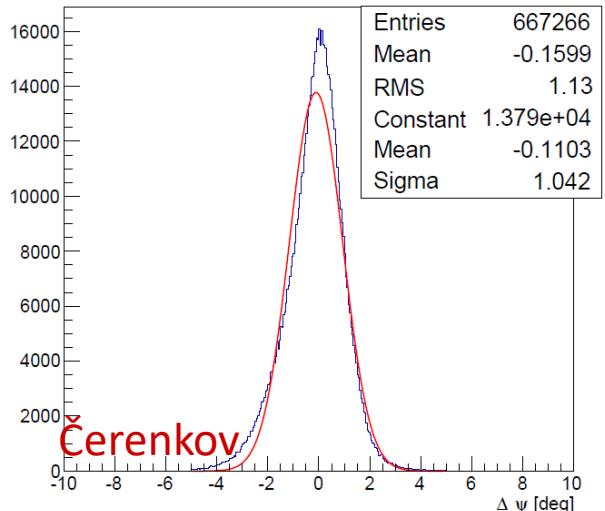
Mixed signal events



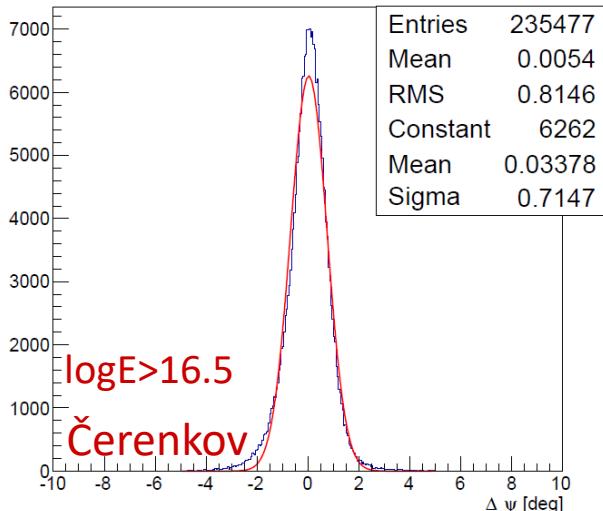
Verification of monocular Ψ resolution



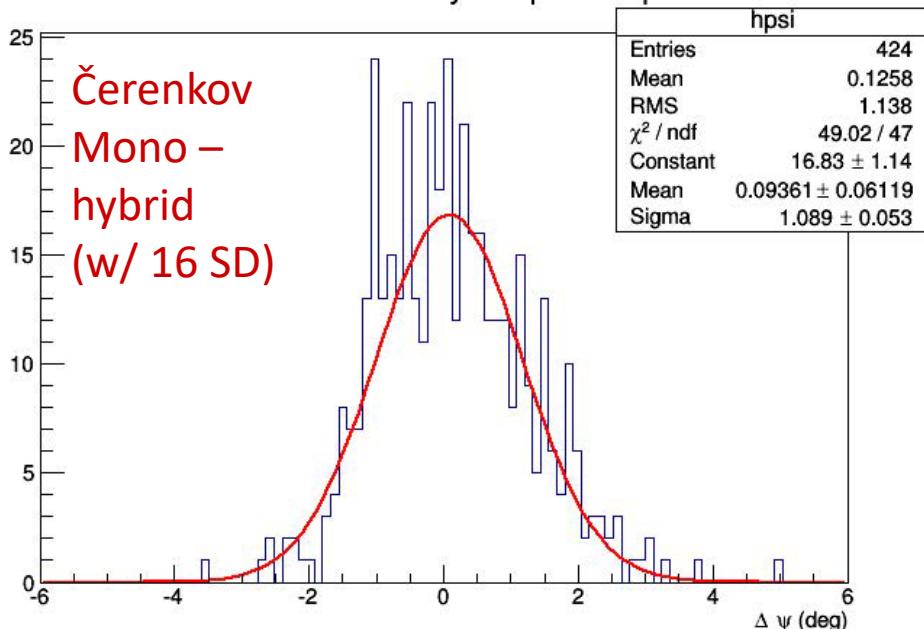
Shower ψ angle: $\Delta \psi$ (deg)



Shower ψ angle: $\Delta \psi$ (deg)



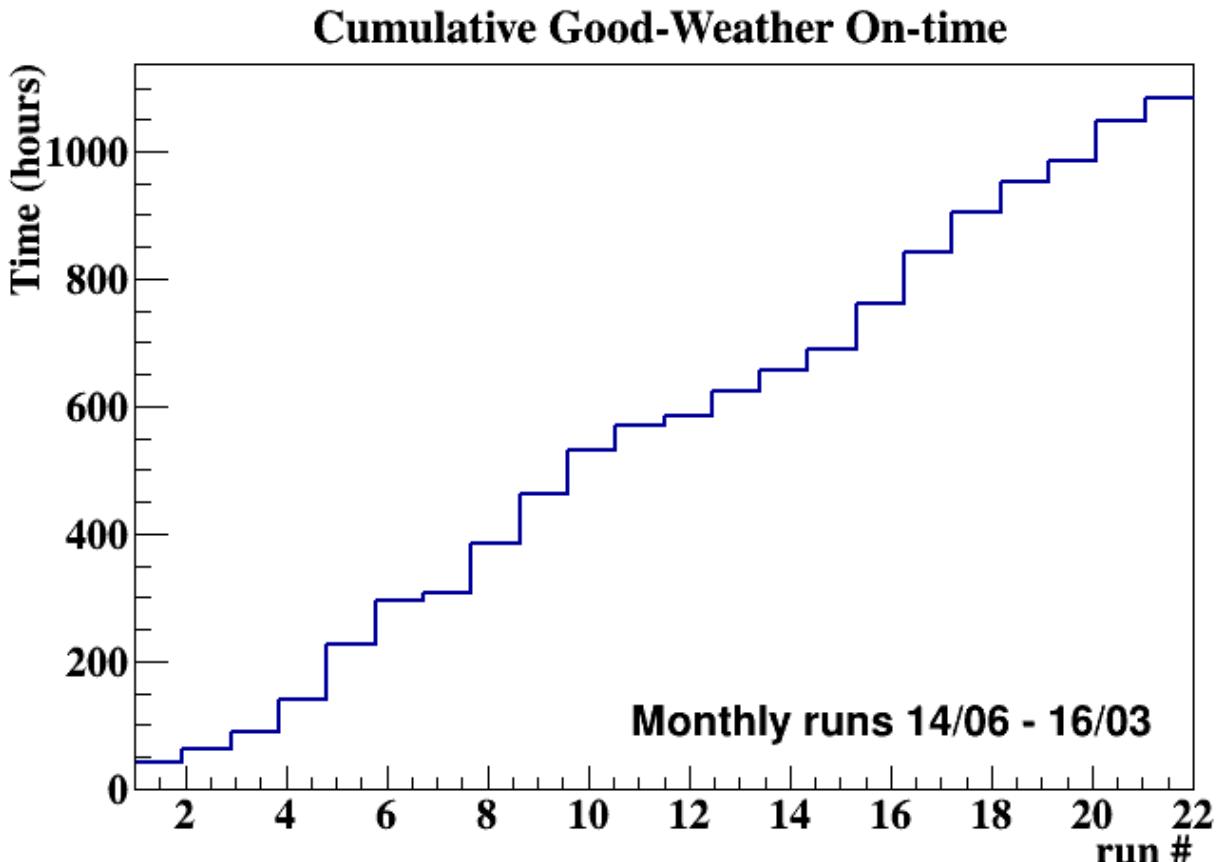
TALE C'kov mono - hybrid psi comparison





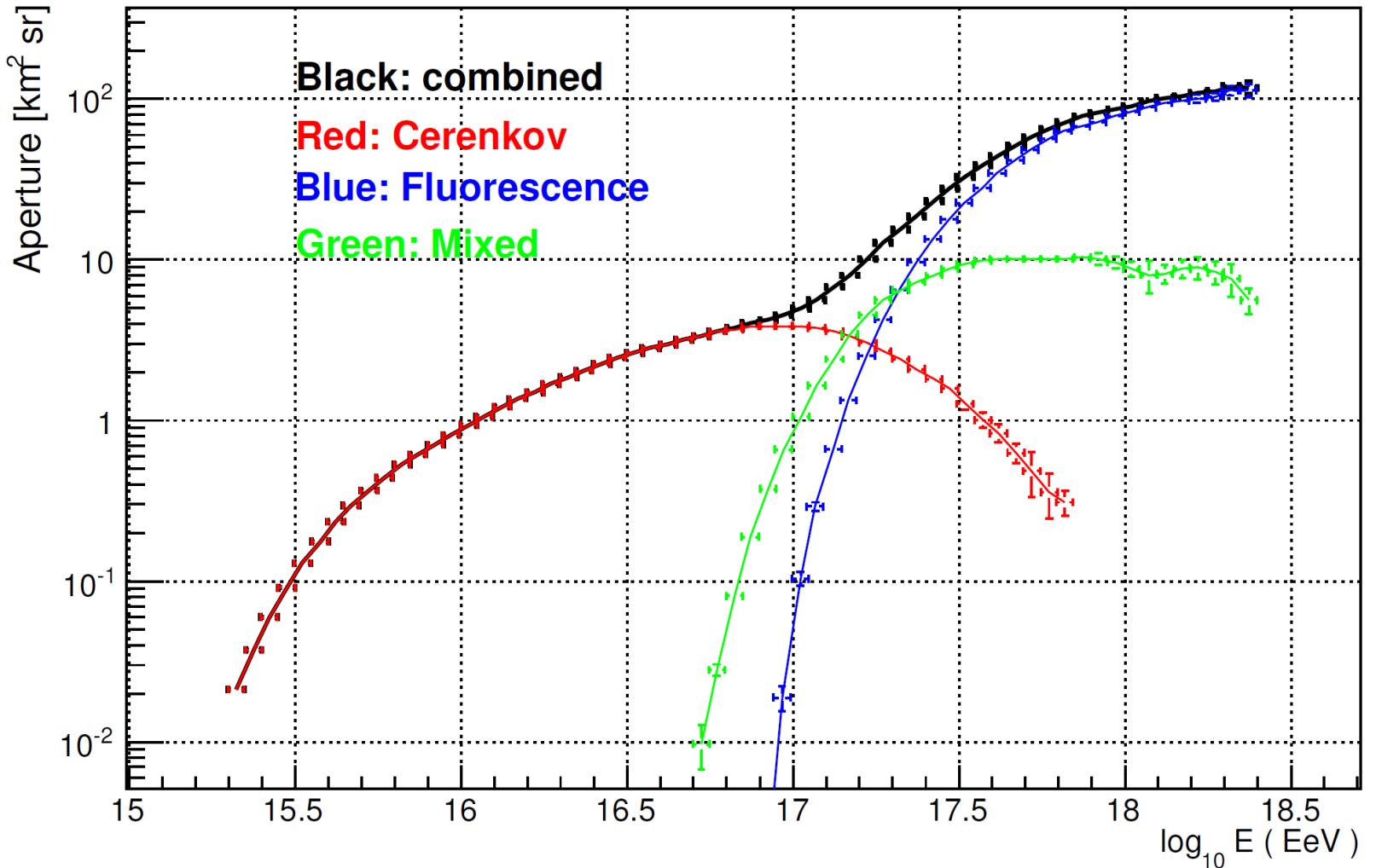
Exposure

- TALE FD data collected from 06/20/2014 to 03/31/2016 (**22 months**).
- Only **good weather** data:
- **Total on-time 1080.0 hours.**





Aperture

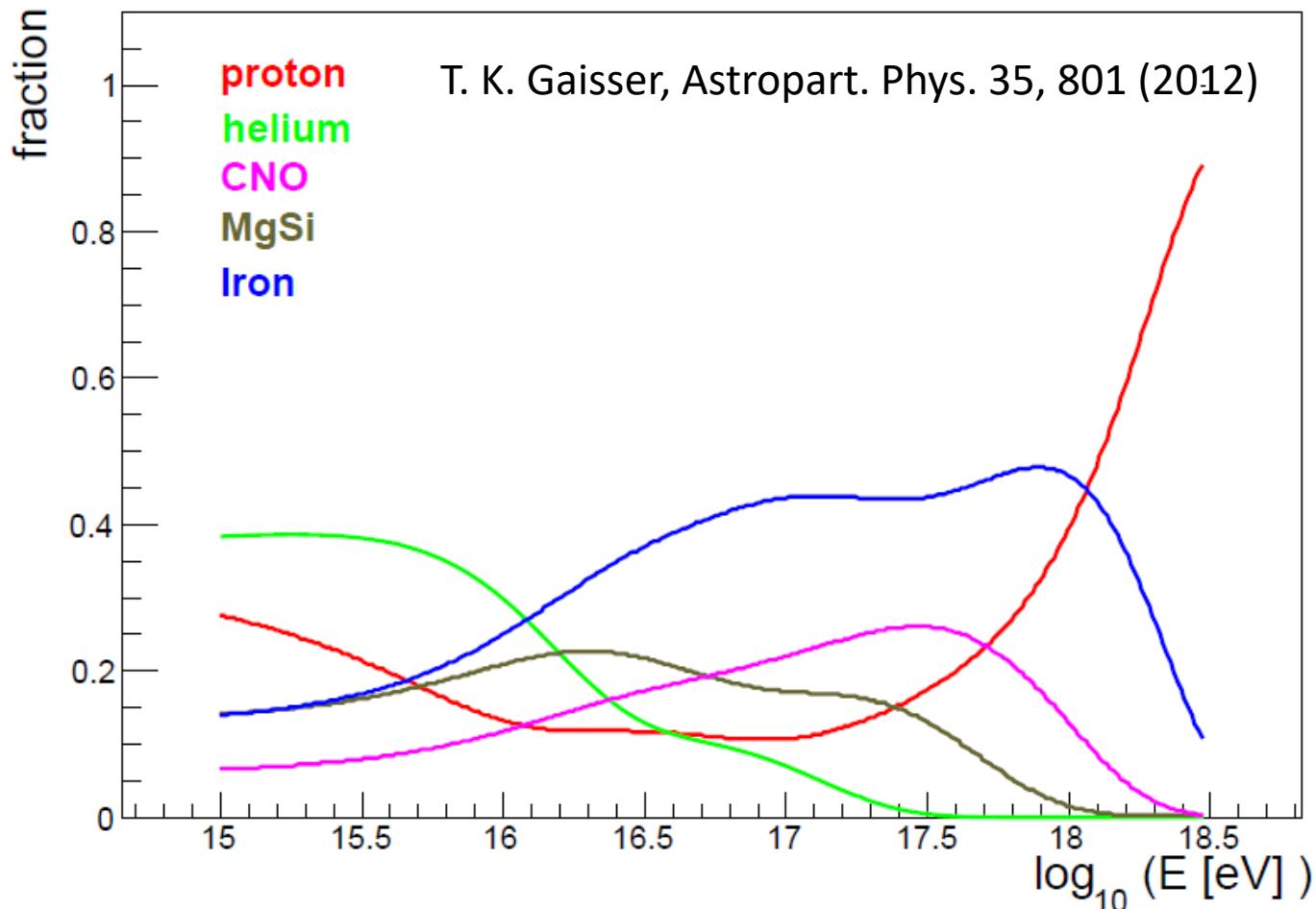




Composition?

