



The role of the diffusive protons in the gamma-ray emission of SNR RX J1713.7-3946

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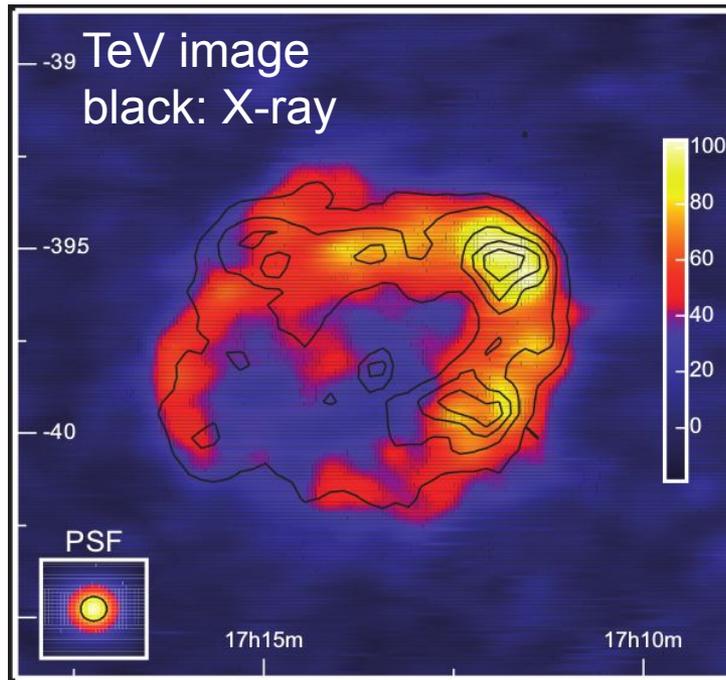
24/02/2017 Reunion, France



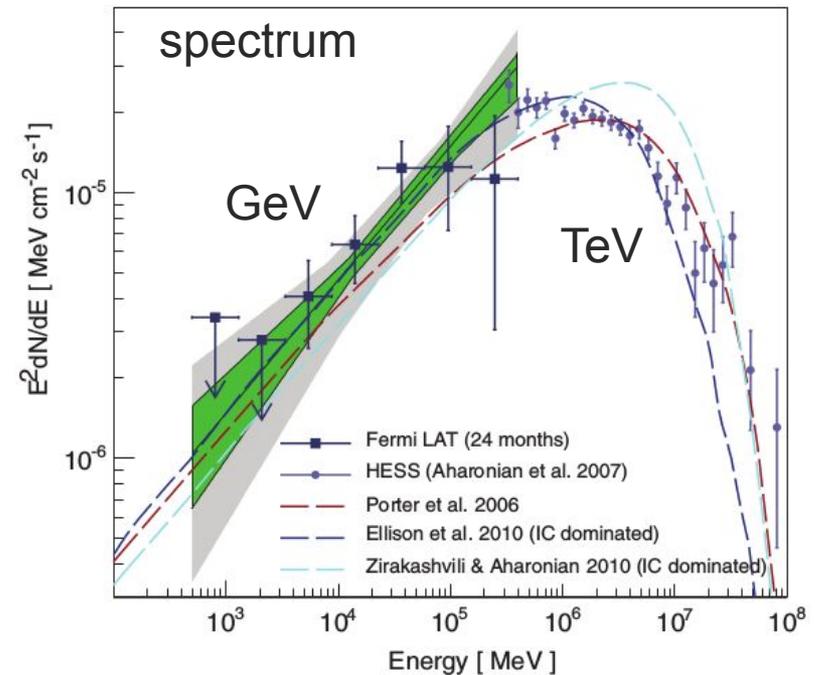
GeV – TeV observations

- TeV image matches X-ray shell
- hard photon index $\Gamma \sim 1.5$

→ purely leptonic origin !



