

# Natural SUSY Dark Matter: A Window on the GUT Scale

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**Is this a natural explanation for dark matter?**

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Ideal candidates for dark matter.



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**This sounds like fine-tuning.**

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... and remember that the MSSM is an **effective** theory.

# Quantifying fine-tuning

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If  $\Delta_a^\Omega = 100$ , a **1%** change in  $a$  gives a **100%** change in  $\Omega_{CDM} h^2$ .

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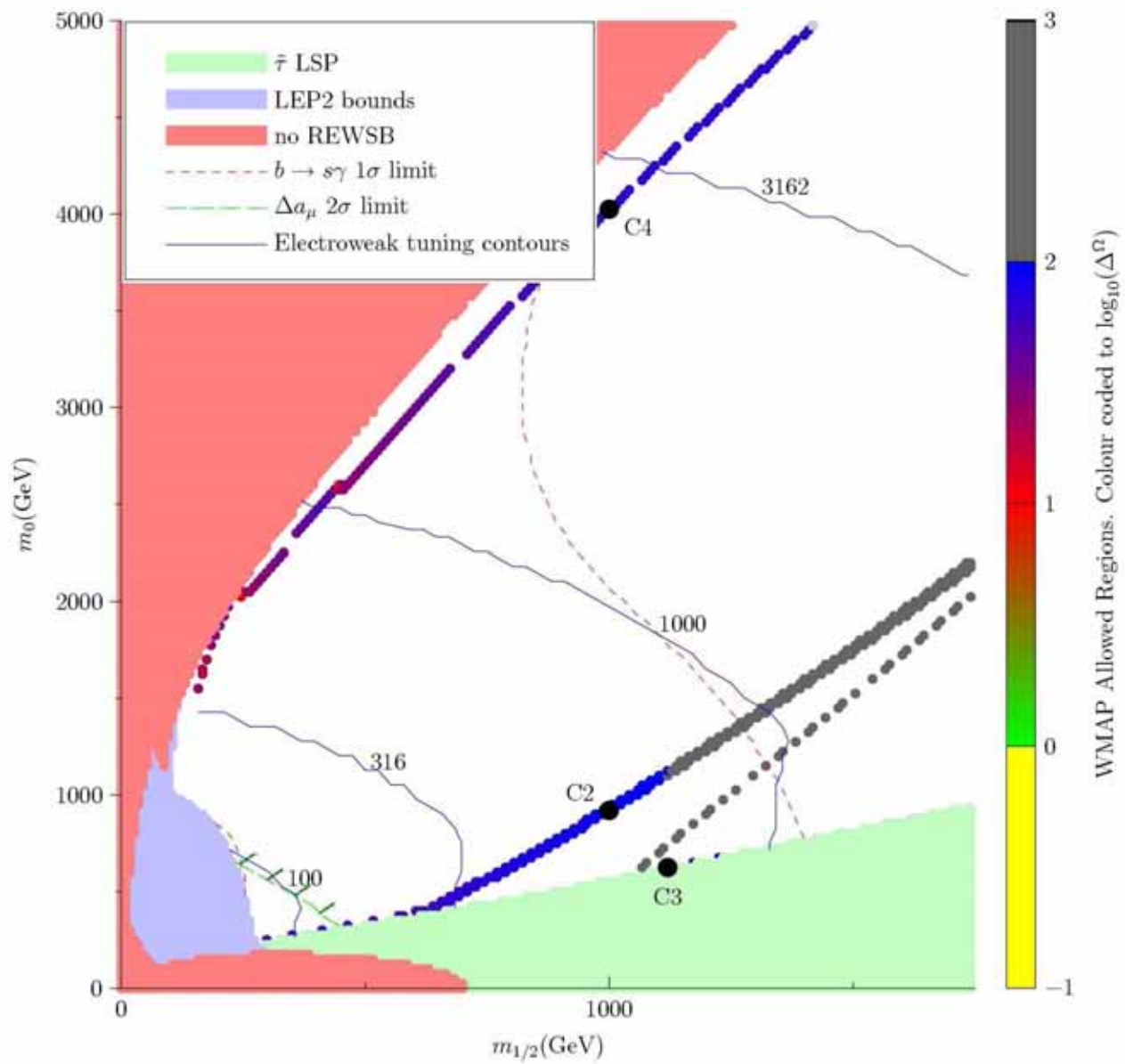
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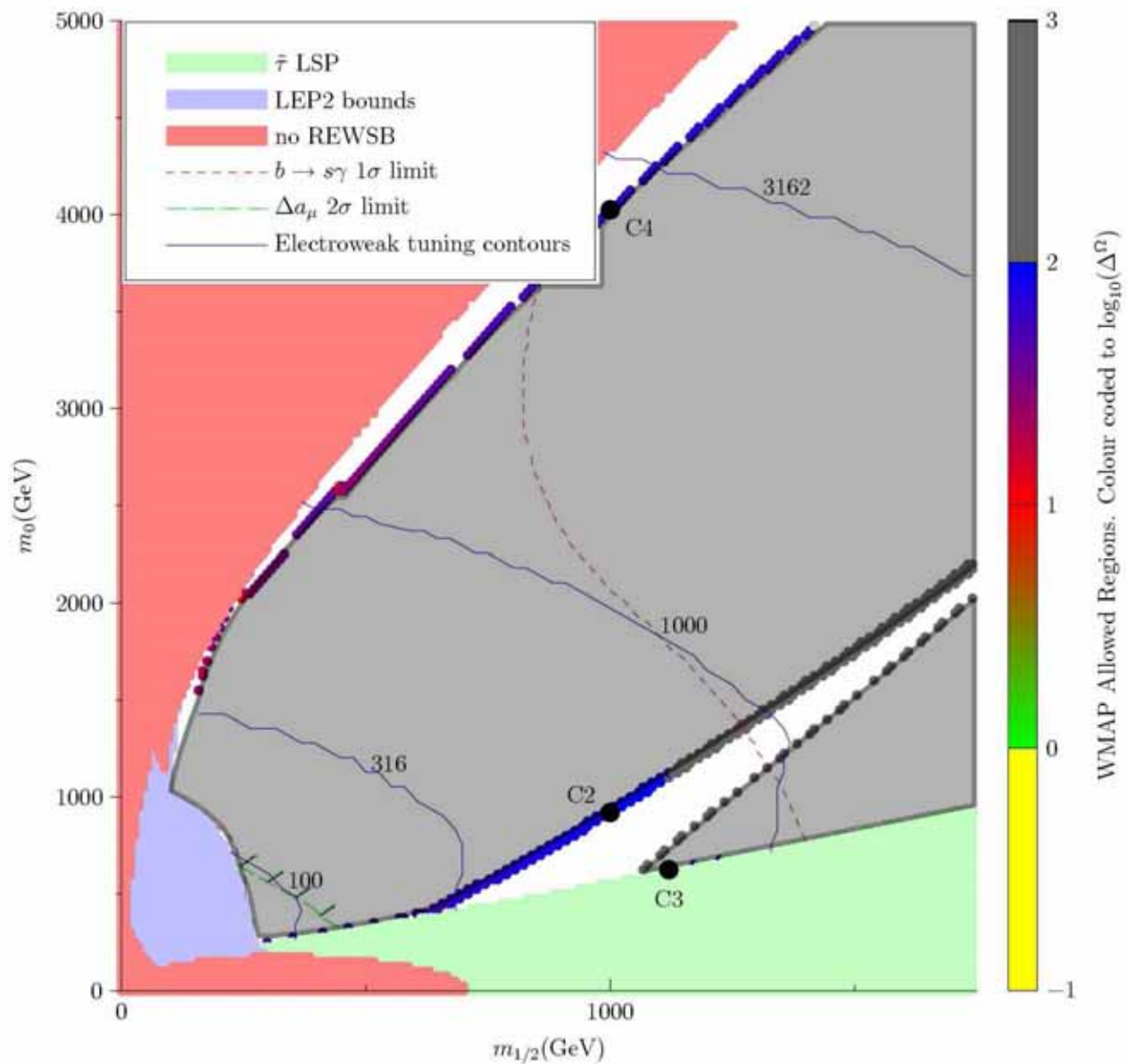
The masses are set at  $m_{GUT}$  and run (using SoftSusy) to  $m_{EW}$ .



