

XENON dark matter constraints: examining the \mathcal{L}_{eff} dependence

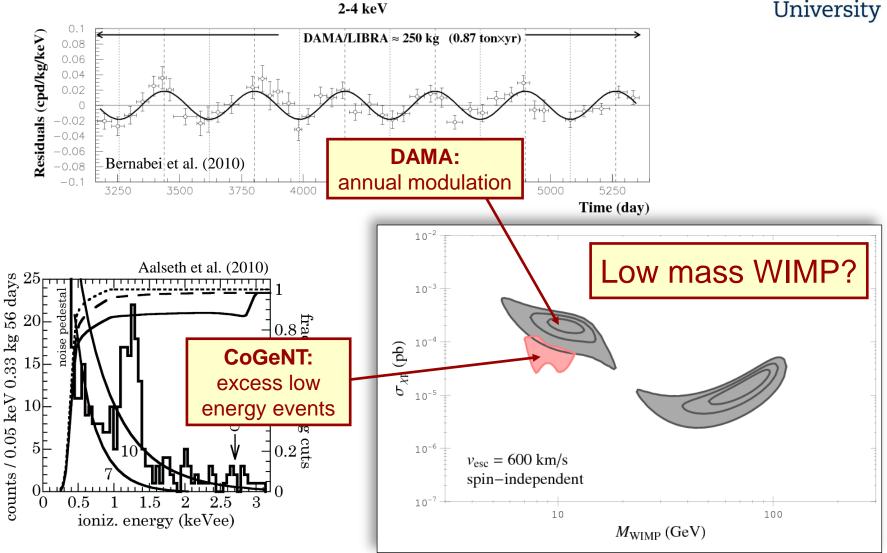
Chris Savage

Oskar Klein Centre for Cosmoparticle Physics Stockholm University

CS, Gelmini, Gondolo & Freese arXiv:1006.0972 + work in progress

Overview





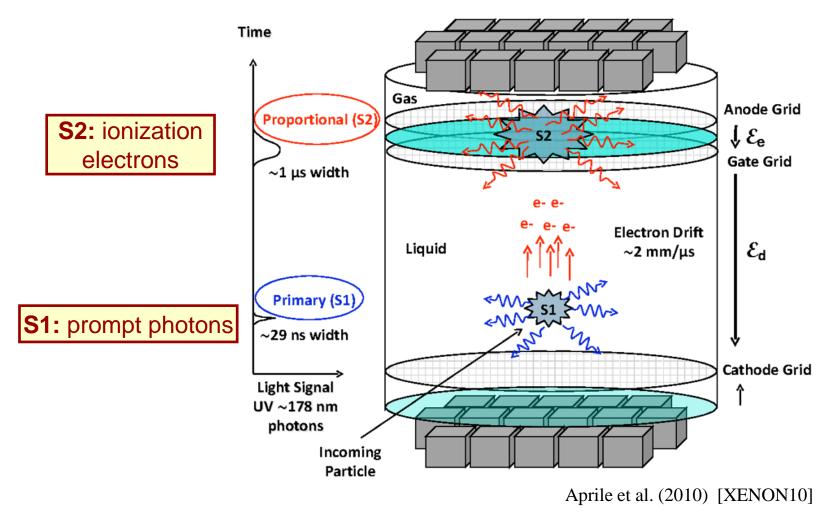
Overview



- What can liquid Xenon (or Argon) detectors tell us about low mass WIMPs?
- Two phase detectors
 - Two scintillation signals: S1 & S2 (allows for background rejection)
 - Analysis: must relate S1 & S2 to nuclear recoil energy E_{nr} ⇒ depends on relative scintillation efficiency factor \mathcal{L}_{eff}
- Outline
 - What is \mathcal{L}_{eff} ?
 - XENON10 and XENON100 constraints

