

Nu_2-WP3: R&D of detectors for future high statistics, high precision experiment
(R&D for neutrinoless Double Beta Decay experiments)

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Double Beta Decay (DBD)

A: Mass Number

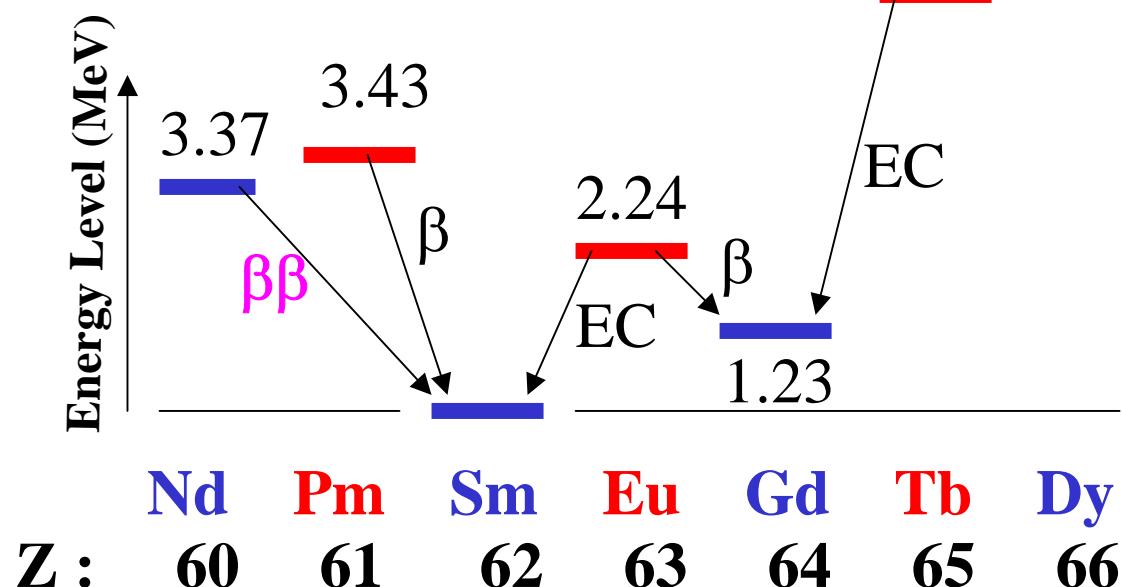
Z: Atomic Number

$\Delta Z = 1$: Beta Decay

$\Delta Z = 2$: DBD

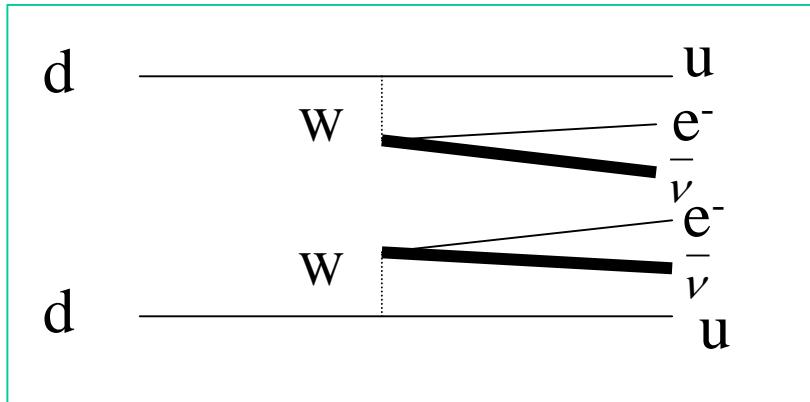
Example
 $A=150$

p n
— —
Odd-Odd
Even-Even



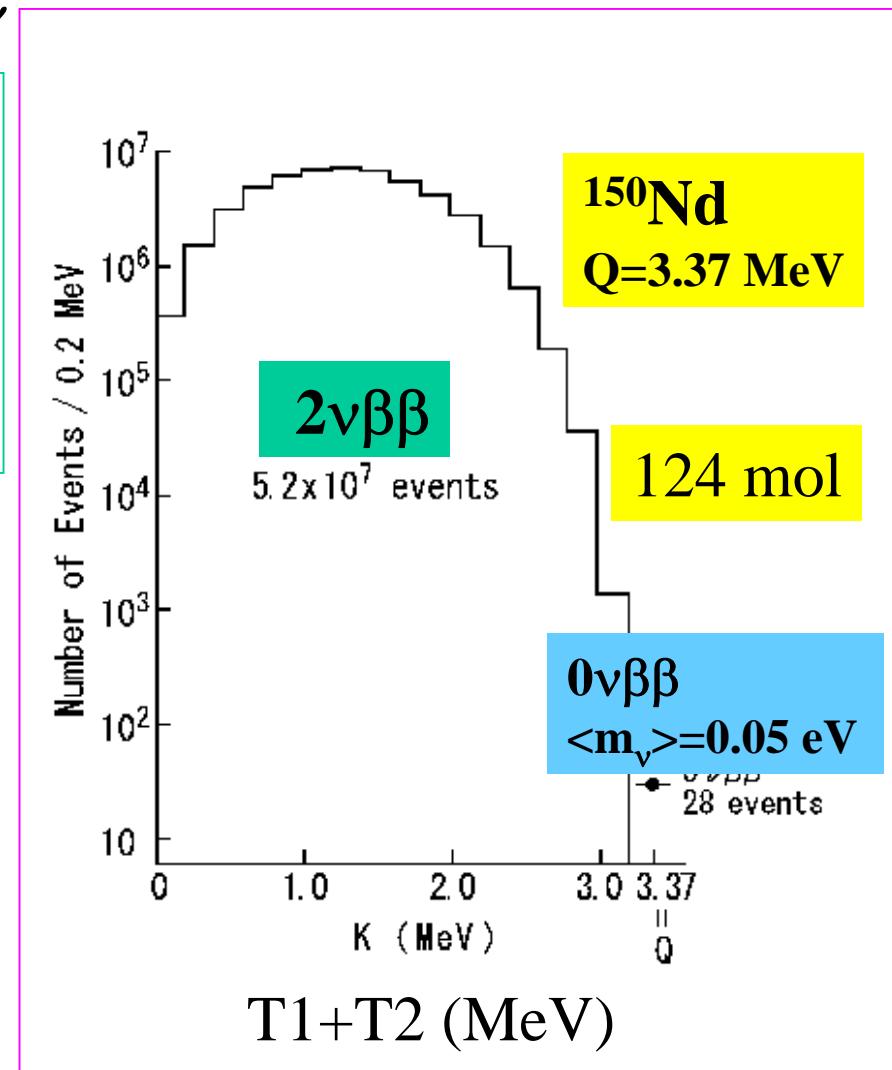
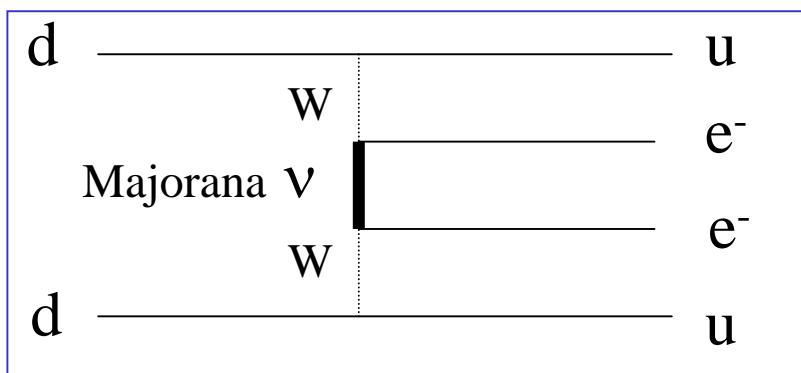
Double Beta Decay (DBD)

$2\nu\beta\beta: (Z, A) \rightarrow (Z + 2, A) + 2e^- + 2\nu$

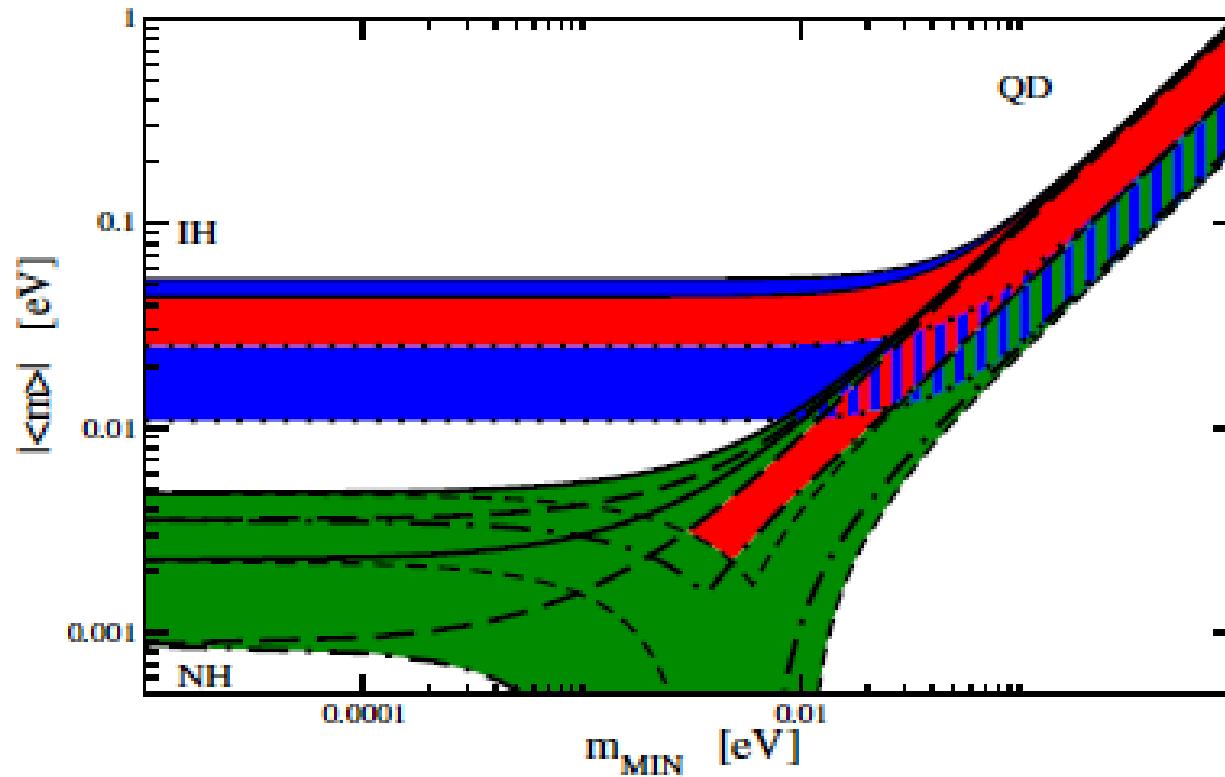


$0\nu\beta\beta: (Z, A) \rightarrow (Z + 2, A) + 2e^-$

Lepton number violation process



Expected effective neutrino mass $\langle m_\nu \rangle$



S. Pascoli, S.T.P., 2006
NDM06

DBD experiment and Sensitivity of Effective Neutrino Mass

Number of events (n) and Half life ($T_{1/2}$)

$$n = (\ln 2)kN_0t / T_{1/2} > \sqrt{BG}$$

k : efficiency, N_0 : number of Nuclei, t : meas. time, BG: background

Half life ($T_{1/2}$) and Effective neutrino mass ($\langle m_\nu \rangle$)
(in the case of mass term dominance)

