

New Signatures of WIMPlless Dark Matter

Jason Kumar
University of Hawai'i

Collaborators

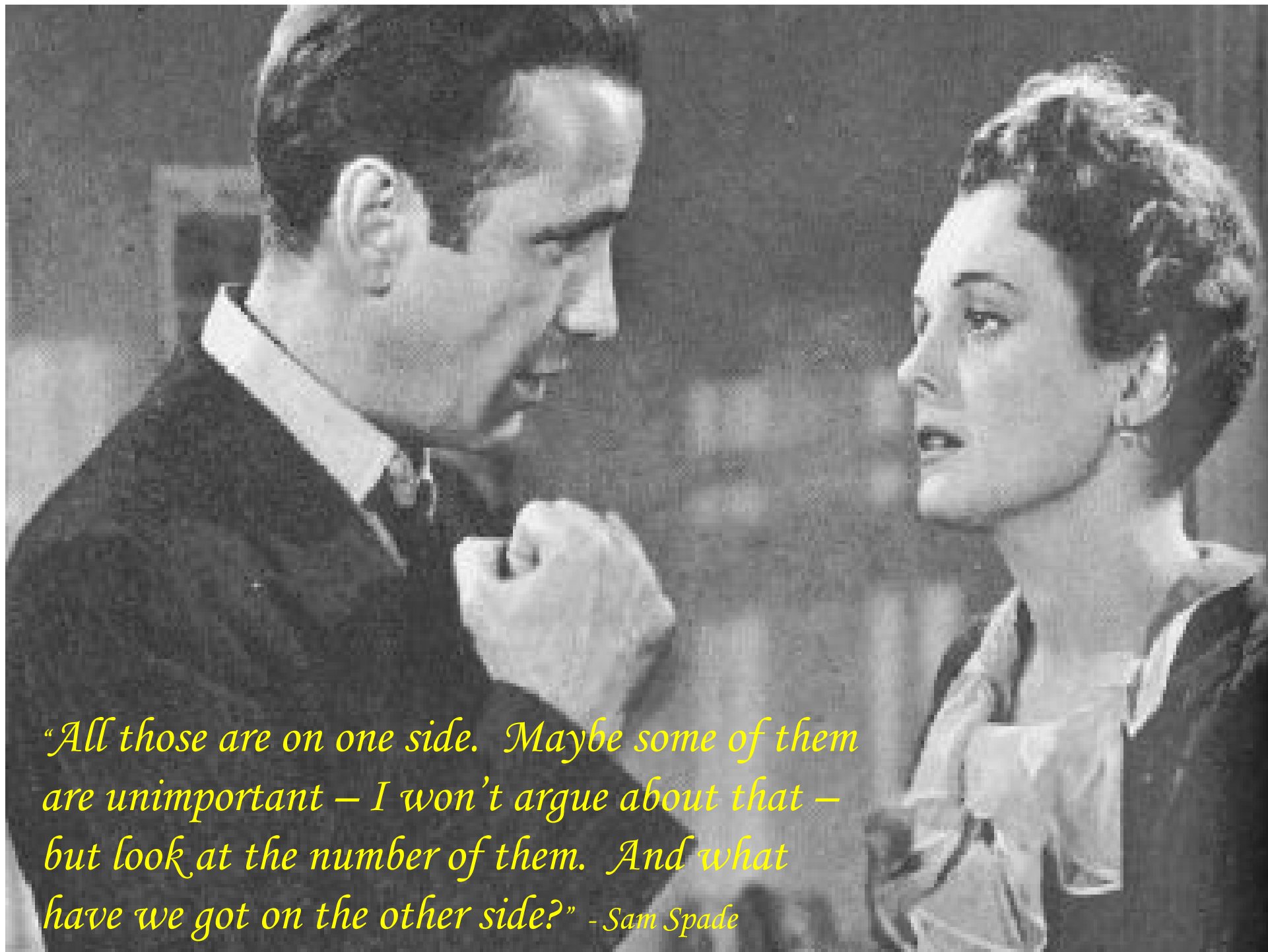
- Johan Alwall
- Vernon Barger
- Jonathan Feng
- John Learned
- Danny Marfatia
- Enrico Sessolo
- Stefanie Smith
- Louis Strigari
- Shufang Su

