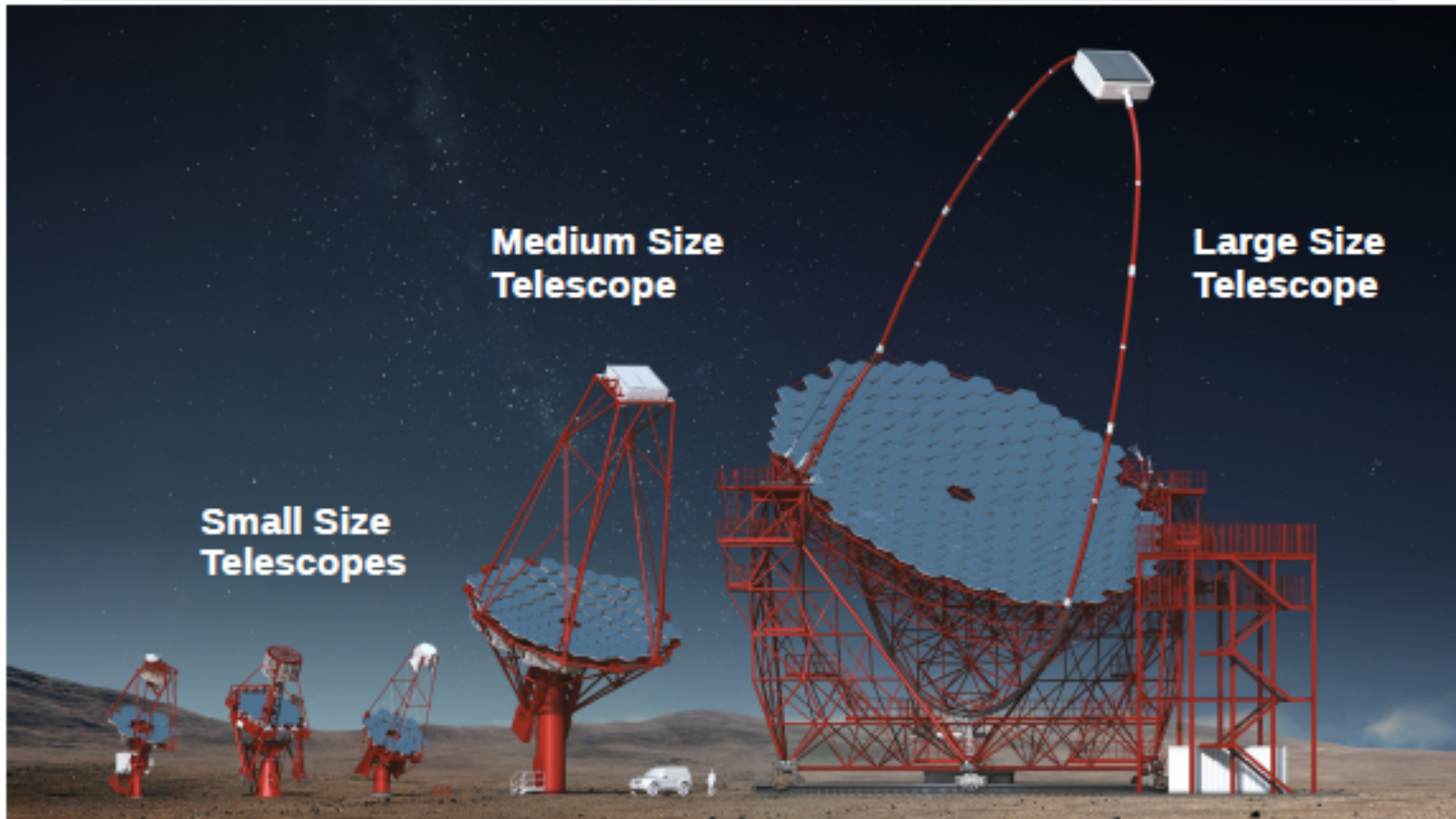


# Activités du groupe HESS/CTA

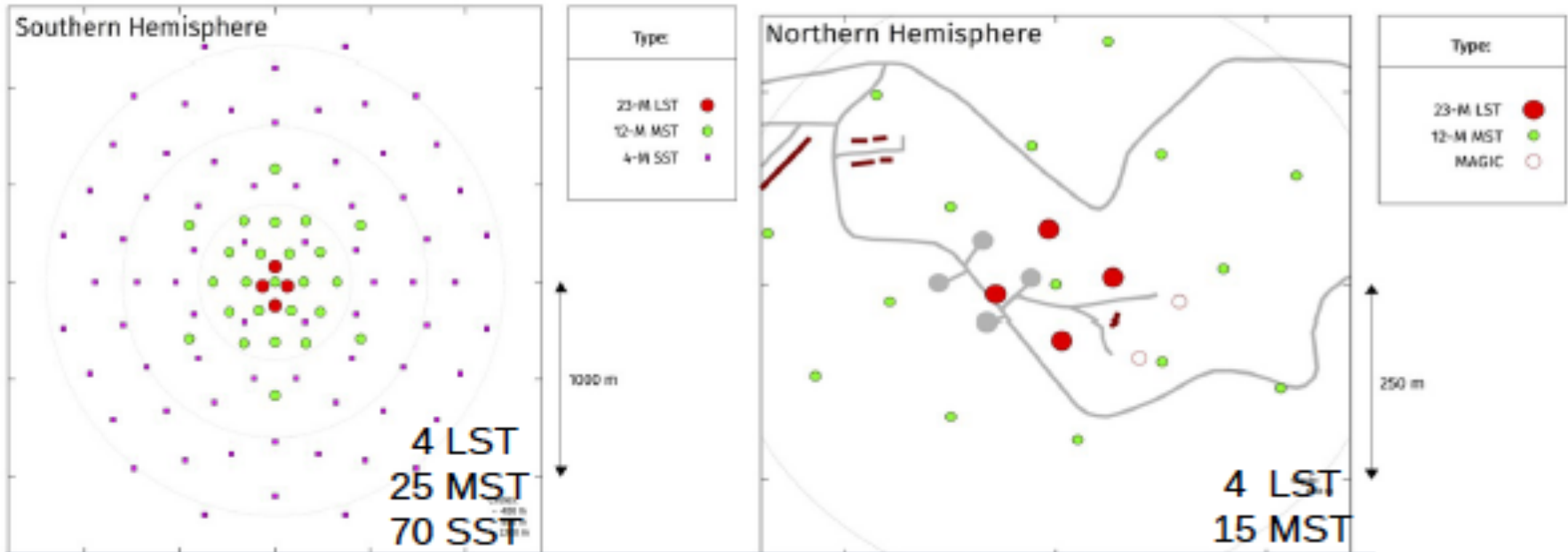


Fonctionne depuis 2004 avec 4 « petits » et depuis 2012 avec le 5eme grand télescope  
Fonctionnement jusqu'en 2019. Discussion en cours pour un prolongement.