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu} |M^{0\nu}|^2 \langle m_\nu \rangle^2,$$

$G^{0\nu}$: phase space, $M^{0\nu}$: nuclear matrix element

$$\langle m_\nu \rangle_{SNS} = \left(\frac{(BG)^{1/2}}{\ln 2 k N_0 t G^{0\nu} |M^{0\nu}|^2} \right)^{1/2}$$

Nuclear sensitivity for neutrinoless DBD

Isotope	A %	$Q_{\beta\beta}$ MeV	$S_N 10^{-24} y^{-1}$ (eV) $^{-2}$	Experiment/collaboration
^{48}Ca	0.187	4.276	0.11	CANDLES ^a
^{76}Ge	7.8	2.039	0.22	MAJORANA ^b GENIUS ^c GERDA ^d
^{82}Se	9.2	2.992	0.86	Super-NEMO ^e
^{100}Mo	9.6	3.034	2.02	MOON ^f
^{116}Cd	7.5	2.804	0.90	COBRA ^g CAMEO ^h
^{130}Te	34.5	2.529	0.73	CUORE ⁱ , COBRA ^g
^{136}Xe	8.9	2.467	0.13	EXO ^j , XMASS ^k
^{150}Nd	5.6	3.368	11.3	DCBA ^l

H. Ejiri, J. Phys. Soc. Jpn. 74 (2005) 2101

Status of neutrinoless DBD

Table VI. Limits on neutrino-less double β^- decays. $Q_{\beta\beta}$: Q value for the $0^+ \rightarrow 0^+$ ground state transition. $G^{0\nu}$: kinematical factor (phase space volume)^a in units of 10^{-14} y^{-1} , $T_{1/2}^{0\nu}$: half-life limits in units of 10^{24} y and $\langle m_\nu \rangle$: limit on the effective ν mass in units of eV.

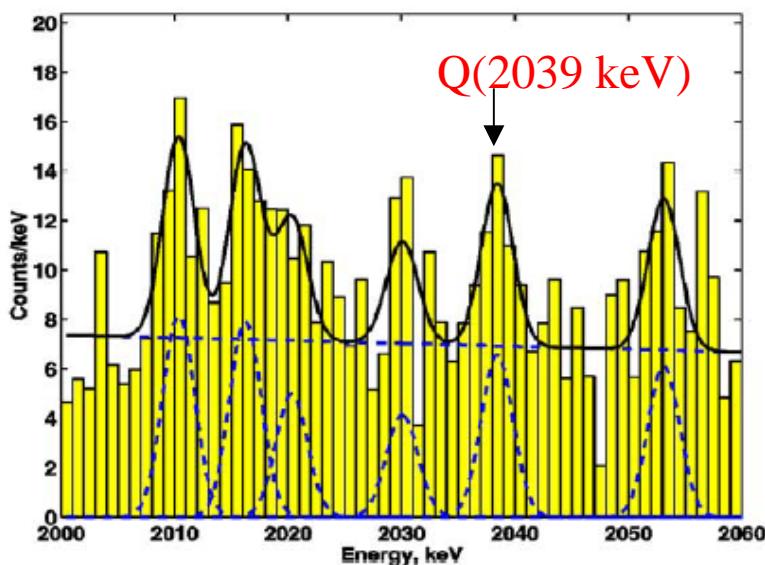
isotope	$Q_{\beta\beta}$ MeV	$G^{0\nu} 10^{-14} \text{ y}^{-1}$	$T_{1/2}^{0\nu} 10^{24}$	$\langle m_\nu \rangle$ eV	Comments
⁴⁸ Ca	4.276	4.46	> 0.014	$< 7.2\text{-}45$	<i>b</i>
⁷⁶ Ge	2.039	0.44	$> 19(12^*)$	$< 0.35(\approx 0.44^*)$	<i>c(c')</i>
⁷⁶ Ge	2.039	0.44	> 16	$< 0.33\text{-}1.35$	<i>d</i>
⁸² Se	2.992	1.89	> 0.19	$< 1.3\text{-}3.2$	<i>e</i>
¹⁰⁰ Mo	3.034	3.17	> 0.35	$< 0.7\text{-}1.2$	<i>f</i>
¹¹⁶ Cd	2.804	3.24	> 0.17	< 1.7	<i>g</i>
¹²⁸ Te	0.876	0.12	> 7.7	$< 1.1\text{-}1.5$	<i>h</i>
¹³⁰ Te	2.529	2.86	> 0.75	$< 0.3\text{-}1.6$	<i>i</i>
¹³⁶ Xe	2.467	3.03	> 0.44	$< 1.8\text{-}5.2$	<i>j</i>
¹⁵⁰ Nd	3.368	13.4	> 0.0012	< 3	<i>k</i>

a: $G^{0\nu} = \ln 2 G^{(0\nu)}$, where $G^{(0\nu)}$ is for $(T_{1/2}^{0\nu})^{-1}$ in ref,²³⁾ *b*: ref.⁹⁷⁾ *c*: ref.⁸¹⁾ *c'*: ref.³¹⁾ *: finite values, *d*: ref.³⁰⁾ *e*: ref.³²⁾ *f*: ref.³²⁾ *g*: ref.⁹⁹⁾ *h*: geochemical method ref.⁹⁰⁾ *i*: ref.¹⁰⁰⁾ *j*: ref.¹⁰¹⁾ *k*: ref.⁸⁵⁾

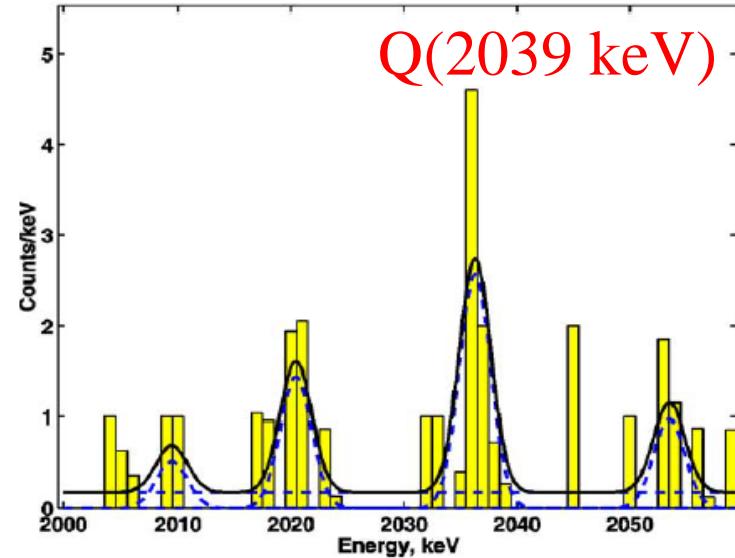
H. Ejiri, J. Phys. Soc. Jpn. 74 (2005) 2101

KKDC claim for neutrinoless DBD of ^{76}Ge from the Heidelberg-Moscow experiment with 14 collaborators

Klapdor,K,D,C, Phys. Lett. B 586 (2004) 198



Sum energy of 2 electrons



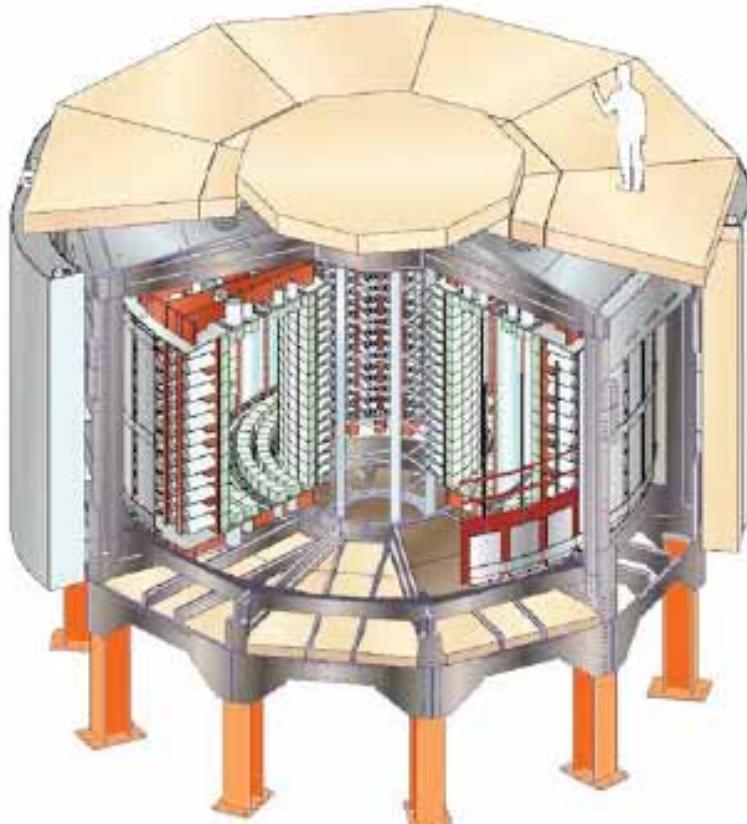
Pulse shape analysis

$(0.69\text{-}4.18)\times 10^{25}$ yr, 1.19×10^{25} yr from data of 1990-2003
 $0.24 \text{ eV} < \langle m_\nu \rangle < 0.58 \text{ eV}$, best value 0.44 eV

The NEMO3 detector

(France, Finlande, Japon, Maroc, République tchèque, R-U, Russie, Ukraine, USA)

@ Frejus Underground Laboratory : 4800 m.w.e.



Source: 10 kg of $\beta\beta$ isotopes
cylindrical, $S = 20 \text{ m}^2$, 60 mg/cm^2

Tracking detector:

99.5 % cells ON

Vertex resolutions :

$$\sigma_{\perp} (\Delta \text{Vertex}) = 0.6 \text{ cm}$$
$$\sigma_{\parallel} (\Delta \text{Vertex}) = 1.3 \text{ cm} \quad (Z=0)$$

Calorimeter:

97% PM+scintillators ON

TDC resolution : 250 ps @ 1MeV

FWHM (1 MeV) 14% (PM 5") 17% (PM 3")

Magnetic field: 25 Gauss

Gamma shield: Iron (18 cm)