– 0803.4196, 0806.3746, 0808.4151, 0908.1768, 1002.3366,
1004.4573



Hints of low-mass dark matter

- **DAMA/LIBRA** (see Belli's talk)
 - 8.9σ annual modulation of event rate
 - potentially explainable by low mass, large σ_{SI} DM (no channeling?)
- **CDMS** (see Saab's talk)
 - two candidate events (~ 1 bg. expected $\rightarrow \sim 77\%$ CL)
 - both have relatively low recoil energy, but issues with timing cuts....
- **CoGeNT**
 - event rate unexplained by **known** backgrounds, fits low mass DM
 - 90% CL region excludes null hypothesis (no DM contribution)
- **CRESST ...?** (see Seidel's talk)
 - excess events in oxygen band which may fit low mass DM...?
- **fit to DAMA/CoGeNT/CRESST** (Collar, Hall, Hooper, McKinsey)
 - fit point $\rightarrow m_\chi \sim 7\text{GeV}$, $\sigma_{\text{SI}} \sim 2 \times 10^{-4} \text{pb}$ (v. dist., form factors, **quenching**)



"All those are on one side. Maybe some of them are unimportant – I won't argue about that – but look at the number of them. And what have we got on the other side?" - Sam Spade

the key to low mass is low recoil energy...

- XENON100
 - claims to rule out CoGeNT/DAMA
 - major questions about L_{eff} .
- XENON10
 - reanalysis by SGGF suggests they can put tighter bounds than XENON100
 - would rule out most of DAMA/CoGeNT region
 - questions about assumptions of efficiency at low recoil energy (Collar)
 - but the plot thickens... analysis using S2 signal could rule out the entire CoGeNT/DAMA region (wish I had seen P. Sorensen's talk)
- CDMS
 - analyzing data using lower threshold
 - issues with energy scale calibration of silicon detectors?



What's a simple theorist to do?

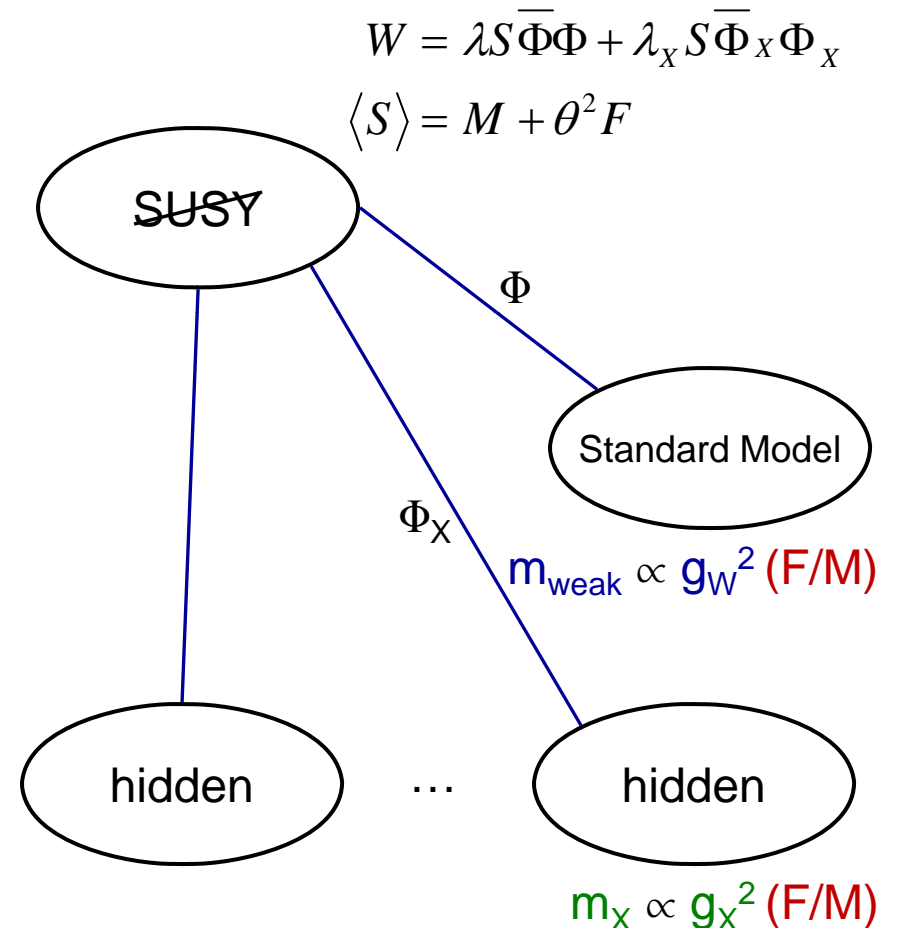
"My interest ... is purely a sporting one." - Rick Blaine