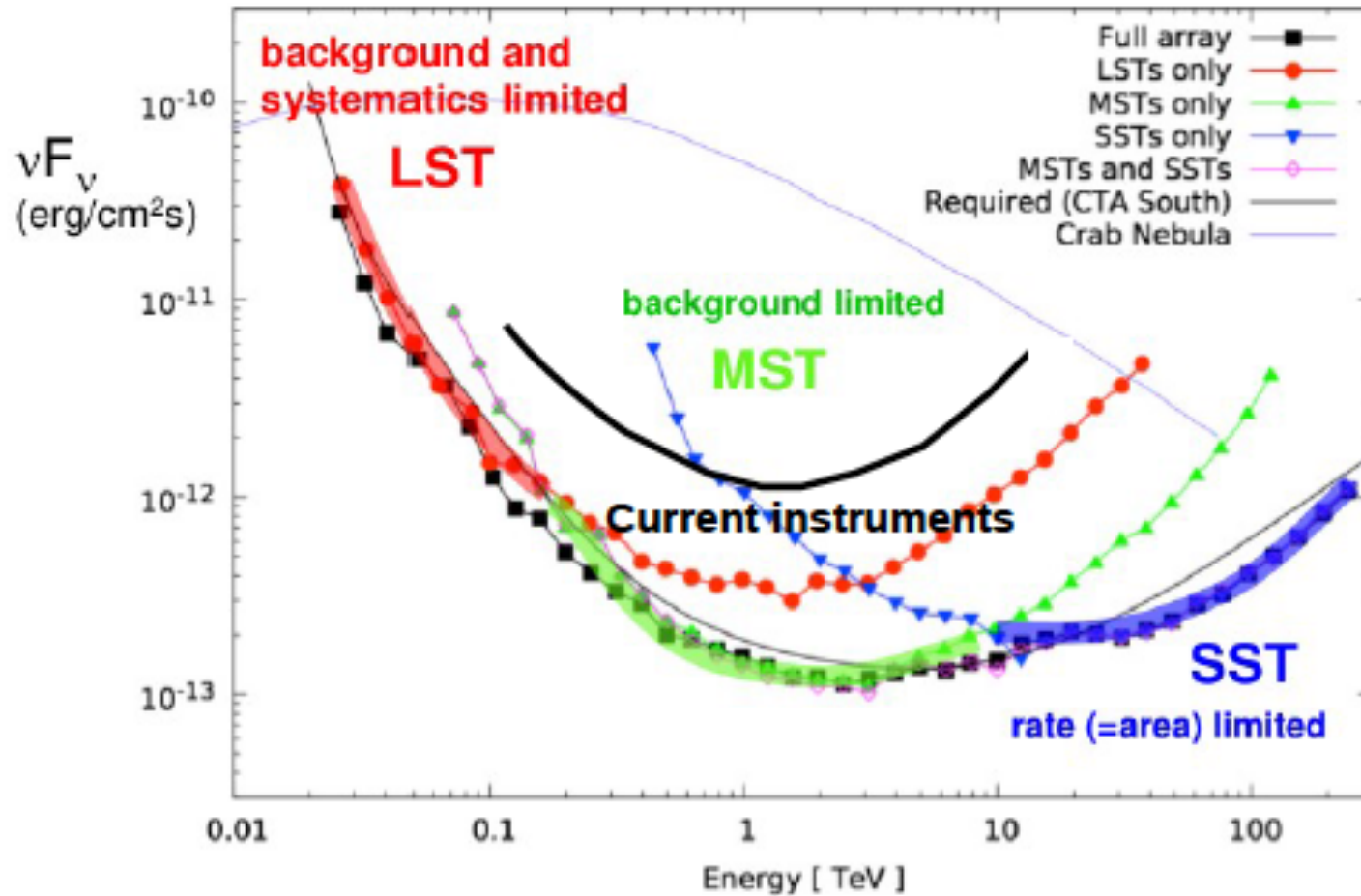


Three different sizes of telescope optimized for three different energy ranges

# CTA – 2 sites



# CTA - sensibilité



# CTA - LST

Grands télescopes de CTA (4 par site)

- 24m de diamètre
- ⇒ Sensible a partir de 20-30 GeV
- Structure légère
- ⇒ Moins de 20s pour pointer une région du ciel

