

Two-Stage Supernova Collapse and Double Neutrino Signal from SN1987A

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Neutrino signal from SN1987A

Detector	Working mass	Number of events	
		2:52 UT	7:35 UT
LSD	C_nH_{2n} 90 t	5	2
	Fe 200 t		
KII	H_2O 2140 t	3	12
IMB	H_2O 5000 t	0	8
Baksan	C_nH_{2n} 200 t	1	6

LSD signal:

fluctuation?

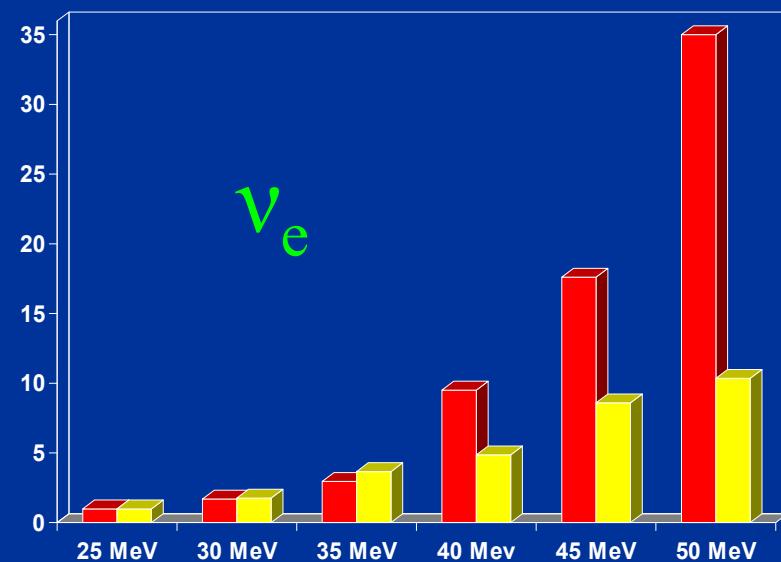
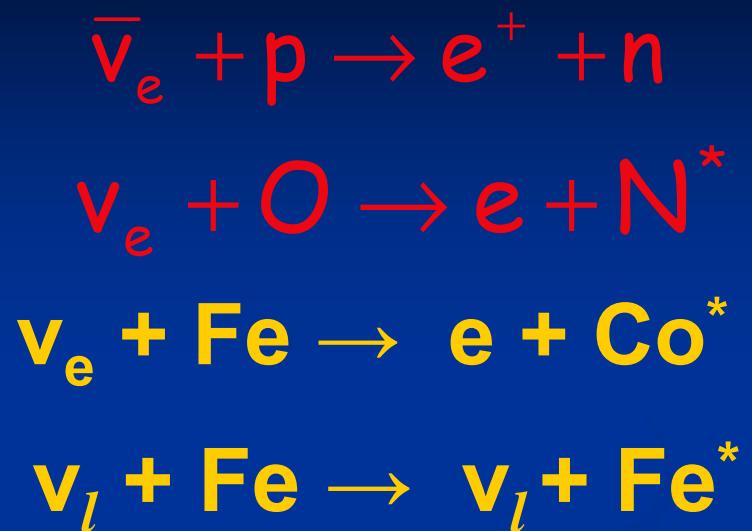
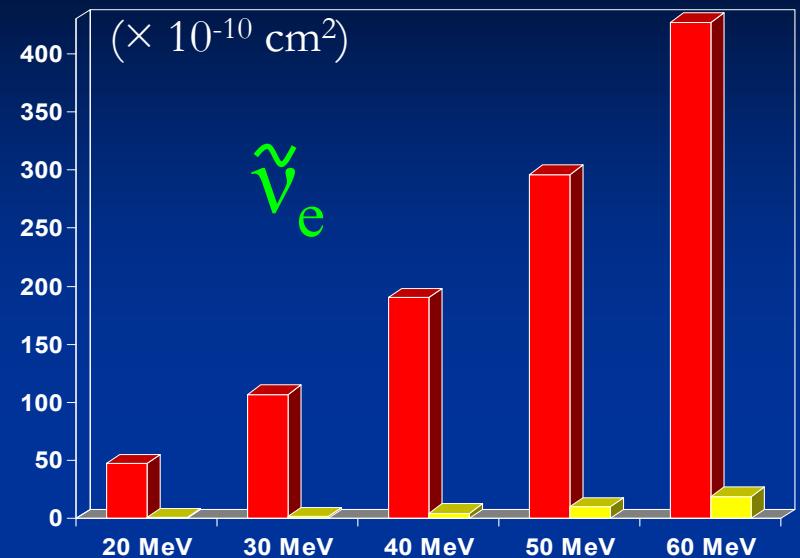
reported probability $< 10^{-3}$

ν from the first stage of 2-stage collapse?

*De Rujula, 1987; Dokuchaev et al., 1987;
Imshennik & Ryazhskaya, 2004*

why other detectors have not seen them?

Kamiokande and LSD sensitivities



Question:

what properties should have a two-stage collapse model to explain the double SN1987A neutrino signal?

