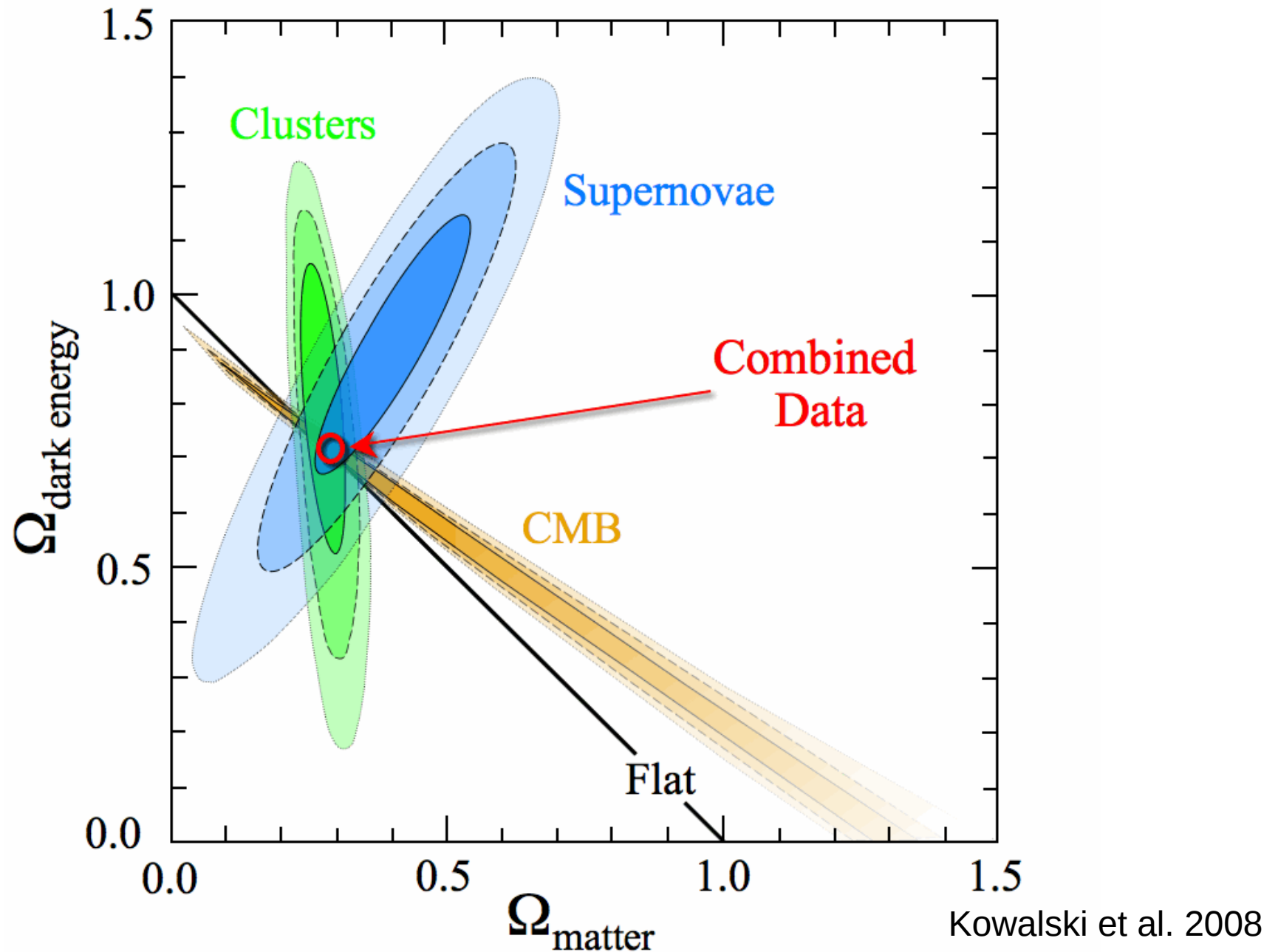


Colloque national Dark Energy
11-13 octobre 2017

The power of probe combination : cosmology with LSS and CMB

Stéphane Ilić
(IRAP/CEICO)

Why combining data ?



Which datasets to combine ?

Which datasets to combine ?

“Categories” of probes :

	“Background evolution” (i.e. expansion)	“Perturbed sector” (i.e. linear growth of structures)
Early Universe		
Late Universe		

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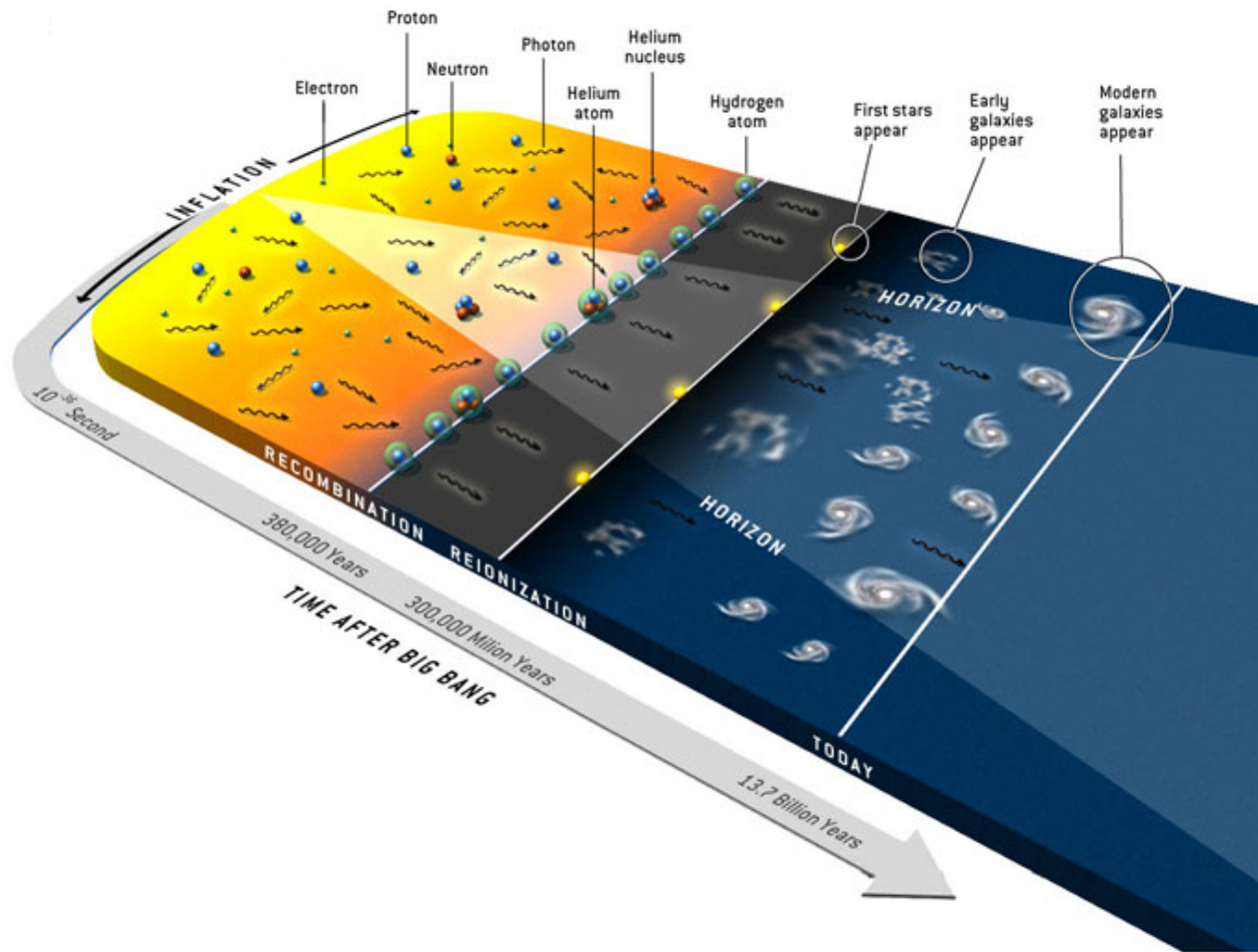
	“Background evolution” (i.e. expansion)	“Perturbed sector” (i.e. linear growth of structures)
Early Universe	Primordial CMB (T & P)	Primordial CMB (T & P)
Late Universe	Standard candles : BAO, SN, Cepheids	LSS surveys : Weak lensing, galaxy clustering (+ Secondary CMB)

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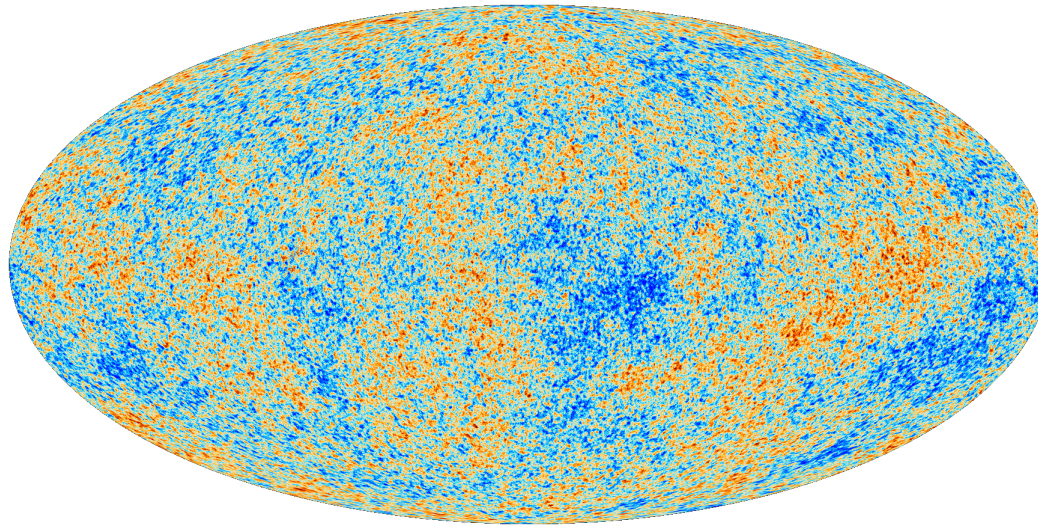
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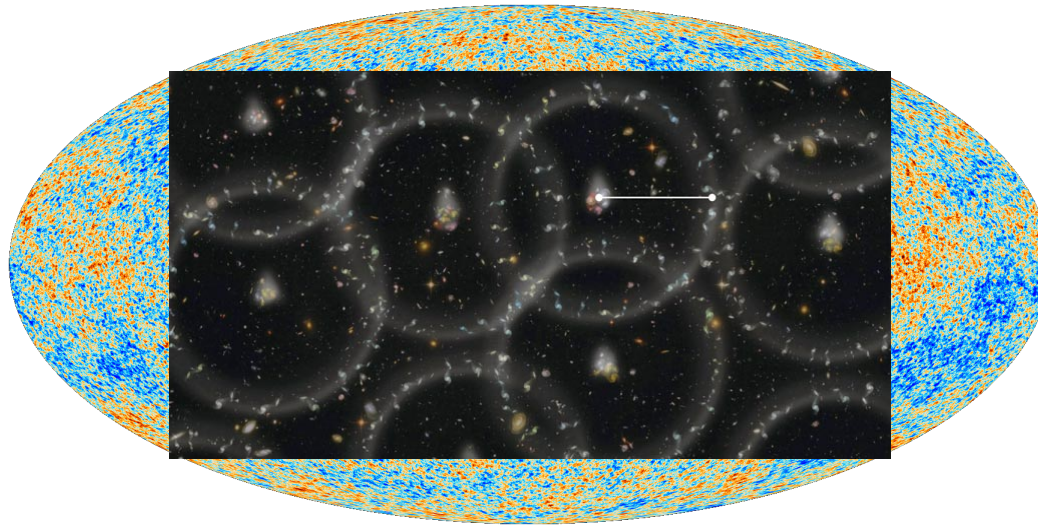
(e.g.  **planck** +  **euclid**)

