

Technology issues in Čerenkov light images

or

Technologies implications for RICH performance

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Content of the talk

Relates implications on technologies choices $\sigma(\theta_C)$ to the physics goals

Considering:

★ Kind of physics measurements

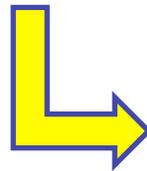
★ Momentum range to be covered

★ Machine environments

Particle density in the final state, operation frequency ,.....



geometry and technologies choices to achieve a given angular resolution



keep all the contributions to the resolution under control during the whole lifetime of the experiment

Many applications of RICH detectors

Hadronic environment

ALICE
LHCb
PANDA
NA62
COMPASS

e+e- environment

BaBar, BELLE
BELLE upgrade (SUPER-B)

Space experiments (on satellite and balloon)

AMS (measures flux of charged particles and light nuclei)
CREAM

Underground

ANTARES, NESTOR, NEMO ,
KM3net, AMANDA, ICECUBE

Nuclear physics

ALICE
JLAB

Astrophysics

A long list...

To measure the Čerenkov angle θ_C

Main contributions to angular resolution $\sigma(\theta_C)$ from :

Chromaticity
 N_γ multiplicity



Radiator

($n(\lambda)$, thickness, transparency...)

$N_{p.e.}$ multiplicity
Spatial localization



Photon detection

(QE, photon collection efficiency, pixel size,...)

Emission point
Photon path



Geometry

(Proximity focus, focussed geometry,...)

Tracking
Multiple scattering
Decays, interactions,..



“External” error

RICH2010 Cassis 3-10 may

Čerenkov detectors performance

The angular resolution per photon:

$$\sigma(\theta_c) = \sqrt{\sigma(\theta_{\text{rad}})^2 + \sigma(\theta_{\text{PD}})^2 + \sigma(\theta_{\text{geom}})^2 + \sigma(\theta_{\text{tr}})^2}$$

And the separating power:

$$\sigma_{\text{ring}}(\theta_c) = \frac{\sigma(\theta_c)}{\sqrt{N_{\text{pe}}}}$$

$$N_{\sigma} \approx \frac{(m_1^2 - m_2^2)}{(2 p^2 \sqrt{n^2 - 1} \sigma(\theta_c))}$$

The number of photo-electrons N_{pe} :

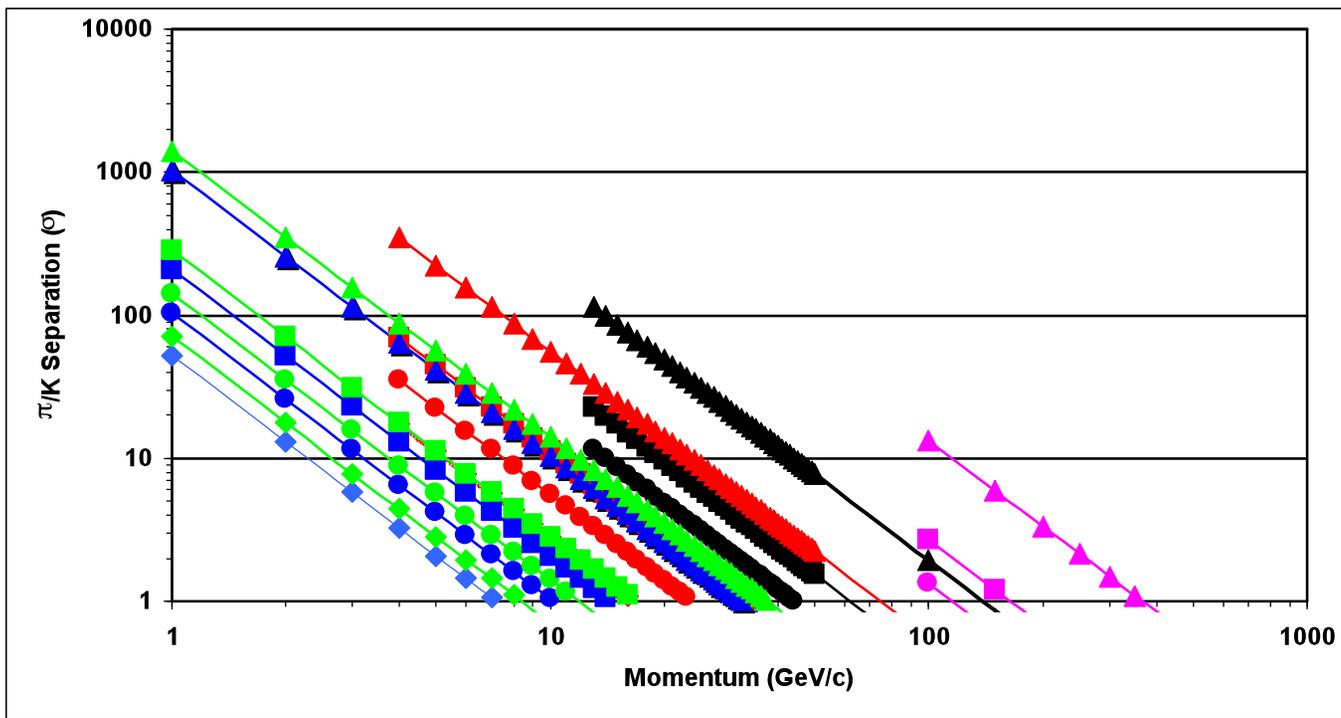
$$N_{\text{pe}} = 370L \int \varepsilon \sin^2 \theta_c dE = L N_0 \sin^2 \theta_c$$

Usually N_0 between ~ 20 and 100

General rule: **minimize** $\sigma(\theta_c)$
maximize N_{pe}

Čerenkov angle resolution and separating power

π/K separation



Refractive Indices

$N=1.474$ (Fused Silica)

$N=1.27$ (C_6F_{14} CRID)

$N=1.02$ (Typical Silica Aerogel)

$N=1.001665$ (C_5F_{12}/N_2 CRID Mix)

$N=1.0000349$ (He)

$\sigma(\theta_c)$

◆ 2 mrad

● 1 mrad

■ 0.5 mrad

▲ 0.1 mrad

(Plot from B.Ratcliff)

RICH detectors by angular resolution

$$\sigma(\theta_C) \approx O(10 \text{ mrad})$$

Ex: ALICE, BELLE, BELLE upgrade, JLAB, CLEO-C,
BaBar and HERMES (closed)

differ by machine environment machine, particle density,
BUT momentum range similar

(...in between the AMS experiment)

$$\sigma(\theta_C) \approx O(1 \text{ mrad})$$

Ex.: COMPASS, LHCb, NA62

RICHes in experiments at hadron accelerators

Example of RICH detectors with $\sigma(\theta_c) \approx 0(10 \text{ mrad})$

ALICE started to operate **at LHC**

The RICHes detectors of HERMES, BaBar DIRC, BELLE, CLEO-c have operated successfully with this range of resolution.