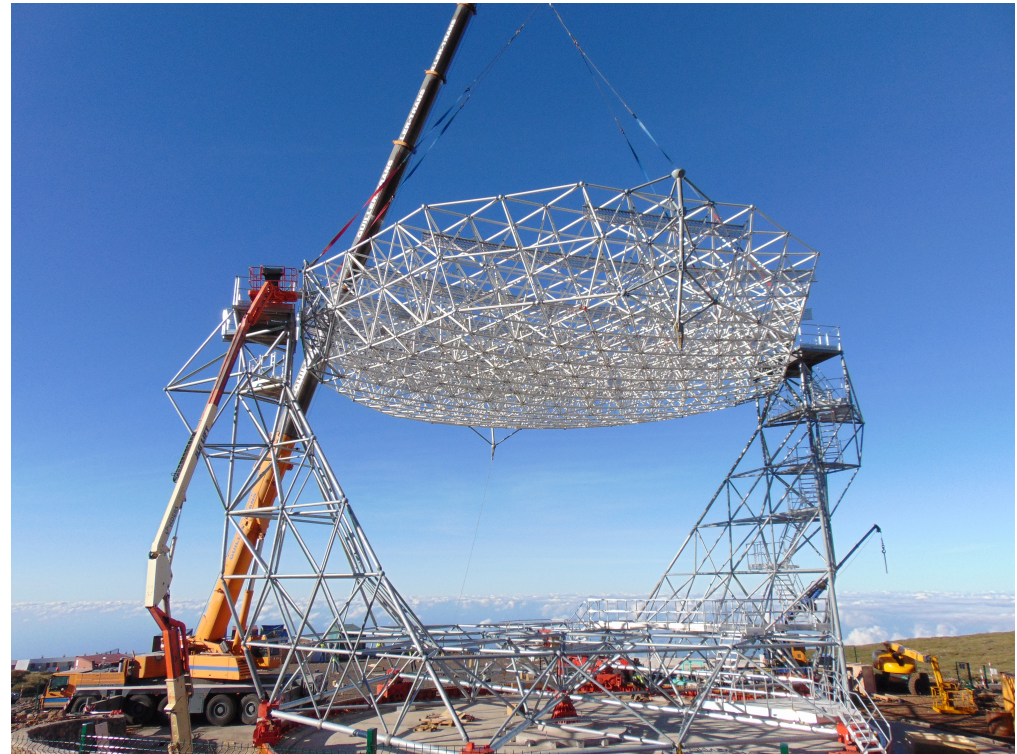
Au LAPP:

- Mécanique: Arche haubanées
- Mécatronique: Drive
- Electronique: Contrôleur de camera
- Informatique: drive (+ gestion des alertes) & SW de contrôle du télescope

**Première lumière du LST1 automne 2018.**

3 autres attendus entre 2020 et 2023.

