SUPERGALACTIC STRUCTURE OF MULTIPLETS

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REASONING

- Hypothesis:
 - Sources are correlated with the supergalactic plane.
 - Magnetic fields are correlated with sources.
 - Multiplets (energy/distance correlated events) may exhibit structure indictive of clustering of sources
 - Random fields will diffuse events perpendicular to their average direction. Multiplets should be in 'wedges' not rectangles or circles.

DATA SUMMARY

• Data:

ame as Energy Spectrum Anisotropy study Abbasi et al 2018, ApJ, 862, 91

- 7 years surface detector (SD) data (from ICRC hotspot) # SD >= 4, Zenith < 55°, Pointing Error < 10°
 - Additional cuts (for low energy zenith distribution):
 - Pointing error < 5°, boundary distance > 1.2 km , Lateral fit χ^2 < 10
 - E ≥ 10^{19.0} eV 3027 events
- Monte Carlo:
 - Data positions used, energy interpolated from fully reconstructed Monte Carlo with statistics restricted to the # of data at all energy thresholds
 - Testing energy direction patterns not densities

MULTIPLET SHAPE

Coherent field deflection $\delta \approx 0.5^{\circ} Z \frac{S}{kpc} \frac{B}{\mu G} \frac{10^{20} eV}{E}$ $E_{2} < E_{1}$

Coherent and random fields



Random field deflection

 E_1

 $\delta_{rms} \approx 0.1^{\circ} Z \frac{B_{rms}}{\mu G} \frac{10^{20} eV}{E} \sqrt{\frac{S}{kpc}} \sqrt{\frac{L_c}{100pc}}$

Brownian Motion on Sphere (Thanks Wikipedia)

Drift-diffusion -> Wedge Multiplets



WEDGE MULTIPLETS

- Ranked correlation of energy-distance in sections of spherical caps
- Highest significance is negative indicating possible magnetic deflection



OVERSAMPLING GRID

SGL (supergalactic longitude) and SGB (supergalactic latitude)



Red – 5 year hotspot



 Larger grid spacing due to large parameter scan
 Single simulation takes >2.5 hours on U of U supercomputers

SCAN PARAMETERS

Scan finds most likely real correlation Scan result corr. significance is not used for final significance Scan space limited by computation time, exposure area, statistics.

- Energy Threshold
 - 10, 15, 20...100 EeV
- Wedge width (spherical cap sections)
 - 10°, 20°, 30°...90° (+5° on each side of center)
- Maximum Distance
 - 15°, 20°, 25°...90°
- Pointing Direction (0° "up" clockwise)
 - 0°, 5°, 10°...355°

Mean scan number (minus less than three events or scans with no change of events)

RANKED ENERGY/DISTANCE CORRELATION

Values are ranked 1 to n. Kendall's correlation is used

 $= \frac{\text{(number of concordant pairs)} - \text{(number of discordant pairs)}}{\frac{1}{2}n(n-1)}$

Any monotonic function F(x,y) results in τ = 1. Removes model assumption



Source: wikipedia

p-Values

- Calculated by permutation
- p-Value is probability correlation is zero

MULTIPLET SEARCH RESULT





Correlation significance Maximized by scan NOT USED IN FINAL SIGNIFICANCE

Strength/sign of correlation USED TO CALCULATE SIGNIFICANCE

Negative correlations of greater significance correlated with supergalactic plane

DATA WEDGE EXAMPLES

• Ranked correlation of energy-distance in sections of spherical caps

4 highest significances are negative correlations as expected



Distance<=80°, **Direction = 270°**



E>=30 EeV, Width = 30°, Dist.<=70°, Dir. = 275°





E>=15 EeV, Width = 10°, Dist<=55°, Dir = 345° E>=20 EeV, Width = 70°, Dist.<=70°, Dir. = 320° 11

SUPERGALACTIC STRUCTURE





Strength/sign of correlation USED TO CALCULATE SIGNIFICANCE

Average value of correlation shows correlation with supergalactic plane

PARAMETER DISTRIBUTIONS

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Energy Threshold [EeV]

