

A new vision of the inert doublet model of dark matter

Laura Lopez Honorez

Universidad Autónoma de Madrid

based on *The inert doublet model of dark matter revisited.*

arXiv:1003.3125

in collaboration with C. Yaguna

IDM 2010 - Montpellier

Why revisiting the Inert doublet model ?

3-body processes can take over 2-body processes

3-body \equiv real + virtual massive particle
e.g. $WW^* \rightarrow W\bar{f}f'$

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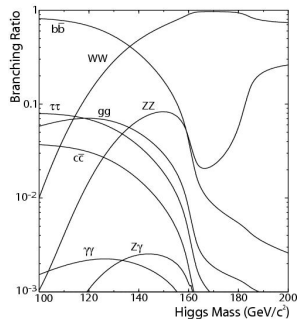
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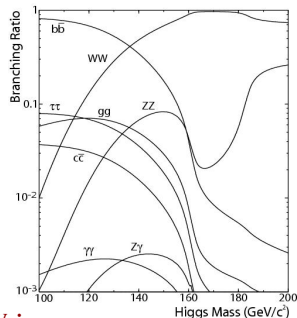
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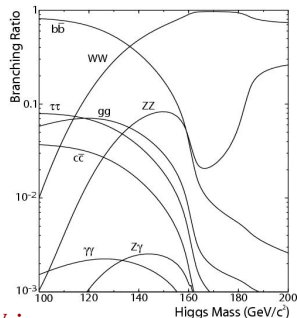
\rightsquigarrow Affect **relic abundance, viable parameter space, detection**

\rightsquigarrow **Significant effect** on : neutralino LSP [Chen & Kamionkowski JHEP '98, Yaguna PRD'10],

gravitino LSP [Choi & Yaguna '1003, & all '1007] (see Yaguna talk ! !),

Higgs DM [Hosotani, Ko & Tanaka PLB'09], singlet scalar DM [Yaguna PRD'10]

\rightsquigarrow **In this talk** : 3bdy annihilation in the **Inert doublet model**



Scalar WIMP as dark matter

Minimal DM spirit : SM + extra $SU(2)_L$ doublet see Cirelli *et al* '05-'09

- DM = **neutral** member of the extra doublet
Higgs : $H_1 \rightsquigarrow h$ and DM : $H_2 \rightsquigarrow H_0, A_0, H^\pm$
- stability \rightsquigarrow extra **Z_2 symmetry**
SM \rightarrow SM, $H_1 \rightarrow H_1$ and $H_2 \rightarrow -H_2$

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Possible Interactions

- (co)annihilation through **known** $SU(2)_L \times U(1)$ gauge bosons exchange
- quartic coupling λ_i to Higgs H_1

$$\lambda_3 |H_1|^2 |H_2|^2 + \lambda_4 |H_1^\dagger H_2|^2 + \frac{\lambda_5}{2} \left[(H_1^\dagger H_2)^2 + h.c. \right]$$

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$$\frac{1}{2} (\lambda_{H_0} H_0^2 + \lambda_{A_0} A_0^2 + 2\lambda_{H_c} H^+ H^-) (2v_0 h + h^2)$$

$$m_\chi^2 = \mu_2^2 + \lambda_\chi v_0^2$$

$\lambda_{H_c} \equiv \lambda_3/2$
 $\lambda_{H_0, A_0} \equiv (\lambda_3 + \lambda_4 \pm \lambda_5)/2$

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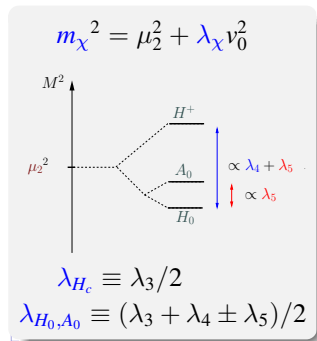
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Here we take H_0 as the LZ2P and $\lambda_L \equiv \lambda_{H_0}$



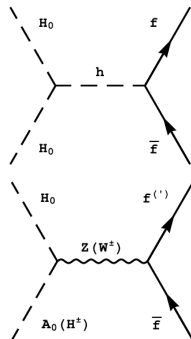
Inert doublet model parameter space

free parameters : $m_{H_0}, m_h, \lambda_L, \Delta m_{A^0}, \Delta m_{H^\pm}$

- $m_{H_0} \lesssim m_W$: GeV range ($\sim 10 - 80$ GeV) :

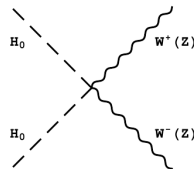
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Barbieri PRD06, LLH JCAP06, Gustafsson PRL07, Cao PRD07, Andreas JCAP08, ...