Aharonian et al. 2007



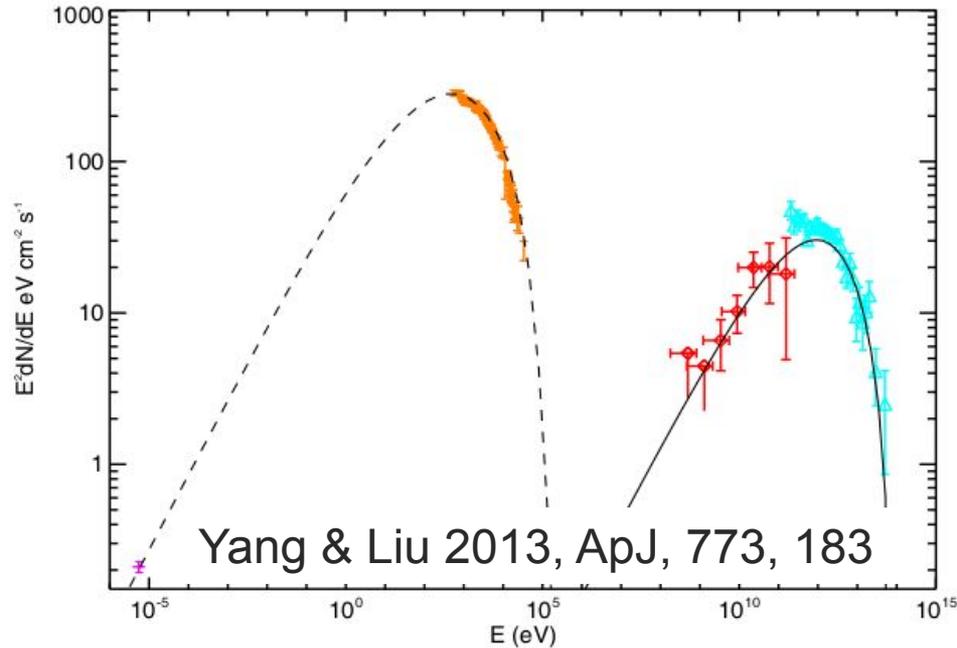
Abodo et al. 2011



Leptonic model and its difficulty



$$\frac{dN}{dE} \propto E^{-\alpha} \exp[-(E/E_{cut})^\delta]$$

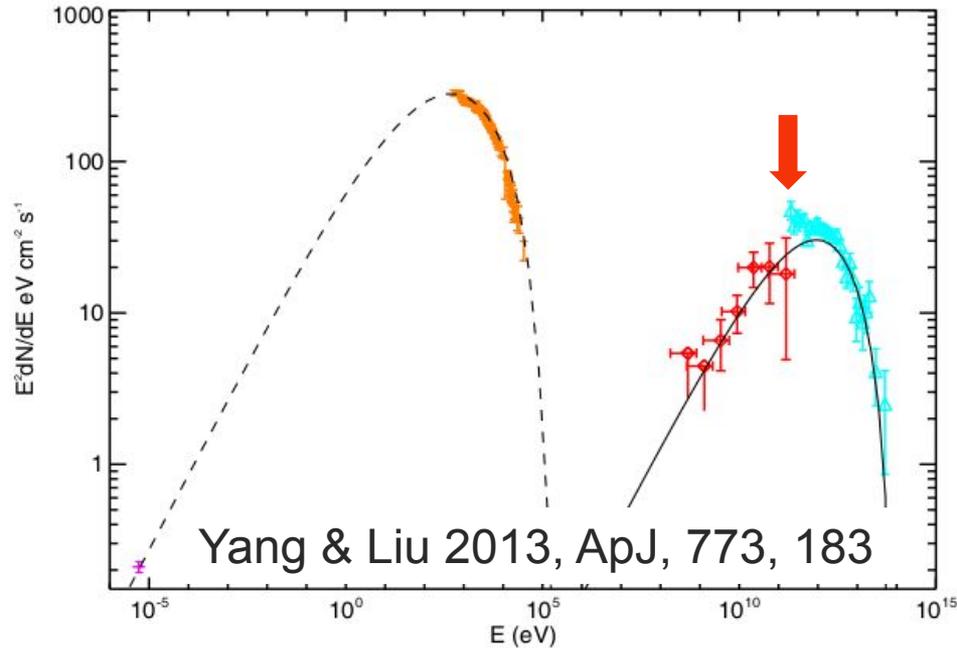




Leptonic model and its difficulty