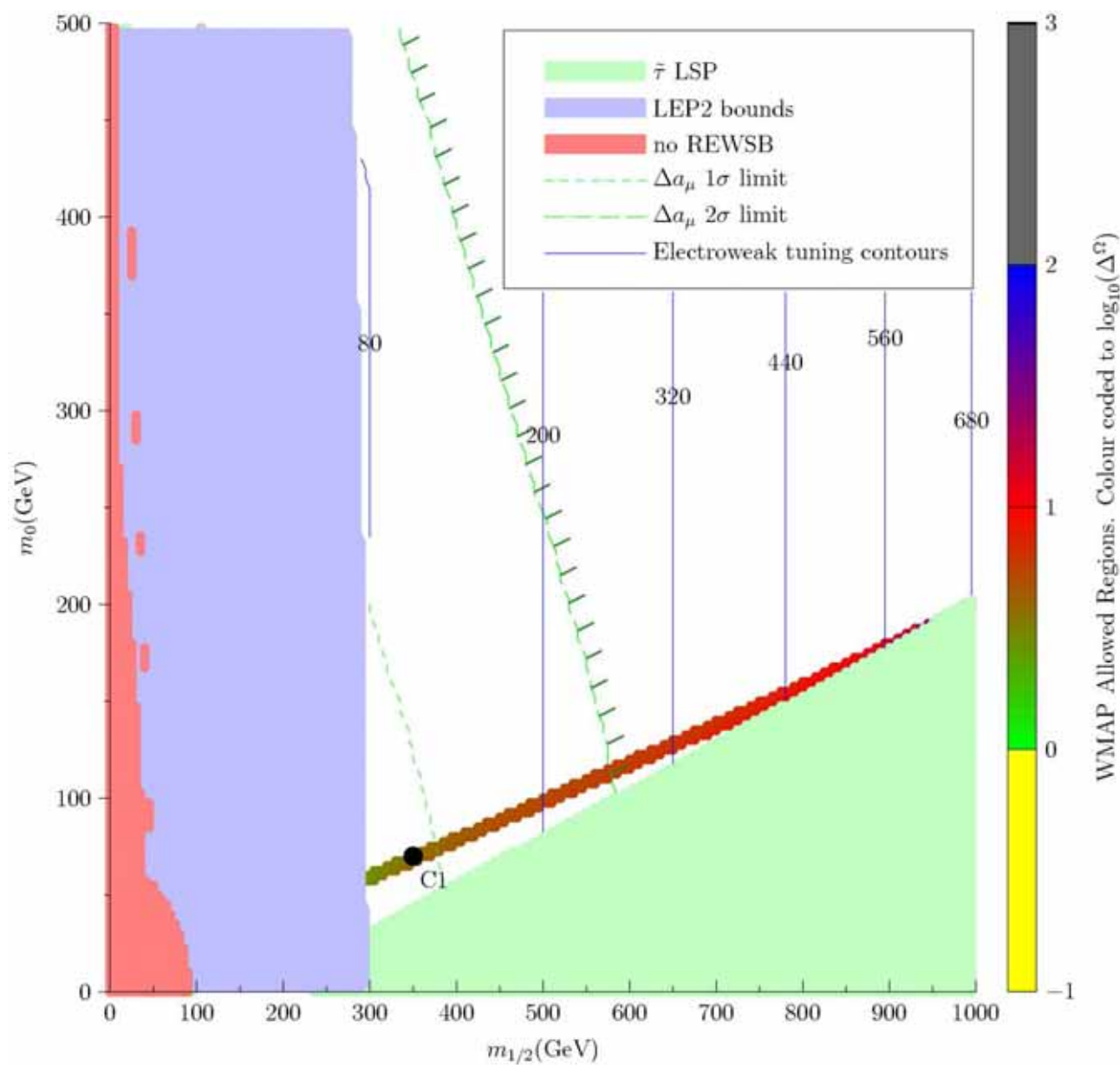
The CMSSM with  $A_0 = 0$ ,  $\tan \beta = 50$  from hep-ph/0609147, S.F.King, J.R.