- $m_{H_0} \gg m_W$: TeV range ($\gtrsim 530$ GeV) :
 $H_0 H_0 \rightarrow ZZ, WW, hh$ and coannihil into bosons

Cirelli NPB06, Hambye JHEP09



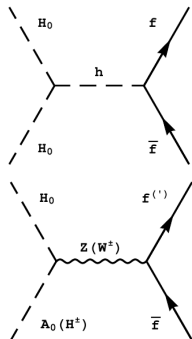
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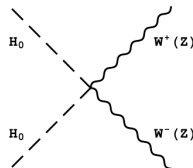
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LARGE MASS GAP DUE TO EFFICIENT
WW AND ZZ ANNIHILATION

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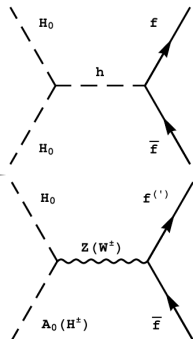
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\rightsquigarrow **Significantly affected by 3bdy annihilation :**

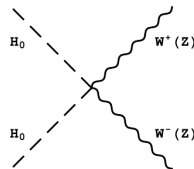
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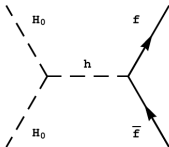
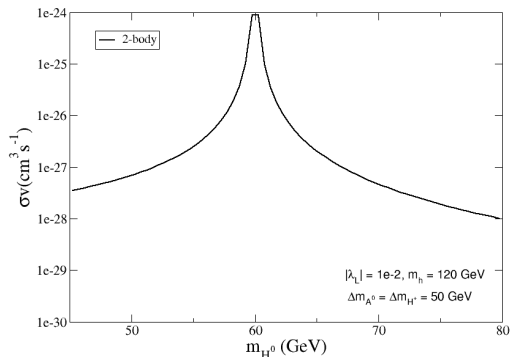
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Analysis for fixed parameters

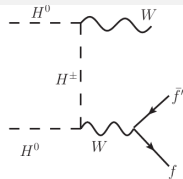
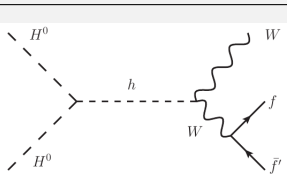
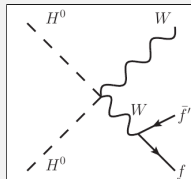
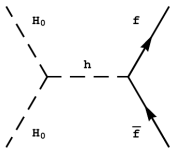
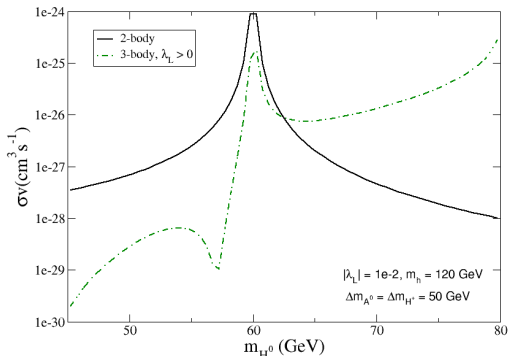
2-3 body annihilation cross section near m_W threshold

