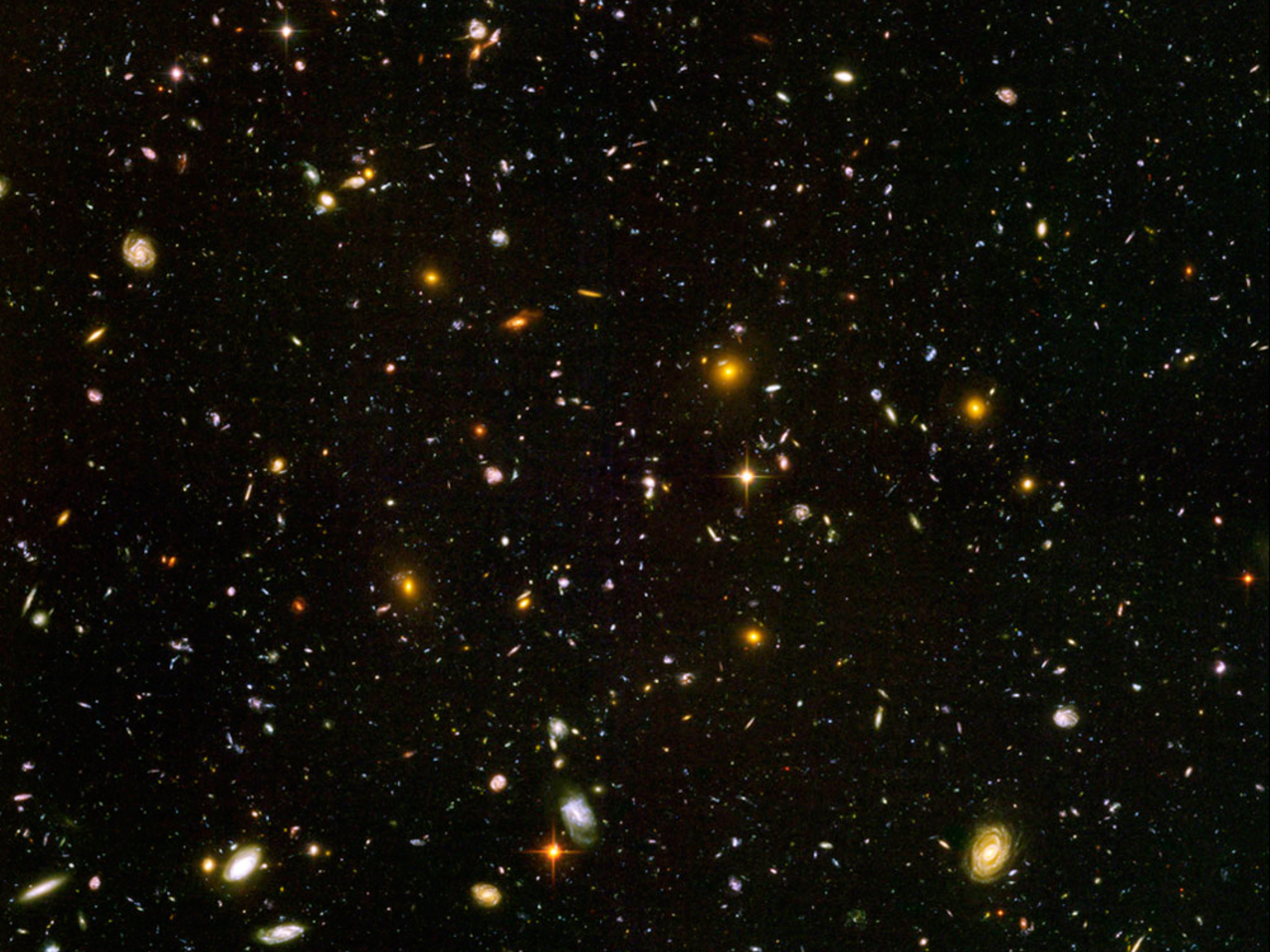




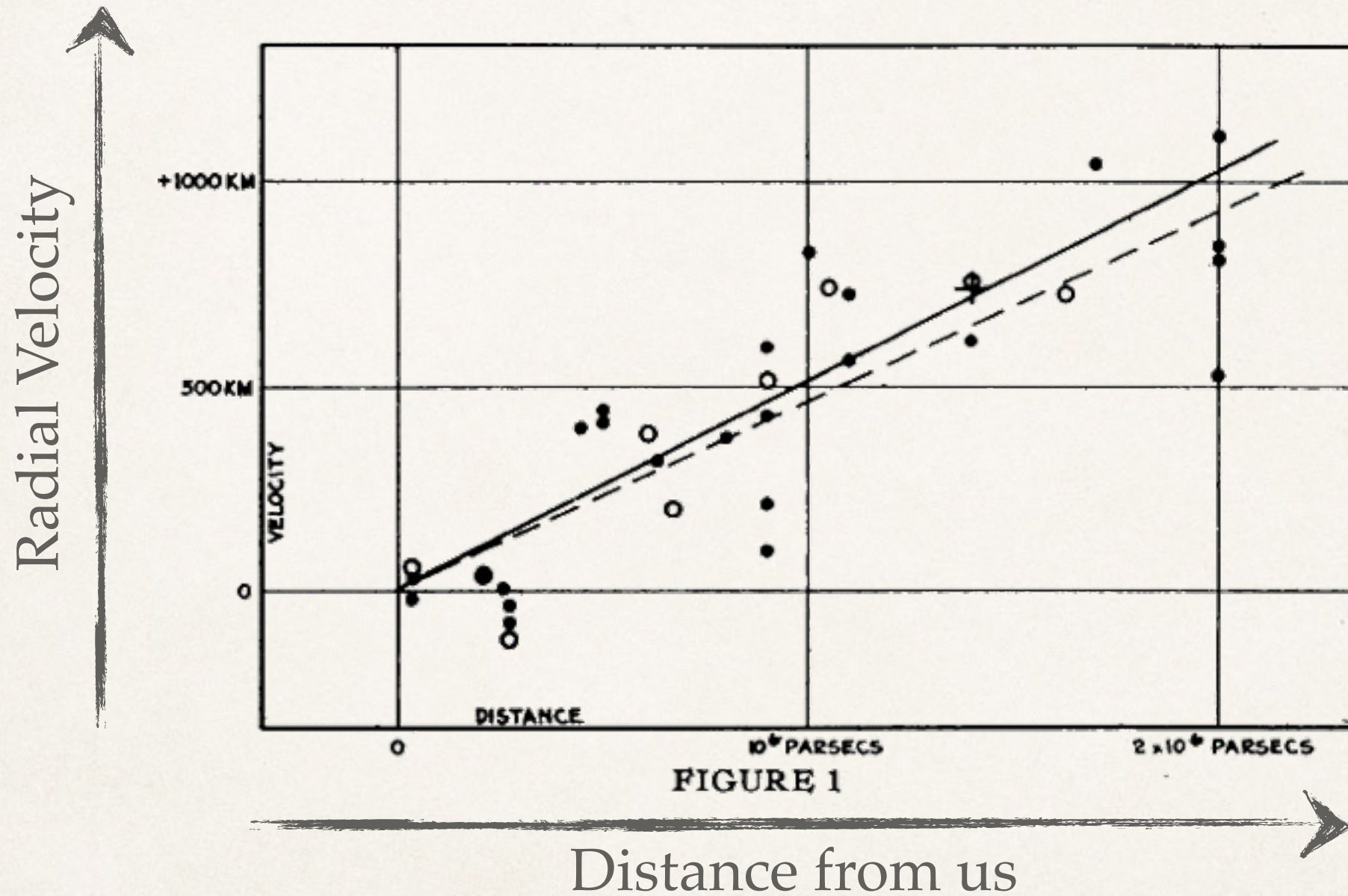
Modern Cosmology: Astrophysical analyses of SNeIa and future surveys

IPHC – Mickael RIGAULT

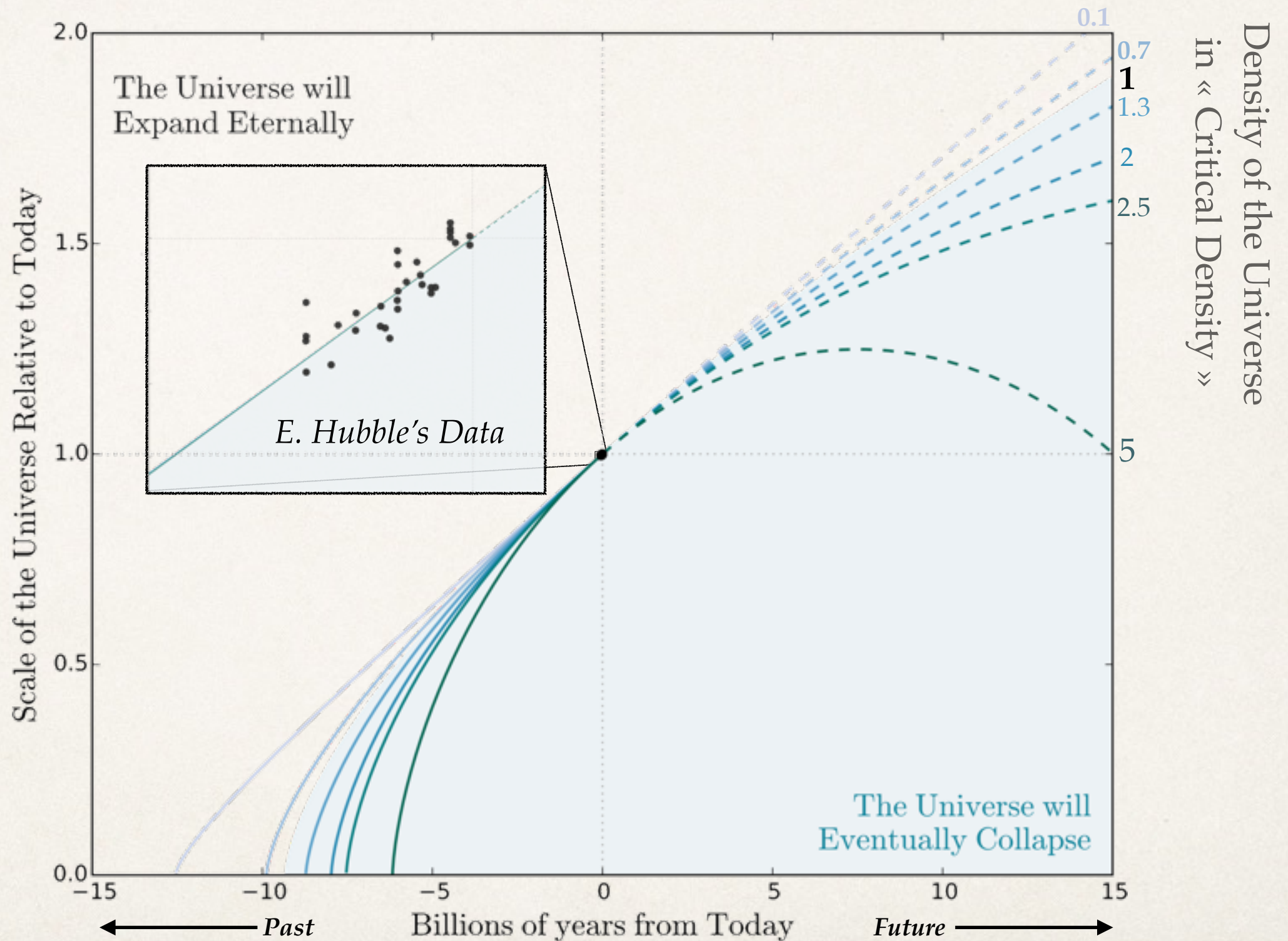


The Universe is Expanding!