- Data scan parameter probability distribution ratios to MC
 - For negative correlations +/- 40 deg from supergalactic plane



4.5

PDF Ratio (Data/MC) .2 5.5

> 0.5 0

0 10 20 30 40 50 60 70 80



- Generally biased to more statistics inside the wedge
- Not small scale correlations

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PARAMETER DISTRIBUTIONS



Histogram of pointing directions for *negative bin averages* cutting at different significances

Most common directions are perpendicular (+/- ~10) to the supergalactic plane (and parallel to the left)

TEST STATISTIC





 $ax^2 + bx + c$

HIGHER CURVATURE CLOSER TO SUPERGALACTIC PLANE

SIGNIFICANCE



Use variable not scanned for: τ correlation

- Fit to parabola $ax^2 + bx + c$
- Find probability that a>=0.000244 (curvature)

According to Gaussian fit to MC distribution data significance is 4.560

According to two passed MC it is 4.210 (large uncertainty)

MC TRIAL EXAMPLES



Passed MC (greater curvature than data)

SIMPLE SIMULATION

- 1/E supergalactic plane deflection for event fraction
 - Random "source" position. Gaussian $\sigma = 5^{\circ}$ plus 1/E deflection
- Isotropic Exposure and uniform SGL
- Total energy distribution matches data







1/E for 10% of events. $\frac{S}{kpc}\frac{B}{\mu G} = 25$



- Isotropic positions
- Published spectrum

Analysis applied to simulation

 $\delta \approx 0.5^{\circ} Z \frac{S}{kpc} \frac{B}{\mu G} \frac{10^{20} eV}{E}$

SIMPLE SIMULATION - HOTSPOT

- 1/E supergalactic deflection for event fraction
- Isotropic Exposure and uniform SGL
- Energy distribution published average



N events inside average 30° equal exposure bins with E>=10^19.75 eV

SIMPLE SIMULATION - COLDSPOT

- 1/E supergalactic deflection for event fraction
- Isotropic Exposure and uniform SGL
- Energy distribution published average

COLDSPOT MORE LIKELY





N events inside average 30° equal exposure bins with E>=10^19.2 eV and E<10^19.75 eV

SIMULATION FIELD ESTIMATE

- 1/E supergalactic plane deflection
- Isotropic exposure and uniform SGL
- Energy distribution published average



1/E for 100-10*Y% of events. $\frac{S}{kpc}\frac{B}{\mu G} = X$

- Light blue outline curvature 'a' like data
- Red boxes hotspots in >1/100 trials
- Blue boxes hot/coldspot in >1/100 trials
- Big purple boxes match all three tests
 - RESULT: 15 < S*B < 20 kpc*μG and 60 to 70% isotropic
 Agrees with data only determination of 19 kpc*μG

 $\delta \approx 0.5^{\circ} Z \frac{S}{knc} \frac{B}{\mu G} \frac{10^{20} eV}{F}$

COHERENT FIELD ESTIMATES

100

90

70

60 50

40

Energy (EeV) 80

• Fit 1/E to straight line





0.03 0.035

 $\delta \approx 0.5^{\circ} Z \frac{S}{kpc} \frac{B}{\mu G} \frac{10^{20} eV}{E}$

E>=30 EeV, Width = 30°, Distance<=80°, Direction = 270°



E>=30 EeV, Width = 30°, Distance<=70°, Direction = 275°

COHERENT FIELD ESTIMATES

50

45

40 () 35 IJ

30 Serenda

25 20

• Fit 1/E to straight line





 $\delta \approx 0.5^{\circ} Z \frac{S}{kpc} \frac{B}{\mu G} \frac{10^{20} eV}{E}$

E>=15 EeV, Width = 10°, Distance<=55°, Direction = 345°





E>=30 EeV, Width = 30°, Distance<=70°, Direction = 320°

COREHENT FIELD ESTIMATES





S*B [kpc* μ B]

R²<0 fits no better than horizontal line (all random field)
S*B = 0

Field directions rotated 90[°] from wedge direction

Apparent sheet and filament?