Initial Assumption

primary fractions (H4a CR composition model)

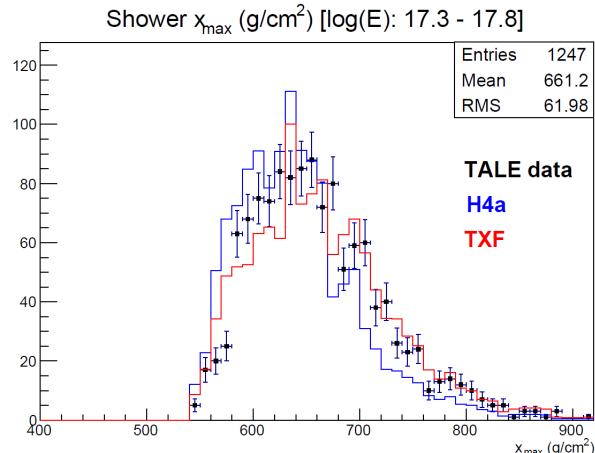
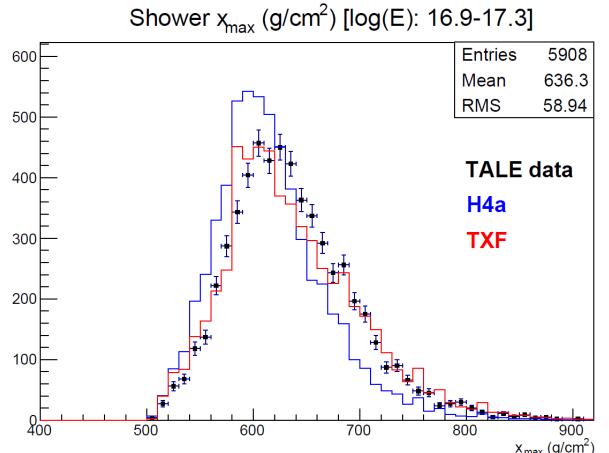
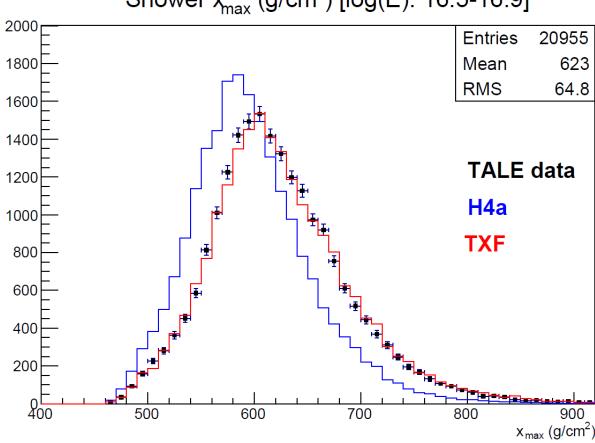
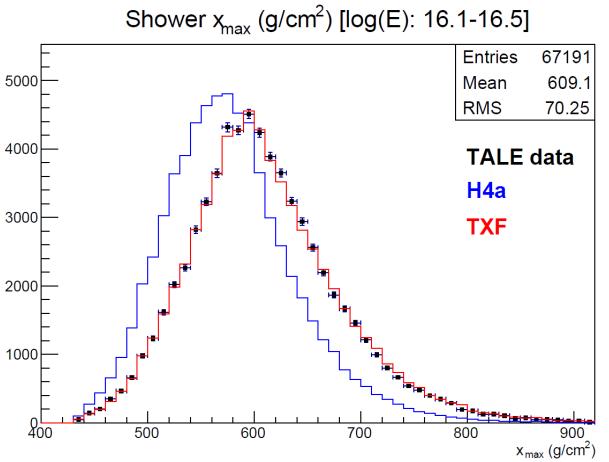
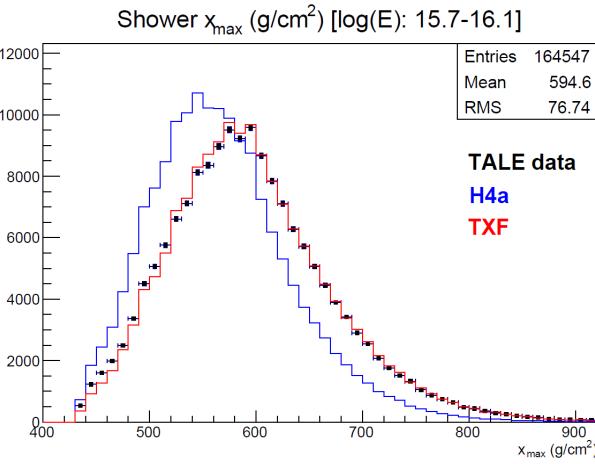
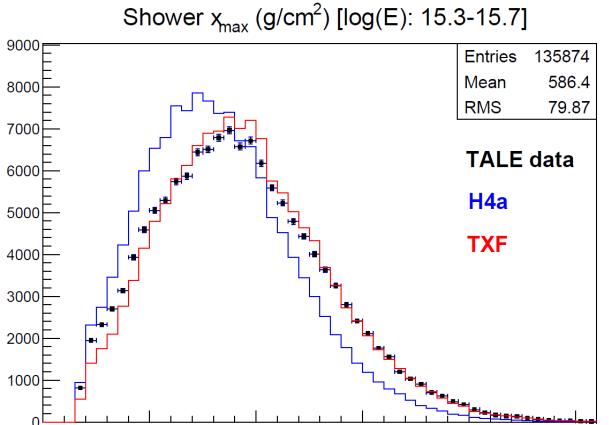




Xmax MC/Data Comparison

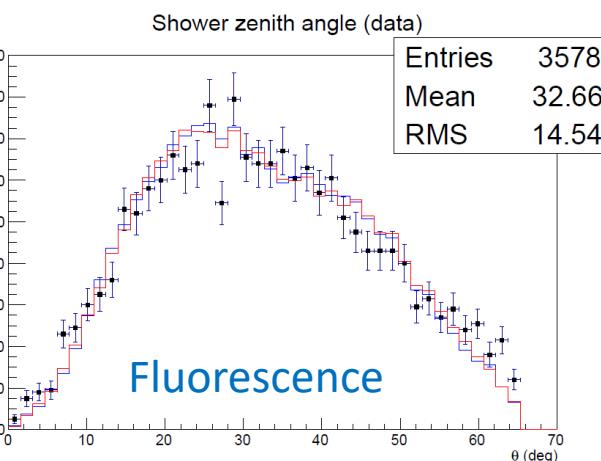
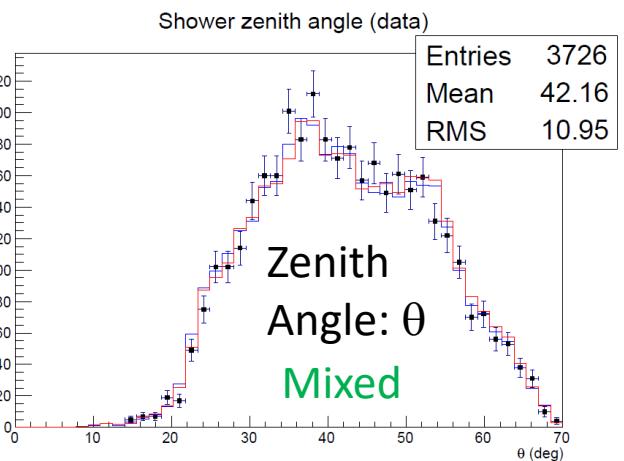
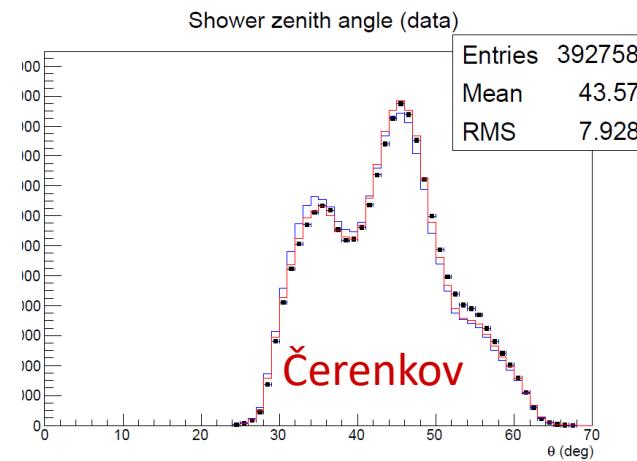
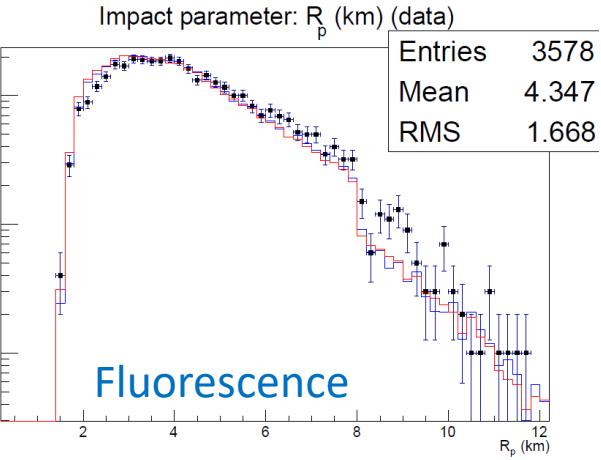
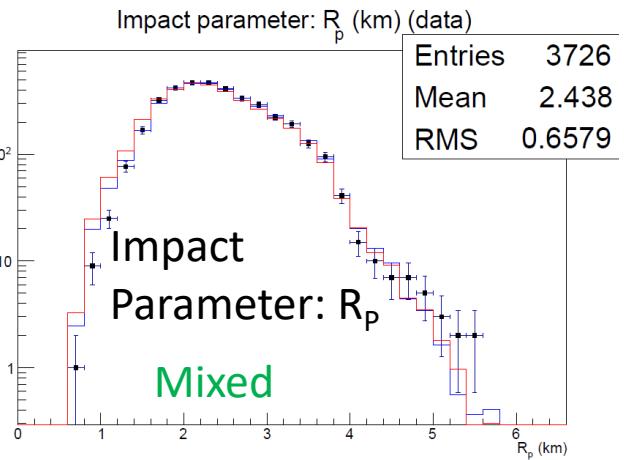
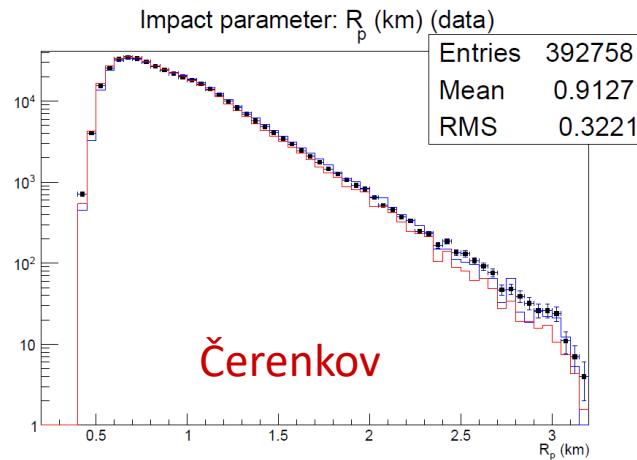
Originally assumed **H4A** Composition (interpreted using QGS_jet II.3) gives an Xmax distribution appeared to be too high in the sky.