Neutron shield: borated water + wood

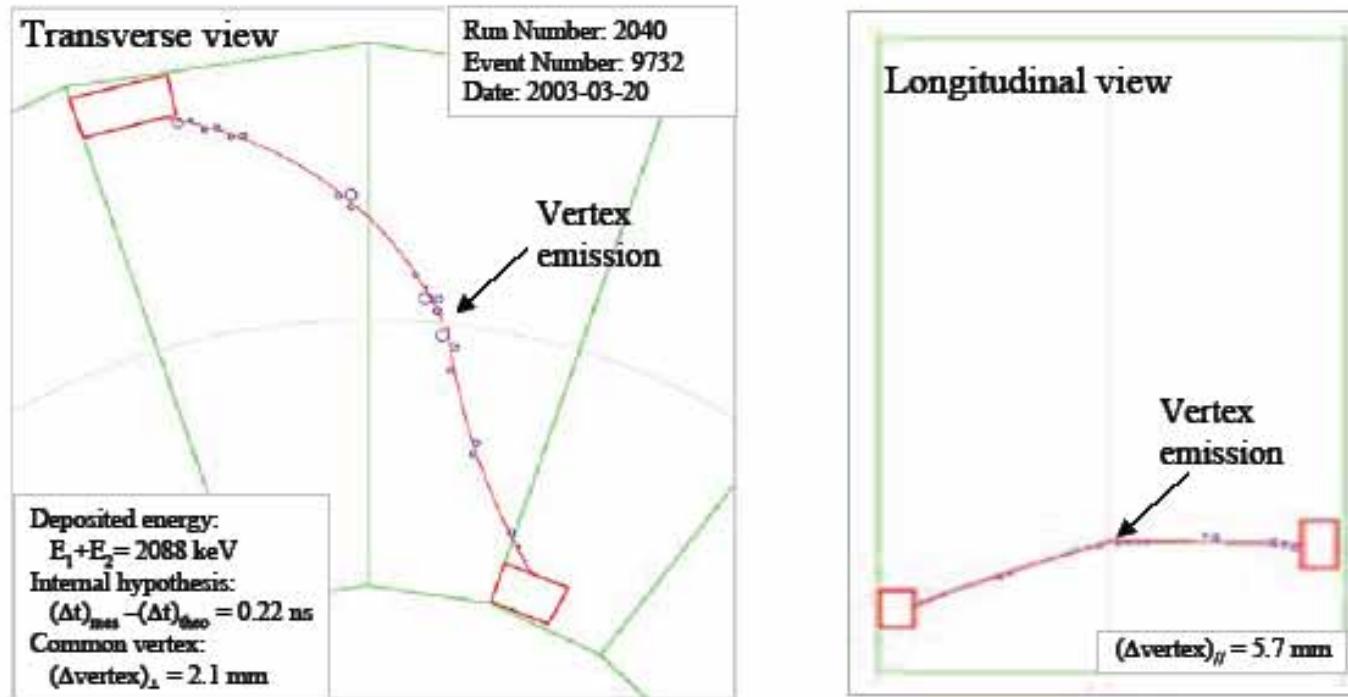


Identification e^- , e^+ , γ and α

X. Sarazin @ NDM06

$\beta\beta$ events selection in NEMO3

Typical $\beta\beta 2\nu$ event observed from ^{100}Mo



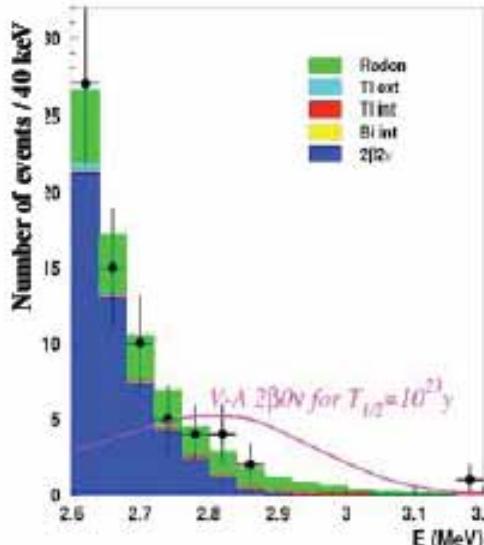
Trigger: 1 PM > 150 keV
3 Geiger hits (2 neighbour layers + 1)
Trigger Rate = 5.8 Hz
 $\beta\beta$ evts: 1 event every 2 minutes

X. Sarazin@NDM06

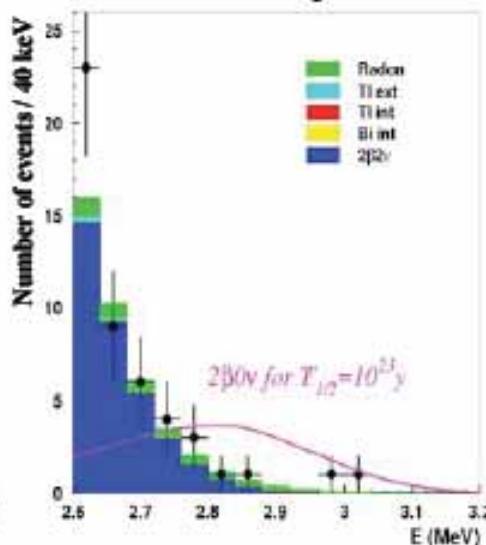
$\beta\beta 0\nu$ results with ^{100}Mo

100Mo, 7 kg

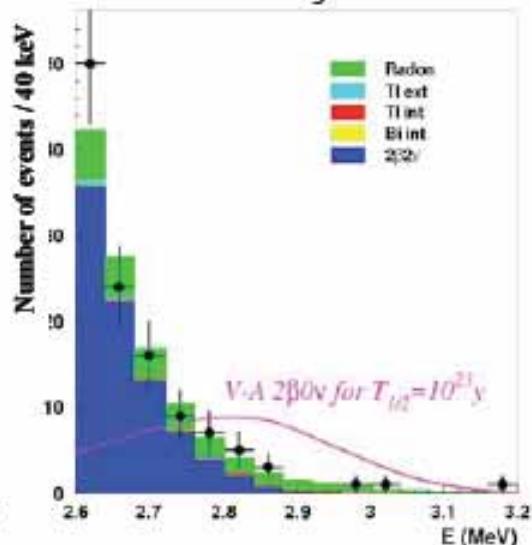
Phase I, High radon
394 days



Phase II, Low radon
299 days



Phase I + II
693 days



Phases I + II

$T_{1/2}(\beta\beta 0\nu) > 5.8 \cdot 10^{23}$ (90 % C.L.)

Expected in 2009

$T_{1/2}(\beta\beta 0\nu) > 2 \cdot 10^{24}$ (90 % C.L.)

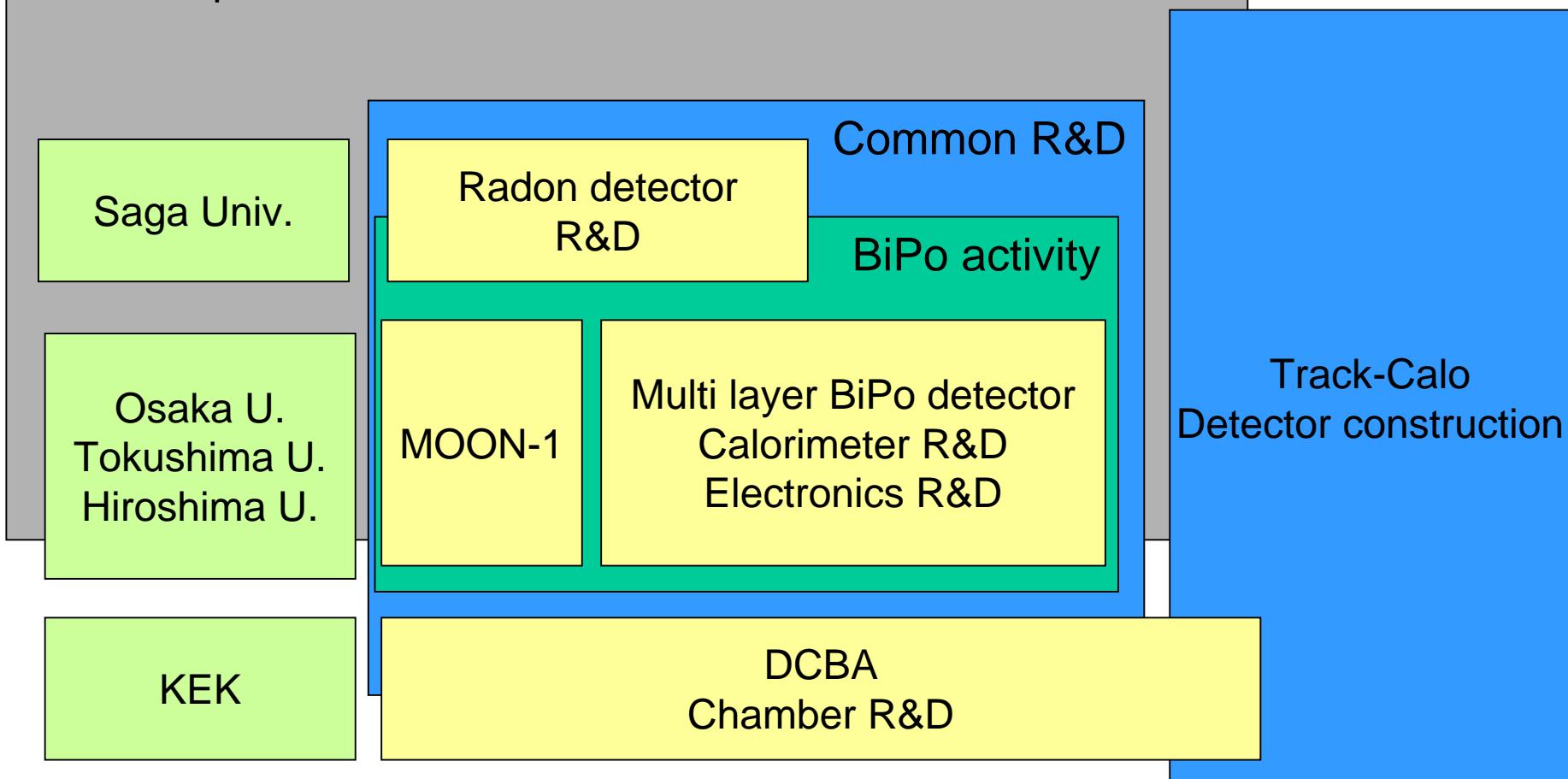
X. Sarazin@NDM06

Double beta decay experiments with tracking detectors

- International collaboration
 - Tracking-calorimetric / Tracking detector
NEMO, MOON, DCBA
- Common R&D for future detectors
 - Enrichment of ^{150}Nd , Energy resolution, BGD reduction
- Collaboration for LIA project
 - France: F. Piquemal, Ch. Marquet, J.S. Ricol / CENBG
X. Sarazin, S. Jullian, D. Lalanne, C. Augier, L. Simard,
J. Argyriades /LAL
 - Japan: N. Ishihara/KEK, M. Nomachi/Osaka-U, H. Ejiri/RCNP-Osaka U,
H. Ohsumi/Saga-U, K. Fushimi/Tokushima-U, Hiroshima-U.

Collaborative work

NEMO/ super NEMO collaboration



Groups interested in Track-Calor detectors in Japan.