The WIMPlless miracle

- non-relativistic thermal dark matter
 - Boltzmann equation $\rightarrow \rho \propto \langle \sigma_A v \rangle^{-1}$
- to get observed DM density need $\langle \sigma_A v \rangle \sim 1 \text{ pb}$
- stable matter with coupling and mass of the electroweak theory would have about right relic density for dark matter
 - WIMP miracle
- $\sigma_A v \propto g^4 / m^2$
- the real miracle is in what controls this ratio
- in GMSB, $m_{\text{soft}} \propto g^2 (F/M_{\text{mess}})$
- if particle at m_{soft} is stable, $\rho \propto (F/M)^2$, regardless of what m_{soft} is
- WIMPlless miracle

WIMPlless setup

- the standard “**low-energy SUSY**” setup (GMSB)
 - **one sector breaks supersymmetry**
 - MSSM receives SUSY-breaking, as well as other hidden sectors
 - **soft scales controlled by one spurion field**
- we add to this **extra gauge sectors**, which behave in a qualitatively similar way
 - assume symmetry **stabilizes particle at SUSY-breaking scale**, which **could be anything**
 - but $\rho \propto (F/M)^2$ is **always approximately right**

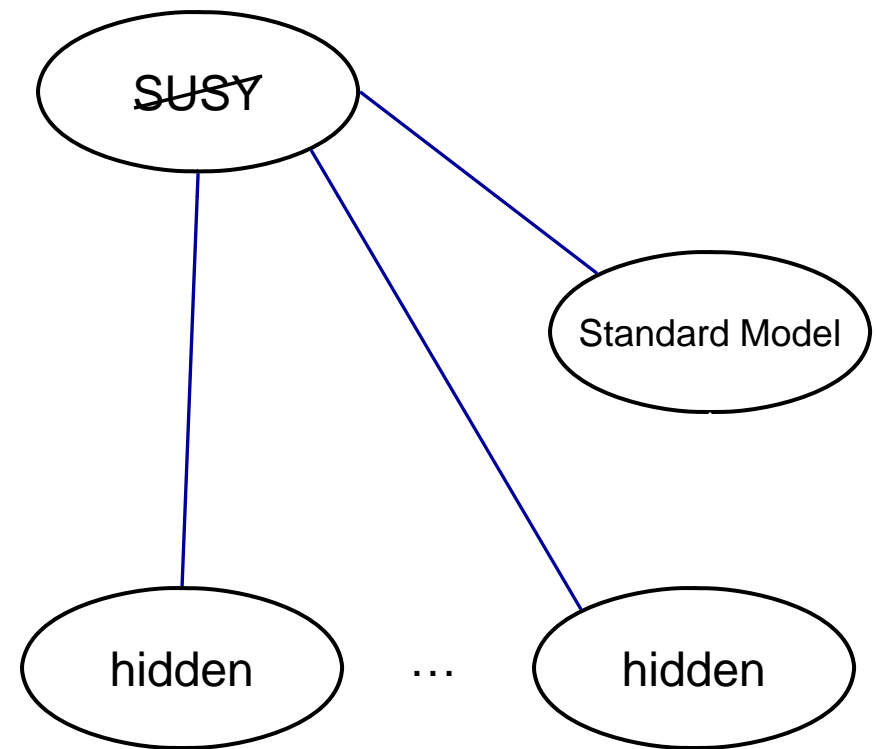


WIMPlless Miracle

- a new, well-motivated scenario for dark matter (scalar or fermion)
- natural dark matter candidates with approximately correct mass density
- unlike “WIMP miracle” scenario, here dark matter candidate can have a range of masses and couplings
- opens up the window for observational tests, beyond standard WIMP range
- implications for collider, direct and indirect detection strategies (with a focus on low mass, but more general)