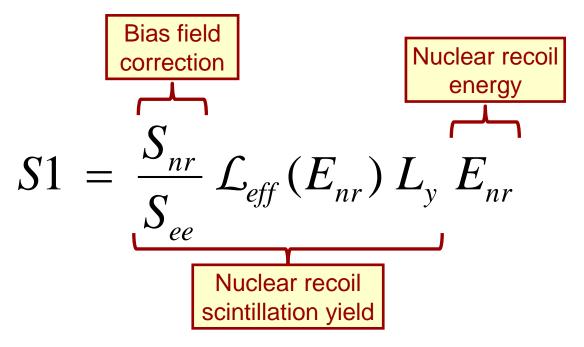
Liquid Xenon/Argon detectors: Concertion





\mathcal{L}_{eff} : Relative scintillation efficiency

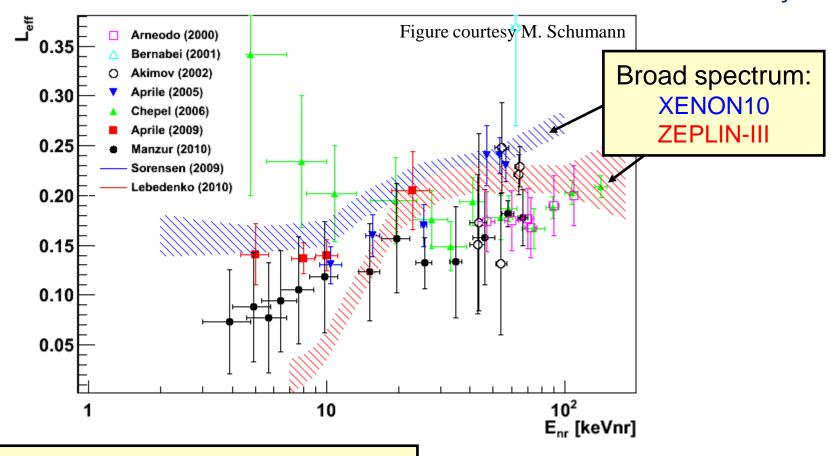




- \mathcal{L}_{eff} : nuclear recoil scintillation efficiency relative to fixed reference point
- L_y: light yield for 122 keV γ-rays (electronic recoils); easily measured

\mathcal{L}_{eff} measurements





Fixed energy (neutron beam):

Aprile et al. (2009) Manzur et al. (2010)

$\mathcal{L}_{\mathsf{eff}}$ issues

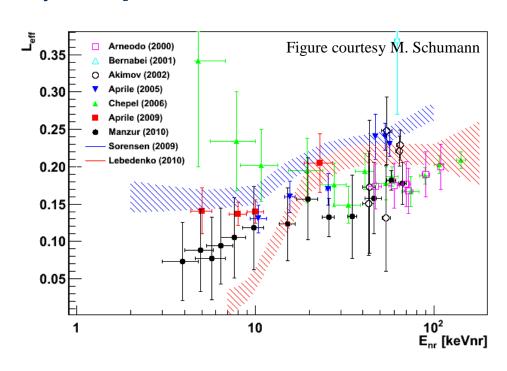


- Which data set(s) to use?
- Collar & McKinsey (2010)
- Lower L_{eff} gives weaker constraints Talk by D. McKinsey

- Conservative case (fixed energy): Manzur et al. (2010) [Note: conservative, not necessarily "best"]
- For discussion, see: A. Manalaysay, arXiv:1007.3746

