

# The $B \rightarrow X_s \gamma$ BF at NNLO

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*XLII Rencontre de Moriond*



# Why a NNLO calculation?

- Exp:

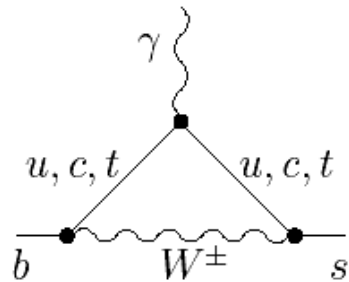
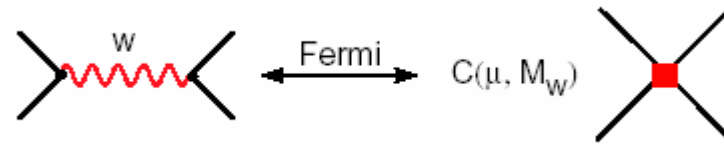
$$\text{BR}(B \rightarrow X_s \gamma) = (3.55 \pm 0.24^{+0.09}_{-0.10} \pm 0.03) 10^{-4}$$

latest HFAG  $E_\gamma > 1.6$  GeV

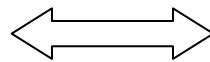
- Until recently: SM prediction at NLO in QCD + leading EW and non-perturbative effects with  $\sim 10\%$  theory error
- Need to match  $\sim 5\%$  exp error at end of B factories: first results of a collective effort for the NNLO calculation in [hep-ph/0609232](https://arxiv.org/abs/hep-ph/0609232)
- **authors:** M. Misiak, H.M. Asatrian, K. Bieri, M. Czakon, A. Czarnecki, T. Ewerth, A. Ferroglia, P. Gambino, M. Gorbahn, C. Greub, U. Haisch, A. Hovhannisyan, T. Hurth, A. Mitov, V. Poghosyan, M. Slusarczyk and M. Steinhauser.

# b $\rightarrow$ s transitions

$$\Lambda_{\text{QCD}} \ll m_b \ll M_W$$



at  $O(1/M_W^2)$



$$O_7 = m_b \bar{b}_R \sigma_{\mu\nu} F^{\mu\nu} S_L$$

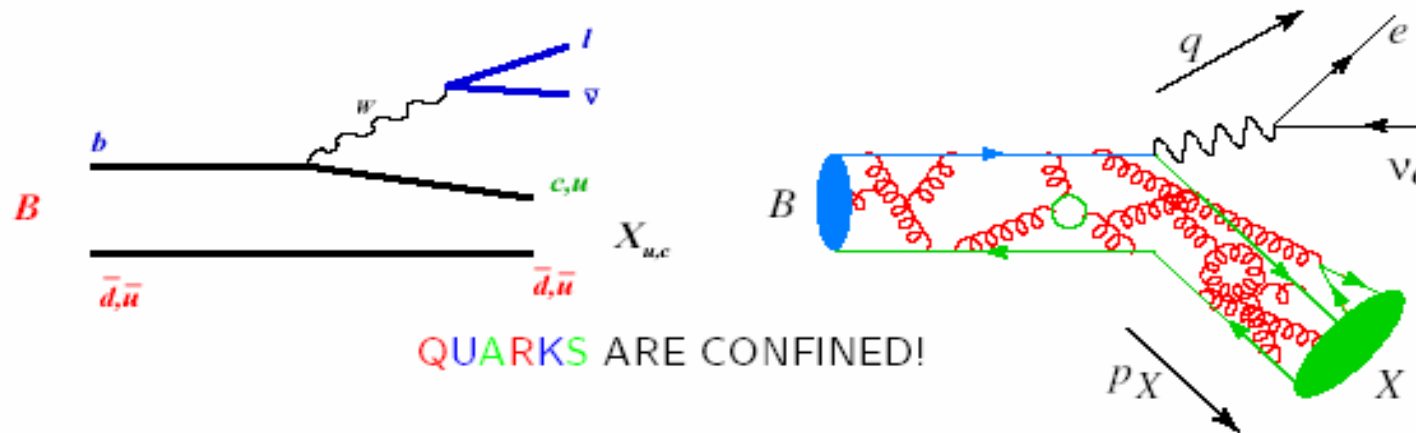
But many more operators appear adding gluons