Examples of RICH detectors with $\sigma(\theta_c) \approx 0(1 \text{ mrad})$

LHCb started to operate **at LHC**

NA62 starting to operate in 2012 **at SPS**

The RICH of ALICE

*See detailed talks by P. Martinengo
at this conference*

Physics aims:

mainly proton ID in the range 0(few GeV), d and α also interesting
physics measurement: inclusive hadron spectra from Pb-Pb collisions
measurement of particle ratios vs pT

For particle ID over the momentum range also dE/dx, TOF, TRD are used

The RICH must cover the range 1-5 GeV/c (1-3 GeV/c for π/k and 2-5 GeV/c for p)

Environment:

Pb-Pb collisions

Density of charged particles about 2000/ rapidity unit

Low rate (< 10KHz)

Geometry:

limited 'radial' space \rightarrow compact detector \rightarrow proximity focus

The RICH of ALICE : the HMPID choice

Proximity focused

Radiator: 15 mm C_6F_{14} (liquid) with $n=1.2989$ @ 175 nm
 $\theta_c = 694$ mrad

Photon converter: Reflective layer of CsI (QE= 25% @ 175 nm)

Photodetectors:

MWPC with CH_4 at atmospheric pressure (4 mm sensitive gap)
analogue pad readout

*VHMPID: upgrade planned to extend PID to 30 GeV/c.
 C_5F_{12} gas radiator (1m) mirror-focused RICH
CsI photocathode + GEM photon detector*

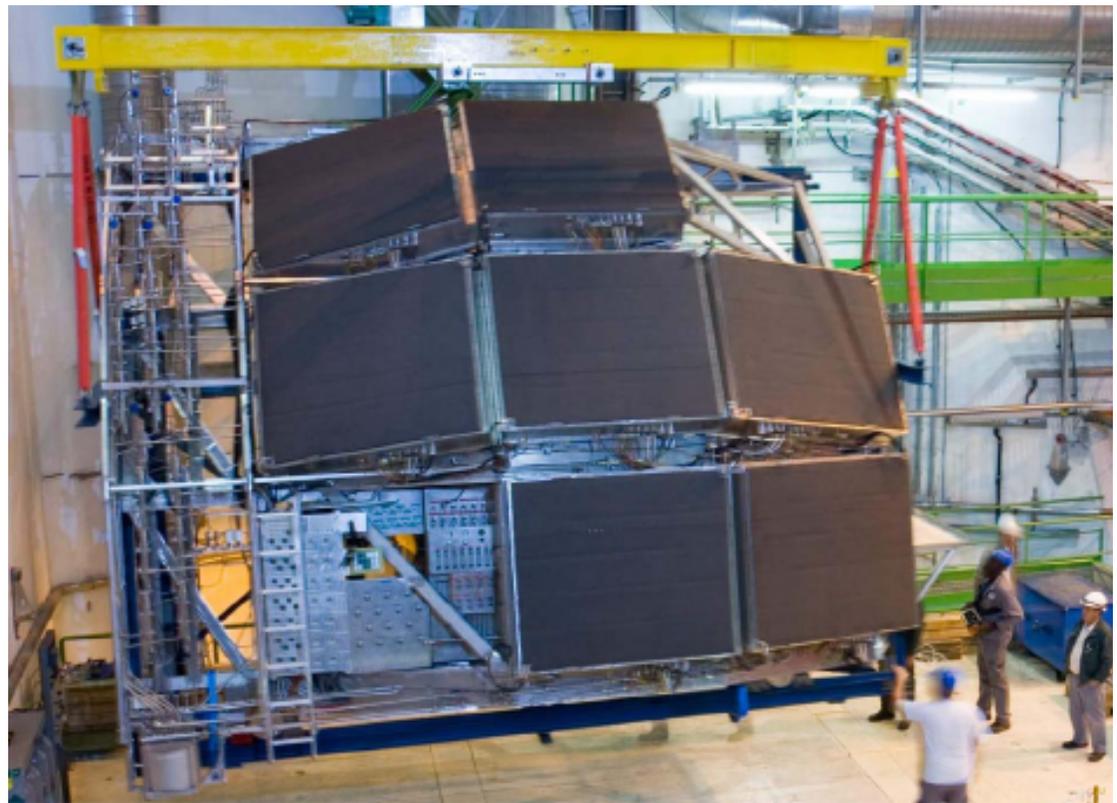
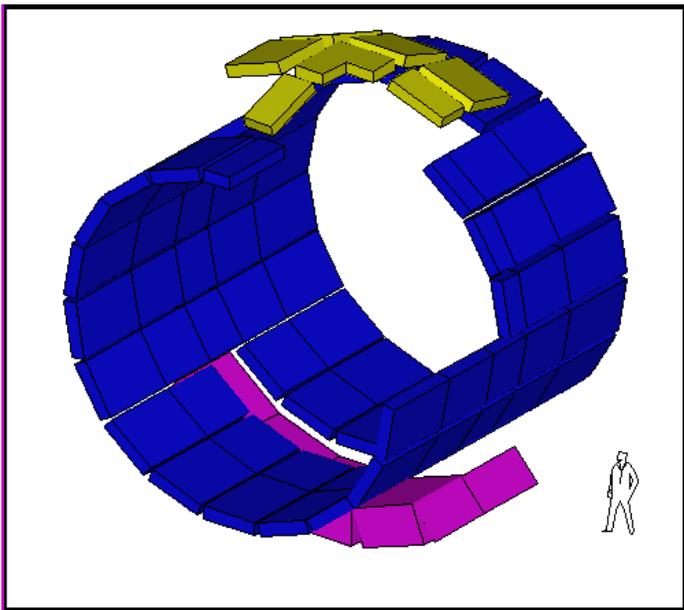
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(talks by A. DiMauro at this conference)