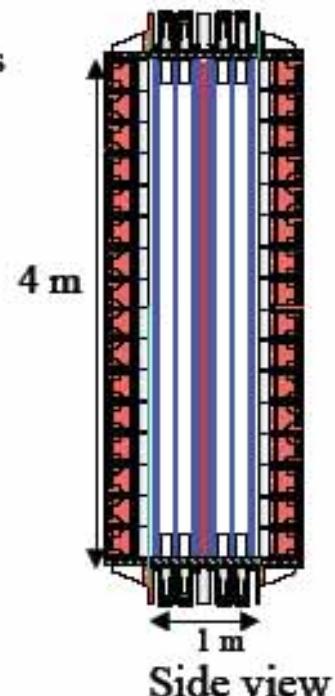
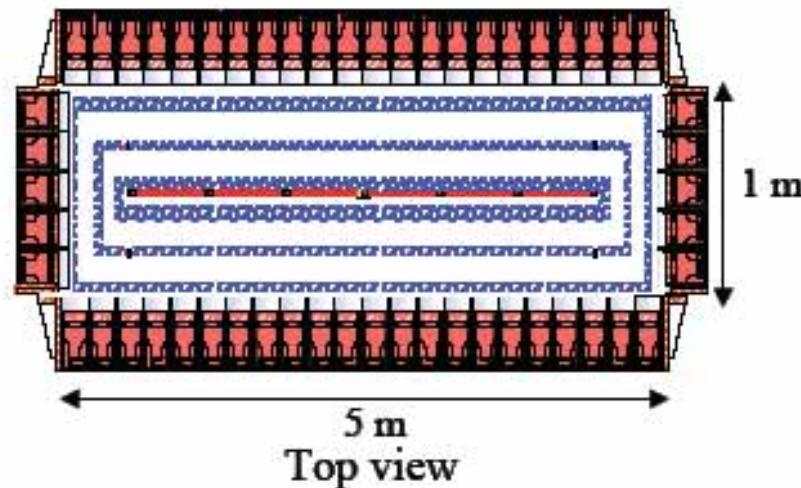
Possible SuperNEMO design

Planar and modular design: ~ 100 kg of enriched isotopes (20 modules \times 5 kg)

1 module: Source (40 mg/cm²) 4 x 3 m²

Tracking : drift chamber ~3000 cells in Geiger mode

Calorimeter: scintillators + PM ~1 000 PM if scint. blocs
 ~ 100 PM if scint. bars



Ch. Marquet @ NDM06

From NEMO3 to SuperNEMO... objectives

$$T_{1/2}(\beta\beta 0\nu) > \ln 2 \times \frac{N_{\text{avo}}}{A} \times \frac{M \times \epsilon \times T_{\text{obs}}}{N_{\text{excl}}}$$

NEMO-3		SuperNEMO
^{100}Mo	Choice of isotope	^{150}Nd or ^{82}Se
7 kg	Isotope mass M	100-200 kg
8 %	Efficiency $\epsilon(\beta\beta 0\nu)$	~ 30 %
$^{208}\text{Tl} < 20 \mu\text{Bq/kg}$ $^{214}\text{Bi} < 300 \mu\text{Bq/kg}$	$N_{\text{excl}} = f(\text{BKG})$ Internal contaminations ^{208}Tl and ^{214}Bi in the $\beta\beta$ foil	$^{208}\text{Tl} < 2 \mu\text{Bq/kg}$ (If ^{82}Se : $^{214}\text{Bi} < 10 \mu\text{Bq/kg}$)
8% @3MeV	Energy resolution FWHM(calorimeter)	4% @3MeV
$T_{1/2}(\beta\beta 0\nu) > 2 \cdot 10^{24} \text{ y}$ $\langle m_\nu \rangle < 0.3 - 1.3 \text{ eV}$		SENSITIVITY
		$T_{1/2}(\beta\beta 0\nu) > 10^{26} \text{ y}$ $\langle m_\nu \rangle < 50 \text{ meV}$

Main R&D tasks:

- 1) $\beta\beta$ source production 2) Radiopurity 3) Energy resolution

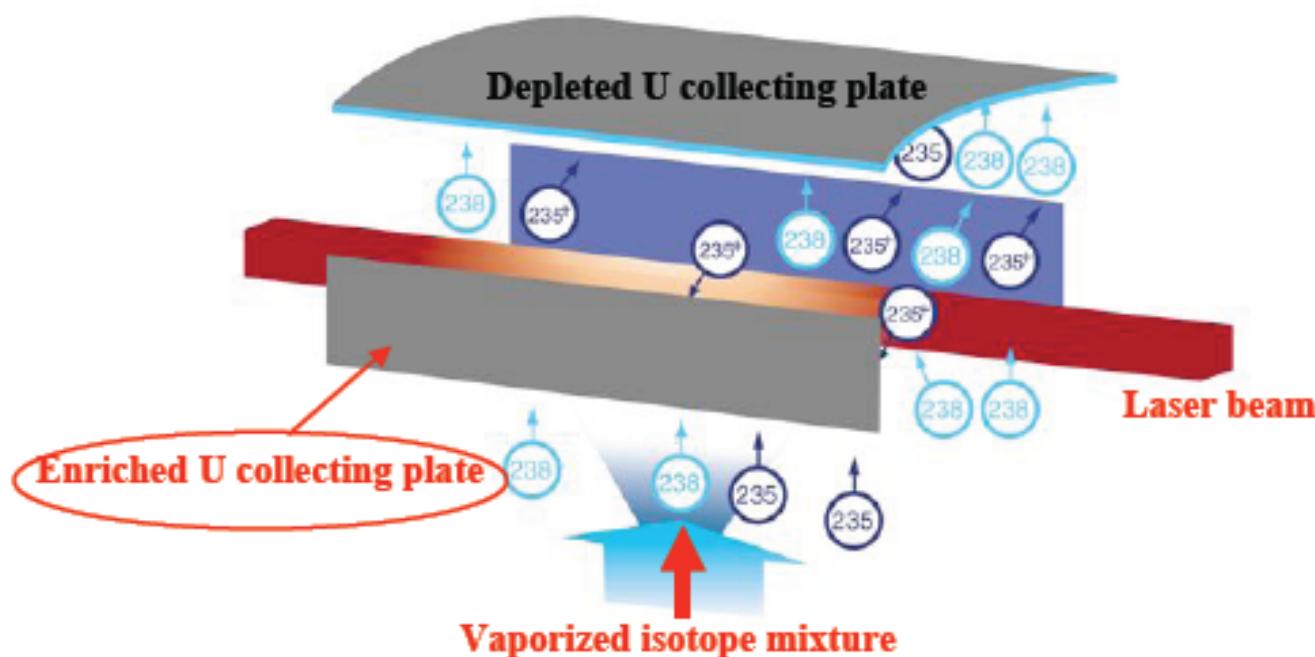
+ Tracking, Electronics, DAQ, Software, Mechanics, Theory

Ch. Marquet – NDM06 – September 2006

^{150}Nd production: The Laser Method (AVLIS)

AVLIS: Atomic Vapor Laser Isotope Separation

Selective photoionization based on :
isotope shifts in the atomic absorption optical spectra
 $\text{U} + 3 \text{ selective photons} \rightarrow ^{235}\text{U}^+ + \text{e}^-$



Ch. Marquet – NDM06 – Septembre, 2006

^{150}Nd production: possibility in France ?

2000-2003: R&D with MENPHIS demonstrator facility (CEA/Pierrelatte - France)



Main results for the process :

- 204 kg of enriched uranium at $\approx 2.5\%$ mean (predicted) value of enrichment
- Production raise: few kg/h for U



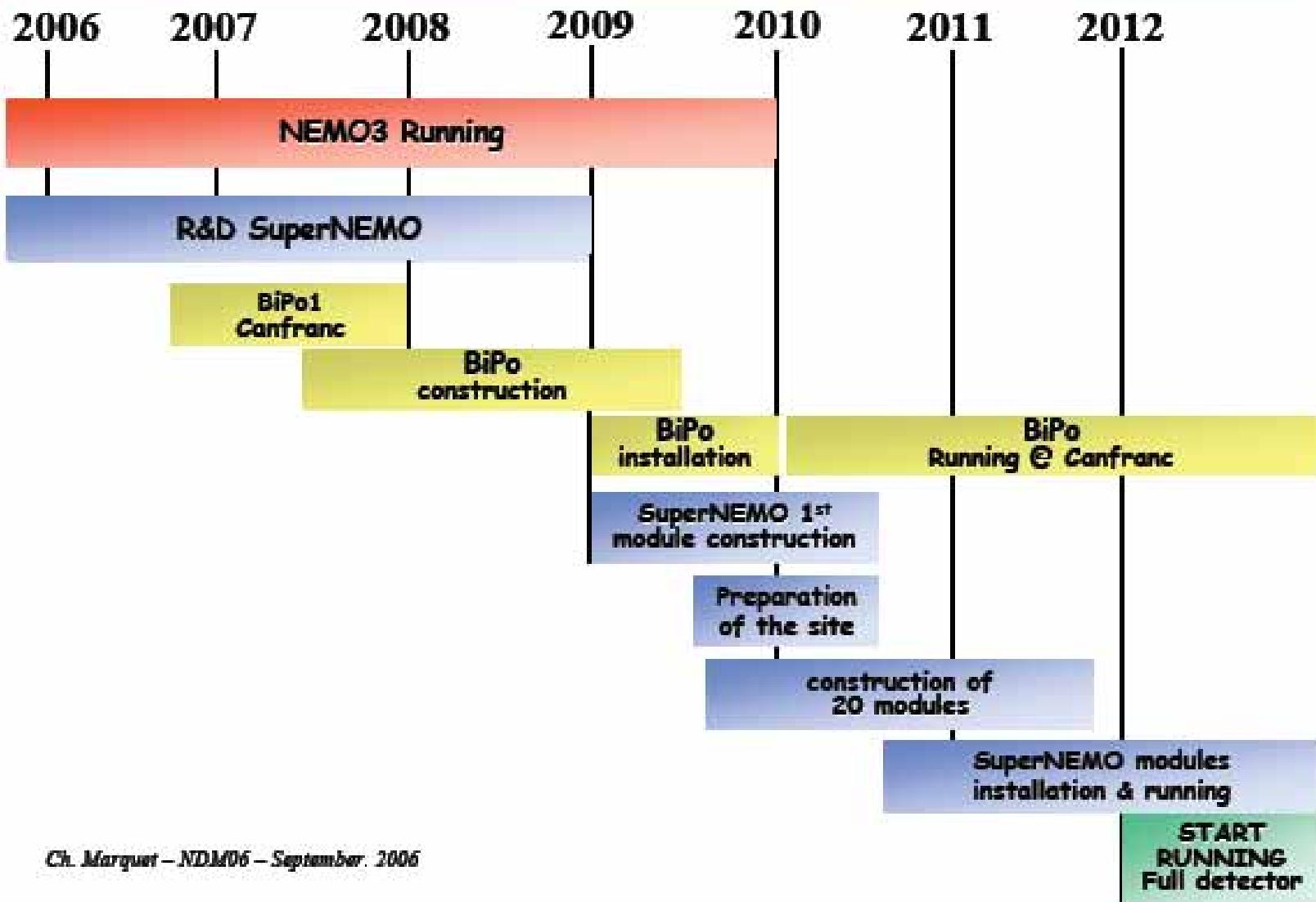
~2000kg natural U evaporated
~400 assay measurements

Restart MENPHIS... for ^{150}Nd ?

- Technically possible: Nd has been already enriched at 60% ^{150}Nd
- Reasonable production raise: few weeks for 100 kg
- Several interested experiments: SuperNEMO, SNO+, MOON, DCBA (*Letter of Interest – July 2006*)

Ch. Marquet – NDM06 – September 2006

Schedule summary



Ch. Marquet - NDM06 - September, 2006

MOON Detector

- **Multi layers module**

- **^{100}Mo foil & Plastic scintillator**

Mo foil is interleaved with PLs.