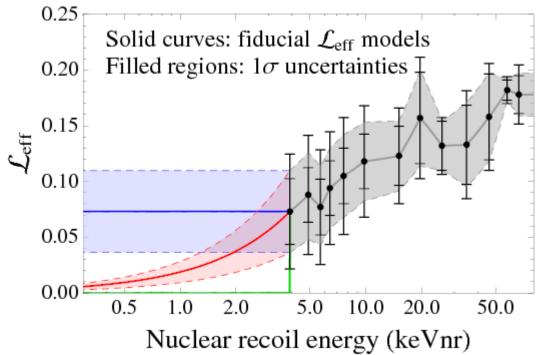


- Flat (constant)
- **Falling**
- Zero



$\mathcal{L}_{\mathsf{eff}}$ models





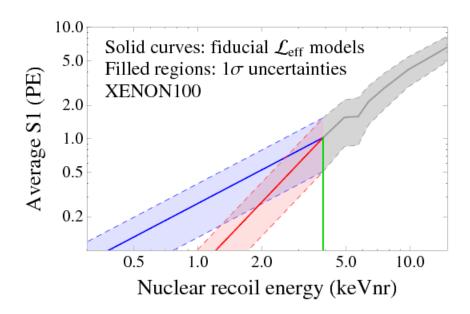
- High energy measurements: Manzur et al. (2010)
- Low energy \mathcal{L}_{eff} extrapolation:

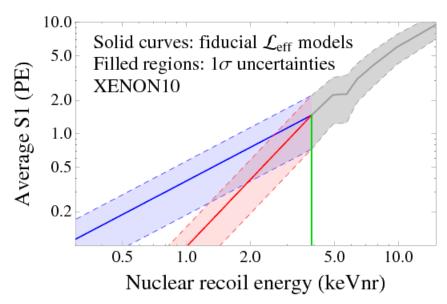
 - ConstantLinearly falling

Zero

\mathcal{L}_{eff} models







XENON100

S1 threshold: 4 PE

XENON10

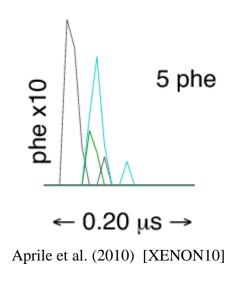
S1 threshold: 2 PE



$$\langle S1 \rangle = \frac{S_{nr}}{S_{ee}} \mathcal{L}_{eff}(E_{nr}) L_{y} E_{nr}$$

- Measured S1 signal: area of peak in electronic readout
 - Due to detector geometry and photocathode efficiency, only a small fraction (~ 10%) of prompt photons produce a photoelectron (PE) in a PMT ⇒ Poisson fluctuations
 - Due to digitization, PMT gain, etc., peak area of single PE varies: 1.0 ± 0.6 PE







Reconstructed nuclear recoil energy E'

$$E' = \frac{S_{ee}}{S_{nr}} \frac{1}{\mathcal{L}_{eff}(E') L_{y}} S1$$

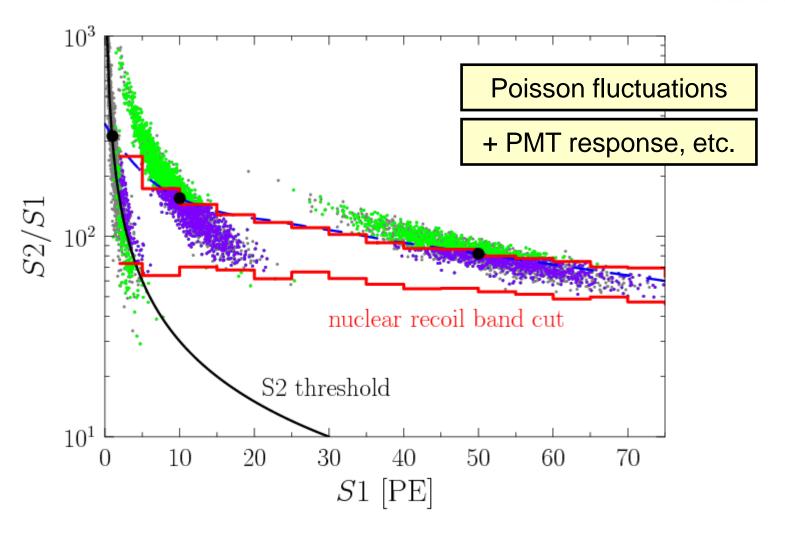
But S1 \neq \langle S1 \rangle , so E' \neq E_{nr}

- High energy recoils: good approximation
 S1 ≈ ⟨S1⟩ ⇒ E' ≈ E_{nr}
- Low energy recoils: poor approximation