$$\frac{dN}{dE} \propto E^{-\alpha} \exp[-(E/E_{cut})^\delta]$$



Radio + Xray



Super exponential: $\delta > 1.0$

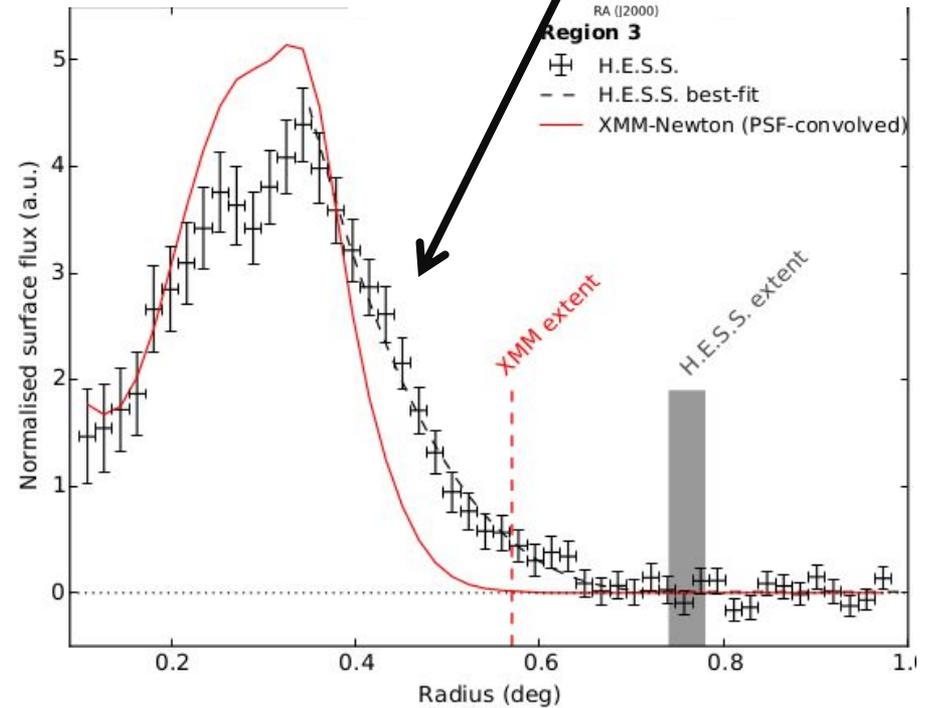
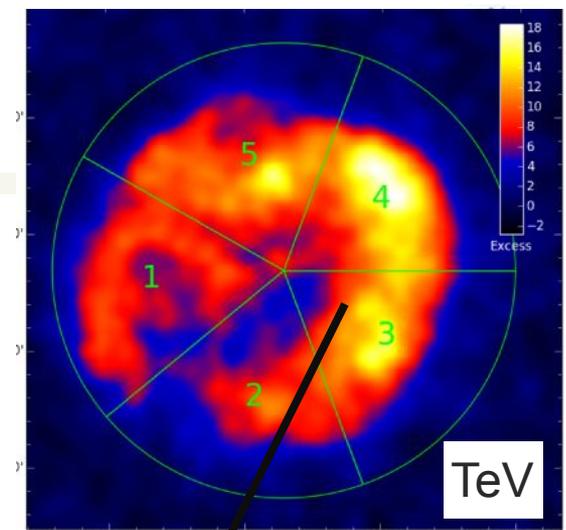
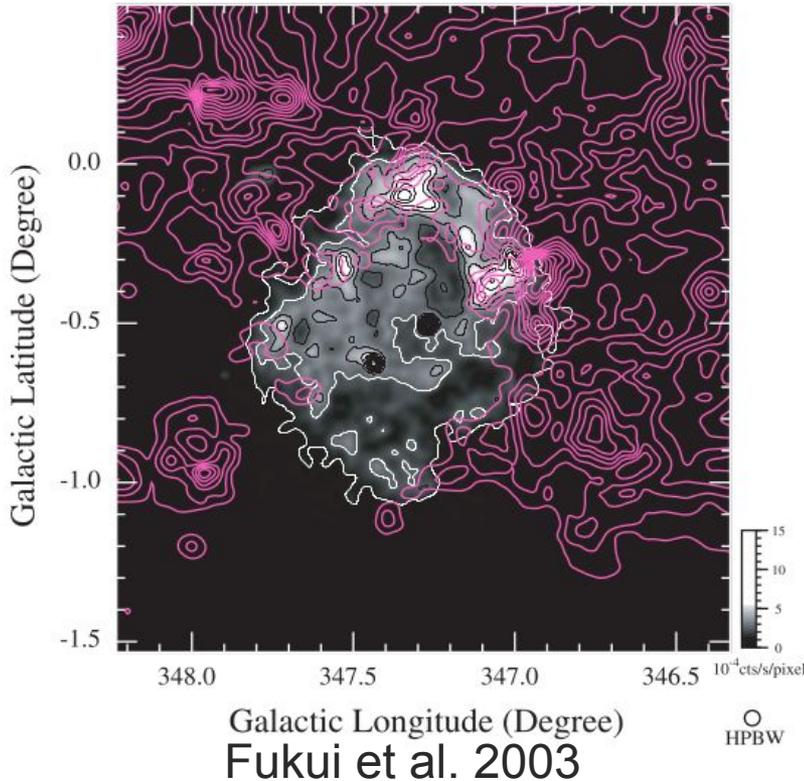
GeV + TeV



Sub exponential: $\delta \sim 0.5$

Hadronic component ?