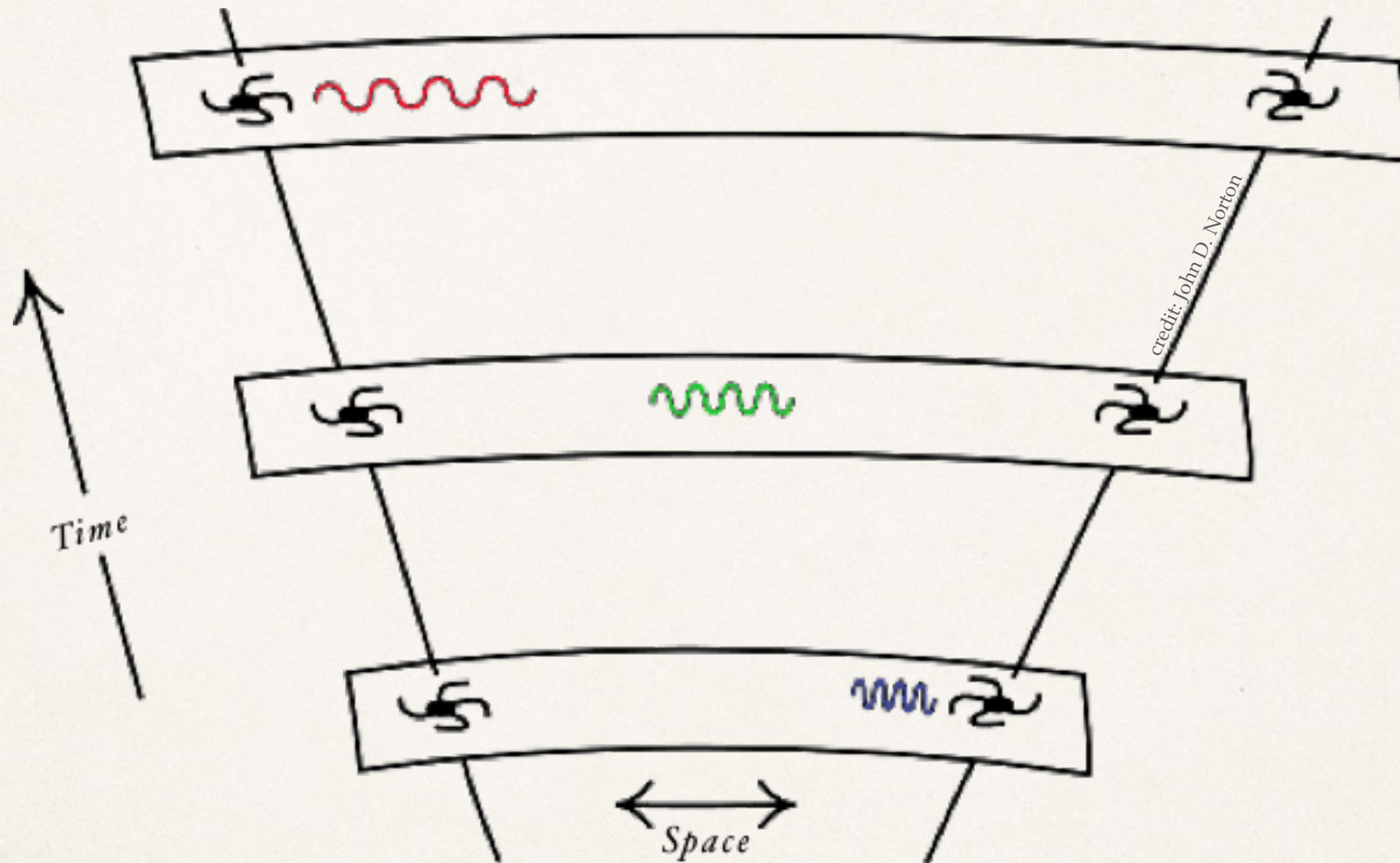
Edwin HUBBLE, 1929



Measuring the Fate of the Universe



The Redshift as an Expansion Tracer



The expansion of the Universe stretches the photon's wavelength

Type Ia Supernovae: Standard Candles

Flux \Leftrightarrow Distance

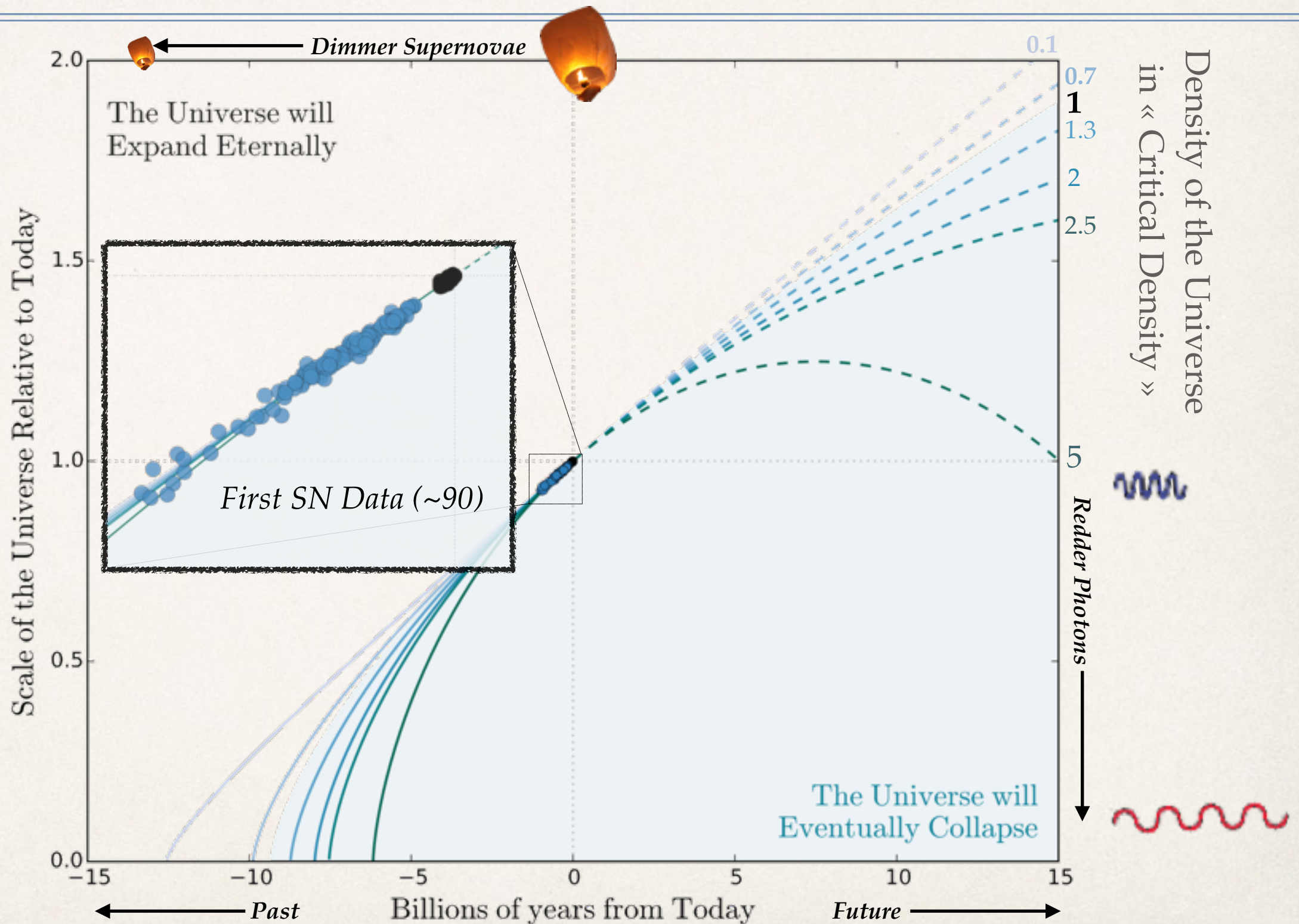


As bright as a galaxy for few days

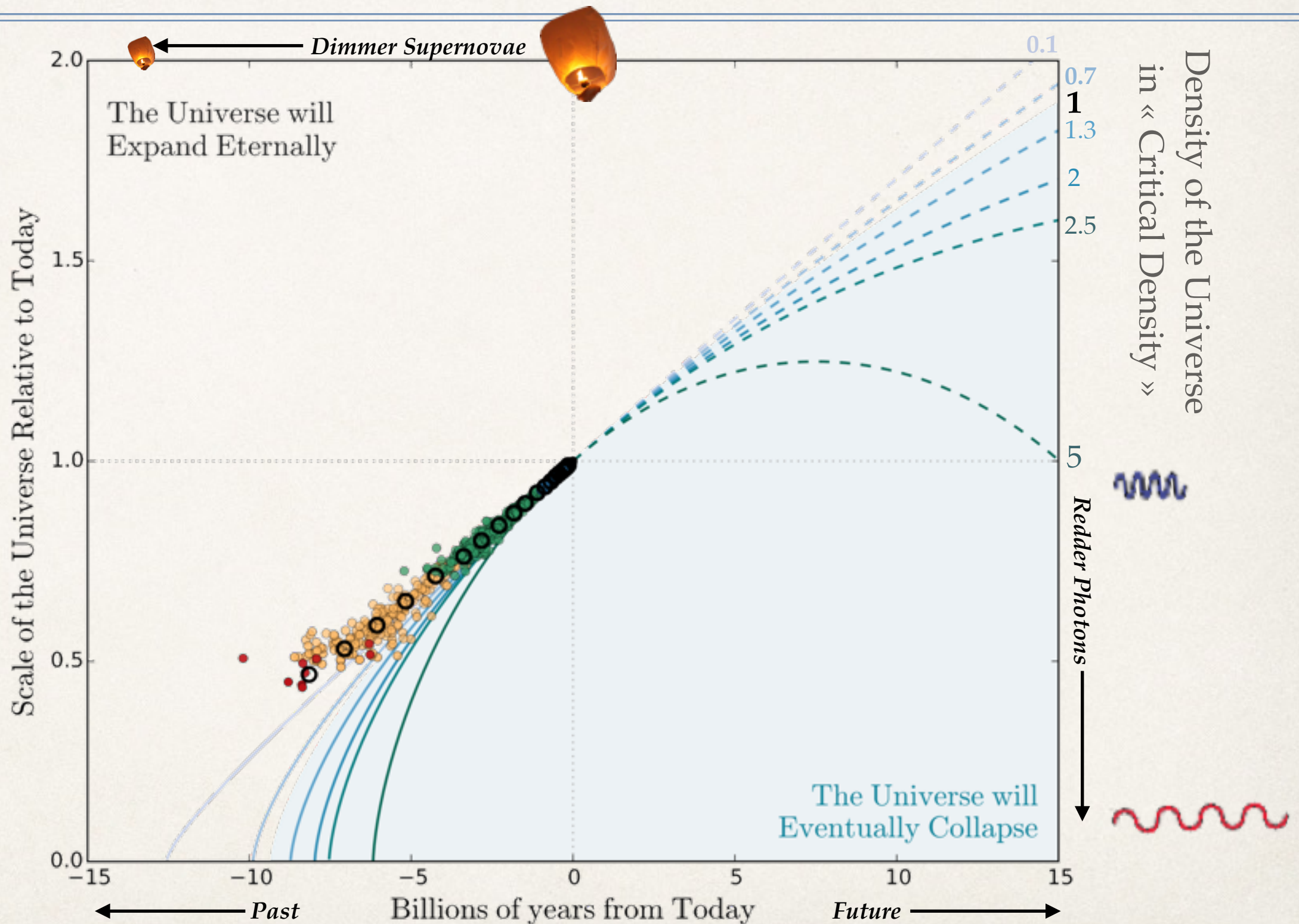


Artist's Concept

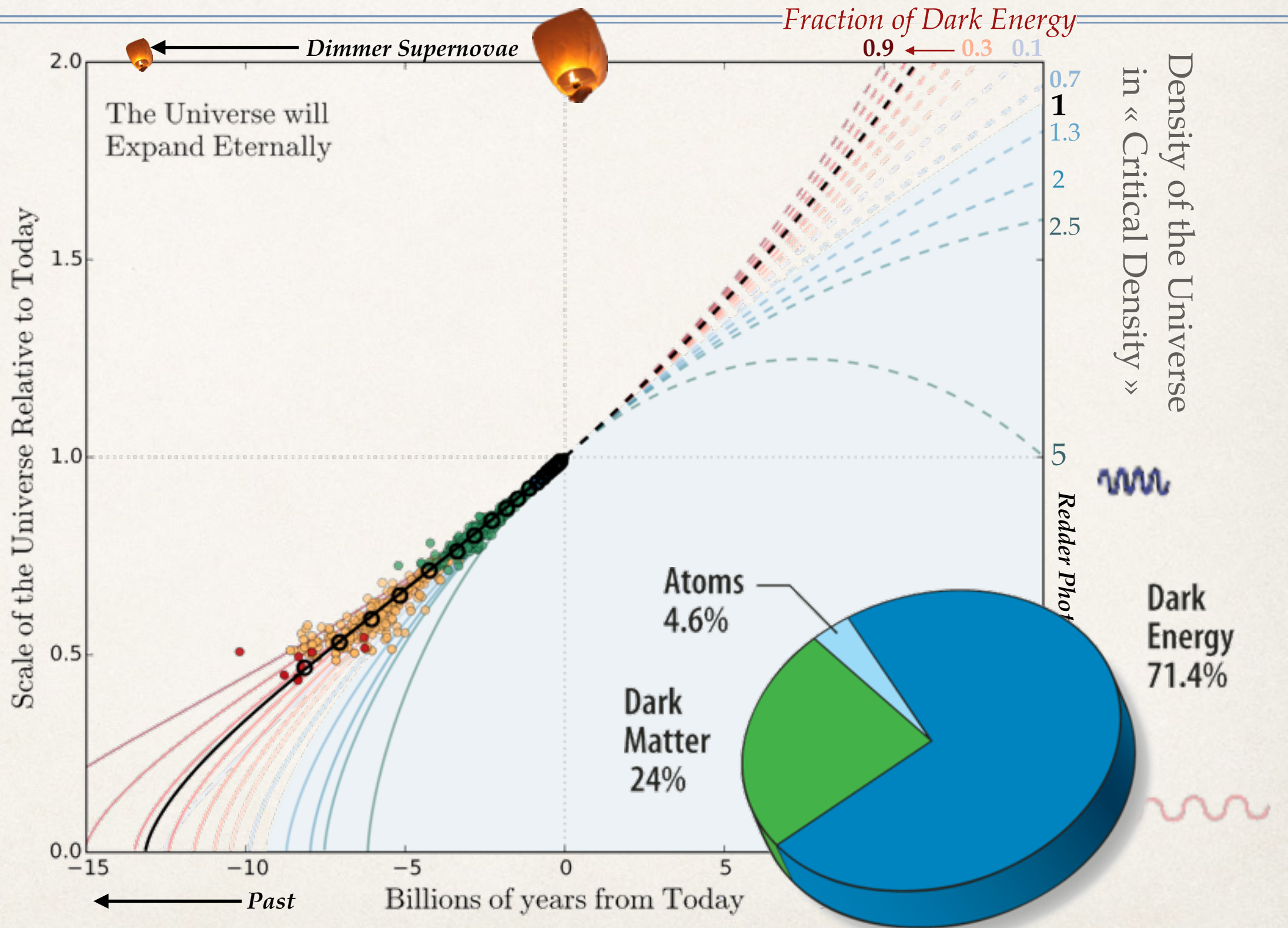
Measuring the Fate of the Universe



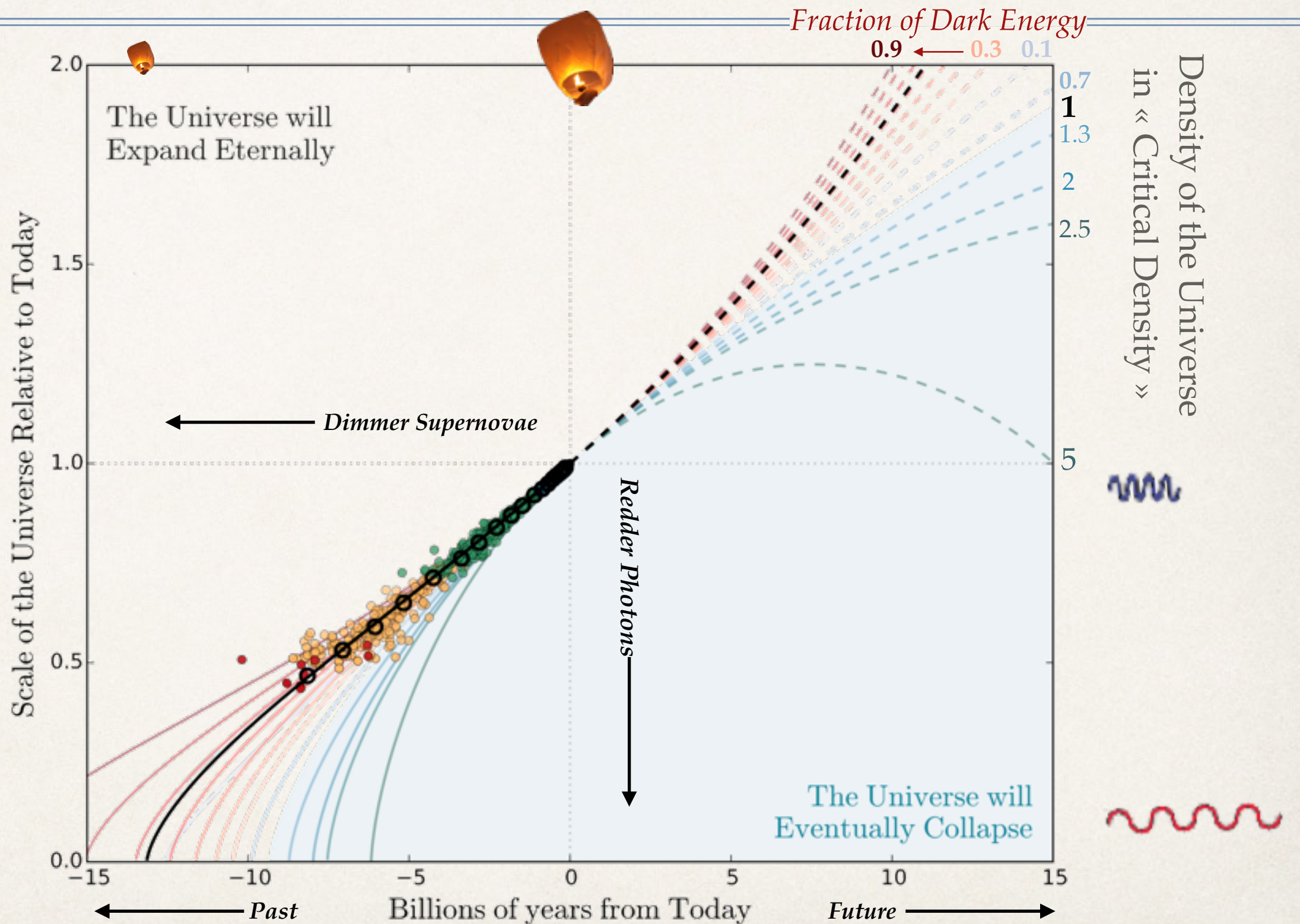
The Data do not match the Predictions



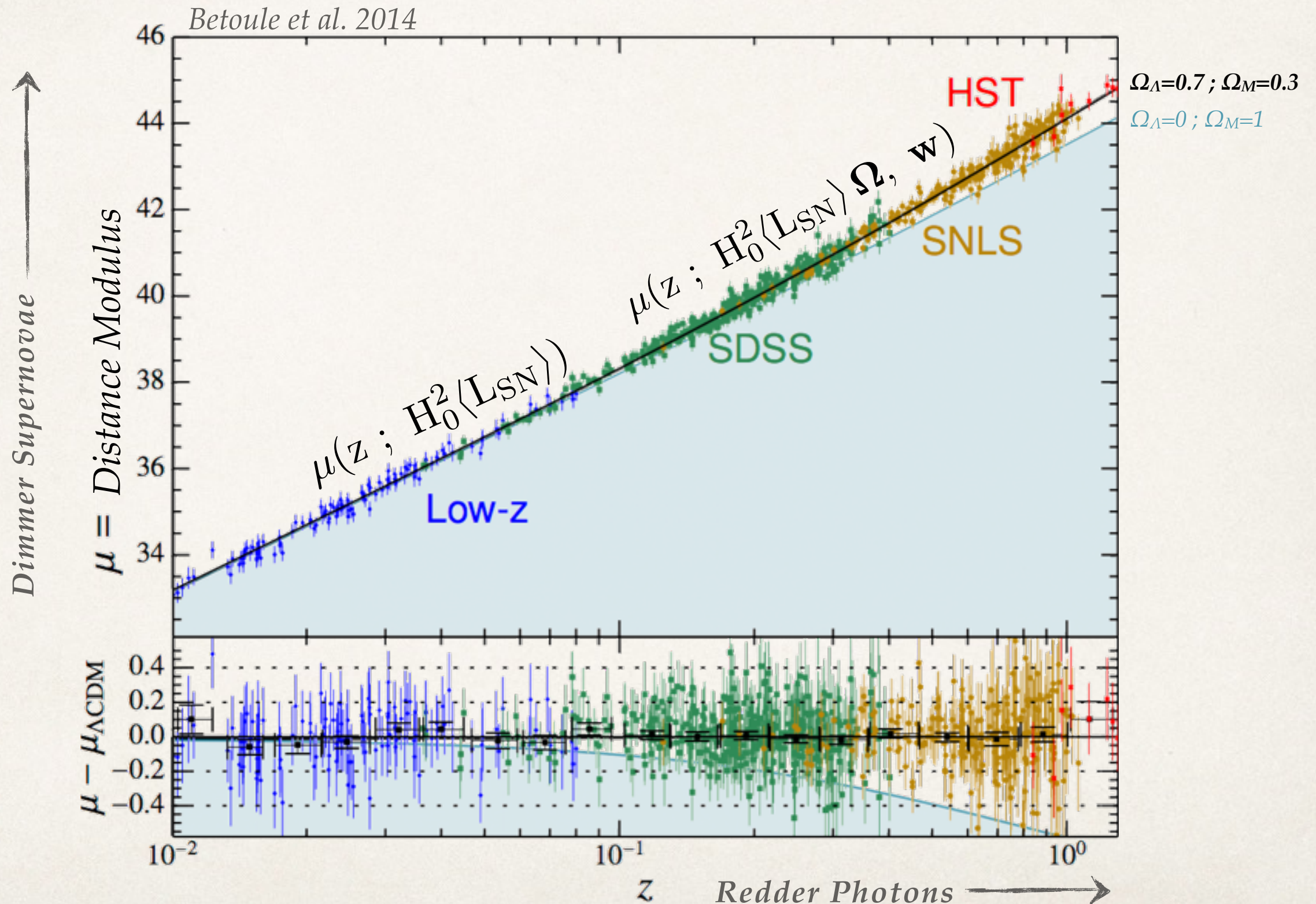
The Universe's Expansion is Accelerating!



