SUPERGALACTIC STRUCTURE OF MULTIPLETS

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## REASONING

- Hypothesis:
- Sources are correlated with the supergalactic plane.
- Magnetic fields are correlated with sources.
- Multiplets (energy/distance correlated events) may exhibit structure indictive of clustering of sources
- Random fields will diffuse events perpendicular to their average direction. Multiplets should be in 'wedges' not rectangles or circles.


## DATA SUMMARY

- Data:
- 7 years surface detector (SD) data (from ICRC hotspot) \# SD >= 4, Zenith < 55 ${ }^{\circ}$, Pointing Error < $10^{\circ}$
- Additional cuts (for low energy zenith distribution):
- Pointing error $<5^{\circ}$, boundary distance $>1.2 \mathrm{~km}$, Lateral fit $\chi^{2}<10$
- $E \geq \mathbf{1 0}^{19.0} \mathrm{eV}-3027$ events
- Monte Carlo:
- Data positions used, energy interpolated from fully reconstructed Monte Carlo with statistics restricted to the \# of data at all energy thresholds
- Testing energy direction patterns not densities


## MULTIPLET SHAPE

Coherent field deflection
$\delta \approx 0.5^{\circ} Z \frac{S}{k p c} \frac{B}{\mu G} \frac{10^{20} \mathrm{eV}}{E}$
$E_{2}<E_{1}$

Random field deflection

$$
\delta_{r m s} \approx 0.1^{\circ} Z \frac{B_{r m s}}{\mu G} \frac{10^{20} \mathrm{eV}}{E} \sqrt{\frac{S}{k p c}} \sqrt{\frac{L_{c}}{100 p c}}
$$

Coherent and random fields


Brownian Motion on Sphere (Thanks Wikipedia)


## WEDGE MULTIPLETS

- Ranked correlation of energy-distance in sections of spherical caps
- Highest significance is negative indicating possible magnetic deflection


E>=30 EeV, Wedge Width $=30^{\circ}$, Distance $<=80^{\circ}$, Direction $=\mathbf{2 7 0}$

## OVERSAMPLING GRID

SGL (supergalactic longitude) and SGB (supergalactic latitude) Red - 5 year hotspot


- Larger grid spacing due to large parameter scan
- Single simulation takes $\mathbf{> 2 . 5}$ hours on $U$ of $U$ supercomputers


## SCAN PARAMETERS

Scan space limited by computation time, exposure area, statistics.

- Energy Threshold
- 10, 15, 20... 100 EeV
- Wedge width (spherical cap sections)
- $10^{\circ}, 20^{\circ}, 30^{\circ} . .90^{\circ}\left(+5^{\circ}\right.$ on each side of center)
- Maximum Distance
- $15^{\circ}, 20^{\circ}, 25^{\circ} . .90^{\circ}$
- Pointing Direction ( $0^{\circ}$ "up" clockwise)

Mean scan number
(minus less than three events or scans with no change of events)
${ }^{\circ} 0^{\circ}, 5^{\circ}, 10^{\circ} \ldots 355^{\circ}$

## RANKED ENERGY/DISTANCE CORRELATION

- Values are ranked 1 to n. Kendall's correlation is used

```
        (number of concordant pairs) - (number of discordant pairs)
    \frac{1}{2}n(n-1)
```

- Any monotonic function $F(x, y)$ results in $\tau=1$.

- Calculated by permutation
- p-Value is probability correlation is zero


## Source: wikipedia

## MULTIPLET SEARCH RESULT



Correlation significance
Maximized by scan
NOT USED IN FINAL SIGNIFICANCE


Strength/sign of correlation USED TO CALCULATE SIGNIFICANCE

## DATA WEDGE EXAMPLES

- Ranked correlation of energy-distance in sections of spherical caps
- 4 highest significances are negative correlations as expected


$$
\mathrm{E}>=30 \mathrm{EeV}, \text { Width }=30^{\circ},
$$

Distance $<=80^{\circ}$, Direction $=270^{\circ}$


E>=15 EeV, Width $=10^{\circ}$, Dist $<=55^{\circ}$, Dir $=345^{\circ}$


E $>=30$ EeV, Width $=30^{\circ}$, Dist. $<=70^{\circ}$, Dir. $=275^{\circ}$


E>=20 EeV, Width $=70^{\circ}$, Dist. $<=70^{\circ}$, Dir. $=320^{\circ}$

## SUPERGALACTIC STRUCTURE



Strength/sign of correlation
 USED TO CALCULATE SIGNIFICANCE

Average value of correlation shows correlation with supergalactic plane

## PARAMETER DISTRIBUTIONS

- Data scan parameter probability distribution ratios to MC
- For negative correlations +/- 40 deg from supergalactic plane


