

Supernova neutrinos flux reconstruction

in Čerenkov detectors

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Gallo Rosso, Vissani, Volpe, (2007) ArXiv:...

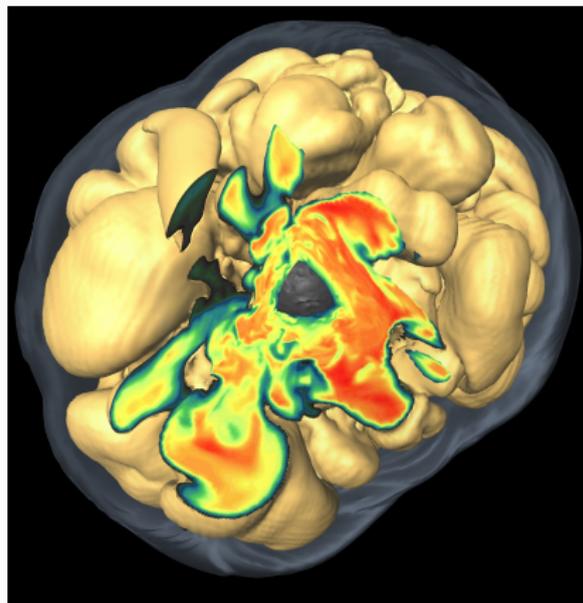
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Introduction

Delayed ν -driven explosion^[1]

1. Instability & collapse
2. Bounce & shock propagation
3. Shock stallation
4. Accretion
5. Cooling

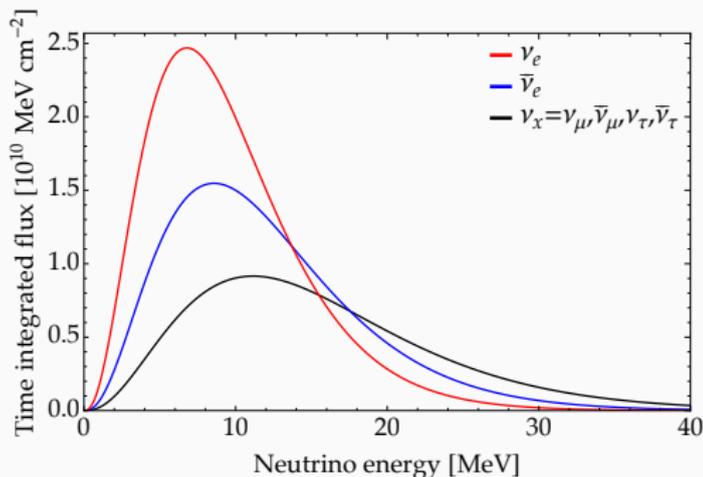


¹H. A. Bethe and J. Wilson, R. Astrophys. J. 295, 14 (1985).

²Image from Leonhard Scheck and H.-Thomas Janka.

Theoretical outcomes

- ~ 10 s signal
- $\sim 10^{53}$ erg
(gravitational)
- 99% in ν
- Equipartition?



Total energy \mathcal{E}	\Leftrightarrow	normalization
Mean energy $\langle E \rangle$	\Leftrightarrow	1 st moment
Pinching α	\Leftrightarrow	width

SN 1987A

- Large Magellanic Cloud (51.4 kpc)
- 25 $\bar{\nu}_e$ neutrino events
- Good agreement with predictions ^[3]
 - $\mathcal{E}_{\bar{e}} \sim 5 \times 10^{52} \text{ erg} \pm 10\%$
 - $\langle T_{\bar{e}} \rangle \sim 4 \text{ MeV} + 50\% - 20\%$
- ν_e, ν_x still missing
 - Future detection is crucial to prove equipartition hypothesis

³F. Vissani, J. Phys. G **42**, 013001 (2015).

Lu *et al.* ^[4]

- JUNO detector
- Importance of combining channels }
 - ◇ equipartition
 - ◇ no oscillation
 - ◇ pinching known
- $\mathcal{E}_{\bar{e}}$ up to 5% @ 90% C.L.
- $\langle E_{\bar{e}} \rangle$ up to 1% @ 90% C.L.