Adjusted Xmax dist. (a 4-component fit to QGS-Jet II.3) to better match TALE FD data: call this Tale Xmax Fit (**TXF**)





MC/Data Comparison



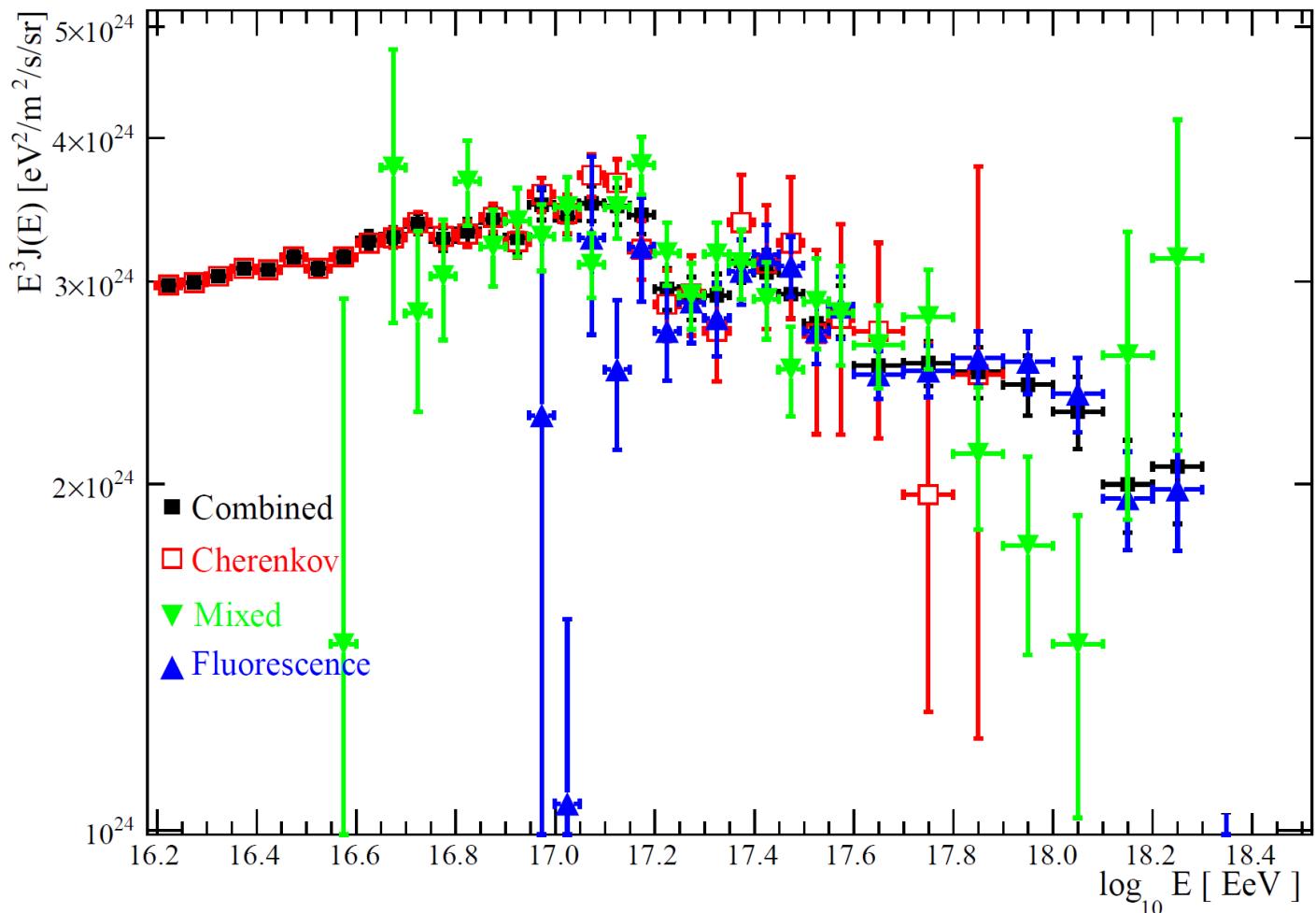
— H4A

— TXF

Validation of Aperture Calculation



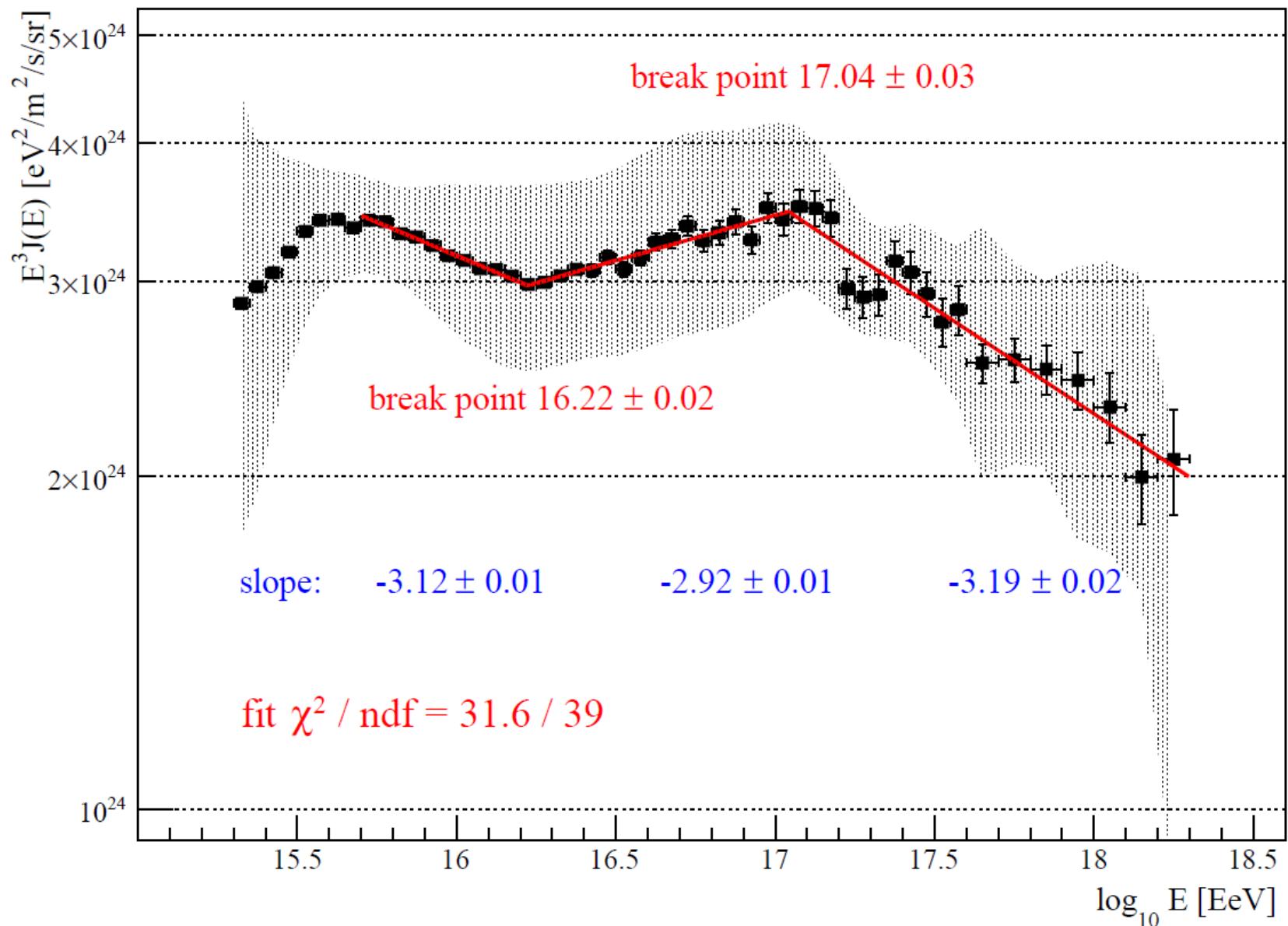
Consistency



Spectra obtained from the three classifications of events appear to be consistent

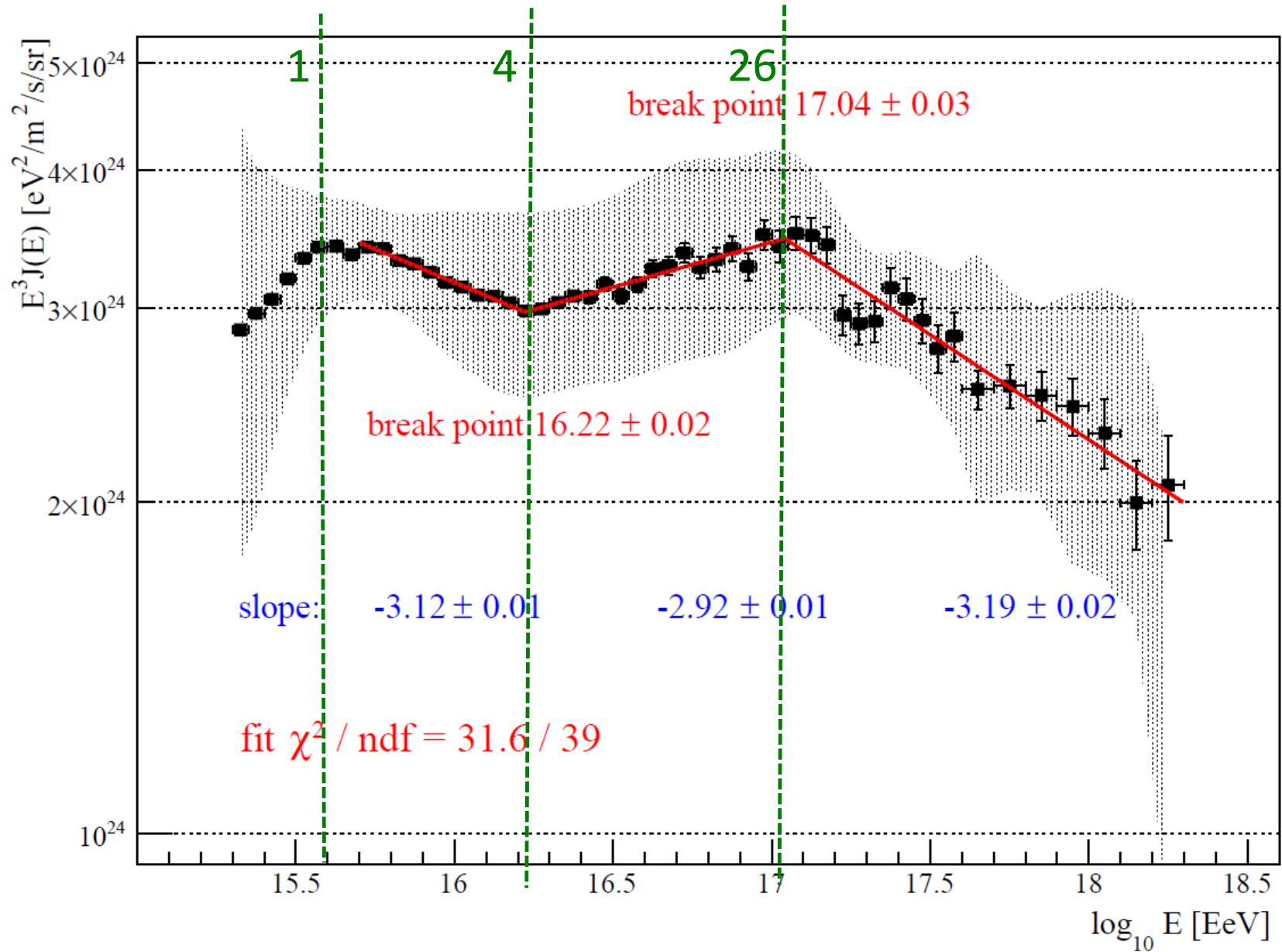


TALE Spectrum





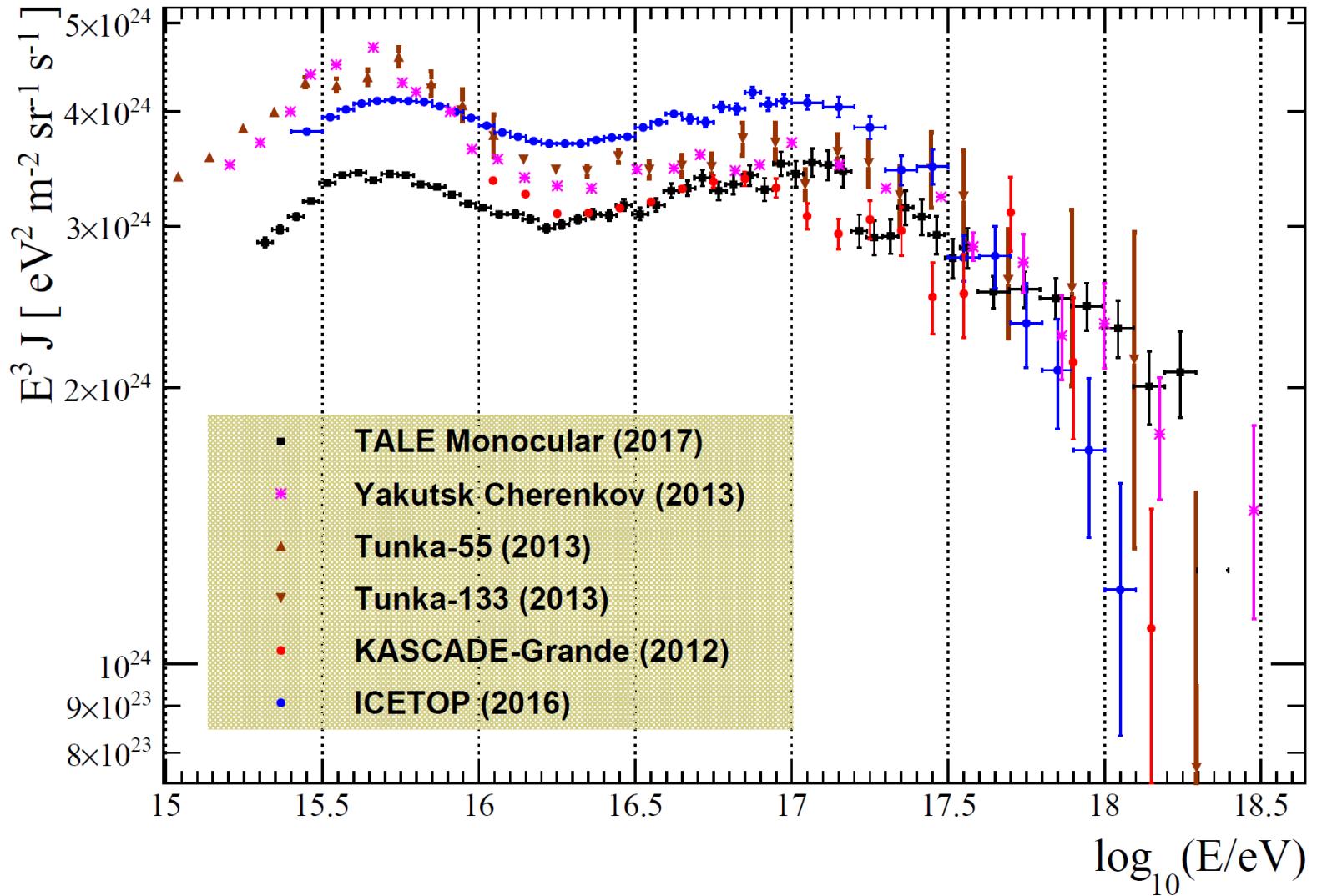
TALE Spectrum



Comparison to other experiments



TALE Spectrum compared to some recent Measurements





Summary

- We have measured the cosmic ray energy spectrum in the range $10^{15.3}$ - $10^{18.3}$ eV
 - Published 2018 September 24.
[The Astrophysical Journal, Volume 865, Number 1](#)
 - Three spectral features are seen: Knee, “dip”, “second knee” at energies at $10^{15.6}$ eVm $10^{16.22}$ eV and $10^{17.04}$ eV
 - The energies of the three features are approximately in the ratio of 1:4:26 (??? proton: beryllium: iron ???)
- Composition results in the near future
- TALE surface detector now operational
 - FD to SD trigger now running



