Model discrimination in the SGWB from Phase Transitions



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Based on

DC, V. Sanz and G. White [arXiv:1806.02332] DC, G. White [arXiv:1803.05438, JHEP]

Model discrimination in the SGWB?



GW spectra from thermal parameters

 $\frac{\beta}{H} = T \frac{\partial}{\partial T} \left(\frac{S_E}{T} \right)$

Normalized rate



Latent heat

Nucleation temperature

 $2\ln g_*$

 $\left|\frac{S_E}{T_N} \sim 177 - 4\ln\frac{T_N}{\text{GeV}} - \right|$

 $p(t_N)t_N^4 = 1 p(T) = T^4 e^{-S_E}/T$

GW spectra from thermal parameters

$$h^{2}\Omega_{\rm sw} = 8.5 \times 10^{-6} \left(\frac{100}{g_{*}}\right)^{-1/3} \Gamma^{2} \bar{U}_{f}^{4}(\alpha) \left(\frac{\beta}{H}\right)^{-1} \underbrace{v_{w} S_{\rm sw}(f)}_{\text{Sw}} \text{Weir, arXiv:1705.01783}$$

$$f_{\rm sw} = 8.9 \times 10^{-7} \text{Hz} \frac{1}{v_{w}} \left(\frac{\beta}{H}\right) \left(\frac{T_{N}}{\text{Gev}}\right) \left(\frac{g_{*}}{100}\right)^{1/6} \underbrace{S_{\rm sw}}_{\text{Sw}} = \left(\frac{f}{f_{\rm sw}}\right)^{3} \left(\frac{7}{4+3\left(\frac{f}{f_{\rm sw}}\right)^{2}}\right)^{7/2}$$

- Two parameters:
 - $\boldsymbol{\Omega}$ and $\boldsymbol{\mathsf{f}}$
 - Peak amplitude and frequency
- How much model discrimination is possible?

Effective models for a first order PT

- Minimal model for a PT: double well potential → three terms in the effective potential, with relative signs
- We consider two limiting cases: 1. $V(h_D, T) = \frac{1}{2}m(T)^2h_D^2 - c_3(T)h_D^3 + \frac{1}{4}\lambda(T)h_D^4$ 2. $V(h_D, T) = \frac{1}{2}m(T)^2h_D^2 - \frac{1}{4}\lambda(T)h_D^4 + c_6(T)h_D^6$
- Models: SSB in a dark gauge sector



PT from renormalizable operators (1)

$$V(H,T) = \Lambda^4 \left[-\frac{1}{2} \left(\frac{h_D}{v} \right)^2 + \frac{1}{4} \left(\frac{h_D}{v} \right)^4 \right] + V_{\rm CW}(h_D)$$
 Zero tempositive tem

Zero temperature potential

inite temperature correction

(see for example Quiros, [hep-ph/9901312])

High temperature expansion of the thermal functions

$$V(H,T) = \Lambda^{4} \left[\left(-\frac{1}{2} + \left(\frac{1}{8} + \frac{N_{G}}{24} \right) \frac{T^{2}}{v^{2}} + \frac{3}{24} N_{GB} \frac{g^{2}}{4} \frac{T^{2} v^{2}}{\Lambda^{4}} + y^{2} N_{f} \frac{T^{2}}{24} \frac{v^{2}}{\Lambda^{4}} \right) \left(\frac{h_{D}}{v} \right)^{2} - \left(N_{GB} \left(\frac{g^{2}}{4} \right)^{3/2} \frac{1}{4\pi} \frac{v^{3} T}{\Lambda^{4}} \right) \left(\frac{h_{D}}{v} \right)^{3} + \frac{1}{4} \left(\frac{h_{D}}{v} \right)^{4} \right]$$

$$\mathbf{v}/\Lambda, \mathbf{g}, \mathbf{N}_{\mathsf{GB}}, \mathbf{N}_{\mathsf{f}} \mathbf{x} \mathbf{y}$$