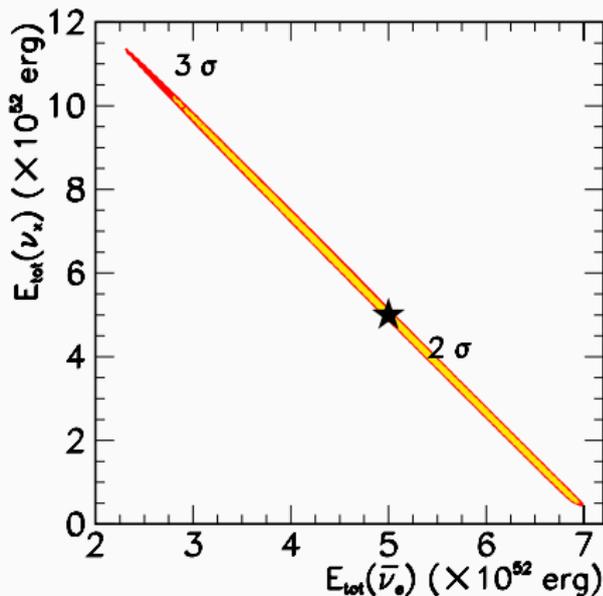
↓ with oscillation (MSW)
w/o equipartition

\mathcal{E}_{tot} known up to 13%
but for spectral shape fully known

⁴Lu *et al.* Phys. Rev. D 94, 023006 (2016)

H. Minakata *et al.* [5]

- Hyper-Kamiokande
- IBD only
- Pinching known
 - $\mathcal{E}_{\bar{e}}$ acc. 15% @ 3σ
 - $\langle E_{\bar{e}} \rangle$ acc. 2% @ 3σ
- Pinching unknown
 - $\mathcal{E}_{\bar{e}}$ acc. 50% @ 3σ
 - $\langle E_{\bar{e}} \rangle$ acc. 4% @ 3σ
 - Parameters degeneracy



⁵H. Minakata *et al.*, JCAP 0812, 006 (2008)

This work

- Importance of adding channels
 - Full 3×3 analysis
- Ansätze released
 - No equipartition
 - Pinching unknown
 - (and oscillation)



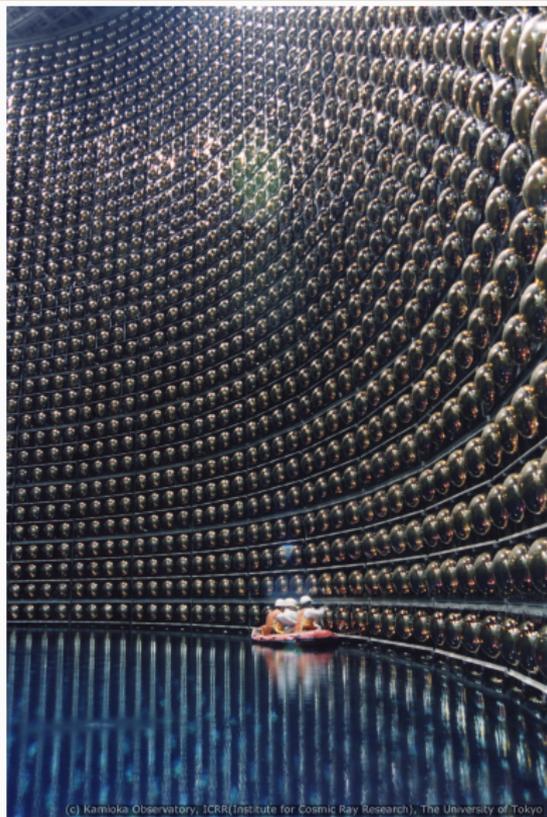
Hypotheses and method

Supernova

- Distance $D^* = 10$ kpc
- Total energy $\mathcal{E}^* = 3 \times 10^{53}$ erg

Super-Kamiokande

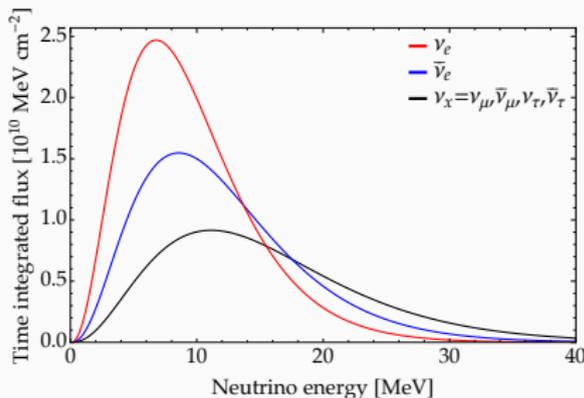
- 22.5 kton of fiducial mass
- 5 MeV threshold
- 100% efficiency
- Channels
 - Inverse Beta Decay (IBD)
 - Elastic scattering on e^- (ES)
 - Neutral-currents on ^{16}O (OS)



