

# (SPMT<sup>+</sup>ABC intro)

JUNO-ABC evaluation  
CENBG (Bordeaux)

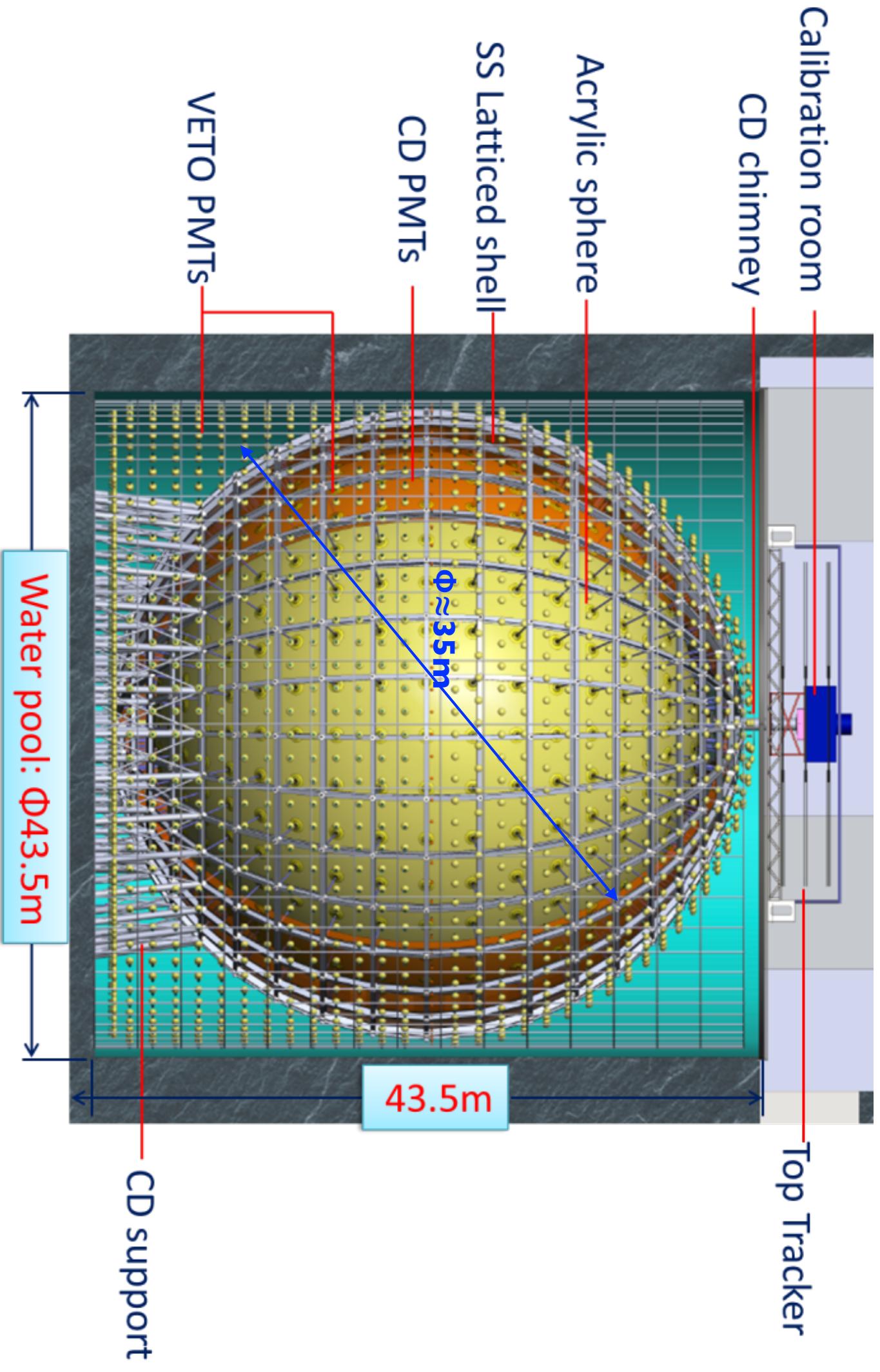
Anatael Cabrera  
on behalf of the SPMT group

CNRS / IN2P3 @ APC (Paris) — LNCA (Chooz)

# (fast physics recap)

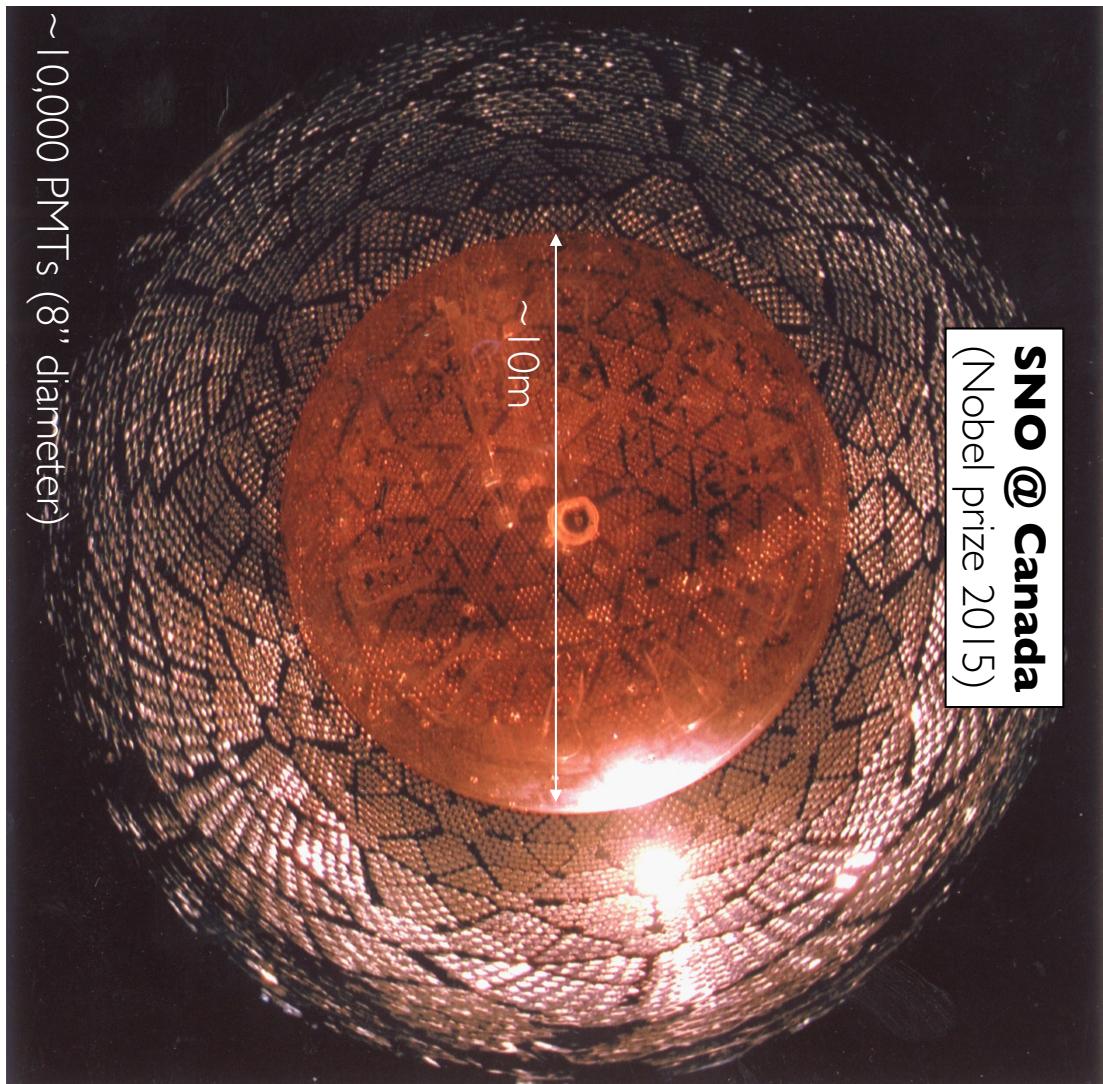
What/Why the SPMT?...