Le LAPP est responsable des plusieurs éléments critiques pour le suivi d'alerte transients



# CTA - Data

Gilles Maurin, Thomas Vuillaume, Florian Gate

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Contribution majeure dans le WG Data Management de CTA

Groupe très actif dans la définition du pipeline d'analyse de CTA

⇒ Définition du framework

⇒ Introduction du calcul haute performance

⇒ Implémentation de méthodes performantes (machine learning...)

Le groupe est en très bonne position pour analyser efficacement les toutes premières données

# La Physique qui nous intéresse

## Galactique

*SNR/MC*  
A. Fiasson

*SFR*  
G. Maurin  
G. Lamanna

## Extragalactique

*AGN – GRB et physique fondamentale  
(EBL...)*

D. Sanchez (deputy convener HESS)  
G. Maurin  
T. Vuillaume  
A. Fiasson  
Q. Piel  
F. Gaté  
A. Carosi

## Matière noire

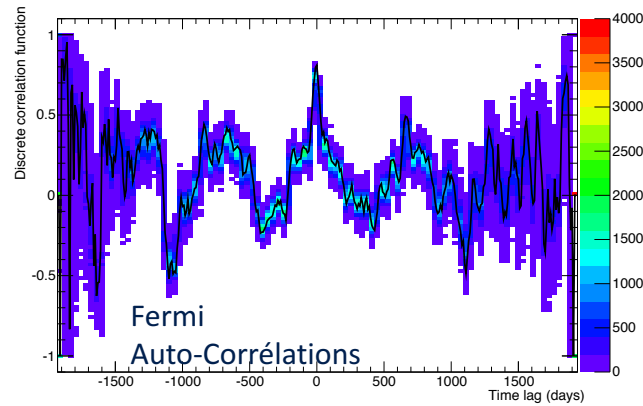
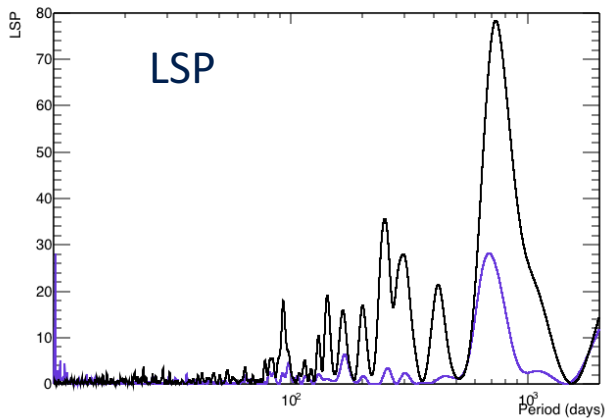
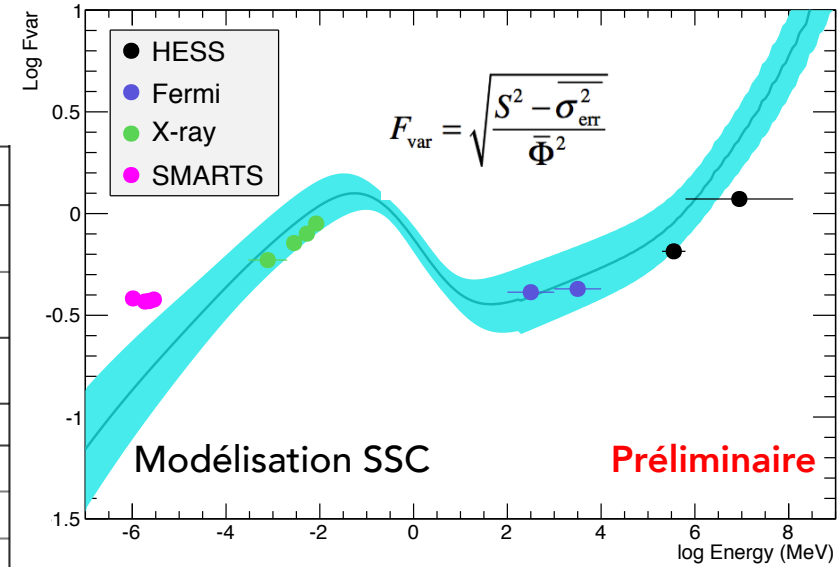
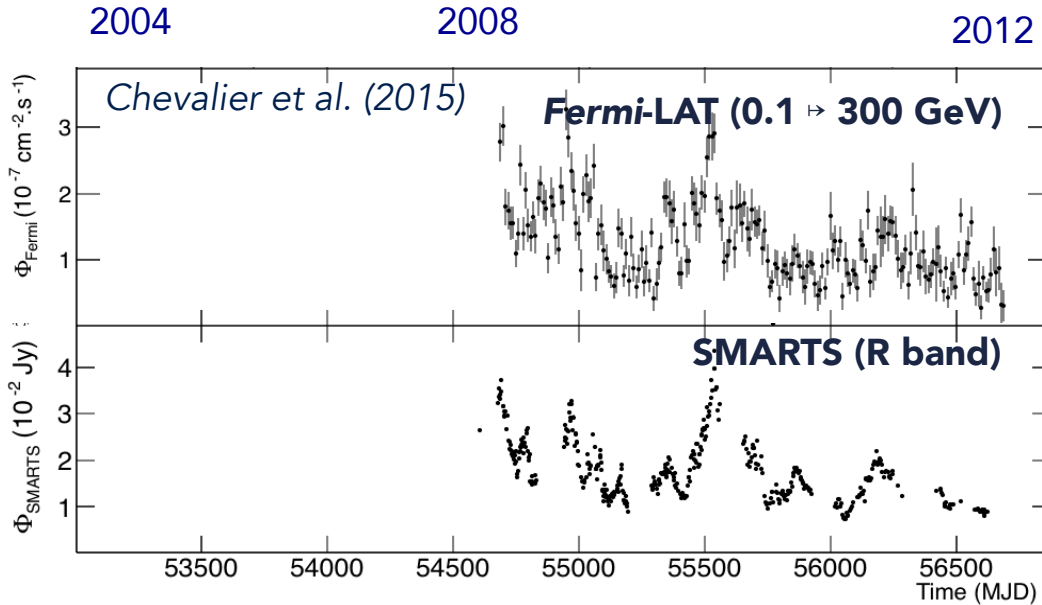
V. Poireau  
C. Armand  
J.P. Lees  
G. Lamanna