- σv_{2bdy} : higgs mediated,
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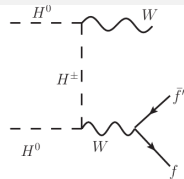
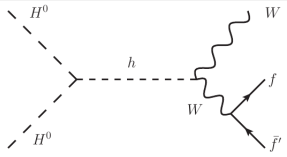
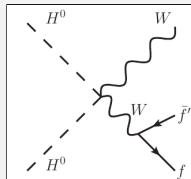
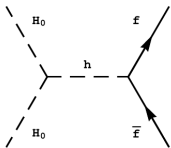
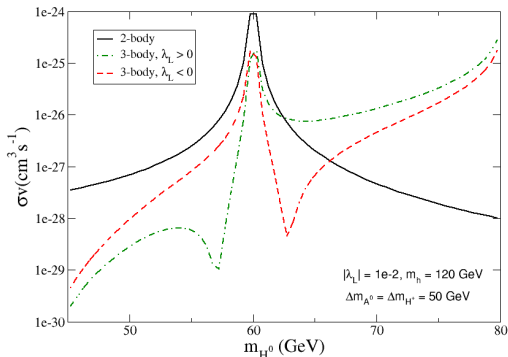
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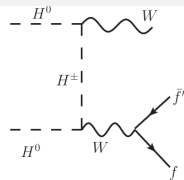
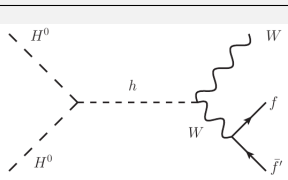
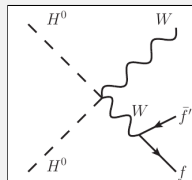
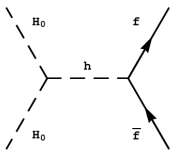
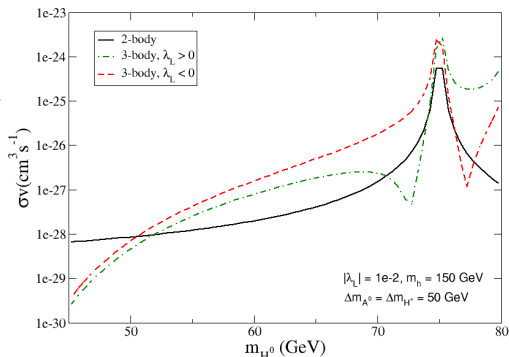
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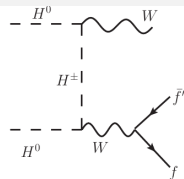
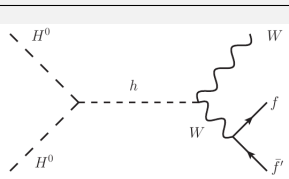
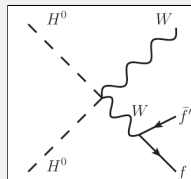
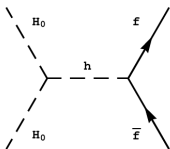
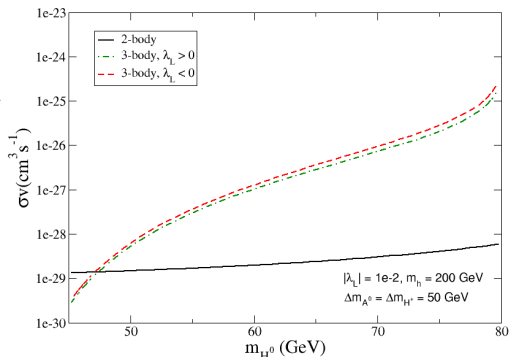
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$\rightsquigarrow \sigma V_{2bdy}$ VS σV_{3bdy} depends on
 m_{H^0} , m_h , λ_L sign and amplitude



Comparing 2-3 body relic density

Ω_{dm} is Δm dependent \leftrightarrow allows coannihilations

In the absence of coannihilations :

- roughly $\Omega_{dm} \propto 1/\langle\sigma v\rangle$ with
 $\langle\sigma v\rangle = \langle\sigma v(2\text{-body})\rangle + \langle\sigma v(WW^*)\rangle$
- We expect
 $\Omega_{dm}(3\text{-body}) \lesssim \Omega_{dm}(2\text{-body})$