# the JUNO detector...



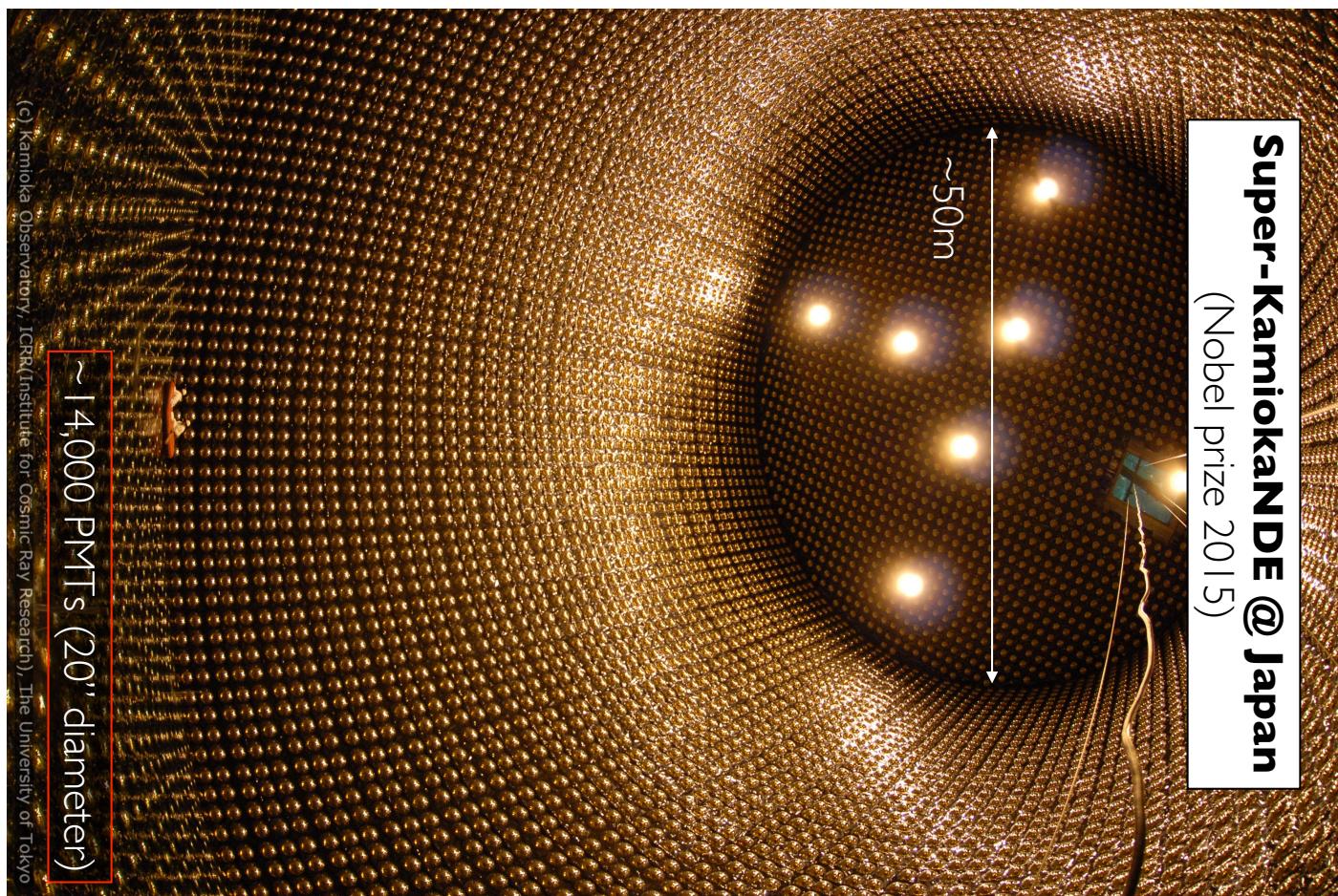
# the JUNO detector (predecessors)...

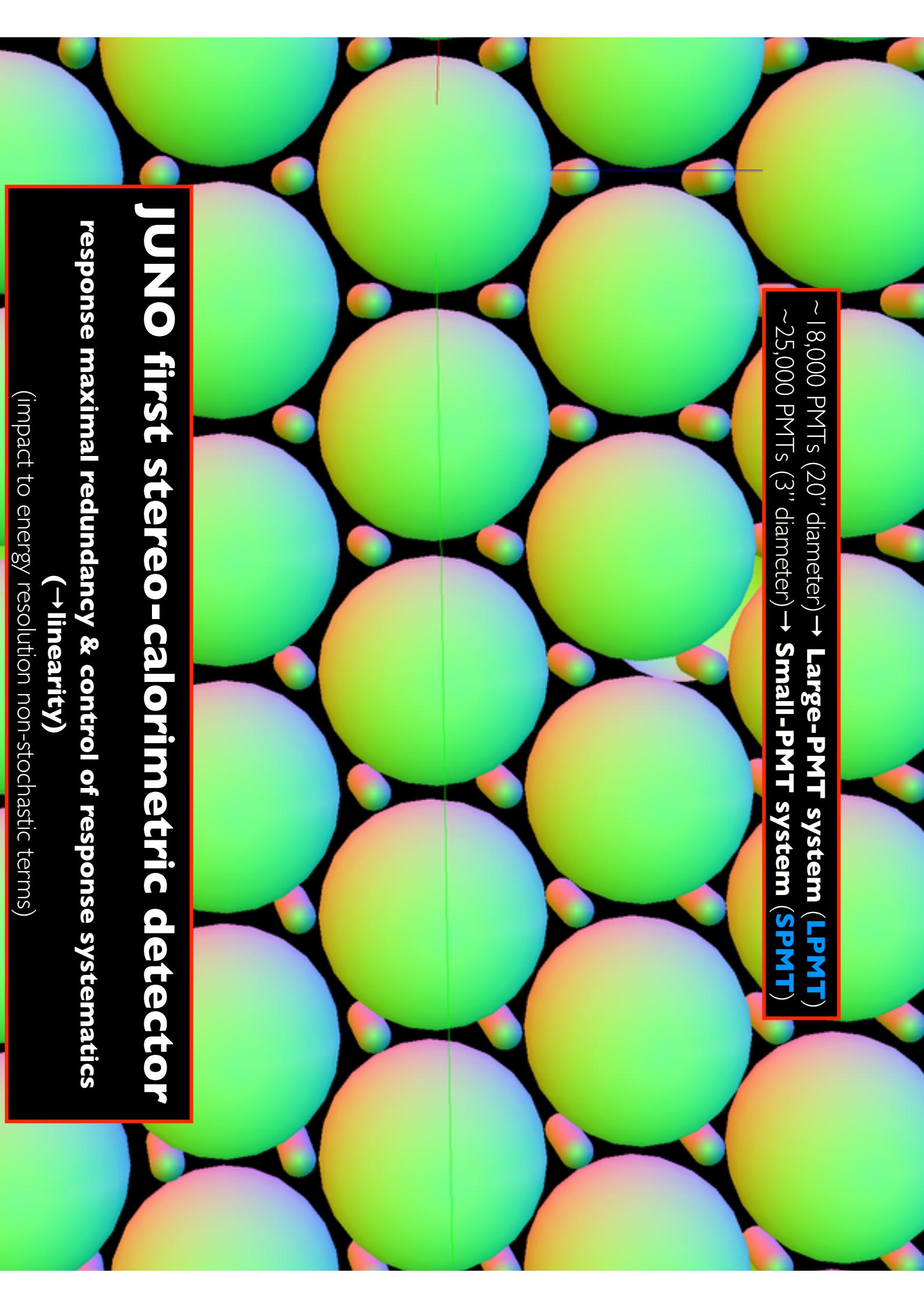
**SNO @ Canada**  
(Nobel prize 2015)



**JUNO can be regarded as a hybrid of both...**  
(filled with liquid-scintillator → **~100x more light**)

**Super-KamiokaND<sup>E</sup> @ Japan**  
(Nobel prize 2015)





$\sim 18,000$  PMTs (20" diameter) → **Large-PMT system (LPMT)**  
 $\sim 25,000$  PMTs (3" diameter) → **Small-PMT system (SPMT)**

# JUNO first stereo-calorimetric detector

response maximal **redundancy & control of response systematics**

(→ **linearity**)

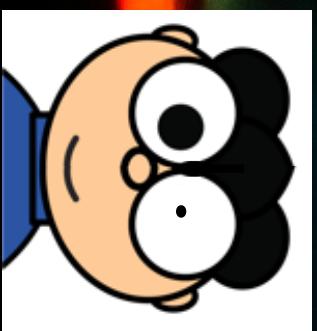
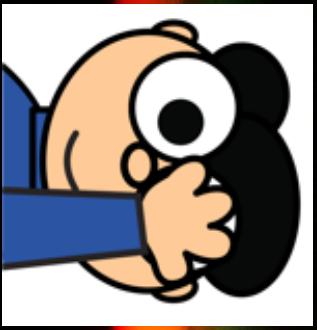
(impact to energy resolution non-stochastic terms)

JUNO upgrade...