# AGN - Étude des blazars et de leur variabilité

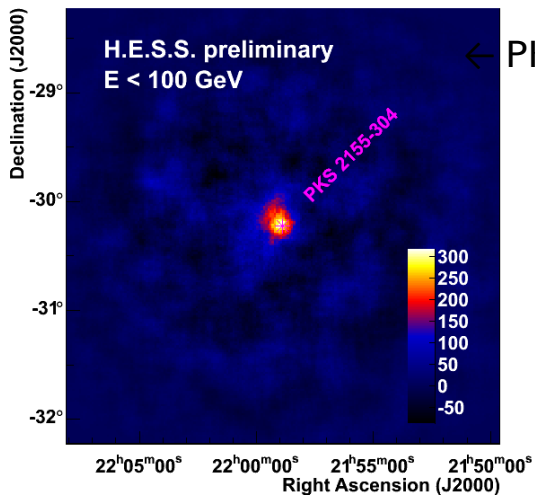
David Sanchez, Jill Chevalier

## Courbes de lumière MWL de PKS 2155-304



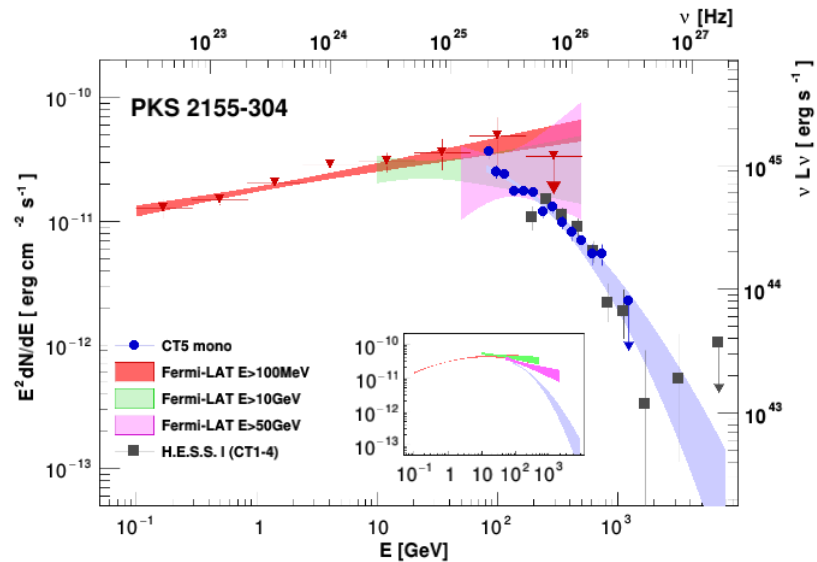
# AGN – Premiers résultats H.E.S.S. II

David Sanchez

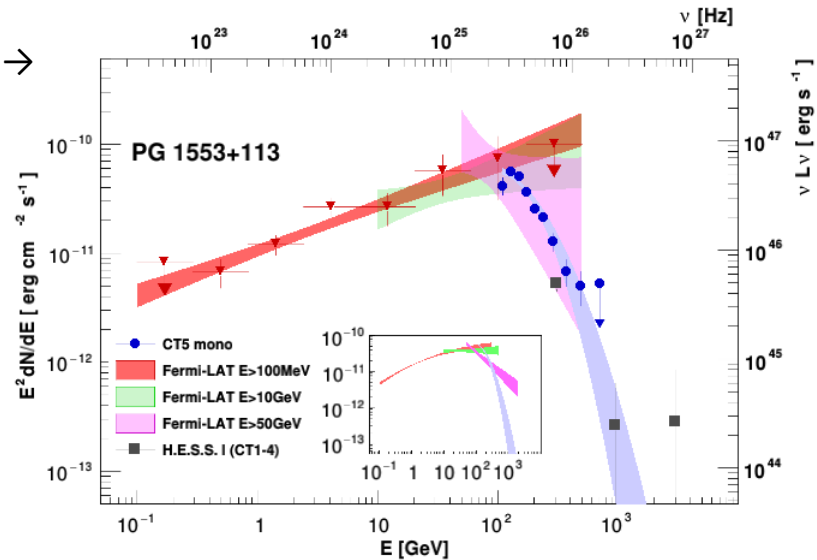


← PKS 2155-304  
E < 100 GeV

PKS 2155-304 →  
2013-2014



PG 1553+113 →  
2013



- ◆ Premiers résultats H.E.S.S. II sur les AGN
- ◆ Papier soumis

# GRB - H.E.S.S.

Quentin Piel, Alessandro Carosi & Armand Fiasson

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- (Re)analysis of GRB observed since 2012

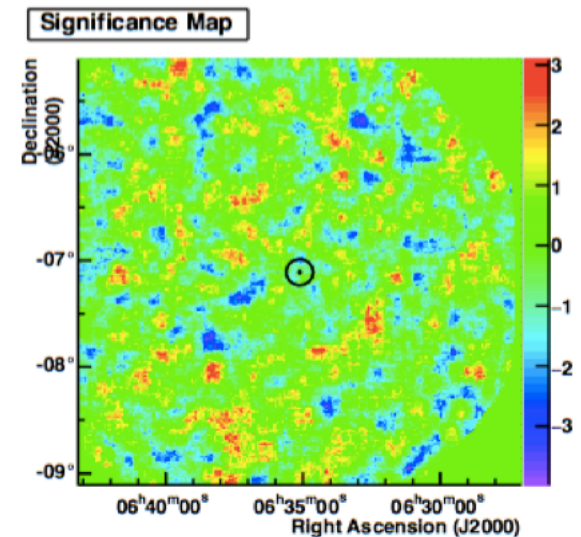
35 GRB observations made by H.E.S.S.

