



25^e Congrès Général de la Société Française de Physique



Étude de l'équation d'état de la matière nucléaire : INDRA-FAZIA, un dispositif expérimental innovant au GANIL

> M. Henri Pour les collaborations INDRA & FAZIA



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INDRA & FAZIA collaborations



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- Nuclear matter:
 - infinite system of interacting protons and neutrons
 - coulomb interaction and surface effects are neglected
- Describe by its Equation of State (EoS): $e(T,\rho,\delta)$
- Dense matter phase diagram:
 - composition of the nuclear matter in the different regions
 - astrophysical and cosmological implications







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- Dense matter phase diagram:
 - composition of the nuclear matter in the different regions
 - astrophysical and cosmological implications
- How to explore this diagram?
 - through Heavy Ion Collisions (HIC)
- Zone of interest:
 - → the one of the nucleus
 - degrees of freedom: the nucleons

Température (en MeV)









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Description of heavy ion collisions :

- Tools = transport models
- Goal = constrain the parameters of the models
 - constrain the EoS of asymmetric nuclear matter

 $^{\rm 112}Sn$ + $^{\rm 112}Sn$ à 50 MeV/nucleon



Ref: M. Colonna et al., Phys. Rev. C, 82, 054613 (2010)









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prediction











Multi-detector of charged particles « resident » at GANIL :

- → First experiment in 1993
- → Today, database of over 40 combinations projectile/target/energy
 - > from 15 to 150 MeV/nucleon
 - \rightarrow from ³⁶Ar + KCl to ¹⁹⁷Au + ¹⁹⁷Au
- → Still working (last experiment April/May 2019)







What is INDRA?

- 336 modules of detections (564 detectors in total)
- Three types of detector:
 - ionisation chamber (IC)
 - silicon (Si)
 - caesium iodide (CsI(TI))
- I detection module = 1 telescope
 - Si Csl(Tl)
 - IC Si Csl(Tl)
 - IC CsI(TI)
- → 17 rings, cylindrical symmetry around the beam axis
- → Wide angular coverage: 90% of 4π



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What are its performances?

- Detection of almost all the reaction products
- Low detection threshold (1 MeV/nucleon IC)
- Identification :
 - \rightarrow in Z up to Z=92
 - in Z & A − up to Z≤4 (1993), Z≤8 (2001)







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Improvements?

FAZIA !







New generation of charged particles detection array:

- → 10 years of R&D (detectors & electronics)
- → First experiment: June 2015, IsoFazia @ LNS Catania (Italie)
- Demonstrator installed in GANIL since the beginning of 2019



FaziaPre experimental setup, ^{40,48}Ca + ¹²C @ 25,40 MeV/nucleon (LNS Catania, Italy)





What is FAZIA?

- Modular multi-detector with a block configuration
- → 1 bloc = 16 telescopes Si (300 µm) Si (500 µm) CsI(Tl) (10 cm)
- Digital electronics, embedded under vacuum (11 cards per block)
 - control of the detectors high voltages, thresholds, ...
 - signals are sampled close to the detectors and their preamplifiers
 - slow control
- → Silicon(2x2 cm²):
 - * tilted with respect to the main crystal axis
 - > uniform resistivity (nTD process), uniform thickness
 - reverse mounted (particles enter by the low electric field)
 to enhance the signal dynamic range





He

n

What are its performances?

- Identification
 - in Z up to Z=54 (heavier projectile used up to now)
 - > in Z & A up to Z \leq 20-25 (depending on the reaction)
- Spatial resolution
 - high granularity : 16 2x2 cm² telescopes
 - very good resolution for correlation studies
- Digital electronics
 - Pulse Shape Analysis (PSA) in all detectors
 - Iower the detection threshold







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What could become a problem:

→ 1 block: (70x10x10) cm³ & 15 kg







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INDRA

Wide angular coverage





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- to know their charge Z and mass A







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What is INDRA-FAZIA?

- INDRA: rings 6 to 17 (14° to 178°)
 - \sim Si CsI(Tl) up to ring 9 (45°)
 - CsI(TI) from ring 10 to 17



- FAZIA demonstrator (12 blocks):
 - Si Si CsI(Tl) from 1.5° to 12.5°







Acquisition coupling:

To construct the physical events we need a universal clock – the CENTRUM (10 ns cadenced clock)



CENTRUM







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Trigger conditions fulfilled – request of a timestamp





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Expected performances (example)?

- Thanks to the INDRA wide angular coverage:
 - information all along the velocity range
- Thanks to the FAZIA isotopic identification:
 - big enhancement of the IMF (Intermediate Mass Fragments) identified fraction respect to INDRA solo







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INDRA-FAZIA is thus, a unique tool to study tiny effects that can be linked to the nuclear matter EoS !!



2018

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Before conclusion, nothing better than experimental data!!

- → E789, first INDRA-FAZIA experiment in GANIL
- Probe the density dependance of the nuclear matter EoS
- → ^{58,64}Ni + ^{58,64}Ni @ 32,52 MeV/nucleon
- → April/May 2019
- 38 BTUs of data taking
- ✤ 8 systems studied, at least 30M events for each one
- Data reduction in progress





2 4 6 8 10 12 14 16 18 20 22 24 Total Multiplicity (INDRA)

> Correlation between charge of fragments detected in FAZIA & violence of collisions (multiplicity) measured by INDRA

Typical Z & A identification obtained with FAZIA during the experiment

(FAZIA)

N





Conclusion

Nuclear matter EoS

- → Describe how nuclear matter evolves in various conditions (T, ρ , δ ,...)
- Important ingredients for astrophysics
- Can be studied through Heavy Ion Collisions (HIC)

INDRA-FAZIA detection arrays in GANIL: a unique tool to study the EoS

- wide angular coverage (80% of 4π)
- ➔ full isotopic identification (Z&A) in FAZIA
- mechanism selection (collisions sorting)





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Other topics of interest than "only" the EoS

- → Hot, dilute, asymmetric nuclear matter
 - > transport properties
 - > clusterization in dilute nuclear medium
- In medium nuclear structure and reactions





FAZIA @ NFS

INDRA & FAZIA are charged particles detectors: what about the neutrons?

- → Approved experiment for FAZIA @ NFS (SPIRAL 2) : E720
 - what is the neutron signal in the Si Si CsI(TI) telescopes of FAZIA?
 - > what is the efficiency of such process for the FAZIA telescopes?



- How to do that?
 - by looking at the signal observed in the CsI(TI) associated to a proton detection

Ref: FAZIA collaboration. Proposal for the GANIL PAC : Measurement of the absolute neutron detection efficiency of FAZIA telescopes. GANIL PAC, 2016

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