- **Compact 1 t detector**

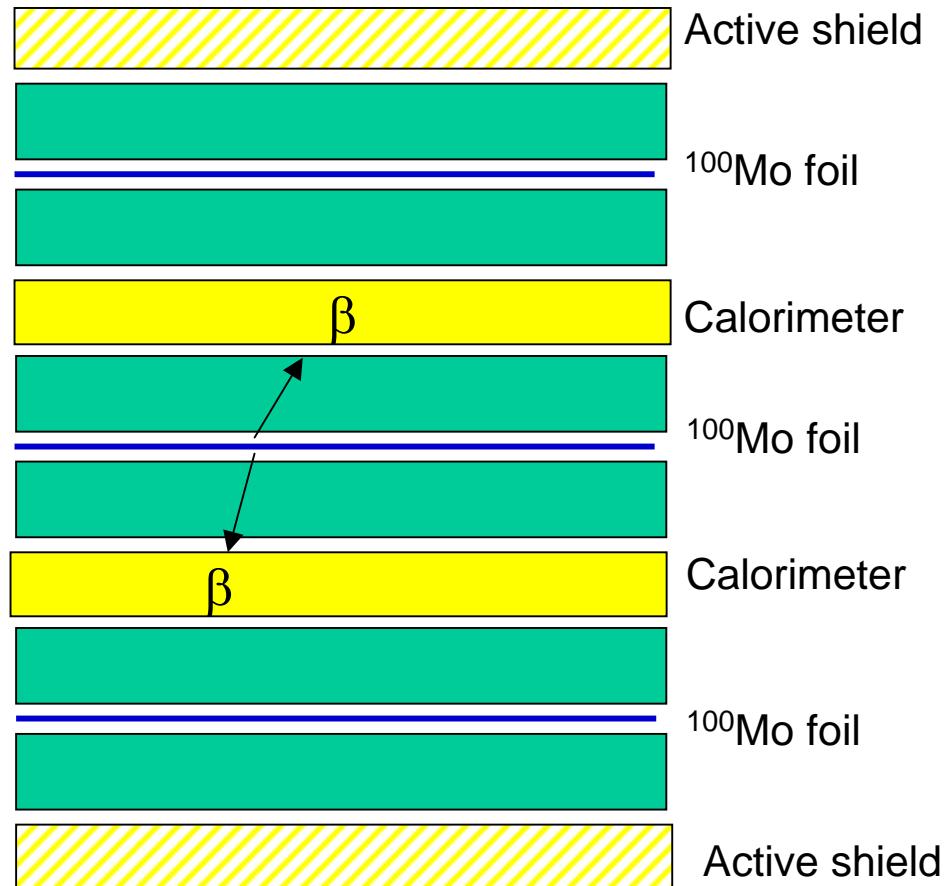
PL works both as calorimeter and as active shield

- **No TOF**

- **Effective gamma veto**

- **Particle ID.**

- Not in MOON-1 prototype detector



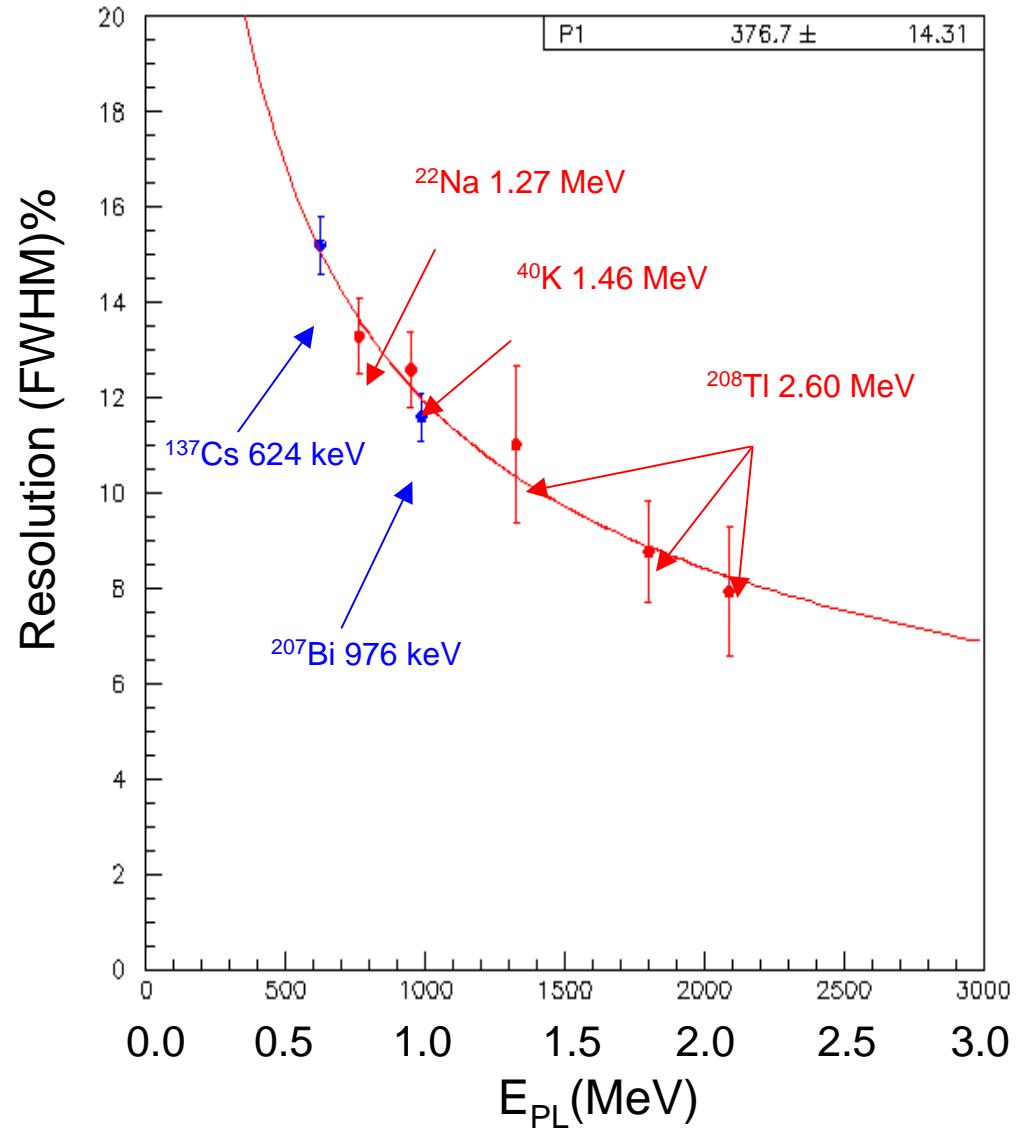
M. Nomachi

MOON-1 Energy resolution

$$R(FWHM) = \frac{(11.9 \pm 0.5)}{\sqrt{E(MeV)}} \%$$

$$R(FWHM) @ 3\text{MeV} = 6.8\%$$

M. Nomachi



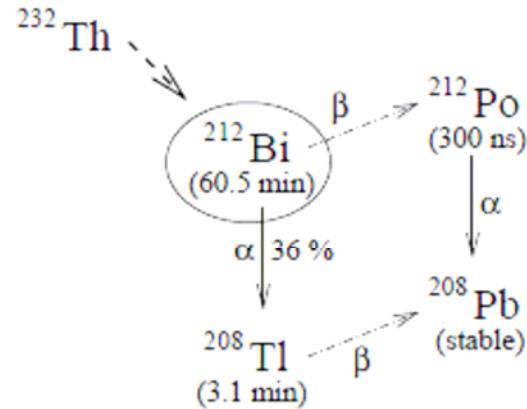
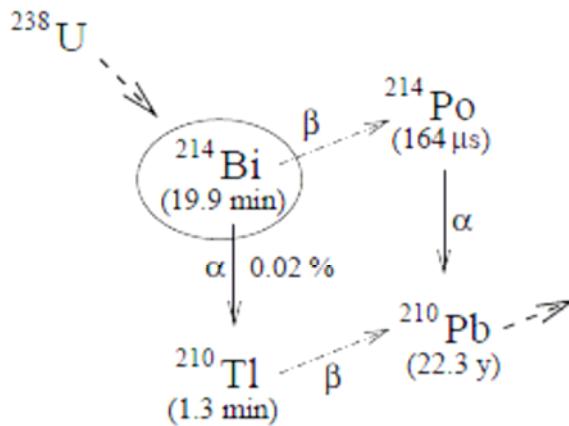
Ultra Low Background Detector

To achieve required purity, the detector with high sensitivity
BiPo-detector (LAL,CENBG,Osaka,Tokushima)

With 5 kg of ^{82}Se source foil ($\sim 12 \text{ m}^2$, 40 mg/cm^2)

$2 \mu\text{Bq/kg}$ of ^{208}Tl \rightarrow 50 (e^- , delay α) ^{212}Bi decays / month
 $\varepsilon \sim 6 \%$ \rightarrow 3 decays / month

Background < 1 events/month is required !



BGD measurement with Radon monitor

– R&D works at Saga Univ. (H. Ohsumi) –

Possible contributions to SuperNEMO

- Radon detectors, (70 litters, Sensitivity 1mBq/m^3)

for Bi-Po experiment (@Canfranc)