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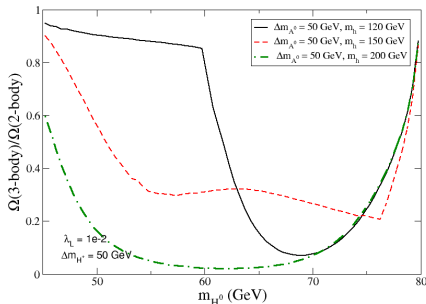
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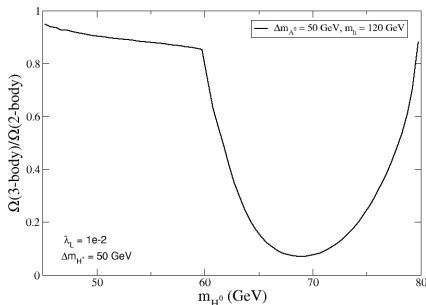
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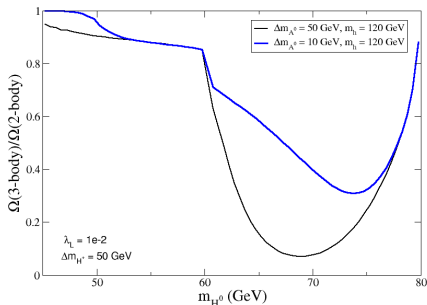
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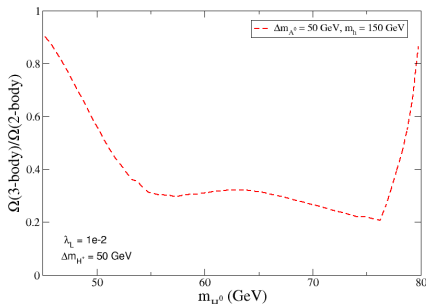
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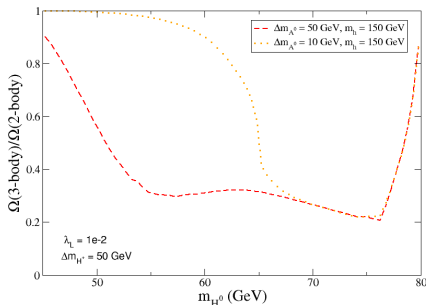
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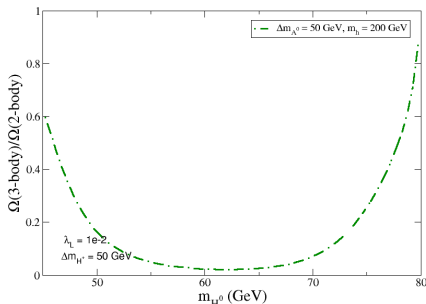
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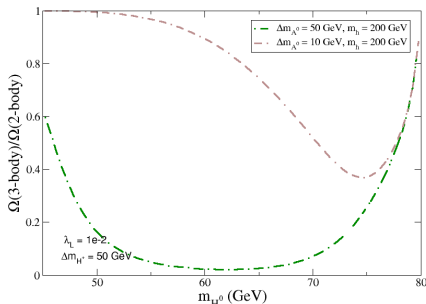
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\rightsquigarrow still 3-body final states significantly affect predictions for Ω_{dm}



Parameters for $\Omega_{H_0} = \Omega_{dm}^{WMAP}$

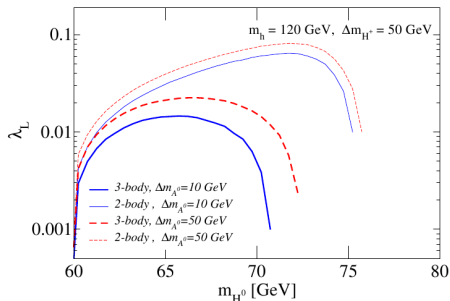
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Going from 2bdy only to 2+3bdy with or without coannihilations :

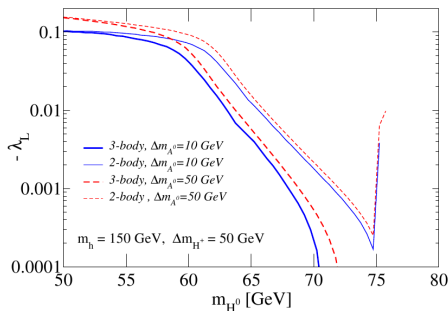


- correct λ_L is reduced up to $\sim \mathcal{O}(10)$.
- viable m_{H^0} range shrinks
 - 2bdy settled by the onset of W^+W^- annihilations
 - 2+3bdy depends on WW^* annihilations

Viable parameter space

Derive the $\lambda_L - m_{H^0}$ compatible with $\Omega_{dm}^{WMAP} h^2 = 0.11$

Going from 2bdy only to 2+3bdy with or without coannihilations :

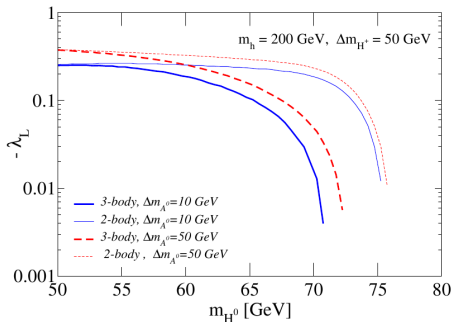


- correct λ_L is reduced up to $\sim \mathcal{O}(10)$.
- viable m_{H^0} range shrinks
 - 2bdy settled by the onset of $W^+ W^-$ annihilations
 - 2+3bdy depends on WW^* annihilations

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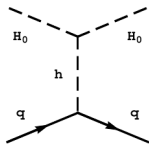
\rightsquigarrow rather generic feature of the Inert doublet model independently of m_h