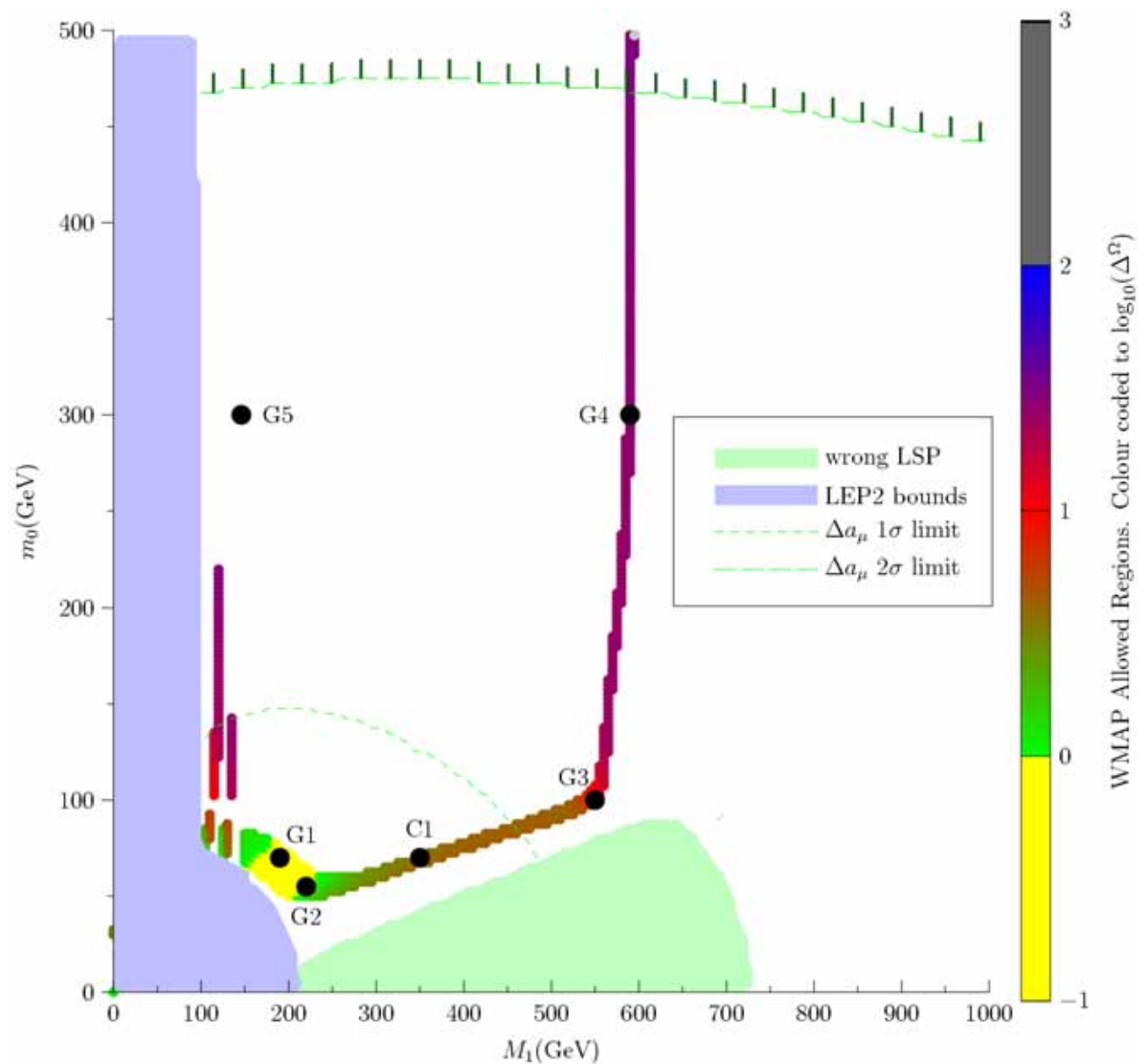
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# The CMSSM with $A_0 = 0$ , $\tan \beta = 10$