End



Reserve Slides

Systematic Errors



Table 5. Estimates of systematic uncertainties in the TALE FD energy scale and spectrum measurement. This uncertainty is approximately constant as a function of energy [*Explanation of change: Added entry: Cherenkov model*]

Energy	Source	value	contribution to spectrum
$< 10^{17}$ eV	photonic scale	10%	20%
$< 10^{17}$ eV	missing energy	10%	20%
$< 10^{17}$ eV	atmosphere	0	0
$< 10^{17}$ eV	Cherenkov model	5%	10%
$< 10^{17}$ eV	fluorescence yield	0	0
$< 10^{17}$ eV	composition (X_{max})	3%	6%
10^{18} eV	photonic scale	10%	20%
10^{18} eV	missing energy	5%	10%
10^{18} eV	atmosphere	2%	4%
10^{18} eV	Cherenkov model	0	0
10^{18} eV	fluorescence yield	10%	20%
10^{18} eV	composition (X_{max})	3%	6%
$< 10^{17}$ eV	total	15%	31%
10^{18} eV	total	15%	31%



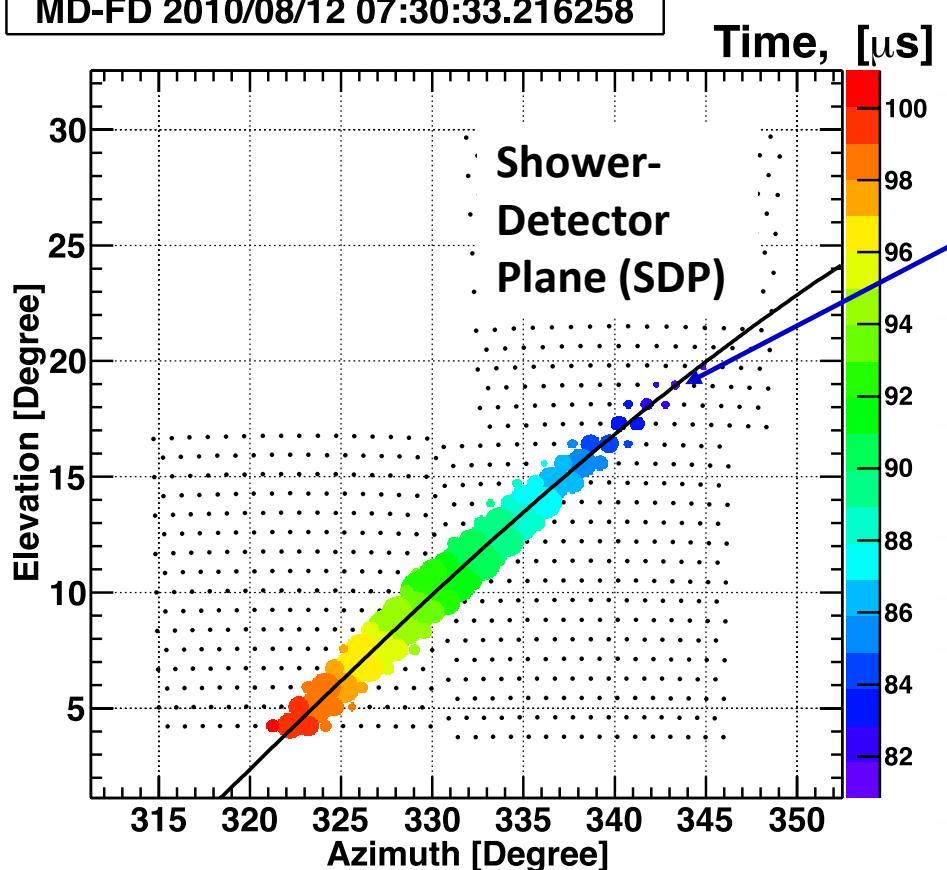
Outline

- Introduction to TALE
- TALE Events and Reconstruction
- Detector Resolutions
- Aperture and Data/MC Comparisons
- Spectrum
- Interpretation of Spectrum
- Continuing work on TALE FD and SD

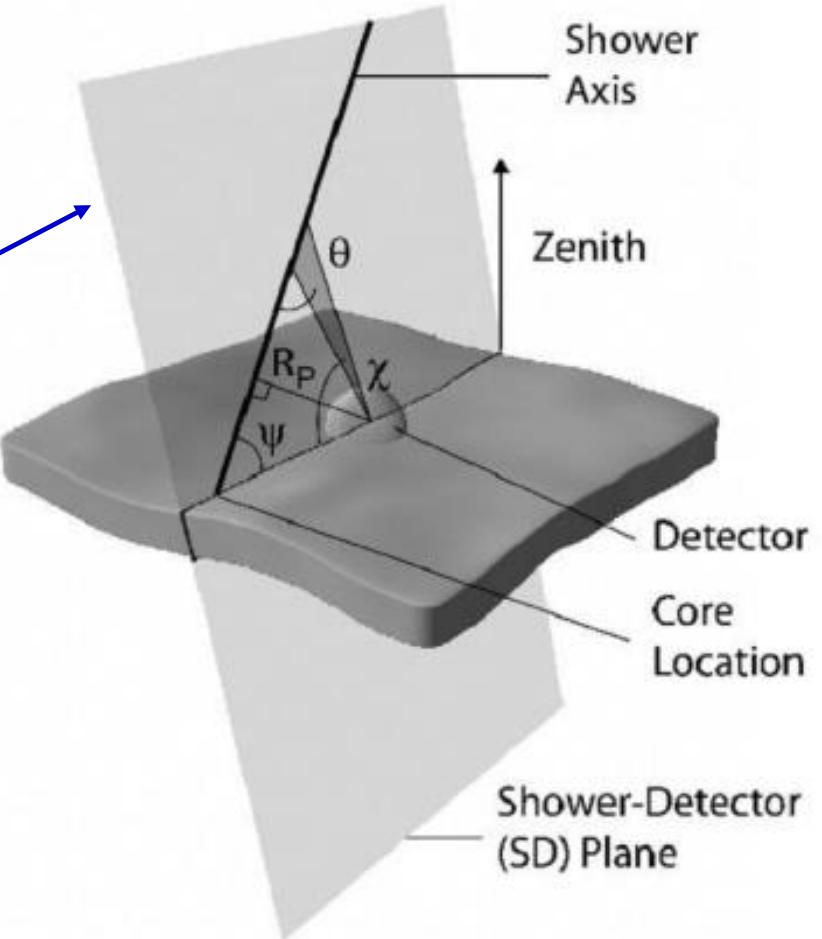
“Reconstruction” of an air shower



MD-FD 2010/08/12 07:30:33.216258



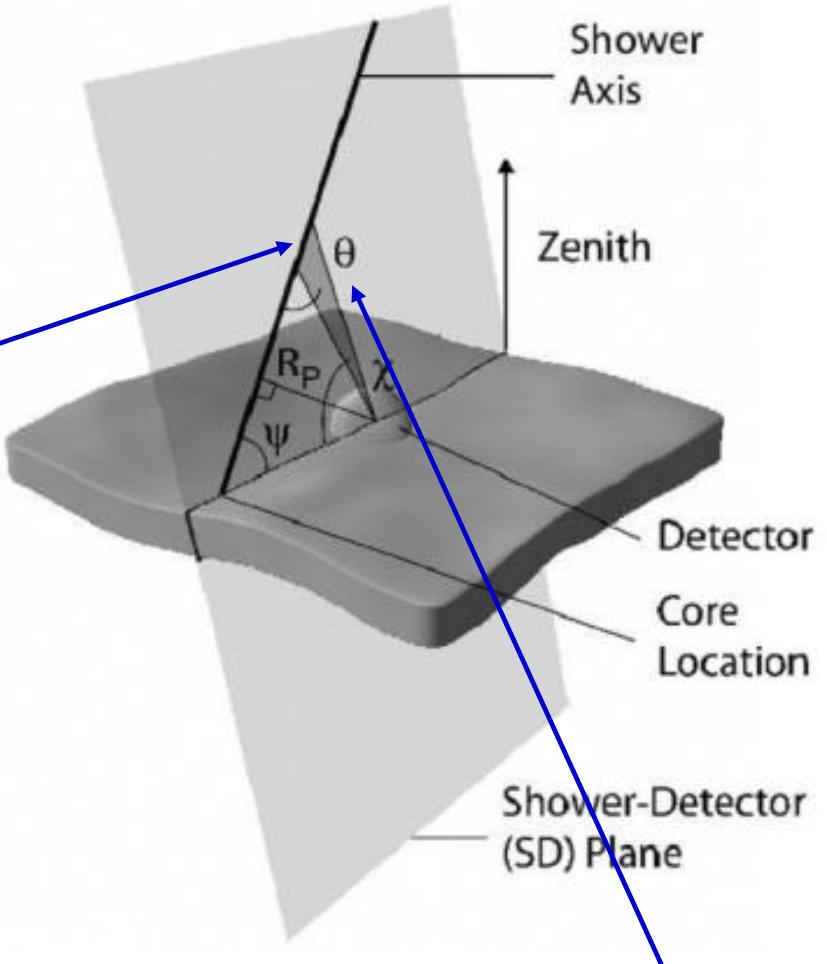
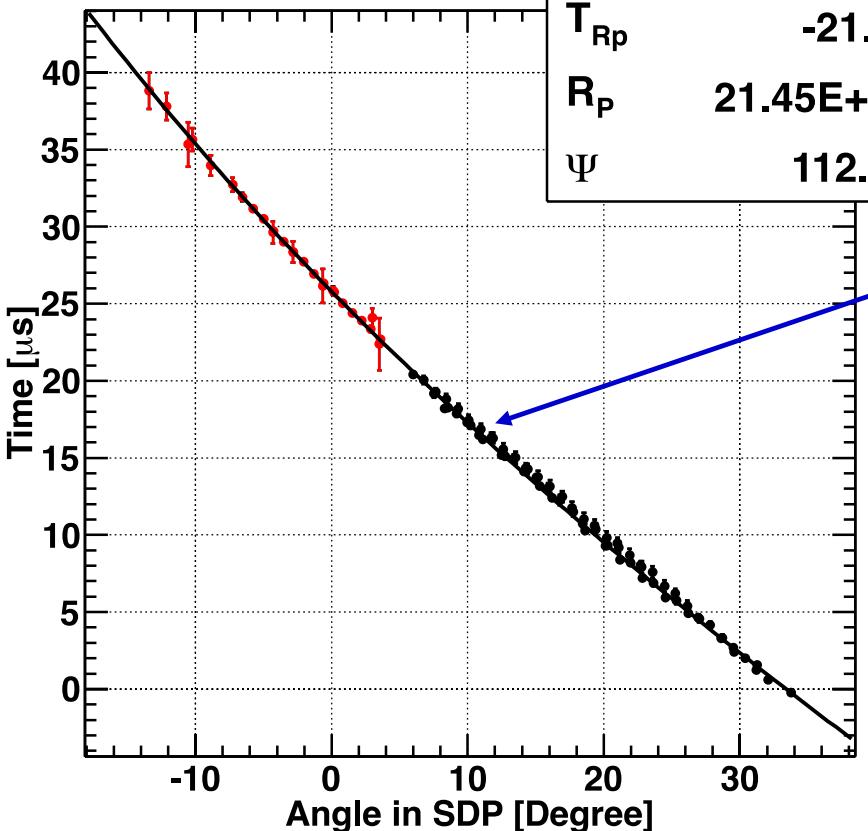
Event Display showing pattern of hit pixels



Direction of hit pixels fitted to a shower-detector plane (SDP)

Timing Fit

Time vs Angle (Hybrid)