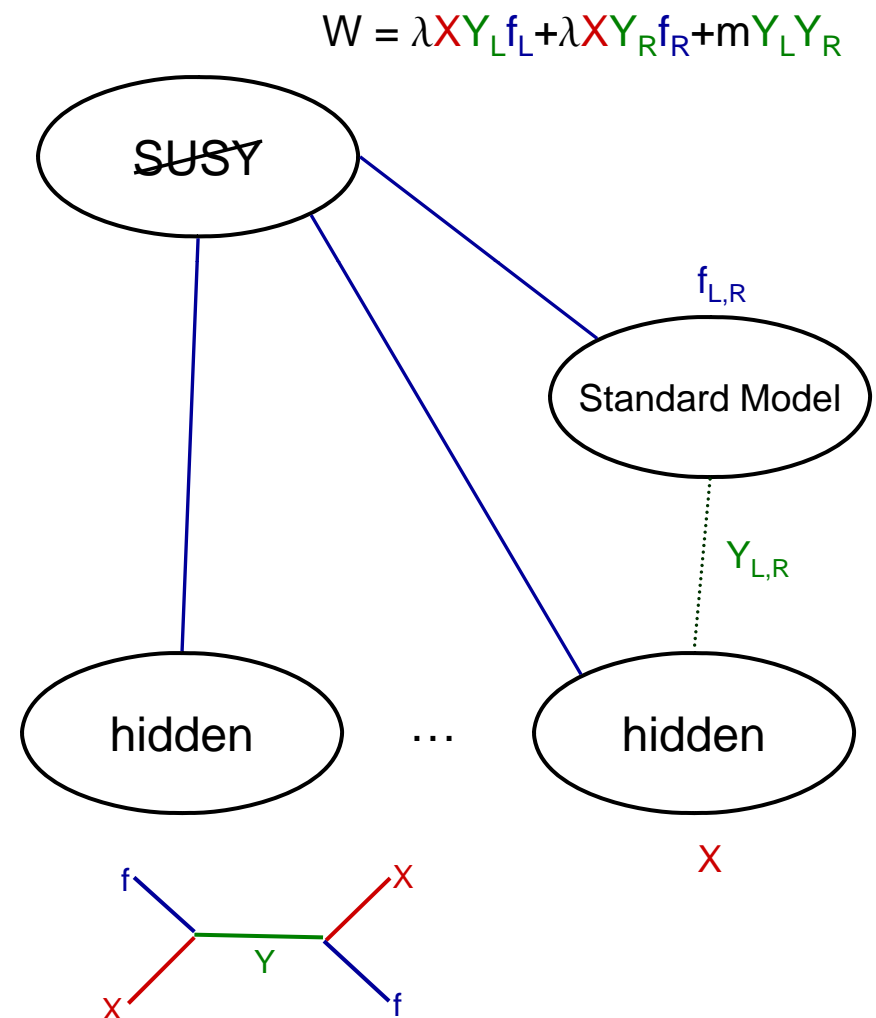
Yukawa coupling to SM

- if no connection between SM and hidden sector...
 - only gravitational effects



Yukawa coupling to SM

- if no connection between SM and hidden sector...
 - only gravitational effects
- but could have connectors between those sectors
 - exotics (Y) charged under both SM and hidden sector
 - exotic 4th generation multiplet
- Yukawa couplings between dark matter, SM matter and exotic connectors
 - get nuclear scattering through light or heavy (loop) quarks
 - annihilation to SM matter



New WIMPlless features for low mass....