\rightsquigarrow modify prospects for DM detection

Implications for Direct Detection

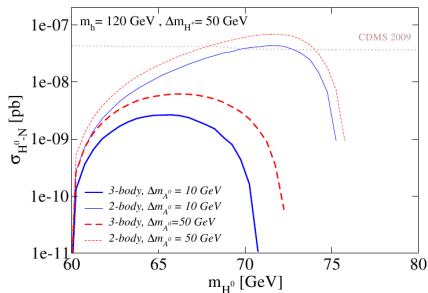
Direct detection through Elastic Scattering

Prospects along the viable parameter space :



$$\sigma_{H_0-N} \propto \left(\frac{\lambda_{H_0}}{M_{H_0} M_h^2} \right)^2$$

\rightsquigarrow predictions for σ_{H_0-N} reduced
up to $\sim \mathcal{O}(100)$

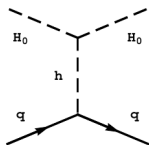


\rightsquigarrow better compatibility with present bounds

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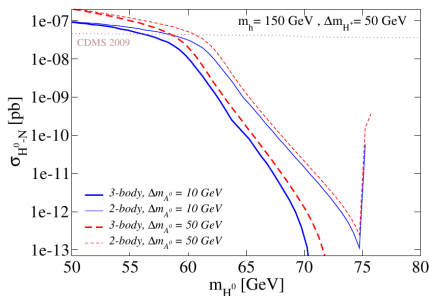
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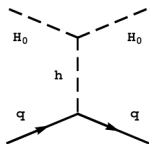


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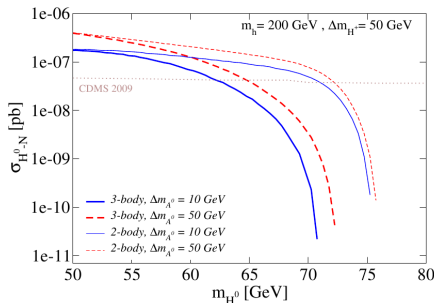
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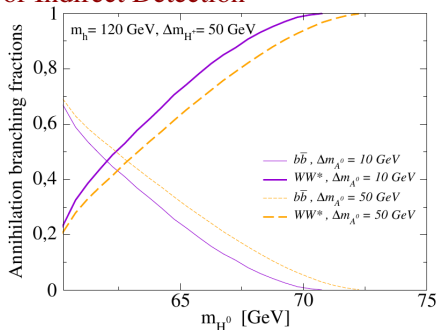
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More Implications

For Indirect Detection

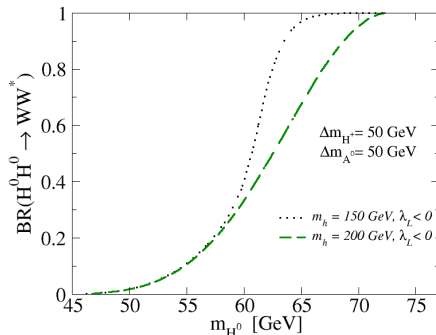


- new parameter space
- annihilations **no more $b\bar{b}$ dominated**

$\leadsto R(H_0 H_0 \rightarrow WW^*) \sim 1$
for m_{H^0} near W threshold

More Implications

For Indirect Detection

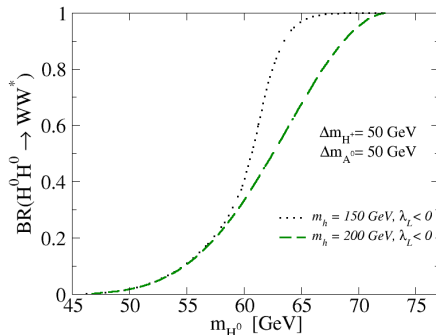


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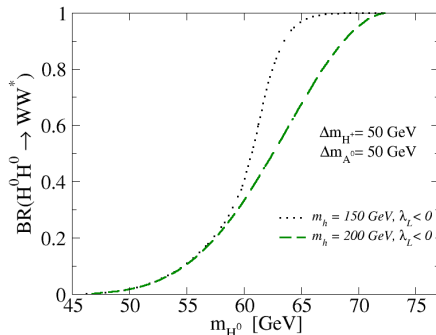