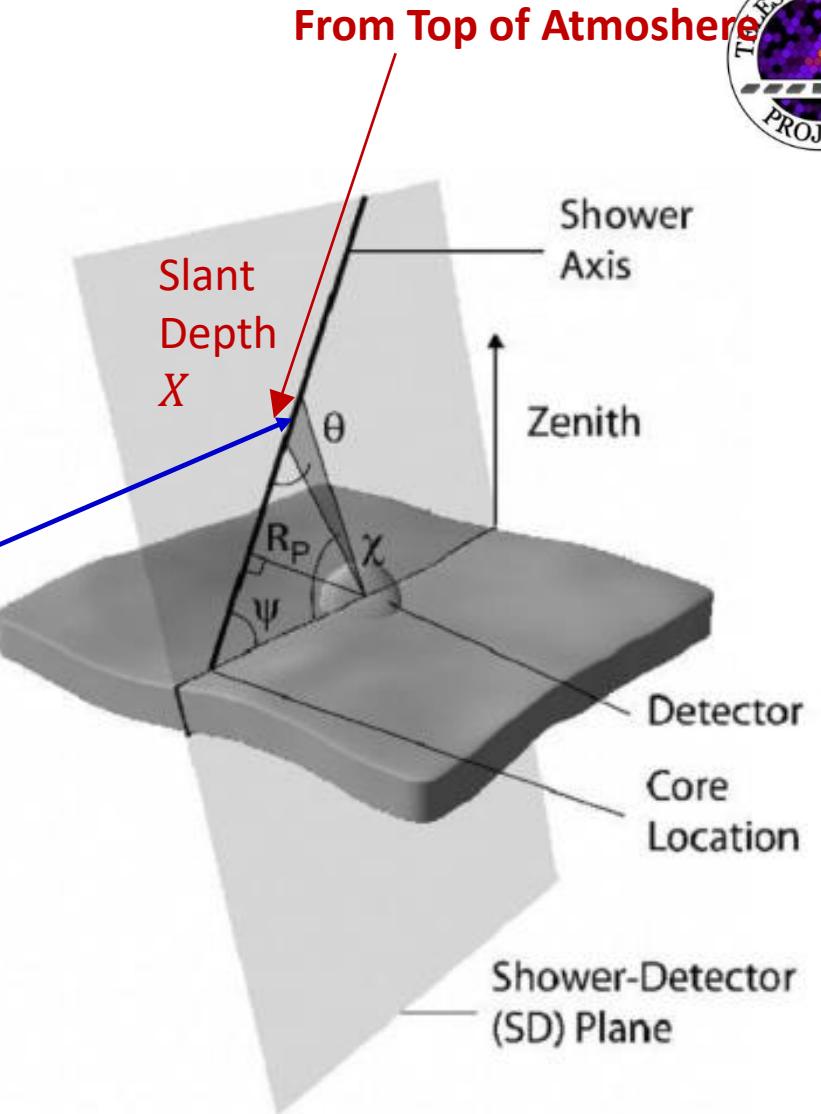
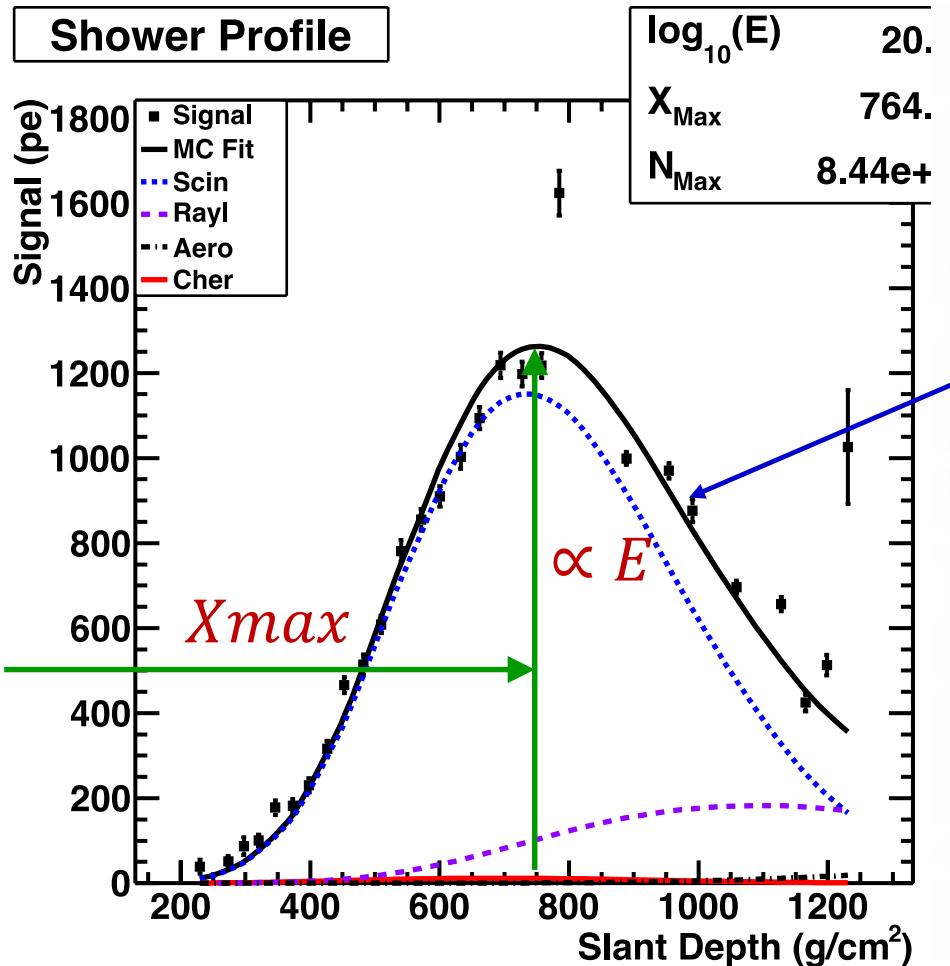


Arrival times of signal light in each pixel is fitted as a function of the SDP θ angles: **Gives direction of primary cosmic ray**

$$t_i = t_0 + \frac{R_P}{c} \tan \frac{\theta_i}{2}$$



Shower Profile Fit



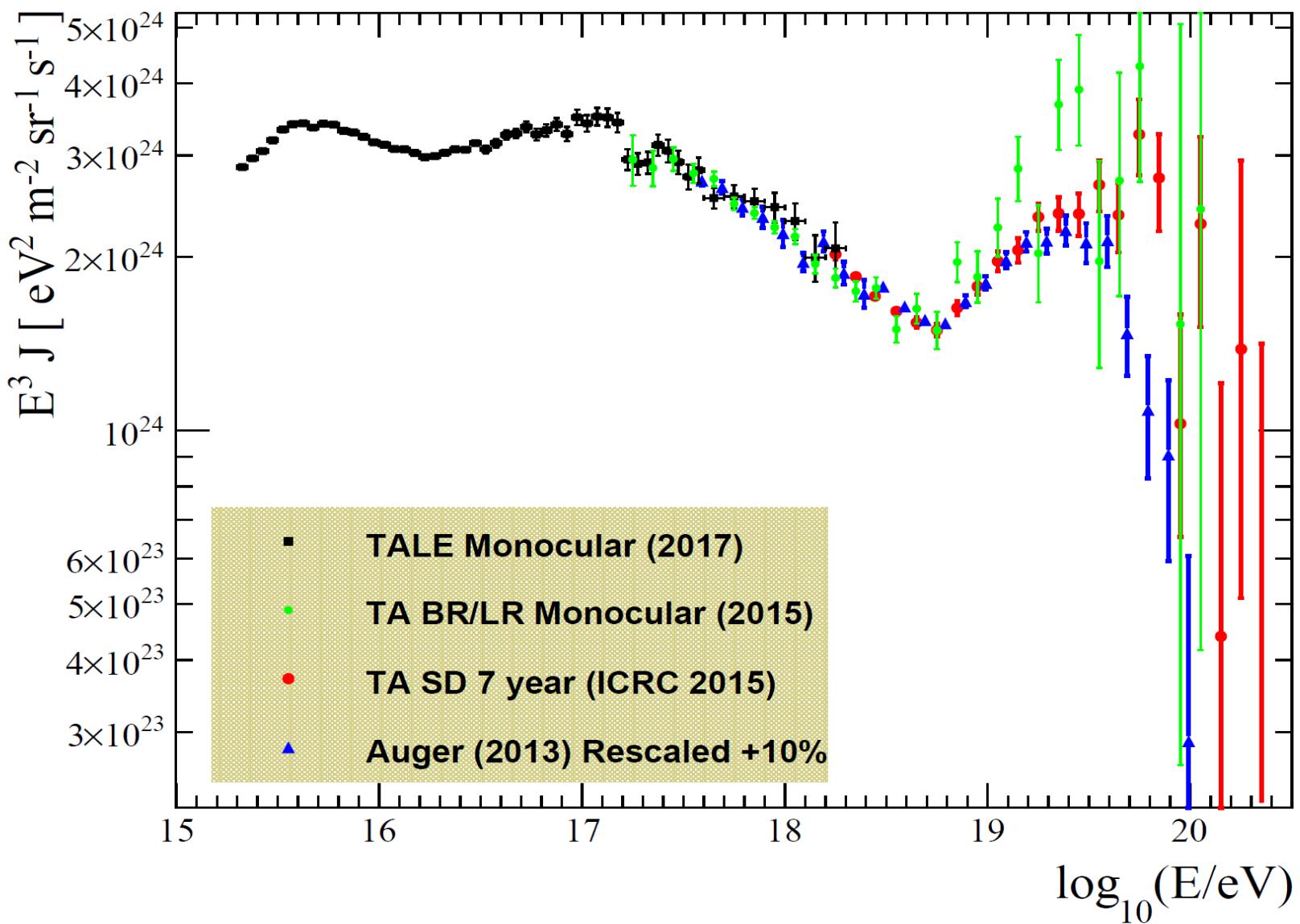
SDP θ angles converted to slant depth.

Light signal fitted to depth to give energy E and X_{max}
(depth of maximum)

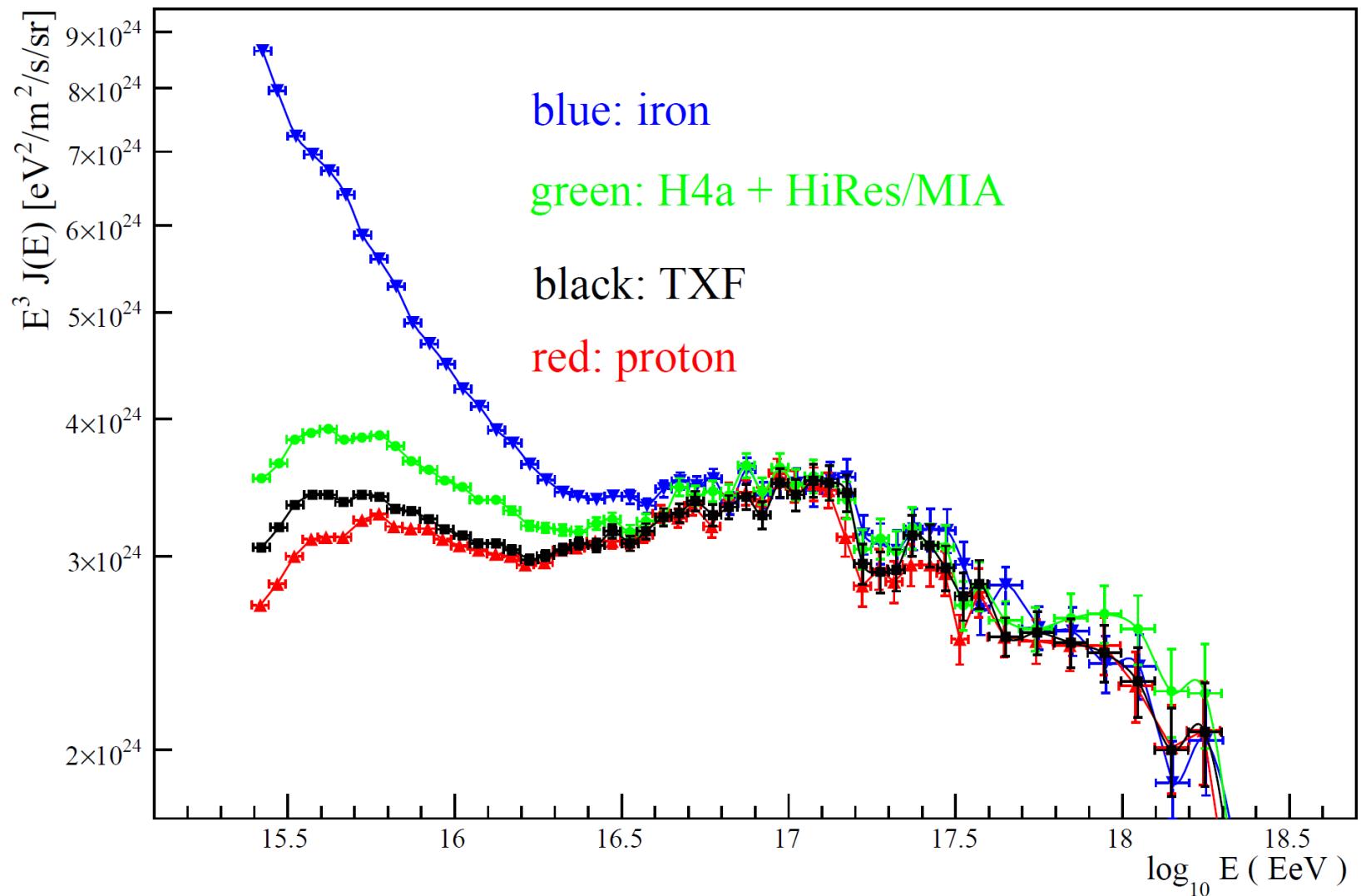


TA Energy Spectrum

TALE Spectrum with Spectra of TA-SD and Auger (Rescaled Energy)



Dependence on Xmax Distribution

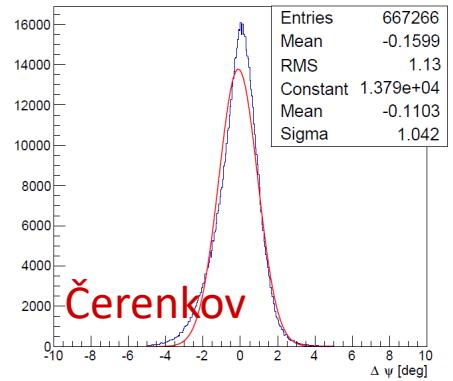


Reconstruction Resolution (1/2)