The tensor current is not conserved and runs between  $M_W$  and  $m_b$   
 We have (at least) 3 scales:  $M_W, m_b, \Lambda_{\text{QCD}}$

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{QCD} \times \text{QED}} + \frac{4G_F}{\sqrt{2}} V_{ts}^* V_{tb} \sum_{i=1}^8 C_i(\mu) Q_i + \dots$$

$L = \ln m_b/M_W$  must be resummed. LO:  $\alpha_s^n L^n$ , NLO:  $\alpha_s^n L^{n-1}$ , NNLO:  $\alpha_s^n L^{n-2}$

# The advantage of being inclusive



The decay of a B meson is not the decay of a free b quark!

**However**  $\Lambda_{\text{QCD}} \ll m_b$  : inclusive decays via OPE admit systematic expansion in  $\alpha_s$  and  $\Lambda_{\text{QCD}}/m_b$  starting with  $\Lambda_{\text{QCD}}^2/m_b^2$ .

➔ Leading term is parton model.

➔ Non-pert corrections are generally small and can be controlled

Dominant (magnetic operator) contributions to inclusive radiative B decays are described by OPE

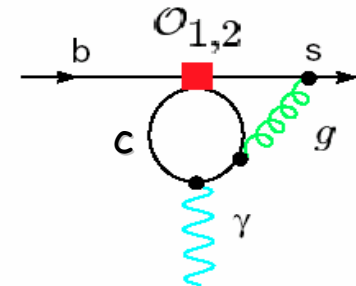
# The main perturbative ingredients

## Process independent:

- **The Wilson coefficients  $C_i$**  (encode the short distance information, initial conditions at scale  $M_W$ )
- **The Anomalous Dimension Matrix** (mixing among operators, determines the evolution of the coefficients, allowing to resum large logs of  $m_b/M_W$ )

## Process dependent:

- **Matrix elements** of  $O_i$ , for ex:

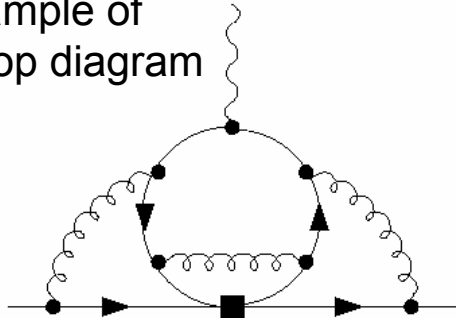


Until recently: QCD at NLO, plus EW & power corrections

# Status of the NNLO enterprise

- **3-loop Wilson coefficients for  $O_7$  and  $O_8$**   
Misiak, Steinhauser, NPB 683 (2004) 277
- **3-loop mixing in  $(O_1, \dots, O_6)$  &  $(O_7, O_8)$  sectors**  
Gorbahn, Haisch, NPB 713 (2005) 291  
Gorbahn, Haisch, Misiak, PRL 95 (2005) 102004
- **4-loop mixing  $(O_1, \dots, O_6)$  into  $O_7$  and  $O_8$**   
Czakon, Haisch, Misiak, hep-ph/0612329
- **2-loop matrix elements of  $O_7$  (real & virt), BLM for others**  
Bieri, Greub, Steinhauser, PRD 67 (2003) 114019  
Blokland, Czarnecki, Misiak, Slusarczyk, Tkachov, PRD 72 (2005) 033014  
Asatryan, Ewerth, Greub, Hurth, Hovhannisyan, Poghosyan, NPB 749 (2006) 325  
Melnikov, Mitov, PLB 620 (2005) 69  
Asatryan, Ewerth, Ferroglia, PG, Greub, hep-ph/0607316  
Asatryan, Ewerth, Gabrielyan, Greub, hep-ph/0611123
- **3-loop matrix elements of  $O_1, O_2$**   
Bieri, Greub, Steinhauser, PRD 67 (2003) 114019  
Misiak, Steinhauser, hep-ph/0609241 (**interpolation**)