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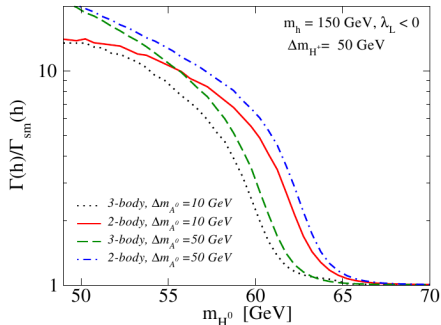
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$\rightsquigarrow \text{R}(H_0 H_0 \rightarrow WW^*) \sim 1$
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Detection at colliders



through **extra contributions to $\Gamma(h)$** due to
 $h \rightarrow A_0 A_0, H_0 H_0$ [Cao PRD07]

\rightsquigarrow The new parameter space
slightly change the prospects

Conclusion

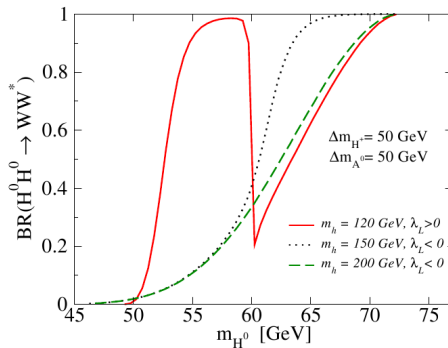
The **Inert Doublet** is a **WIMP** with a rich **Scalar** DM phenomenology

- We have shown that for $m_{H_0} \lesssim m_W$, annihilation into **3 bdy final states** (WW^*) **MUST be taken** into account
- The **genuine DM viable parameter space** next to W threshold was derived
 - relevant coupling to higgs λ_L is **reduced** up to $\mathcal{O}(10)$
 - the viable range in m_{H^0} **shrinks** by several GeVs
- **Prospects for DM detection are modified**
 - Direct detection : σ_{H_0-N} **decreases** by a factor $\mathcal{O}(100)$
 - DM annihilation is no more $\bar{b}b$ dominated,
 $\text{BR}(WW^*) \sim 1$ for $m_{H_0} \lesssim 60 \text{ GeV}$

This is the End
Thank you for your attention !!

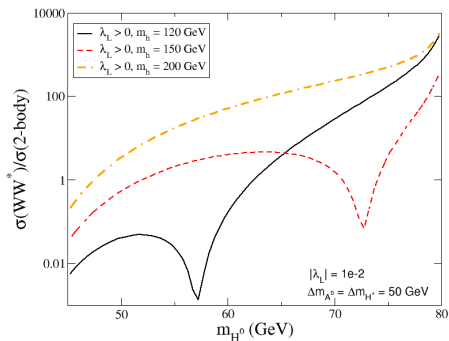
Backup

branchings



Comparing 2-3 body annihilation cross section

3bdy annihilation dominates over 2 bdy on

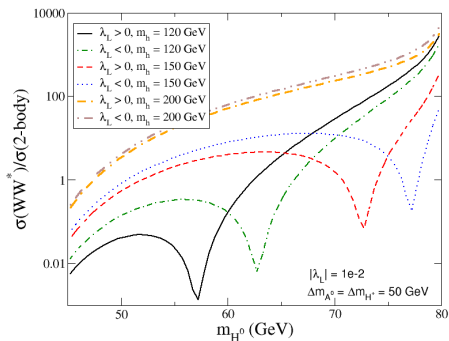


• $\lambda_L = 10^{-2}$

... a significant range of the parameter space, depend on m_h

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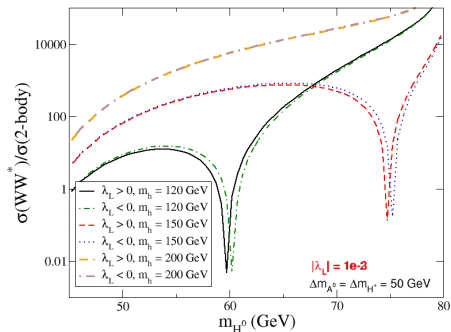


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Comparing 2-3 body annihilation cross section

3bdy annihilation **dominates** over 2 bdy on



- $\lambda_L = 10^{-2}$
... a **significant** range of the parameter space, depend on m_h
- $\lambda_L = 10^{-3}$
... the entire mass range independently of m_h
but not representative for $H_0 \equiv \text{DM}$