The RICH of ALICE

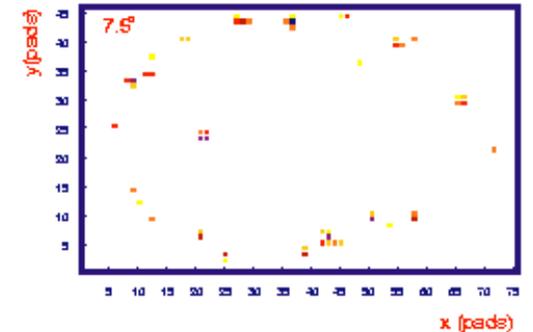
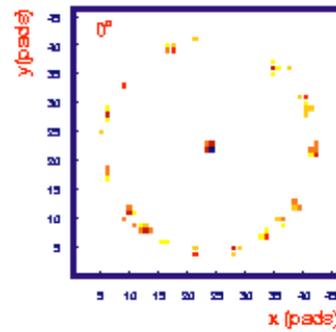
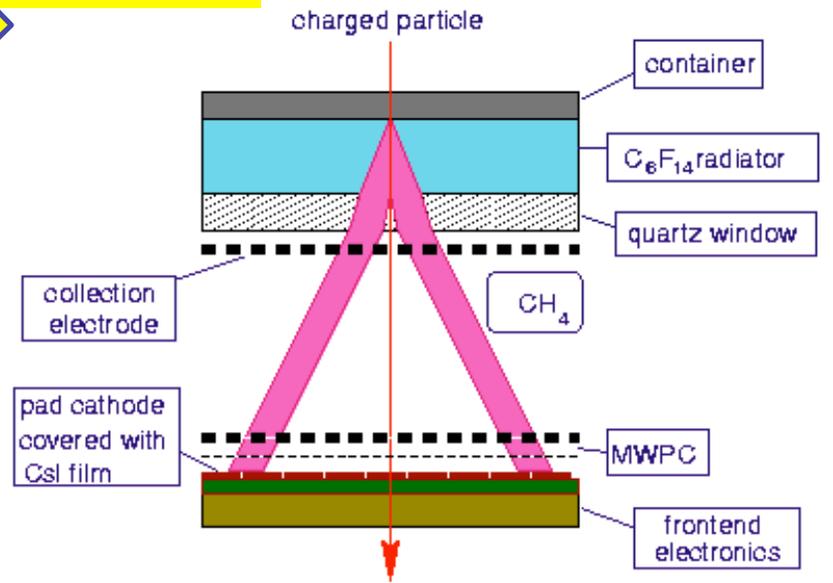
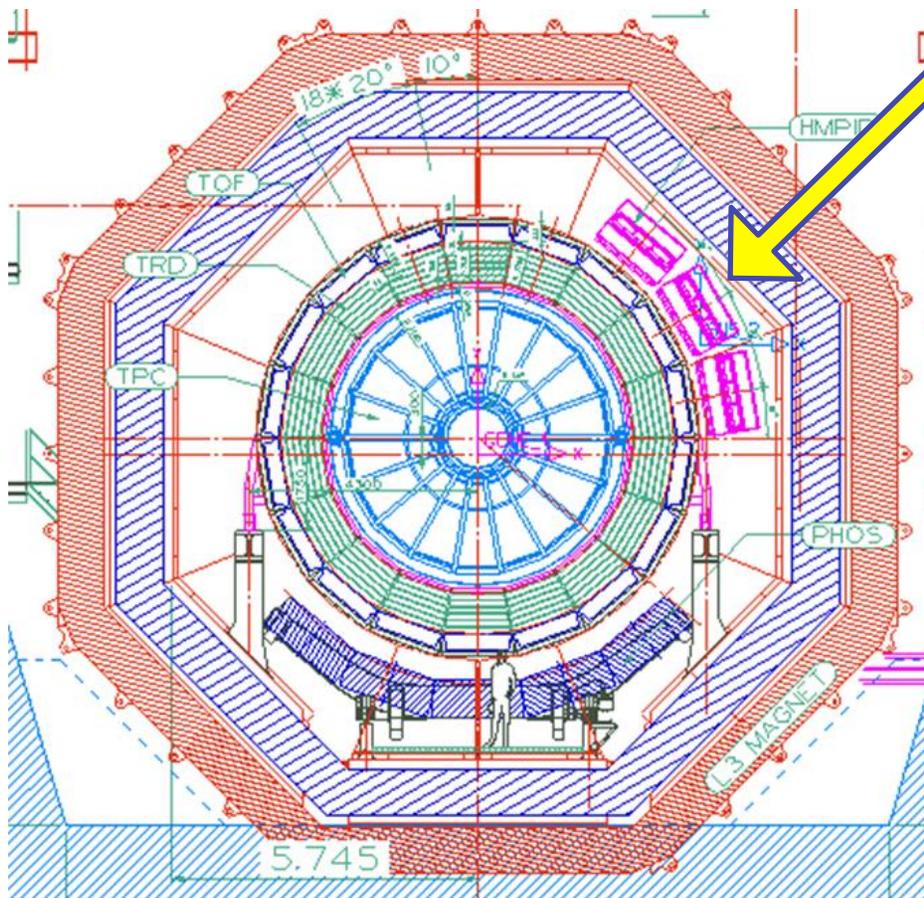
The HMPID RICH identifies hadrons $\pi/K/p$ in the range 1/3/5 GeV/c

7 modules of 1.5m x 1.5m (5% of barrel)



The RICH of ALICE

7 RICH modules
5 m from the collision



The RICH of ALICE : the resolution

Contributions to the angular resolution (per single photon and $\beta=1$, $\theta_C = 694$ mrad):

1. **Chromaticity** determined by the basic property of dispersion of radiator, convoluted with the media on the UV path and the QE of CsI
 $\rightarrow \sim 10$ mrad
2. **Spatial error** : determined by the granularity of the photodetector
 $\rightarrow \sim 5$ mrad
3. **Geometry**: determined by the length of the radiator and the proximity gap thickness
 $\rightarrow \sim 4$ mrad
4. **Track error**: negligible but requires investigation

Dominant contribution to sigma: chromaticity of the radiator.

$$N_{p.e.} \approx 18 \text{ for } \beta=1$$

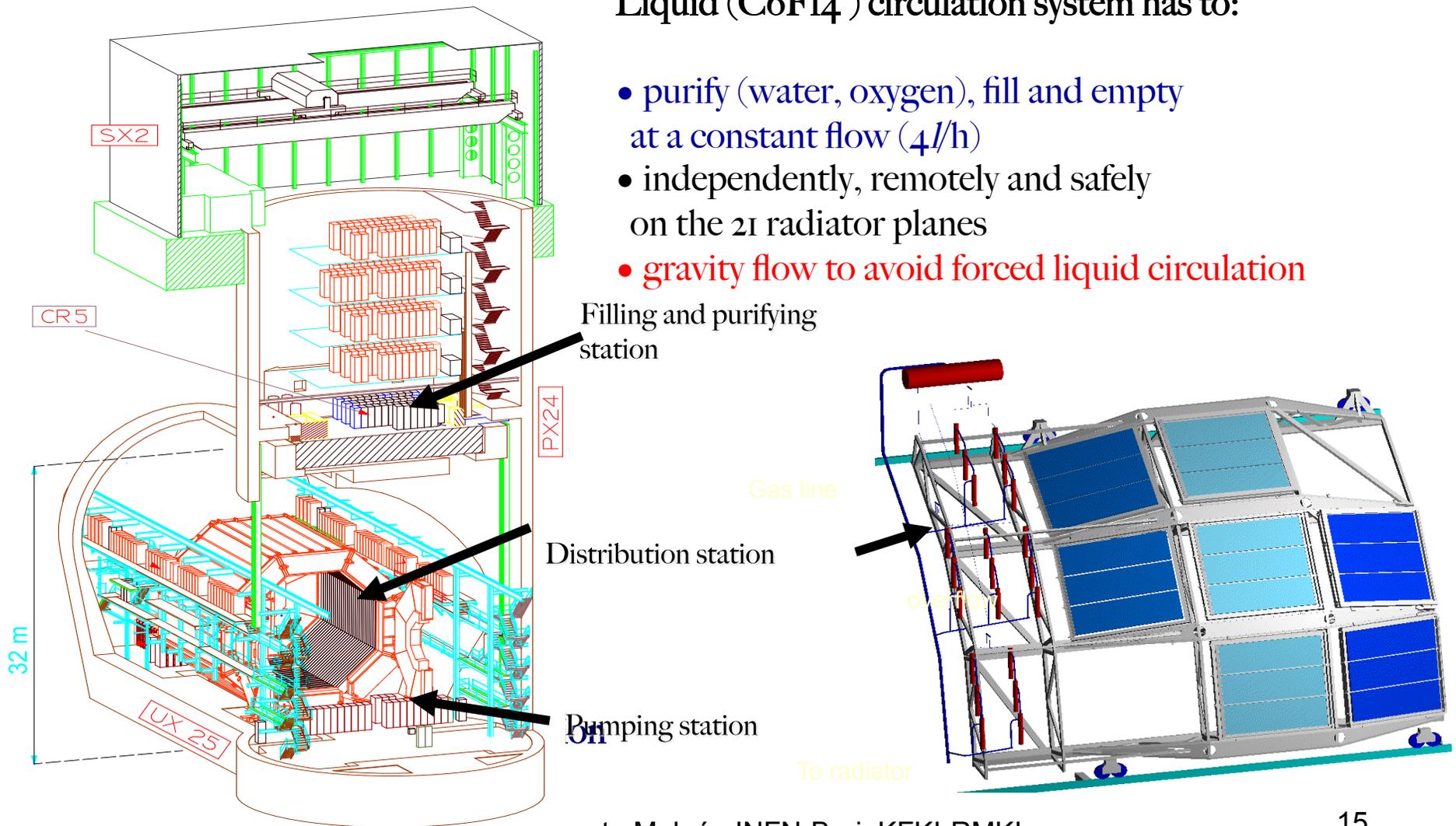
The RICH of ALICE

(taken from L. Molnar)

The C_6F_{14} circulation system

Liquid (C_6F_{14}) circulation system has to:

- purify (water, oxygen), fill and empty at a constant flow (4l/h)
- independently, remotely and safely on the 21 radiator planes
- gravity flow to avoid forced liquid circulation



The RICH of LHCb

*See detailed talks by C.Blanks, F.Muheim,
R.Young, A.Powell at this conference*

Physics aims:

separate $K/\pi/p$ in the range 2-100 GeV/c to
reconstruct rare (and less rare) B decays
(ex. $B \rightarrow KK$ and $K\pi$, $B \rightarrow D_s K$ and $D_s \pi$, ...)