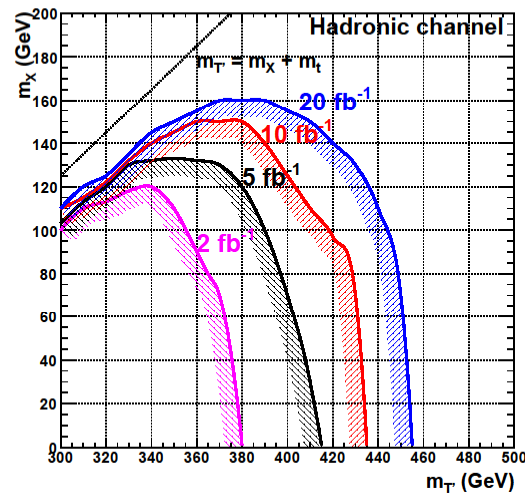
- scalar WIMPlless DM
 - can have larger σ_{SI} than expected for neutralinos
 - for σ_{SI} , need to couple to $f_L^\dagger f_R$
 - need light quark mass or squark mixing insertion
 - chirality suppression
 - with scalar DM, chirality flip from m_Y
- can fit near CHHM region ($\lambda_b \sim 0.8$, $m_\chi \sim 6-7$ GeV, $m_Y \sim 400$ GeV)
 - assuming hierarchical Yukawa coupling to 3rd generation only (simple FCNC solution)
 - can be tested with near term data
 - collider QCD production of 4th generation quark connectors (Y)
 - $Y \rightarrow X + \text{jets} \rightarrow$ always missing E_T (hidden sector charge)
 - distinctive signature can be seen with Tevatron and near-term LHC data
 - DM annihilation to SM $\rightarrow \nu$ can be see at Super-K

Upshot

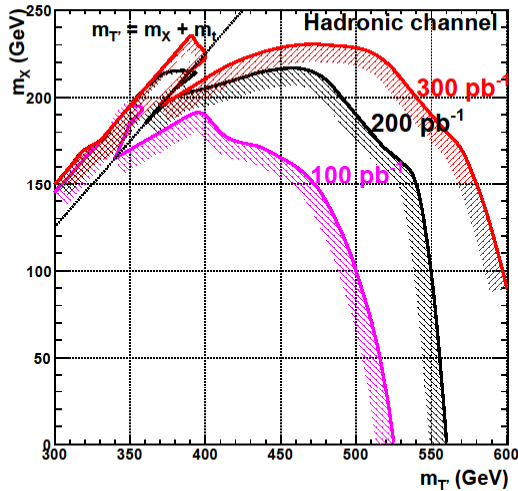
$Y=T'$
 $m_{T'} \sim 300-600 \text{ GeV}$