example of  
4loop diagram



## Results of 4-Loop Mixing




- numerical effect of  $\mathcal{O}(\alpha_s^3)$  mixing of  $Q_{1-6}$  into  $Q_7$  on  $\mathcal{B}_{\text{SM}}$  is  $-2.4\%$  to  $-4.4\%$  depending on scale
- impact of remaining  $\mathcal{O}(\alpha_s^3)$  mixing of  $Q_{1-6}$  into  $Q_8$  is found to be  $\approx 10$  times smaller

$$\gamma_{1-6,7}^{(2)} = \begin{pmatrix} \frac{150994745}{1062882} + \frac{1272596}{6561} \zeta_3 \\ \frac{138336202}{177147} - \frac{2713672}{2187} \zeta_3 \\ -\frac{58397866}{177147} + \frac{3236560}{2187} \zeta_3 \\ -\frac{5108749081}{2125764} + \frac{2007886}{6561} \zeta_3 \\ \frac{5824017302}{177147} + \frac{112180720}{2187} \zeta_3 \\ \frac{3603565835}{531441} + \frac{15361912}{6561} \zeta_3 \end{pmatrix}$$

Czakon, Haisch, Misiak

Ulrich Haisch, University of Zurich, "Flavour in the era of the LHC" workshop, 3<sup>rd</sup> meeting, 15-17 May, 2006, CERN

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Czakon, Haisch, Misiak in writing
- **2-loop matrix elements of  $O_7$  (real & virt), partially of  $O_8$**   
Bieri, Greub, Steinhauser, PRD 67 (2003) 114019  
Blokland, Czarnecki, Misiak, Slusarczyk, Tkachov, PRD 72 (2005) 033014  
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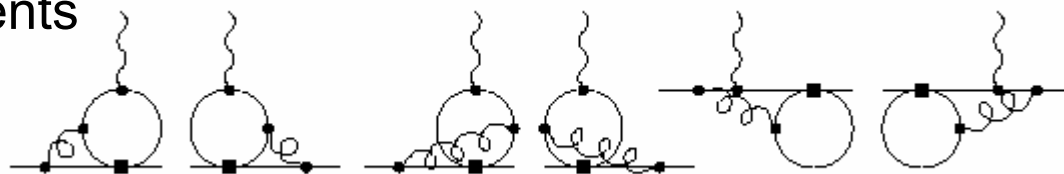


# The charm mass problem

$m_c$  enters the phase factor  
due to normalization

$$C = \left| \frac{V_{ub}}{V_{cb}} \right|^2 \frac{\Gamma[\bar{B} \rightarrow X_c e \bar{\nu}]}{\Gamma[\bar{B} \rightarrow X_u e \bar{\nu}]} \approx 0.58$$

and the NLO matrix elements



LO diagrams vanish: **definition of  $m_c$  is a NNLO issue.**

**Numerically very important** because these are large at NLO:

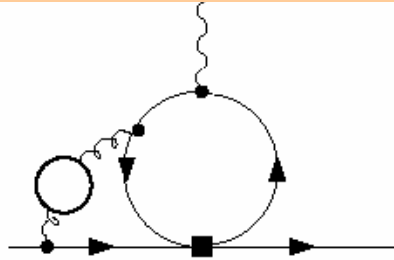
$$m_c(m_c) = 1.25 \pm 0.10 \text{ GeV} \quad m_c(m_b) = 0.85 \pm 0.11 \text{ GeV} \quad m_c(\text{pole}) \sim 1.5 \text{ GeV}$$

But **pole mass has nothing to do with these loops**

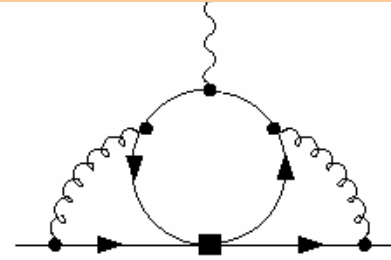
Changing  $m_c/m_b$  from 0.29 (pole) to 0.22 (MSbar) increases  $BR_\gamma$  by 11%