# Relaxing the CMSSM: non-universal gauginos



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Bulk region (t-channel $\tilde{f}$ exchange)	$< 1$
slepton coannihilation (low $M_1, m_0$ )	3 – 15
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Therefore the MSSM allows for **natural dark matter**.

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If we **minimise** the coefficients, we **minimise** the dark matter tuning.

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- ③ The same measures of **sensitivity** can be used to relate LHC data to  $\Omega_{CDM}h^2$ .
  - By studying sensitivity to electroweak scale SUSY parameters we can find the sensitivity required of LHC measurements to disprove SUSY.

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- We can identify models that provide the most natural explanation of the observed phenomena.
- We can then make novel predictions for both the LHC, ILC and dark matter detection experiments to test the theory.

# Backup Slides

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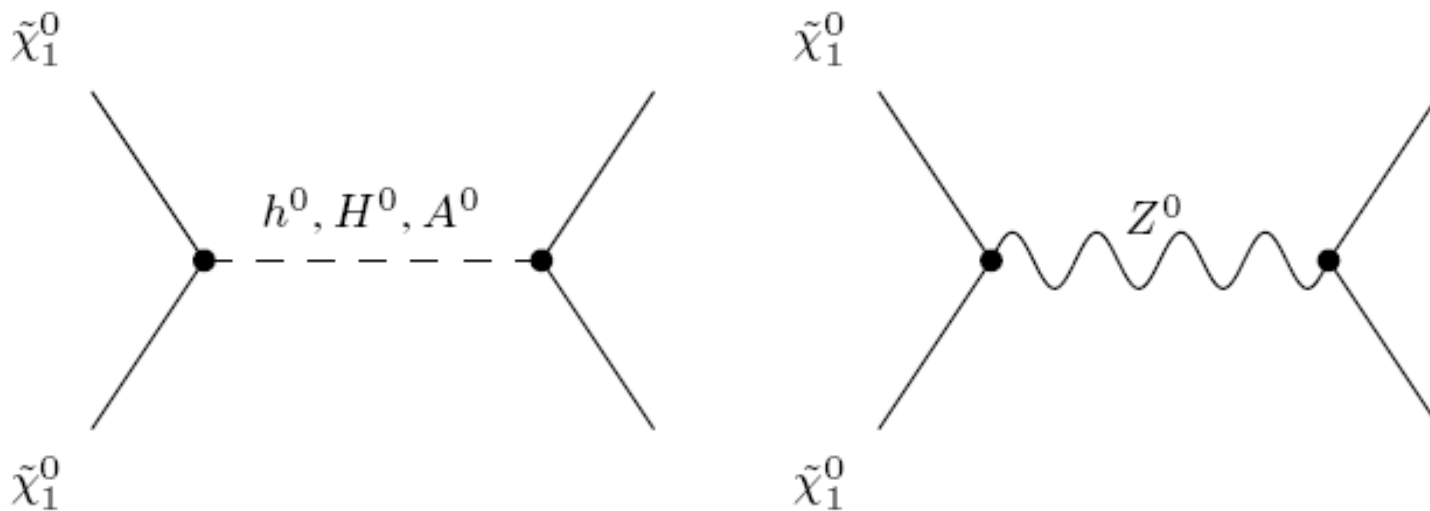
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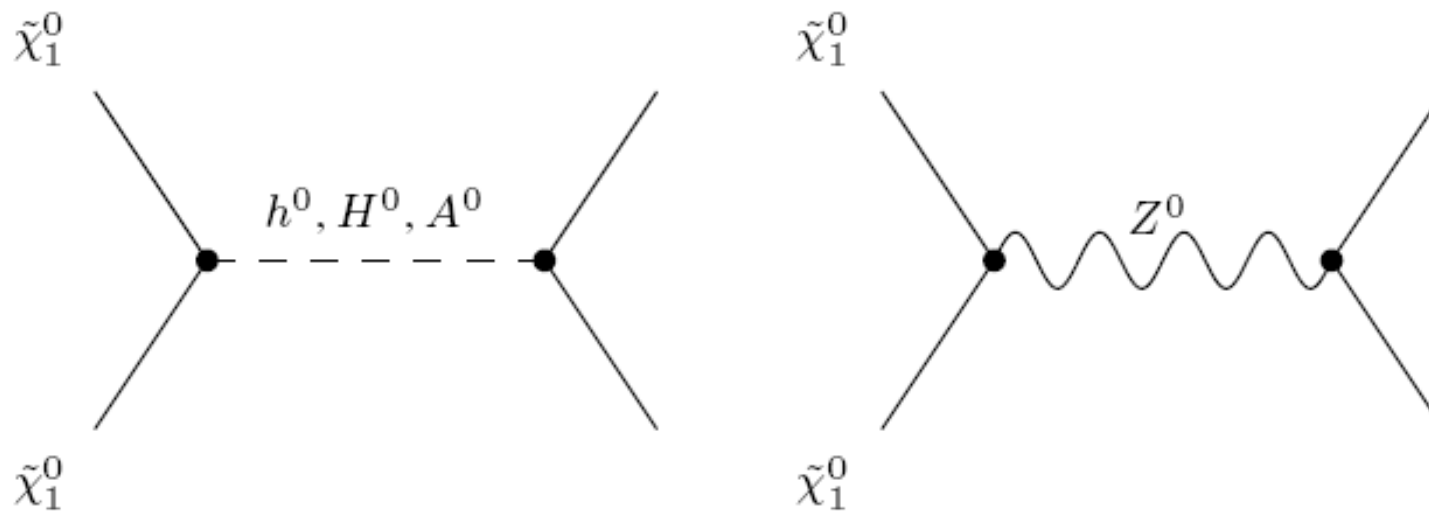