- Study of low background materials used for SuperNEMO

measurement of radon/thoron emanation

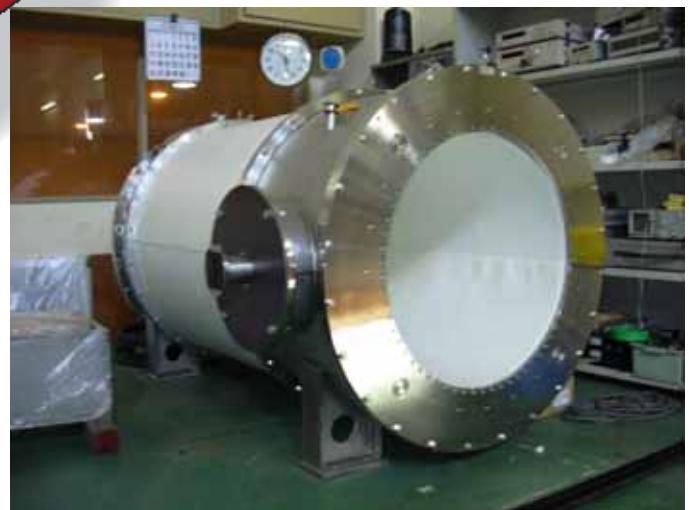
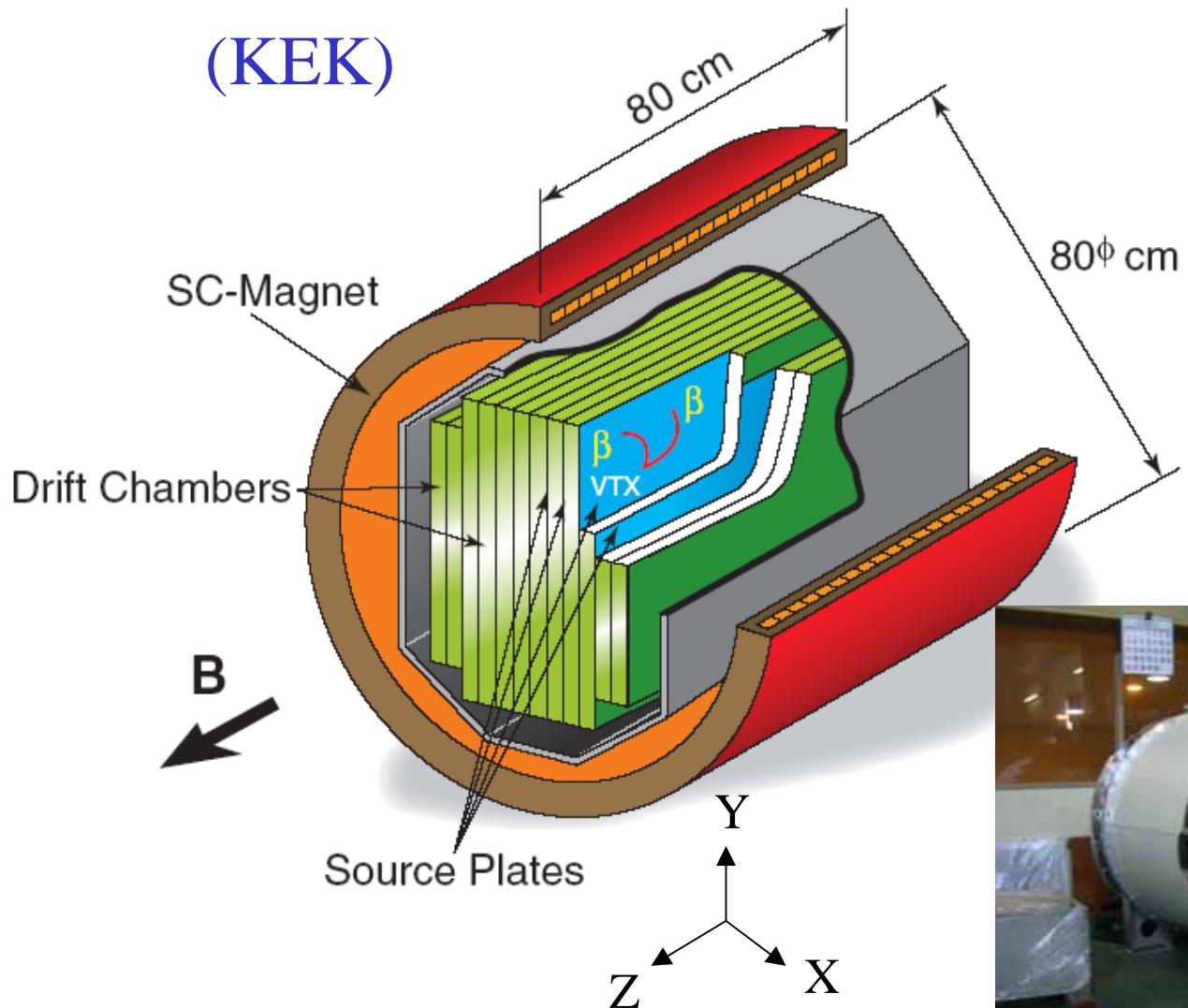
collaboration with CENBG Bordeaux.

- Next generation high sensitive radon detector (Next goal -- 0.1mBq/m^3)

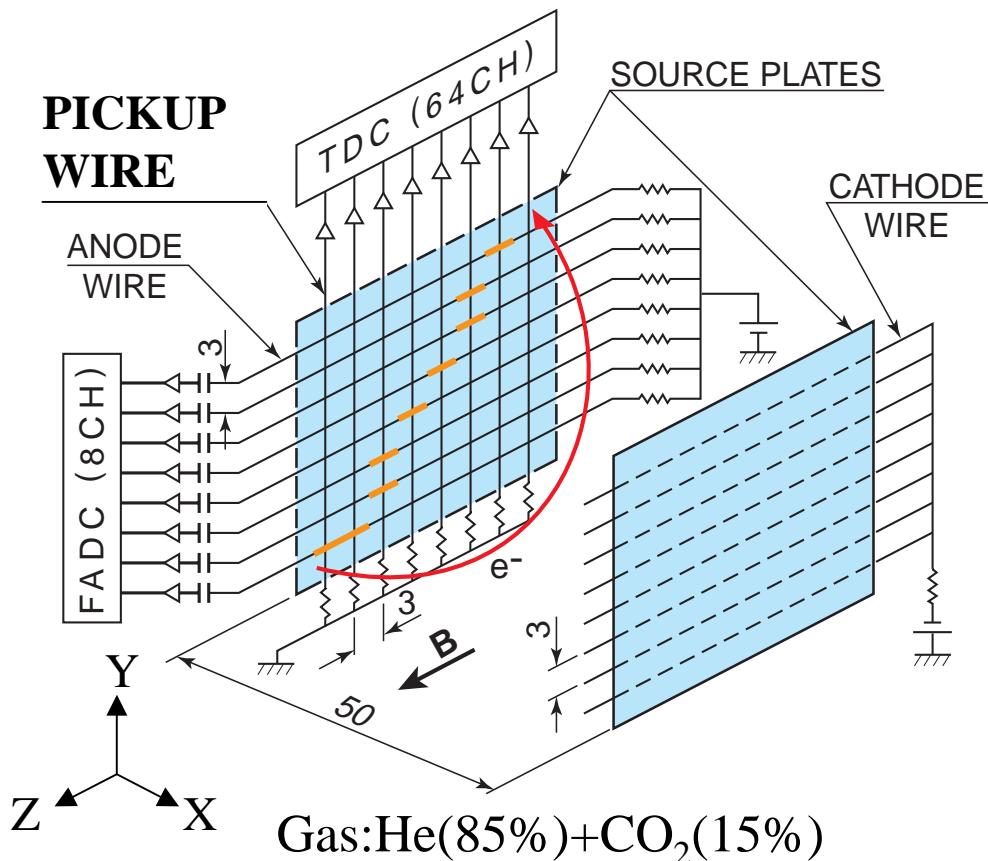
Bigger size volume, or new technique, using low background photodiode etc.

DCBA-T3 (Drift Chamber Beta-ray Analyzer)

(KEK)



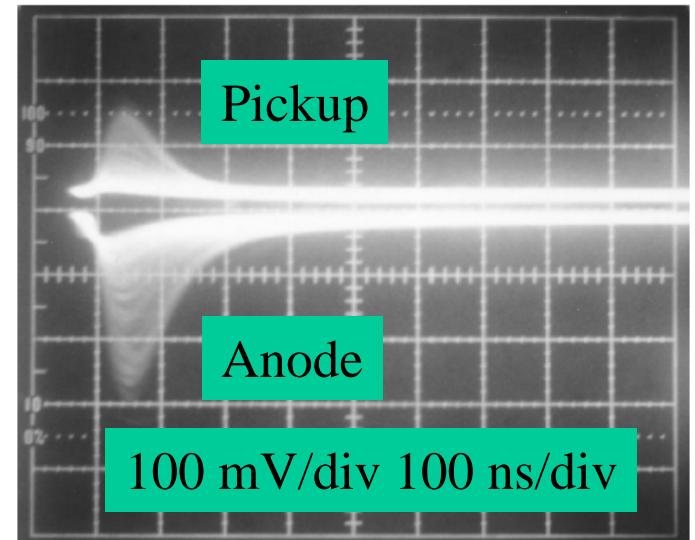
Principle of DCBA



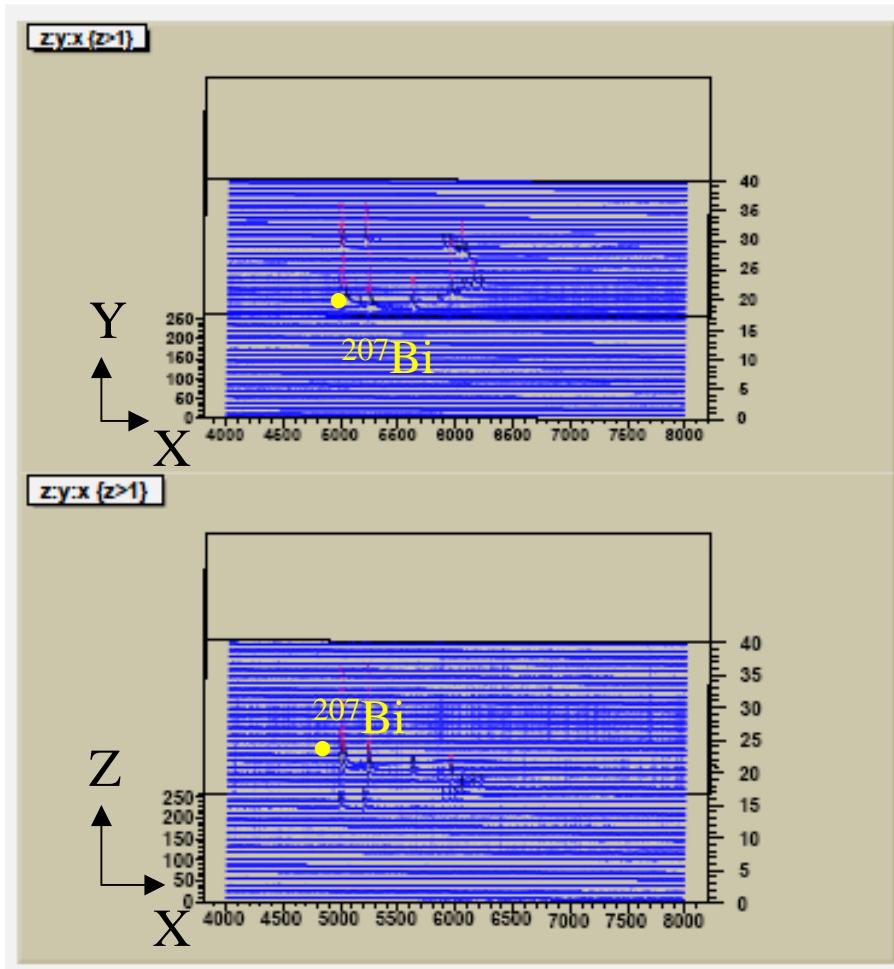
$$p \cos \lambda = 0.3rB,$$

$$T = (p^2 + m_e^2)^{1/2} - m_e$$

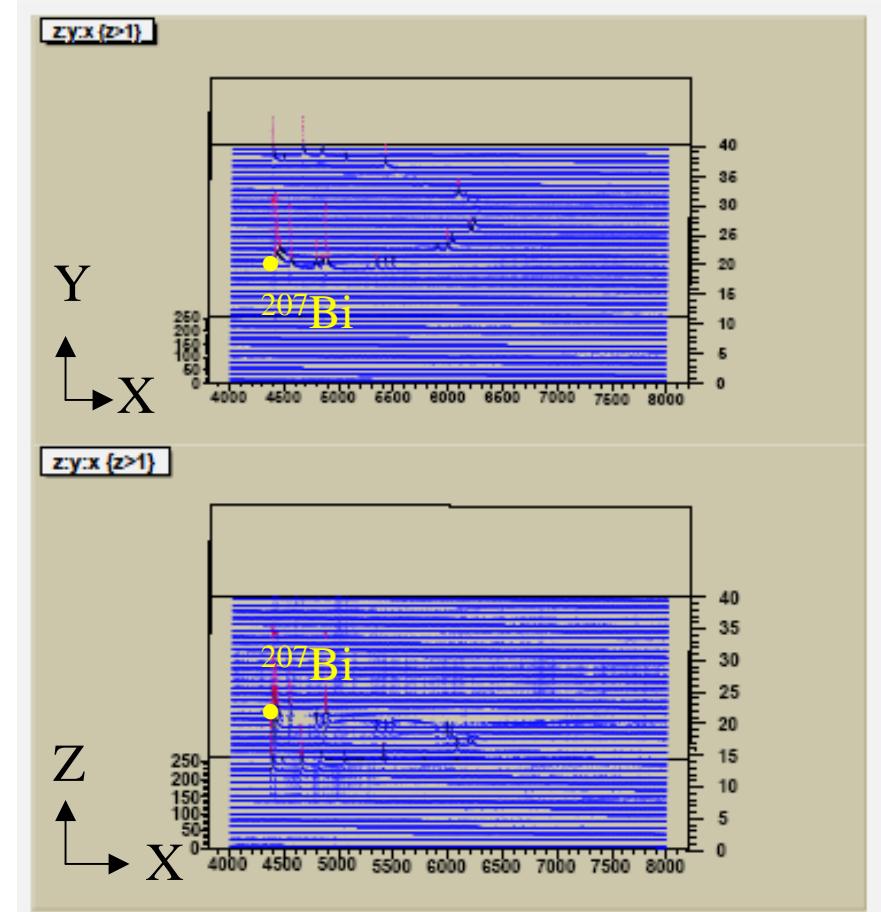
p (MeV/c): momentum, r (cm): radius,
 λ : pitch angle, B (kG): magnetic field,
 m_e (MeV/c²): electron mass