# JUNO (before)

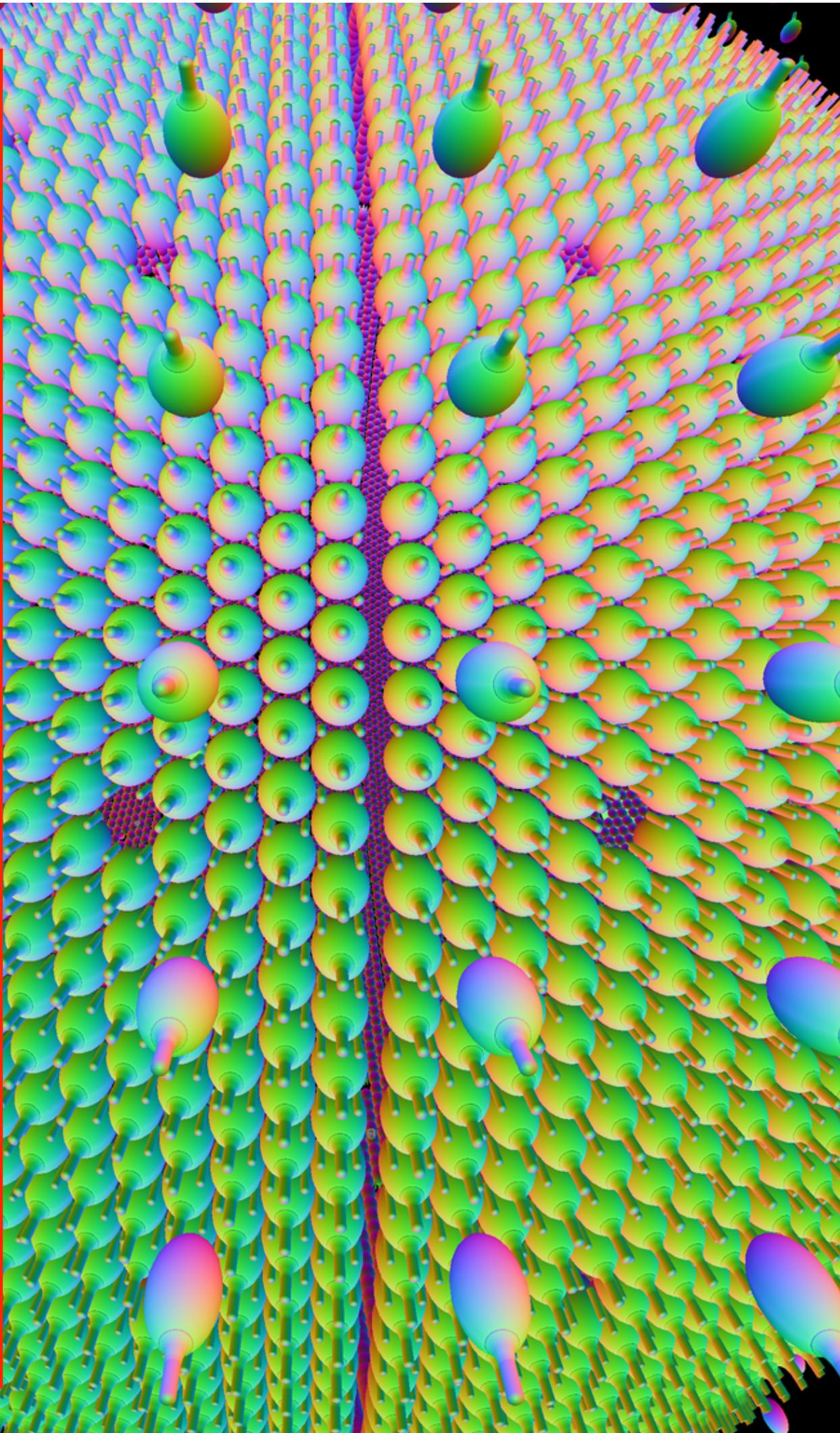
# JUNO (now)

single-calorimetric  
stereo-calorimetric



# JUNO: a photo-cathode collos

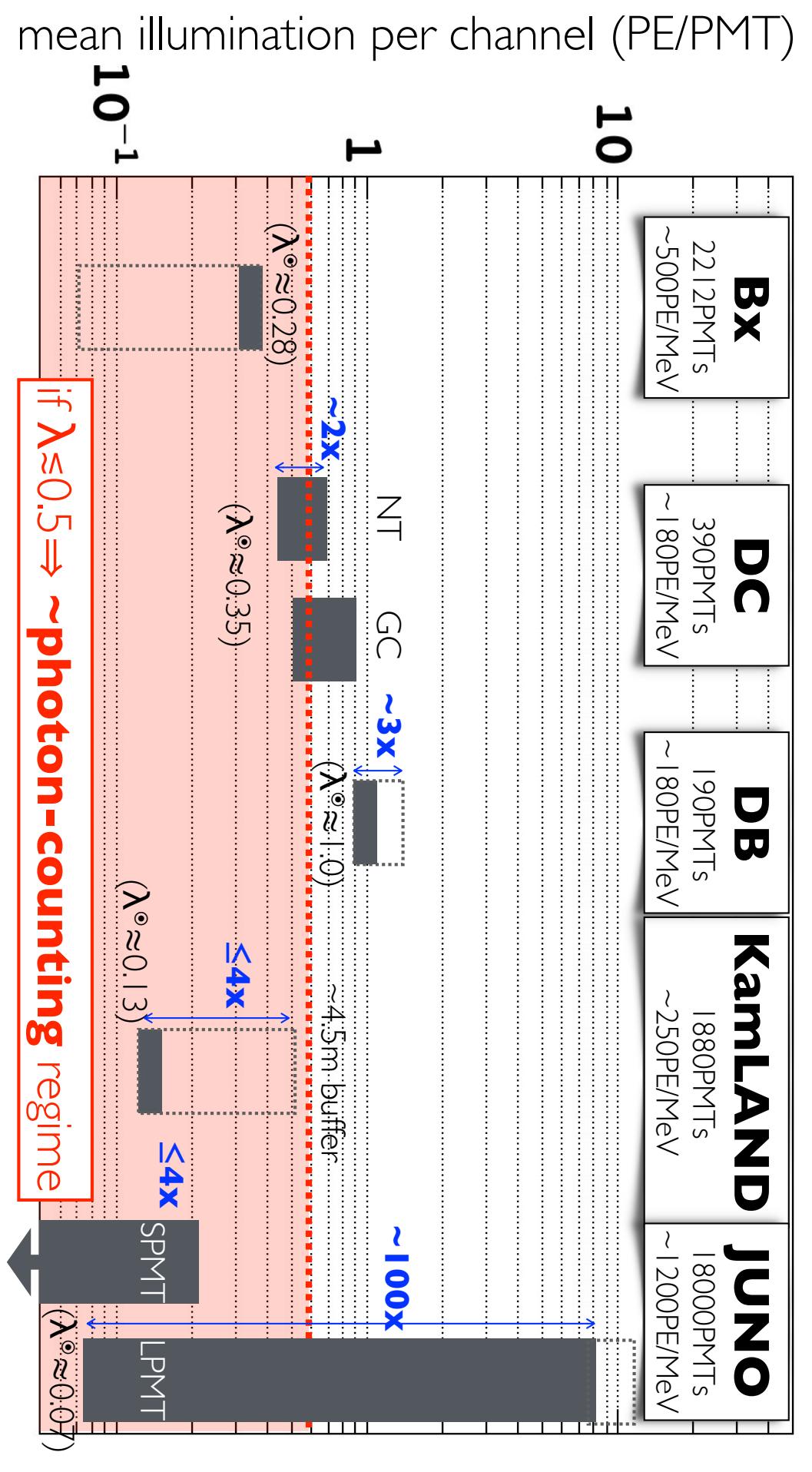
**highest light level ever made ( $\geq 1000$  PE/MeV)**  
**highest “calorimetry” control of systematics**