IDM : Mass Ranges

Mass Ranges	main contributions to σ_{eff}	mass splittings	main Refs
$m_{H_0} \ll m_W (\mathcal{O}(GeV))$	$H_0 H_0 \rightarrow h^* \rightarrow \bar{f} f$	$\Delta m_{ij} \gtrsim m_Z - m_{H_0} \sim 90 \text{ GeV}$	Andreas <i>et al</i> '08
$m_{H_0} \lesssim m_W$	$H_0 H_0 \rightarrow h^* \rightarrow \bar{f} f$ $H_0 A_0(H^+) \rightarrow Z^*(W^*) \rightarrow \bar{f} f^{(')}$	$\Delta m_{ij} \gtrsim m_Z - m_{H_0} \gtrsim 7 \text{ GeV}$	Barbieri <i>et al</i> '06 LLH <i>et al</i> '06
$m_{H_0} \gg m_W (\mathcal{O}(TeV))$	$H_0 H_0 \rightarrow ZZ, WW, hh$ coannihil into bosons	$\Delta m_{ij} \lesssim 17.6 \text{ GeV}$	Hambye <i>et al</i> '09

IDM : Potential - constraints

- Full Potential

$$V(H_1, H_2) = \mu_1^2 |H_1|^2 + \mu_2^2 |H_2|^2 + \lambda_1 |H_1|^4 + \lambda_2 |H_2|^4 \\ + \lambda_3 |H_1|^2 |H_2|^2 + \lambda_4 |H_1^\dagger H_2|^2 + \frac{\lambda_5}{2} \left[(H_1^\dagger H_2)^2 + h.c. \right]$$

- Dark scalars couplings to Higgs and masses :

$$\frac{1}{2} (\lambda_{H_0} H_0^2 + \lambda_{A_0} A_0^2 + 2\lambda_{H_c} H^+ H^-) (2v_0 h + h^2) \\ m_h^2 = 2\lambda_1 v_0^2, \quad m_i^2 = \mu_2^2 + \lambda_i v_0^2.$$

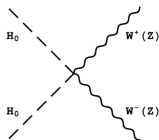
- Stability constraint

$$\lambda_{1,2} > 0, \\ \lambda_{H_0}, \quad \lambda_{A_0}, \quad \lambda_{H_c} > -\sqrt{\lambda_1 \lambda_2}.$$

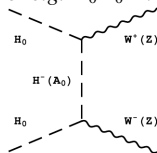
- EWPT measurements : $\Delta T \approx \frac{1}{12\pi^2 \alpha v^2} (m_{H^+} - m_{A_0})(m_{H^+} - m_{H_0})$

IDM in the High mass regime : Relic abundance

Quartic couplings ON : extra Higgs processes and $m_{H_0} \neq m_{A_0} \neq m_{H_c}$

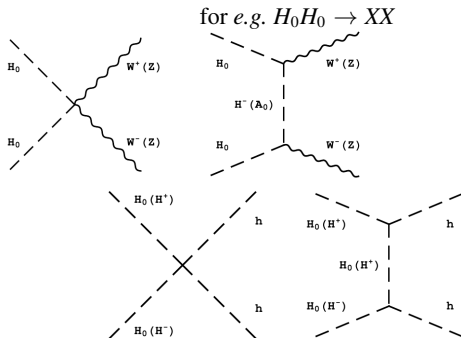


for *e.g.* $H_0H_0 \rightarrow XX$



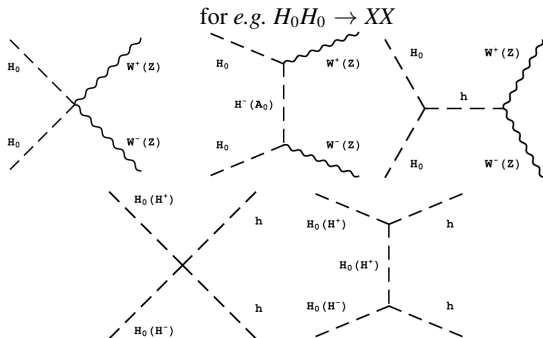
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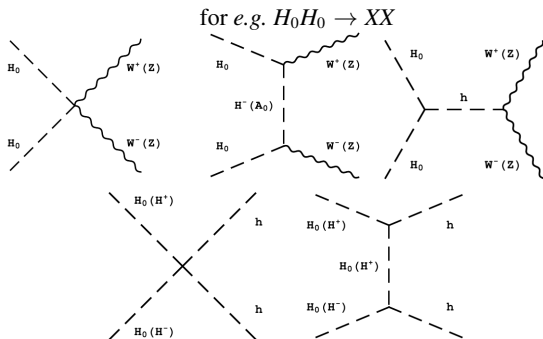