$0.22 \pm 0.04$  **DOMINANT 6% theory error at NLO**

# Interpolation in $m_c$ (Misiak & Steinhauser)



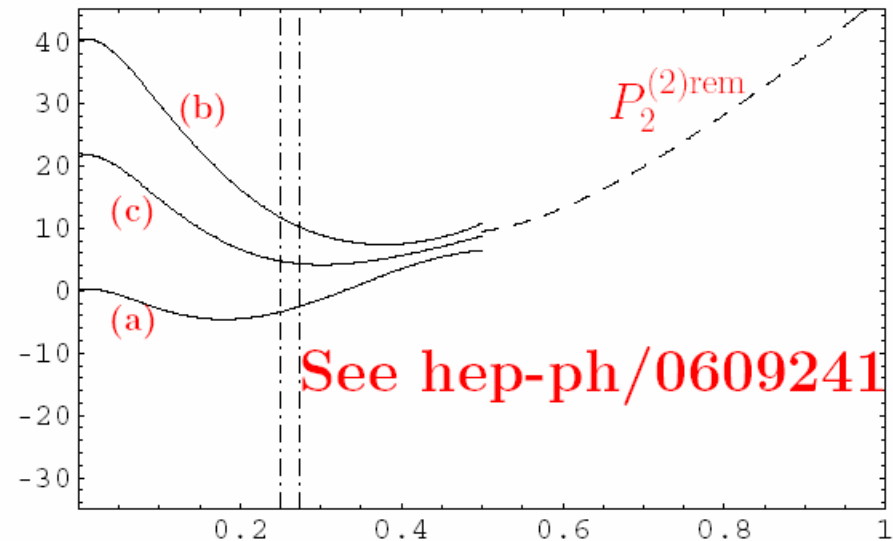
$O(\alpha_s^2 n_f)$  diagrams  $\rightarrow$  BLM contribution, known at small and large  $m_c$  Bieri et al