# the JUNO challenge...

@ **IMeV**

$\lambda^\odot = \text{mean illumination per channel @ center}$

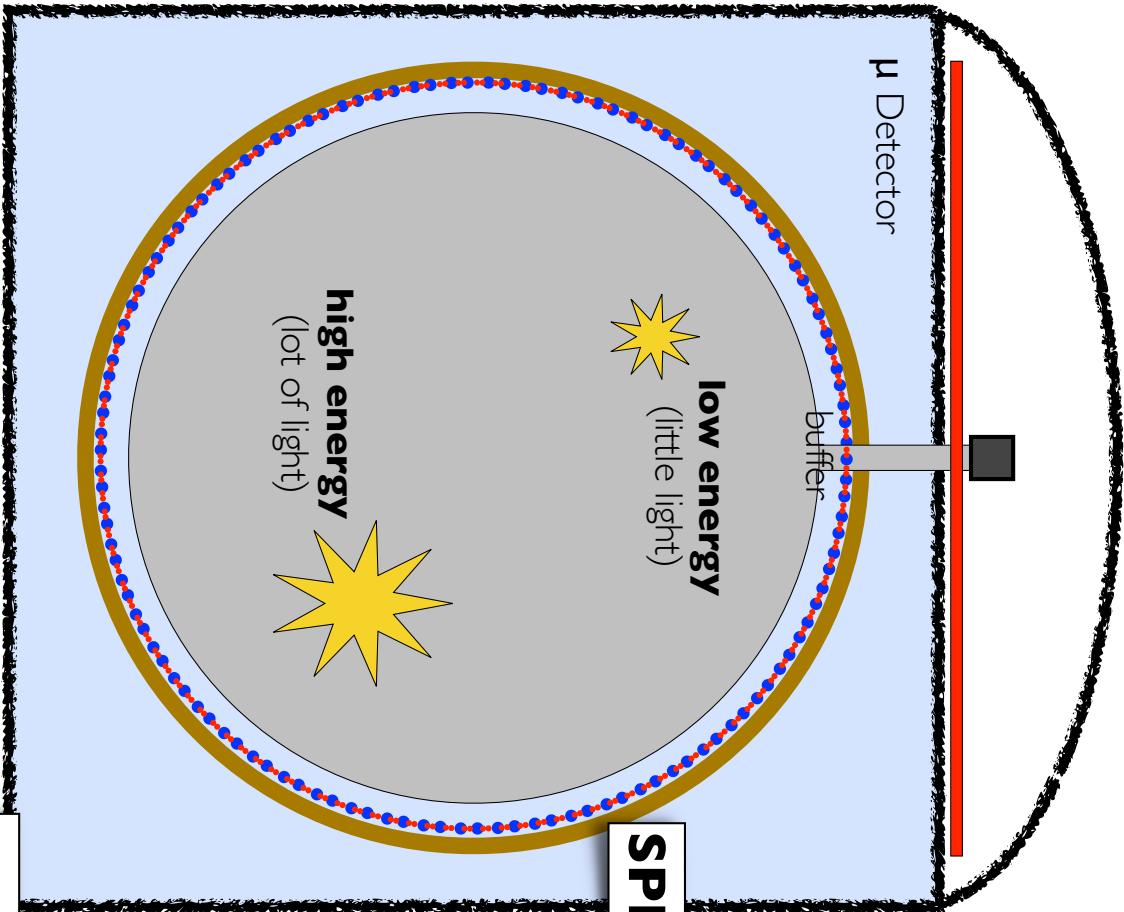


**HIGHEST** precision calorimetry ( $\leq 3\%$  @  $1\text{ MeV}$ )

$\oplus$

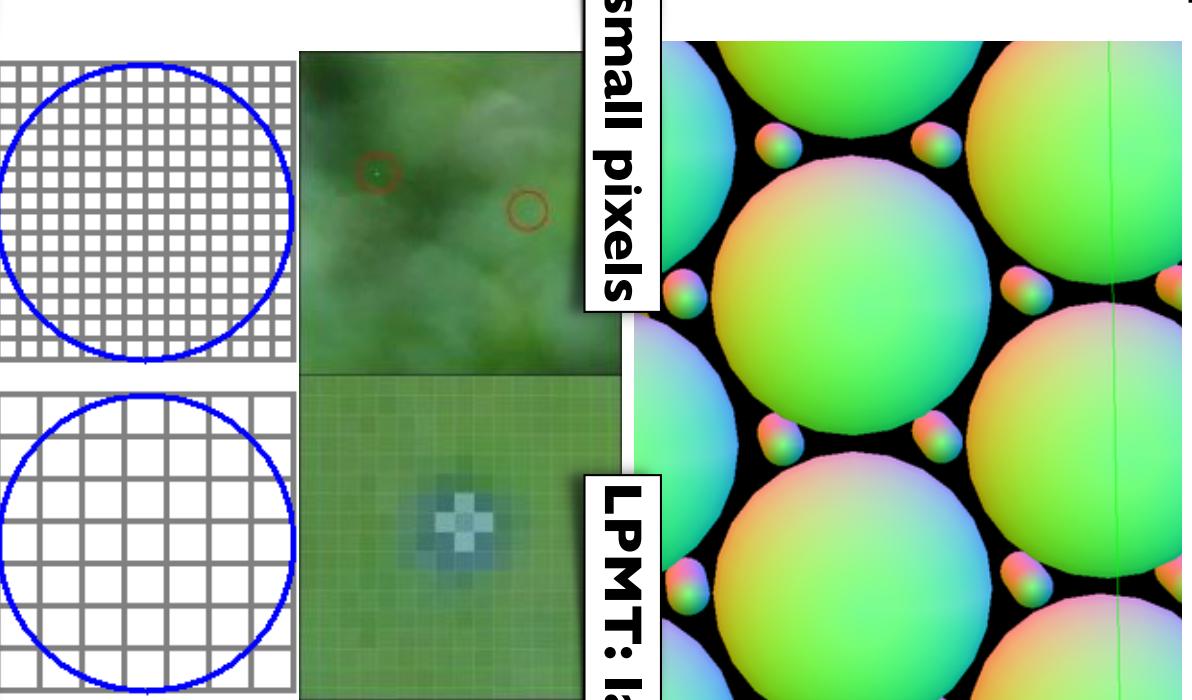
**LARGEST** dynamic range in calorimetry (channel-wise) [ $\Rightarrow$  **uniformity** $\oplus$ **linearity** $\oplus$ **stability**]