Aim

Having a detection or a catalog of upper limits

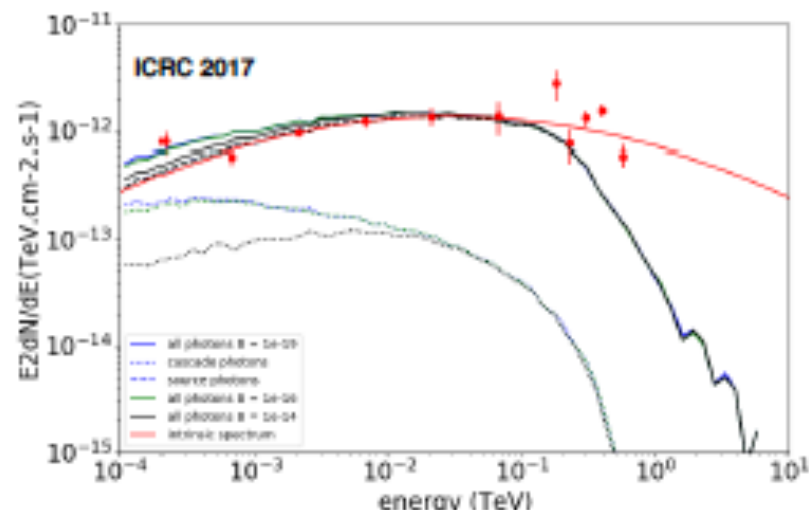
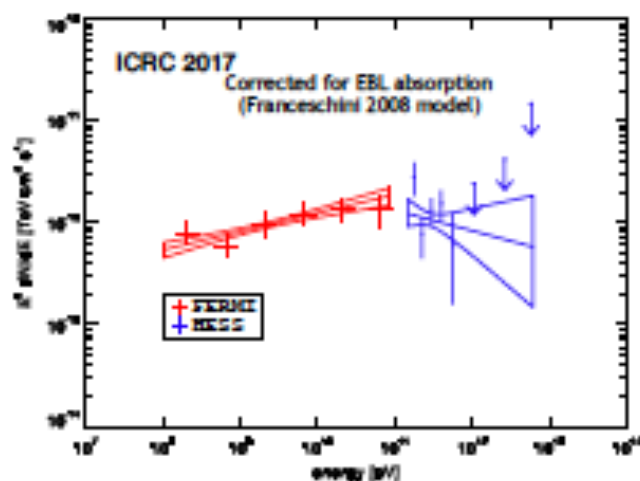
- PI of the GRB observation program of 2018

In charge of the proposal writing and GRB experts



- HESS activities: propagation task force (Gaté, Sanchez)

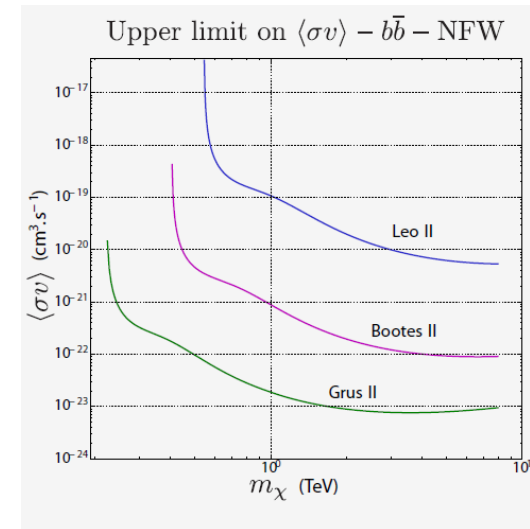
- Study observed spectra modification for IGMF assumption with HESS sources.
- Constrain superior limit on IGMF

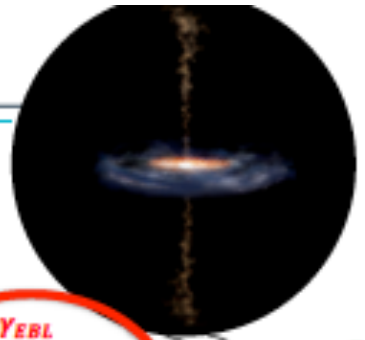


- Study ongoing on large number of sources
- Proposal for high opacity sources of interest

# DARK MATTER

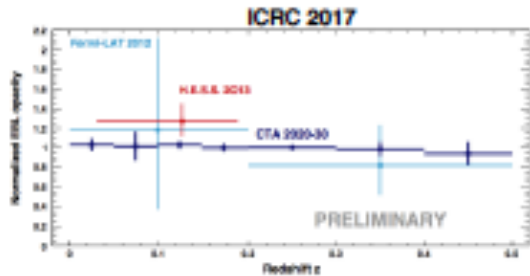
- Group for indirect dark matter search using HESS
  - Vincent Poireau, Céline Armand, Jean-Pierre Lees, Francesca Calore (LAPTh)
  - Céline : thesis 2017-2020, 80 % theoretical work, 20 % experimental work
- Indirect search for dark matter with gamma rays using DM rich regions
- Dwarf spheroidal galaxies
  - Three dwarves observed in 2016 already analyzed
  - Analysing now 2017 data: three dwarves for 30 h of observation
  - Proposal for 2018 for 70 h of observation
- Irregular galaxies
  - Proposal for 2018 for 30 h of observation



