

Why do we care about LLPs?

Long lived particles are generic consequence of theories with:

- Small couplings
- Scale (or loop) hierarchies
- Phase space suppression

SM provides template weak decays: K_L or n!



Encompasses a very broad theory space. In terms of portals, can see these theories' 'LLP footprints' through e.g.

Kinetic mixing Higgs mixing Lepton yukawa

LHC coverage

- ATLAS/CMS will set best limits for charged, colored and/or heavy LLPs
- LHCb: $\mathcal{O}(\text{GeV})$ neutral LLPs with large muon BR and $c au \sim 1\,\text{m}$
- In some cases, ATLAS, CMS, LHCb coverage is already complementary



- Large hadronic BGs: lighter, neutral, longer-lived LLPs are hard to see
- There are few good theory priors. Hard to build a single detector to cover all these cases: Complementarity is crucial!

Dean Robinson dean.robinson@uc.edu

CODEX-b proposal

Pre-Run 3 (2020): Data AcQuistion will be moved to surface

largest shielded space at LHC near an IP!



CODEX-b proposal

Pre-Run 3 (2020): Data AcQuistion will be moved to surface

largest shielded space at LHC near an IP!



Why is CODEX-b a good idea?

- Instrumented volume is small: Consider more ambitious detection technologies such as calorimetry or time-of-flight compared to other proposals. (More later)
- Close to IP: Only \sim 4 bunch crossing times for relativistic objects. Integrate CODEX-b into the DAQ & readout, and treat as subdetector
 - $\circ~$ Identification and at least partial reconstruction of the LLP event.
 - $\circ\,$ E.g. tag a VBF jet for Higgs decays, or an associated $K^{(*)}$ for B decays.
- Backgrounds are controllable
 - Ample space for additional shielding
 - Also decorrelated wrt other experiments
- Competitive NP Reach!

my (+ repetion requires in chill and and an inf again in any of an and Malun wind wind longo (n. Sola The art and all needed after stand " (Chief and the mail and Guiltin apringhi - Clair of

ברביאה הואישאה מימי למין ויאווירי מויי למווארמיות Benchmarks & Portals

1:15

BG Study He (the remain of buse of our I for for lange to בין וואשורייה ליולביות אומר אלא ולשכו ליוא אייליי איי לייון איי ה Outlook

Prelude & Sat Prelude & Setup

In the second second second second second and and second s

Benchmark Scenarios

Five benchmark LLP scenarios

- Massive dark photon γ_d (kinetic mixing)
- Light $\mathcal{O}(\text{GeV})$ scalar φ (Higgs mixing)
- Heavy neutral lepton, N (mixing with the active neutrinos)
- $h \rightarrow \text{dark glueballs}$ (twin Higgs mixing portal)
- QCD coupled ALP $m_a < 3\pi$ (diphoton channel!)

180x.xxxxx [Evans, Knapen, Papucci, Ramani, DR]

$b \rightarrow s \varphi$: Higgs-scalar mixing

Single parameter portal: Higgs-scalar mixing angle, θ ,

- In the mass basis, effective mixing $\sin \theta \varphi y_f \overline{f} f$
- For $m_{\varphi} \sim 1 \text{ GeV}$, dominant production from $b \rightarrow s$ decays, leveraging huge $\sigma_{b\overline{b}} \sim 500 \,\mu\text{b}$
- Production and decay controlled by same portal





- LHCb reach scaled from $B o \mathcal{K}(arphi o \mu\mu)$ results, assuming no BG
- Reach/size exceeds most other proposals!
- NB large theory uncertainties for $m_arphi\gtrsim 1\,{
 m GeV!}$ (cf. 1708.08503 [Evans])

Application: Sterile Coannihilating DM

DM χ mixes with a heavier state ψ that coannihilate to the Higgs mixing scalar φ . 1803.02901 [Tito D'Agnolo, Mondino, Ruderman Wang]



Heavy Neutral Lepton: N

A heavy 4th (Dirac or Majorana) neutrino: effective mixing via extended PMNS matrix elements: U_{eN} , $U_{\mu N}$, or $U_{\tau N}$.

Production channels

- $c \rightarrow s \ell N$, $D_s \rightarrow \ell N$ $(\sigma_{c\overline{c}} \sim 5 \text{mb!})$
- $b \to c \ell N$, $B_c \to \ell N$
- $W/Z \rightarrow \ell/\nu N$ (Large GB sample vs fixed target!)