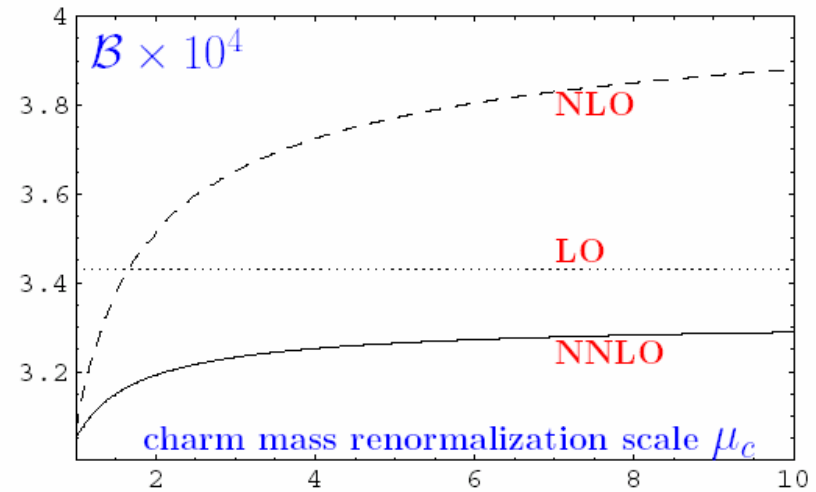
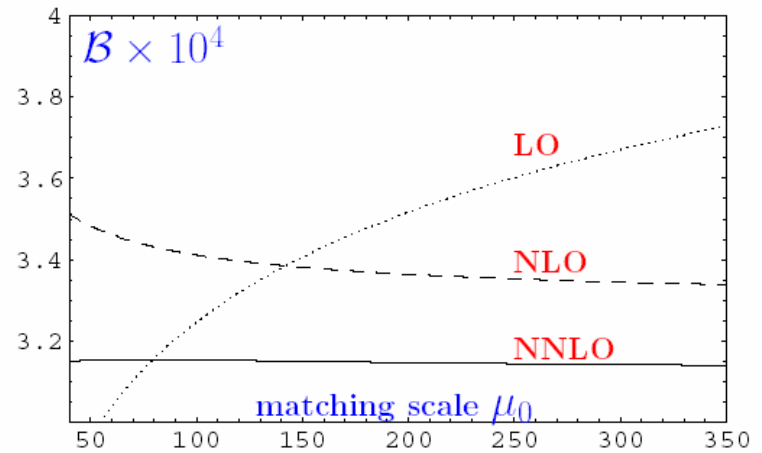
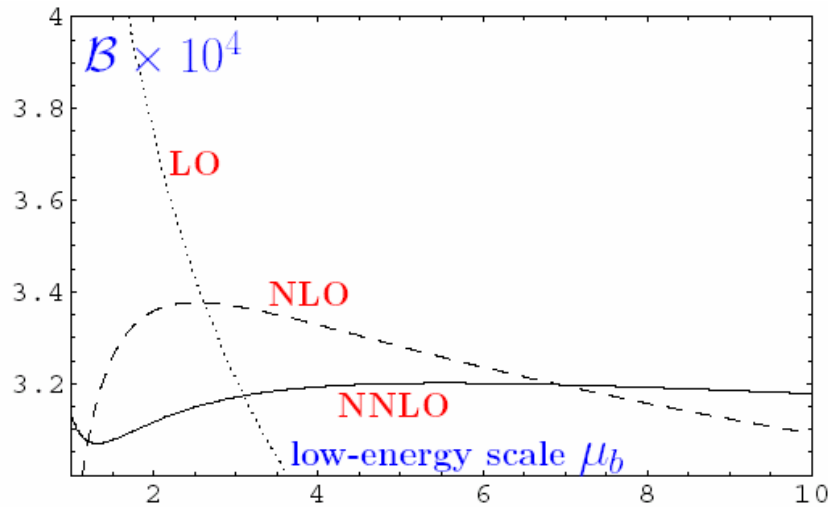
non-BLM contribution only for  $m_c \gg m_b/2$  Misiak & Steinhauser

BLM contributions generally dominant. Since small-heavy  $m_c$  interpolation works very well at NLO and for NNLO-BLM use it at NNLO as a first step

Three ansatz for matching non-BLM terms at  $m_c=0$  are employed and averaged:  
**3% uncertainty on the BR**



# Residual scale dependence



Clear reduction of scale dependence  
Charm mass dependence remains visible

**Higher orders uncertainty ~3%**

central values:  $\mu_b=2.5\text{GeV}$   
 $\mu_c=1.5\text{GeV}$  and  $\mu_0=m_t$

# NNLO “estimate” hep-ph/0609232

Including known power corrections, pure OPE result is at the moment

$$\mathcal{B}(\bar{B} \rightarrow X_s \gamma)_{E_\gamma > 1.6 \text{ GeV}}^{\text{NNLO}} = (3.15 \pm 0.23) \times 10^{-4}$$

**Total error ~7% = ± 3% (interpolation) ± 3% (parametric)  
± 3% (higher orders) ± 5% (non-pert)**

At NLO:  $\text{BR}(E_g > 1.6 \text{ GeV}) = 3.58 \cdot 10^{-4}$  (Misiak, PG 2001) and  
the NNLO error was 6% ( $m_c$  scale)+4% (other NNLO)

The present NNLO result (given the same inputs) is not far  
from the edge of that range (-12%)

Various effects (notably but not only BLM, 4loop ADM) lower  
the BR; charm scale is not set. Two scales ( $\mu_c$  and  $\mu_b$ ) interplay.