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$$\sigma_{\text{eff}} = \sum_{ij} \left(\sigma_g^{ij} + \sigma_\lambda^{ij} \right) \propto \frac{1}{m_{H_0}^2}$$

where $\sigma_\lambda^{ij} = \frac{\Lambda^{ij}}{m_{H_0}^2}$ with $\Lambda^{ij} \propto \lambda * \lambda$ and $\Lambda^{ij} > 0$



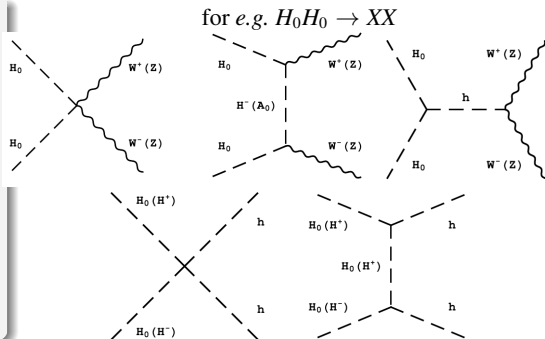
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- $m^* \sim 534 \text{ GeV}$ is minimal to satisfy WMAP



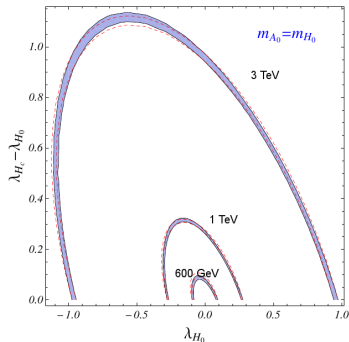
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- for fixed m_{H_0} and σ_{eff} λ_i lie on ellipsoid surface



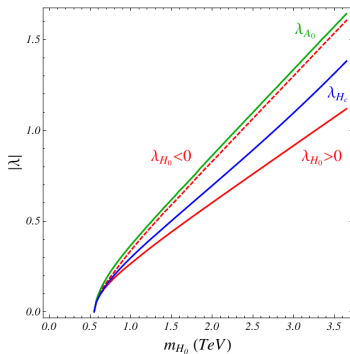
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 $\rightsquigarrow \lambda_i$ are **bounded**,
 $\lambda_i^{\text{max}} \propto m_{H_0}$ at high mass.



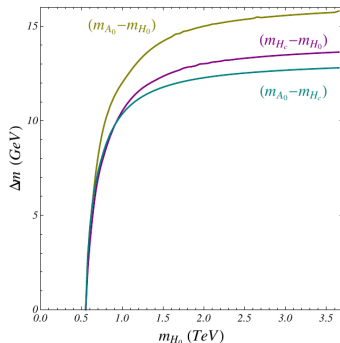
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 λ_i lie on ellipsoid surface
 $\rightsquigarrow \lambda_i$ are bounded,
 $\lambda_i^{\text{max}} \propto m_{H_0}$ at high mass.
- mass splittings are also bounded as
 $m_i - m_j \propto (\lambda_i - \lambda_j)/m_{H_0}$



3bdy effect on DM ?

3-body processes can enhance DM annihilation :

- supersymmetric dark matter :

- neutralino LSP : Chen & Kamionkowski JHEP '98 study $\sigma v_{v \rightarrow 0}$ and impact on ν detection from annihilation in the Earth bellow WW and $\bar{t}t$ mass threshold

Yaguna PRD'10 demonstrate up to 10% effect on Ωh_χ^2 for bino-like χ including $\bar{t}t^*$ (usually 2-bdy $\bar{b}b$ dom)

- gravitino LSP : Choi & Yaguna '1003 W^*l and $Z^*\nu$ give significant (up to 90%) to \tilde{G} decay (usually 2-bdy $\gamma\nu$ dom)

Choi, Restrepo, Yaguna & Zapata '1007 gamma+antimatter signal [see Yaguna talk ! !]

- scalar DM

- Higgs DM : Hosotani, Ko & Tanaka PLB'09 (gauge-Higgs unification)
 $\Omega_{DM} \rightsquigarrow m_{DM} = 75 \text{ GeV}$ (2bdy only) $\Rightarrow m_{DM} = 70 \text{ GeV}$ (including 3bdy)
- singlet scalar DM : Yaguna PRD'10, $SS \rightarrow h \rightarrow WW^*$ enhance $\sigma v_{v \rightarrow 0}$ and reduce Ω_{DM} independently of S-higgs coupling

This is the End