Decay channels

- $N \rightarrow 3\nu$ (Z channel; always present!)
- $N \rightarrow \nu \ell \ell$ (Z or W channel)
- $N \rightarrow \ell$ had (W channel)

Consider U_{eN} , $U_{\mu N}$, $U_{\tau N}$ one at a time



mass range $0.1 \le m_N \le 10 \, {
m GeV}$

HNL: U_{eN}



- Reach dominated by charm decays, B decays at high masses
- Strong limits set by current data
- cf. similar analysis by 1803.02212 [Heo, Hirsch Wang]

HNL: $U_{\mu N}$



- Reach dominated by charm decays; *B* decays at high masses; *W*, *Z* in short lifetime regime
- Strong limits set by current data
- cf. similar analysis by 1803.02212 [Heo, Hirsch Wang]

HNL: $U_{\tau N}$



• Only weak limits set by current data!

HNL: $U_{\tau N}$



- Only weak limits set by current data!
- Reach dominated by B, B_c decays; W, Z in short lifetime regime

$h ightarrow \gamma_{\rm d} \gamma_{\rm d}$

Higgs-Dark photon portal:

 $yhF'_{\mu\nu}F'_{\mu\nu}+\epsilon F'_{\mu\nu}B_{\mu\nu}$

- $Br[h \rightarrow \gamma_d \gamma_d]$ and γ_d lifetime controlled by separate parameters
- $\gamma_{\rm d}$ branching ratios fixed by e^+e^- data
- GF Higgs production, simulated with Pythia8



$h \rightarrow \gamma_{d} \gamma_{d}$: Higgs-Dark photon portal $m_{\gamma_{d}} = 0.5 \text{ GeV}$ $m_{\gamma_{d}} = 10 \text{ GeV}$



- For $m_{\gamma_d} = 10$ GeV, ATLAS disp. dijet reach dominates 1605.02742 [Coccaro, Curtin, Lubatti, Russell, and Shelton] 1412.0018 [Curtin, Essig, Gori, and Shelton]
- For $m_{\gamma_d} = 0.5 \,\text{GeV}$, ATLAS reach is sys limited: Assume factor of 5 improvement. (Could scale as much as \sqrt{L})

Dark glueballs: G^0 (preliminary)

'Fraternal' Twin Higgs (neutral naturalness) setup:

- Heavy-light Higgs mixing via $f/v \gtrsim 3$.
- One heavy flavor dark QCD sector: glueball lightest state

• Decays to SM, $G^0 \rightarrow qq$, via higgs portal. Loop and yukawa suppressed, $\Gamma \sim m_0^7/f^4 m_h^2!$

 G^0



- ATLAS/ CMS reach limited by VH production: trigger on 1DV + 1 lepton/ associated hard object 1506.06141 [D. Curtin, C. Verhaaren]
- Can access lighter glueball regime, over broad range of f values!

Axion-like Particles, a (preliminary)

Well-motivated over broad range of hidden valley type theories

- UV coupling to $aG\widetilde{G}$
- Induces αaFF̃ operator and mixing with SM π⁰!
- For m_a < 3π, production through π⁰-a mixing, decay to γγ!



- Reach leverages huge π^0 production rate, modulo exponential fragmentation penalty (a la Lund: $\sim e^{-(m_a^2 m_\pi^2)\pi/\kappa}$)
- Reach supplements beam dump in vital coupling/mass regime!
- Effective reach even for $a \rightarrow ee\gamma$ [Br $\sim 1\%$] But can CODEX-b see photons?!?

mille actation values by chill and unforthe same and anay Malun anin bane lando A Store and all news a first work " (Checking of a mail of the Guitif aparque - (Coli of

A Stand Branch and and a far and provided Prelude & Setup הריקה הנחדקיה מחיי היי (יאוורי מויי להווידיהה Annune Bounda erguale allarge ment

1:25

Benchmarks & Portals

BG Study minant of bung of come the finance בין וואשוראי ליינראי אור אלי ולים אירו איילוא איילוא איין איי

Outlook

Backgrounds

Modest amount of additional shielding is needed for CODEX-b



- Conceptual study:
 - Passive Pb shield attenuates muon & neutral hadron BGs 0
 - A thin veto layer rejects secondary production within the shield. 0
- To estimate BGs we use a (preliminary) Geant4 simulation, validated with simplified propagation model
- BGs are controlled in this proof-of-concept with active veto eff $\sim 10^{-5}$

Data-driven BG calibration

- Data-driven BG measurement in the LHCb cavern
 - Measure background charged flux behind the UXA wall (later: different shield thicknesses/materials)
 - $\circ\;$ Allows us to begin to properly calibrate MC estimates



- Happening this year!
 - A BG calibration with scintillator+PMT: Plan to install behind UXA wall for engineering run

BG Calibration

Wrapped scintillator panel: $30 \times 30 \times 2$ cm. Measure charged flux through UXA wall



Courtesy of Raphael Dumps



mille actation values by chill and unforthe same and anay Malan main marching Store and all news a first work " (Checking of a mail of the Guiltin Apringhi - Clair of