Study the Acceleration

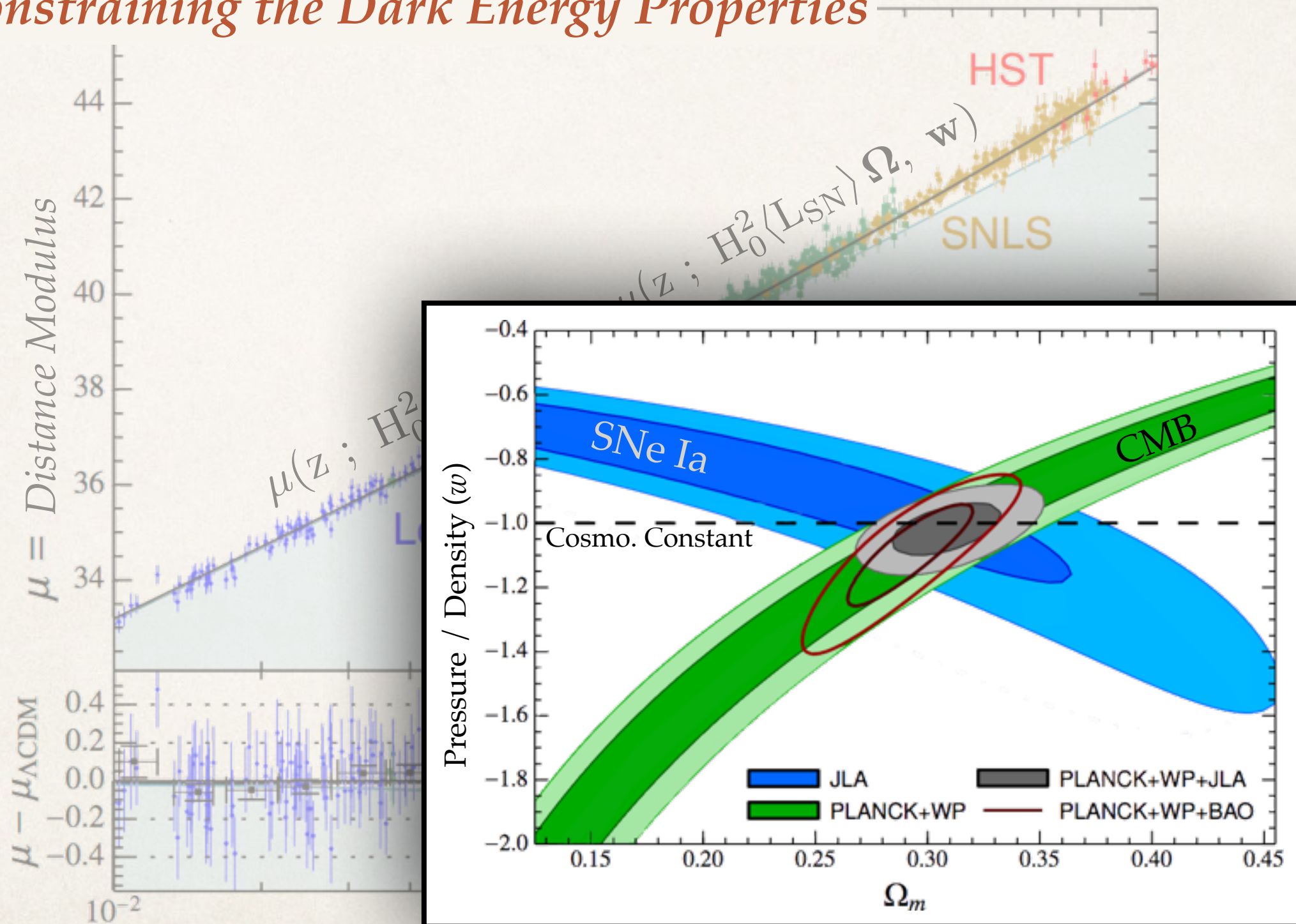


Cosmology using Type Ia Supernovae



Cosmology using Type Ia Supernovae

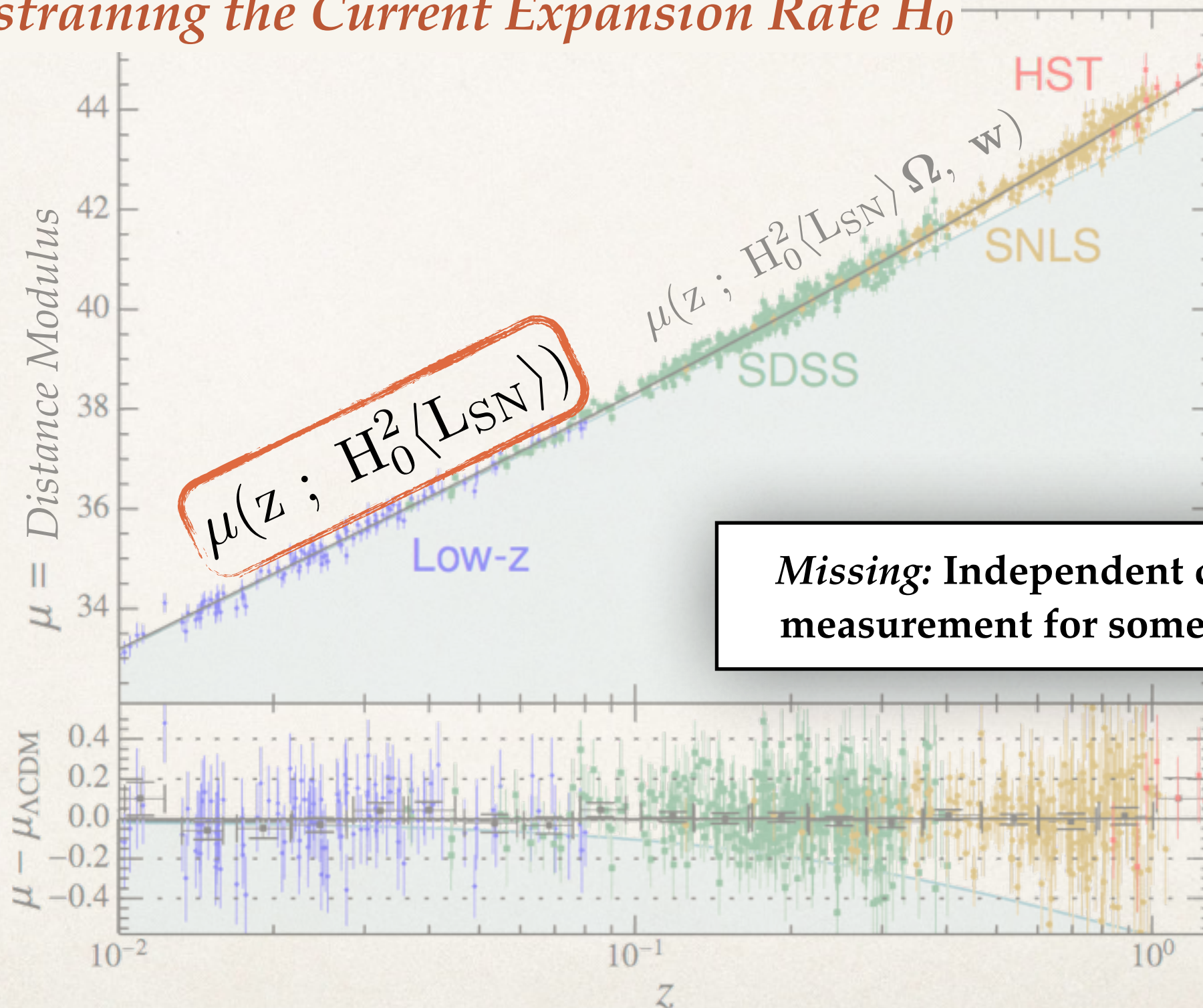
Constraining the Dark Energy Properties



Betoule et al. 2014

Cosmology using Type Ia Supernovae

Constraining the Current Expansion Rate H_0

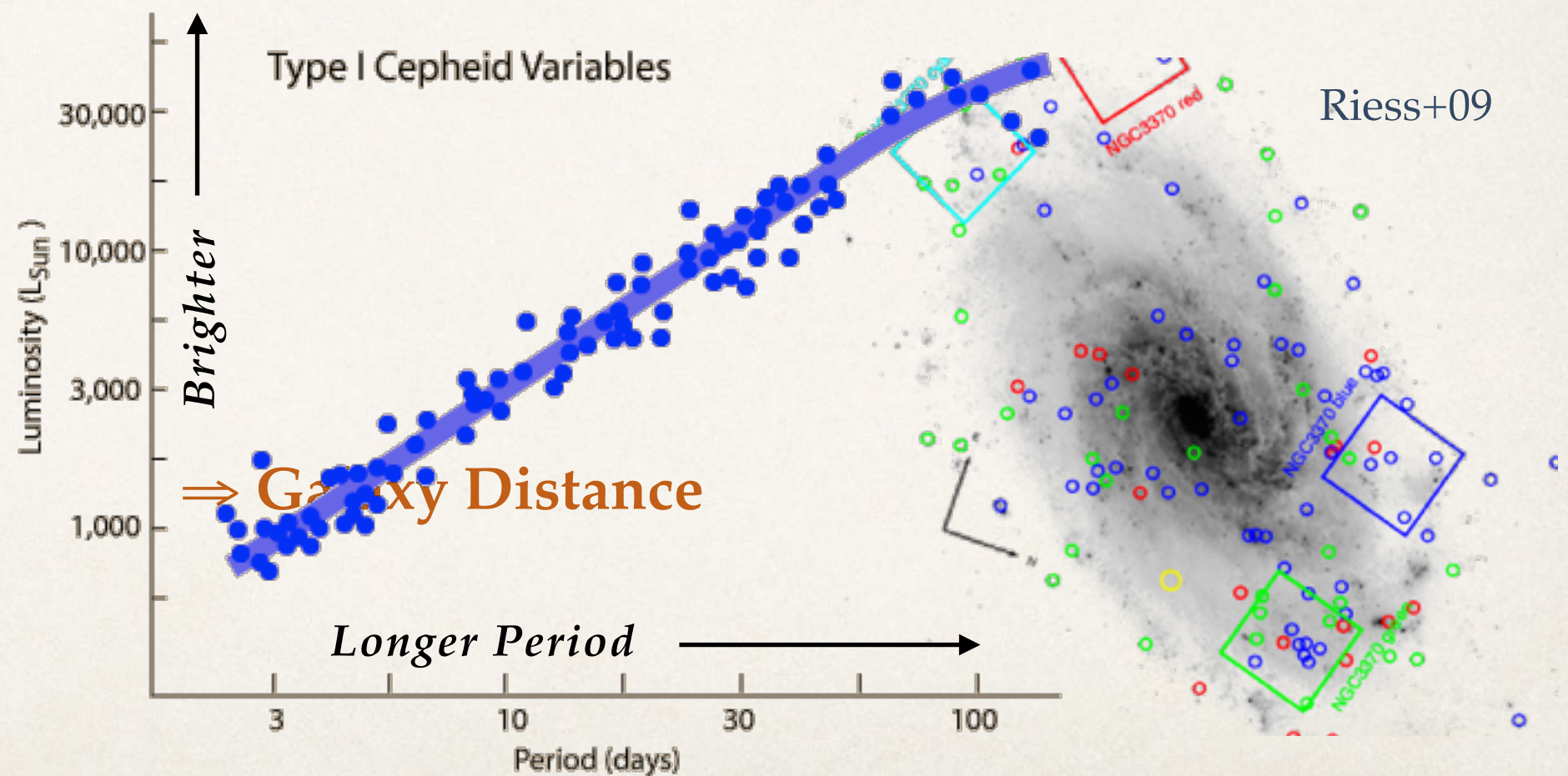


Disentangle H_0 and the SN luminosity

Independent distance measurement for some SNe Ia

Best case: Cepheids

Cepheids: *Bright Young Stars with a Pulsation—Luminosity relation*

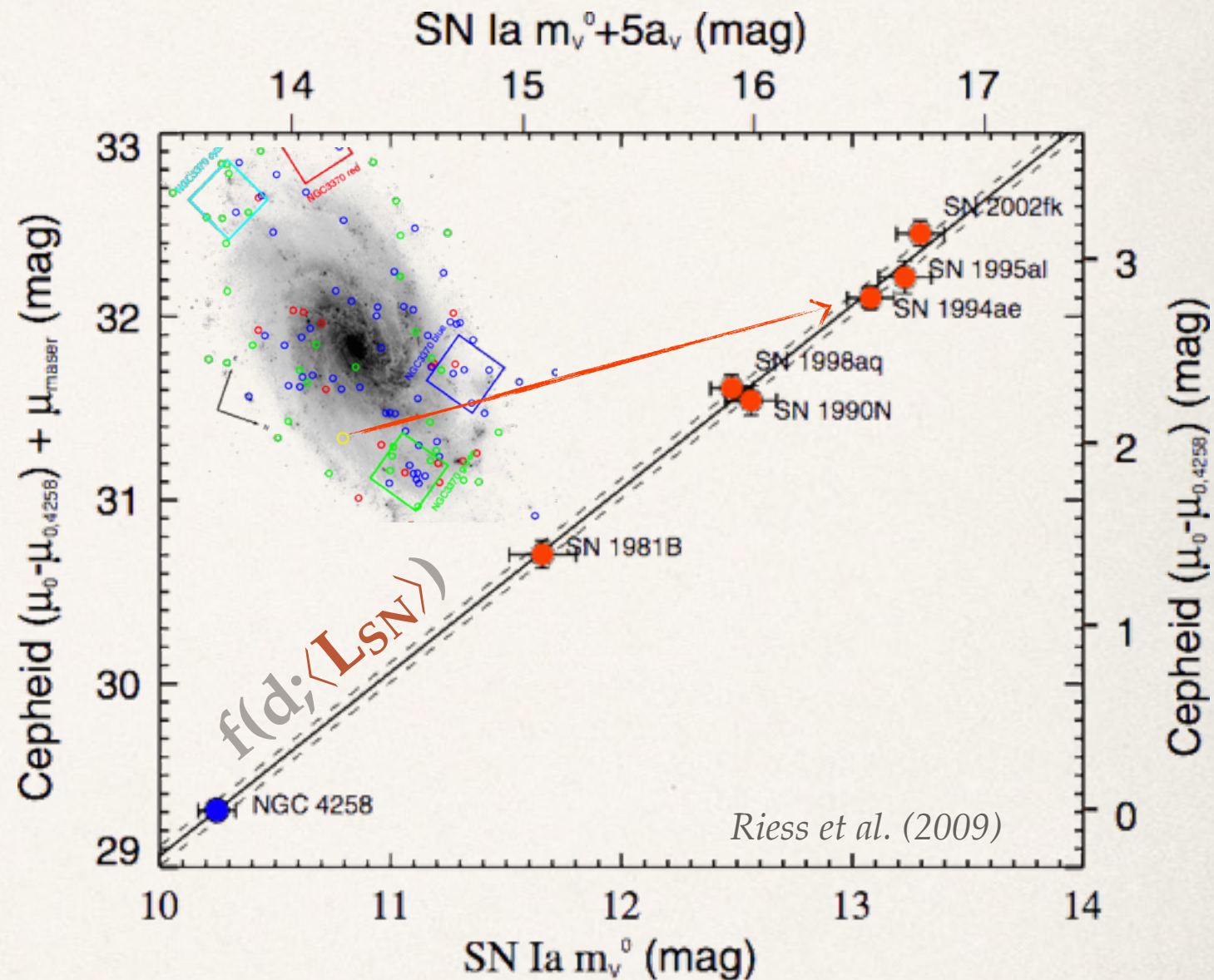


Disentangle H_0 and the SN luminosity

Independent Distance Measurement

— Distance & Flux \Rightarrow Luminosity —

Cepheids: Bright young stars with a pulsation-luminosity relation



$$H_0 = 73.2 \pm 1.8 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

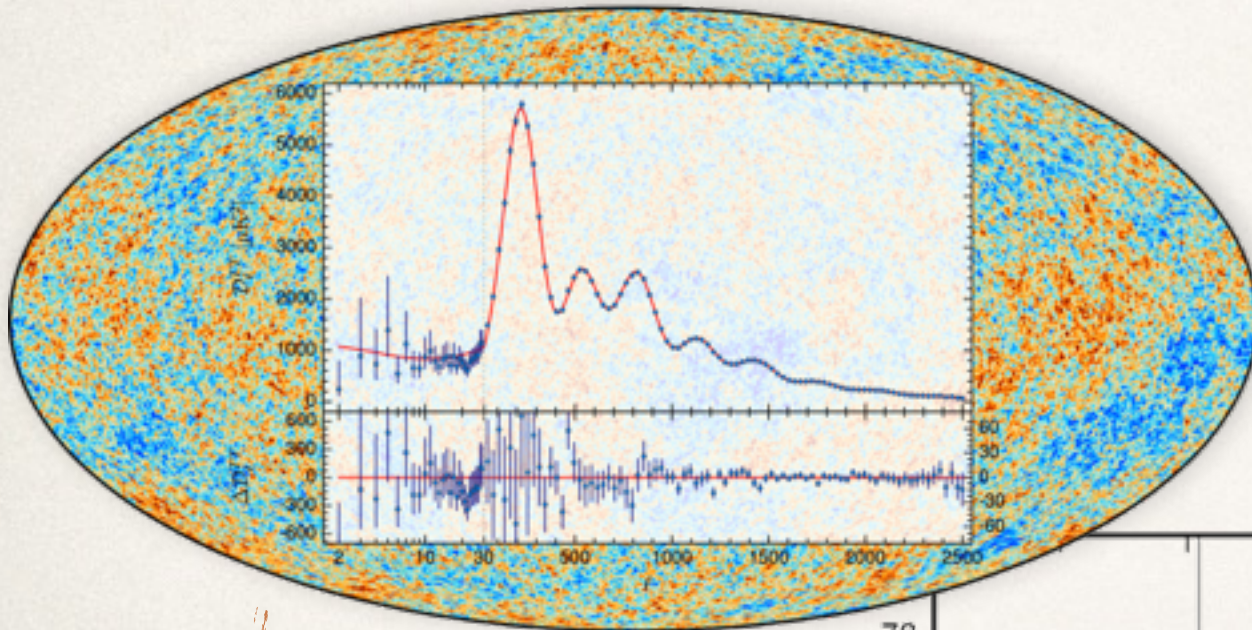
(2.5% ; Riess et al 2016)

The Hubble Constant

Planck 2015 — Cosmological Results

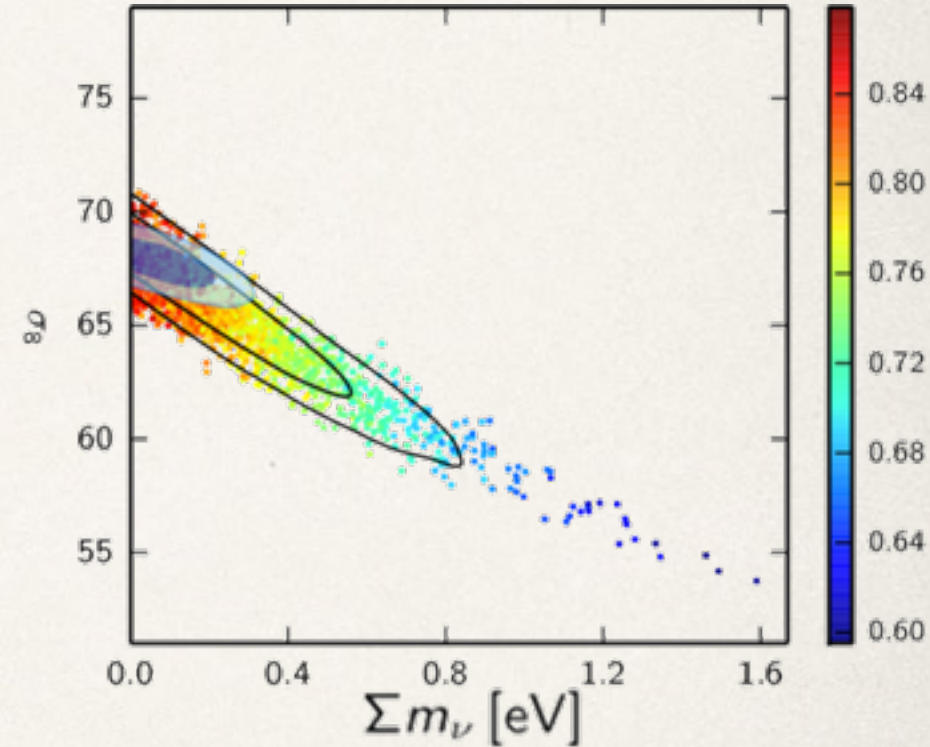
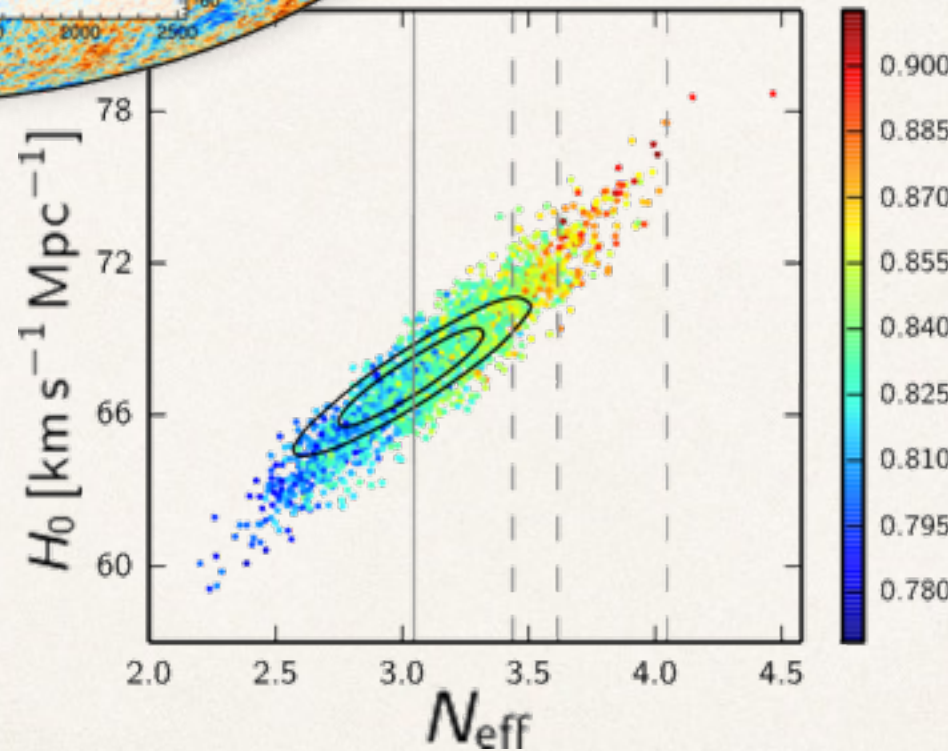
*Test the Standard Model
of Cosmology*