I – Combining CMB & LSS data

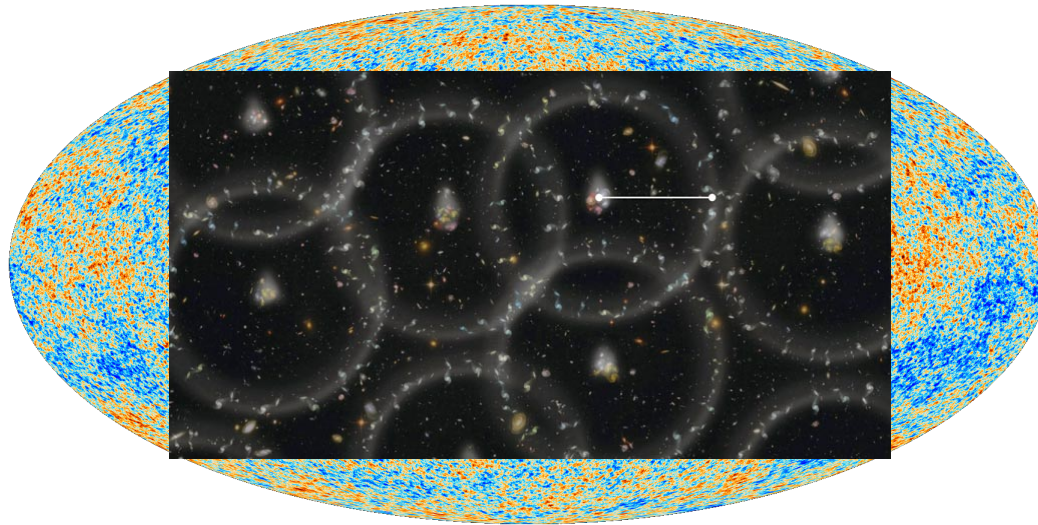
How to combine datasets ?



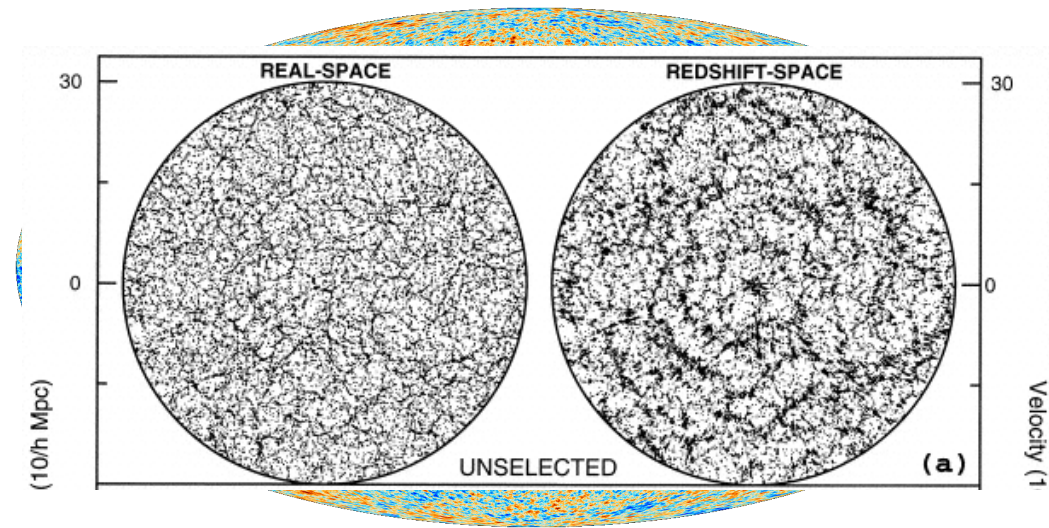
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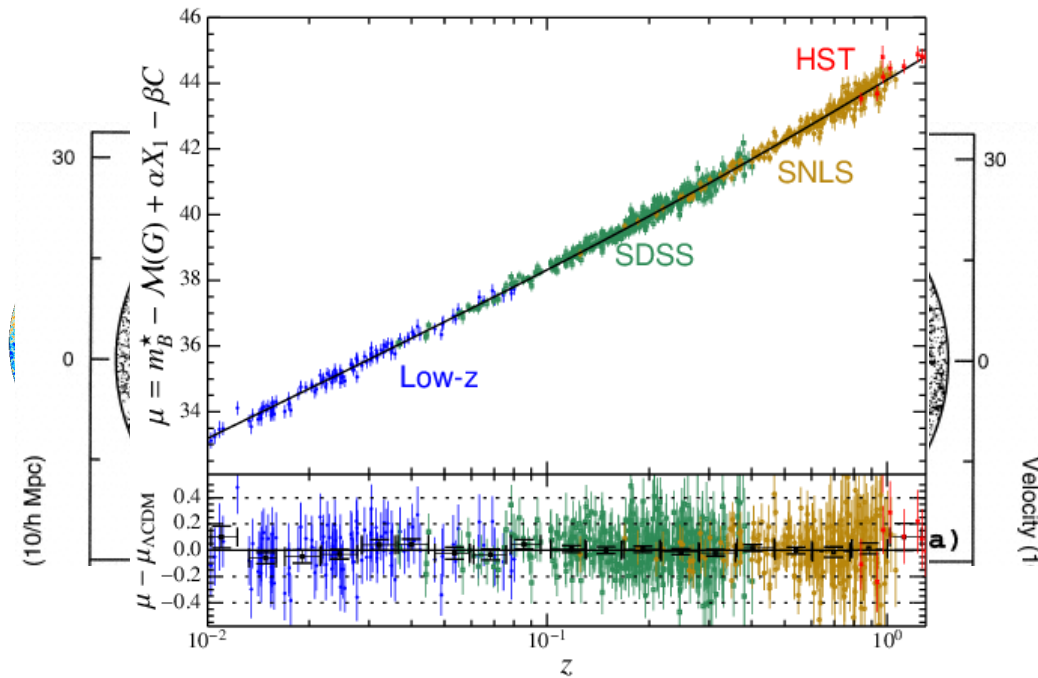
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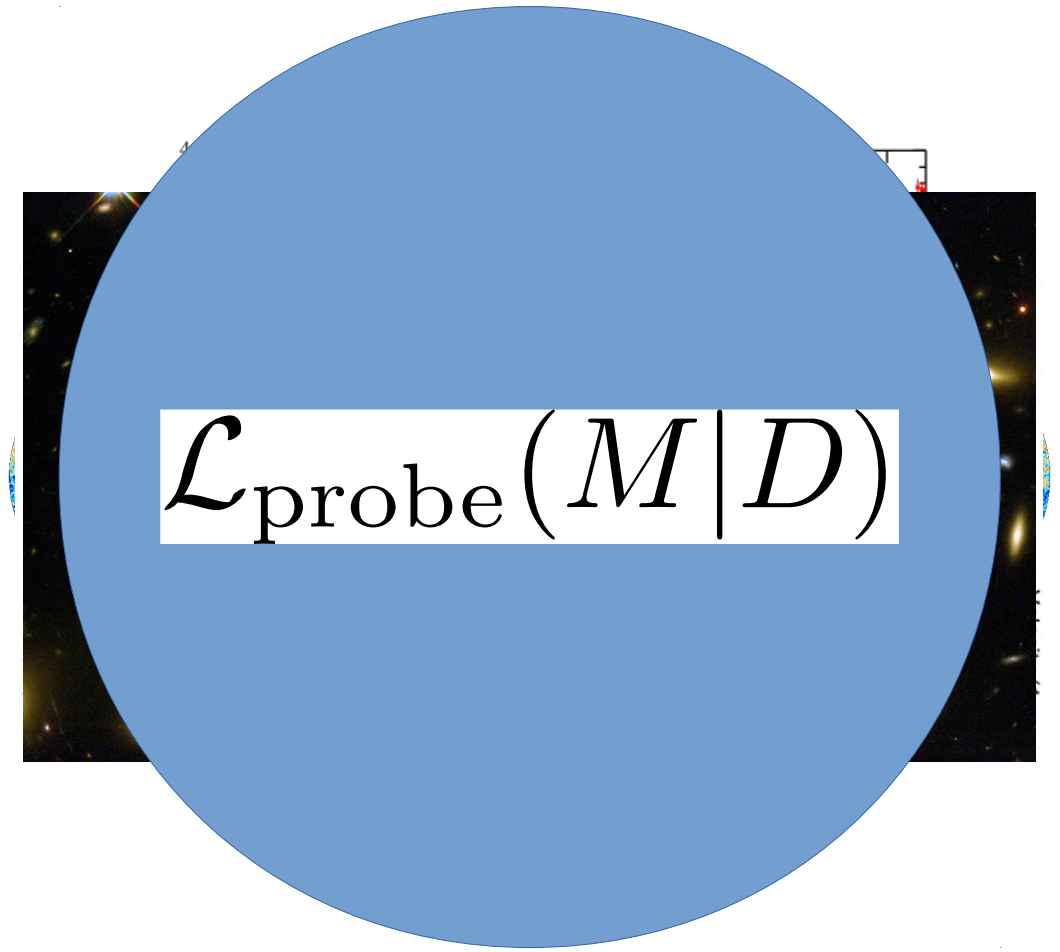
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$$\mathcal{L}_{\text{probe}}(M|D)$$

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(correlations ~ 0)

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Fisher formalism :

- Choice of “fiducial” model : $\Theta_{\text{fid}} = \{\theta_{1,\text{fid}}, \theta_{2,\text{fid}}, \theta_{3,\text{fid}}, \dots\}$

- Assumption : likelihood $\sim$ Gaussian :

$$\mathcal{L} \propto \exp \left[-\frac{1}{2} (\Theta - \Theta_{\text{fid}})^T \mathcal{F} (\Theta - \Theta_{\text{fid}}) \right]$$

- Fisher matrix :

$$\mathcal{F} = \begin{pmatrix} -\frac{\partial^2 \ln \mathcal{L}}{\partial \theta_1^2} \Big|_{\text{fid}} & -\frac{\partial^2 \ln \mathcal{L}}{\partial \theta_1 \partial \theta_2} \Big|_{\text{fid}} & \dots \\ -\frac{\partial^2 \ln \mathcal{L}}{\partial \theta_1 \partial \theta_2} \Big|_{\text{fid}} & -\frac{\partial^2 \ln \mathcal{L}}{\partial \theta_2^2} \Big|_{\text{fid}} & \\ \vdots & & \ddots \end{pmatrix} = \mathbf{C}^{-1}$$

Forecasted errors
on parameters

- Then...