Theory: $E_{nr} \Leftrightarrow \langle S1 \rangle$

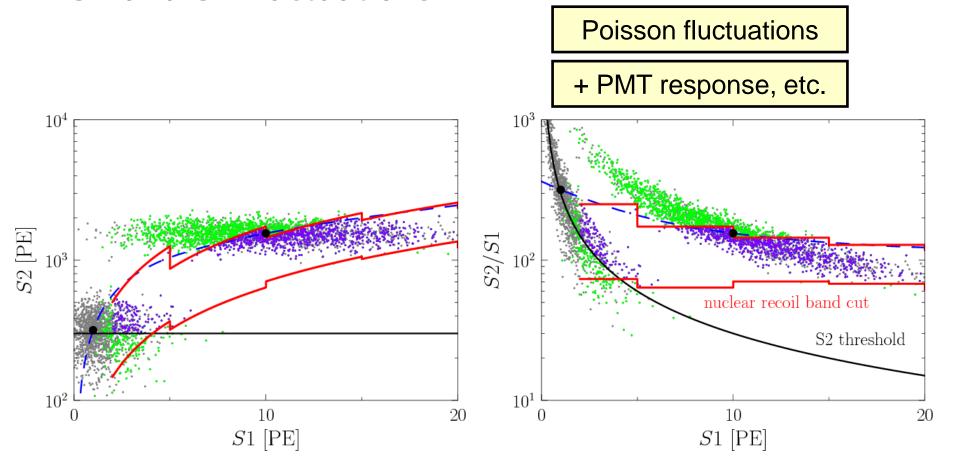
Observations: $E' \Leftrightarrow S1$





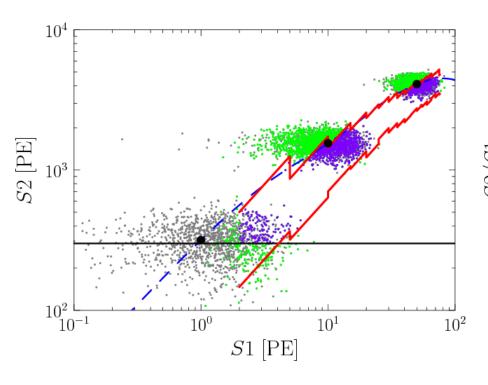


S1 and S2 fluctuations



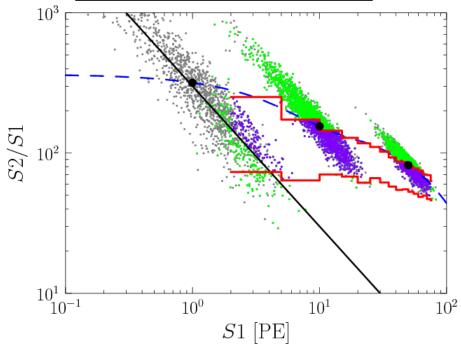


S1 and S2 fluctuations: relative size increases at low recoil energies



Poisson fluctuations

+ PMT response, etc.



Efficiencies: monte carlo

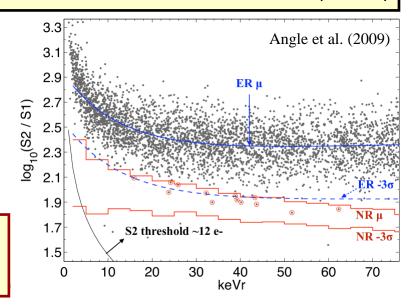
See also P. Sorensen, arXiv:1007.3549

Stockholm University

- For each (S1), generate S1 & S2 signals for a large number of random events
 - Poisson fluctuations in ionization electrons, photoelectrons
 - Fluctuations in PMT peak areas (digitization, gain, etc.)
- Efficiency: fraction that pass all cuts
 - S1 > 2 PE
 - S2 > 300 PE
 - S2/S1: nuclear recoil band cut
 - η_{S1} : peak finding efficiency factor
 - ε_p: PMT hit pattern
 - ε_f: S1 pulse shape

XENON100: no monte carlo apply $\langle S1 \rangle \geq 1$ PE cutoff (conservative)