Environment:

Works at hadronic machine (LHC), high particle density
Works at 1 MHz
Must reject pion better than at the percent level

Geometry:

focussed, 2 RICHes with 3 different radiators

The RICH of LHCb : the choices

Focussed geometry

Radiators: 5 cm aerogel $n = 1.03$ @ 400 nm
 95 cm C_4F_{10} $n=1.0014$ @ 400 nm
 180 cm CF_4 $n=1.0005$ @ 400 nm

Mirrors : 4 spherical ($f= 135$ cm)+ 16 plane ($R>600m$) in RICH1
 52 spherical ($f= 430$ cm)+ 40 plane ($R=80$ m) in RICH2

Photodetectors:

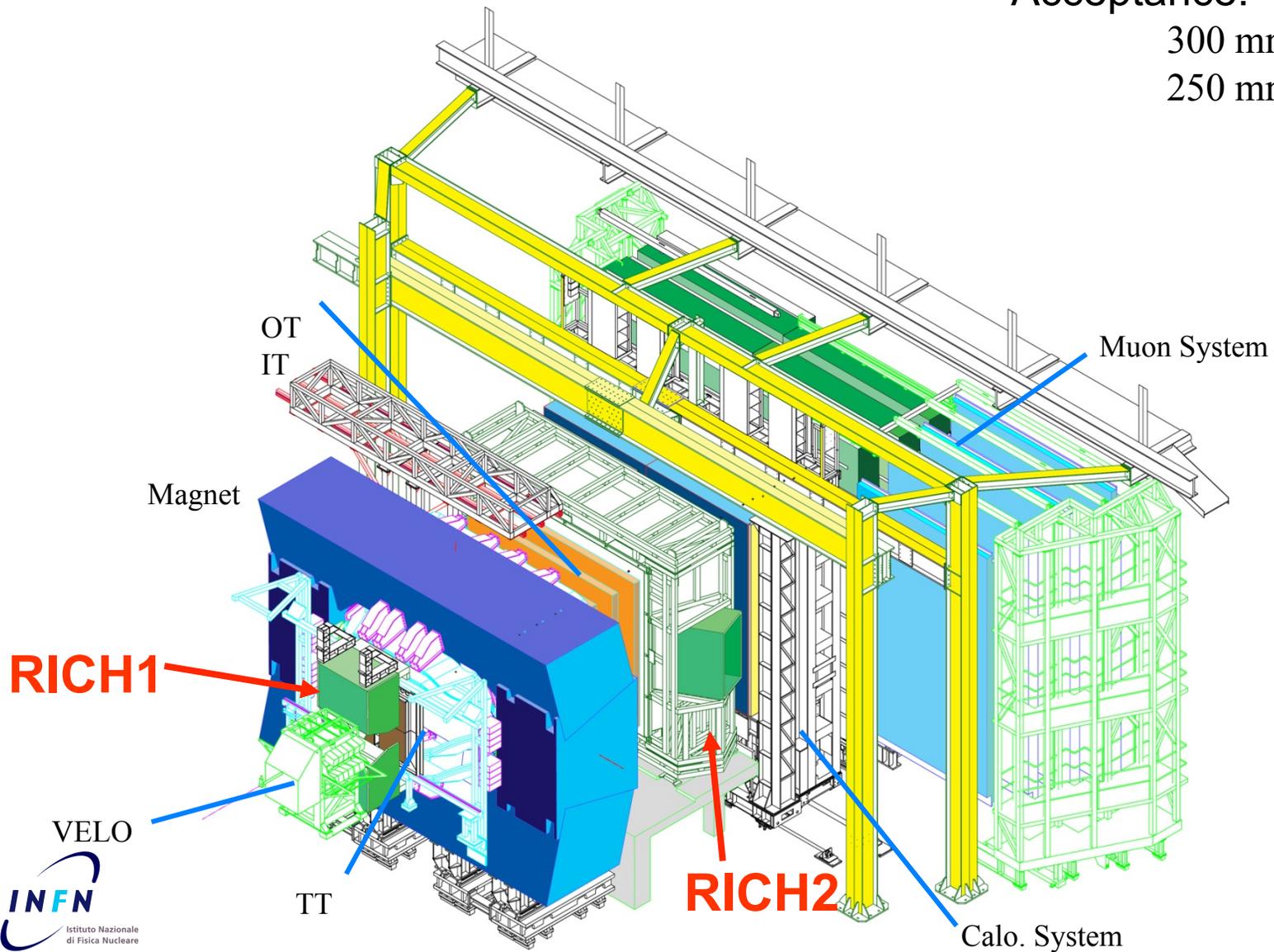
484 Hybrid Photon detectors (HPD) granularity 2.5 mm at the photocathode level

The RICH of LHCb

Acceptance:

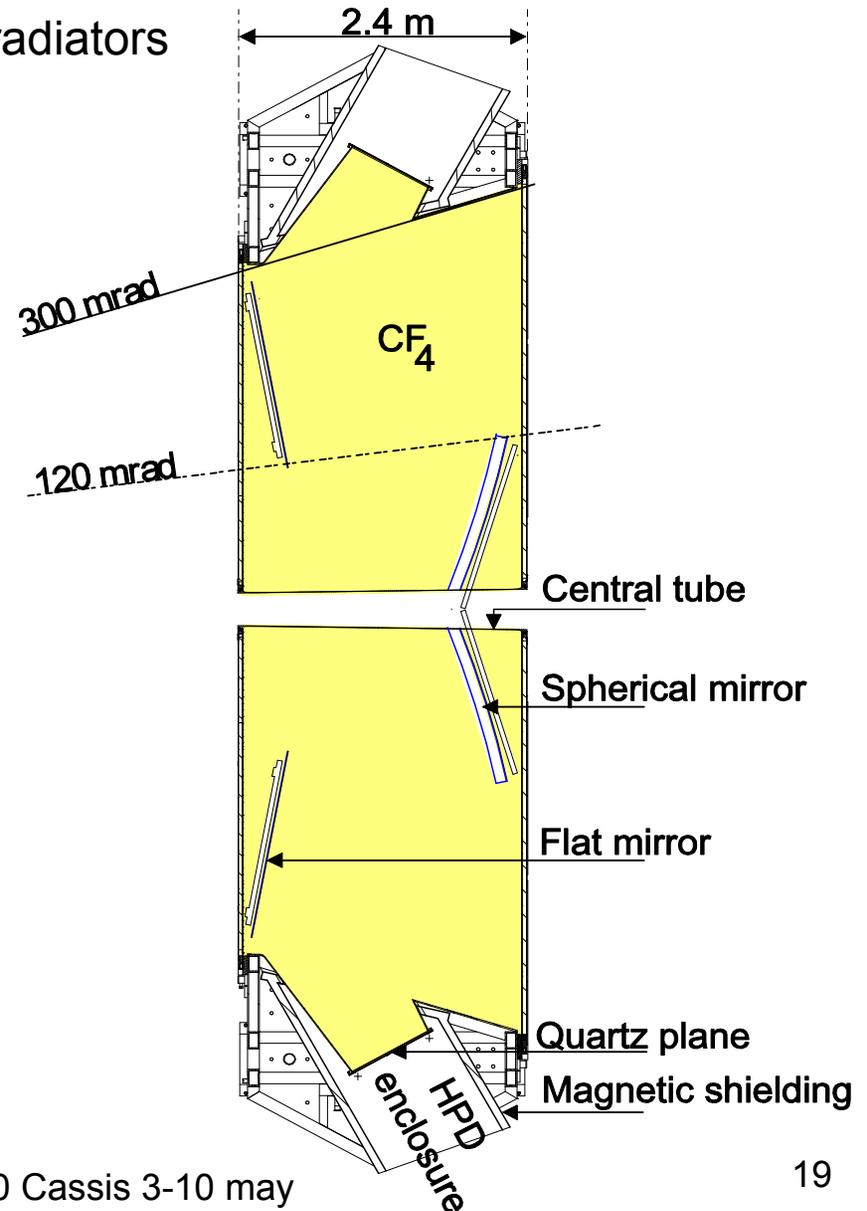
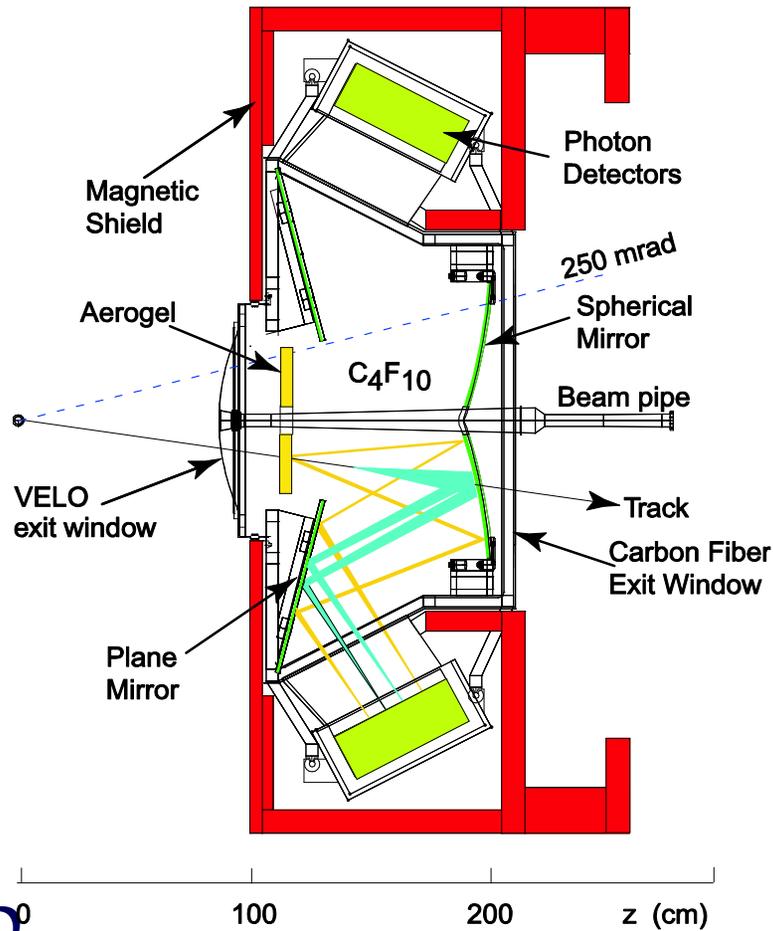
300 mrad horizontal

250 mrad vertical



The RICH of LHCb

The solution of LHCb: 2 RICHes with 3 radiators



The RICH of LHCb

RICH-1 vessel



RICH-2 vessel



The RICH of LHCb : the resolution

Needs a resolution In the range of O(1 mrad) , sub mrad in RICH2

Units : mrad

RICH-1 **RICH-2**

	Aerogel	C ₄ F ₁₀	CF ₄
Emission	0.4	0.8	0.2
Chromatic	2.1	0.9	0.5
HPD	0.5	0.6	0.2
Track	0.4	0.4	0.4
Total	2.6	1.5	0.7

Expected $N_{pe} \approx$ 6.5 30 22

The RICH of LHCb: the resolution

Needs to control:

Radiators:

Composition of gas radiators (some air, N₂, CO₂ contamination)

gas composition measured by chromatography to calibrate $n-1$

Control P and T continuously for correcting automatically the density ρ_{gas}

Geometry:

Mirror alignment with data. Down to 0.1 mrad

Spatial precision:

Monitor ageing of PD (HPD)

→ *see talk by R. Young on tuesday*

Corrections for magnetic distortion

→ *see poster by F. Xing*

Alignment of HPDs

Tracking:

must be well described by the Montecarlo. $\sigma(\theta_C)$ relies on track information also for alignment.

The RICH of LHCb : the resolution

The alignment of the mirrors is crucial → see talk by C. Blanks and MDMS corrections (poster by P. Xing)

Monitor on-line: from the behaviour of the hardware to the PID performance

After several millions of pp collision events :

	C4F10	aerogel	CF4
Achieved resolution	2.2	8.0	0.9
Expected resolution (from simulation)	1.5	2.6	0.7



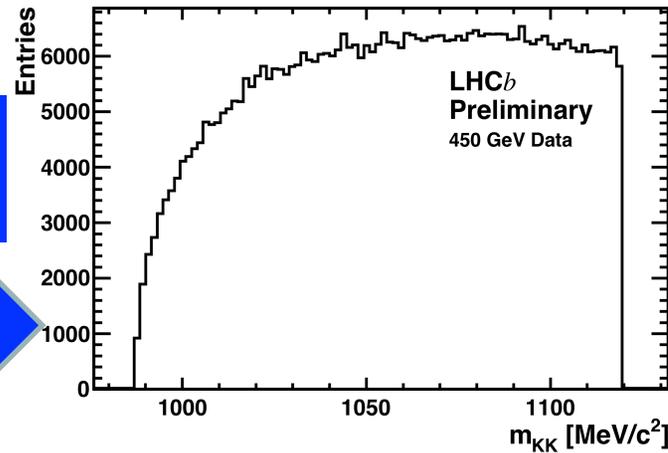
*Mirrors and HPD hit not yet aligned
(and C4F10 absorption has degraded $\sigma(\theta)$)*

What does the RICH of LHCb sees in the very first data?

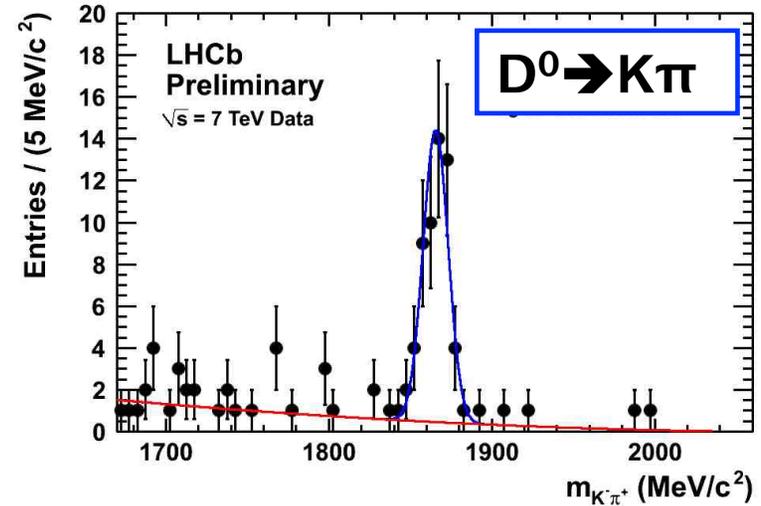
Observation of $\phi \rightarrow K^+K^-$