- Generally biased to more statistics inside the wedge
- Not small scale correlations



## PARAMETER DISTRIBUTIONS




Histogram of pointing directions for negative bin averages cutting at different significances

Most common directions are perpendicular (+/-~10) to the supergalactic plane (and parallel to the left)

## TEST STATISTIC


$x^{2}+b x+c$


HIGHER CURVATURE CLOSER TO SUPERGALACTIC PLANE

## SIGNIFICANCE



Use variable not scanned for: $\tau$ correlation

- Fit to parabola $a x^{2}+b x+c$
- Find probability that $a>=0.000244$ (curvature)

According to Gaussian fit to MC distribution data significance is

According to two passed MC it is

## MC TRIAL EXAMPLES



## Passed MC

(greater curvature than data)

## SIMPLE SIMULATION

- 1/E supergalactic plane deflection for event fraction
- Random "source" position. Gaussian $\sigma=5^{\circ}$ plus $1 / E$ deflection
- Isotropic Exposure and uniform SGL

$$
\delta \approx 0.5^{\circ} Z \frac{S}{k p c} \frac{B}{\mu G} \frac{10^{20} \mathrm{eV}}{E}
$$

- Total energy distribution matches data

$1 / E$ for $10 \%$ of events. $\frac{S}{k p c} \frac{B}{\mu G}=25$


3027 events E>=10^19.0

- Isotropic positions
- Published spectrum


Analysis applied to simulation

## SIMPLE SIMULATION - HOTSPOT

- 1/E supergalactic deflection for event fraction
- Isotropic Exposure and uniform SGL
- Energy distribution published average

- Hotspots not expected on supergalactic plane
- Average ( $\sigma>5.24$ ) location $18.9^{\circ}$ from SGP ( $9.3^{\circ}$ uncertainty). Hotspot is $17^{\circ}$.

N events inside average $30^{\circ}$ equal exposure bins with $\mathrm{E}>=10^{\wedge} 19.75 \mathrm{eV}$

## SIMPLE SIMULATION - COLDSPOT

- 1/E supergalactic deflection for event fraction
- Isotropic Exposure and uniform SGL
- Energy distribution published average

COLDSPOT
MORE LIKELY


N events inside average $30^{\circ}$ equal exposure bins with $E>=10^{\wedge} 19.2 \mathrm{eV}$ and $\mathrm{E}<10^{\wedge} 19.75 \mathrm{eV}$

## SIMULATION FIELD ESTIMATE

- 1/E supergalactic plane deflection
- Isotropic exposure and uniform SGL
- Energy distribution published average

$$
\delta \approx 0.5^{\circ} Z \frac{S}{k p c} \frac{B}{\mu G} \frac{10^{20} \mathrm{eV}}{E}
$$



- Light blue outline curvature 'a' like data
- Red boxes hotspots in $>1 / 100$ trials
- Blue boxes hot/coldspot in >1/100 trials
- Big purple boxes match all three tests
- RESULT: $15<\mathrm{S}^{*} \mathrm{~B}<20 \mathrm{kpc}^{*} \boldsymbol{\mu G}$ and 60 to $\mathbf{7 0 \%}$ isotropic

1/E for $100-10 *$ Y\% of events. $\frac{S}{k p c} \frac{B}{\mu G}=\mathrm{X}$

## COHERENT FIELD ESTIMATES

- Fit 1/E to straight line


$$
\delta \approx 0.5^{\circ} Z \frac{S}{k p c} \frac{B}{\mu G} \frac{10^{20} \mathrm{eV}}{E}
$$

- If $\mathrm{Z}=1$ then $\frac{S}{k p c} \frac{B}{\mu G}=49.24$
- If S = 3.7 Mpc (M82)
- If $\mathrm{Z}=1$ then $\frac{S}{k p c} \frac{B}{\mu G}=40.16$
- If S = 3.7 Mpc (M82)
$\mathrm{E}>=30 \mathrm{EeV}$, Width $=30^{\circ}$, Distance $<=70^{\circ}$, Direction $=275^{\circ}$


## COHERENT FIELD ESTIMATES

- Fit $1 / E$ to straight line



$$
\delta \approx 0.5^{\circ} Z \frac{S}{k p c} \frac{B}{\mu G} \frac{10^{20} \mathrm{eV}}{E}
$$

- If $\mathrm{Z}=1$ then $\frac{S}{k p c} \frac{B}{\mu G}=15.77$
- If S = 3.7 Mpc (M82)
- If $\mathrm{Z}=1$ then $\frac{S}{k p c} \frac{B}{\mu G}=22.56$
- If S = 3.7 Mpc (M82)


## COREHENT FIELD ESTIMATES



- $R^{2}<0$ fits no better than horizontal line (all random field)
- S*B = 0
- Field directions rotated $90^{\circ}$ from wedge direction
- Apparent sheet and filament?