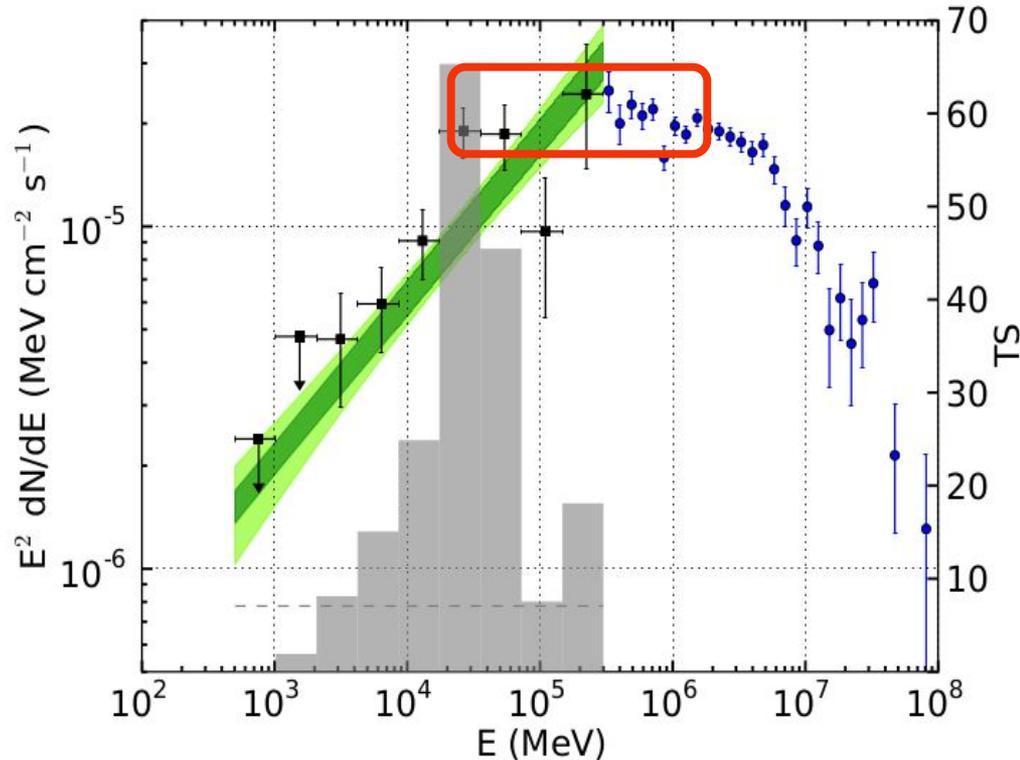
- CO: located in a MC cavity
- TeV radial profile beyonds X-ray's



de Naurois 2015



The spectral break/platform?