I a stand of the solution of the sound of the branch of 1:25 Prelude & Setup ברייקה הנהדבאה כוחוי לחי ויאוריי כולי להואימות Summer Mounta Adjust allarga menu Benchmarks & Portals BG Study Hale time remains of trippe of since IT for lague of Outlook

Thoughts & Questions

CODEX-b significantly enhances the NP reach and capabilities of the LLP programme, complementing or exceeding the reach of other LHC experiments, with largely decorrelated backgrounds

Timeline:



Questions for the workshop:

- $a \rightarrow \gamma \gamma$: Is there a plausible calorimetry solution at CODEX-b?
- More reach: If LHCb dismantles the shield wall to bring out DELPHI, can the shield wall be rebuilt closer to LHCb to buy even more decay volume?
- Tracking: Is a NOVA based techonology a better approach?

Thank you!

and on the providence of the second s

Extras and Details: CODEX-b

norme Koren al ner es binnel binnel and foren and ner es binnel binnel della ner sacion binne binne binne della ner sacion binne binne allo la binnel binne hanne binne allo la binnel binnel and binne allo binnel binnel and binne binne binnel binnel and binne binne binnel binnel and binne binnel binnel binnel and binne binnel binnel binnel binnel binne binnel binn

Primary Backgrounds

• Primary muons may scatter on air. Attenuated with extra shielding, and remainder vetoable by front tracking faces



• Primary neutral hadrons suppressed with additional shielding



(In practice Pb is not ideal for neutrons: other materials to be considered)

 ν-air inclusive inelastic σ: ~ O(3) events, but actual fake rate likely much smaller.

Dean Robinson dean.robinson@uc.edu

Secondary Backgrounds

 Muon or neutron secondary production in shield can be large! Vetoable – 'reducible' – by active veto in shield



 Active veto placed so that active veto eff/rejection rate is minimized, while neutral 'irreducible' BGs are suppressed



Geant4 $(20 + 5)\lambda$ simulation

Estimates validated with simplified propagation model using muon CSDA and kaon scattering length, and scattering cross-sections, from data

	Partic				
BG species	irreducible by shield veto	reducible by shield veto	Baseline Cuts		
$\overline{n+\overline{n}}$	7	$5\cdot 10^4$	$E_{\rm kin} > 1{ m GeV}$		
K_L^0	0.2	$9 imes 10^2$	$E_{\rm kin} > 0.5{\rm GeV}$		
$\pi^{\pm} + K^{\pm}$	0.5	$3\cdot 10^4$	$E_{\rm kin} > 0.5{\rm GeV}$		
$\nu + \overline{\nu}$	0.5	$2 \cdot 10^6$	E > 0.5 GeV		





Tracking Efficiencies

Implement a tracking simulation for above geometry:

- Six hits required for a track
- Assume sensitivity down to 600 MeV momentum

	$c\tau$ (m)	$m_{\varphi} \ [B o X_s \varphi]$		$m_{\gamma_{ m d}} \; [h o \gamma_{ m d} \gamma_{ m d}]$				Dominated by opening		
Dominated by assumption > 600 MeV. Needs proper imulation of turn-off.		0.5	1.0	2.0	0.5	1.2	5.0	10.0	20.0	angle resolution. Requires optimization using station spacing and granularity
	0.05	-	-	-	0.39	0.48	0.50	_	_	
	0.1	-	-	-	0.48	0.63	0.73	0.14	-	
	1.0	0.71	0.74	0.83	0.59	0.75	0.82	0.84	0.86	
	5.0	0.55	0.64	0.75	0.60	0.76	0.83	0.86	0.88	
	10.0	0.49	0.58	0.74	0.59	0.75	0.84	0.86	0.88	
	50.0	0.38	0.48	0.74	0.57	0.75	0.82	0.87	0.88	
	100.0	0.39	0.45	0.73	0.62	0.77	0.83	0.87	0.89	
	500.0	0.33	0.40	0.75	-	_	-	-	-	

 Lesson: Proof-of-concept tracking effs are O(1). Can be further optimized.

Boost Resolution

• Reconstruct parent boost from the measured decay vertex (no timing!), assuming 2-body decay with relativistic products (only need spatial info!)



- The resolution is < 1% dominated by distance to first measured point, not detector granularity
- Boost distribution is dominated by the spread of boosts, not resolution.

Boost and Mass Reco

Using only geometry of tracks, can reconstruct parent:



Mild resolution even for $b \rightarrow s\varphi!$

For $b \rightarrow s\varphi$, use time-of-flight to reconstruct LLP mass. Assume 100 ps and 50 ps resolution per hit. LLPs slow enough for mild mass reconstruction!



Dark glueballs Lifetime