Change the model, change H_0



*THE MODEL
CONSTRAINS H_0*

$z \sim 1000$

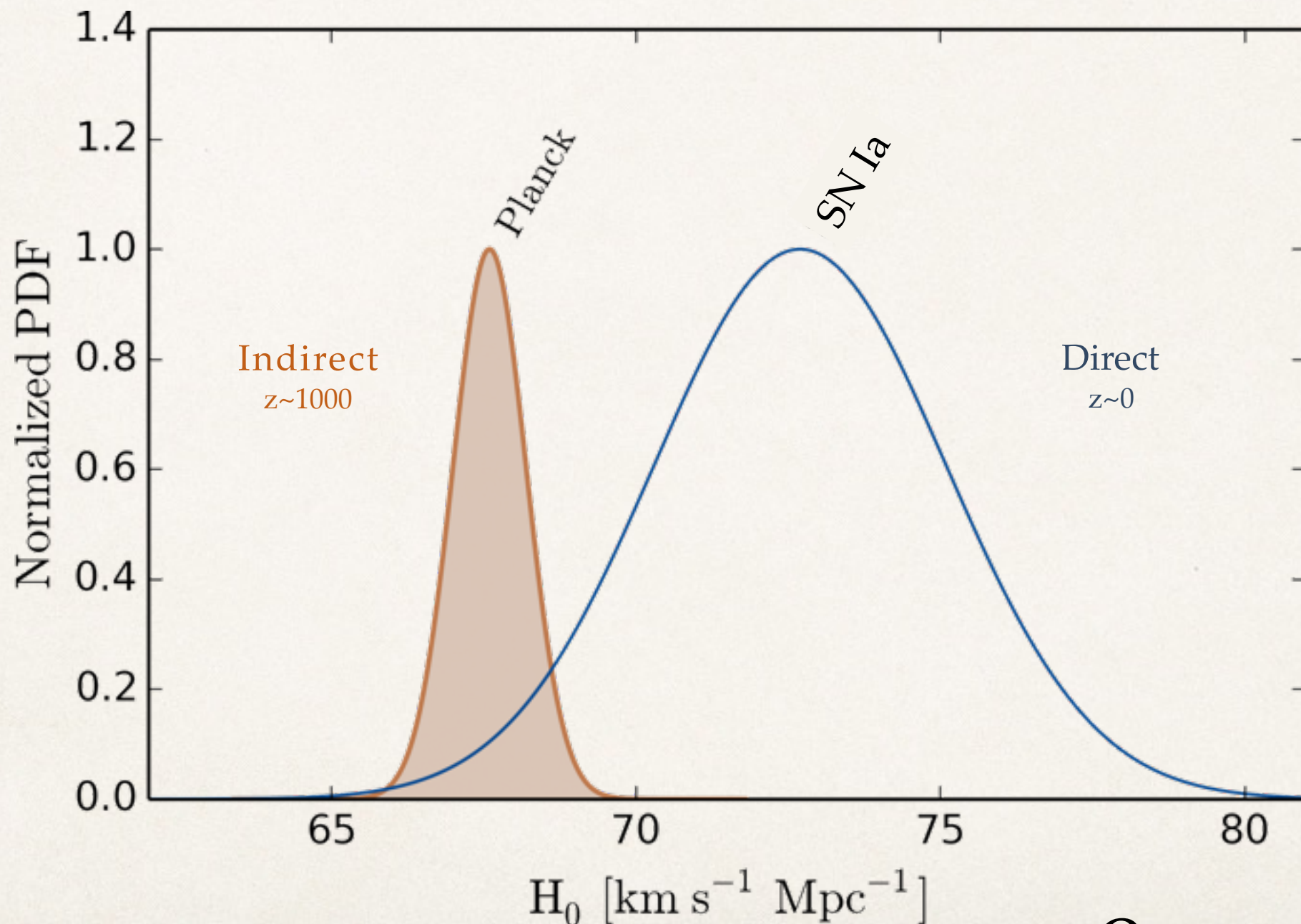


$z \sim 0$

$H_0 = 67.8 \pm 0.9 \text{ km s}^{-1} \text{Mpc}^{-1}$
— Derived —

Tension in the Standard Model of Cosmology

Could it be an indication for a new (massless) particle ?



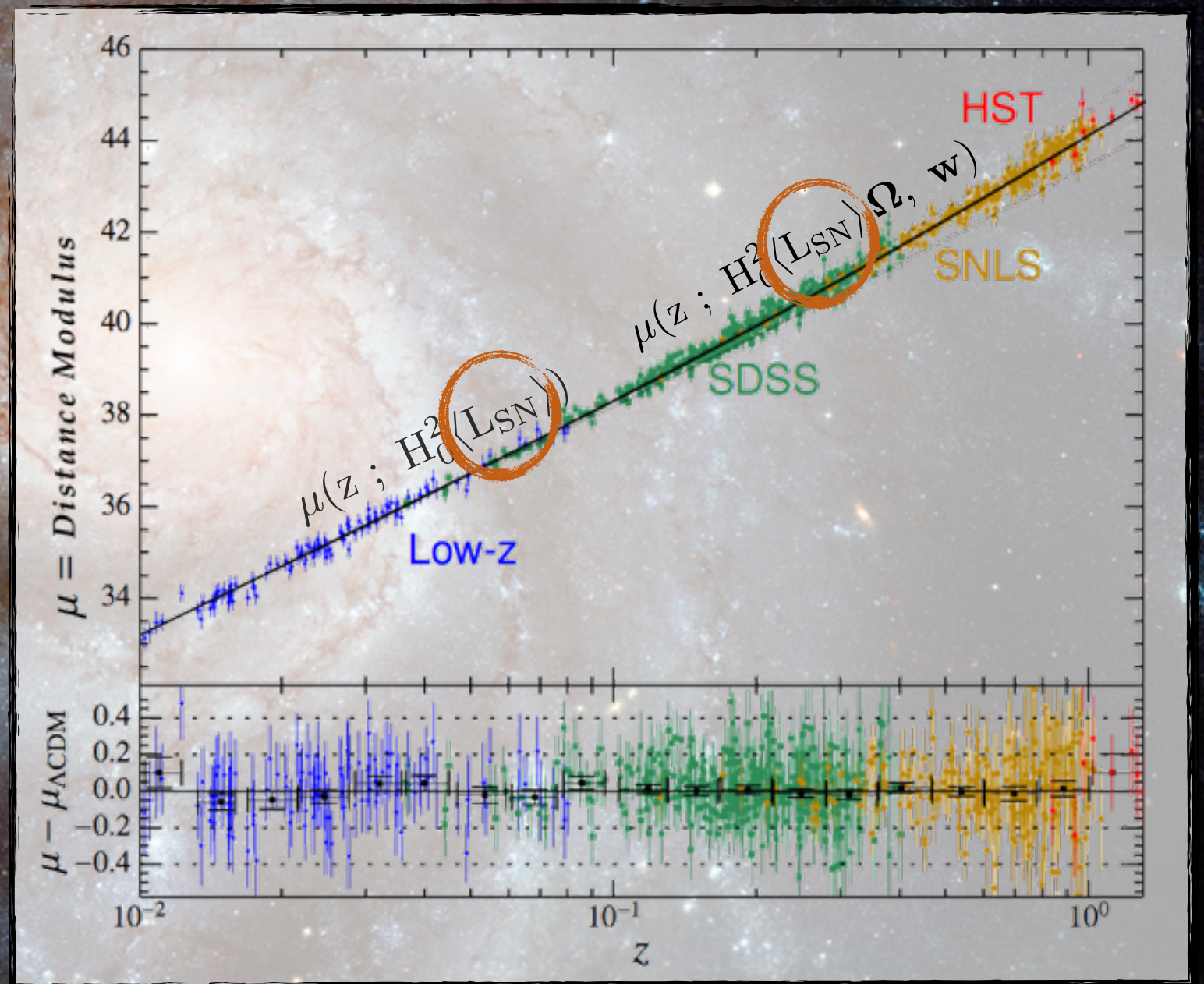
$\sim 2.5\sigma$

...Or a systematic error

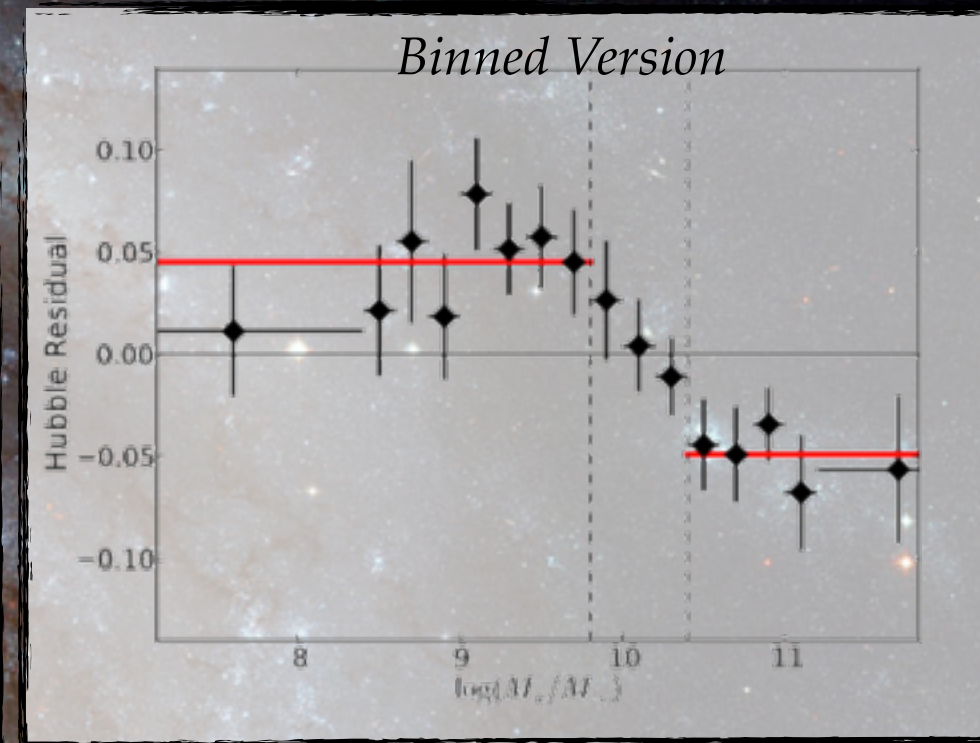
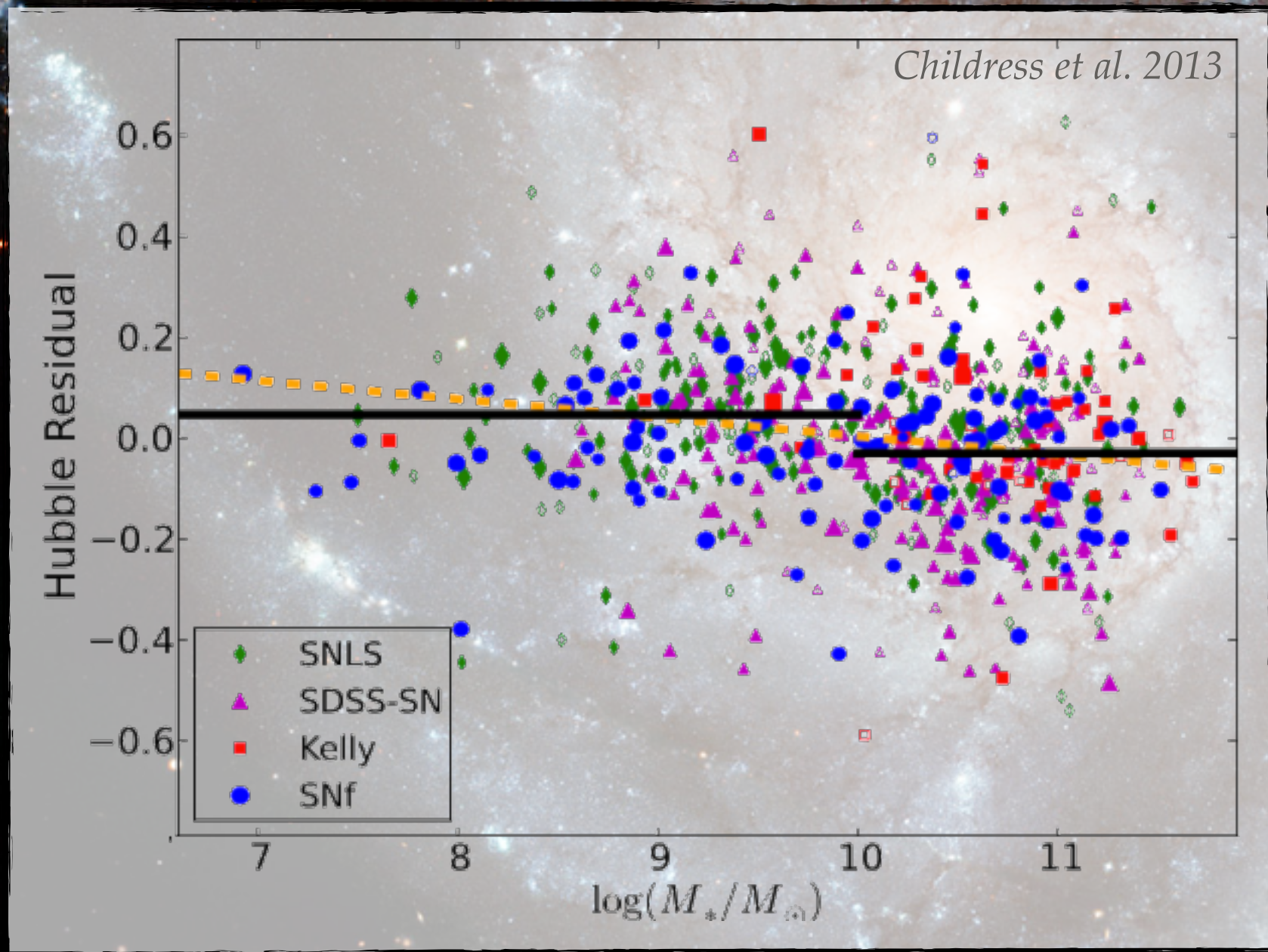
Precise Astrophysics for Accurate Cosmology