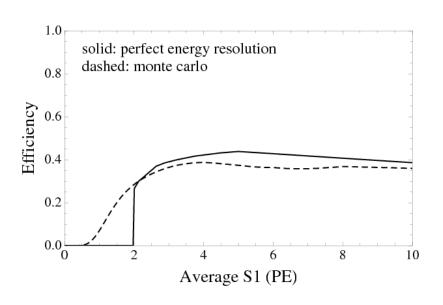
XENON10: PRD 80, 115005 (2009)

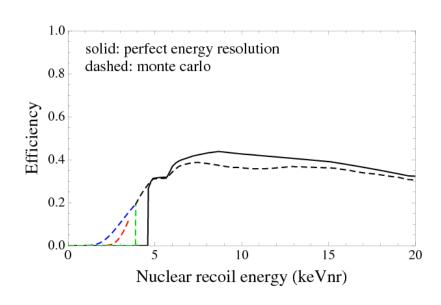


Efficiencies: monte carlo



Fractions of events passing all cuts in XENON10 ($2 \le S1 \le 75$ PE)



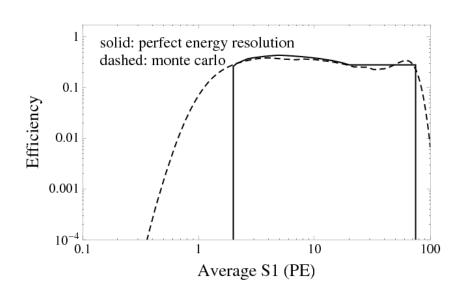


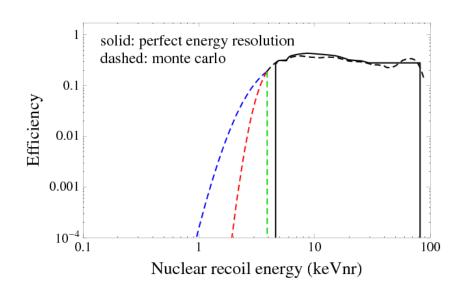
constant/falling/zero \mathcal{L}_{eff}

Efficiencies: monte carlo



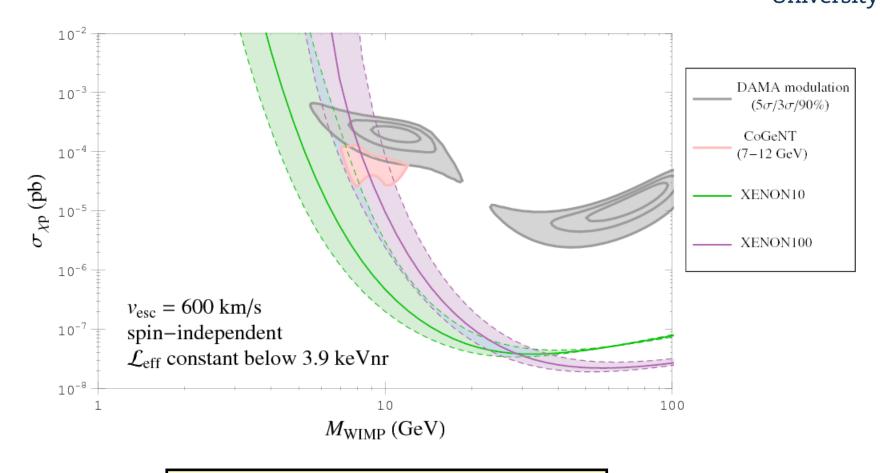
Fractions of events passing all cuts in XENON10 ($2 \le S1 \le 75$ PE)