See talk by F. Muheim at this conference

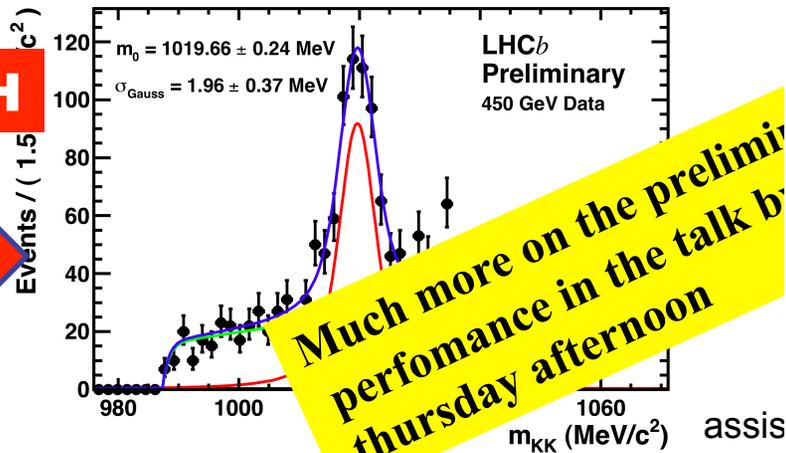
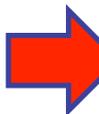
Only tracking



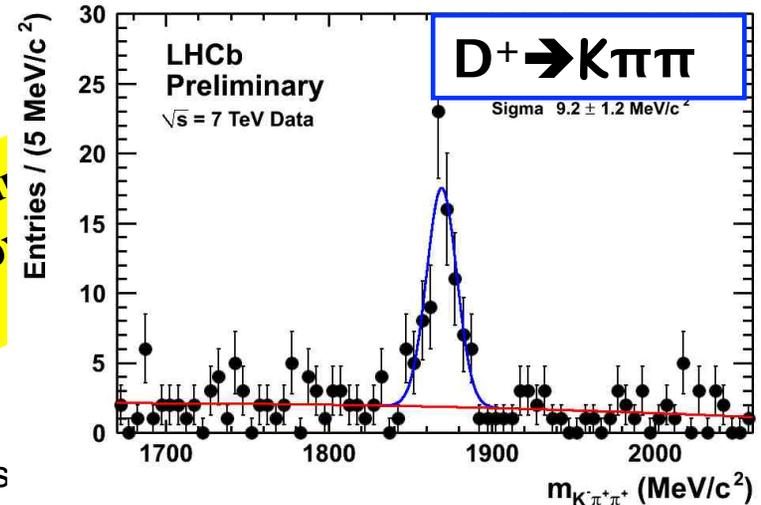
Observation of D^0 and D^+



With RICH



Much more on the preliminary performance in the talk by thursday afternoon



The RICH of NA62

See detailed talks by M. Lenti at this conference

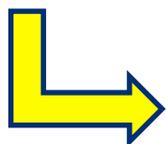
Physics aims:

measure $BR(K^+ \rightarrow \pi^+ \nu \nu)$ expected in the

Standard Model to be $O(10^{-11})$ at 10% precision

Present result: $1.73 (+1.15 -1.05) \times 10^{-10}$ (BNL E787/E949)

Dominant Background : $K^+ \rightarrow \mu^+ \nu$ ($K_{\mu 2}$ largest BR: 63.4%)



3σ π - μ separation (15-35 GeV/c)

Need $\sim 10^{-12}$ rejection factor of which from Particle ID: 10^{-2}
(Kinematics: 10^{-5} and Muon Veto: 10^{-5})

The RICH of NA62

See detailed talks by M. Lenti at this conference

Environment:

Kaon beam at 800 MHz

Needs to match a pion (10 MHz rate) with a kaon seen by the beam spectrometer (800 MHz rate)



measure the pion crossing time at 100 ps level

Geometry:

focussed

The RICH of NA62 : the choice

Based on SELEX RICH idea

Focussed geometry

Radiator:

17 m Neon ($n-1=62 \times 10^{-6}$ @ 300 nm) at 1 atm
 $\theta_C = 11.3$ mrad (π thresh.: 12 GeV/c)

Mirrors:

spherical (20 exagonal elements with 17 m focal length)

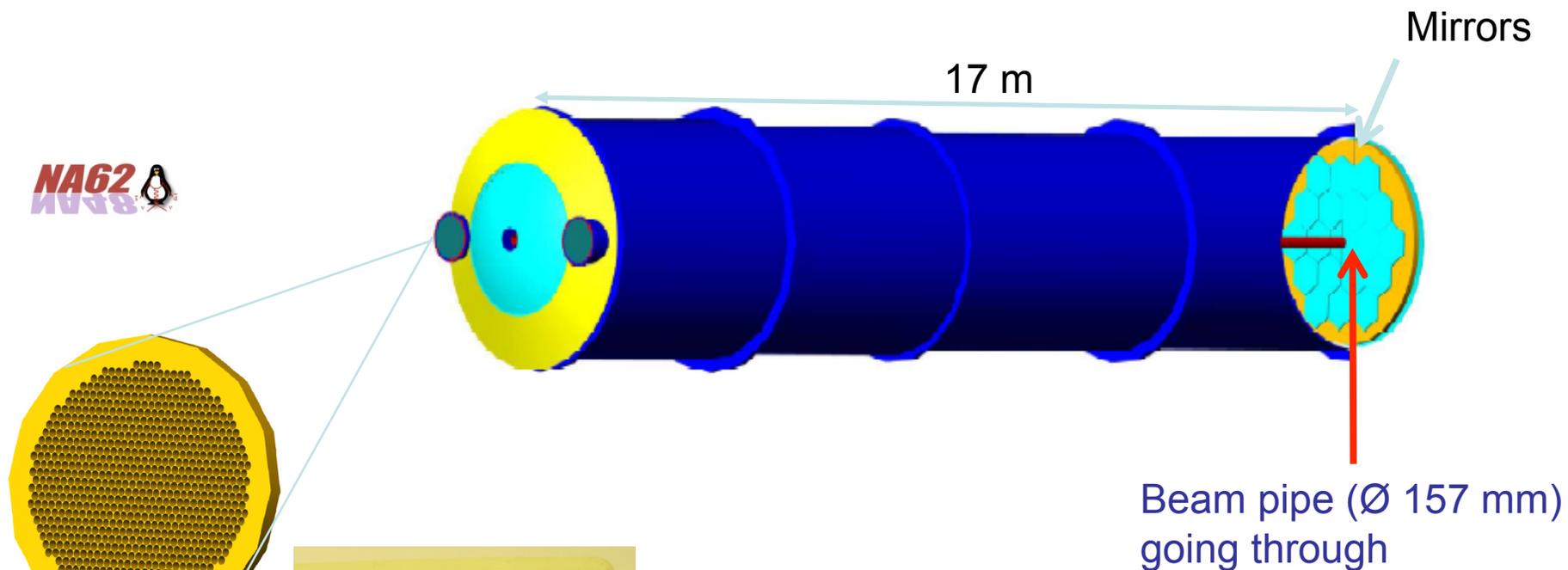
Photon detector:

2000 PMT Hamamatsu R7400-U03 with granularity 18 mm
and time resolution better than 100 ps.

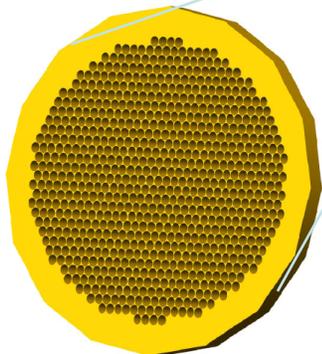
The RICH of NA62

Vessel volume: 200 m³ , 17 m long

(between straw tubes and liquid Kr calorimeter)