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- Then...

$$\begin{aligned} \mathcal{L}_{\text{probe1+probe2}} &= \mathcal{L}_{\text{probe1}} \times \mathcal{L}_{\text{probe2}} \\ \mathcal{F}_{\text{probe1+probe2}} &= \mathcal{F}_{\text{probe1}} + \mathcal{F}_{\text{probe2}} \end{aligned}$$

Current + “future” datasets ?

Question : how to make forecasts for next-gen LSS surveys + CMB ?

Current + “future” datasets ?

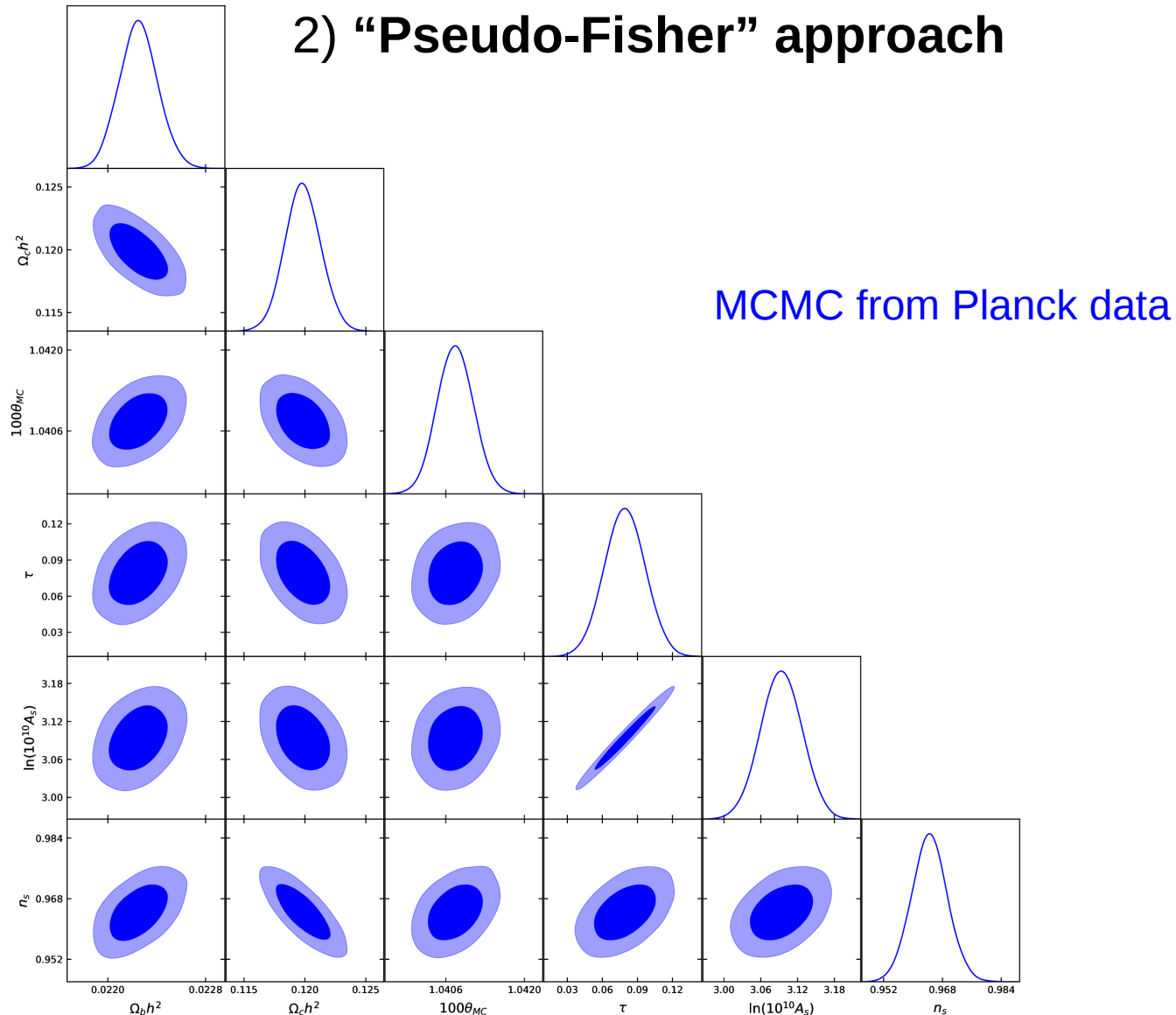
Question : how to make forecasts for next-gen LSS surveys + CMB ?

- 1) Choose a fiducial, compute a “fake Planck” Fisher
→ but : ignores the actual data (and fiducial !)

Current + “future” datasets ?

Question : how to make forecasts for next-gen LSS surveys + CMB ?

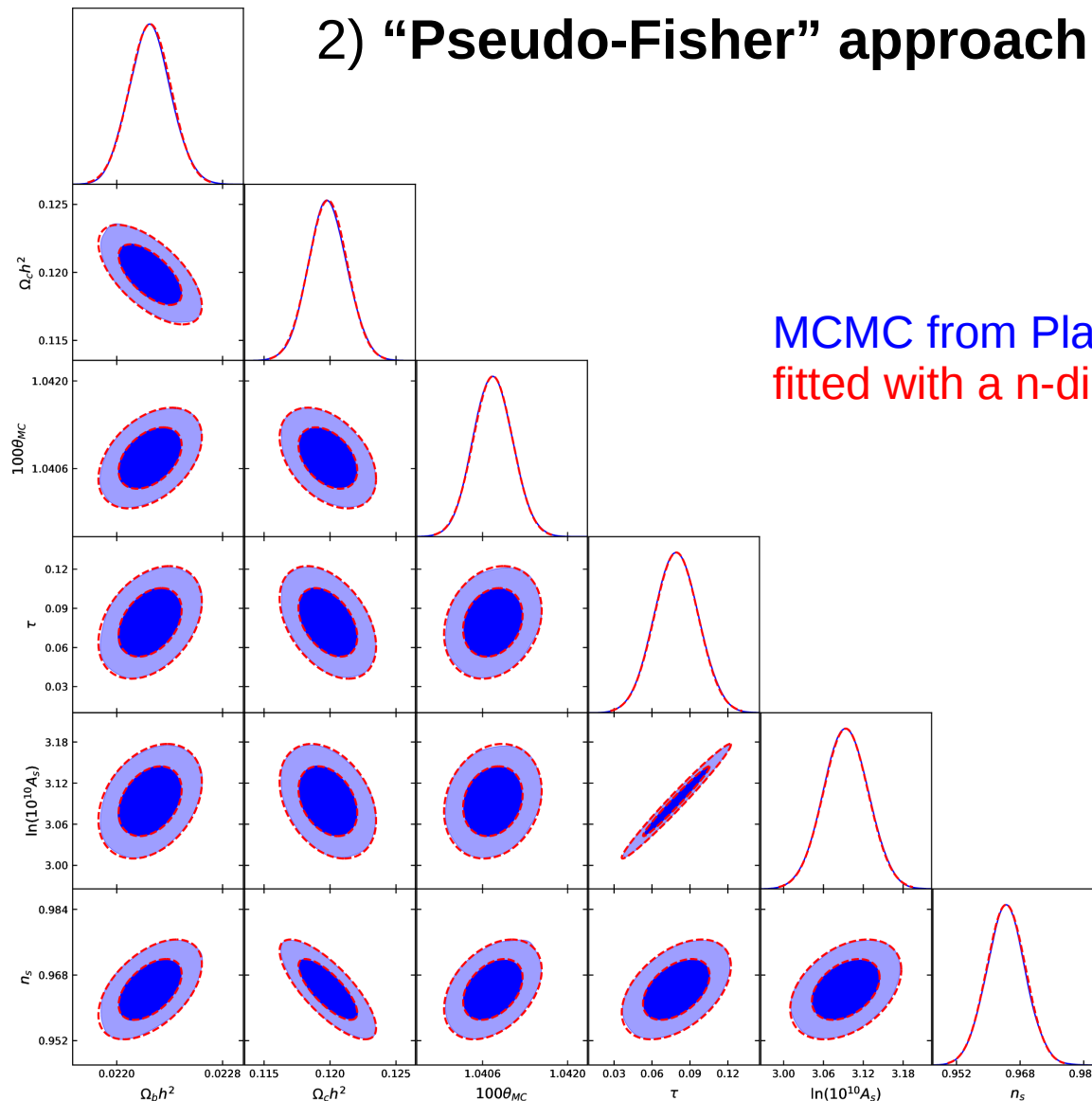
2) “Pseudo-Fisher” approach



Current + “future” datasets ?

Question : how to make forecasts for next-gen LSS surveys + CMB ?

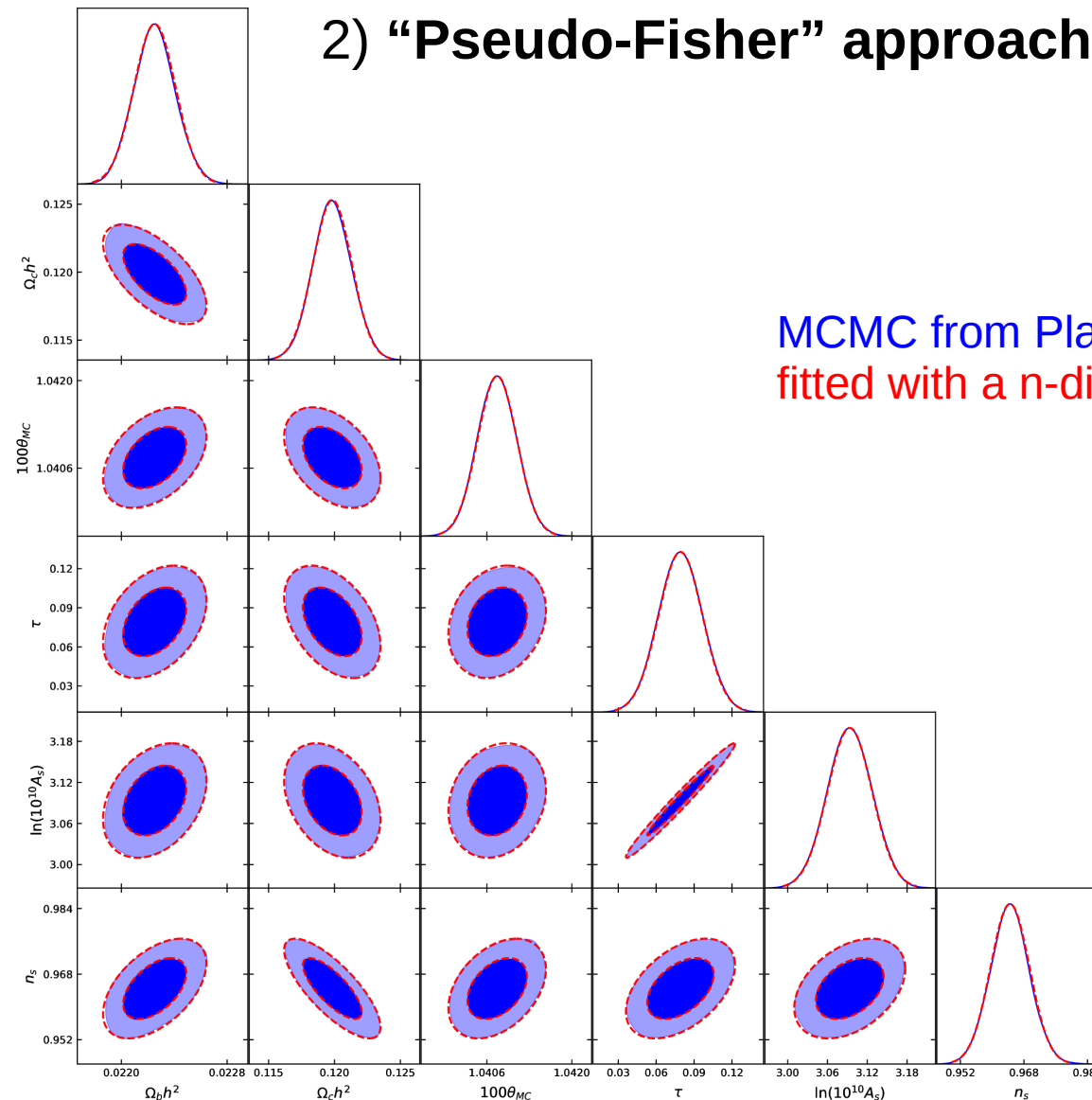
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Current + “future” datasets ?