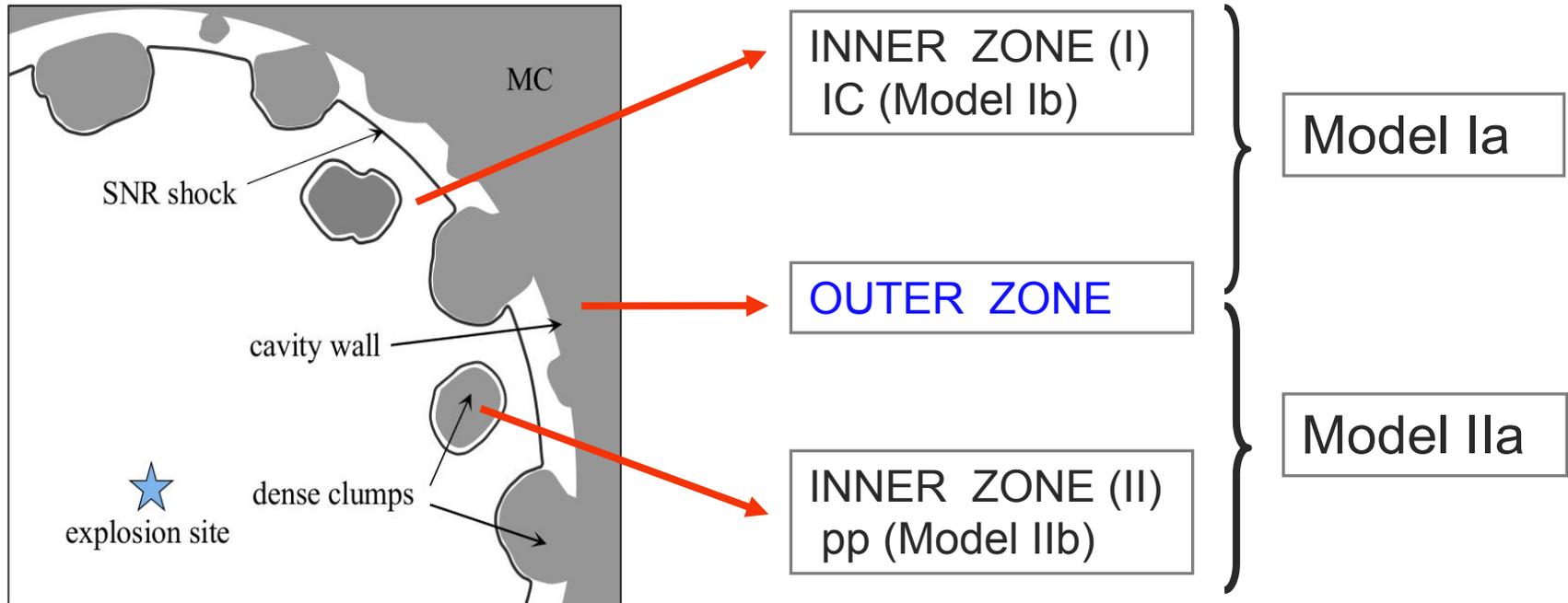
Federic et al. 2015 A&A

not a simple power law plus cutoff ?

Two-zone model



- Due to $V_{sh} \sim 4000$ km/s and $n \sim 0.01$ cm⁻³, we assume that the shock doesn't extensively hit upon the MC but may sweep up some dense clumps.





Particle distributions



- INNER ZONE (I) pow law + cutoff
- OUTER ZONE

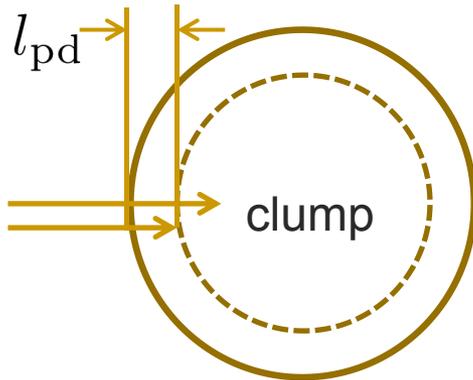
$$\frac{\partial}{\partial t} f(E_p, R, t) = \frac{D(E_p)}{R^2} \frac{\partial}{\partial R} R^2 \frac{\partial}{\partial R} f(E_p, R, t) + Q(E_p, R, t).$$