**NNLO SM Prediction**  
 $3.15 \pm 0.23 \times 10^{-4}$   
 hep-ph/0609232

CLEO Phys. Rev. Lett. 87, 251807 (2001)

BELLE Phys.Lett. B 511, 151 (2001)

BELLE Phys.Rev.Lett.93:061803,2004

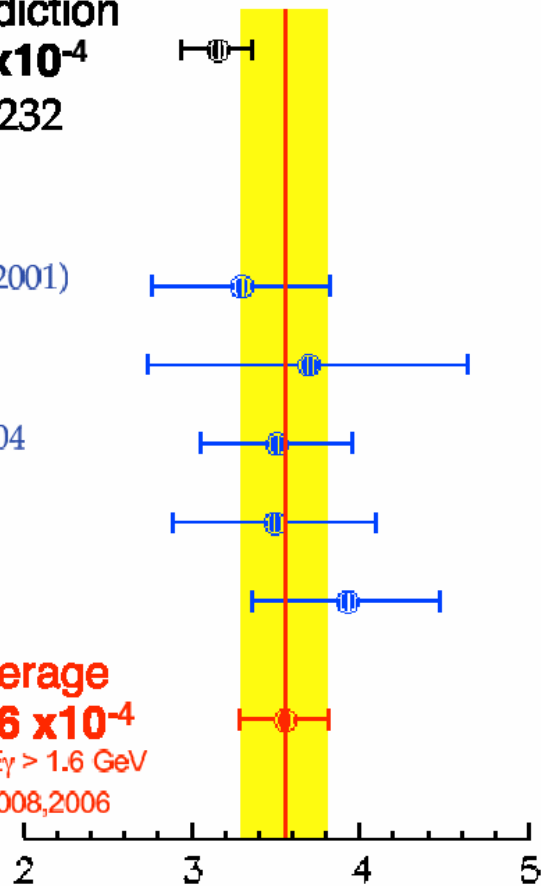
BABAR PRD 72, 052004 (2005)

BABAR hep-ex/0507001

**HFAG Average**

$3.55 \pm 0.26 \times 10^{-4}$

Extrapolation to  $E_\gamma > 1.6$  GeV  
 from PRD73:073008,2006



$BR(b \rightarrow s\gamma)_{E_\gamma > 1.6 \text{ GeV}} \times 10^{-4}$

# The disadvantage of cuts

The OPE breaks down close to boundaries:  
unreliable unless cut is low enough.

In the endpoint region  $E_\gamma \sim 2$  GeV new dynamics:  
distribution function (Shape Function=SF)  
and Sudakov resummation.

Nobody questions validity of OPE for  $E_{\text{cut}} \sim 1-1.2$   
GeV. Experiments cut at 1.8-1.9 GeV.  
**What happens in between?**

In 2001 we proposed  $E_{\text{cut}} = 1.6$  GeV as a  
reference point because SF studies showed  
 $O(<1\%)$  SF effects for  $E_{\text{cut}} < 1.6$  GeV

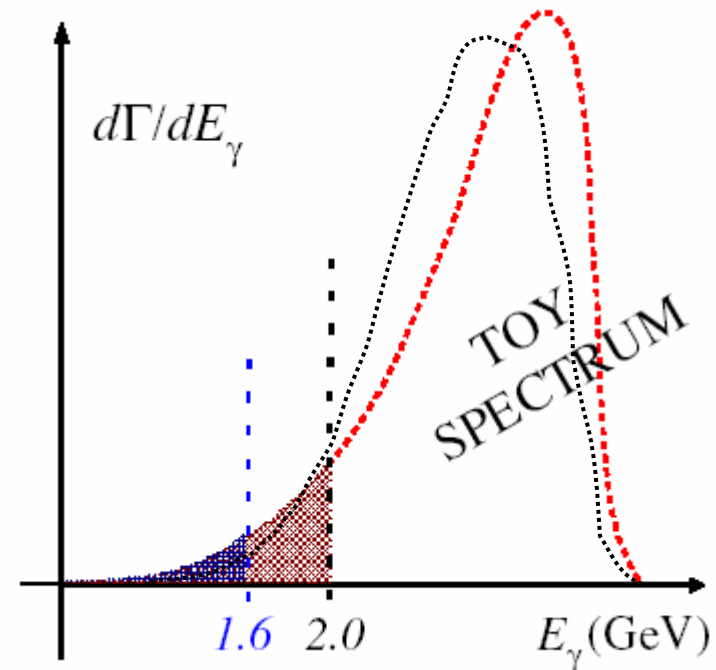
Kagan&Neubert

Confirmed later: we stick to that view

Benson, Bigi, Uraltsev, Andersen& Gardi

But possible effects of scale  $\Delta = m_b - 2E_{\text{cut}}$  would  
lower  $BR(E_{\text{cut}} < 1.6 \text{ GeV})$  by  $\sim 3\%$ . Total tail  
effect from 1.0 to 1.6 GeV would be  $\sim 7 \pm 4\%$  at  
NNLO!