Question : how to make forecasts for next-gen LSS surveys + CMB ?

2) “Pseudo-Fisher” approach



MCMC from Planck data
fitted with a n-dimensional Gaussian

$$C^{-1} = \mathcal{F}$$

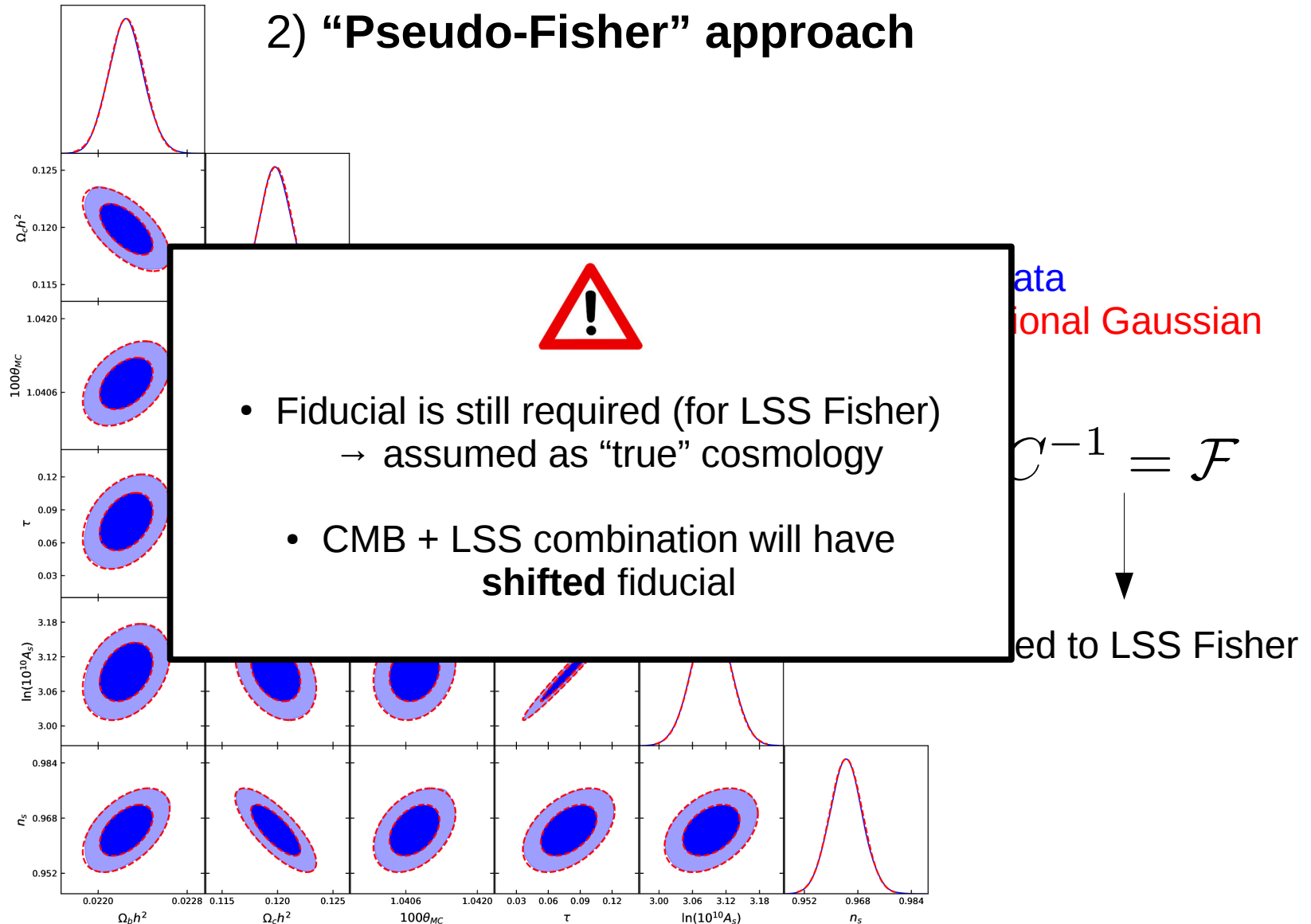


Added to LSS Fisher

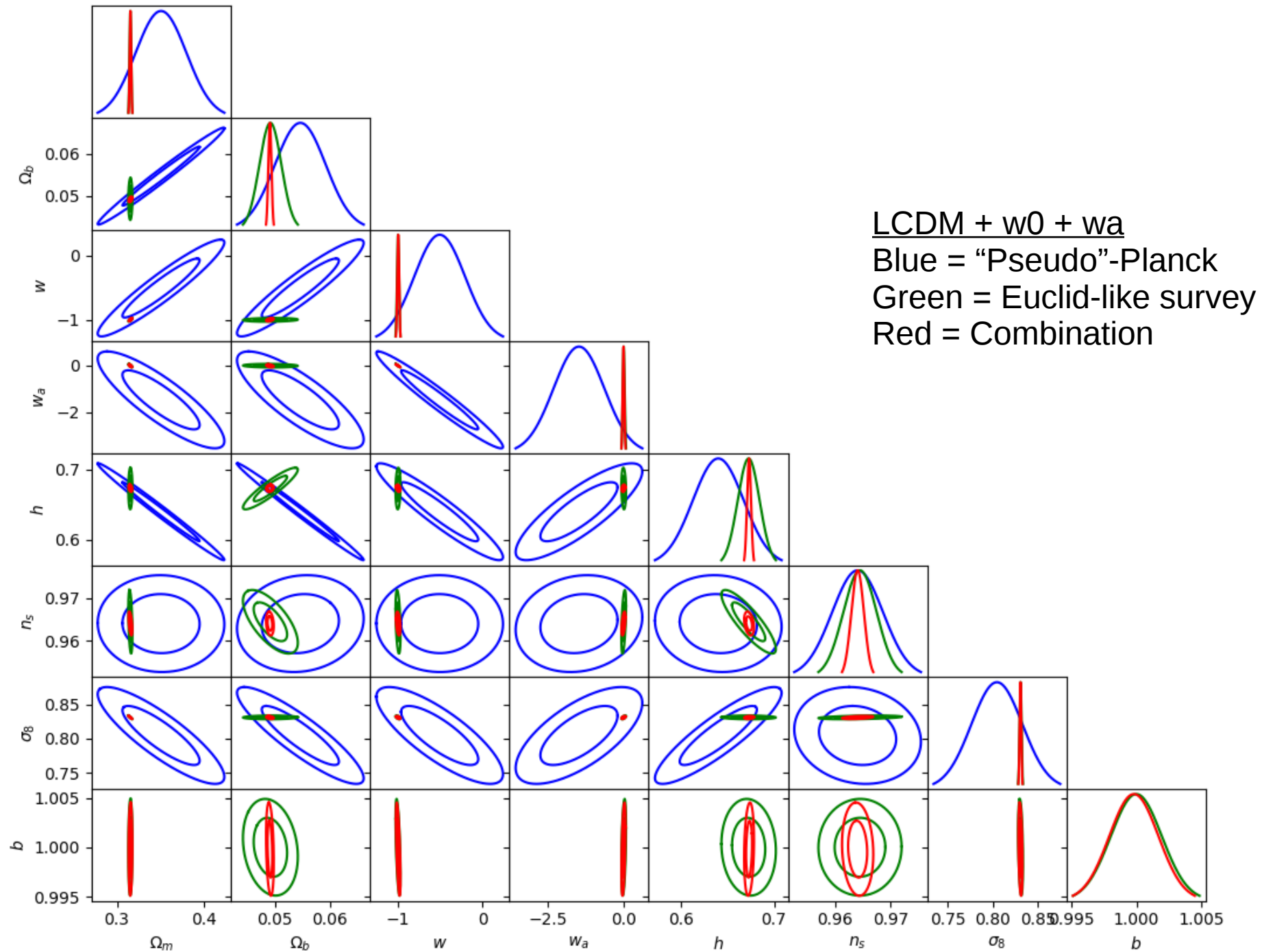
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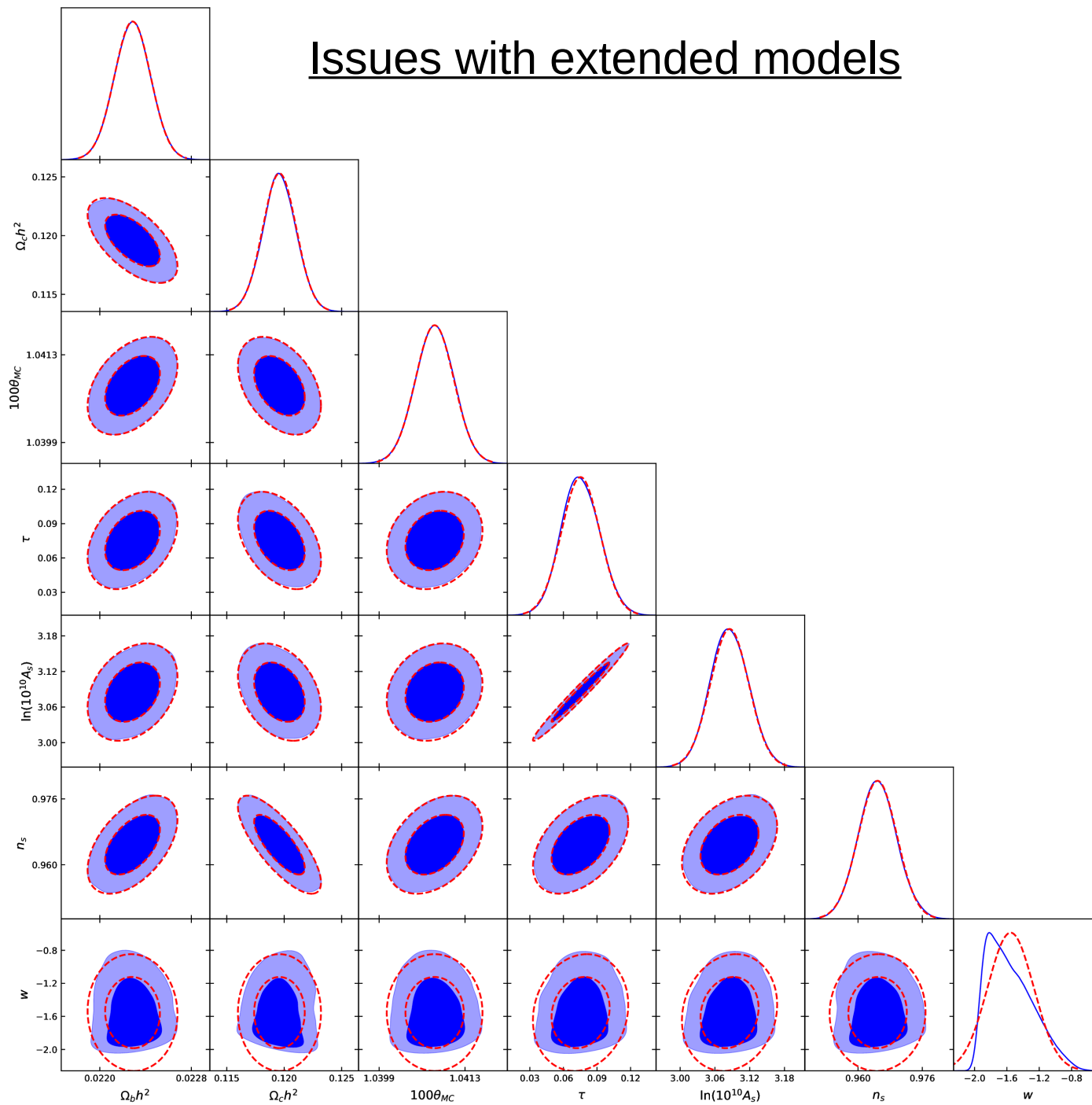
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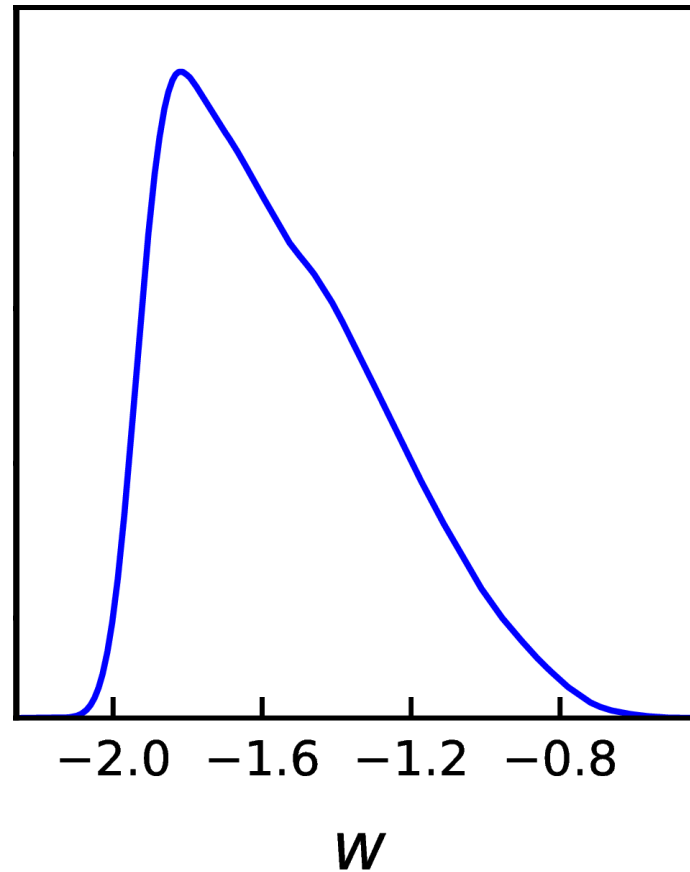