$$Q(E, R, t) = Q_0 E^{-\alpha} \delta(R - R_{\text{esc}}(t)) / 4\pi R^2$$

- INNER ZONE (II)

The penetration depth (Inoue et al. 2012) :

$$l_{\text{pd}} \approx (\kappa_d t)^{1/2} \propto E^{1/2}$$



$$N_\gamma(E) dE \propto M(E) E^{-p} dE$$

$$\propto l_{\text{pd}}(E) E^{-p} dE$$

$$\propto E^{-p+1/2} dE$$

$$N_\gamma(E) \propto E^{-1.5} \text{ for } p = 2$$

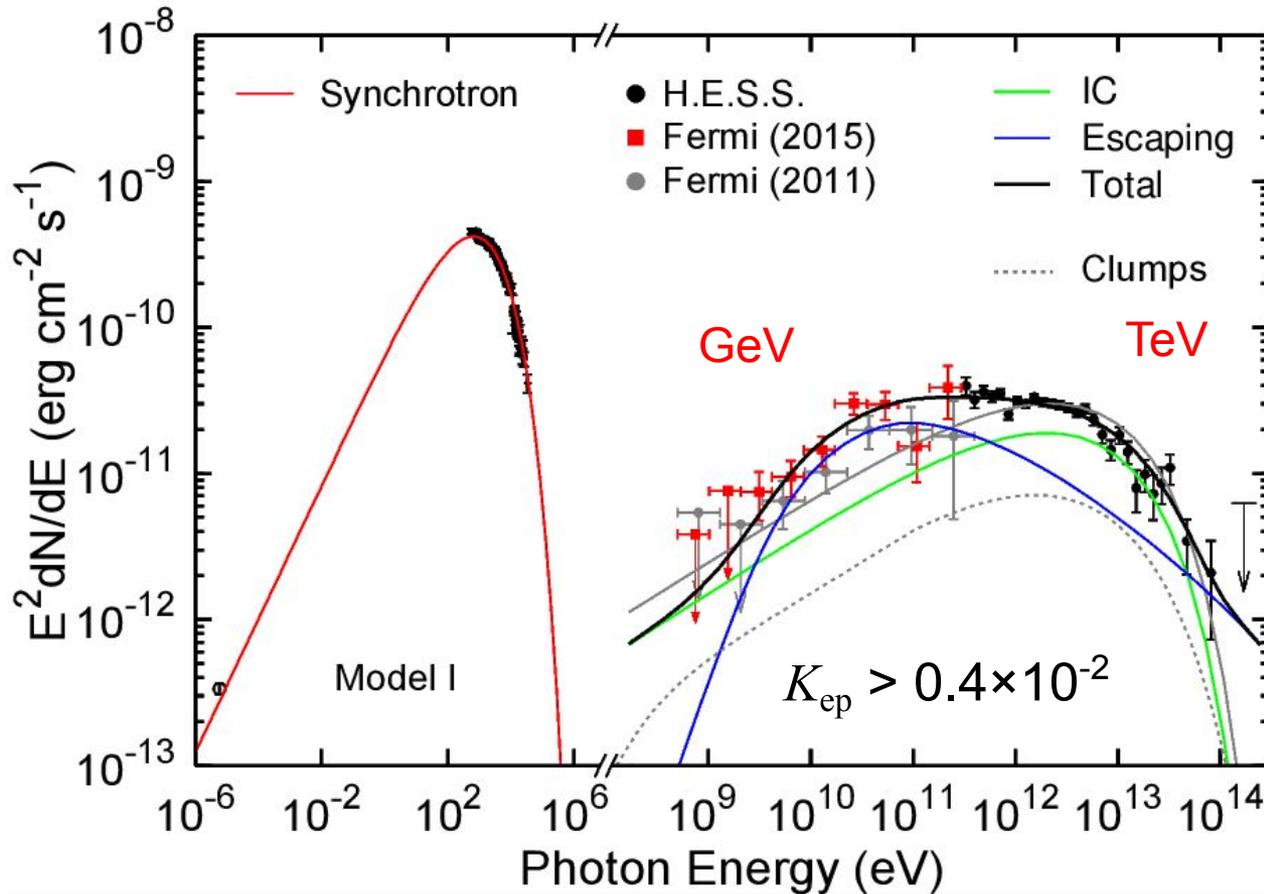
The mass of the shock clumps can be found in Sano et al. 2013