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In this case, annihilation is usually **too efficient**:  $\Omega_{CDM} h^2 \ll \Omega_{CDM}^{WMAP} h^2$   
 except on the edges of the resonance.

# Coannihilation

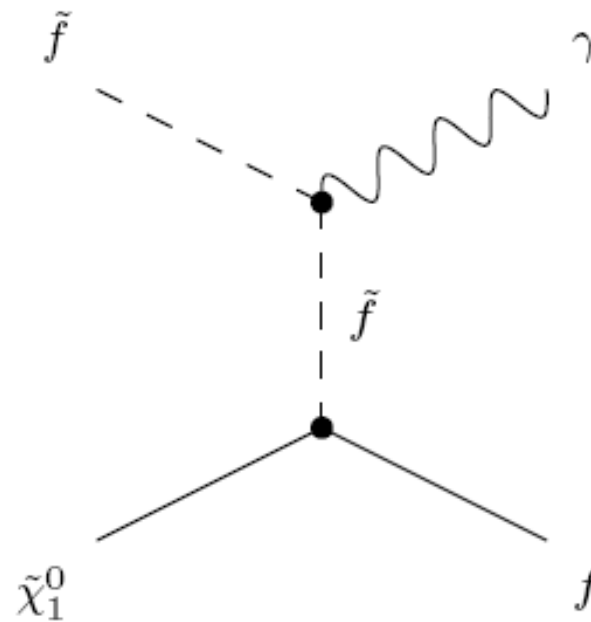
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We must account for annihilations with these particles:



# Stau running in the coannihilation region