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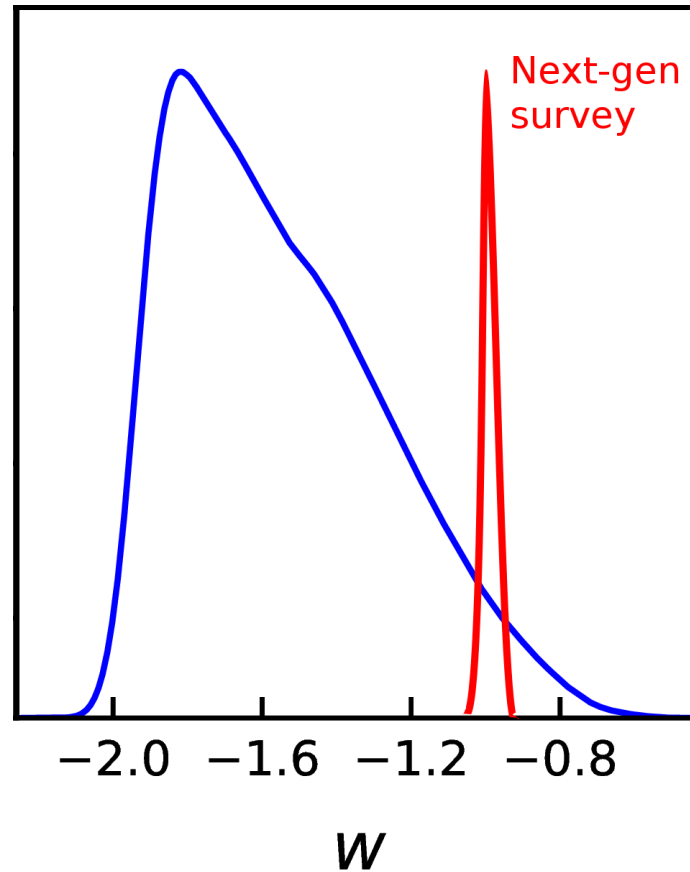
Current + “future” datasets ?

Solution :



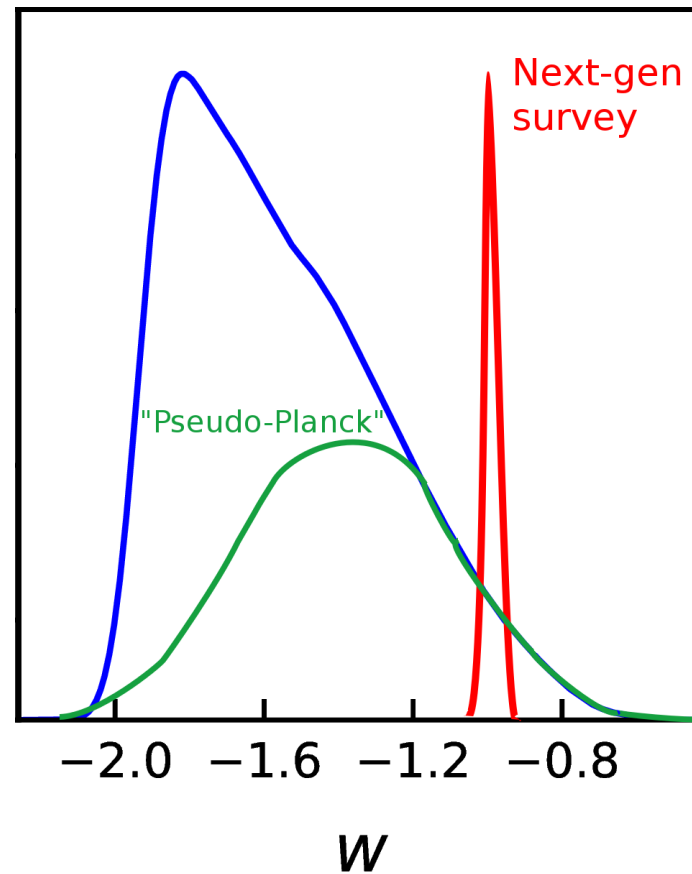
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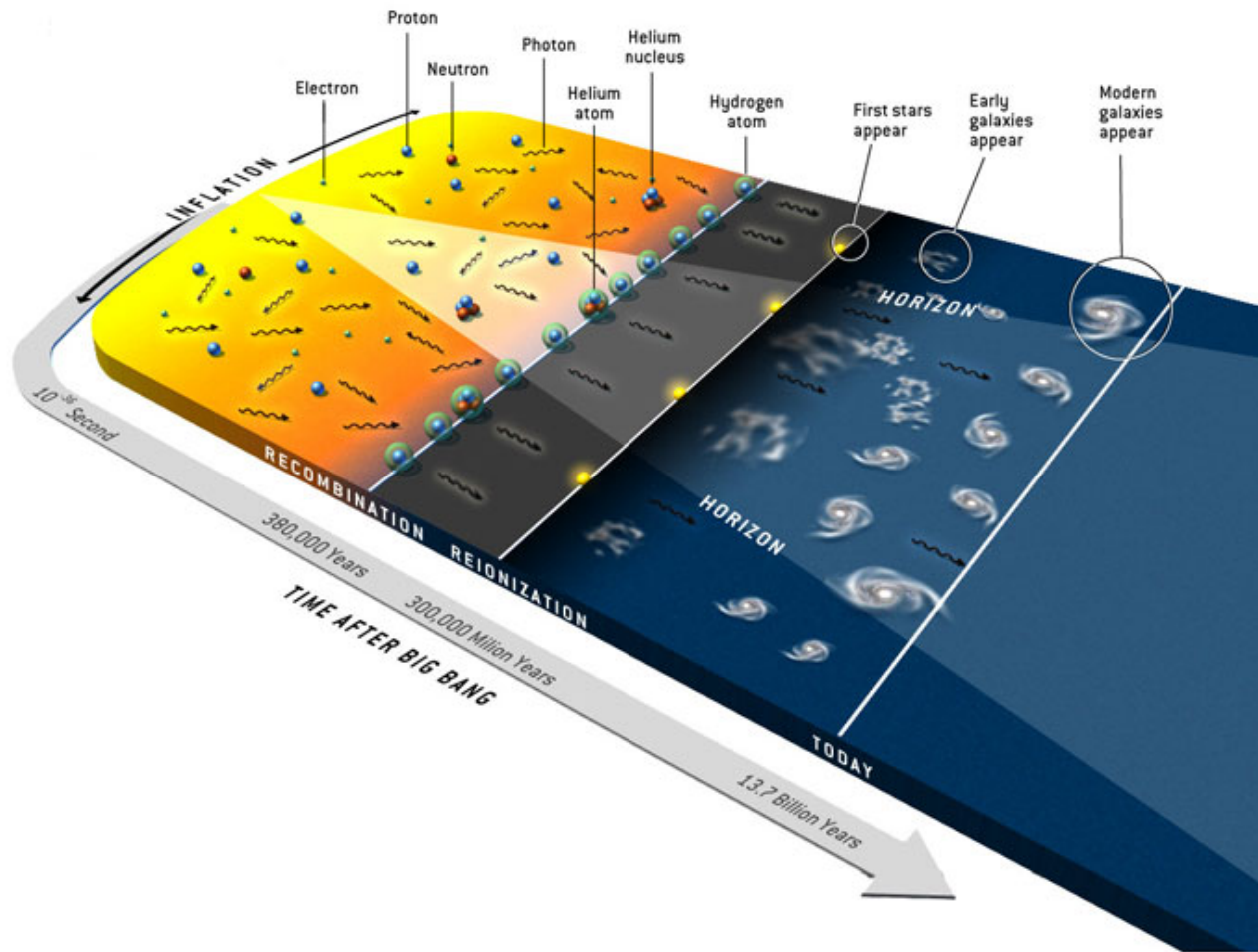
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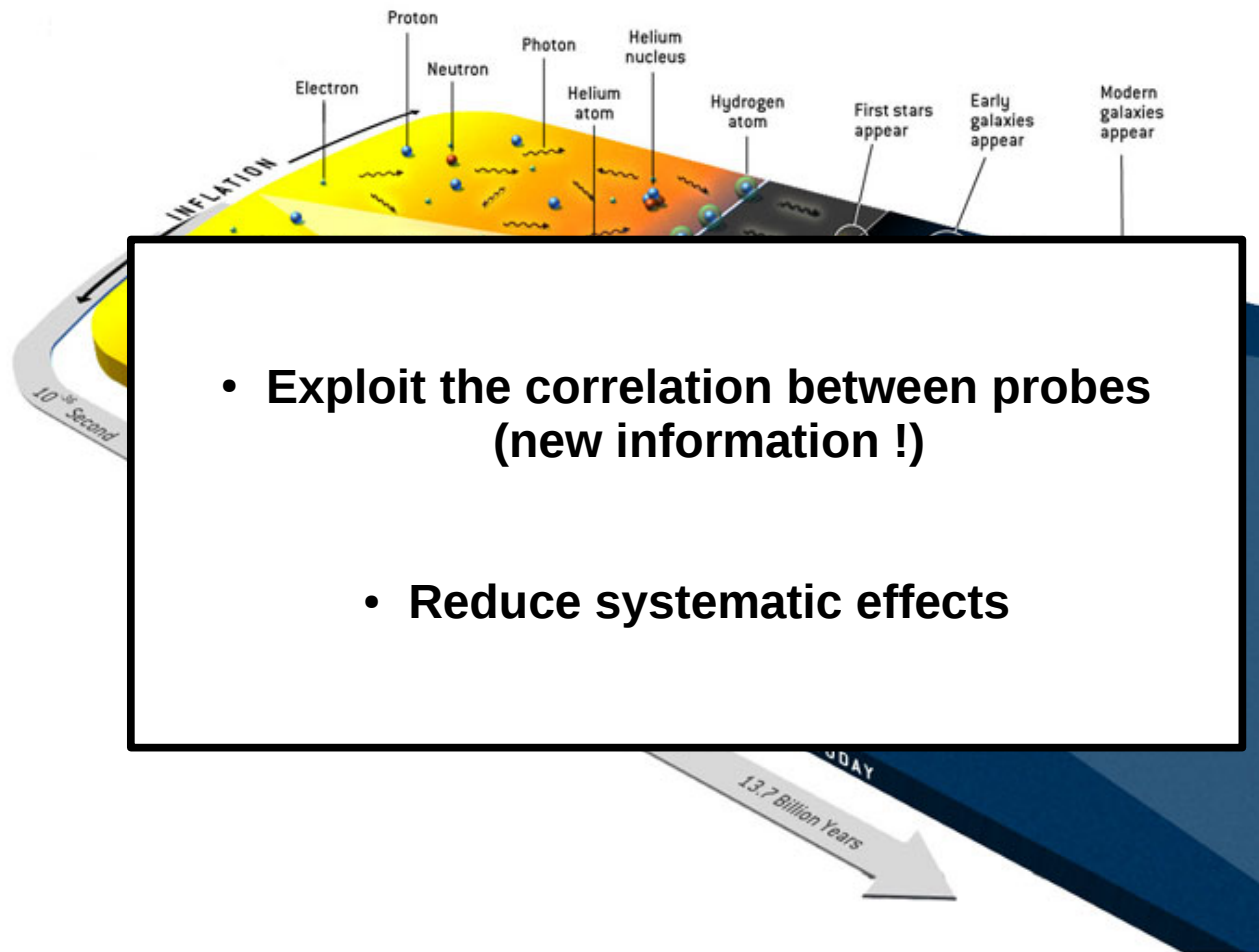