## SUPERGALACTIC FIELD CORRELATION



- Average field in equal solid angle bins
- Fit shows field is consistent with S*B $\approx 19 \mathrm{kpc} * \mu \mathrm{G}$ at supergalactic plane

- R-Square is proportional to random component
- Maximum random field is at hot/coldspot


## POSSIBLE FILAMENT?



Apparent filament?


$$
\nabla \times \mathbf{B}=\mu_{0} \mathbf{J}+\mu_{0} \varepsilon \frac{\partial \widetilde{s}}{\partial t}
$$



$$
I=\frac{\sum(\nabla \times B) * 10^{-10} * 2 \pi\left(1-\cos 25^{\circ}\right)}{4 \pi 10^{-7} \mathrm{Tm} / A}
$$

## POSSIBLE FILAMENT?



## CONCLUSION

- Multiplet structure correlates with the supergalactic plane
- Suggests sources, coherent, and random fields correlated with SGP
- A measure of magnetic fields using UHECR
- Consistent with supergalactic sheet field of $\sim 19 \mathrm{kpc} * \mu \mathrm{G}$ (estimate of data and simulation)
- Simulation exhibits multplet curvature and hot/coldspot behavior



## ADDITIONAL

## Steps

- Data: tight cuts for $10^{\wedge} 19 \mathrm{eV}$ resolution (like energy spectrum anisotropy).
- Variables needed: energy and pointing direction in supergalactic coords.
- Field of view $2^{\circ}$ equally spaced grid (down to -16 Dec).
- Rotated into supergalactic coords (supergalactic hypothesis).
- Wedges:
- Spherical caps centered on grid point (like hotspot analysis).
- Angular distance Haversine formula: distance = $\mathbf{2}^{*}$ atan2(sqrt(a),sqrt(1-a)); a = $\sin ((\operatorname{lat2-lat1)/2).\wedge 2+\operatorname {cos}(lat1).*~}$ $\cos (\operatorname{lat} 2) . * \sin ((\operatorname{lon} 2-\operatorname{lon} 1) / 2) . \wedge 2 ;$
- Take section of some angular size out of spherical cap. Rotation angle is clockwise from 90deg latitude.
- direction = atan2(cos(lat2) .* $\sin (l o n 2-l o n 1), \cos (l a t 1) . * \sin (l a t 2)-\sin ($ lat1 $) ~ . * ~ \cos (l a t 2) .{ }^{*} \cos ($ lon2-Ion1));
- Point wedge section in some direction. Defined by grid point and center of wedge,


## Steps

Define scan space for parameters.

- Wedge distances (Haversine distance formula)
- Wedge widths
- Wedge directions
- Minimum energy thresholds
- Calculate how event energies inside wedge change with distance from grid point.
- Rank correlations used to calculate likelihood and strength of functional dependence between two variables. Robust and model independent.
- At each grid point scan parameters for wedge that maximizes correlation significance.
- Resulting correlations can be negative or positive of any strength.
- Find correlation between strength of correlations and the supergalactic plane.
- Bin field of view into equal solid angle bins of supergalactic latitude (each containing equal number of grid points).
- Calculate mean correlations in each FOV bin.
- Parabolic dependence of correlations expected by hypothesis and simulation. Fit latitude correlation mean to parabola. Find how often this happens in isotropic simulation.


Removing hotspot area (144, 41 equatorial)


Analysis result


Scrambling energies of hotspot events E>57 EeV

Configuration of events E>57 EeV within hotspot area matters

## EACH YEAR OF DATA




Five years


Six years


Seven years

## FIRST FOUR AND LAST FOUR YEARS



First Four Years


Last Four Years

Seven Years Data (one year overlap)

## RANDOM MONTE CARLO EXAMPLE



Significance of correlation
NOT USED IN FINAL SIGNIFICANCE


Strength/sign of correlation USED TO CALCULATE SIGNIFICANCE

## CORR. FILTER SHAPE

- Test of magnetic field configuration



## DIRECTION OF CORRELATIONS


$1 / 3$ grid points greater than 3 sigma

## DIRECTION OF CORRELATIONS



Grid points greater than 4.5 sigma