# the competition SPMT vs LPMT ...



**SPMT: small pixels**

**LPMT: large pixels**



**one photon per pixel**  
 $\Rightarrow$  “0” dynamic range  
**[no non-linearity]**

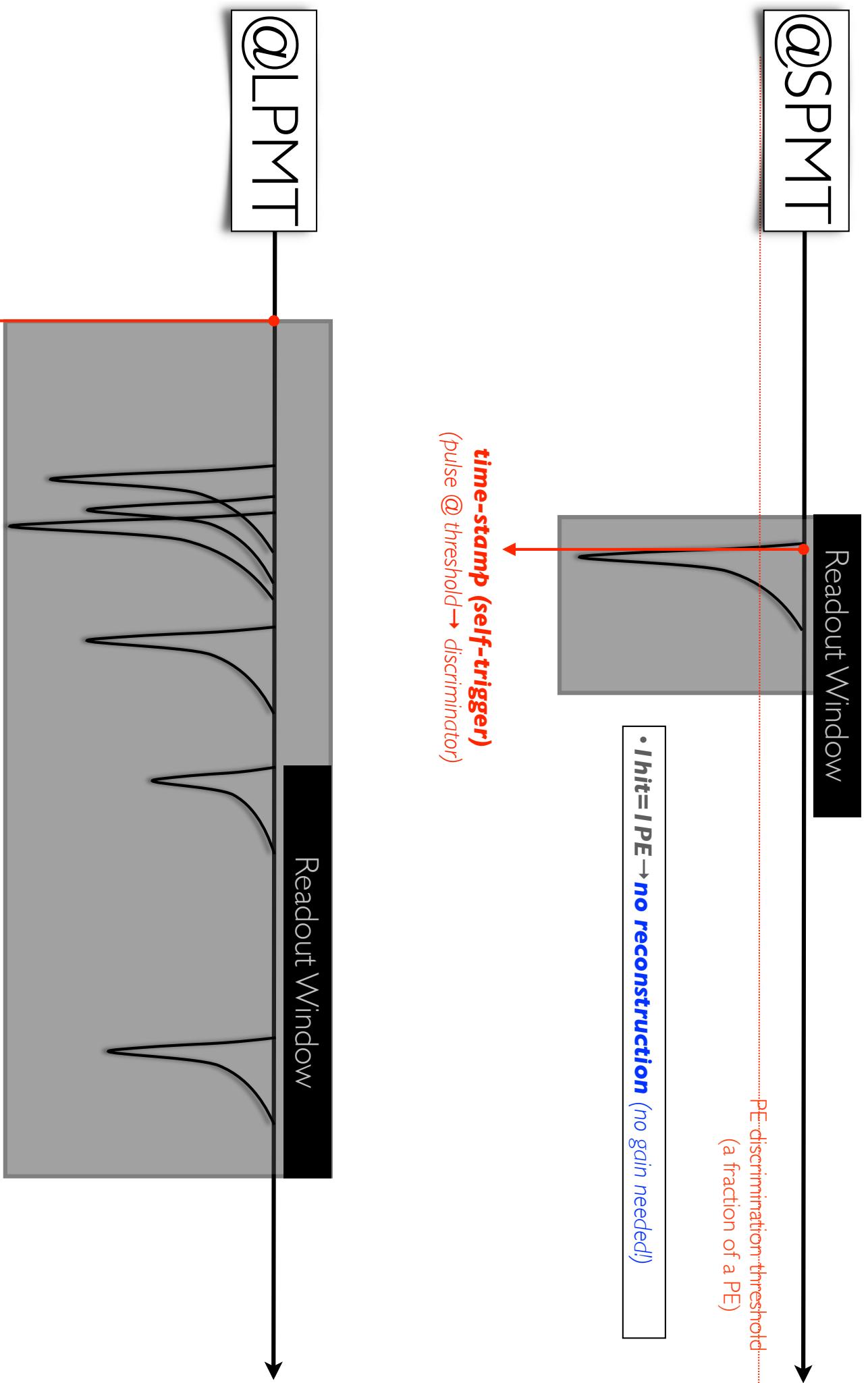
**many photons per pixel**  
 $\Rightarrow$  large dynamic range!  
**[non-linearity control]**

SPMT regime...

# Photon Counting (Photon-Statistics)

$$I_{\text{hit}} = |P_E|$$

# Photon-Counting vs Charge-Integration...



PE discrimination threshold  
(a fraction of a PE)

Readout Window

@SPMT

Readout Window

@LPMT

**time-stamp (global trigger)**

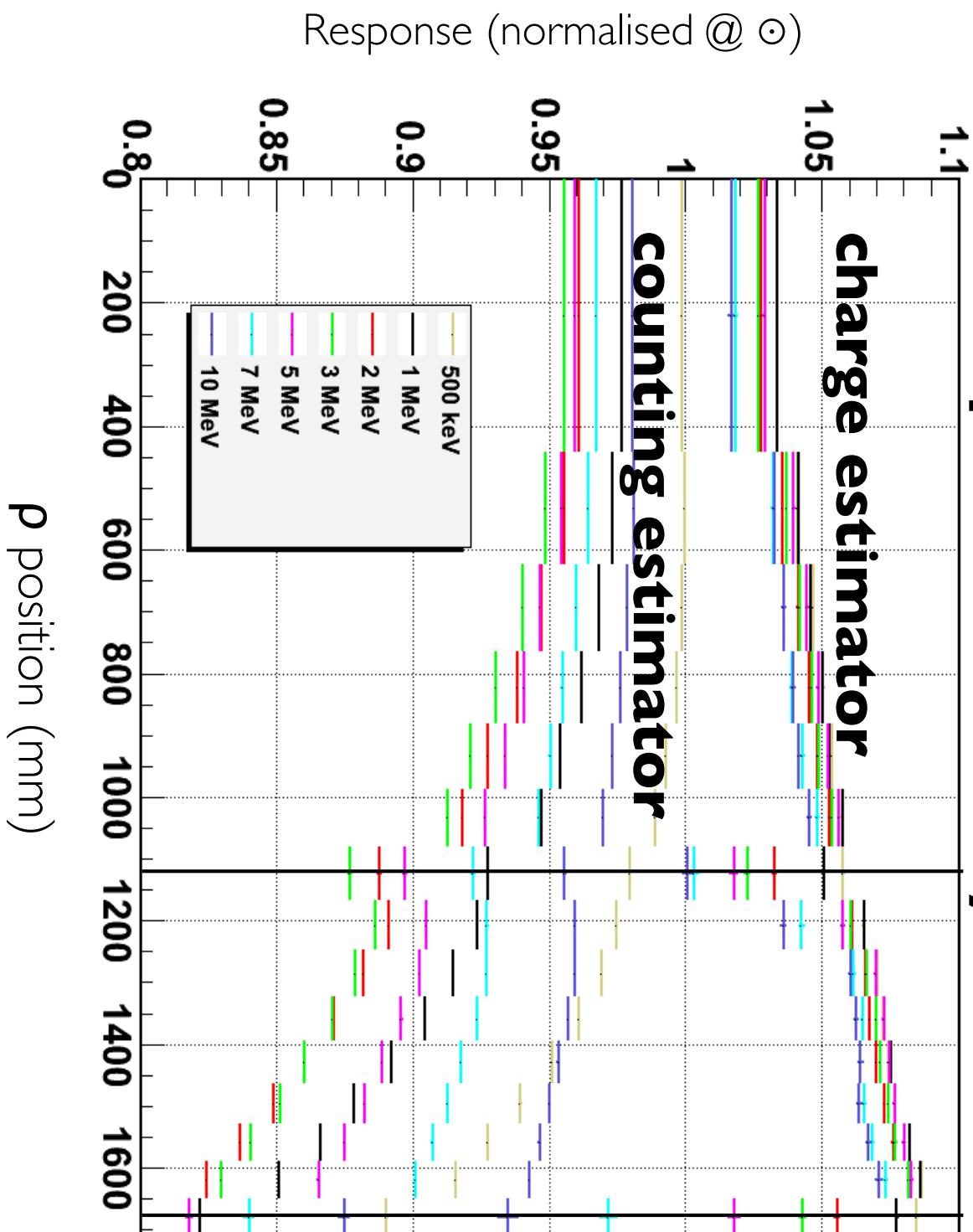
(readout-window sampled, example @ start)

**I<sub>hit</sub> ≠ I<sub>PE</sub> → reconstruction is a must! [QI-reco]**

**I<sub>hit</sub> = I<sub>PE</sub> → no reconstruction (no gain needed!)**

# 12 Photon Counting vs Charge Integration in action...

## response uniformity



**charge & counting both useful → highly complementary!!**

# the cooperation SPMT $\oplus$ LPMT ...

**SPMT**

one photon per pixel  
 $\Rightarrow$  “0” dynamic range

[no non-linearity]

**LPMT**

many photons per pixel  
 $\Rightarrow$  large dynamic range!