SUPERGALACTIC FIELD CORRELATION



NOT SCANNED FOR (signed ranked correlations)

- Average field in equal solid angle bins
 - Fit shows field is consistent with
 S*B ≈ 19 kpc* μG at supergalactic plane





- R-Square is proportional to random component
- Maximum random field is at hot/coldspot

POSSIBLE FILAMENT?



Apparent filament?

0.06

0.04

0

-0.02

-0.04 -0.06

[kpc*Amperes] 0.02

s,



Apparent current flowing towards us inside 25 $^{\circ}$ bin

 $\partial \mathbf{E}$ $abla imes {f B} = \mu_0 {f J} + \mu_0 {f e_0} {f \partial t}$



$$=\frac{\sum (\nabla \times B) * 10^{-10} * 2\pi (1 - \cos 25^{\circ})}{4\pi 10^{-7} Tm/A}$$

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POSSIBLE FILAMENT?





CONCLUSION

- Multiplet structure correlates with the supergalactic plane
- Suggests sources, coherent, and random fields correlated with SGP
 - A measure of magnetic fields using UHECR
- Consistent with supergalactic sheet field of ~19 kpc*μG (estimate of data and simulation)
 - Simulation exhibits multplet curvature and hot/coldspot behavior





ADDITIONAL

Steps

- Data: tight cuts for 10^19 eV resolution (like energy spectrum anisotropy).
 - Variables needed: energy and pointing direction in supergalactic coords.
- Field of view 2° equally spaced grid (down to -16 Dec).
 - Rotated into supergalactic coords (supergalactic hypothesis).
 - Wedges:

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- Spherical caps centered on grid point (like hotspot analysis).
 - Angular distance Haversine formula: distance = 2 * atan2(sqrt(a),sqrt(1 a)); a = sin((lat2-lat1)/2).^2 + cos(lat1).* cos(lat2).* sin((lon2-lon1)/2).^2;
- Take section of some angular size out of spherical cap. Rotation angle is clockwise from 90deg latitude.
 - direction = atan2(cos(lat2) .* sin(lon2-lon1), cos(lat1) .* sin(lat2) sin(lat1) .* cos(lat2) .* cos(lon2-lon1));
- Point wedge section in some direction. Defined by grid point and center of wedge.



Steps

- Define scan space for parameters.
 - Wedge distances (Haversine distance formula)
 - Wedge widths

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- Wedge directions
- Minimum energy thresholds
- Calculate how event energies inside wedge change with distance from grid point.
 - Rank correlations used to calculate likelihood and strength of functional dependence between two variables. Robust and model independent.
- At each grid point scan parameters for wedge that maximizes correlation significance.
 - Resulting correlations can be negative or positive of any strength.
- Find correlation between strength of correlations and the supergalactic plane.
 - Bin field of view into equal solid angle bins of supergalactic latitude (each containing equal number of grid points).
 - Calculate mean correlations in each FOV bin.
- Parabolic dependence of correlations expected by hypothesis and simulation. Fit latitude correlation mean to parabola. Find how often this happens in isotropic simulation.

HOTSPOT AREA CONTRIBUTION



Removing hotspot area (144, 41 equatorial)



Removing events and grid 25° around hotspot

HOTSPOT AREA SCRAMBLING





Analysis result

Scrambling energies of hotspot events E>57 EeV

Configuration of events E>57 EeV within hotspot area matters





First year

Two years

Three years

Four years



FIRST FOUR AND LAST FOUR YEARS



Seven Years Data (one year overlap)

RANDOM MONTE CARLO EXAMPLE

MC created using data positions

with random energies from spectrum





Significance of correlation NOT USED IN FINAL SIGNIFICANCE

Strength/sign of correlation USED TO CALCULATE SIGNIFICANCE

CORR. FILTER SHAPE

Test of magnetic field configuration

THESE ARE ALL DATA



Wedges (spherical cap sections)

- Test of uniform and random fields
- Spherical caps
 - Test of random fields

Squares

Test of uniform fields
0.5 to 3 deg wide

DIRECTION OF CORRELATIONS



1/3 grid points greater than 3 sigma

DIRECTION OF CORRELATIONS



Grid points greater than 4.5 sigma