Model I



- The MCMC method is employed to constrain the parameters.
- The electron to proton number ratio: $K_{ep} = 0.01$



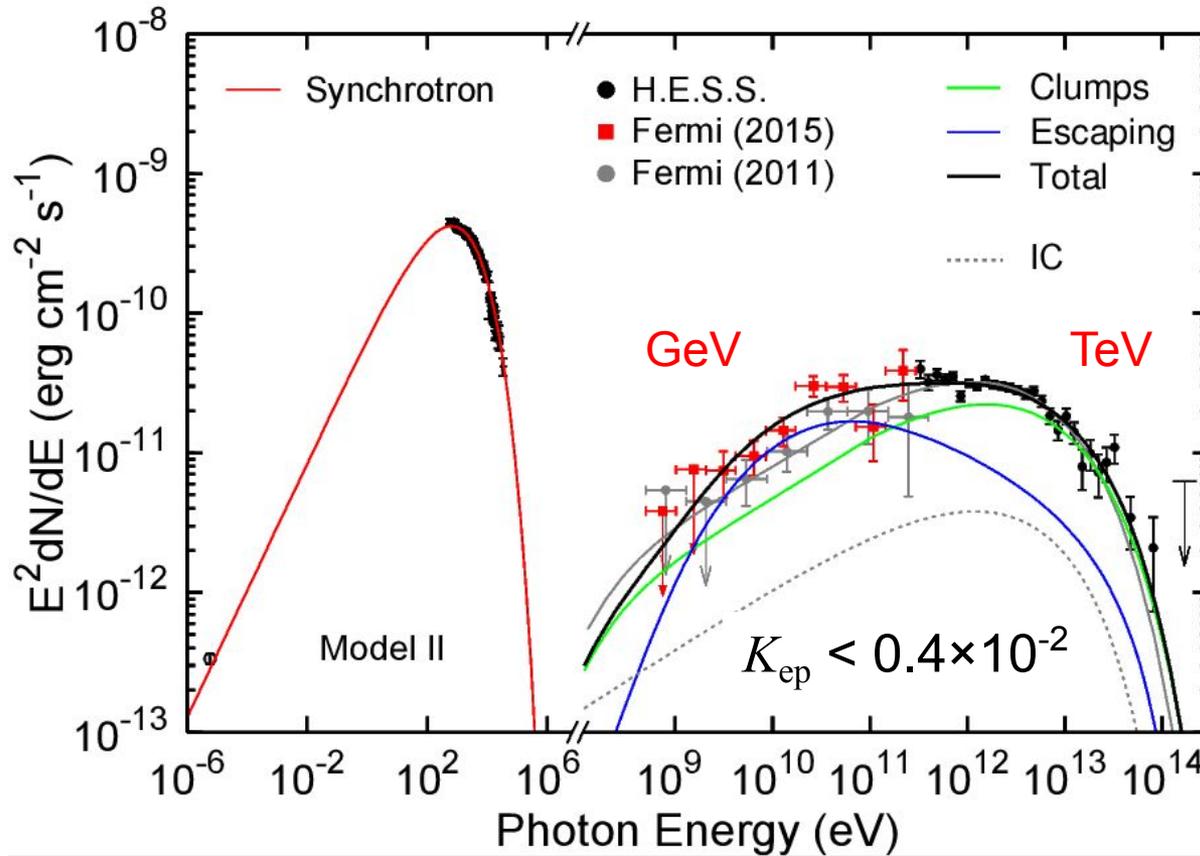
Zhang & Chen 2016



Model II



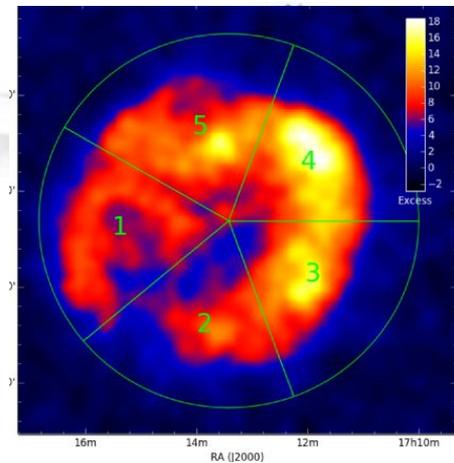
- The electron to proton number ratio: $K_{ep} = 0.001$