[large non-linearity]

(key point)

**BOTH (SPMT and LPMT) see the SAME events**

**BOTH see the SAME ENERGY but estimated differently!**

(if one was non-linear, the other will see it in relative)

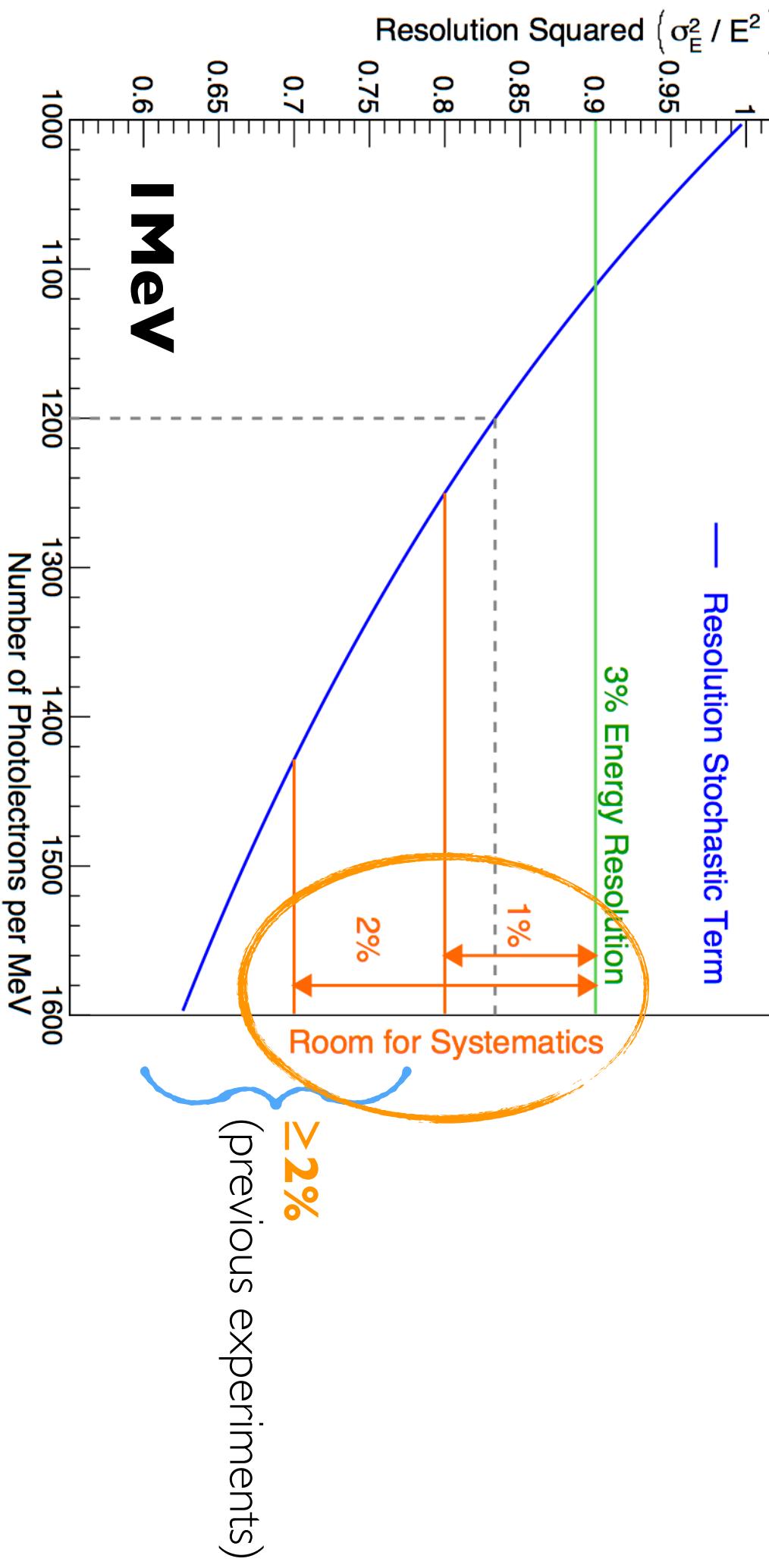
SPMT $\oplus$ LPMT (together) = precision!



# JUNO calorimetry condition...

**lot of light** is a necessary but not sufficient condition

$$\frac{\sigma(E)}{E} = \sqrt{\frac{\sigma_{\text{Stoch}}^2}{E} + \sigma_{\text{Non-Stoch}}^2(E)} \leq 3\% @ 1\text{ MeV}$$



## challenging calorimetry systematics control

# LPM<sup>T</sup> + SPMT complementarity . . .

## LPM<sup>T</sup>

- linear (sensitive to tiny deviations)
- uniform (increase away from centre)
- stability (sensitive to both gain & zeroes)

## SPMT

- non-linear (robust due to Poisson)
- uniform (decrease away from centre)
- stability (sensitive to gain (less) but zeroes)

## LPM<sup>T</sup>⊕SPMT: differences use to aide LPM<sup>T</sup> response control

(correct systematics beyond stability⊕linearity⊕uniformity)

LPM's view...

lot of them but hard to control!!

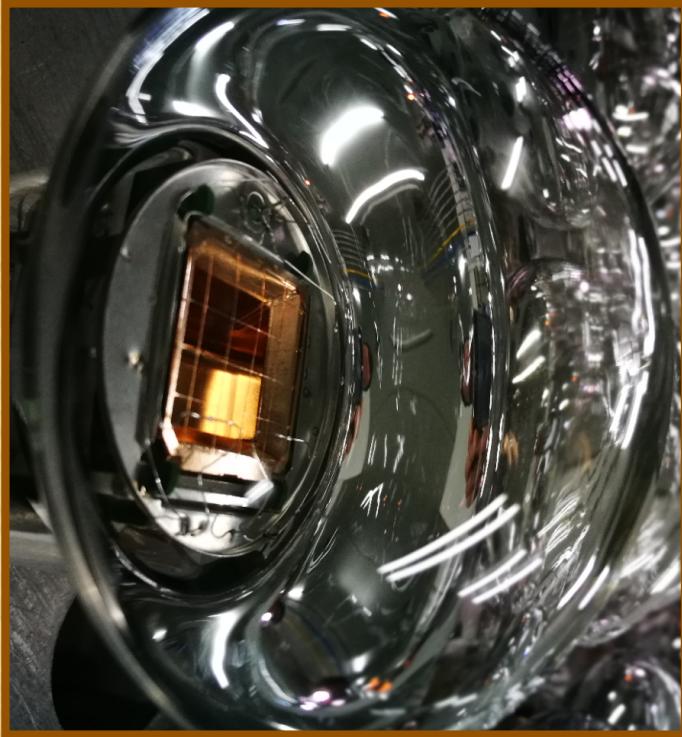


now precision physics!



**REWARD**

much neutrino physics ⊕ calorimetry!!



(don't break them!)

~~REWARD~~ **REWARD**

rich physics...

- **high precision stereo calorimetry ( $\text{LPMT}^\oplus \text{SPMT}$ )**
- **high precision neutrino oscillation physics, etc**
- **high rate supernova detection (acceptance  $\geq 90\%$  Galaxy)**
- **4 $\pi$   $\mu$ -tracking & after muon spallation products (n-activation)**

and more... (skip here)

# physics → detector requirements..

- **high precision stereo calorimetry (LPMT<sup>⊕</sup>SPMT)**

**constraint: [1,10]MeV (photon-statistics), low rate, charge resolution**

- **high precision neutrino oscillation physics, etc**

**constraint: same as stereo-calorimetry (byproduct)**

- **high rate supernova detection (acceptance  $\geq 90\%$  Galaxy)**

**constraint: [1,30]MeV, hugh rate(!!), charge resolution**

- **4π μ-tracking & after muon spallation products (n-activation)**

**constraint: ~[5,40]PE, time resolution (new PMT), linearity, high rate**

trigger rate driven by dark-noise ( $\sim 1\text{kHz}$ ) & muons ( $\sim 1\text{kHz}$ )

Why the SPMT design?