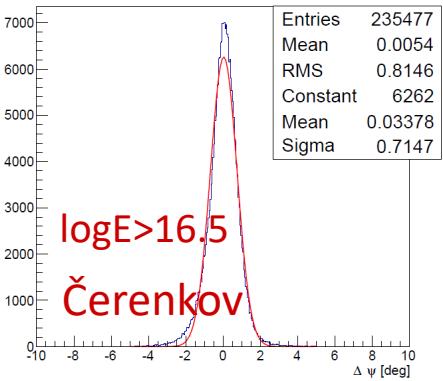


Ψ Angle

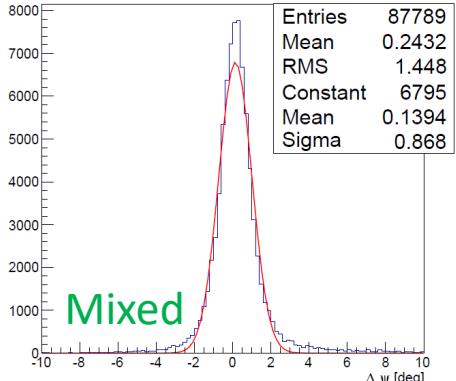
Shower ψ angle: $\Delta \psi$ (deg)



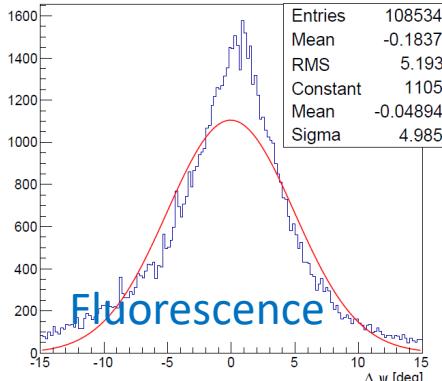
Shower ψ angle: $\Delta \psi$ (deg)



Shower ψ angle: $\Delta \psi$ (deg)

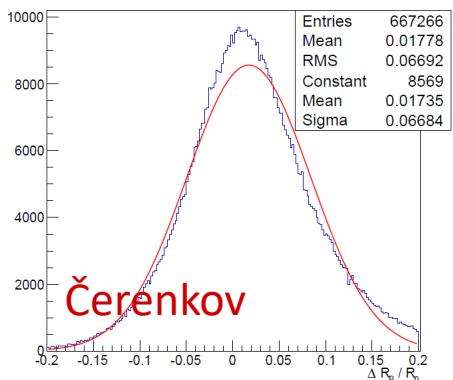


Shower ψ angle: $\Delta \psi$ (deg)

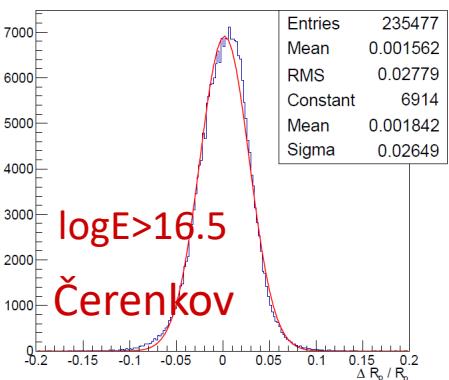


R_p

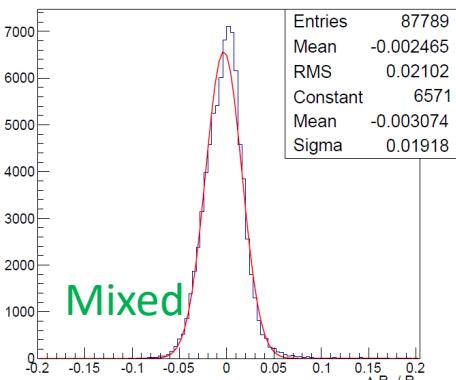
Impact parameter: $\Delta R_p / R_p$



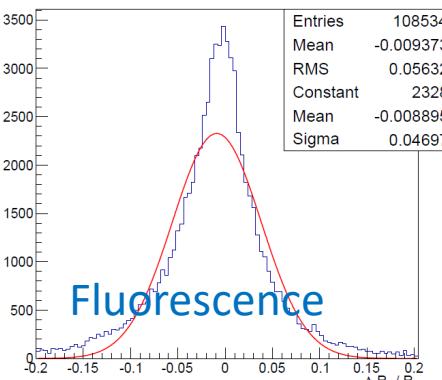
Impact parameter: $\Delta R_p / R_p$



Impact parameter: $\Delta R_p / R_p$



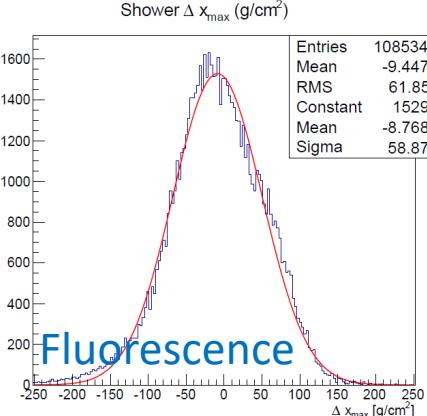
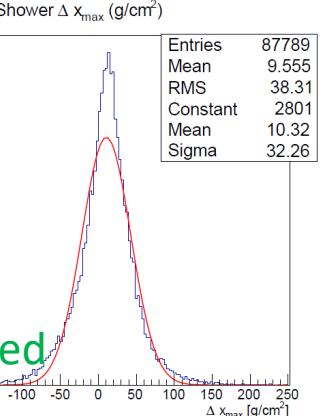
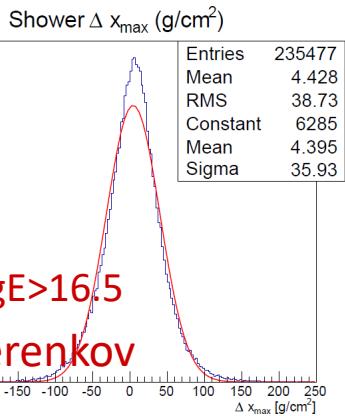
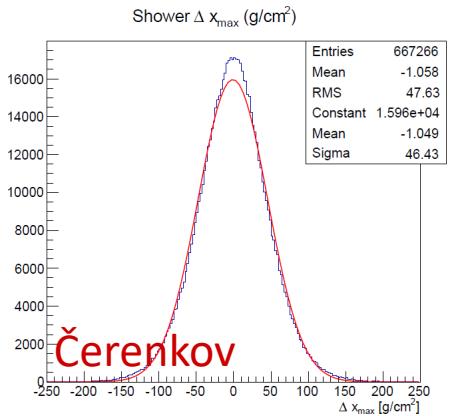
Impact parameter: $\Delta R_p / R_p$



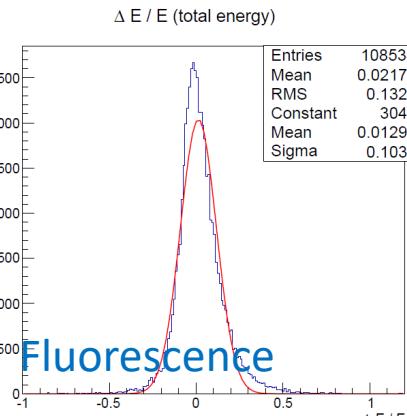
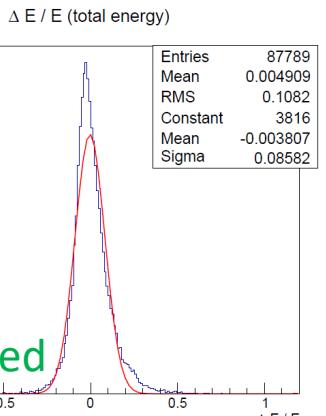
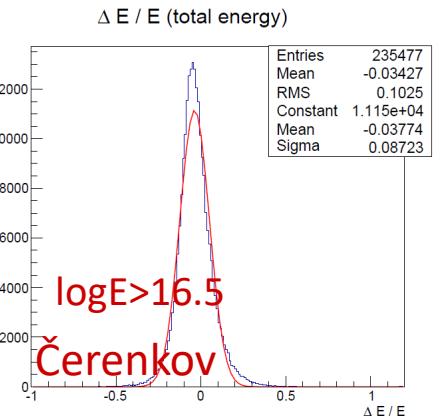
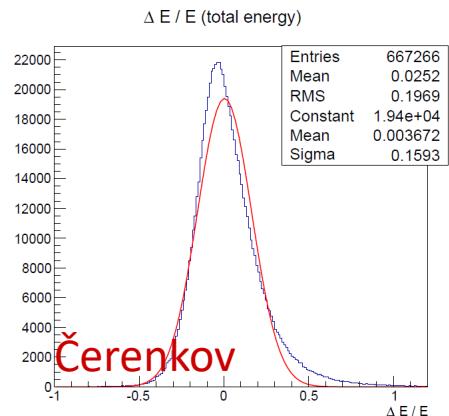
Reconstruction Resolution (2/2)



X_{MAX}



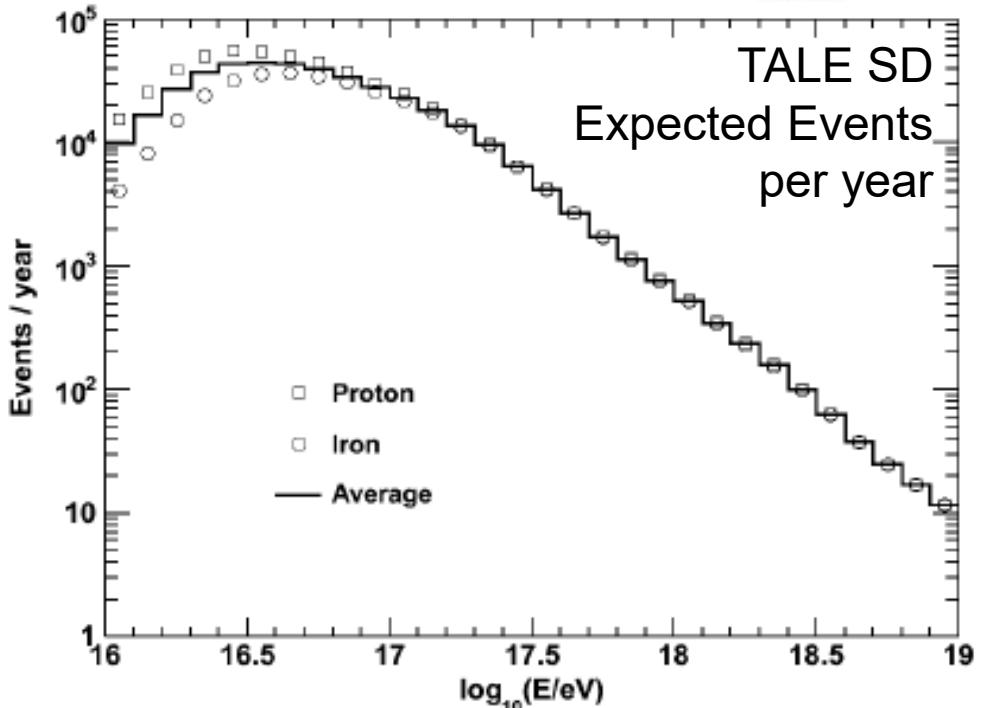
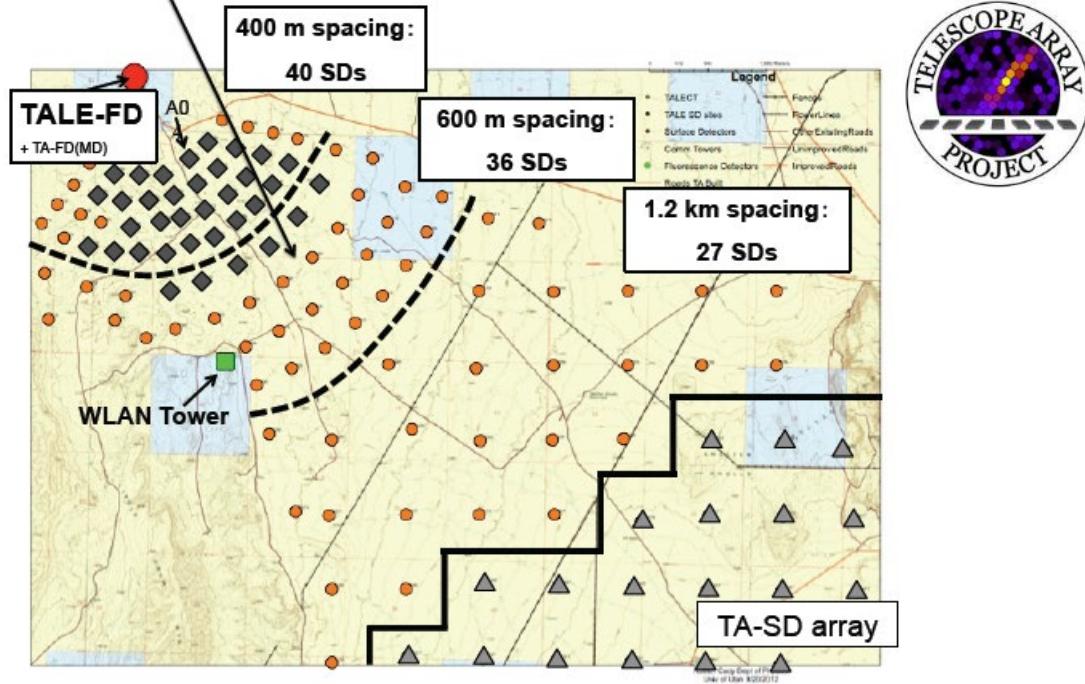
Energy



TALE “Infill”

Surface Detector

- Construction and deployment funded by the Gov't of Japan
- Add infill array (400m and 600m spacing) for hybrid and stand-alone observation.
- Also add counters to build out main TA SD array (1200m separation).