Becher & Neubert



Extrapolation to 1.6 GeV is now  
performed by HFAG using  $b \rightarrow clv$   
input  $\rightarrow$  small 2-3% error counted  
in the experimental budget

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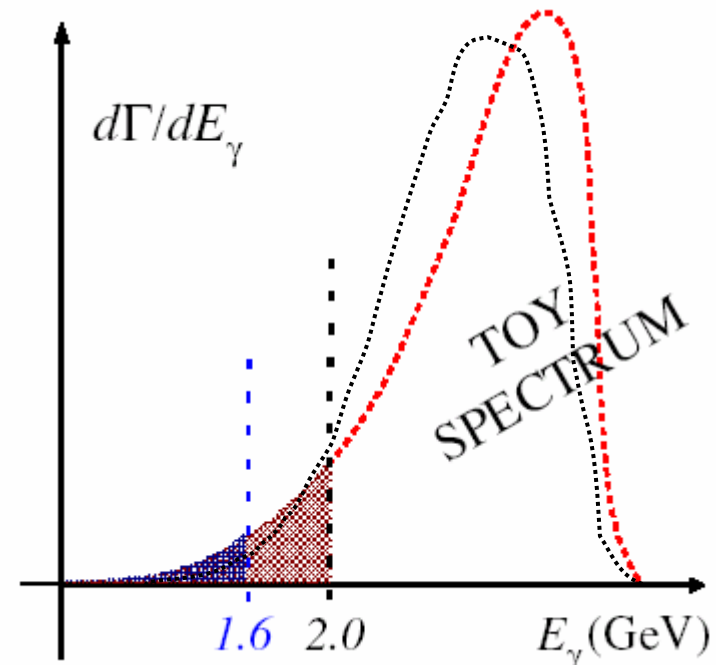
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effect from 1.0 to 1.6 GeV would be  $\sim 7+5-3\%$   
at NNLO!

Becher & Neubert



In the future better to  
compare predictions  
for  $E_{\text{cut}} = 1.8 \text{ GeV}$ !

# More non-perturbative problems

There is **another problem with OPE**: it cannot be applied to the contributions of operators other than  $O_7$  if the photon couples to a light quark.

Some of the additional non-perturbative effects are related to  $O_{1,2}$  matrix elements and scale like  $1/m_c^2$  (they are included)

Voloshin, Ligeti et al, Buchalla et al

Similar  $O(\alpha_s \Lambda/m_b)$  effects in  $O_7$ - $O_8$  have been recently estimated by VIA: 0--3%

Lee, Neubert, Paz

In the presence of a hard gluon, however, the suppression of non-OPE effects in  $O_{1,2}$  m.e. is only  $\alpha_s \Lambda/m_b$  No estimate at present.