NA62



Hamamatsu R7400 U03

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The RICH of NA62

- vessel under construction (steel)
- max overpressure: 150 mbar
- 4 m wide (upstream), 3.4 m wide (downstream)
- thin aluminum entrance and exit windows

Contaminants < 1%

CO₂ used to purge the vessel

- The gas is then circulated in closed loop, and the Neon is introduced while absorbing the CO₂ in a molecular sieve filter.
- At the end the vessel is valve closed

The RICH of NA62 : the resolution

Contributions to the angular resolution (per single photon and $\beta=1$, $\theta_c = 11.3$ mrad):

1. Chromaticity determined by the basic property of dispersion of radiator,

$$\rightarrow \sim 125 \text{ } \mu\text{rad}$$

2. Spatial error : determined by the granularity of the photodetector

$$\rightarrow \sim 265 \text{ } \mu\text{rad}$$

3. Geometry: emission point, mirrors

$$\rightarrow \sim 15 \text{ } \mu\text{rad}$$

4. Track error: $\approx 55 \text{ } \mu\text{rad}$ (35 GeV/p)

The RICH of NA62

Needs to control :

The gas radiator

- monitor n through $n=1+(n_0-1) \rho/\rho_0$
with ρ is the gas density at operating conditions of T and P
Neon density stability $< 1\%$
- leak rate $< 1 \times 10^{-2}$ Std.cc/s
(if not achieved needs a purifier module)
- Contaminants $< 1\%$

Mirror alignment is important : with data and with laser to a level of $O(50 \mu\text{rad})$

Photocathode QE

RICHes IN SPACE EXPERIMENTS

In space: stability is mandatory (essentially no maintenance).
Solid radiators are more suitable.

Proximity focus (no optical element to align etc.,)
What could change: optical quality of the radiator,
QE of photocathode to count photons.

The RICH of AMS

See detailed talks by R.Pereira at this conference

Physics aims:

Cosmic ray spectrum, search for antimatter and dark matter.
Must measure particle velocity β and charge

Environment:

Operates in space (on satellite) for a period of at least 3 years

Geometry:

proximity focus

The RICH of AMS : the choice

Proximity focussed

Solid radiators :

2.5 cm aerogel $n=1.05$

0.5 cm NaF (sodium fluoride) crystal $n=1.334$

Conical reflector around

Photodetectors:

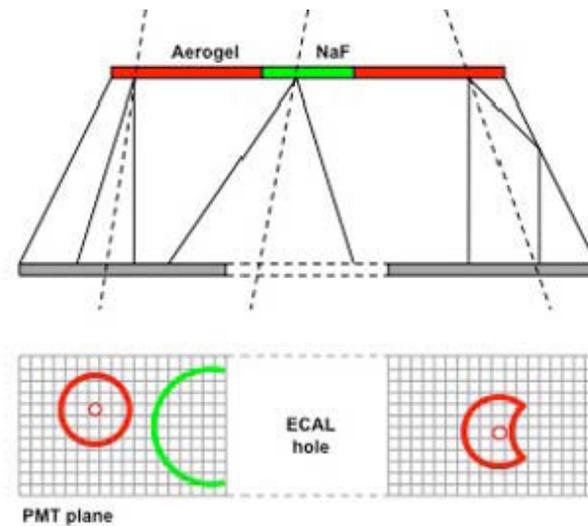
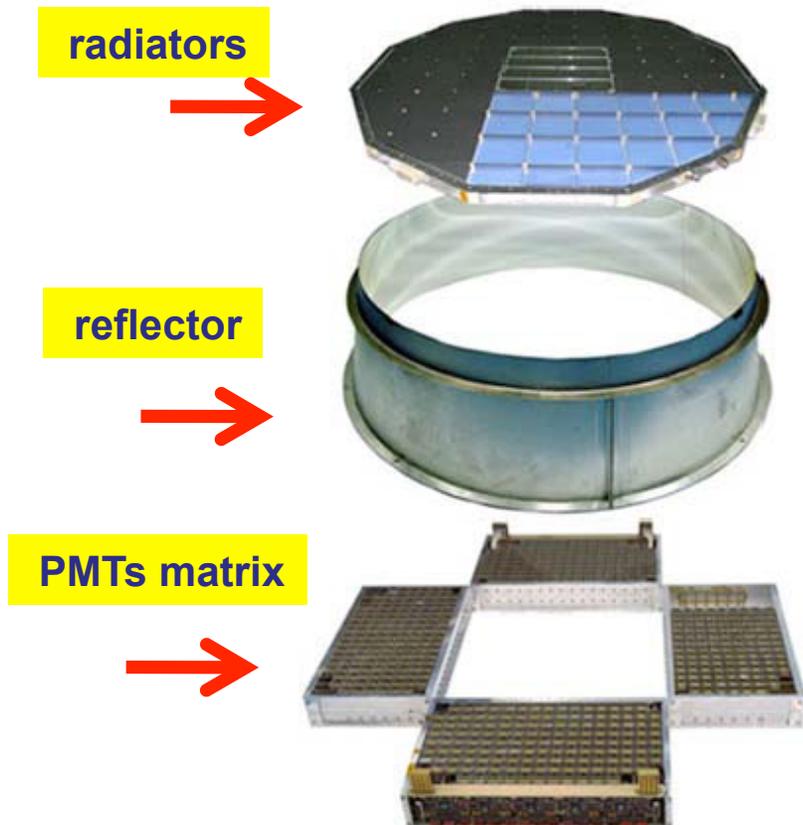
680 PMT Hamamatsu R7600-M16 with plastic light guide

Pixel size : 8.5 mm

Each PMT individually shielded from the stray field

(up to 300 Gauss)

The RICH of AMS



Scheme of the radiators and ring images

The RICH of AMS: the velocity resolution

Aim: must measure β with $\sigma(\beta)/\beta \sim 0.1\%$ for charge 1

Velocity measured from $\beta = 1/n \cos \theta_C$

$$\text{with } \sigma(\beta)/\beta = \tan\theta_C \frac{\sigma(\theta_C)}{\sqrt{N_{pe}}}$$

Contributions to the resolution:

Radiator chromaticity

Radiator thickness

Pixel size (8.5 mm)

The RICH of AMS: the velocity resolution

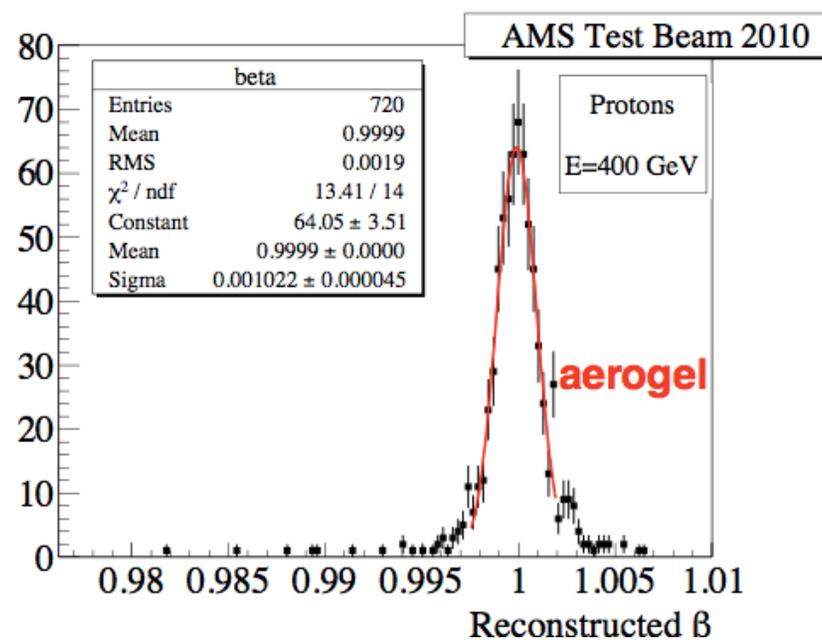
Contributions to the resolution (Units: mrad)

	Aerogel	NaF
Radiator chromaticity	3.2	4.8
Radiator thickness	3.3	0.3
Pixel size (8.5 mm)	4.6	0.6
	6.5	4.8
 $\sigma(\theta_c)$		
$\sigma(\beta)/\beta \approx$	2×10^{-3}	4×10^{-3}

Possible degradation from natural ageing of aerogel ?