(c) Kamioka Observatory, ICRR (Institute for Cosmic Ray Research), The University of Tokyo

Fluences

- Quasi-thermal alpha-fit
- Oscillations: MSW effect (NH)
- Self-interaction neglected



$$\mathcal{F}_i^0(E_\nu) = \frac{dF_i}{dE_\nu} = \frac{\mathcal{E}_i}{4\pi D^2} \frac{E_\nu^{\alpha_i} e^{-E_\nu/T_i}}{T_i^{\alpha_i+2} \Gamma(\alpha_i+2)} \quad (i = \nu_e, \bar{\nu}_e, \nu_x)$$

$$\begin{cases} \mathcal{F}_e = \mathcal{F}_x^0 \\ \mathcal{F}_{\bar{e}} = |U_{e1}|^2 \cdot \mathcal{F}_{\bar{e}}^0 + (1 - |U_{e1}|^2) \cdot \mathcal{F}_x^0 \end{cases}$$

3 × 3 parameters

- Energy emitted \mathcal{E}_i
 - ⇒ $\mathcal{E}_i^* = 0.5 \times 10^{53}$ erg
 - ⇒ **Initial** equipartition hypothesis
- Mean energy $\langle E_i \rangle = (\alpha_i + 1) T_i$
 - ⇒ $\langle E_e \rangle^* = 9.5$ MeV
 - ⇒ $\langle E_{\bar{e}} \rangle^* = 12$ MeV
 - ⇒ $\langle E_x \rangle^* = 15.6$ MeV
- Pinching parameter α_i
 - ⇒ $\alpha_i^* = 2.5$

Monte Carlo

- Expected: $N_{IBD}^* = 4572$ — $N_{ES}^* = 213$ — $N_{OS}^* = 555$
- Extracted: $n_{IBD} = 4706$ — $n_{ES} = 207$ — $n_{OS} = 554$

Caveat on OS

- Signal in window of $4 \div 9$ MeV
 - Cannot disentangle from IBD and ES
 - Cross section uncertainty
 - $\sigma_{OS}(E_\nu) \approx \varepsilon \cdot \sigma_0 \cdot (E_\nu/\text{MeV} - 15)^4$
 - systematic $\sim \text{Gauss}(\varepsilon^* = 1, \sigma_\varepsilon = 0.1)$
- } \Rightarrow 10th parameter ε

Likelihood

$$\mathcal{L}_j(\text{param.}) \propto \prod_{i=1}^{N_{\text{bin}}} \frac{\nu_i^{n_i}}{n_i} e^{-\nu_i} \quad \text{with } j = \text{IBD, ES}$$

$$\mathcal{L}_{OS}(\text{param.}) \propto \exp \left[-\frac{(n_{OS} - N_{OS})^2}{2N_{OS}} - \frac{(\varepsilon - 1)^2}{2\sigma_\varepsilon^2} \right]$$

Analysis

1. IBD ($\mathcal{L} = \mathcal{L}_{\text{IBD}}$)
2. IBD + ES ($\mathcal{L} = \mathcal{L}_{\text{IBD}} \times \mathcal{L}_{\text{ES}}$)
3. IBD + ES + OS ($\mathcal{L} = \mathcal{L}_{\text{IBD}} \times \mathcal{L}_{\text{ES}} \times \mathcal{L}_{\text{OS}}$)