Potential Astrophysical Biases of SNe Ia



The first observed effect: The Mass Step

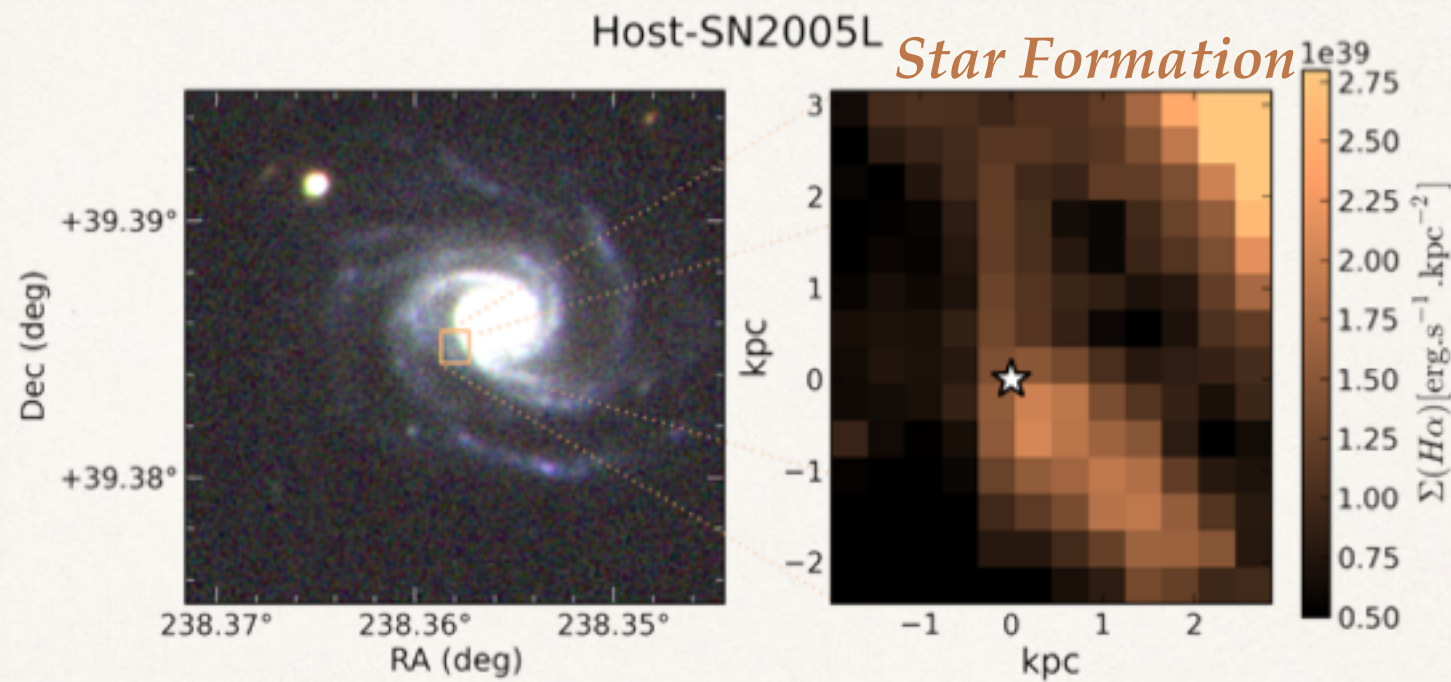




The Local Perspective

Rigault et al. 2013

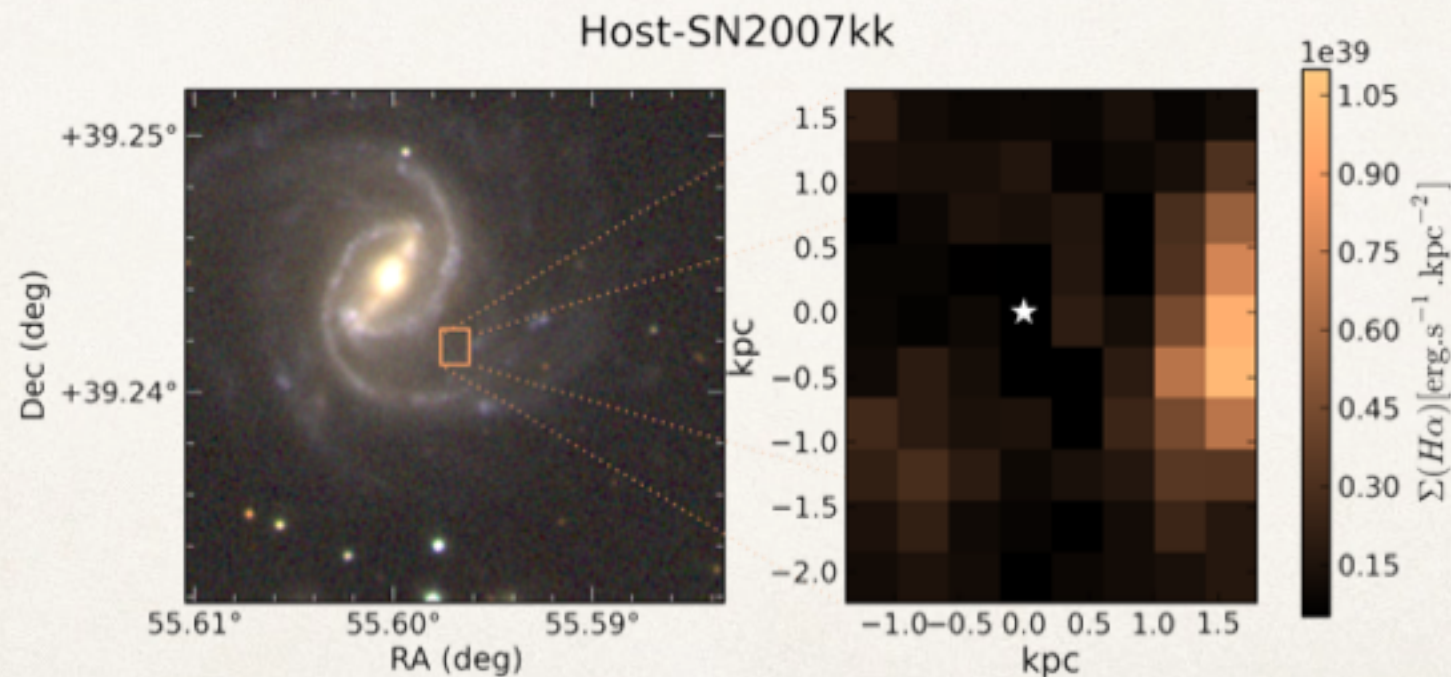
GLOBAL



LOCAL

Star Formation
—
Young Stars

Spiral, Star Forming,
host galaxies

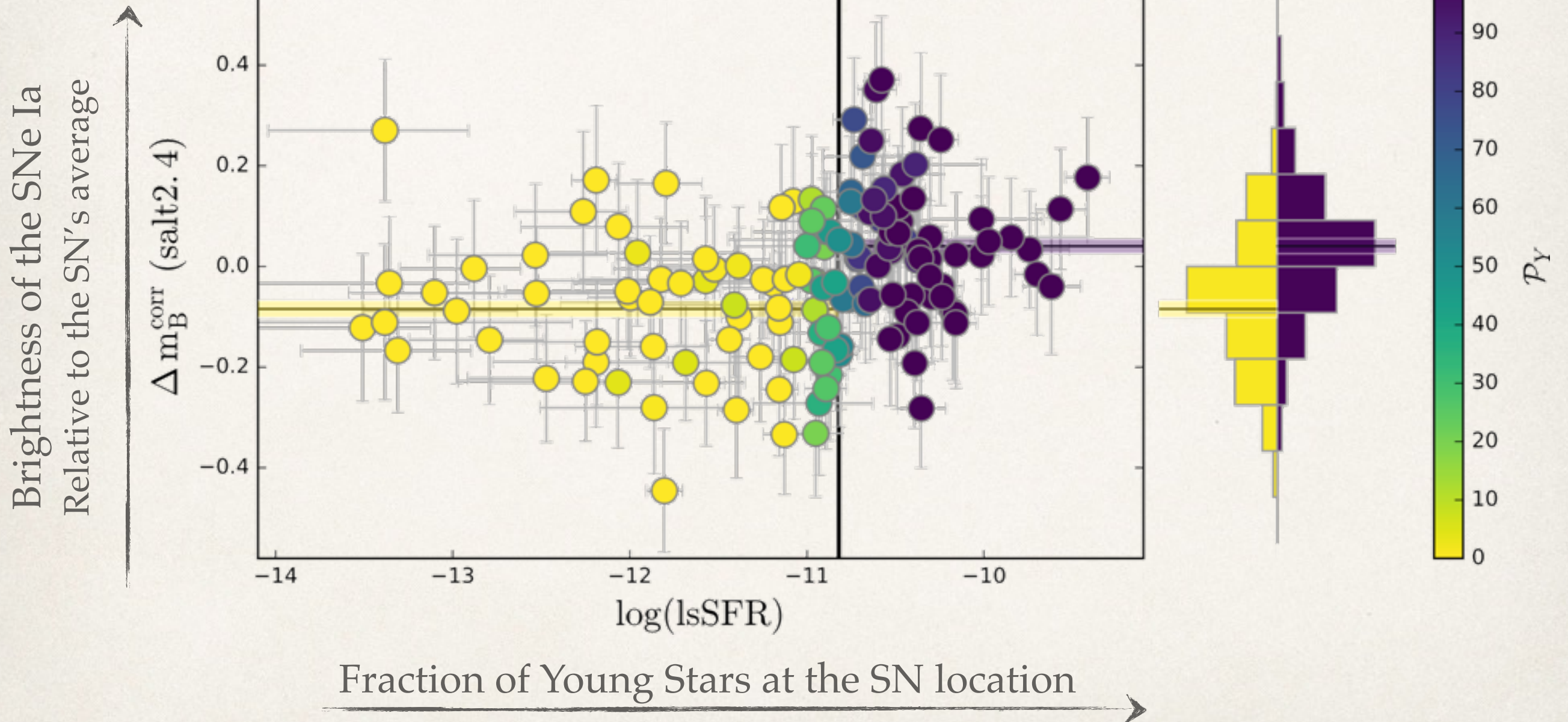


A Star Formation Bias

Rigault et al. 2013, 2015

Rigault et al. 2017*

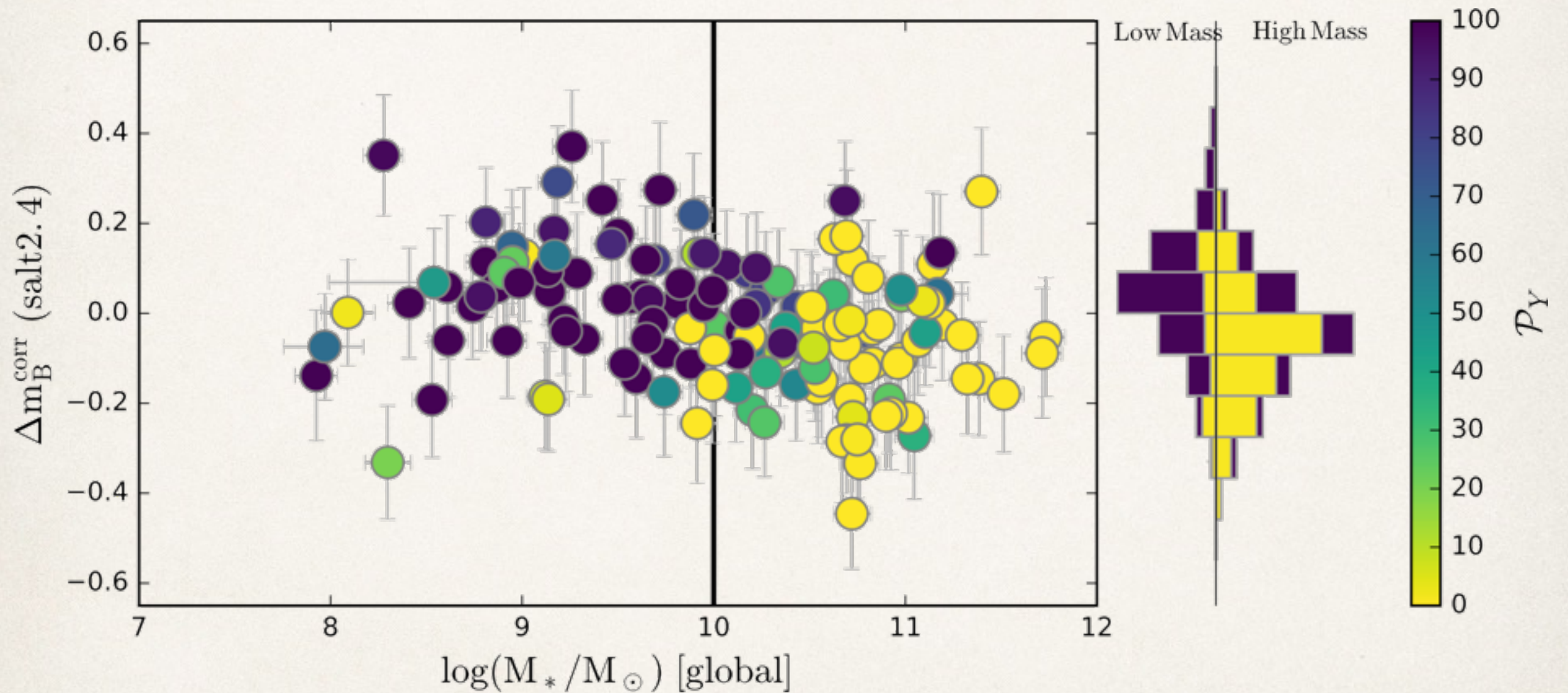
~6 σ detection level



Explaining the Mass Step

Rigault et al. 2017*

More massive galaxies have a lower fraction of young stars



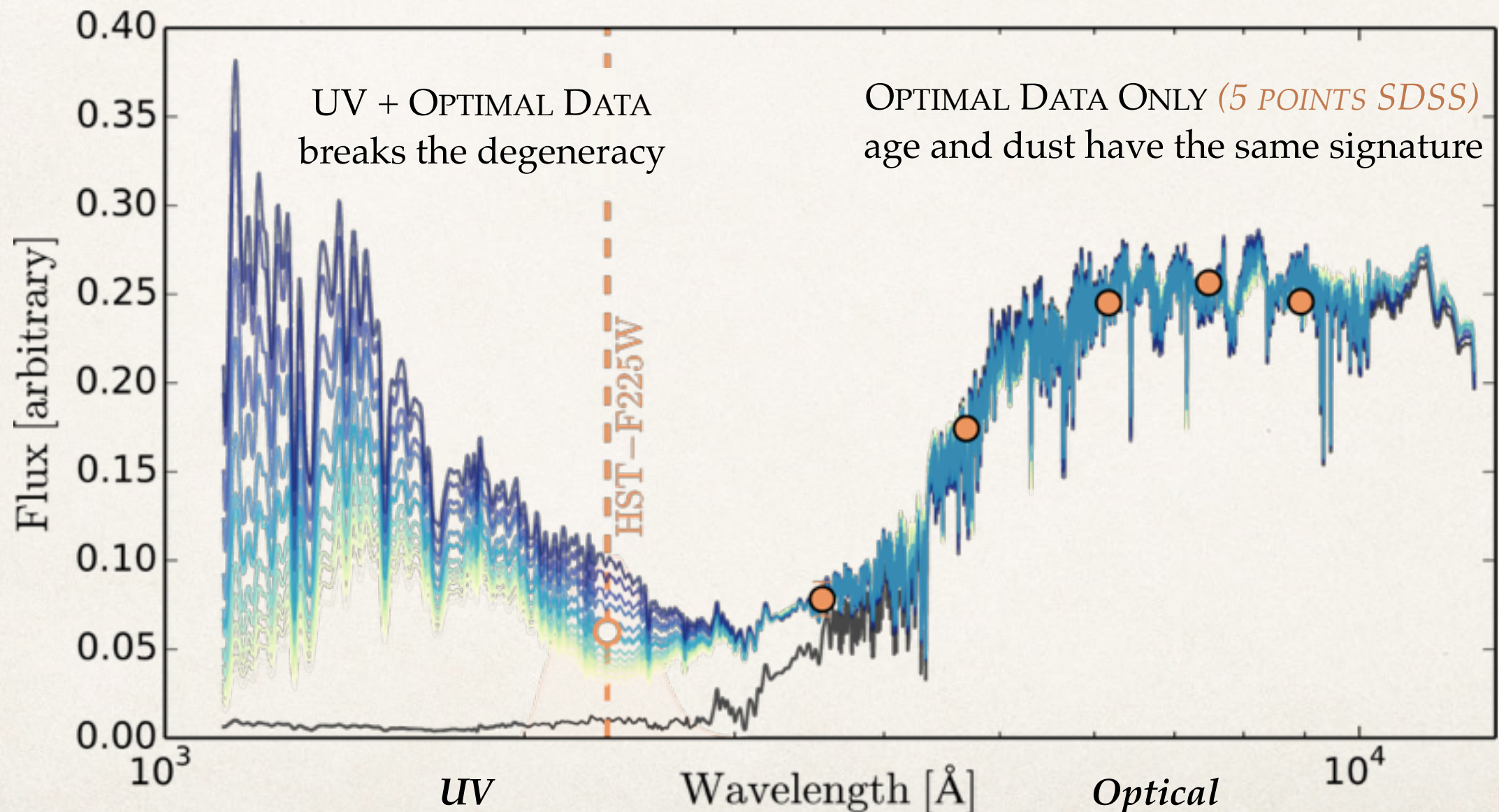
The Next Step

Understand the physical origin of the effects

PI: Rigault ; GO14163

Projet HST GO14163 (PI: Rigault) enables:

- to get the dust extinction along the line of sight
- to get the expected age of the progenitor

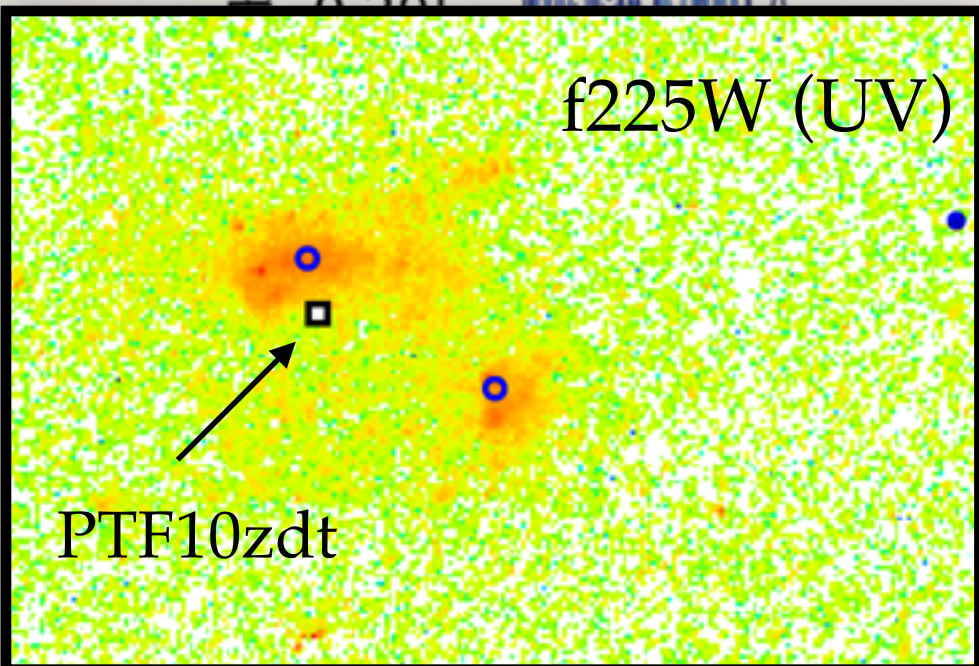
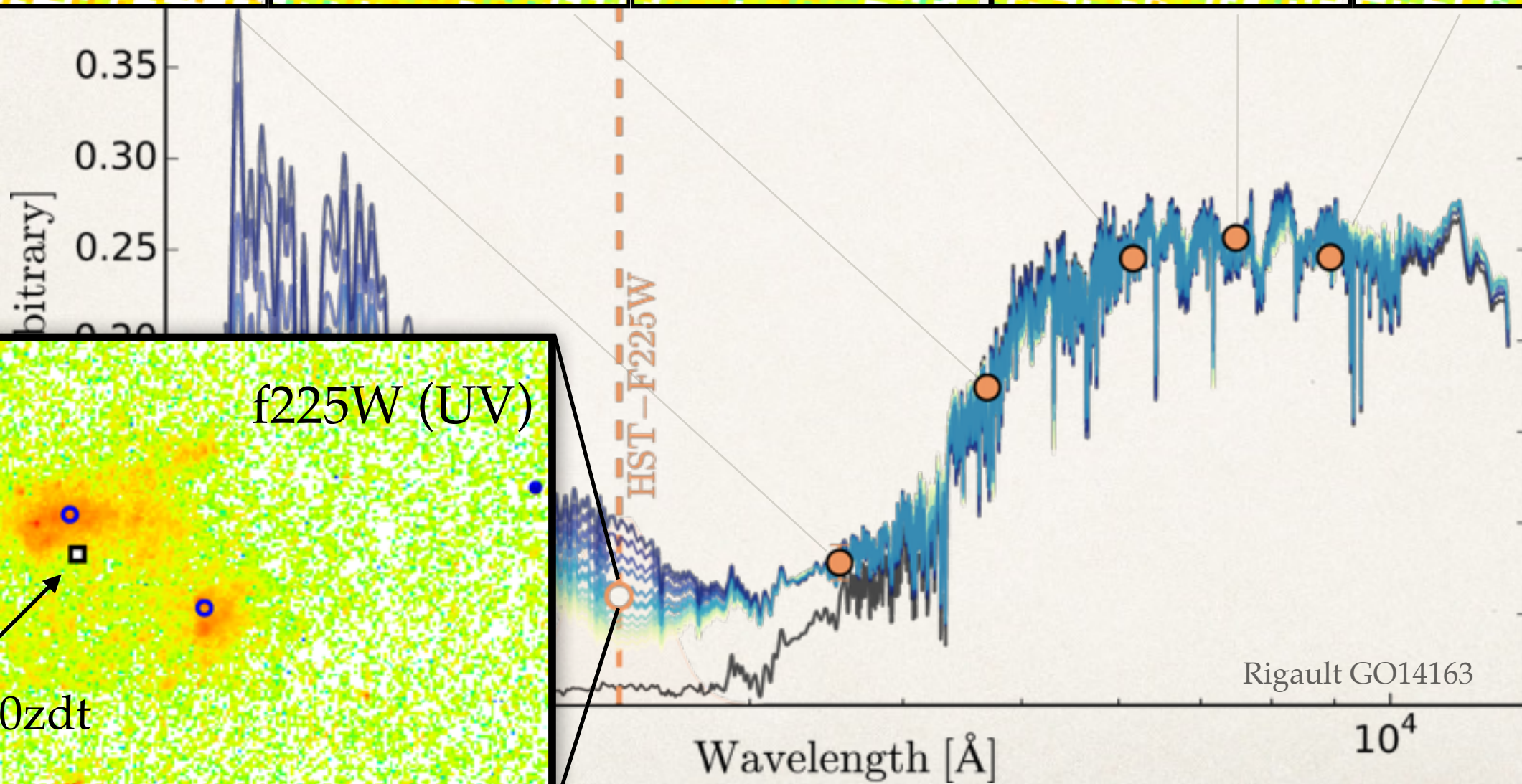
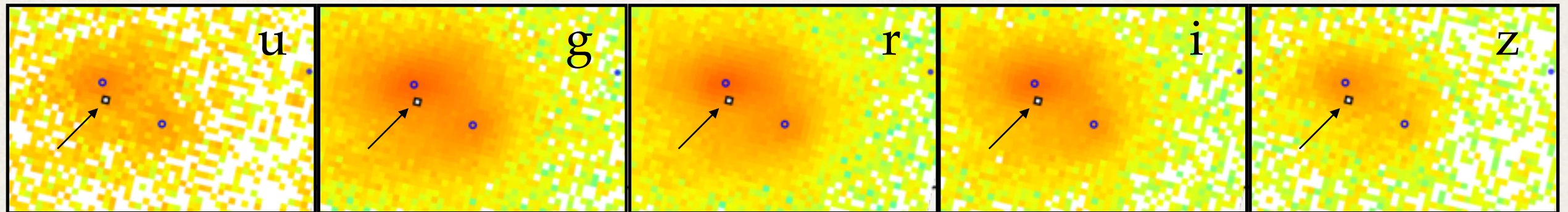