(UNO+physics constraints)

# SPMT into JUNO...

“perturbative” JUNO re-design



low cost



up to 36 000 channels (25 000 now)



tight schedule

**[details in Cedric's & Cayetano's talk]**

# SPMT into JUNO...

‘perturbative’ JUNO re-design



(**very**) low cost



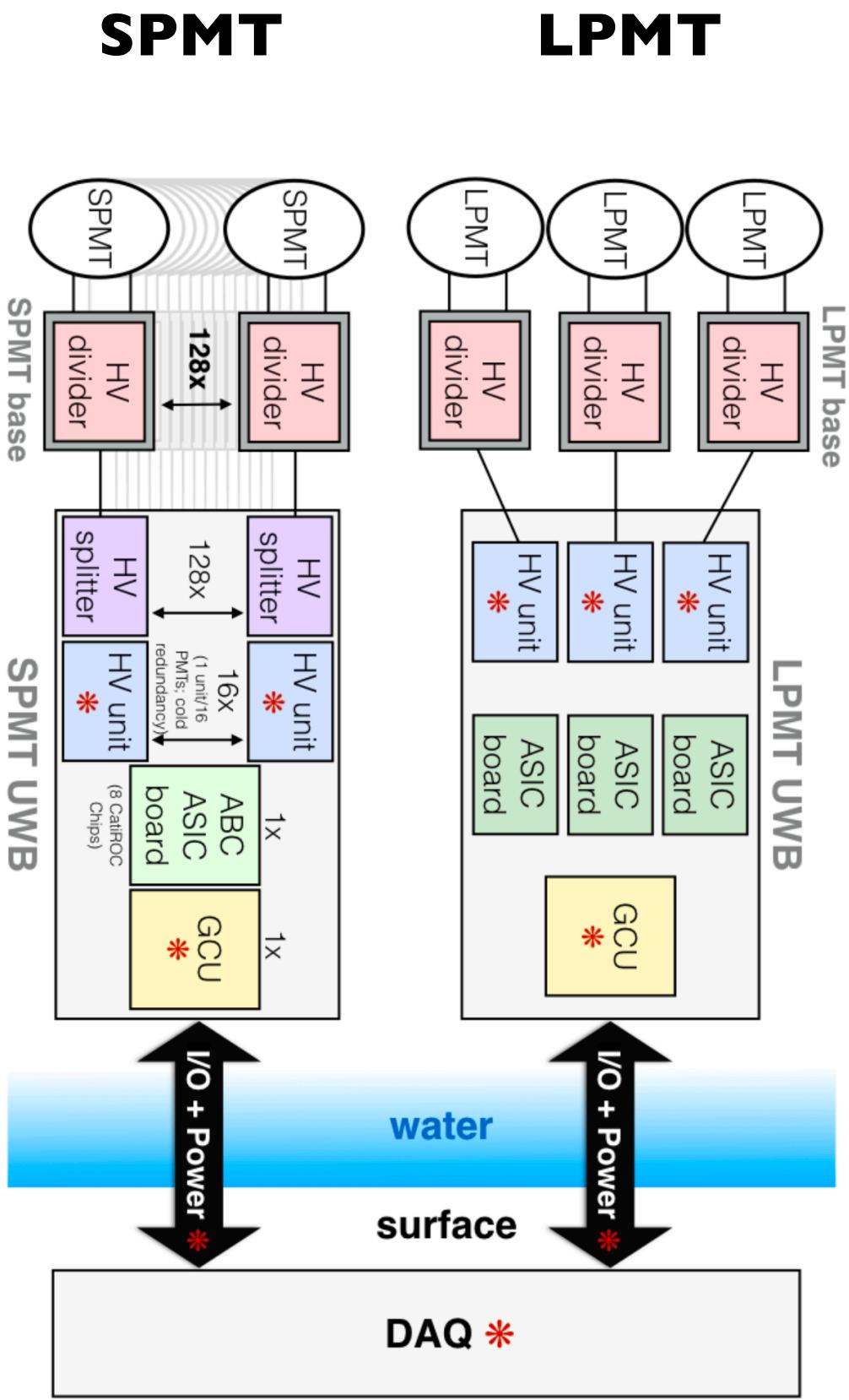
up to 36 000 channels (25 000 now)



(**very**) tight schedule

[**details in Cedric’s & Cayetano’s talk**]

# LPMT & SPMT commonalities...



**LPMT & SPMT design very similar → influence each other strongly**  
 (maximise common elements → **reduce cost & smooth integration**)

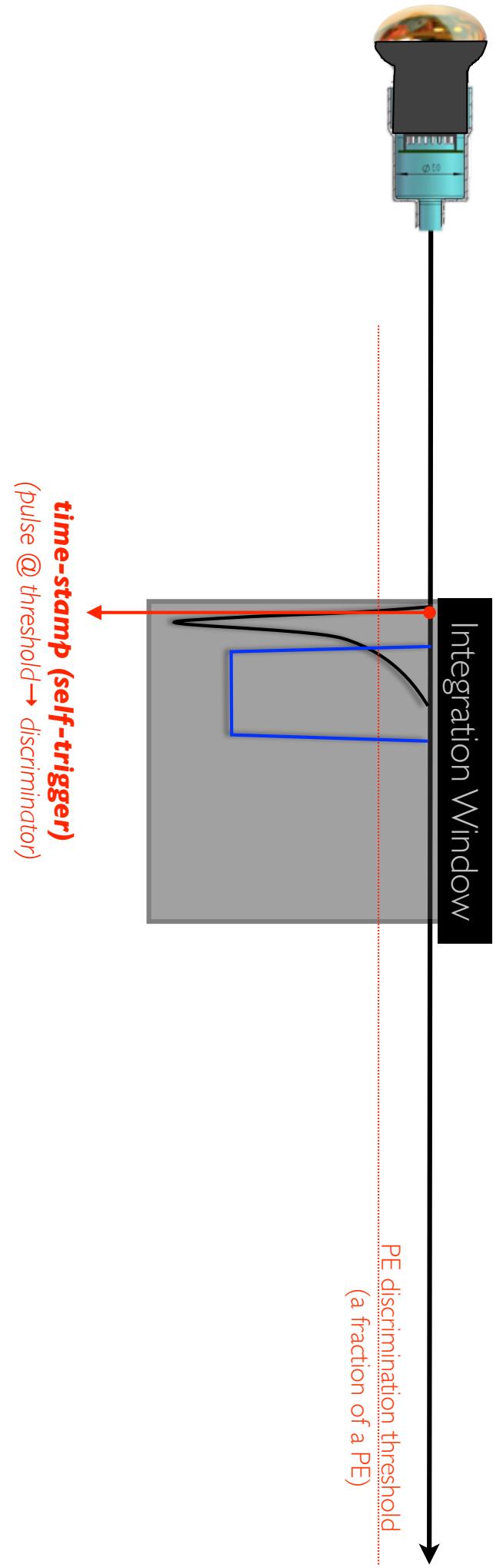
# (most important) JUNO-SPMT specifications...

**CatiROC**

**3" PMT**

	goal	critical	comment
<b>physical diameter</b>	$\leq 80\text{mm}$	yes	light level (photon-counting tuning)
<b>SPE width</b>	$\leq 40\%$	yes	SPE discrimination efficiency
<b>dark count (@1/4 PE)</b>	$\leq 1.5\text{k/s}$	yes	DAQ/readout rate (dominant)
<b>TTS (sigma)</b>	$\leq 1.5\text{ns}$	yes	position reconstruction
<b>HV nominal (<math>g=10^6</math>)</b>	$[0.8, 1.3]\text{kV}$	no	signal size (discrimination for noise)
<b>QE@420nm (average)</b>	$\geq 25\%$	yes	light level (photon-counting tuning)
<b>non-linearity [0,5]PE</b>	$\leq 1\%$	yes	linearity in physics regime (redundancy)
<b>non-linearity [5,100]PE</b>	$\leq 10\%$	no	linearity above main physics
<b>current</b>	$\leq 10\mu\text{A}$	yes	many PMTs on 1HV channel
<b>time resolution</b>	$\leq 0.5\text{ns}$	yes	negligible wrt PMT
<b>charge resolution</b>	$> 1/10 \text{ PE}$	yes	negligible wrt PMT & SPE discrimination
<b>pre-amp gain</b>	$[2, 10]$	ok	compensate (channel-wise) PMT gain
<b>(non supernova) max rate</b>	$10\text{kHz}$	yes	non-supernova physics rate capability
<b>deadtime</b>	$\leq 10\mu\text{s}$	ok	limits ADC maximal rate
<b>supernova max rate</b>	$10\text{MHz}$	yes	supernova physics rate capability

# Photon-Counting detection...



## Light Detection via Photon-Statistics...

- $I_{hit} = I_{PE}$  (multi- $PE$  not possible)
  - **no reconstruction** (extreme: no gain needed!)
  - multi- $PE$  contamination constraint by low-light level (i.e.  $\lambda_{(Poisson)} \leq 0.1$ )
  - **time-stamp** (example: discriminator on rising edge)
  - **charge info** (no need → but do something funky!)
  - **deadtime-less**