Constraints to be considered:

$M_{\text{core}} < 2.2 M_{\text{Sun}}$  Total neutrino energy $< 10^{54}$ erg

Flavour conservation in production processes:

$$N(\nu_e) - N(\bar{\nu}_e) = N(e), \quad N(\nu_\mu) = N(\bar{\nu}_\mu), \quad N(\nu_\tau) = N(\bar{\nu}_\tau)$$

Flavour transformation in the matter of the star (MSW)

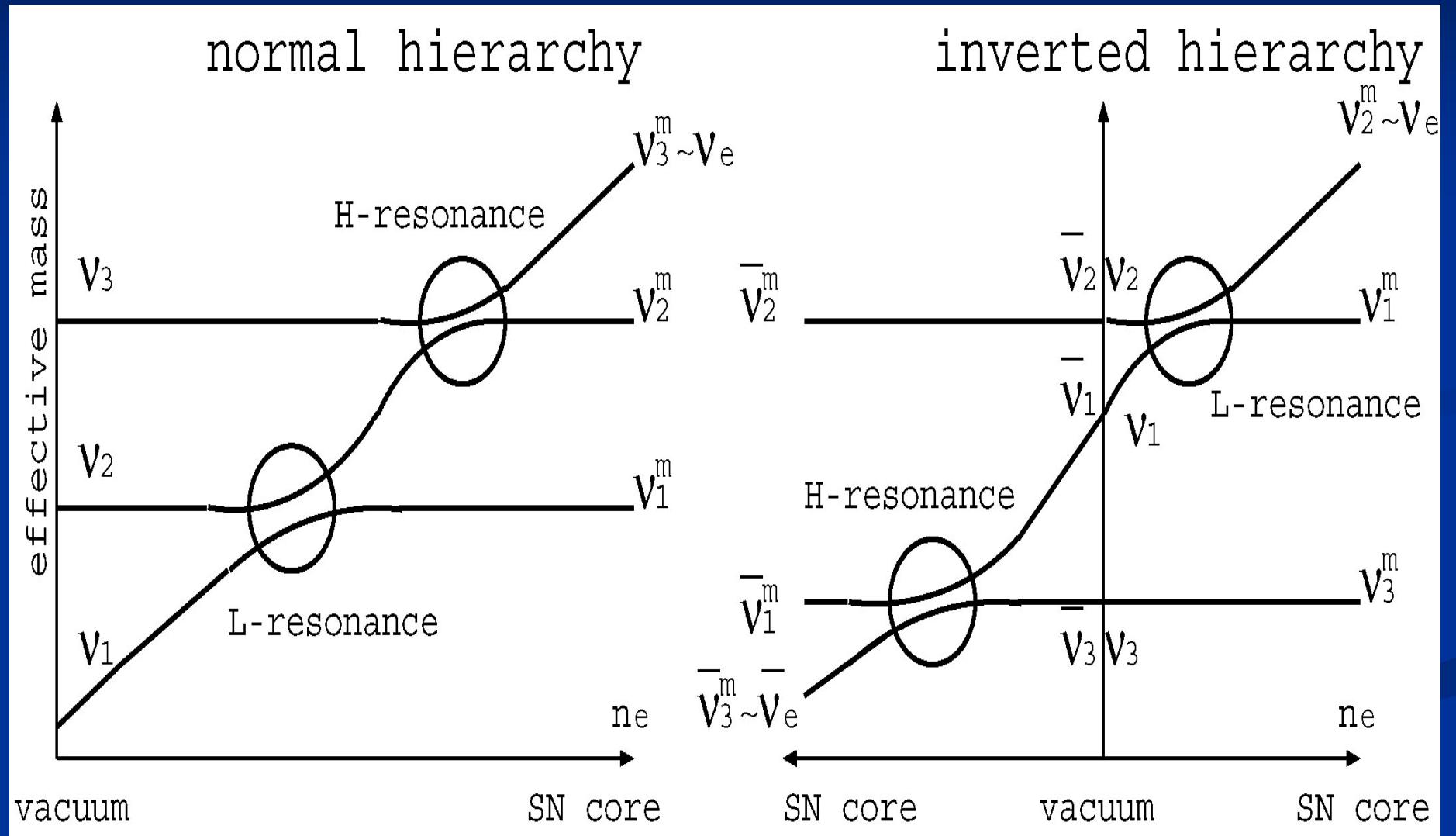
Results

- Production of non-electron neutrinos in the core should be suppressed \longrightarrow no thermal equilibrium!
- $\theta_{13} < 0.003$ \longrightarrow $\tilde{\nu}_e$ production suppressed,
 $45 \text{ MeV} < E_\nu < 50 \text{ MeV}$
- $\theta_{13} > 0.03$, normal neutrino mass hierarchy
 $\tilde{\nu}_e$ production suppressed, $E_\nu > 60 \text{ MeV}$
- $\theta_{13} > 0.03$, inverted neutrino mass hierarchy
 $\tilde{\nu}_e$ production allowed, $30 \text{ MeV} < E_\nu < 45 \text{ MeV}$

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Appendix



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$$F_e = p F_e^0 + (1-p) F_x^0$$

$$F_{\bar{e}} = \bar{p} F_{\bar{e}}^0 + (1-\bar{p}) F_x^0$$

$$4F_x = (1-p) F_e^0 + (1-\bar{p}) F_{\bar{e}}^0 + (2+p+\bar{p}) F_x^0$$

$$p \equiv |U_{e2}|^2 P_H + |U_{e3}|^2 (1 - P_H) \approx |U_{e2}|^2 P_H$$

$$\bar{p} \equiv |U_{e1}|^2 \bar{P}_H + |U_{e3}|^2 (1 - \bar{P}_H) \approx |U_{e1}|^2 \bar{P}_H$$