Prior

$$0.2 \times 10^{53} \text{ erg} \leq \mathcal{E}_i \leq 1.0 \times 10^{53} \text{ erg}$$

$$5.0 \text{ MeV} \leq \langle E_i \rangle \leq 30 \text{ MeV}$$

$$1.5 \leq \alpha_i \leq 3.5$$

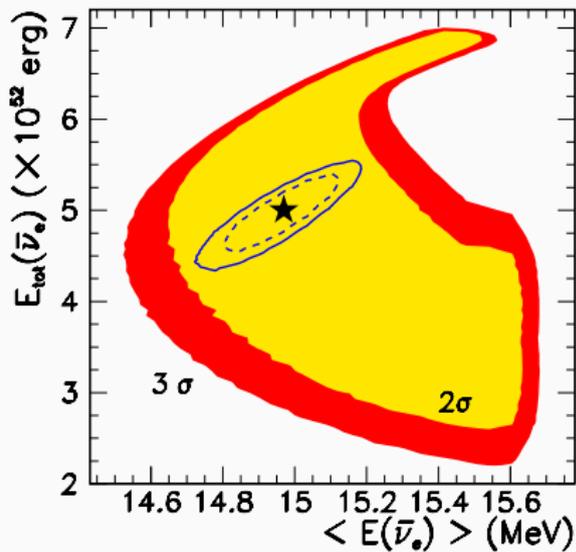
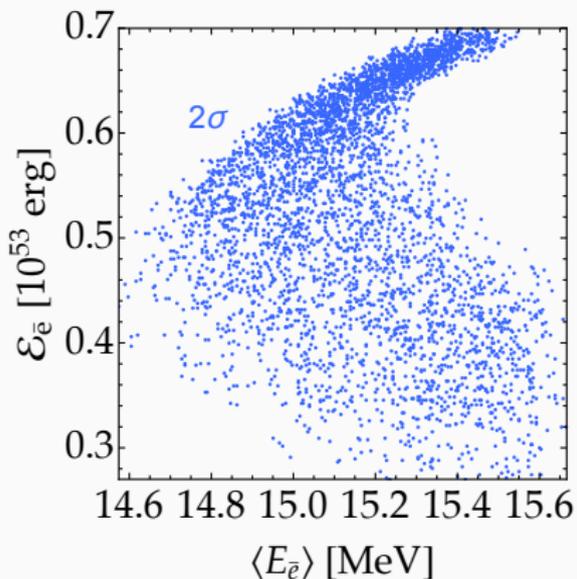
$$0.8 \leq \varepsilon \leq 1.2$$

Condition

$$\log \mathcal{L} \geq \log \mathcal{L}_{\max} - \frac{1}{2} A_{\text{dof,CL}} \quad \text{with} \quad \int_0^A \chi_{\text{dof}}^2(z) \, dz = \text{C.L.}$$

Results

Results

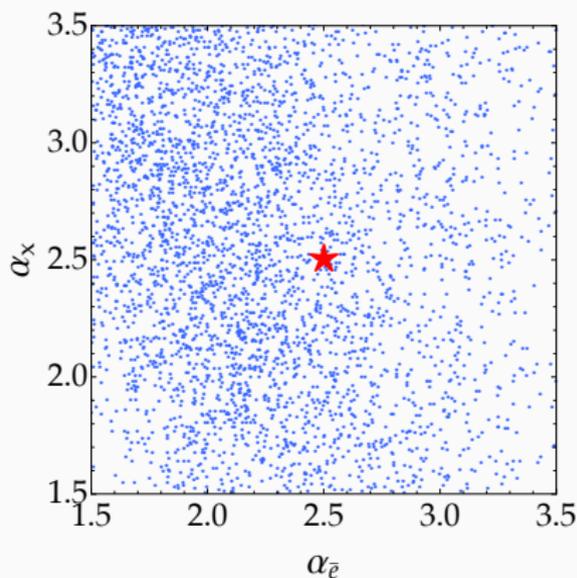
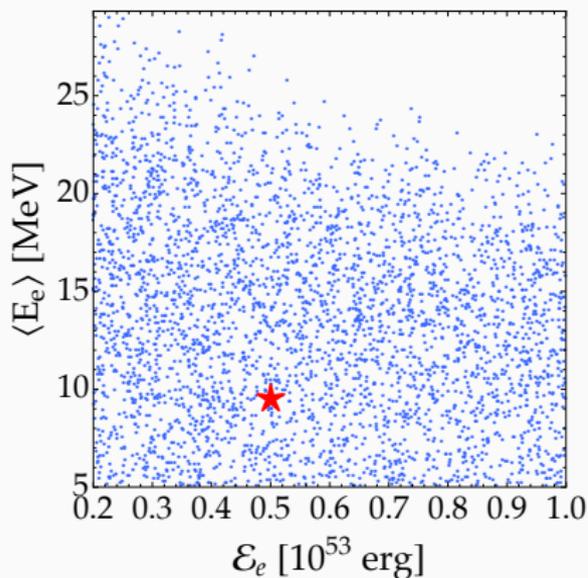