The Next Step

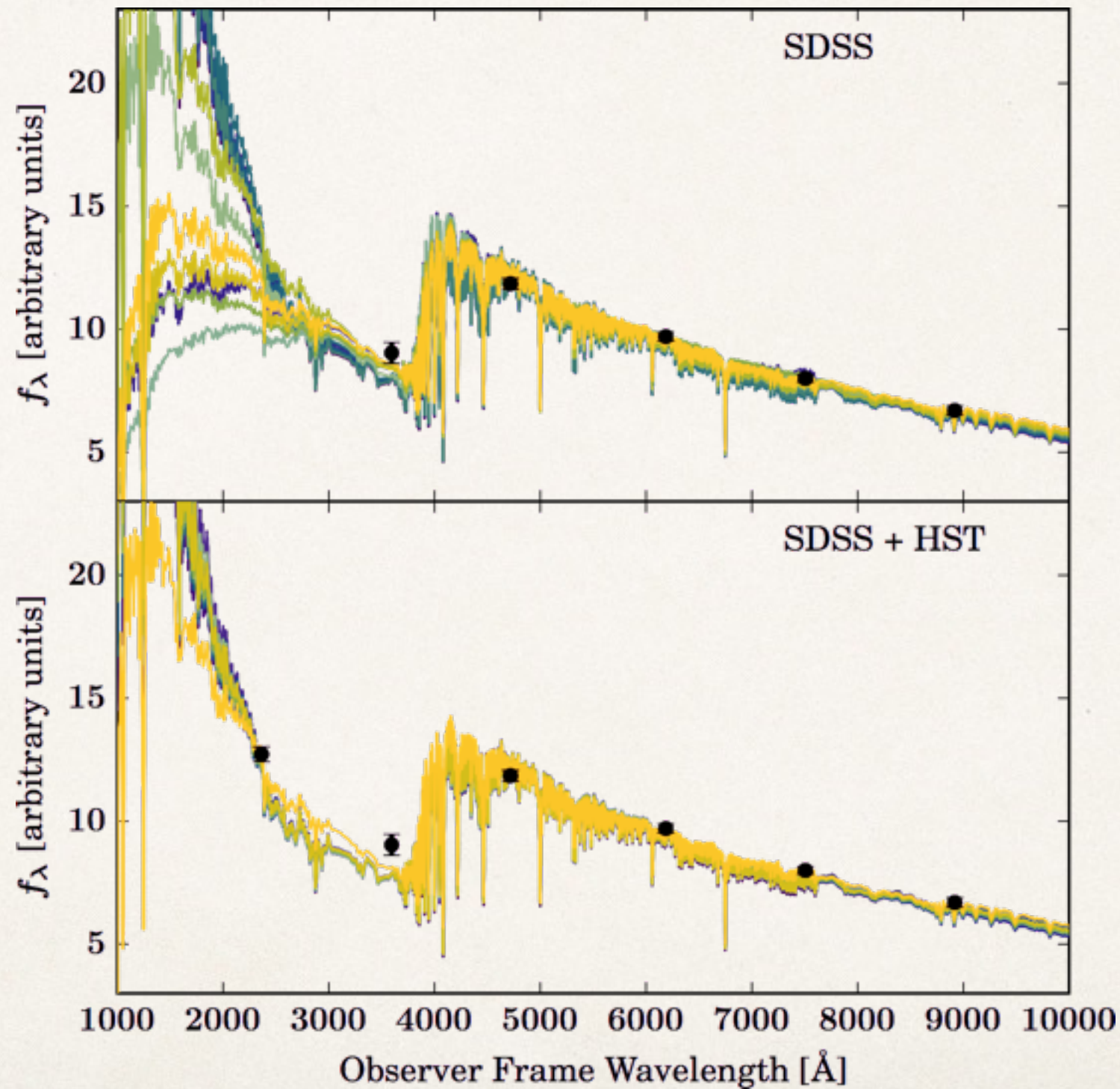
>60 SNeIa

Understand the physical origin of the effects

PI: Rigault ; GO14163



Data reduced and photometry extracted

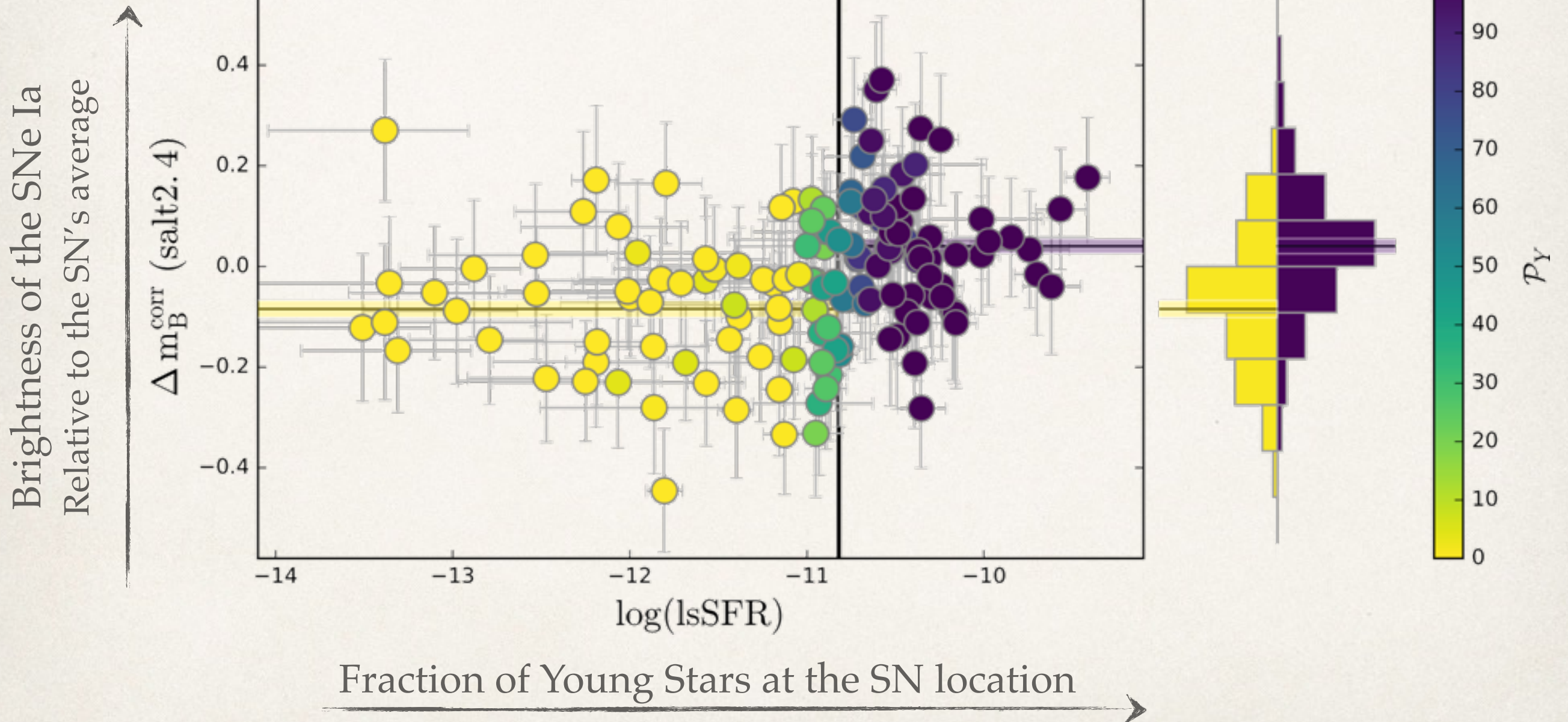


A Star Formation Bias

Rigault et al. 2013, 2015

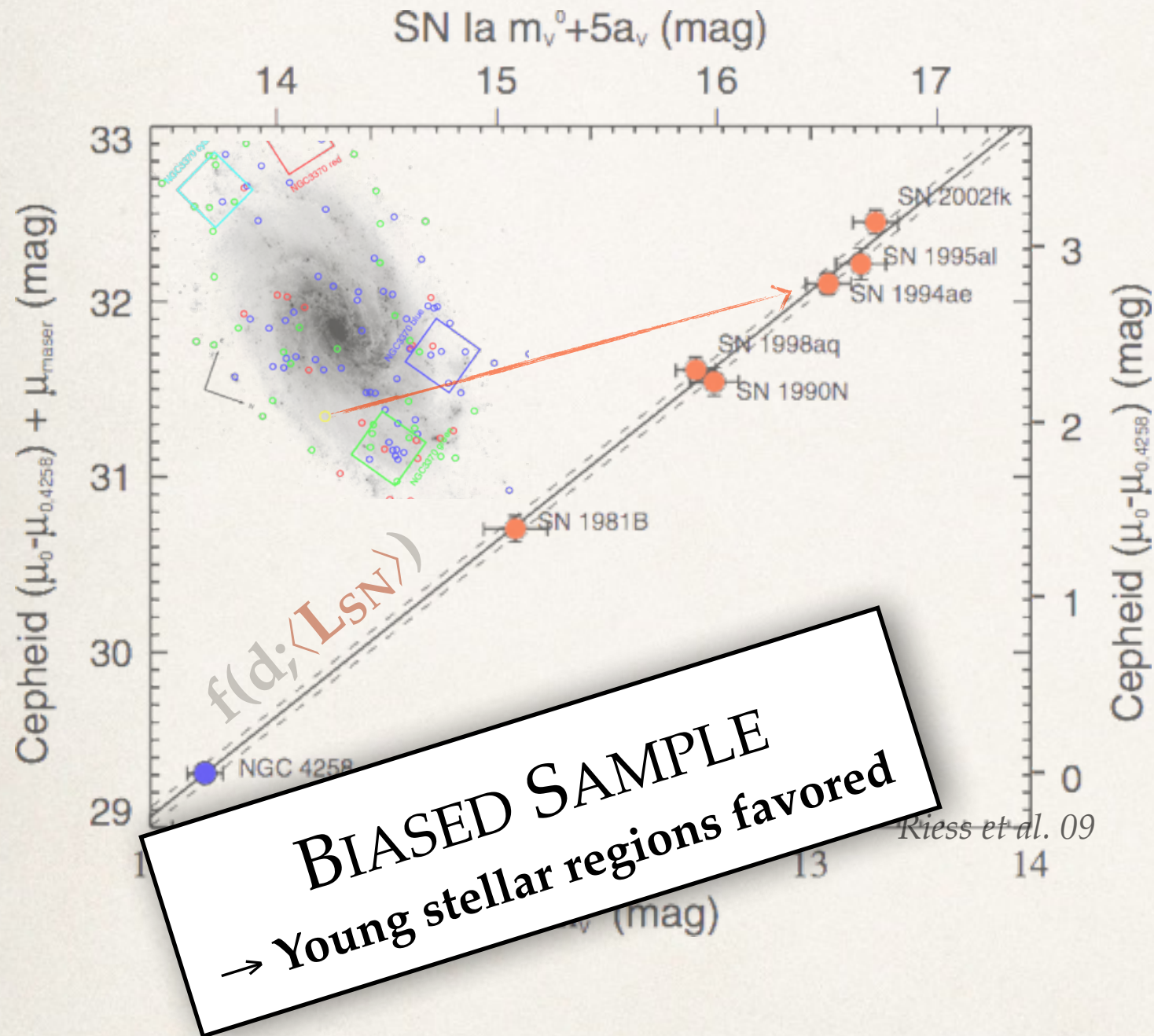
Rigault et al. 2017*

~6 σ detection level



Disentangle H_0 and the SN luminosity

Sample Selection — Impact on H_0



Cepheids: Bright young stars with a pulsation-luminosity relation

MAESURES

SN w Cepheids : ~100 % Young
 SN in Hubble Flow : ~50 % Young

H_0 : OVERESTIMATED BY ~3%

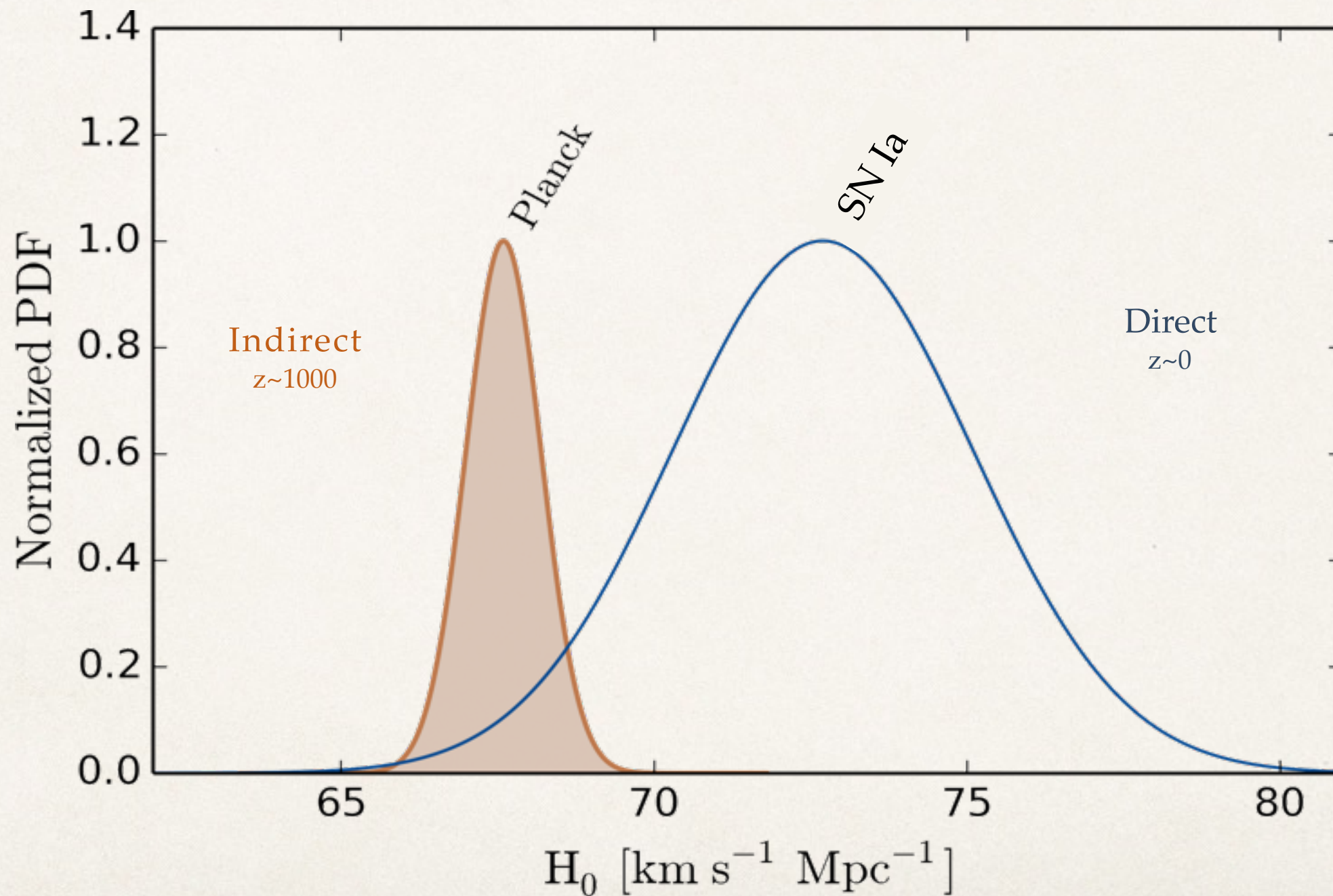
Rigault et al. 2015

$$H_0 = \del{73.2} \pm 1.8 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

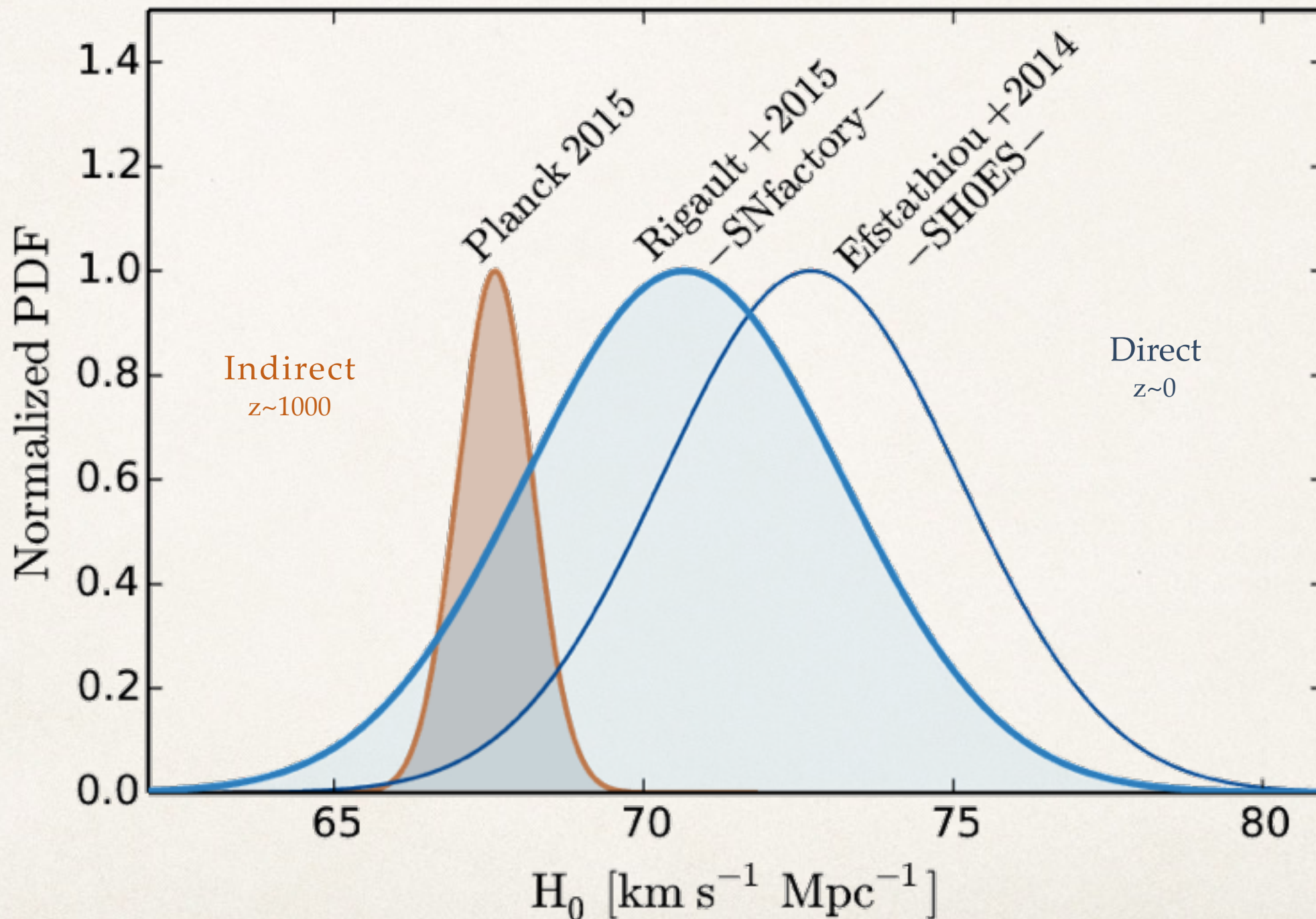
(2.5% ; Riess et al 2016)