Electrons from ^{207}Bi



$T = 458 \text{ keV}$



$T = 971 \text{ keV}$

energy total

Energy Resolution of DCBA-T2

energy total

Entries 281

Mean 0.7967

RMS 0.2316

count

Including BGD

Preliminary

45
40
35
30
25
20
15
10
5
0

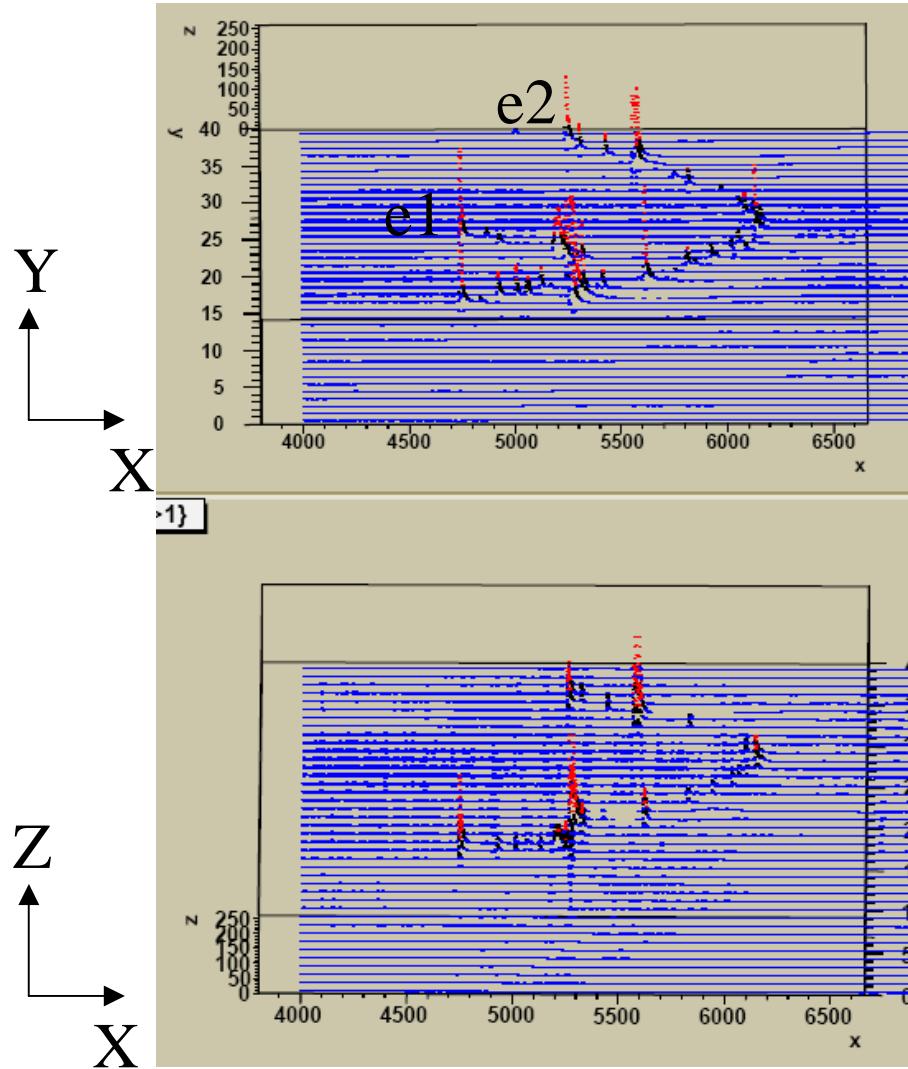
0 0.2 0.4 0.6 0.8 1 1.2 1.4 1.6 1.8 2

0.48
(1.5%) 0.56
(0.6%) 0.98
(7.0%) 1.05
(2.4%)FWHM ≈ 0.15 MeV

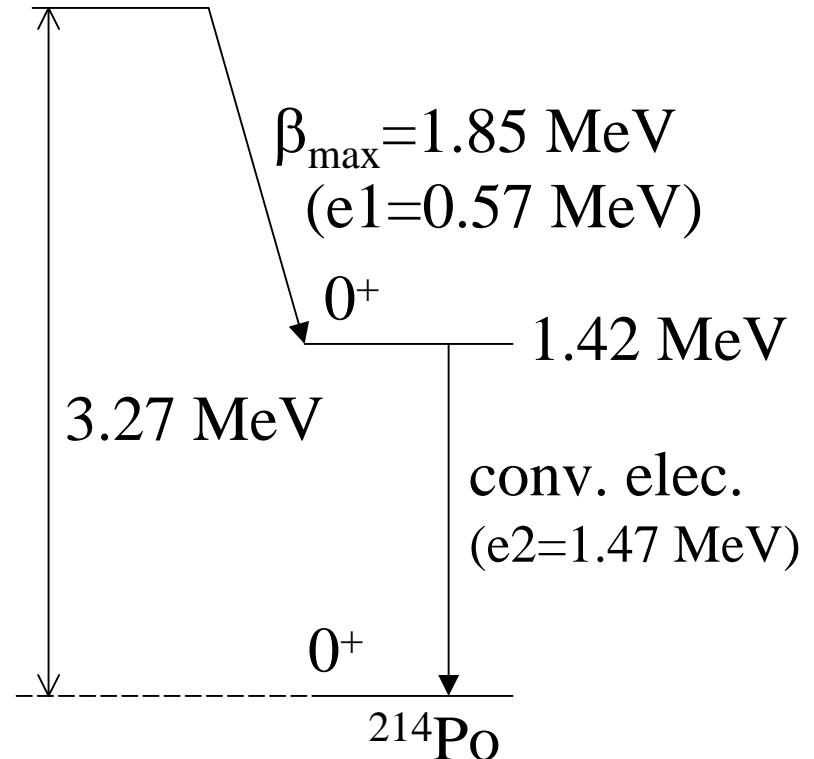
@ 0.98 MeV

FWHM(E_{sum}) ≈ 0.21 MeV
$$\frac{\text{FWHM}(\text{E}_{\text{sum}})}{Q} \approx 6.2\%$$
Aiming at 4%
in DCBA-T3

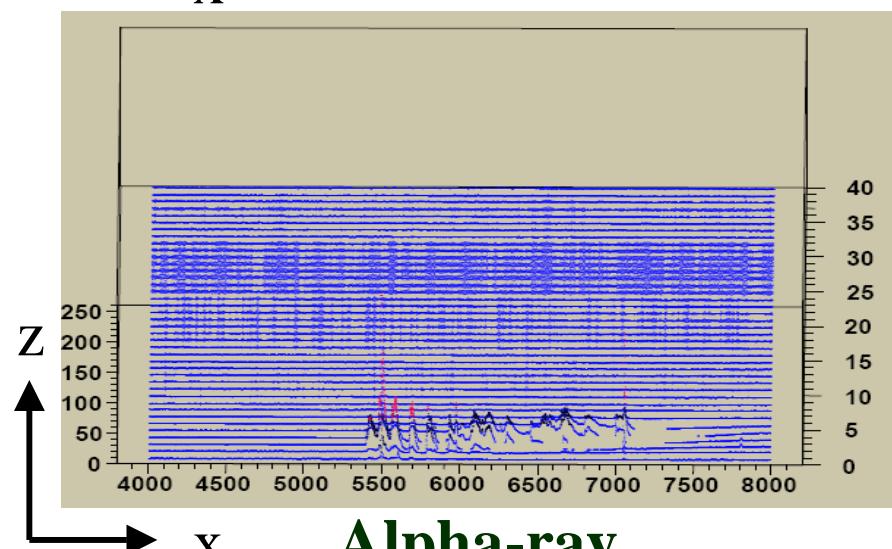
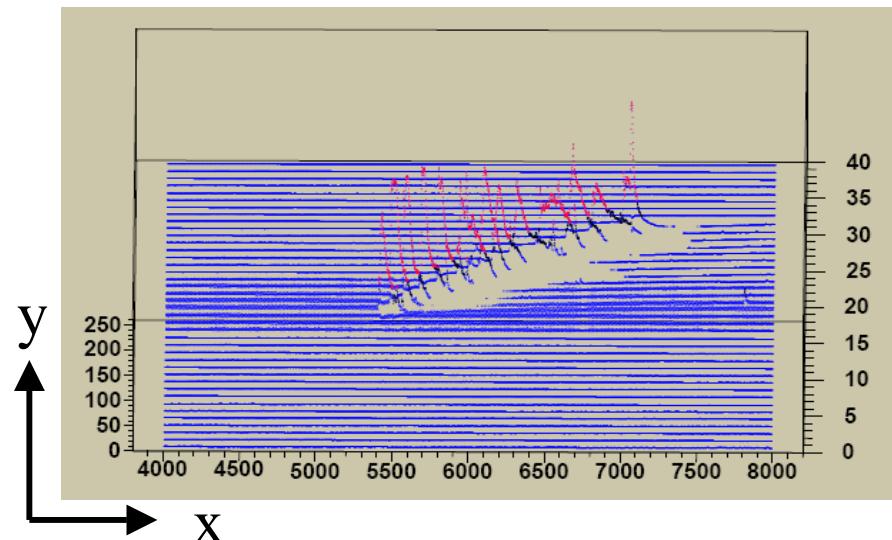
Background 2-electron event