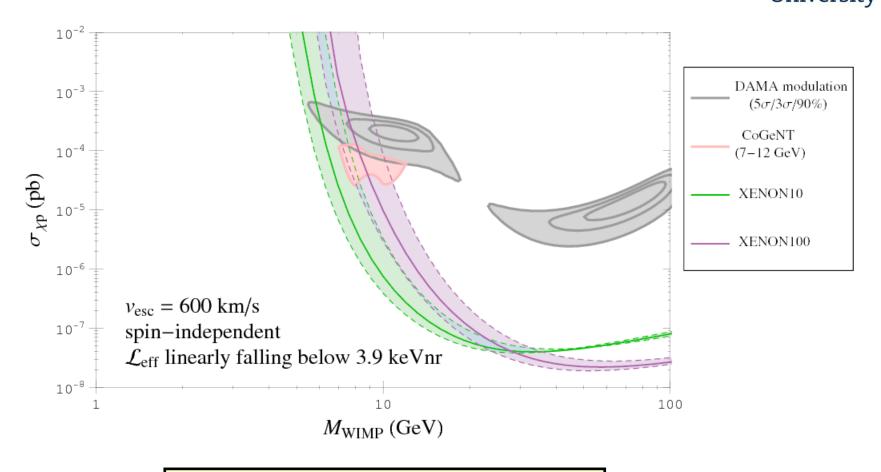
constant/falling/zero \mathcal{L}_{eff}





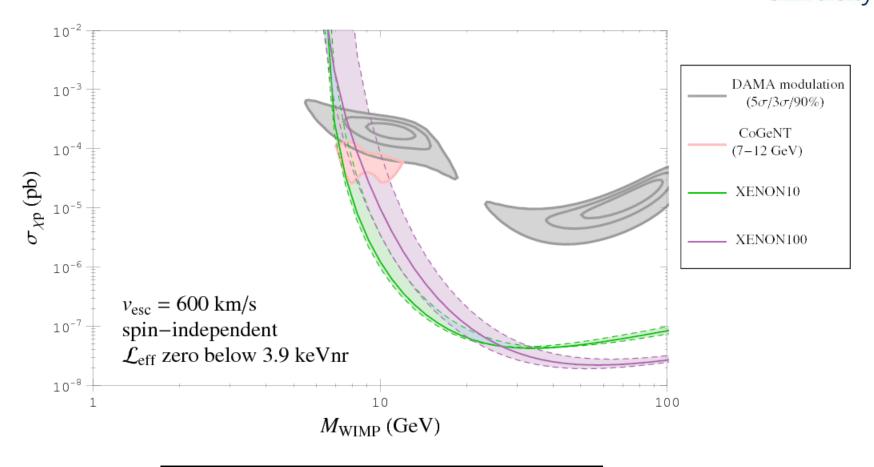
 \mathcal{L}_{eff} constant below 3.9 keVnr





 \mathcal{L}_{eff} linearly falling below 3.9 keVnr





 \mathcal{L}_{eff} zero below 3.9 keVnr



0.5

0.2

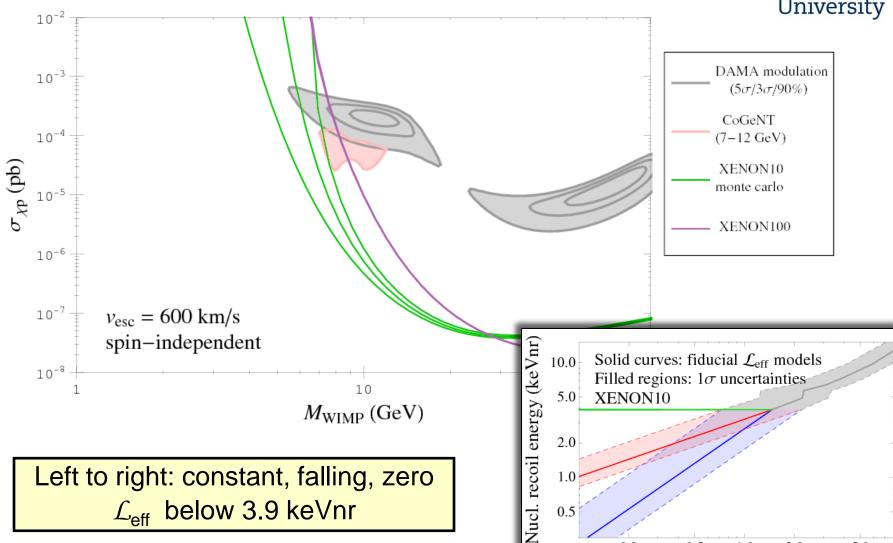
1.0

Average S1 (PE)