Tension in the Standard model of Cosmology

$\sim 2.5\sigma$



~~Tension~~ in the Standard model of Cosmology



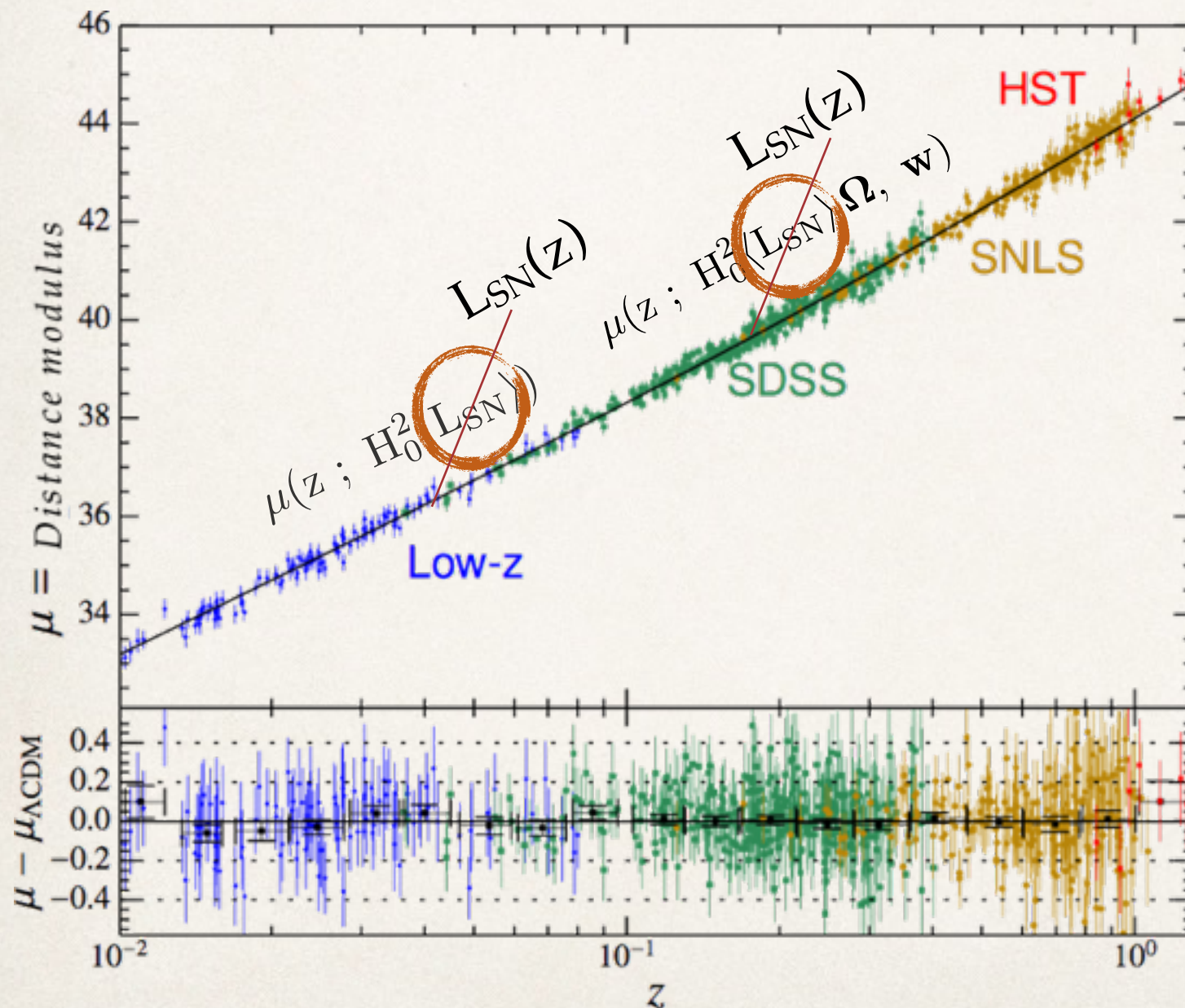
~~$\sim 2.5\sigma$~~

$\sim 1\sigma$

Consequence on the Cosmology — w

Rigault et al. 2013, 2017

Redshift evolution — Impact on w

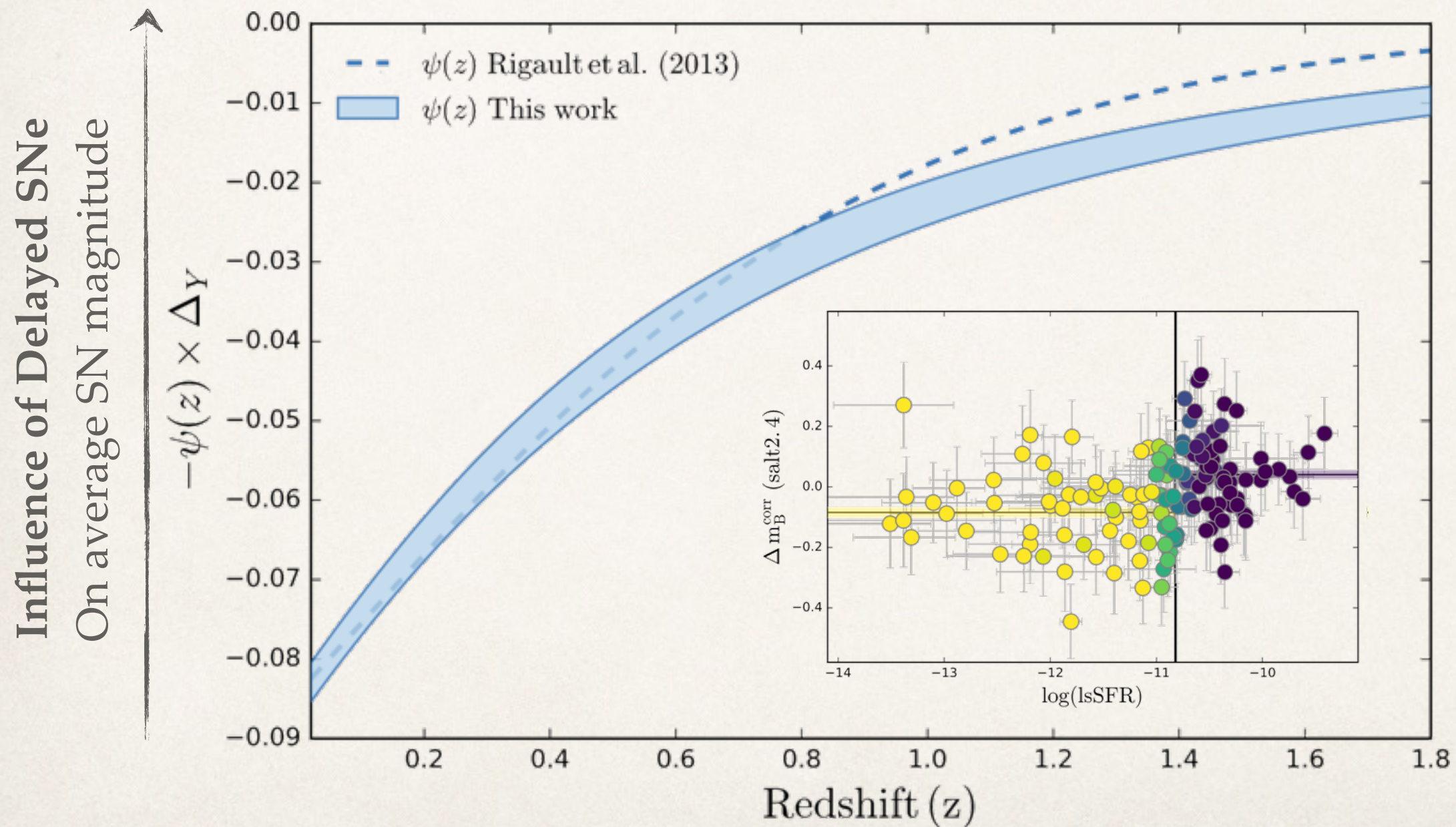


GALAXIES ARE MORE STAR FORMING AT HIGHER REDSHIFT

Could impact w by few percents
Goal of modern Cosmology: 1%

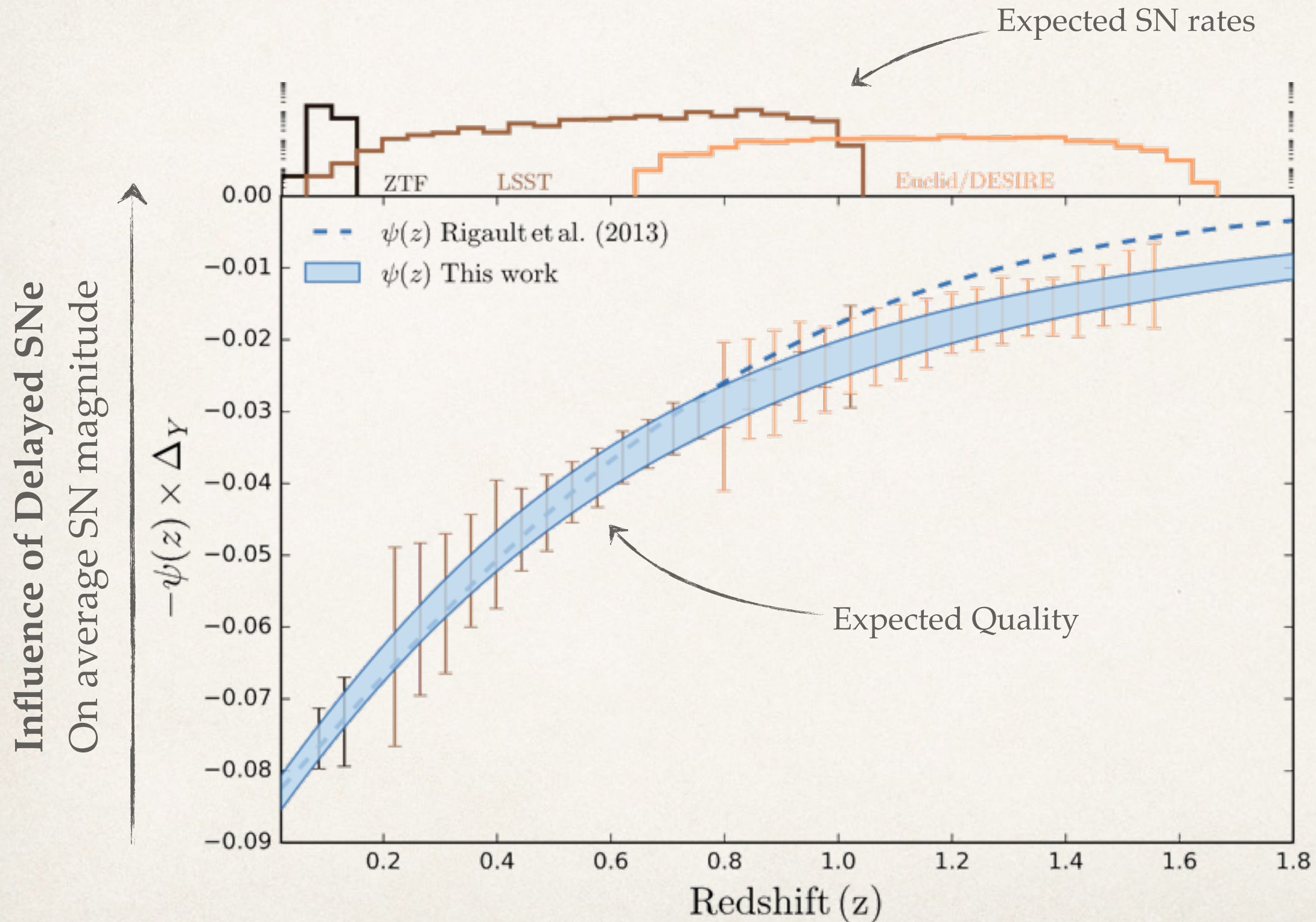
Astrophysical evolution or exotic dark energy?

Rigault et al. 2017*



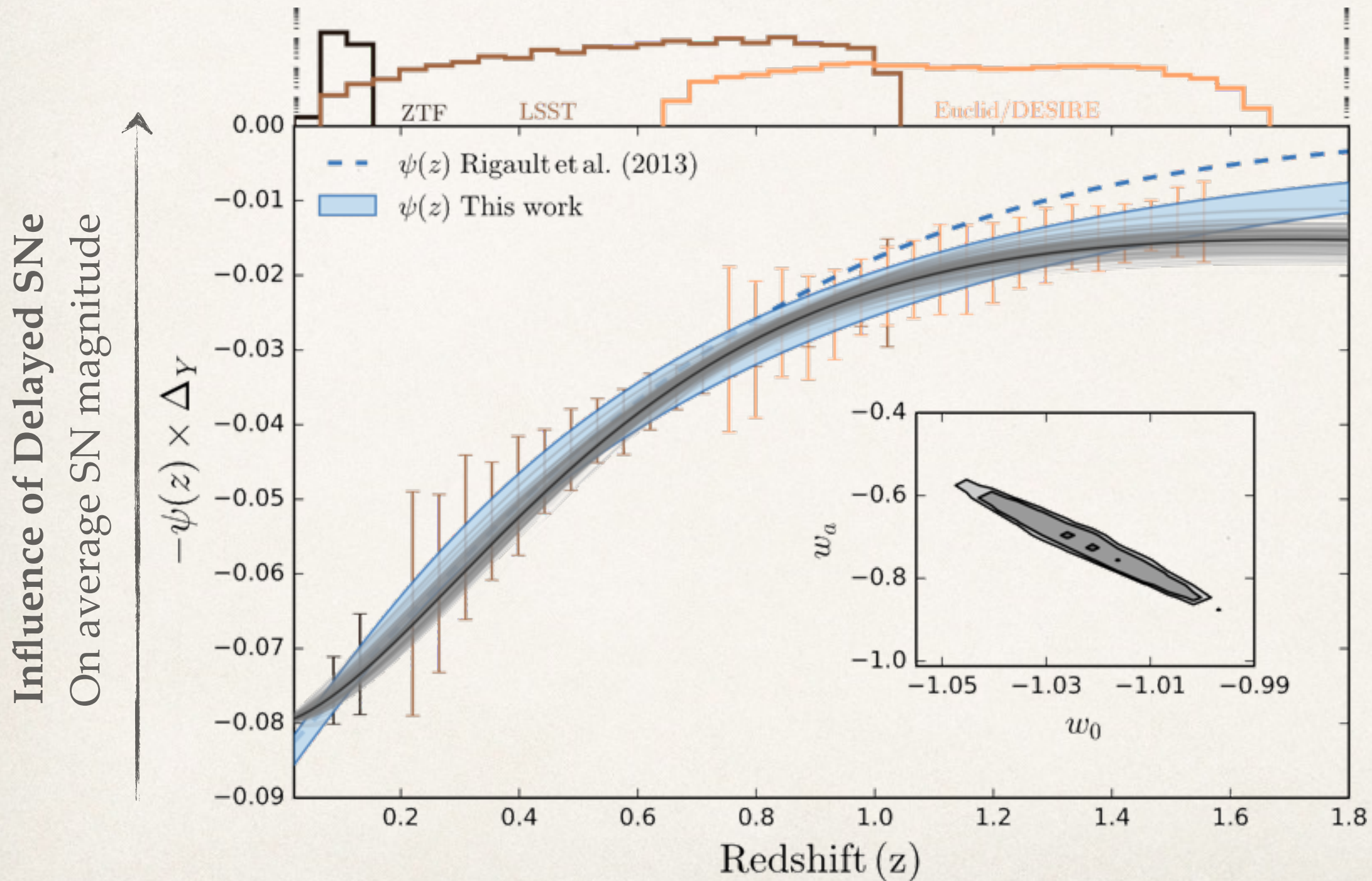
Astrophysical evolution or exotic dark energy?

Rigault et al. 2017



Astrophysical evolution or exotic dark energy?