**they could be important! → 5% uncertainty in BR**



# What's next

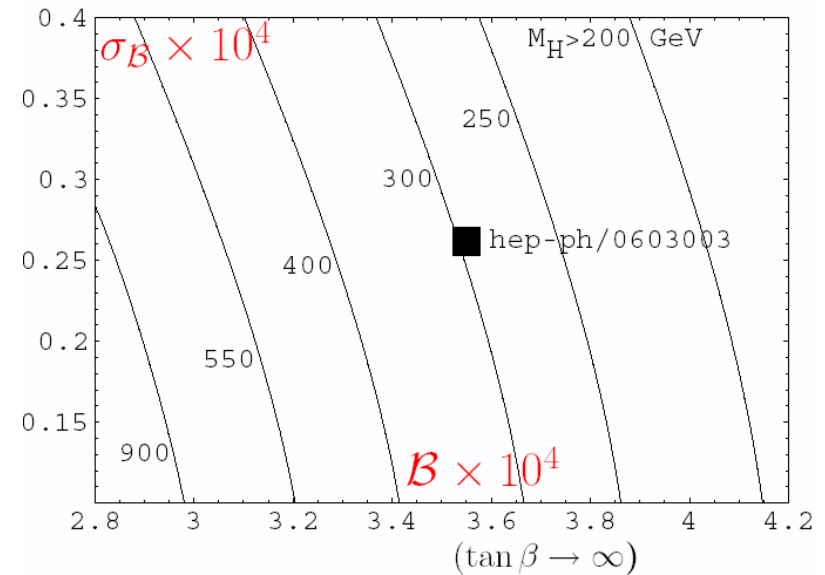
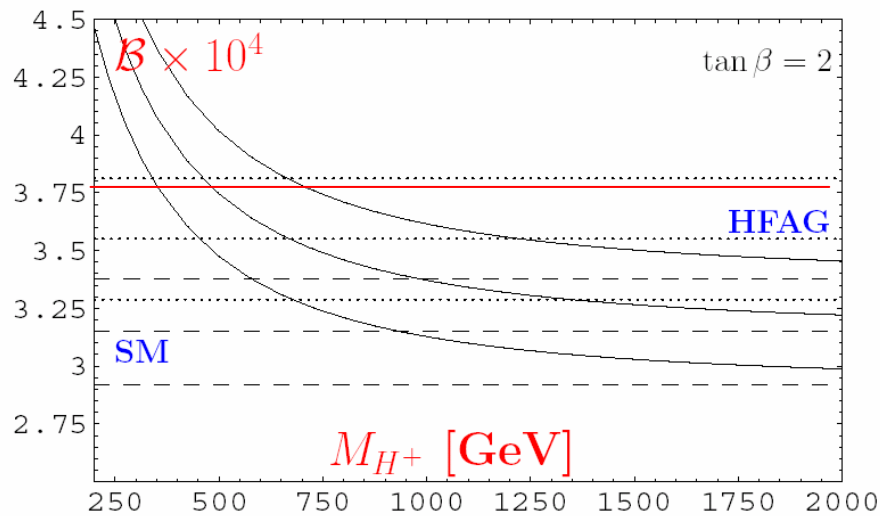
- **Full  $m_c$  dependence** of the NNLO contribution to BR beyond BLM (most difficult step:  $O_{1,2}$  matrix elements)
- **Study of all  $O(\alpha_s \Lambda/m_b)$  contributions**: any idea beyond Vacuum Insertion Approximation by Lee, Neubert, Paz?
- Understanding whether the tail of the spectrum can really be computed precisely and how.
- Calculation of **other missing NNLO** matrix elements (78, 88 contributions). Some known BLM corrections (Ligeti et al) to the spectrum are not yet included in our result

# Impact on New Physics?

- Fortunately, no need for NNLO calculations! but higher order effects sometimes important
- However, status of theory predictions in models of new physics **not always satisfactory**. NLO calculations completed only in a few cases: 2HDM, LR,...
- Recently also MSSM with Minimal Flavor Violation (Degrassi, PG, Slavich, hep-ph/0601135): important dependence on scale at which quark/squark alignment is imposed, relevant effects for light superpartners

# Charged Higgs mass in 2HDM-II

Strongest indirect bound from radiative decays, much stronger than direct ones  
 Mild dependence on  $\tan\beta$ , tighter bounds for small  $\tan\beta$



Central values favor  $M_H \sim 650-700$  GeV

$M_H > 295$  GeV @ 95% CL

remains stringent

# Summary

- NNLO calculation of inclusive radiative B decays is quite advanced. New results have decreased the theory uncertainty to level of the experimental one.
- There are a few NNLO contributions still to be computed: most important 3loop charm diagrams.
- **Non-perturbative effects are at the moment the dominant source of uncertainty**
- Bounds on new physics remain stringent. New NLO calculation in MSSM with MFV, more numerical results soon.