II – Cross-correlating CMB & LSS data

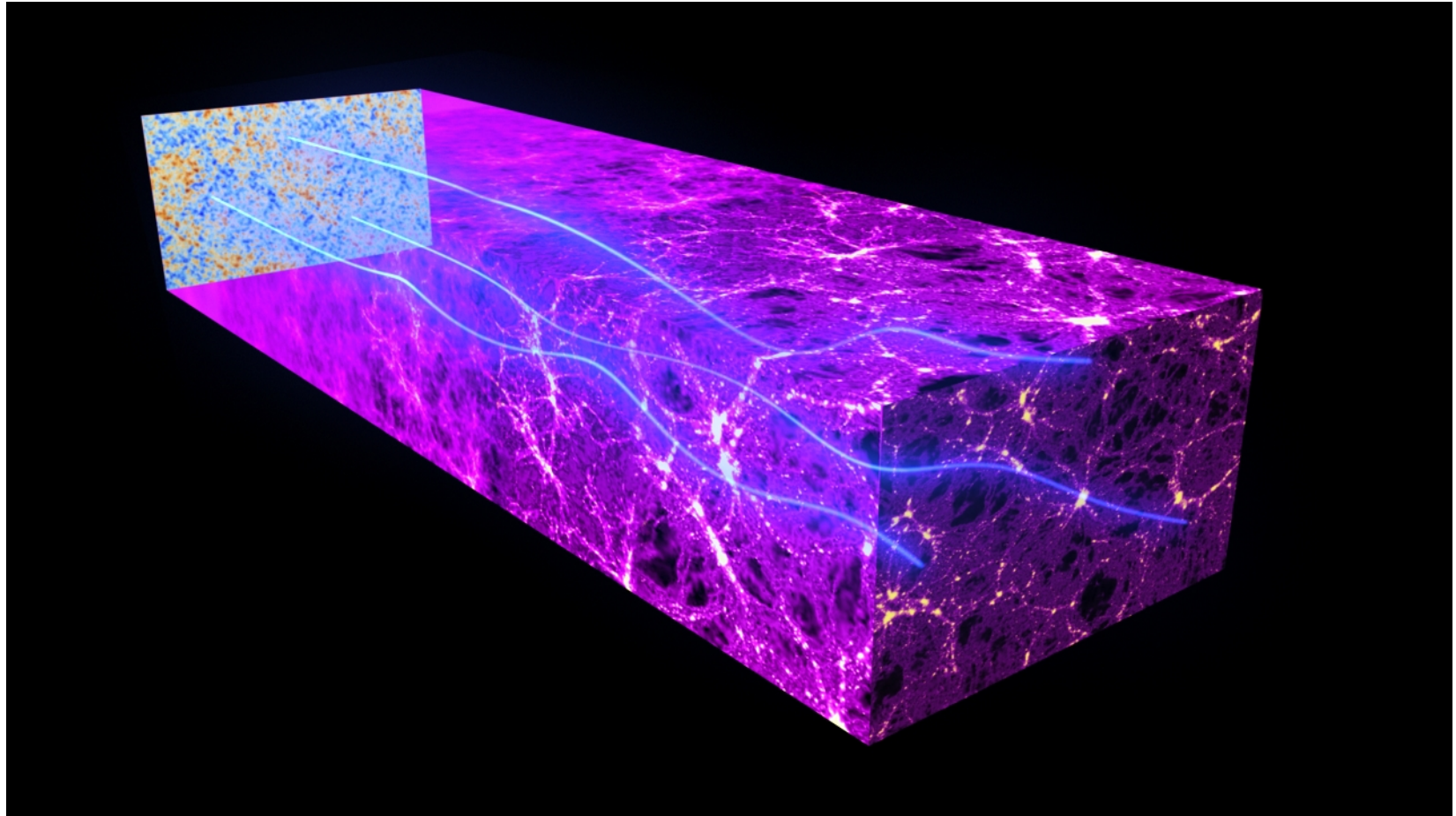
Why cross-correlating ?