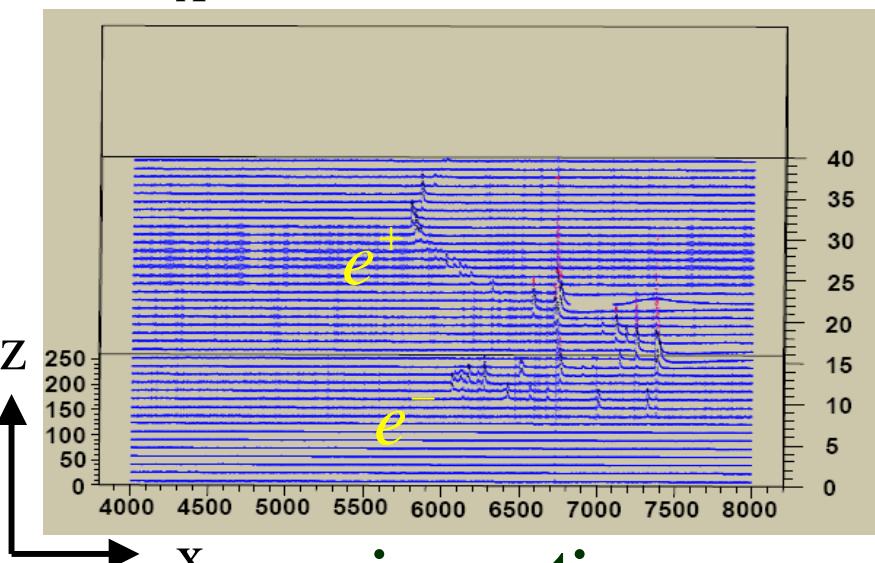
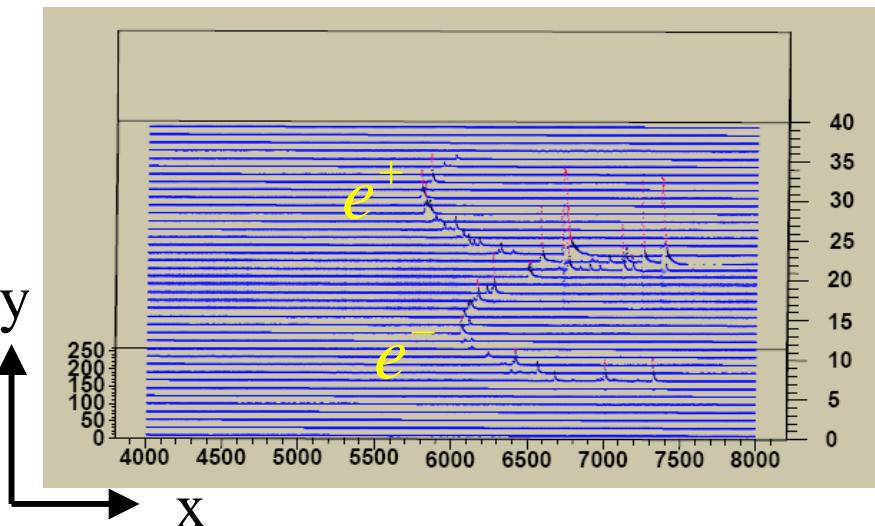
^{214}Bi (Uranium decay series)



OTHER BACKGROUND EVENTS



Alpha-ray



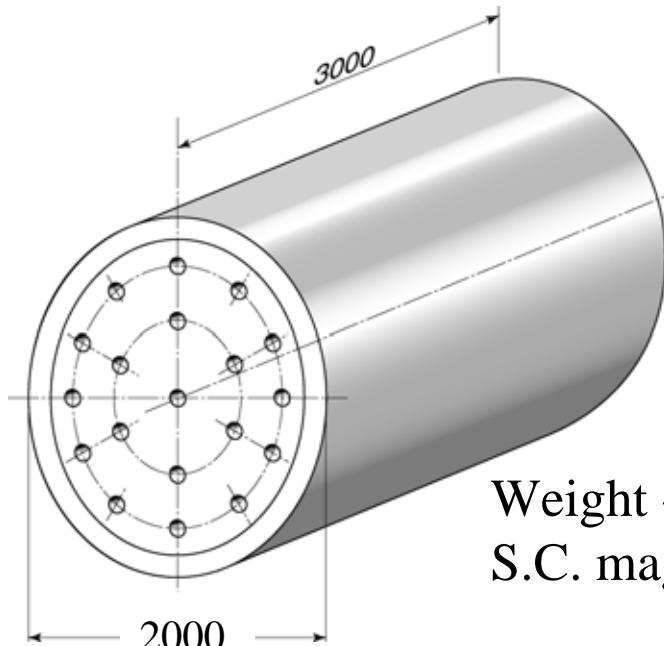
pair creation

Future plan MTD (Magnetic Tracking Detector) Temporary

Source plate: 84 m²/module

Thickness: 15 (40) mg/cm²

Weight: 12.6 (33.6) kg/module



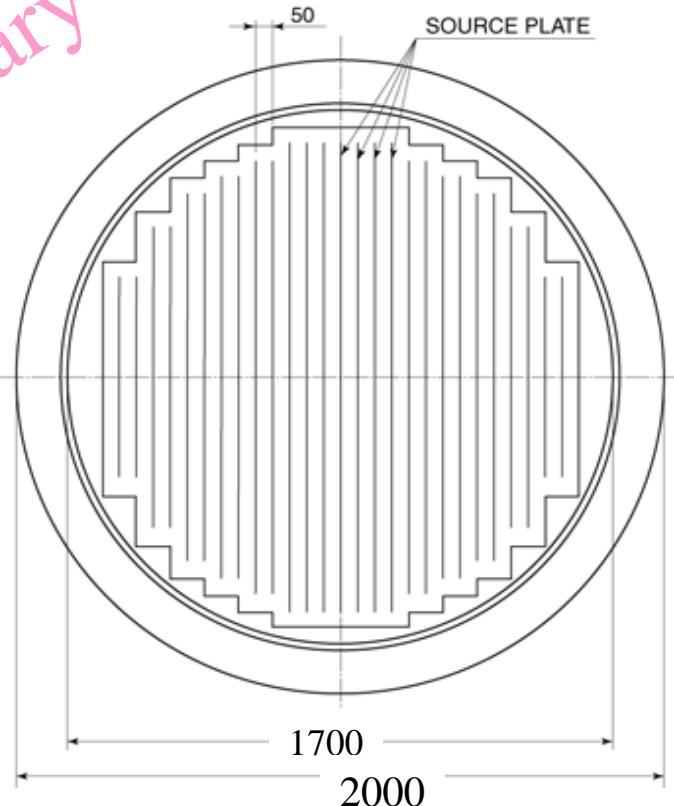
Weight ~ 20 t
S.C. magnet

$\langle m_\nu \rangle \approx 0.8 \text{ eV}$ (0.5 eV) for natural Nd/mod.yr

$\langle m_\nu \rangle \approx 0.2 \text{ eV}$ (0.1 eV) for 60% ^{150}Nd /mod.yr

Anode wire: 10600/module
Pickup wire: 26200/module

Preliminary

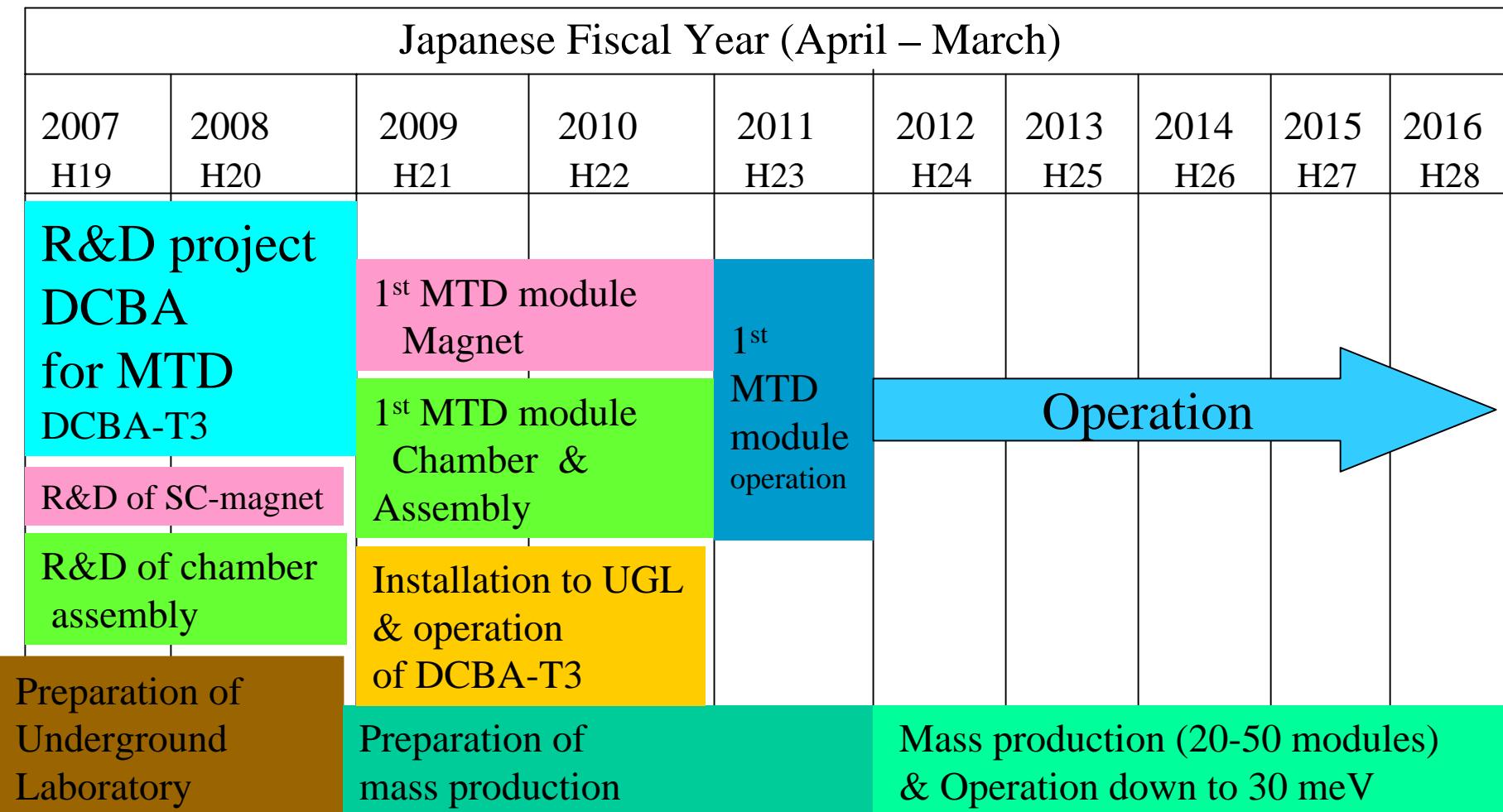


Half-life and Effective Mass Sensitivities of MTD for ^{150}Nd , ^{100}Mo and ^{82}Se (Tentative)

	Natural Nd (5.6% ^{150}Nd)	^{150}Nd (80% enr.)	^{100}Mo (90% enr.)	^{82}Se (90% enr.)
MTD Amount (mol)	190	2700	5400	6600
(600 kg : 50 modules of 15 mg/cm ²)				
$T_{0\nu}^{1/2}$ sens. (yr)	9×10^{24}	1×10^{26}	2×10^{26}	3×10^{26}
$\langle m_\nu \rangle$ sens. (eV)	0.06	0.02	0.07	0.04

Nucl. Matrix Element: A. Staudt et al. Europhys. Lett. 13 (1) (1990) 31

Optimistic Schedule of DCBA/MTD (depending on financial support)



Concluding remarks

1. Observation of $0\nu\beta\beta$ events in ^{76}Ge claimed by KKDC should be confirmed by other experiments using different decay sources.
2. NEMO, MOON and DCBA can use several kinds of decay sources, such as ^{150}Nd , ^{100}Mo and ^{82}Se . They have tracking detectors which are very useful to reduce backgrounds.
3. These groups have made efforts at obtaining enriched ^{150}Nd , improving energy resolution and reducing backgrounds, in order to investigate Majorana nature and the effective mass down to 30 meV, which is predicted in the inverted hierarchy spectrum.
4. Collaboration will be made by exchange of experimental physicists between different test benches in Japanese and French Laboratories.