The RICH of AMS: the velocity resolution

Test beam measurement in 2010 with 400 GeV protons



The RICH of AMS: the charge resolution

Aim: must measure Z (also measured with TOF, dE/dx in Si tracker)
with $\Delta Z = 0.2$ for electric charge

➔ Charge measured by $Z^2 \propto (N_{pe}/\epsilon) 1/\sin^2 \theta_C$

ϵ = acceptance and photon detection efficiency

Contributions to the resolution ΔZ :

Statistical error on N_{pe}

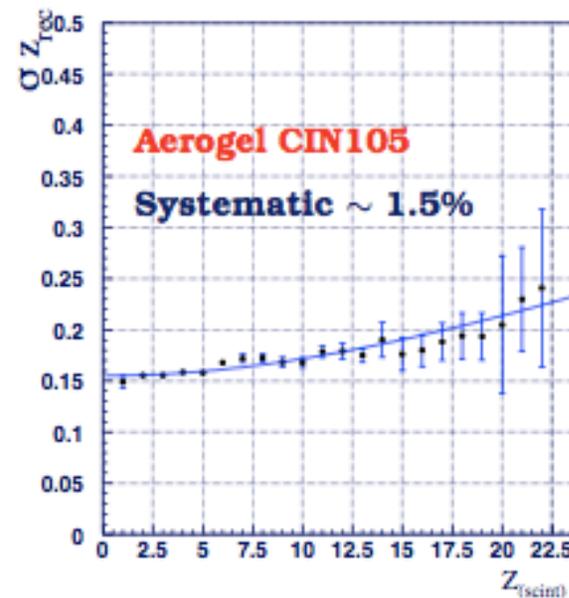
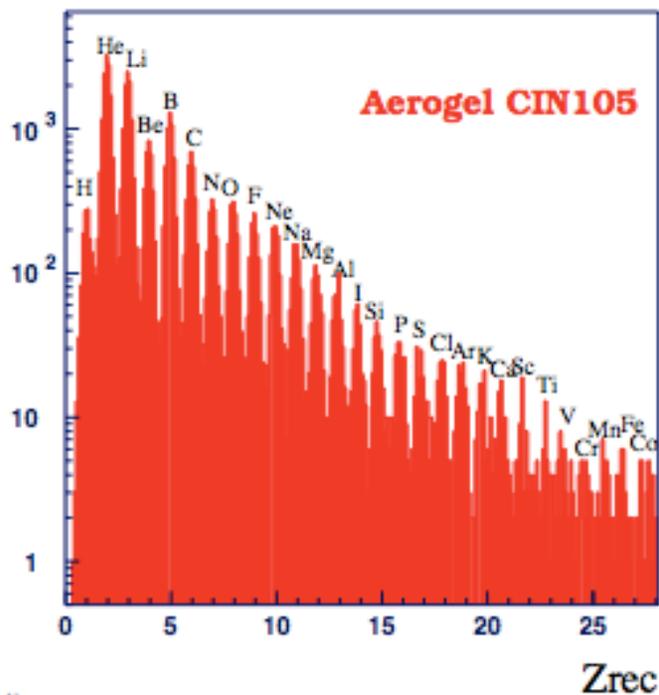
Systematics from non-uniformity of

- radiator (n , thickness, clarity,...)
- photon detection (PMT, temperature effects,...)

The RICH of AMS: the charge resolution

$$\Delta Z = \frac{1}{2} \sqrt{\frac{1 + \sigma_{p.e}^2}{N_0} + Z^2 \left(\frac{\Delta E}{E}\right)^2}$$

Results from test beam 2003 with fragmented ions :



A good precedent : **the RICH of CREAM**

A collaboration of
US, Korea, Italy, Mexico, France, NASA

4 successful flights. Launched from US McMurdo base in Antarctica

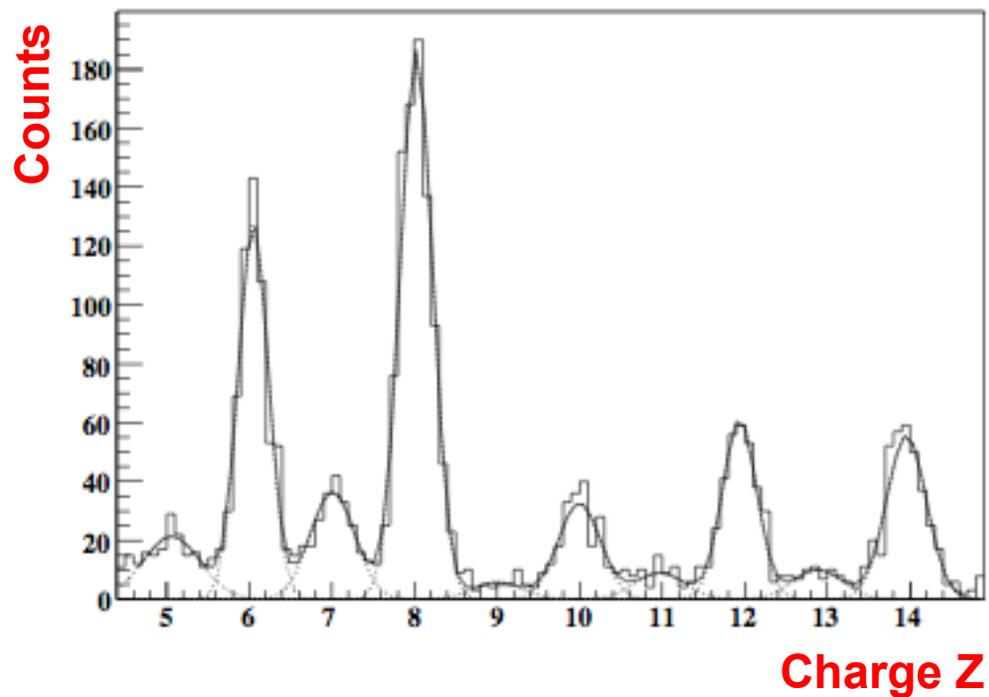
The RICH was proposed, designed, built
in less than 2 years
by a **Mexican-French** collaboration.

200 Aerogel tiles + 1600 PMT Photonis XP1232

Measure charge from $N_{ph} \propto \sin^2\theta_C Z^2$ with similar systematics requirements as in AMS (uniformity of thickness, optical index dispersion of aerogel tiles and in each tile,....)

A good precedent : **the RICH of CREAM**

Measurement of **charge** by CREAM during the second balloon flight



Concluding comments

- ★ RICH technique is extremely powerful and widely used for PID in different environments
- ★ Choices of technologies make flexible RICH designs for different applications. Stability is often to be favoured.
- ★ Technological developments in Photodetectors sector will even improve performance (ex. high time resolution, high QE)
- ★ BUT: RICH detectors are in general sophisticated tools and need important effort to keep under control the different components of the Čerenkov angle resolution
- ★ And, not least, powerful software tools are mandatory to translate the detector response into physics measurements.



W. Kandinski

Ring imaging can be a piece of art