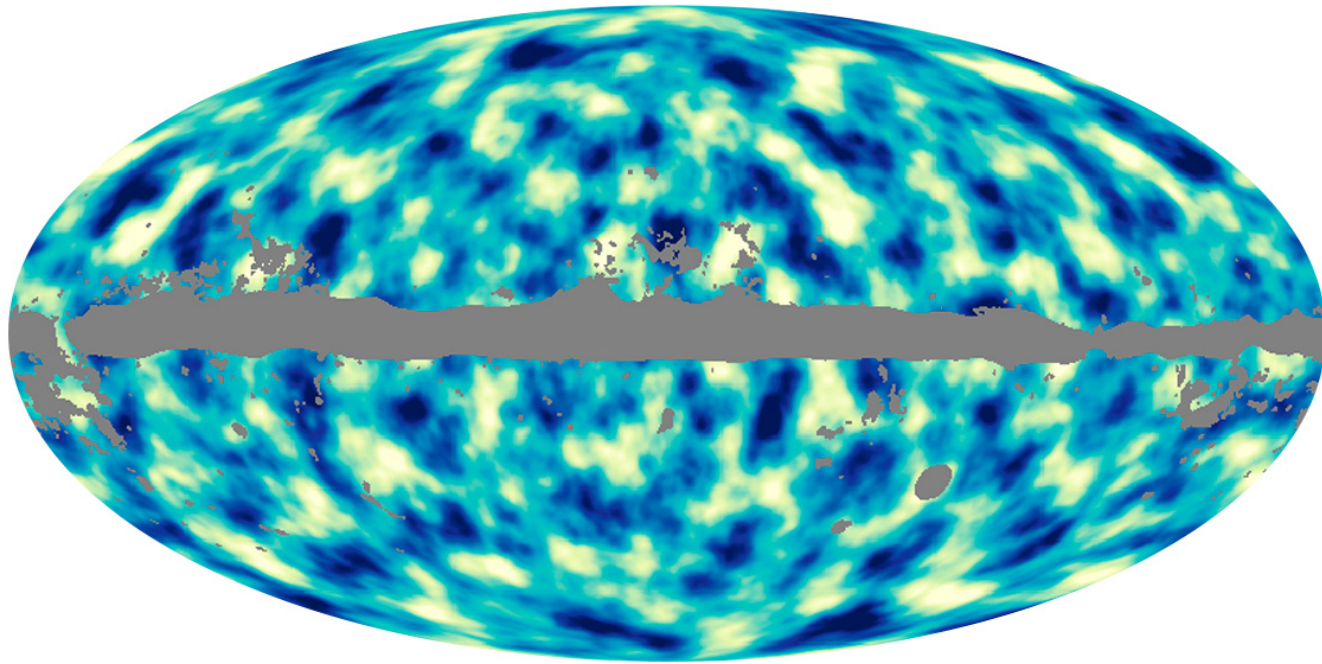
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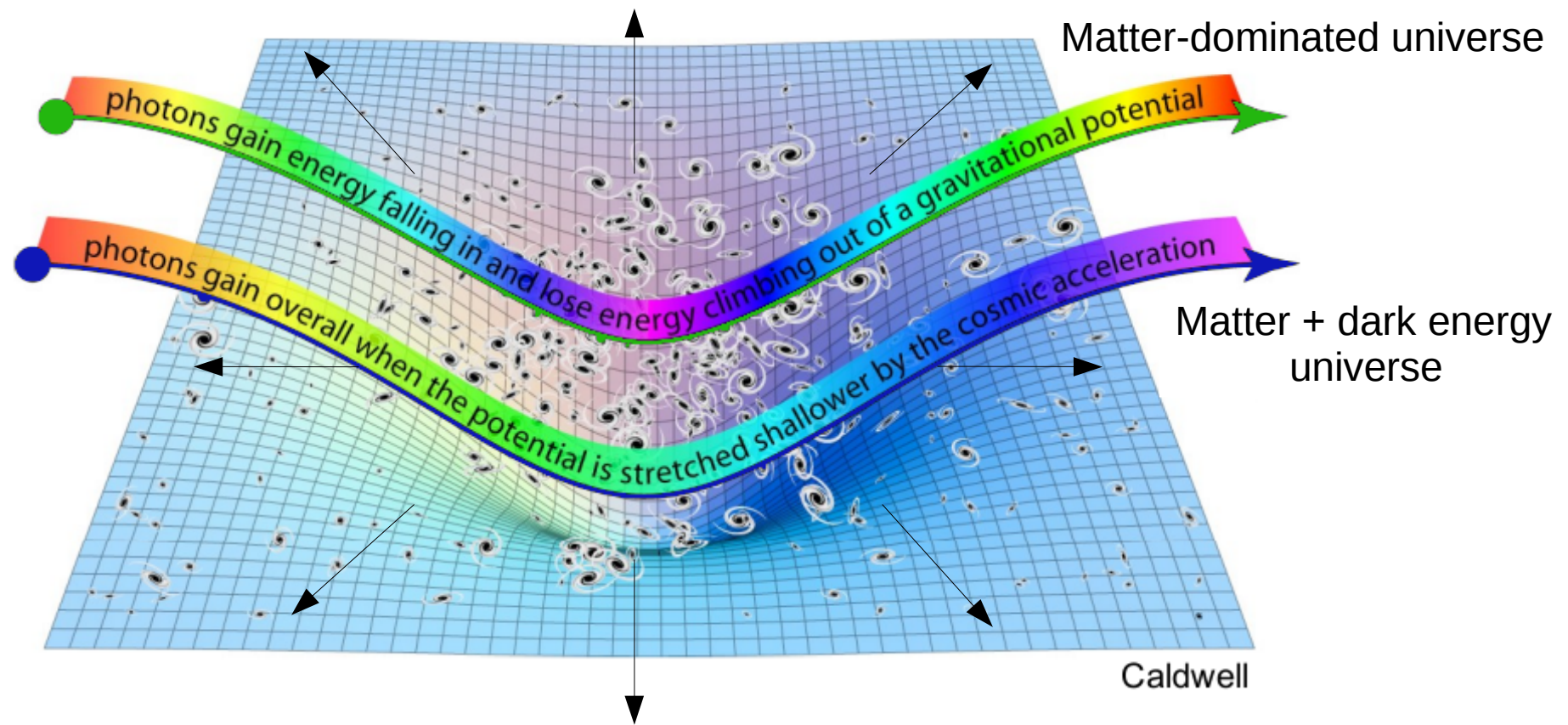
CMB lensing



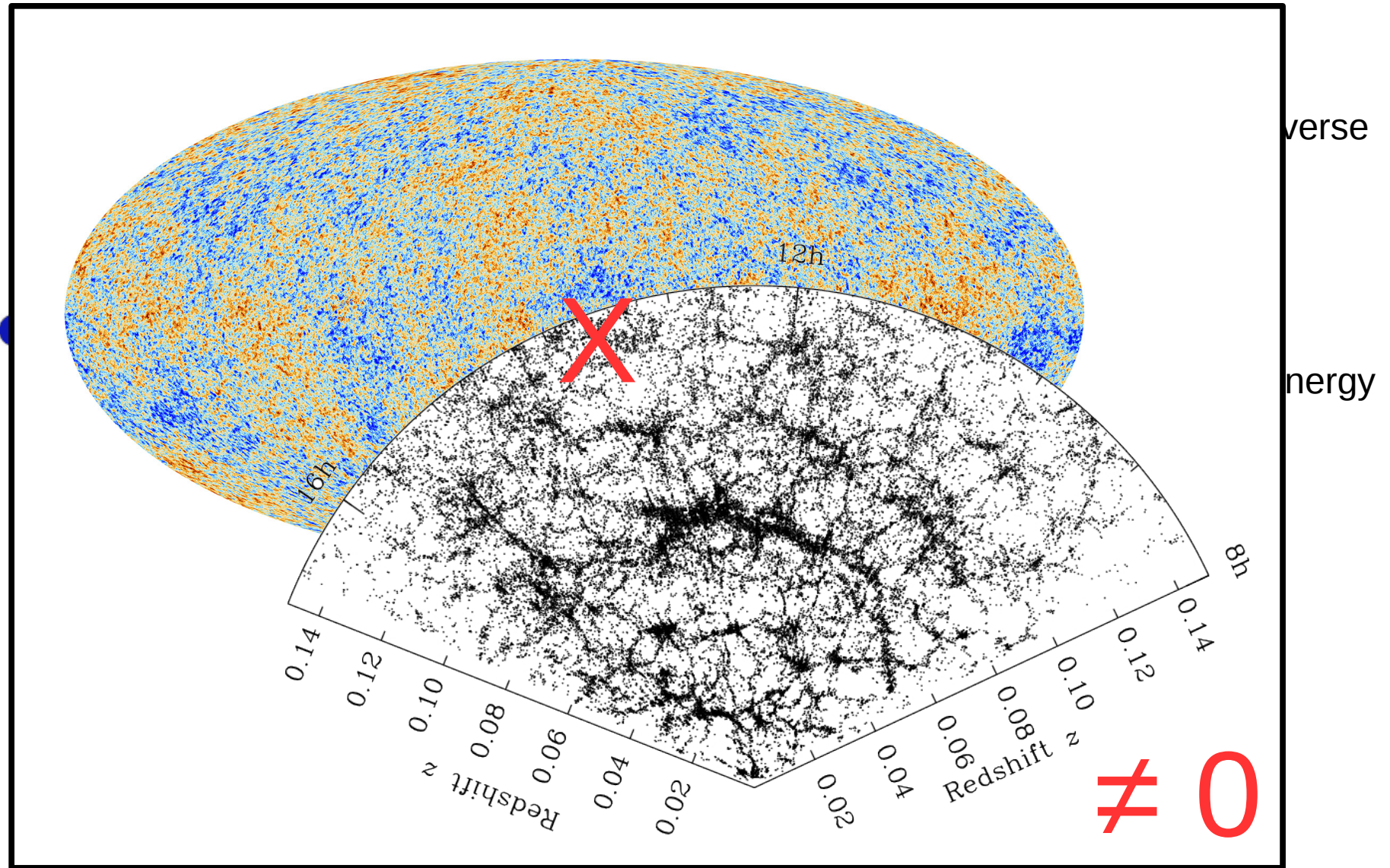
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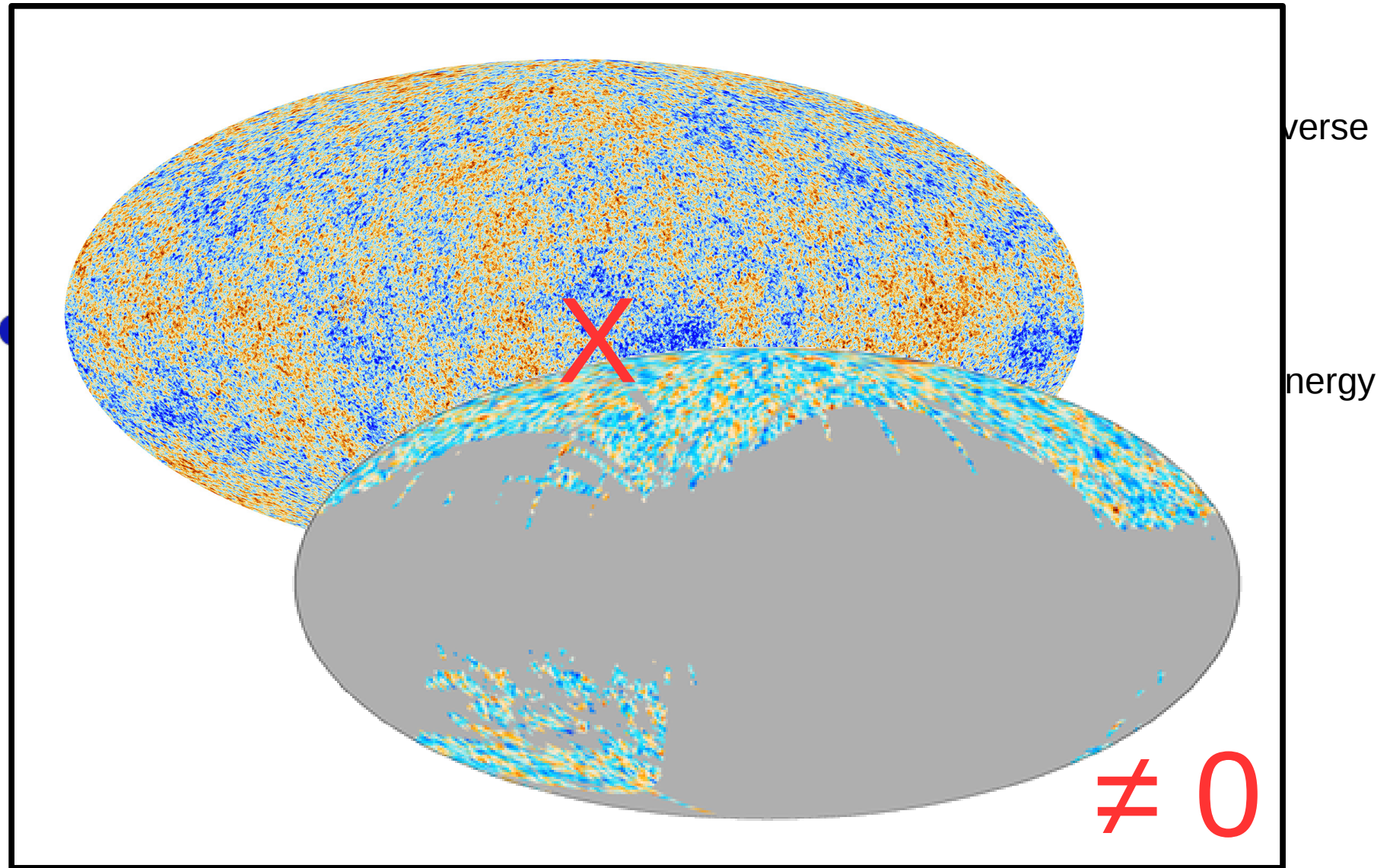
The integrated Sachs-Wolfe effect