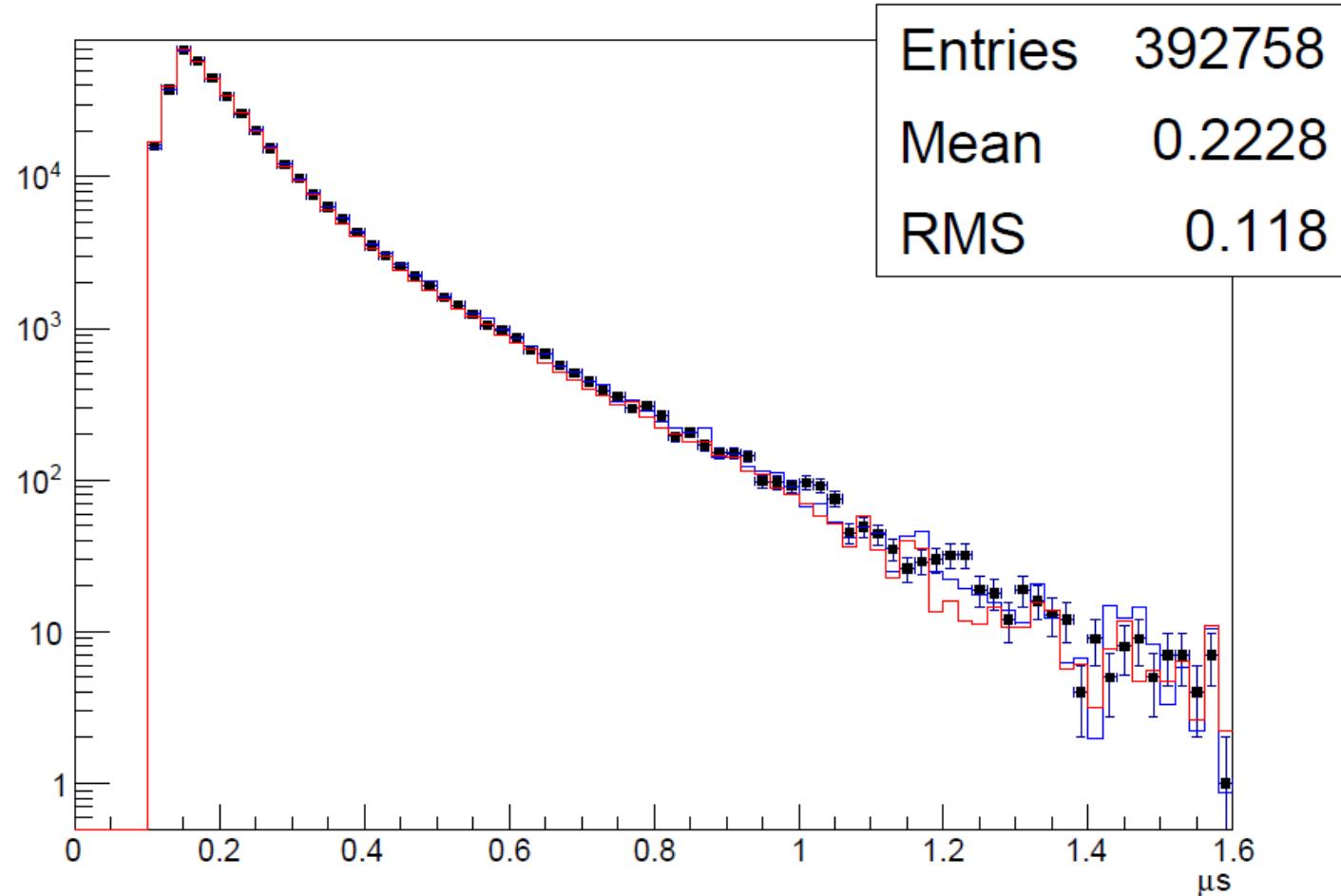
TALE Deployment: Summer 2017



- 103 counters are in place as of fall, 2017
- TALE SD is in shake-down mode

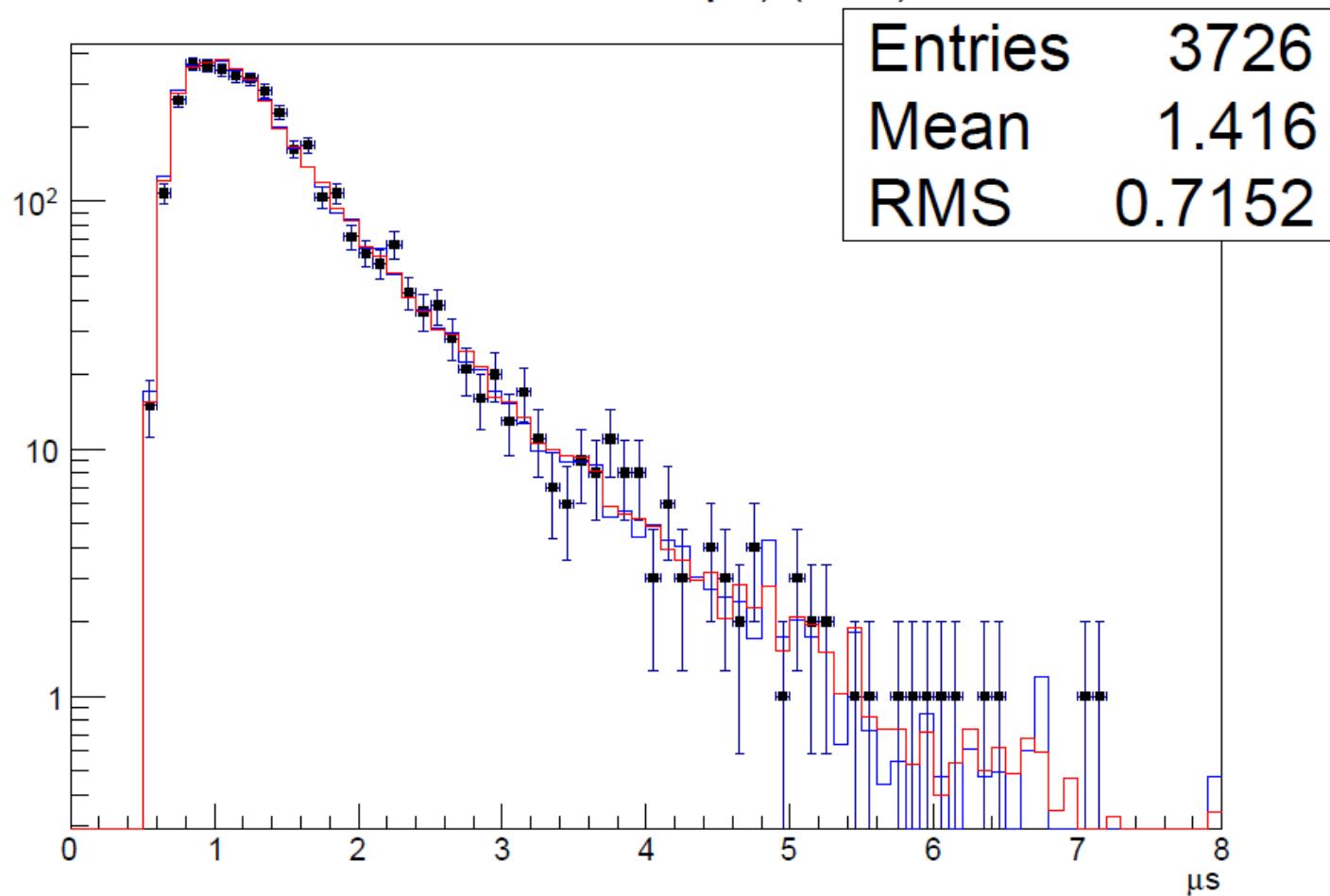


event duration (μ s) (data)





event duration (μ s) (data)



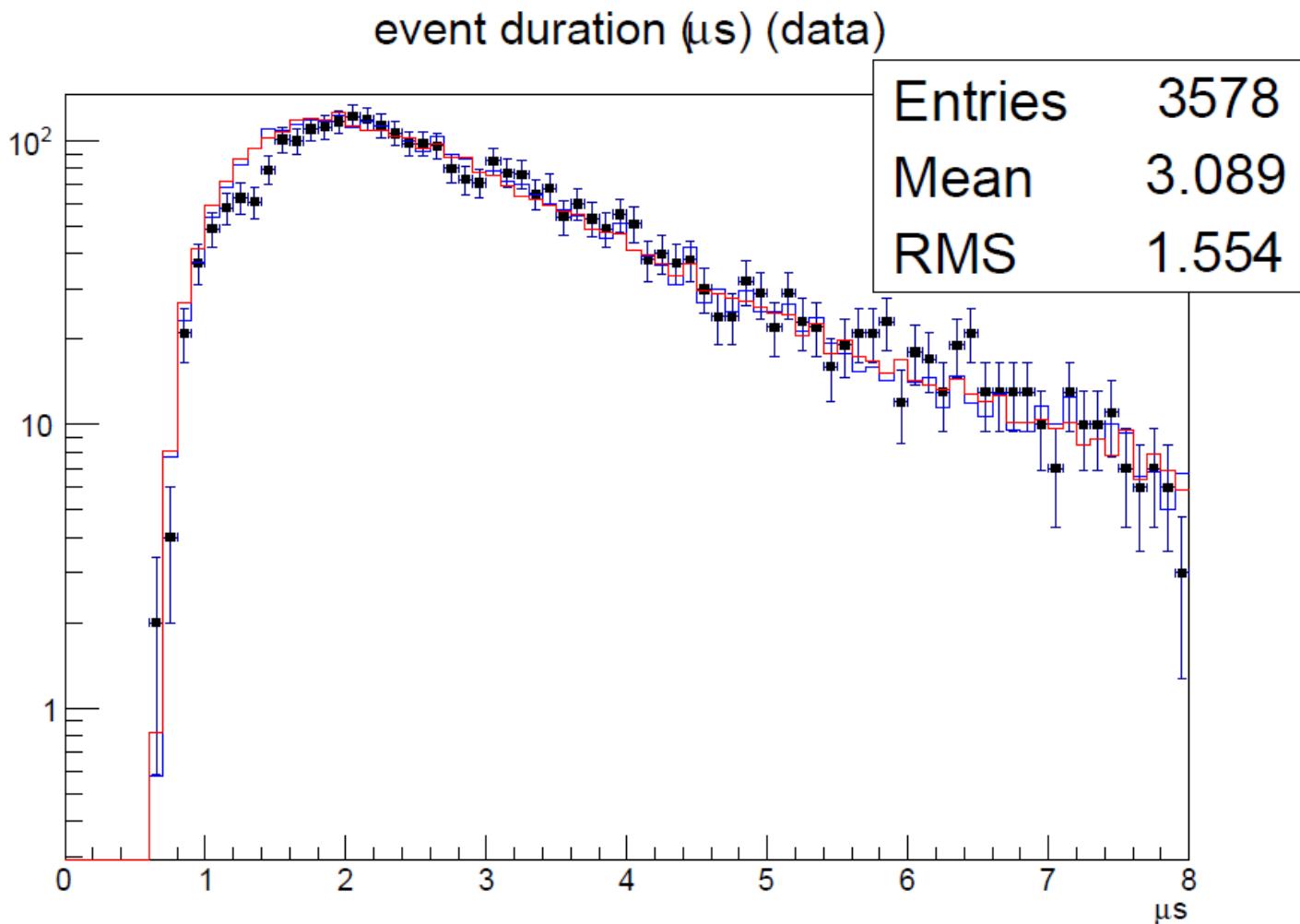
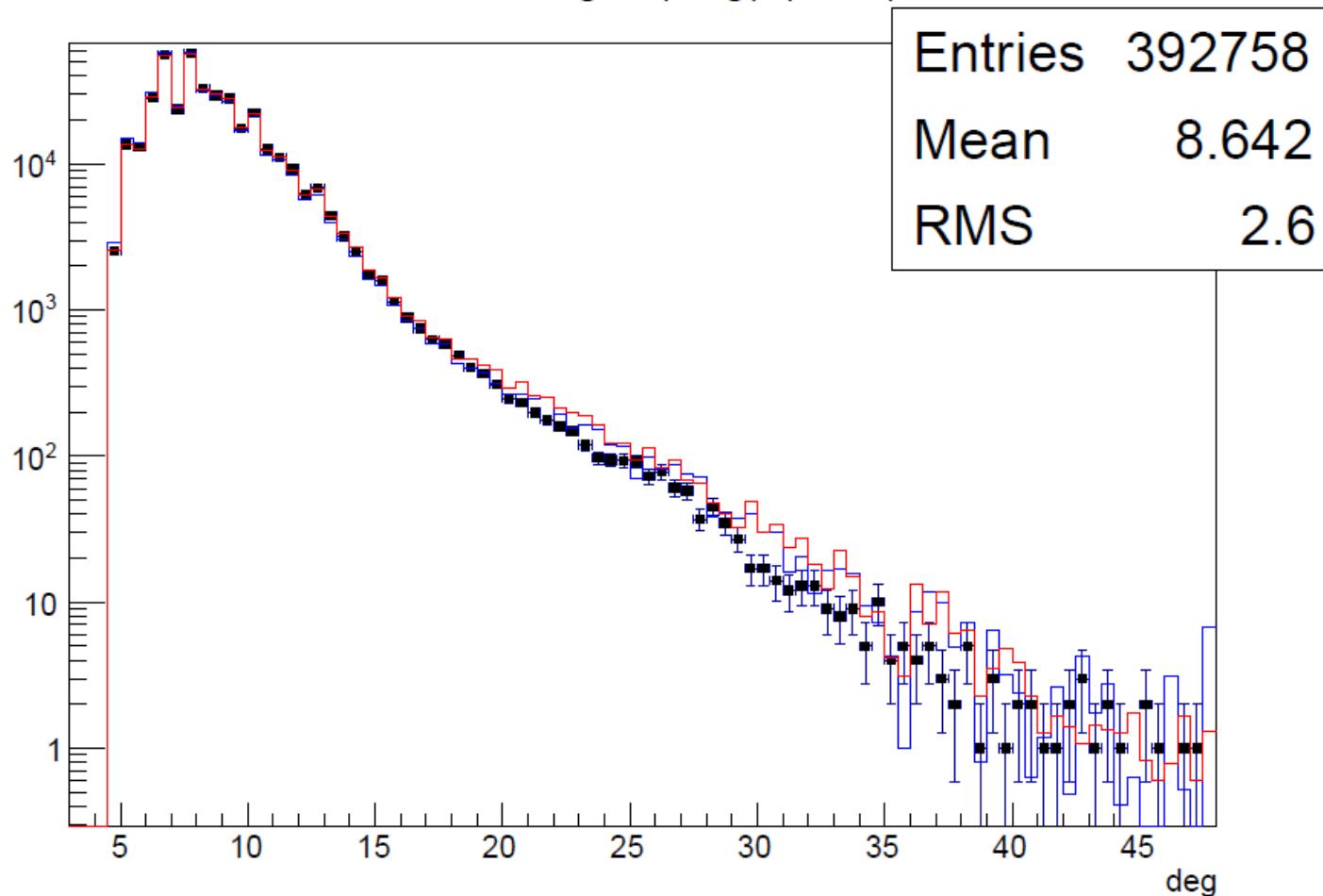


Figure 13: Total event duration (μs), for Cherenkov (left), Mixed (center), and fluorescence events (right). Black points are data, blue / red histograms are MC with mixed composition (H4a / TXF respectively).