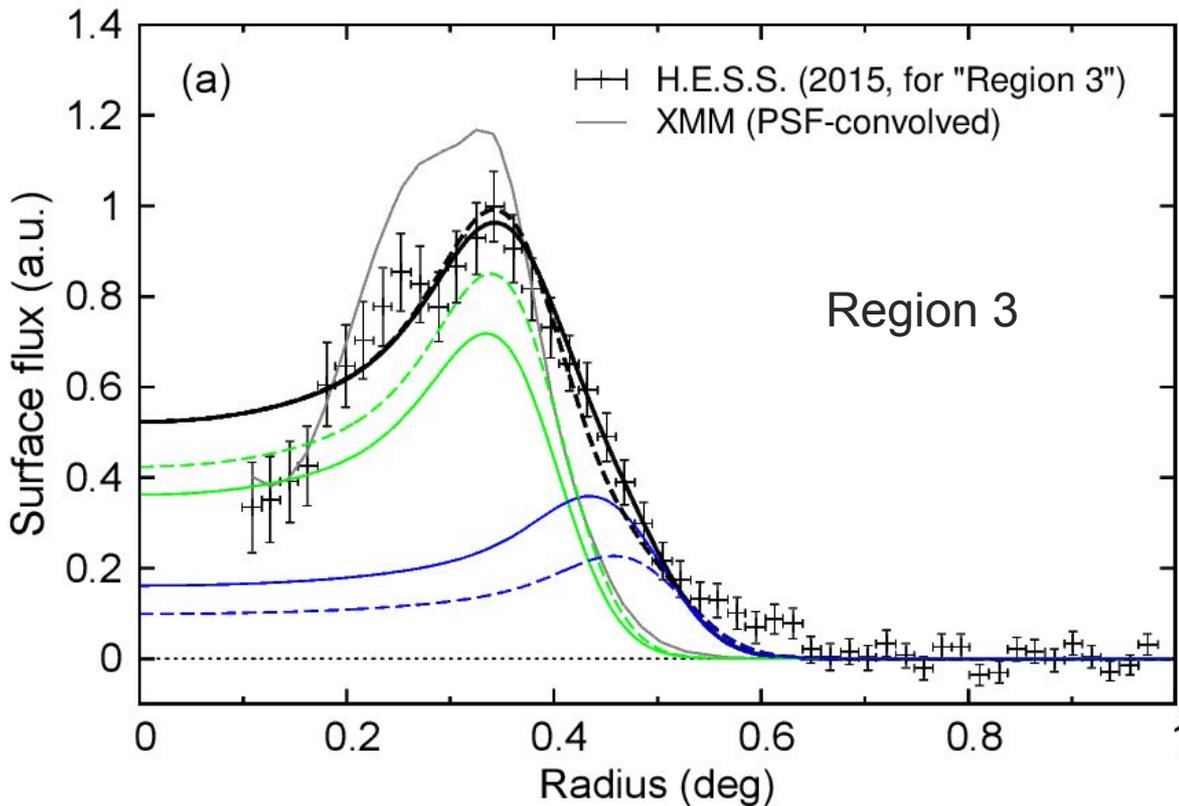
Zhang & Chen 2016



The radial profile at the TeV band



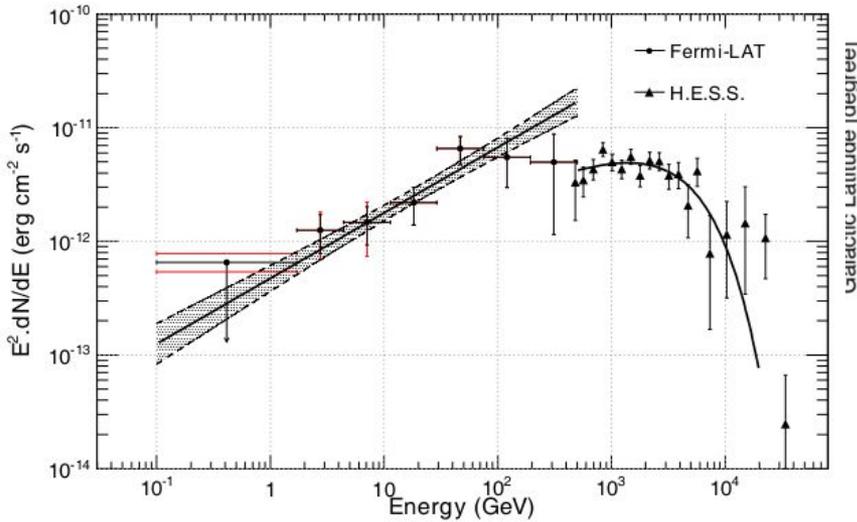
$$J(R) = 2 \int_{r_{1,\min}}^{r_{1,\max}} \frac{\sigma_1 r dr}{\sqrt{r^2 - R^2}} + 2 \int_{r_{2,\min}}^{r_{2,\max}} \frac{\sigma_2 r dr}{\sqrt{r^2 - R^2}}$$



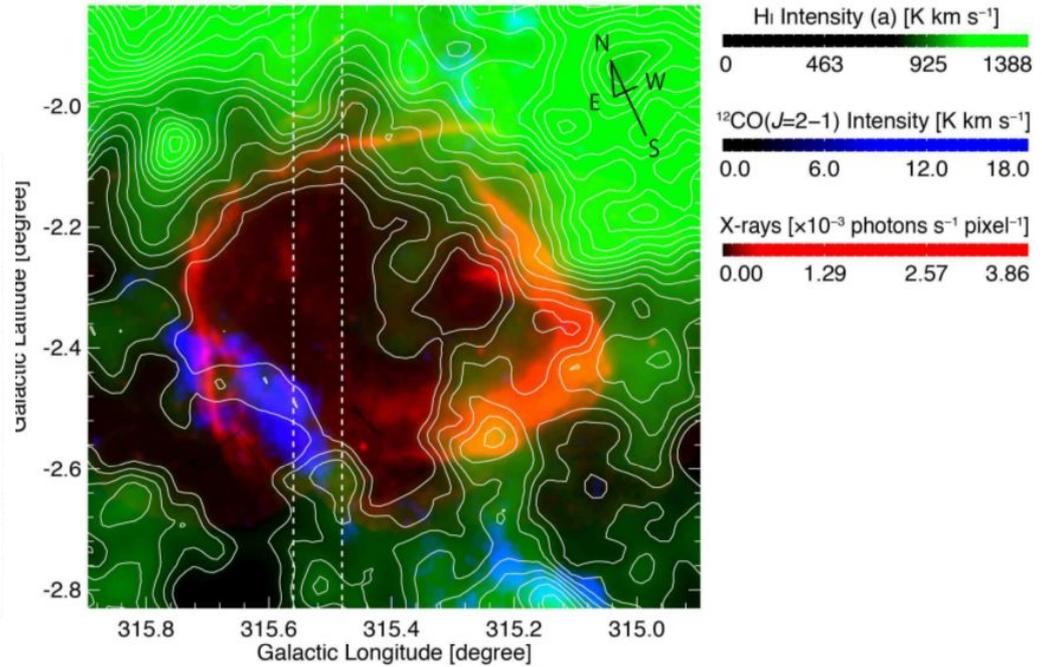
solid: Model Ia
dashed: Model IIa
green: inner
blue: outer



RCW 86: another case ?



Ajello et al. 2016



Sano et al. 2016



Summary



- ✓ The broadband fluxes can be well explained by the two-zone model, in which the gamma-ray emission from inside governs the TeV band, while the outer emission component substantially contributes to the GeV gamma-rays.
- ✓ The two-zone model can simultaneously reproduce the TeV gamma-ray radial brightness profile that significantly extends outside the nonthermal X-ray-emitting region.
- ✓ For the inner region, dominated by the IC process for $K_{ep} > 4 \times 10^{-3}$ and by pp in the shocked clumps for $K_{ep} < 4 \times 10^{-3}$.

Thanks!