Rigault et al. 2017

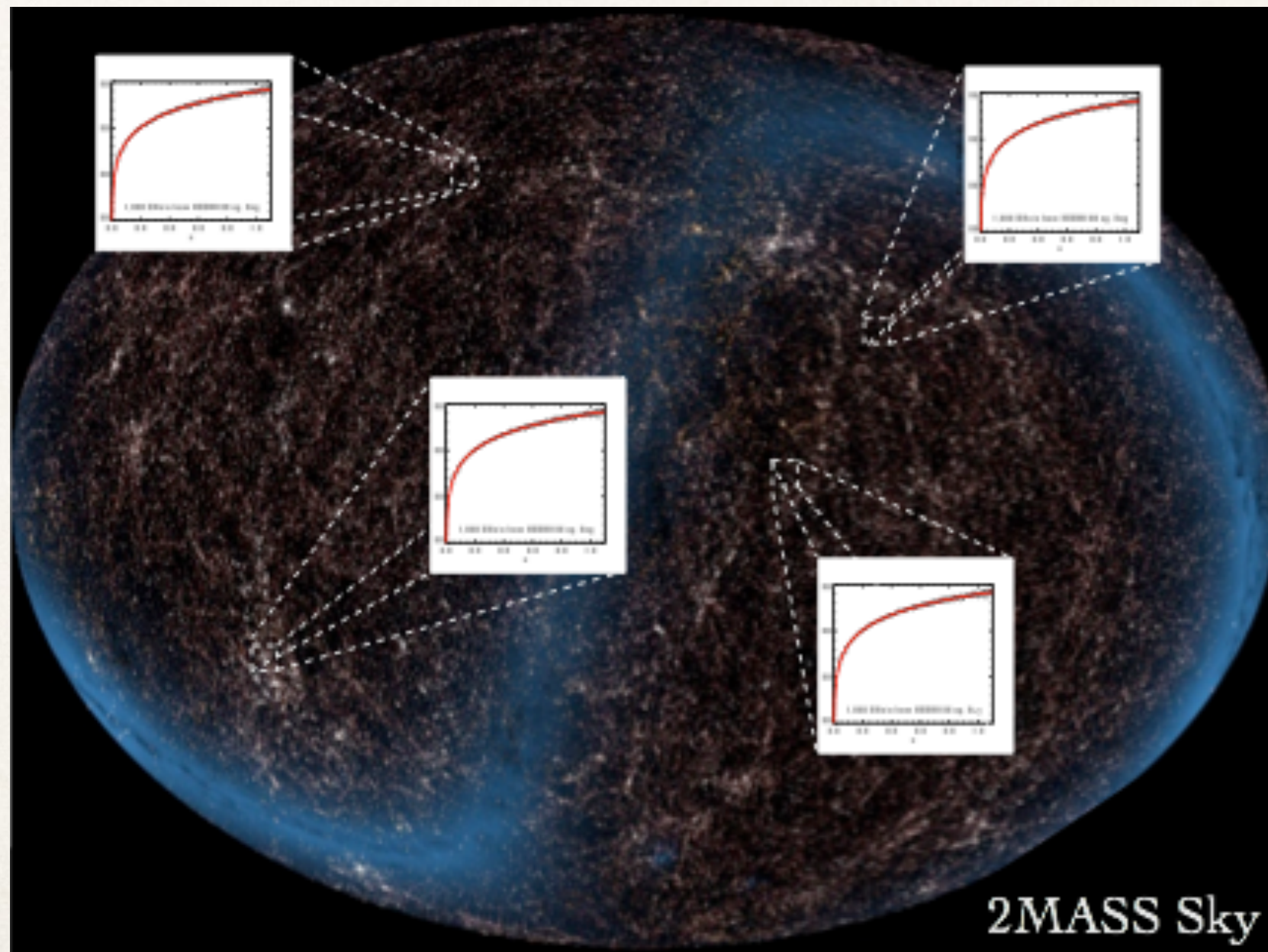




LSST

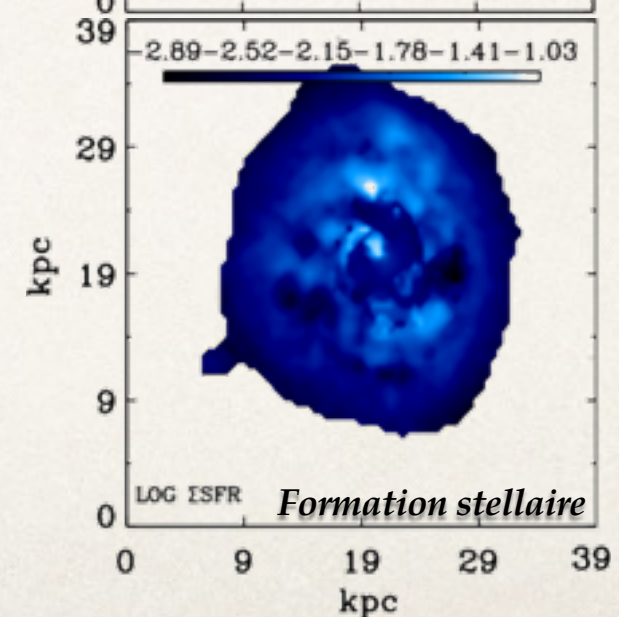
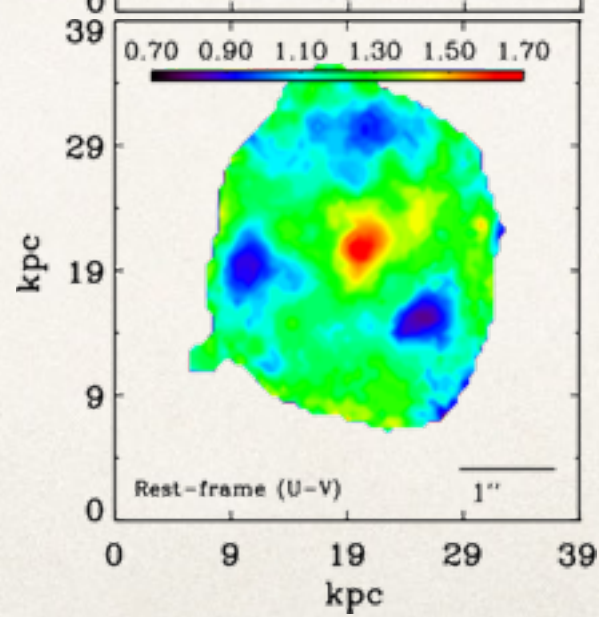
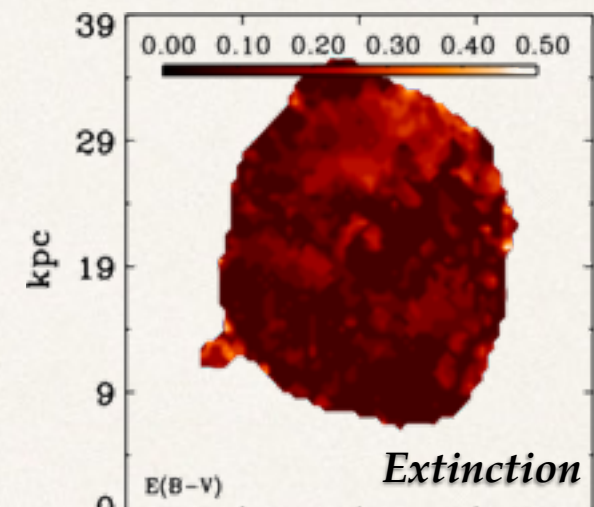
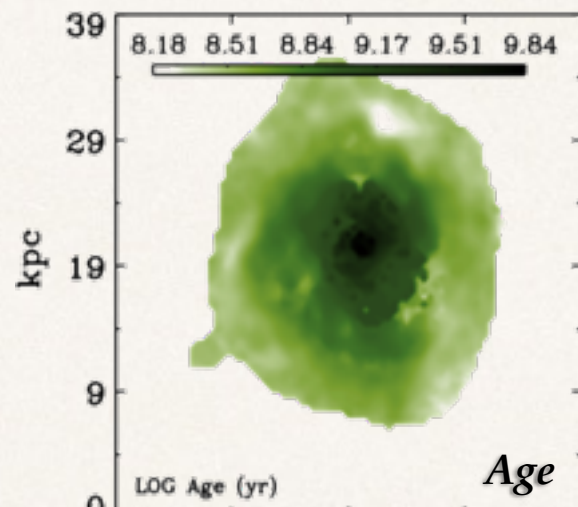
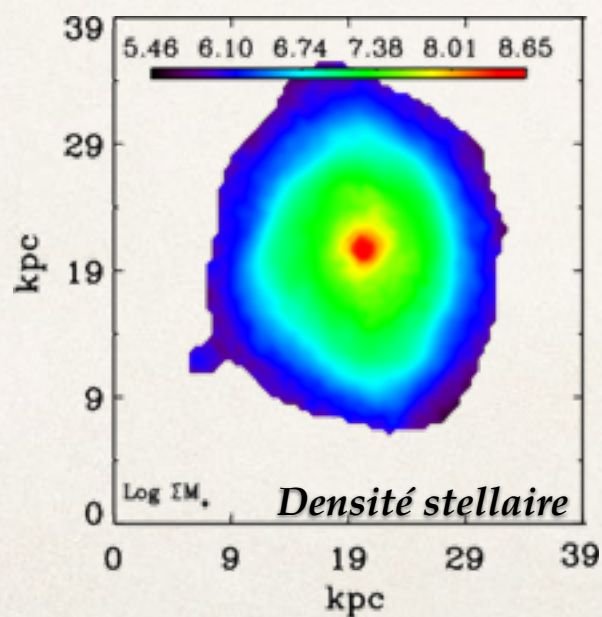
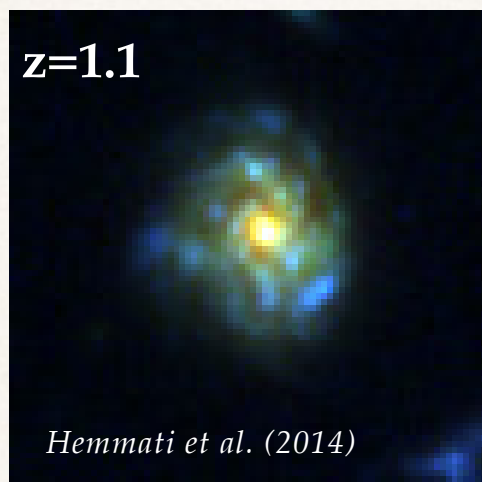
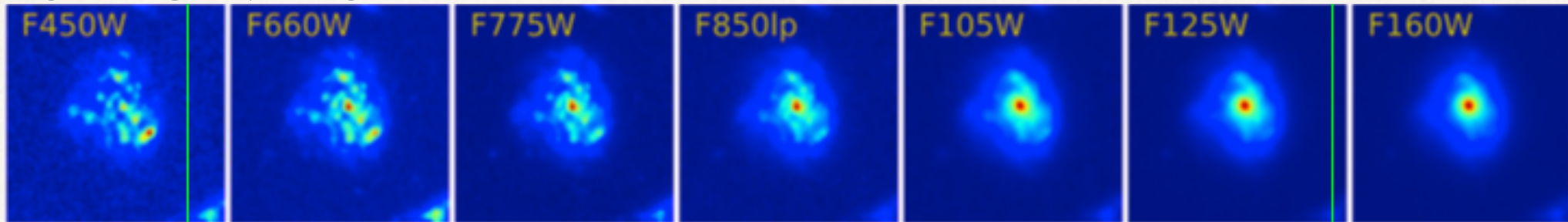
$\sim 10^6$ SNeIa!

+ other probes: Weak lensing, BAO, Clusters



Local Analyses at high-redshift

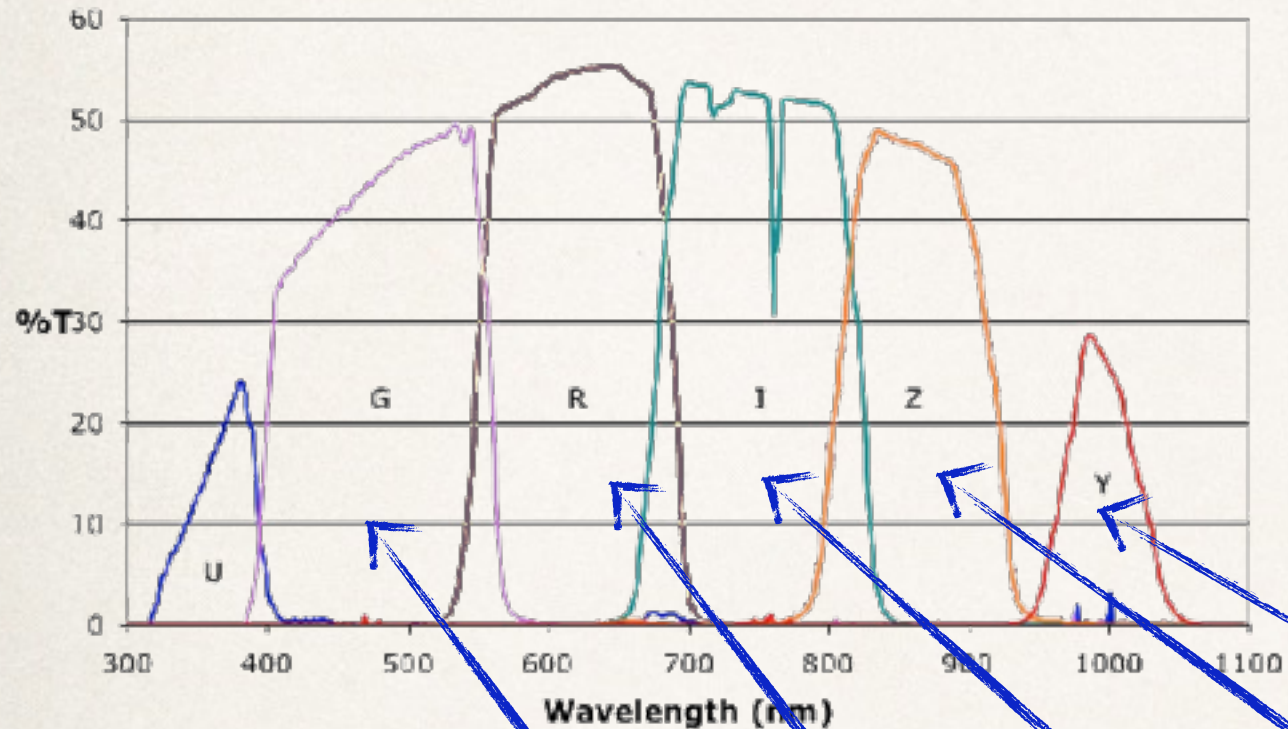
HST — CANDELS



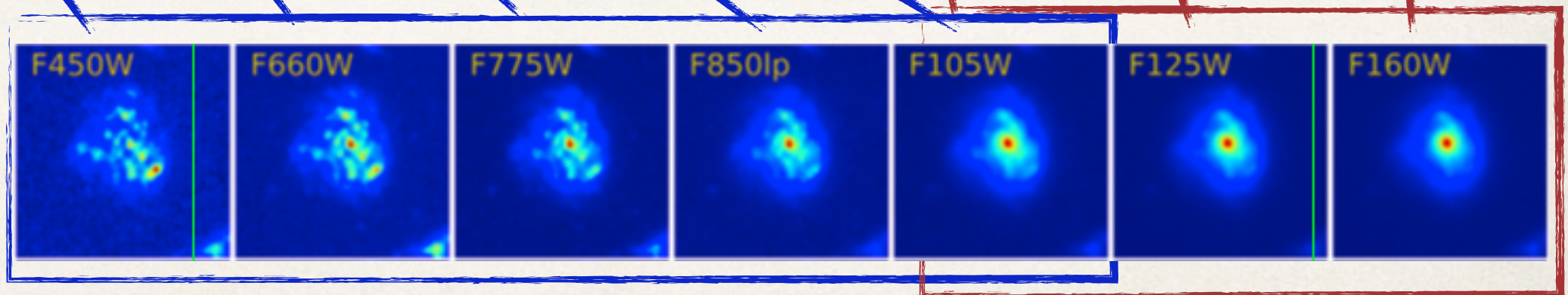
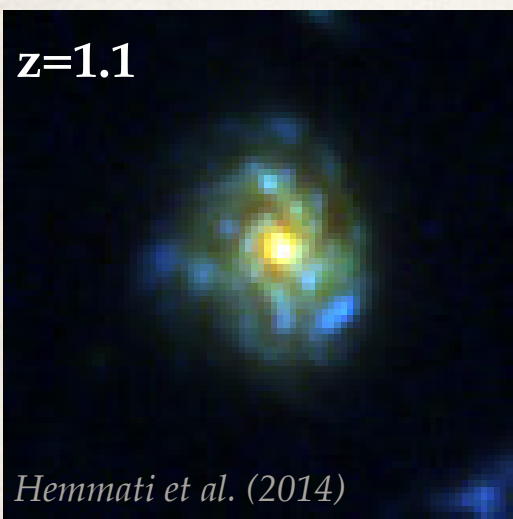
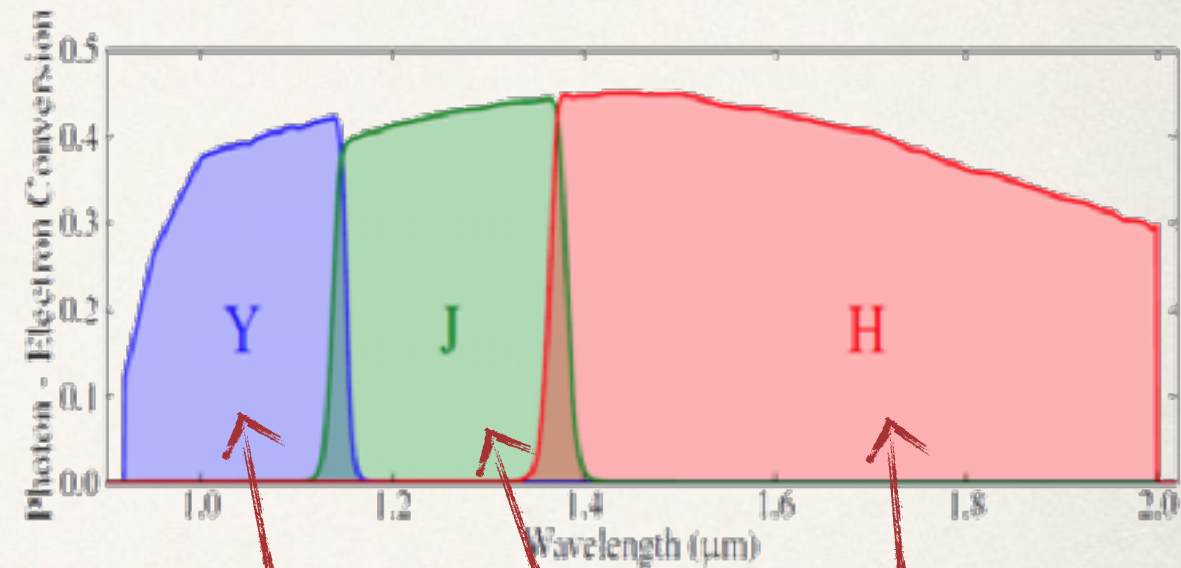
LSST & EUCLID

Getting the local environments of SNeIa

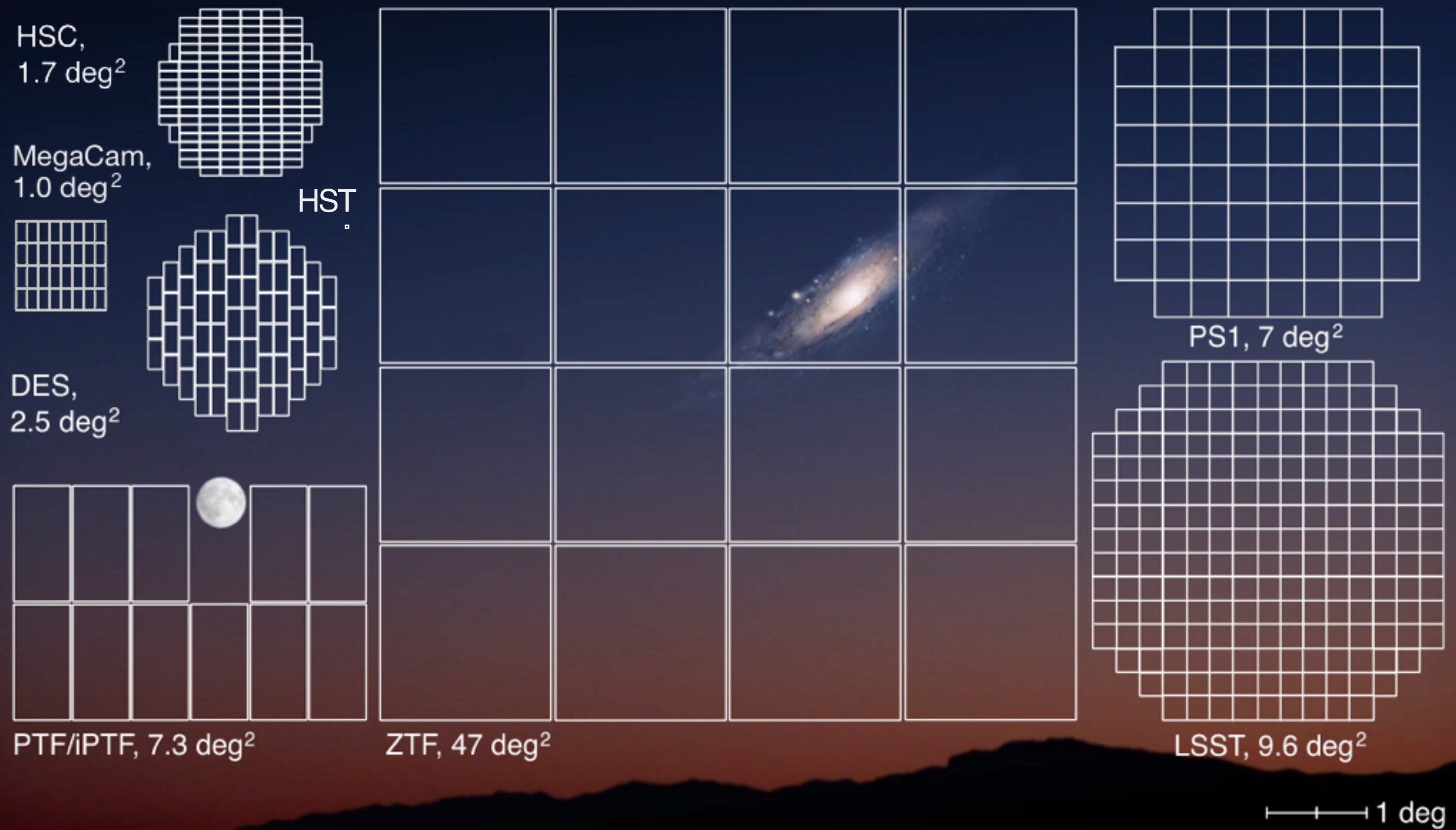
LSST ugrizY Filter Set



EUCLID Y, J, H



New Surveys and Time Domain Astronomy



1990 → 2000 ~ 100 SNe Ia

2000 → 2015 ~ 10³ SNe Ia

2017 → 2030 ~ 10⁶ SNe Ia

Discovery of Dark Energy

Looks like a Cosmological Constant

Is it really?

Conclusion

Type Ia Supernova are key to understand dark energy

Astrophysical effects have to be taken into account

We (will) have all the data we need to do so!