PT from non-renormalizable operators (2)

$$V(h_D, T) = \frac{1}{2}m(T)^2h_D^2 - \frac{1}{4}\lambda(T)h_D^4 + c_6(T)h_D^6 \xrightarrow{\frac{N^4}{4}} \underbrace{\sum_{v(\phi) = 0}^{N^4} \frac{1}{v^2}}_{v(\phi) = (2 - 3\kappa)\frac{\Lambda^4}{v^2}} \\ \lambda(0) = 4\frac{\Lambda^4}{v^4} \\ c_6(0) = \kappa\frac{\Lambda^4}{v^6} \underbrace{\sum_{v(\phi) = 0}^{N^4} \frac{1}{v^2}}_{v(\phi) = 1} \underbrace{\sum_{v(\phi) =$$

Aside: EWPT

• Up to dimension-6 operators, models of the EWPT can be captured by the previous potential

$$V_{6}(h,T) = \left(a_{T}T^{2} - \frac{\mu^{2}}{2}\right)h^{2} + \left(b_{T}T^{2} - \frac{\lambda}{4}\right)h^{4} + \frac{1}{8\Lambda_{6}^{2}}h^{6}$$

$$a_{T} = \frac{y_{t}^{2}}{8} + 3\frac{g^{2}}{32} + \frac{g'^{2}}{32} - \frac{\lambda}{4} + \frac{v_{0}^{2}}{\Lambda_{6}^{2}}4$$

$$b_{T} = \frac{1}{4\Lambda_{6}^{2}}$$
Scale of new physics; a singlet, an extra doublet, ...

$$f_{\rm sw} = 8.9 \times 10^{-7} \text{Hz} \frac{1}{v_w} \left(\frac{\beta}{H}\right) \left(\frac{T_N}{\text{Gev}}\right) \left(\frac{g_*}{100}\right)^{1/6}$$











$f_{\rm sw} = 8.9 \times 10^{-7} \text{Hz} \frac{1}{v_w} \left(\frac{\beta}{H}\right) \left(\frac{T_N}{\text{Gev}}\right) \left(\frac{g_*}{100}\right)^{1/6}$





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Romano, Thrane; arXiv:1310.5300

$$h^2 \Omega_{\rm sw} = 8.5 \times 10^{-6} \left(\frac{100}{g_*}\right)^{-1/3} \Gamma^2 \bar{U}_f^4(\alpha) \left(\frac{\beta}{H}\right)^{-1} v_w S_{\rm sw}(f)$$





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Some qualitative lessons

- Most thermal parameters are sensitive to the ratio v/Λ
- Extra fermions ($N_f x y^2$), and a larger gauge group (N_G),
 - Enhance the amplitude of the signal
 - Sensitive to heavier scalars (effective zero temperature mass)
- The effective non-renormalizable potential yields better detection prospects @LISA

Exotic spectra from simultaneous PT



$$\Omega_{GW} = \sum_{i} w_{i} \,\Omega_{GW}^{(i)} \left(\Omega_{0}^{(i)}, f^{(i)}, \Upsilon_{i}, v_{w}^{(i)}, \frac{\beta^{(i)}}{H}, T_{N}^{(i)}\right)$$

$$1 \quad (0,0) \rightarrow (v_{A},0) \qquad \text{``First transition''}$$

$$2 \quad (v_{A},0) \rightarrow (v_{A}, v_{B}) \qquad \text{``Bubble in a bubble''}$$

$$3 \quad (0,0) \rightarrow (v_{A}, v_{B}) \qquad \text{``Immediate transition''}$$

Observations:

- First PT does not reach $f_1(T) = 1$
- Transitions 2) and 3) correspond to different thermal parameters



To conclude,

- Gravitational waves are new independent probes of particle physics!
- Degeneracy is introduced in the thermal parameters, and again in the GW spectra
- However, some general qualitative lessons can be learned
- More exotic spectra may result from simultaneous PT
- Things to do:
 - Beyond the high-T approximation
 - Beyond one loop
 - HDOs and more general effective potentials

Thank you!

Extra slides

Fermions