95%

Exclusion for $T' \bar{T'} \rightarrow t X \bar{t} X$ at the Tevatron

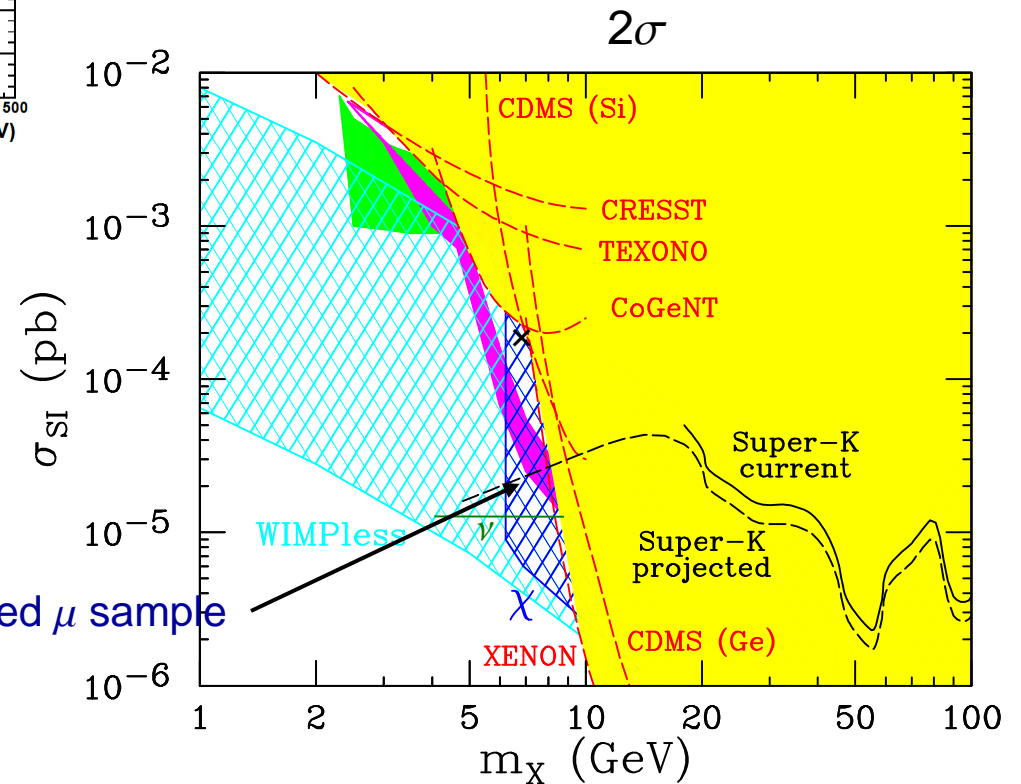


Exclusion for $T' \bar{T'} \rightarrow t X \bar{t} X$ at 10 TeV LHC



fully-contained μ sample

projected Super-K bounds using
 fully-contained events and 3000
 live days, plus WIMPless ($0.3 < \lambda_b < 1.0$) and
 neutralino (Bottino, et al) predictions



New WIMPlless features for Majorana fermion DM....

- Majorana fermion WIMPlless DM
 - for Majorana fermion DM, $\sigma_{\text{SI}}=0$, but σ_{SD} is non-zero
 - so not targeting low mass DAMA/CoGeNT/CRESST now....
 - most models will be seen first through σ_{SI} , σ_{SD} can confirm
 - Majorana fermion WIMPlless DM is only found through σ_{SD}
- IceCube/DeepCore will soon have among the best bounds on σ_{SD}
 - like SK, uses DM annihilation to SM in the sun, followed by ν shower
 - here, consider coupling only to 1st gen. quarks (nuclear spin)
 - best annihilation channels are for superpartners, or τ
 - upshot \rightarrow 3σ evidence possible in 5 yr. ($\lambda_{u,d} \sim 0.5$)

Conclusion

- new WIMPless theoretical scenario for dark matter
 - large range of masses and couplings
- possible explanation for results of DAMA/LIBRA, CoGeNT
- interesting searches at Tevatron and LHC
- signals possible at Super-Kamiokande and IceCube/DeepCore

Mahalo!

Back-up slides

Collider cuts

- Tevatron (hadronic)
 - precuts
 - no isolated leptons
 - jets ≥ 5 ($p_T > 20$ GeV)
 - missing $E_T > 100$ GeV
 - isolation (jet from missing p_T)
 - $\Delta\phi > 90^\circ$ for leading jet
 - $\Delta\phi > 50^\circ$ for second jet
 - additional cuts
 - missing E_T
 - 150, 200, 250 GeV
 - $H_T = \Sigma |p_T|$
 - 300, 350, 400 GeV
 - jets ≥ 6 ($p_T > 20$ GeV)
- LHC (hadronic)
 - precuts
 - no isolated leptons
 - jets ≥ 5 ($p_T > 40$ GeV)
 - missing $E_T > 100$ GeV
 - isolation
 - $\Delta\phi > 11.5^\circ$ for first 3 jets
 - additional cuts
 - missing E_T
 - 150, 200, 250, 300 GeV
 - H_T
 - 400, 500 GeV
 - jets ≥ 6 ($p_T > 40$ GeV)

IceCube/DeepCore

- superpartner channel
 - spectrum from Dimopoulos, Thomas, Wells
 - $m_{\text{stau}} = 137 \text{ GeV}$
 - $m_{\text{sneutrino}} = 111.5 \text{ GeV}$
 - $m_{\chi} = 94.5 \text{ GeV}$
- assume 1° angular acceptance
- IC E_μ -threshold = 100 GeV
- DC E_μ -threshold = 35 GeV
- account for matter effects in sun and vacuum oscillation
 - including τ -regeneration