2.0

5.0





Identification of Dark Matter 2010

7/30/2010





DAMA modulation

 $(5\sigma/3\sigma/90\%)$

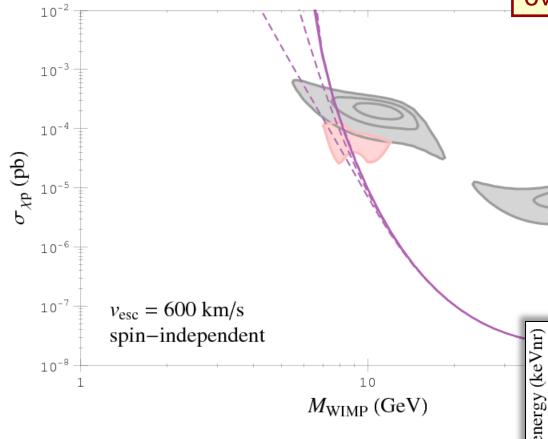
CoGeNT

(7-12 GeV)

XENON100 cutoff: $\langle S1 \rangle \ge 1.0$

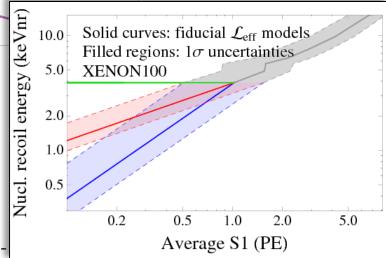
XENON100 no <S1> cutoff lm

Ignores NR band issues: overly optimistic constraints



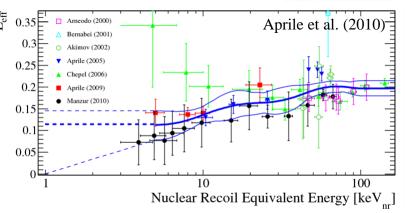
Left to right: constant, falling, zero \mathcal{L}_{eff} below 3.9 keVnr

Identification of Dark Matter 2010

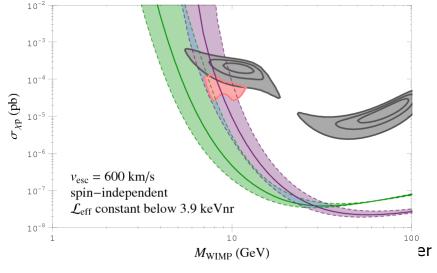


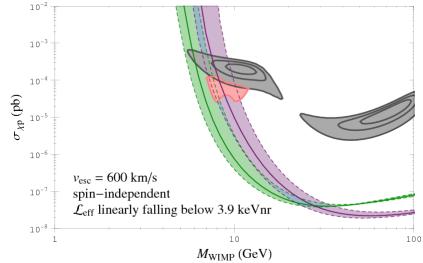
\mathcal{L}_{eff} measurements: fixed energy comparison

- Aprile et al. (2009) measurements
 - Left half of Manzur 1σ band is right half of Aprile 1σ band (approx.)
- XENON100 analysis:
 weighted average of fixed energy
 measurements Talk by M. Schumann



- Central curve: similar to Manzur 1σ upper bound, constant \mathcal{L}_{eff}
- Lower curve: similar to Manzur central values, linearly falling \mathcal{L}_{eff}





Summary



Dependence on $\mathcal{L}_{\mathsf{eff}}$:

- Low energy \mathcal{L}_{eff} extrapolation not significant for XENON100 when using Manzur data, choice of data sets more important
- Both L_{eff} data set and extrapolation important for XENON10

Comparing experiments:

- XENON10 highly incompatible with DAMA regions
- XENON10 strongly constrains CoGeNT region

Possibilities for compatibility:

Quenching factors

Hooper, Collar, Hall & McKinsey (2010)

- Halo models
- Spin-dependent
- · Inelastic scattering
- Etc.