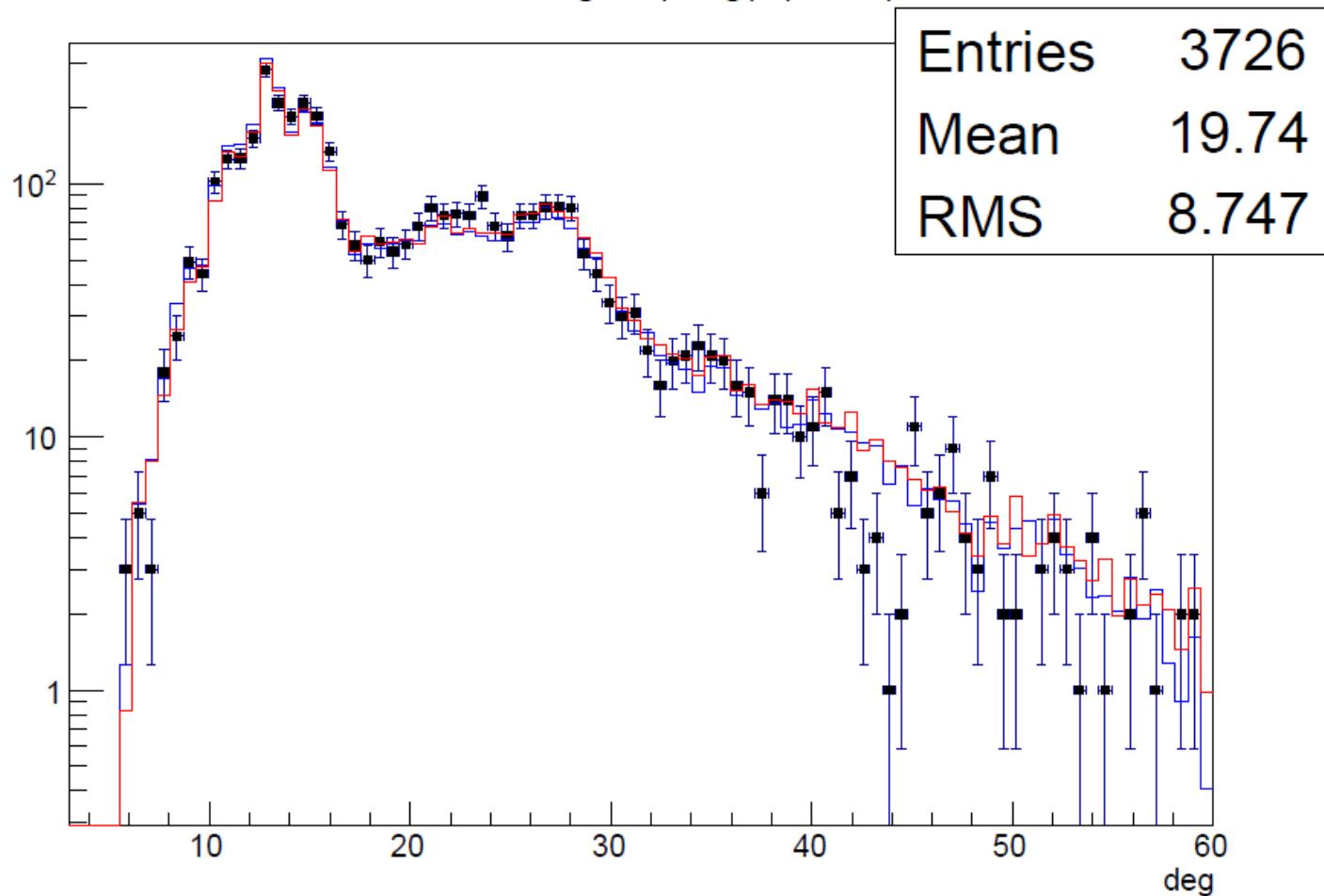


Tracklength (deg) (data)





Tracklength (deg) (data)





Tracklength (deg) (data)

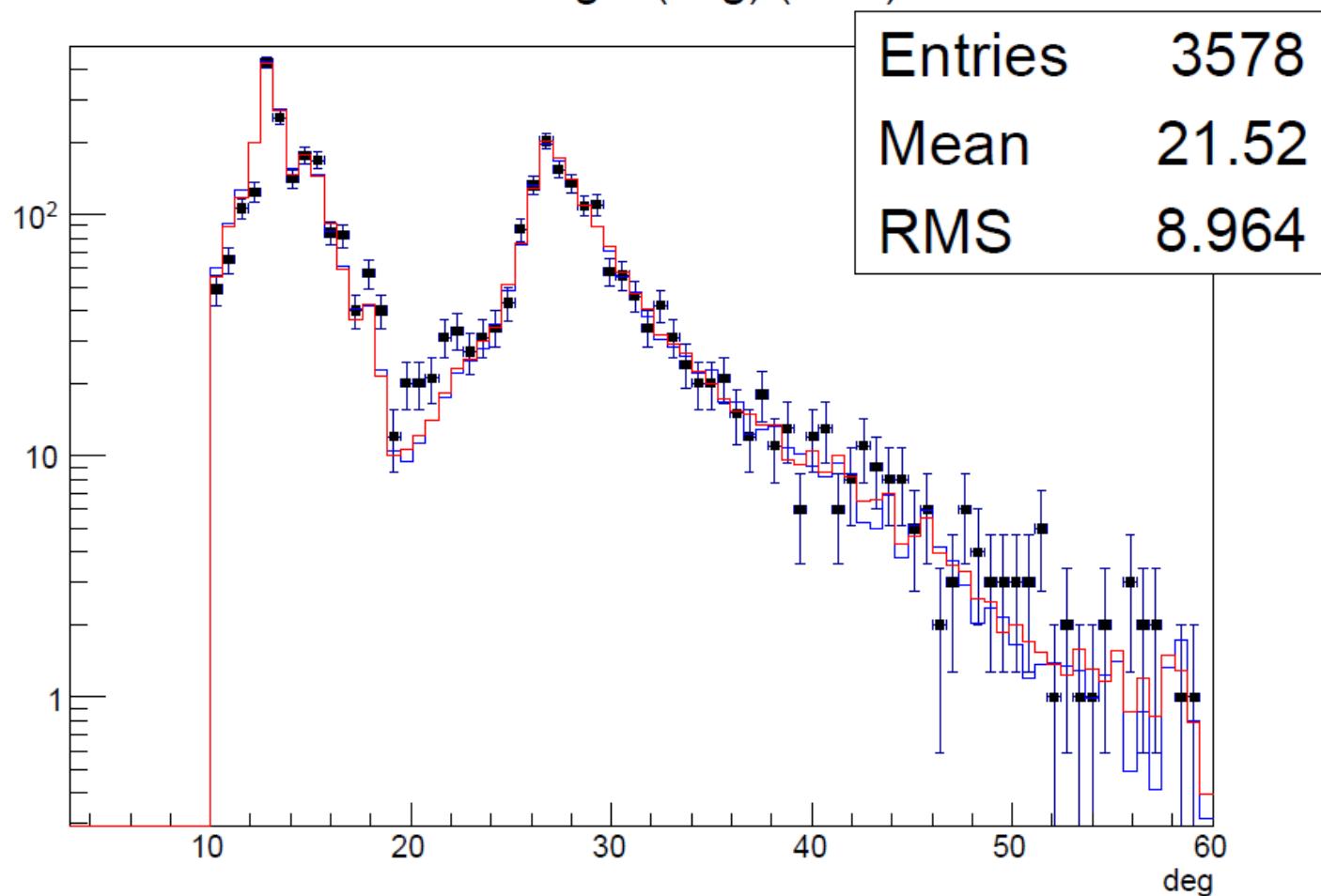


Figure 14: Angular track-length (deg), for Cherenkov (left), Mixed (center), and fluorescence events (right). Black points are data, blue / red histograms are MC with mixed composition (H4a / TXF respectively).

Ratio of calorimetric energy to total energy (QGSII-03, CONEX)

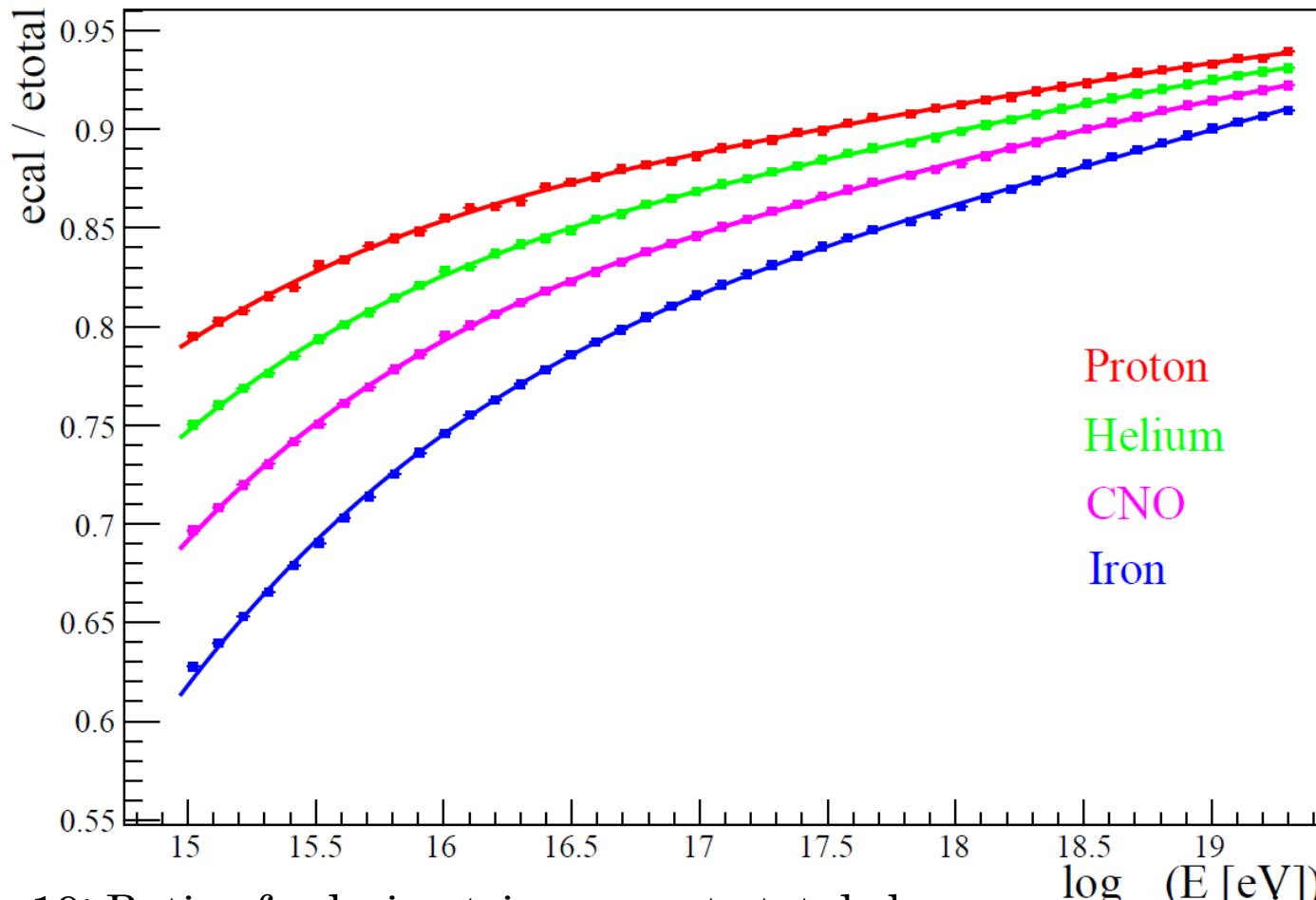


Figure 19: Ratio of calorimetric energy to total shower energy as given by Conex simulations. Simulation sets of mono-energetic showers were used to calculate the ratio. Each point in the figure represents a simulation set and the curves represent a 4-th degree polynomial fit to the point