Validation

Good agreement with Minakata *et al.* (2008)

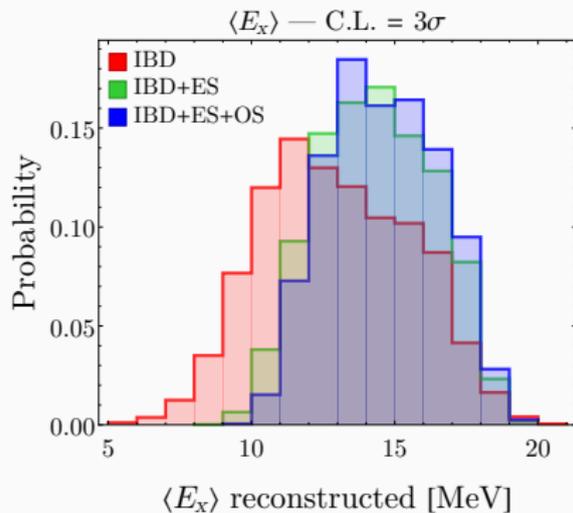
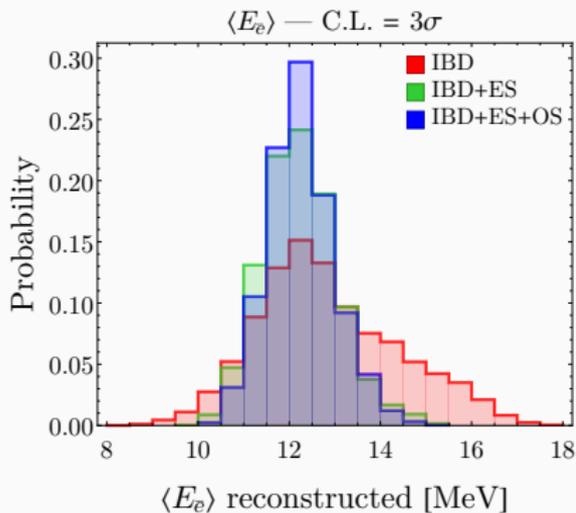
Results

IBD+ES+OS, 10 param., no equipartition, 3σ C.L.



- Param. undetermined for ν_e
- Pinching undetermined for all

Results



$\bar{\nu}_e$	$\langle E \rangle$ [MeV]	Acc. %
IBD	12.9 ± 1.5	12
IBD+ES	12.2 ± 0.8	7
IBD+ES+OS	12.3 ± 0.7	6

ν_x	$\langle E \rangle$ [MeV]	Acc. %
IBD	13.0 ± 2.6	20
IBD+ES	14.3 ± 2.0	14
IBD+ES+OS	14.6 ± 1.9	13

Total energy reconstruction

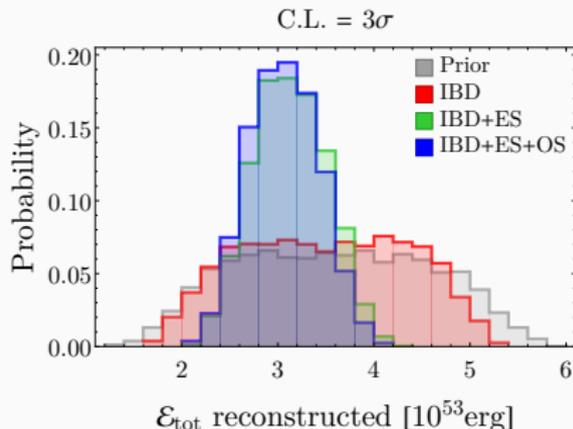
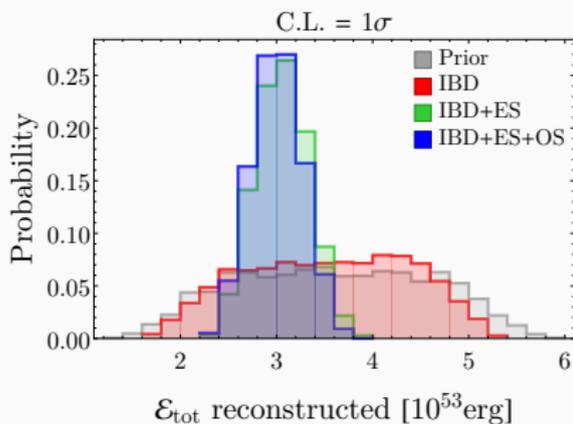
- Point P within a certain C.L.
- $\mathcal{E}_{tot}|_P = \mathcal{E}_{e,P} + \mathcal{E}_{\bar{e},P} + 4\mathcal{E}_{x,P}$