XENON and low mass WIMPS: prospects





Sensitivity to low mass WIMPs possible

- Low thresholds necessary (more background?)
- lacktriangle Credibility of low mass limits at issue until $\mathcal{L}_{\mathrm{eff}}$ becomes better characterized at low recoil energies
- Conservative limits still possible, but want better
 - \Rightarrow New \mathcal{L}_{eff} measurements in progress

Better analysis for low masses: S2 only?

- No discrimination: more background
- Lower threshold: more signal
 DAMA & CoGeNT regions predict 10³ 10⁴+ low energy events in XENON detectors
- Strong constraints even with background

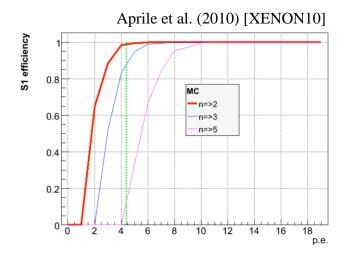
Talk by P. Sorensen



Backup Slides

Effect of η_{S1}

 Peak finding efficiency factor: probability of tagging at least two PMT hits



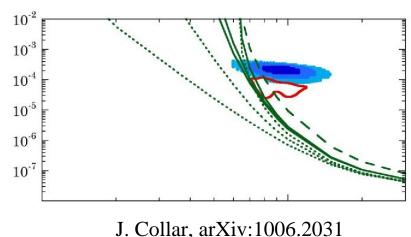
Example:

- Recoil energy that gives an average S1 of (S1) = 0.5 PE
- Efficiency of those events that produce S1 = 2 PE (Poisson fluctuations):

 $\eta_{S1}(0.5)$ [incorrect]

DAMA modulation 10- $(5\sigma/3\sigma/90\%)$ CoGeNT 10 (7-12 GeV) $\sigma_{\chi p} \, (pb)$ XENON10 10-XENON10 10-€ 10-7 spin-independent 10^{-8} 10^{2}

 $\eta_{S1}(2.0)$ [correct]



Does not account for NR band cut (larger effect)

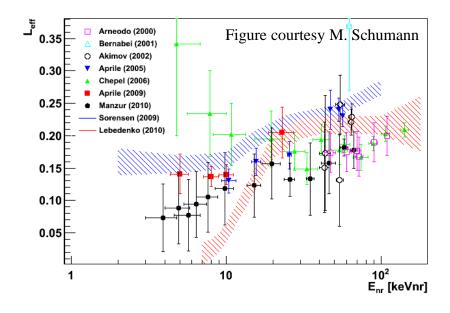
 M_{WIMP} (GeV)

\mathcal{L}_{eff} measurements: consistent with zero?

ZEPLIN-III broad spectrum fit: \mathcal{L}_{eff} falls to zero at $E_{nr} \sim 8 \text{ keVnr}$?

- Not 1σ or 90% CL band
- Does not include systematic uncertainties
- Manzur and Aprile data highly incompatible with $\mathcal{L}_{\text{eff}} \approx 0$

See Manalaysay, arXiv:1007.3746 for discussion of potential issues



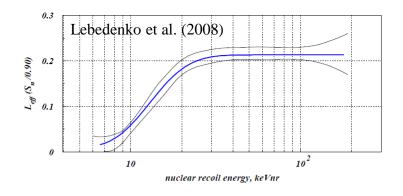


FIG. 15: The derived energy-dependent behaviour of $L_{eff} \cdot S_n$. The thick curve shows the best fit to the data, but other curves producing very similar goodness-of-fit indicators are obtained within the envelope shown. The constraints become very weak outside the energy range shown.

\mathcal{L}_{eff} measurements: consistent with zero?

Manzur data

- Naïve estimate: within ~ 2σ of zero?
- χ^2 of fit grows rapidly \mathcal{L}_{eff} as decreases
- Significant amount of scintillation events seen not easily attributable to any background