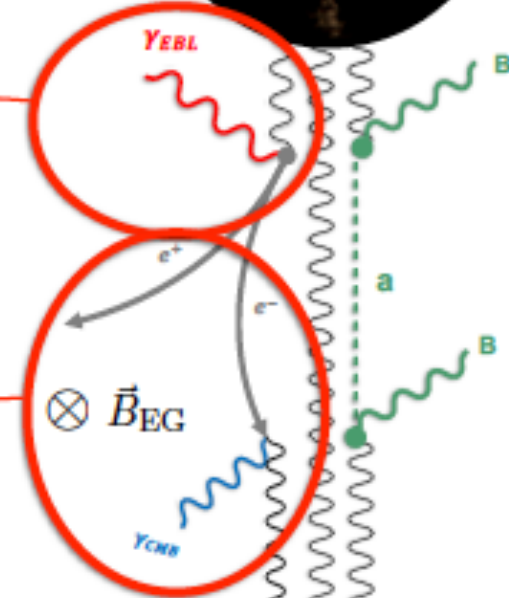
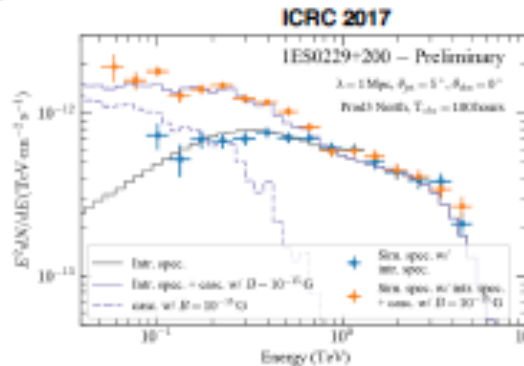


• CTA activities: propagation task force (Gaté, Sanchez, Piel)

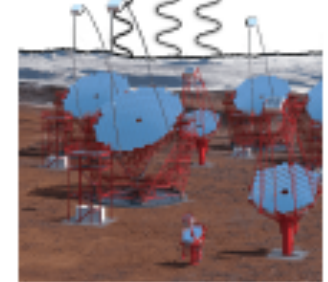
- Constrain Extragalactic Background Light opacity



- Probe extragalactic magnetic fields



- Sources simulation: CTA IRFs, layout, ...
- CTA capabilities to probe IGMF, EBL...

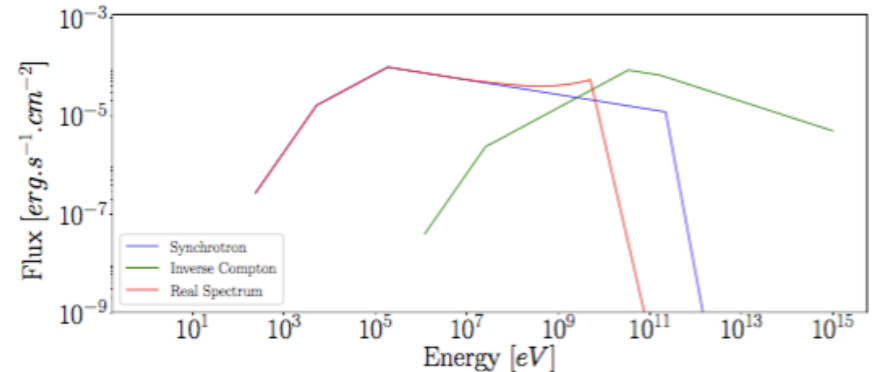


# GRB - CTA

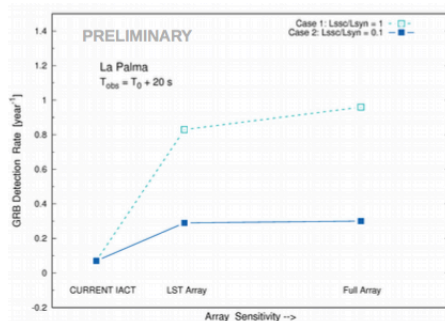
Quentin Piel, Alessandro Carosi & Armand Fiasson

## Simulation of Gamma-Ray Bursts

Prompt and Afterglow phases  
(simple Synchrotron+SSC processes)



## GRB detection rate prediction with CTA



## GRB detection for different set of physical parameters