Warning

For IBD only $\mathcal{E}_{e,P}$ not measured

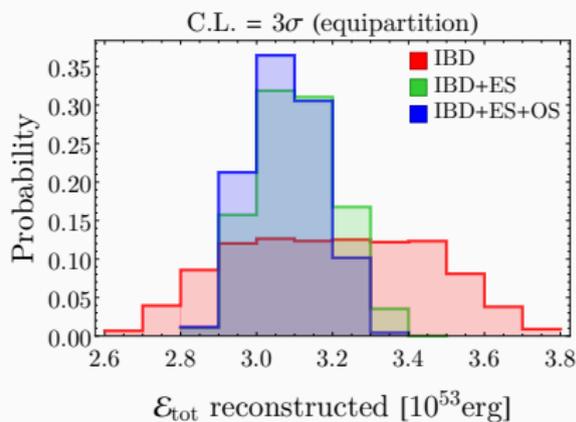
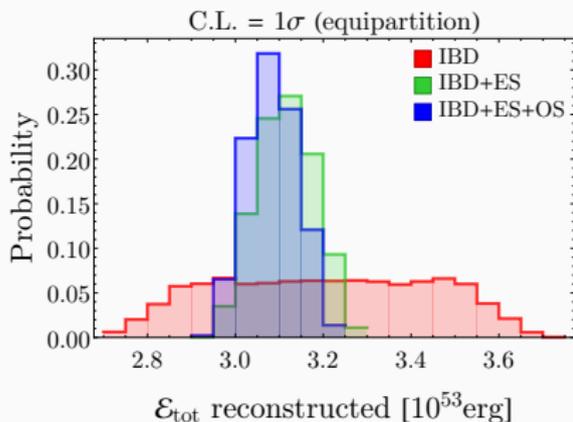
\Rightarrow random uniform in prior $[0.2, 1] \times 10^{53}$ erg

Results: total energy



	1σ C.L.		3σ C.L.	
	\mathcal{E} [10^{53} erg]	$\Delta\mathcal{E}/\mathcal{E}$	\mathcal{E} [10^{53} erg]	$\Delta\mathcal{E}/\mathcal{E}$
IBD	3.53 ± 0.84	24%	3.50 ± 0.85	24%
IBD+ES	3.05 ± 0.28	9%	3.13 ± 0.38	12%
IBD+ES+OS	3.00 ± 0.26	9%	3.06 ± 0.36	12%

Results: total energy



Equipartition	1σ C.L.		3σ C.L.	
	\mathcal{E} [10^{53} erg]	$\Delta\mathcal{E}/\mathcal{E}$	\mathcal{E} [10^{53} erg]	$\Delta\mathcal{E}/\mathcal{E}$
IBD	3.20 ± 0.24	7%	3.20 ± 0.25	8%
IBD+ES	3.12 ± 0.07	2%	3.11 ± 0.10	3%
IBD+ES+OS	3.08 ± 0.06	2%	3.08 ± 0.09	3%

Summary

Summary

- Ansätze of crucial importance
- Parameters
 - ✗ If pinching is unknown it remains unknown
 - ✗ ν_e properties undetermined
 - ✓ $\langle E_{\bar{\nu}_e} \rangle, \langle E_x \rangle$ known within $\sim 10\%$ @ 3σ C.L. (IBD+ES+OS)
- Total energy
 - ✓ Known within $\sim 10\%$ @ 3σ C.L. (IBD+ES+OS)
 - ✓ 1% goal achievable if equipartition holds
- Next: Hyper-K, JUNO, ...