The integrated Sachs-Wolfe effect



The integrated Sachs-Wolfe effect



Strong probe of growth of structures

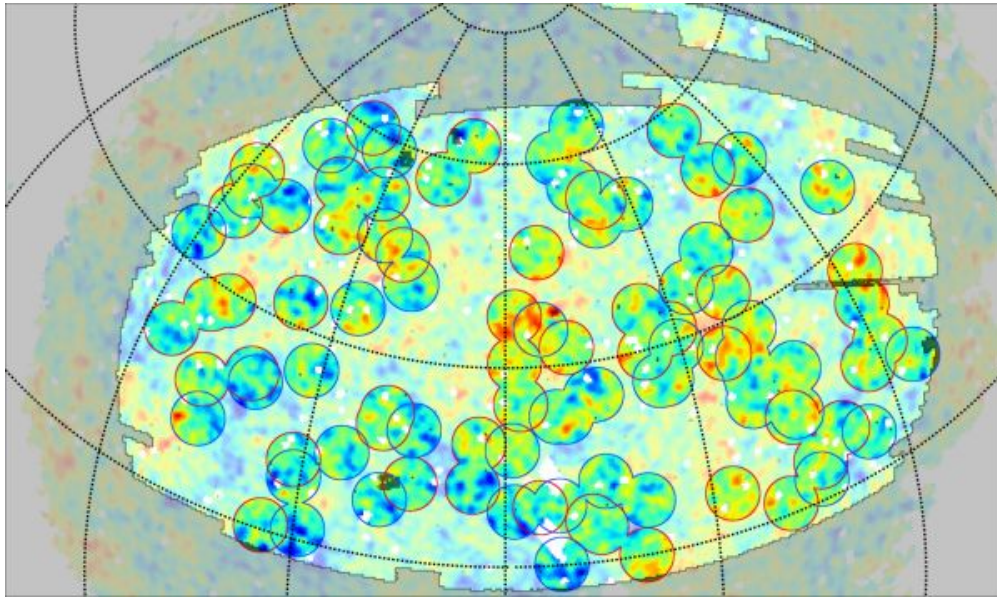
The integrated Sachs-Wolfe effect

From arXiv:1707.02263

Galileon Gravity in Light of ISW, CMB, BAO and H0 data

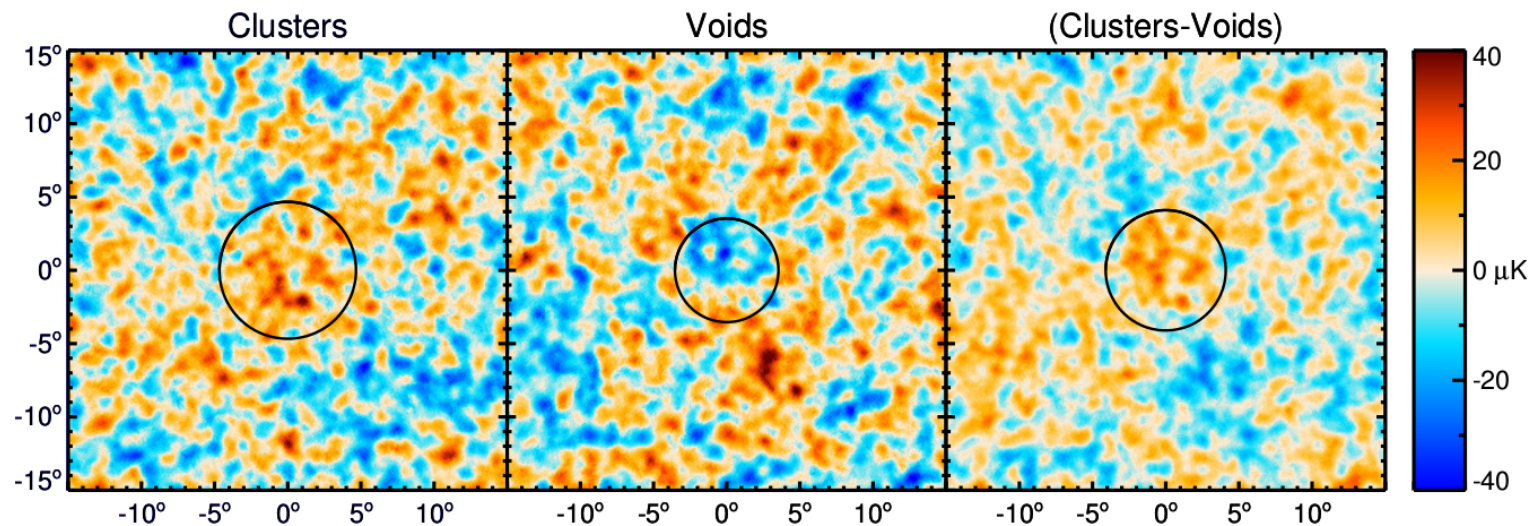
the galaxy sample. It is positive if the potential decays (like in Λ CDM), negative if it deepens. We constrain three subsets of Galileon gravity separately known as the Cubic, Quartic and Quintic Galileons. The cubic Galileon model predicts a negative C_ℓ^{Tg} and exhibits a 7.8σ tension with the data, which effectively rules it out. For the quartic and quintic models the ISW data also rule out a significant portion of the parameter space but permit regions where the goodness-of-fit is comparable to Λ CDM. The data prefers a non zero sum of the neutrino masses ($\sum m_\nu \approx 0.5\text{eV}$) with $\sim 5\sigma$ significance in these models. The best-fitting models have

The integrated Sachs-Wolfe effect

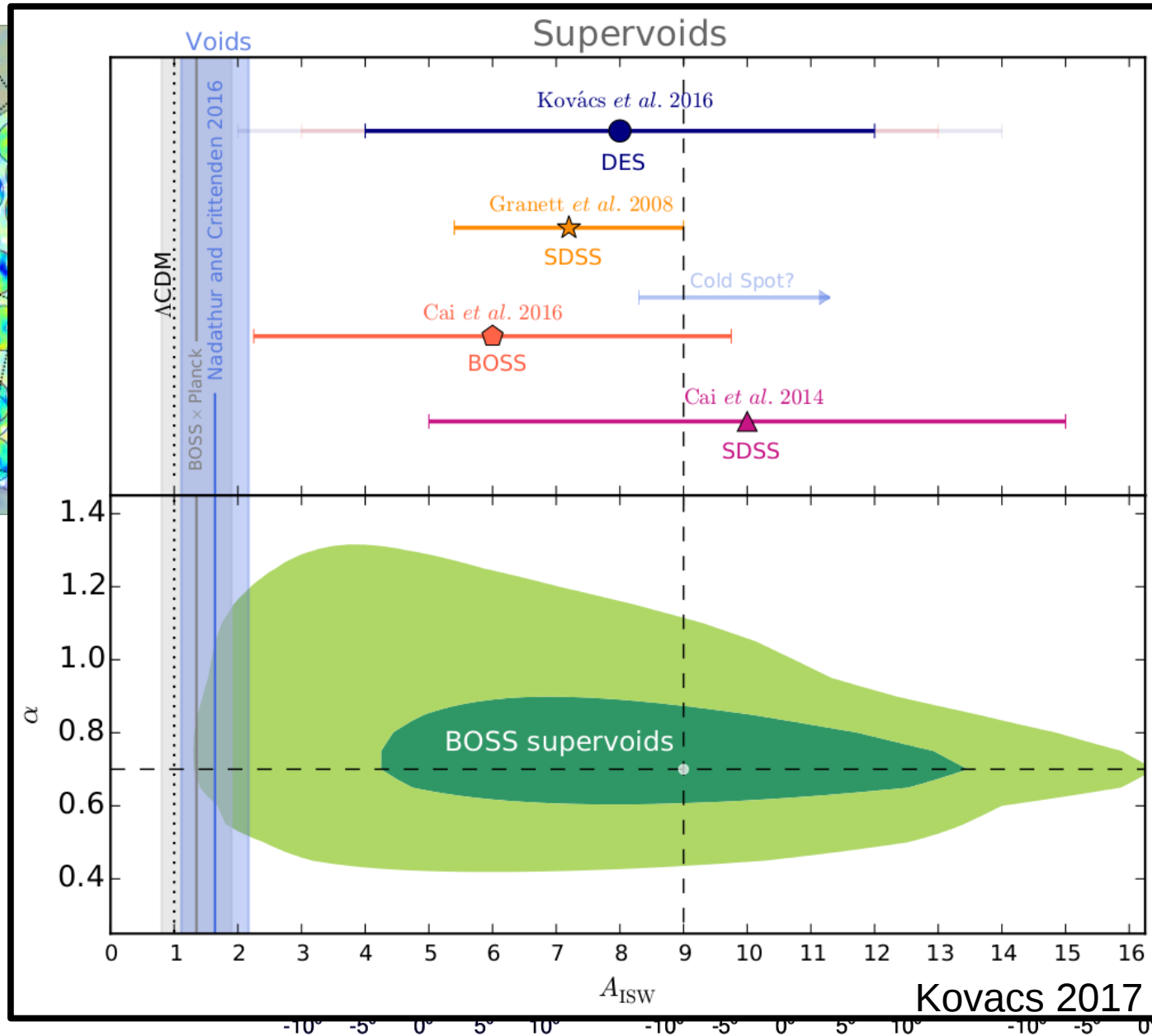
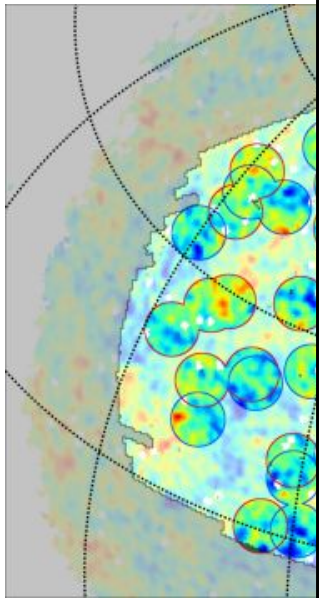


Stacking approach : ISW imprint of individual superstructures

- **(Planck 2013 & 2015, incl. Ilić)**
- Rigorous analysis of sample of 100 structures
- 3 - 4 sigmas signal
- Discrepancy with Λ CDM expectations



The integrated Sachs-Wolfe effect

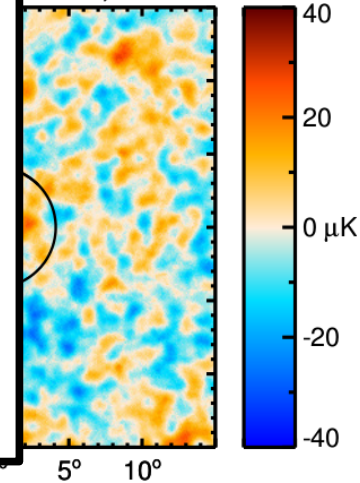


print of

cl. Ilić)

expectations

Voids)



**Thank you for
your attention !**