→ “DIGITAL” READOUT LIMIT

## Light Detection via Charge Measurement...

- $I_{hit} \geq I_{PE}$  (multi- $PE$  possible)
  - **calibration gain needed**
  - multi- $PE$  contamination constraint by low-light level (i.e.  $\lambda_{(Poisson)} \leq 0.1$ )
  - **time-stamp** (TDC-like)
  - **charge info** (ADC)
  - **deadtime-full**

→ “ANALOGUE” READOUT LIMIT

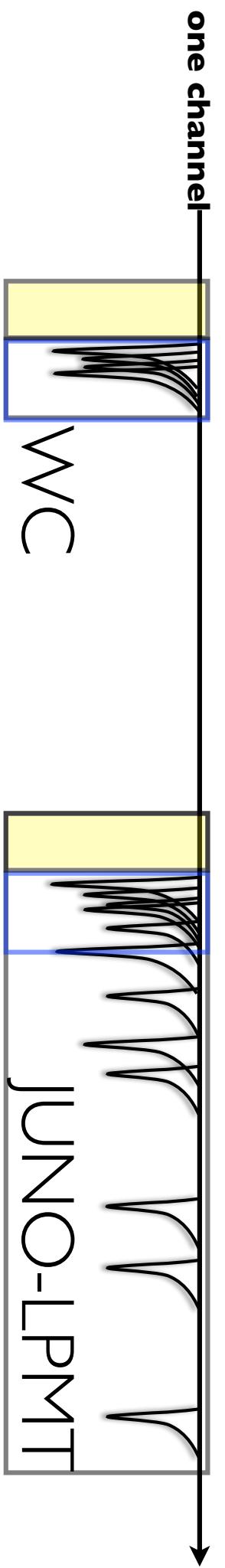
CatiROC $\oplus$ ABC can do both

CatIROC not good for JUNO

the challenge . . .

$\Delta t(T_{\text{ToF}}) \leq 200\text{ns}$   
(per channel)

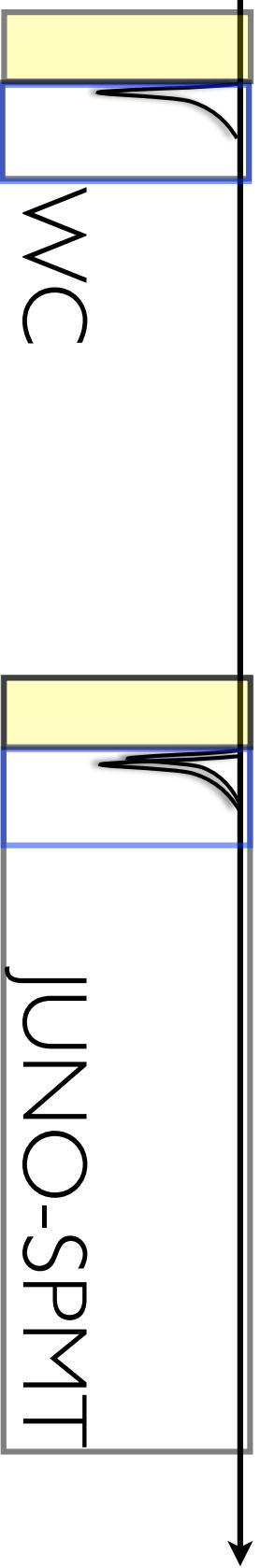
designed for Water-Cherenkov...



CatiRoC acceptance window  $\leq 200\text{ns}$  (per channel) → truncate some scintillation light

SPMT works in Photon-Statistics (low light level: IPE per trigger)

one channel



$WC \approx SPMT-JUNO$

good news . . .

SPMT-JUNO (low light)



PE self-trigger

**CatiROC→JUNO!**

(non-trivial)

but not enough...

“non-deterministic” deadline

&

maximal rate  $\leq \text{look/s}(\cdot)$

→ **success only via ABC**

(even less trivial)

today's evaluation...

What to expect?

We think... but do you agree?

## system CatIROC<sup>⊕</sup>ABC powerful SPM<sub>T</sub> readout

**constraints:**

tight schedule  $\oplus$  underwater

[no room for (big) mistakes]

# ABC card strategy...

## ABC-v0: first prototype

(goal: basic multi-ASIC readout)

**this evaluation here!**

(i.e.  $v0 \rightarrow v1$ )

## ABC-v1: final card

(goal:  $v0 \oplus$  extra functionalities)

bad news (sorry) . . .

**ABC-v0:** available only since JAN-2018  
(production delays)

- ✓ single-ASIC demonstration
- ?? multi-ASIC demonstration  
(FW development still ongoing)

**IMPORTANT: any HW defaults?**

your attention (please) . . .

**demonstration-0:** CatiROC work in JUNO?  
(we think so→**do you disagree?**)

**demonstration-1:** ABC work in JUNO?  
(ongoing effort→**feedback please**)

**demonstration-2:** SPMT readiness for JUNO?  
(complete strategy→**feedback please**)

# milestones 2018...

(3x) **ABCv0 → PMT testing system (China)**

[25,000 PMTs to be tested]

(3x) **ABCv1 ready for production & validation**

(220x) **ABCv1 production**

(~2000) **CatiRoCs production**

[early 2019: **prepare within 2018**]

**final validation & full integration**

[late 2019 — some details here]

**ready for detector installation**

[2020 — no details here]

SPMT system approved . . .

**mid-2016**

(most work 2017)

the SPMT team...

## 20 laboratories (9 countries) so far...

# Our (international) team...

### Armenia

- Yerevan Physics Institute (Yerevan)

### Brasil

- FABC (Sao Paulo)
- PUC (Rio de Janeiro)

### Belgium

- UCL (Brussels)

### Chile

- PUC (Santiago) ([project/physics coordination](#))

### China

- IHEP (Beijing) ([project/physics coordination](#))
- SYSU (Guangzhou)

### France

- APC (Paris) ([project/physics coordination](#))
- CENBG (Bordeaux) ([technical coordination](#))
- CPPM (Marseille)
- LLR (Paris)
- OMEGA (Paris)
- SUBATECH (Nantes)

### Italy

- Padova-INFN (Padova)

### Russia

- Moscow State University (Moscow)
- Institute of Nuclear Research & Russian Academy of Science (Moscow)

### Taiwan

- National Taiwan University NTU (Taipei)
- National Chiao Tung University NCTU (Hsinchu)
- National United University NUU (Miaoli)



# ABC-electronics related team (CNRS/IN2P3) . . .



**coordination:**

- Anatael CABRERA (APC, France) [**electronics**]
- Miao HE (IHEP, China) [**PMT**]
- Pedro OCHOA (PUC-Santiago, Chile) [**HV+under-water boxes**]

**project manager:**

- Cedric CERNIA (CENBG, France) [**all<sup>⊕</sup>interfaces + connectivity**]

**engineers:**

- Selma CONFORTI (OMEGA, France) [**CatiROC**]
- Frederic DRUILLOLE (CENBG, France) [**ABC testing**]
- Amelie FOURNIER (CENBG, France) [**under-water-box**]
- Cedric HUSS (CENBG, France) [**ABC routing**]
- Alexis NOURI (APC, France) [**ABC design — left**]
- Abdel RABII (CENBG, France) [**ABC FW**]
- Cayetano SANTOS (APC, France) [**ABC FW**]
  - (collaboration) Andrea TRIOLSSI (CERN, Switzerland)
- Guillaume VANROYEN (SUBATECH, France) [**ABC DAO**]

**physicists @ SPMT-CNRS/IN2P3 (support):**

- Clement BORDEREAU (CENBG, France) [**PhD**]
- Jose BUSTO (CPPM, France)
- Yang HAN (APC, France) [**PhD**]
- Cecile JOLLET (CENBG, France)
- Frederic PERROT (CENBG, France) [**simulation**]
- Mariangela SETTIMO (SUBATECH, France) [**DAQ**]
- Christophe de la TAILLE (OMEGA, France) [**CatiROC**]
- Frederic YERMA (SUBATECH, France)

**strong international collaboration behind [previous page]**

SPMT@IN2P3...

# 5 laboratories working together [APC+CENBG+CPPM+Ω+SUBATECH]



# questions...

- **SPMT huge system: anything but “small”**
- one of the largest PMT articulation (so far)
- many challenges → our ABC electronics @ core!
- this evaluation: help us see beyond ourselves!

merci...

thank you...

谢谢...